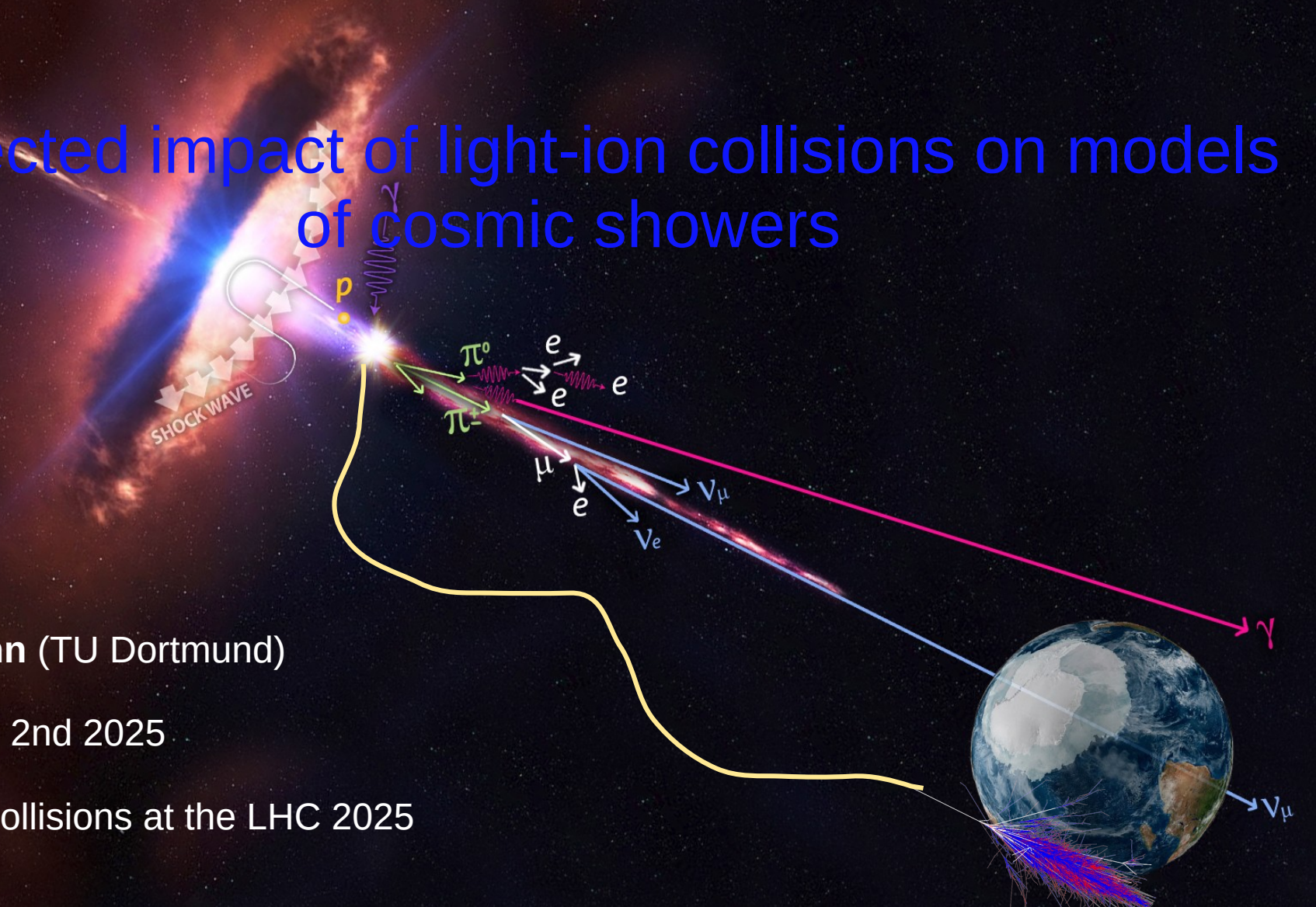


# Expected impact of light-ion collisions on models of cosmic showers



Felix Riehn (TU Dortmund)

December 2nd 2025

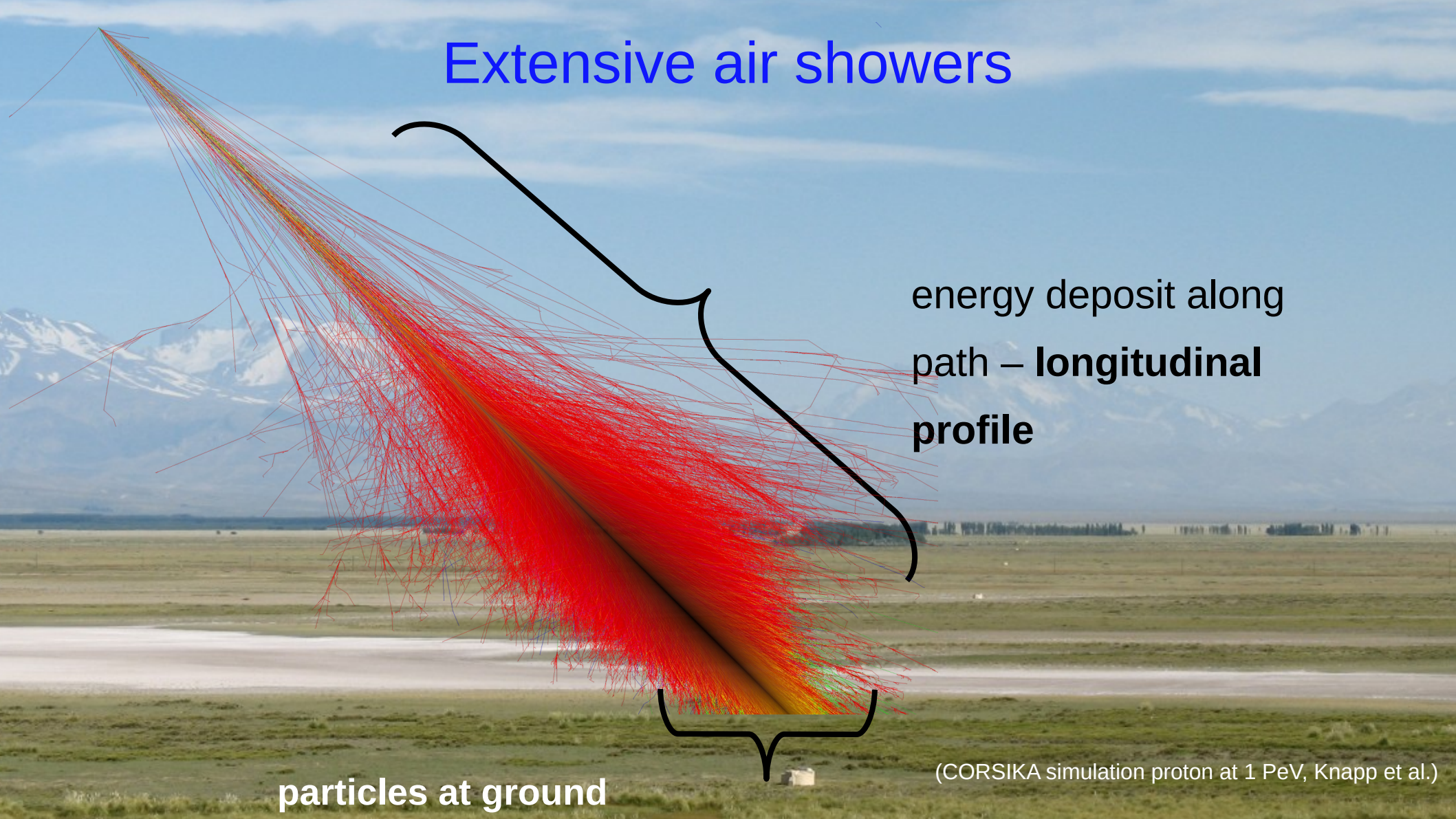
Light-ion collisions at the LHC 2025

# Extensive air showers

energy deposit along path – **longitudinal profile**

**particles at ground**

(CORSIKA simulation proton at 1 PeV, Knapp et al.)



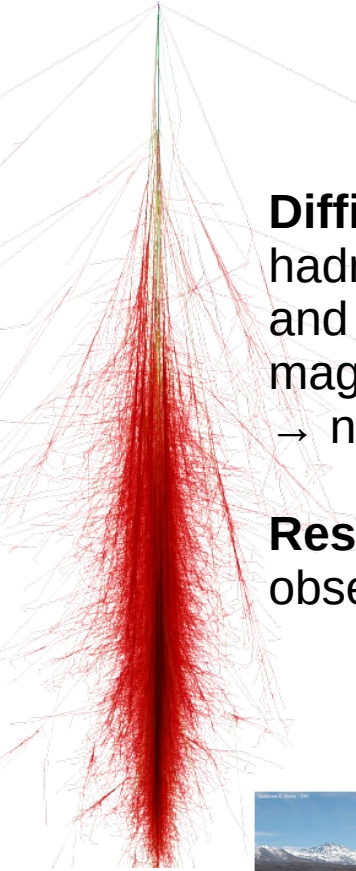
# The problem: air showers not correctly modeled

## Problem I: modeling individual air showers

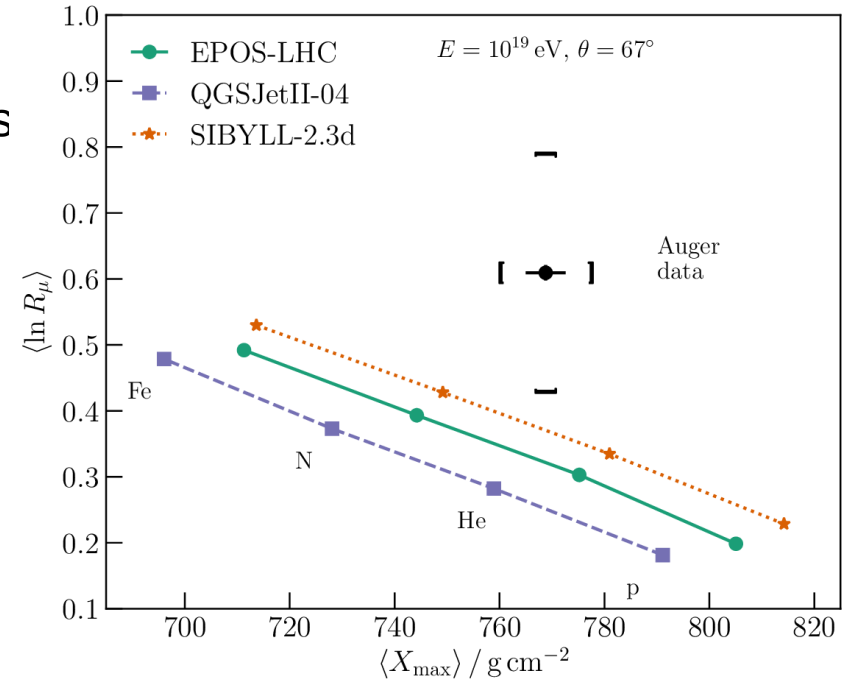
**Difficulty:** Need a fully inclusive model of hadronic interactions of pions, kaons, protons and nuclei with nuclei over 10 orders of magnitude in energy !

→ non-perturbative QCD

**Result:** inconsistent results between observables, in-particular ground based

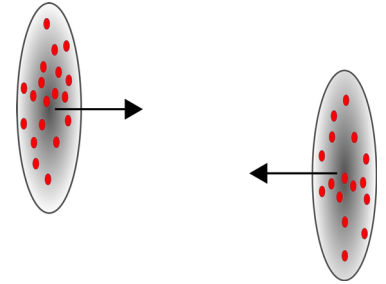
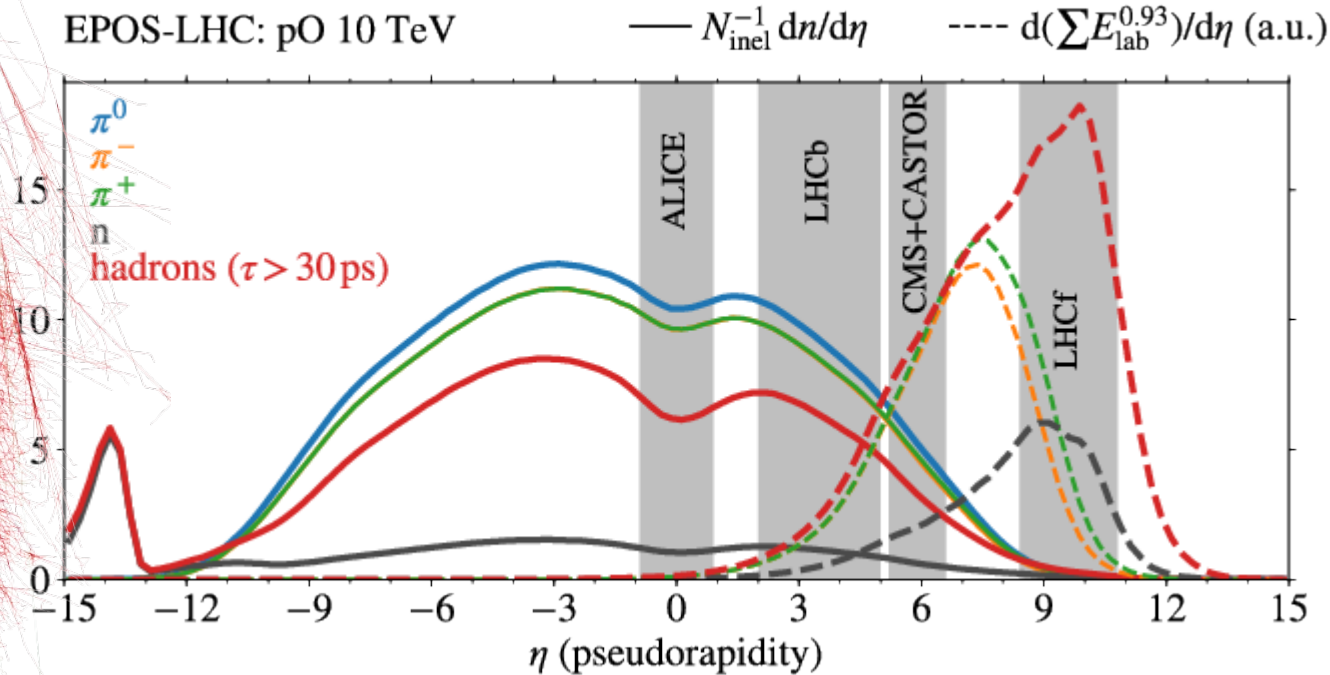


(Auger, PRL 126, 152002 (2021))



# Phase space considerations

(Dembinski et al, Astrophys.Space Sci. 367 (2022) 3, 27)



$$\Delta x \Delta p_x \sim 1$$

$$R' = R/\Gamma = R m_p / E_p$$

$$p_{\perp} \sim \Delta p_{\perp} \sim \frac{1}{R} \sim 200 \text{ MeV}$$

$$p_{\parallel} \sim \Delta p_{\parallel} \sim \frac{1}{R'} \sim \frac{1}{5} E_p$$

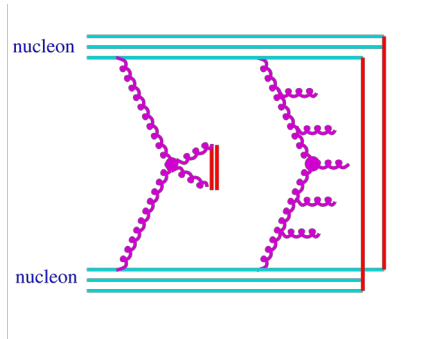
Air shower == chaining average interactions, weighted by energy  
 ==> pT is negligible, **forward particle production**

# CR models

**SIBYLL** (R. Engel, A. Fedynitch, T.K Gaisser, FR, T. Stanev)

- DPM based (related to DPMJET)
- factorization between hard&soft MPI
- charm production
- phenomenologic
- light ions only ( $A < 18$ )

*"bare bones of QCD"*



**EPOS LHC** (T. Pierog, K. Werner)

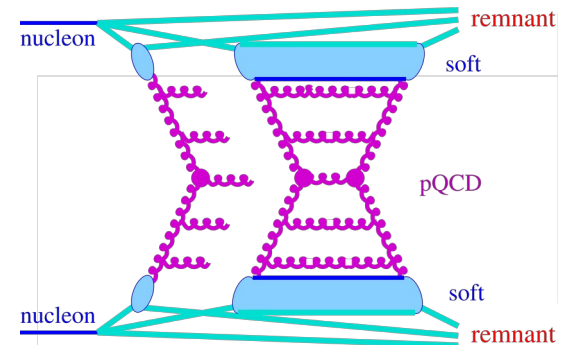
- Gribov-Regge
- energy sharing at amplitude level
- effective collective effects

*"heavy-ion model"*

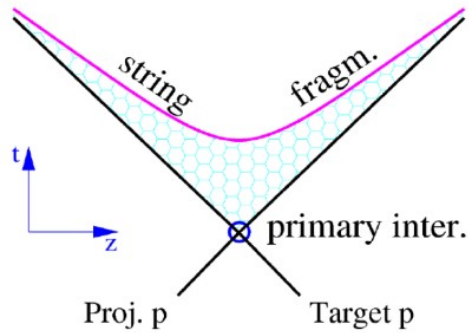
**QGSJET III/III** (S. Ostapchenko)

- Gribov-Regge
- non-linear effects
- higher twist

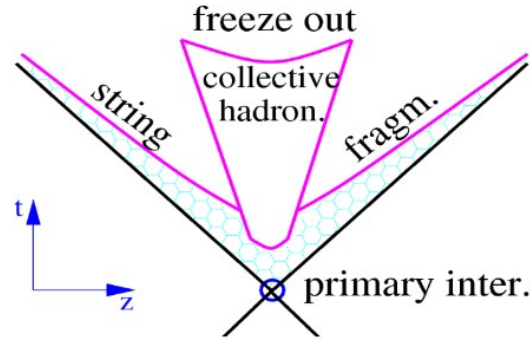
*"rigorous model"*



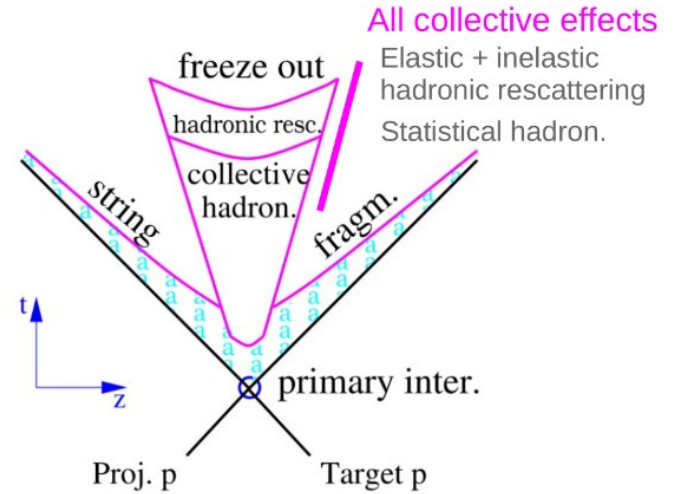
# CR models space-time picture



Sibyll, QGSJET



EPOS LHC  
(no rescattering)

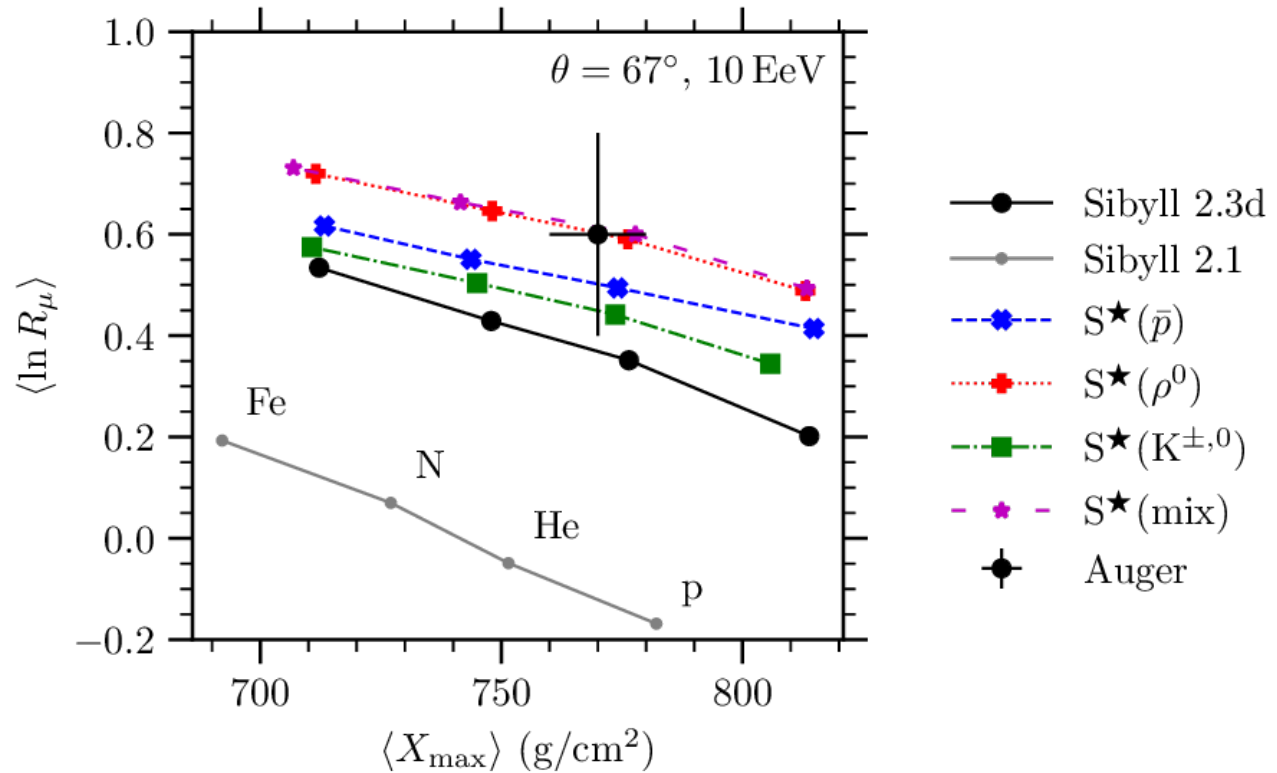


EPOS.LHC-R  
(with hadronic rescattering)

(See T.Pierog ICRC2025)

# Possible solutions

(Astropart.Phys. 160 (2024) 102964)

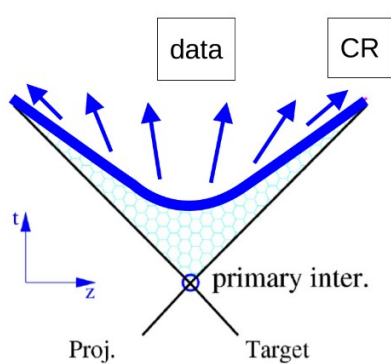


- \* more strangeness
- \* more baryons
- \* more resonances in pion interactions

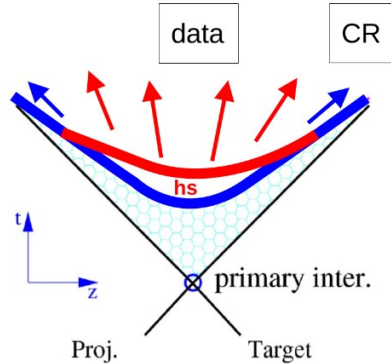
# The EPOS LHC R solution

- \* increased **baryon** and **strangeness** production from collective hadronisation
- \* change ratio between pi and rho

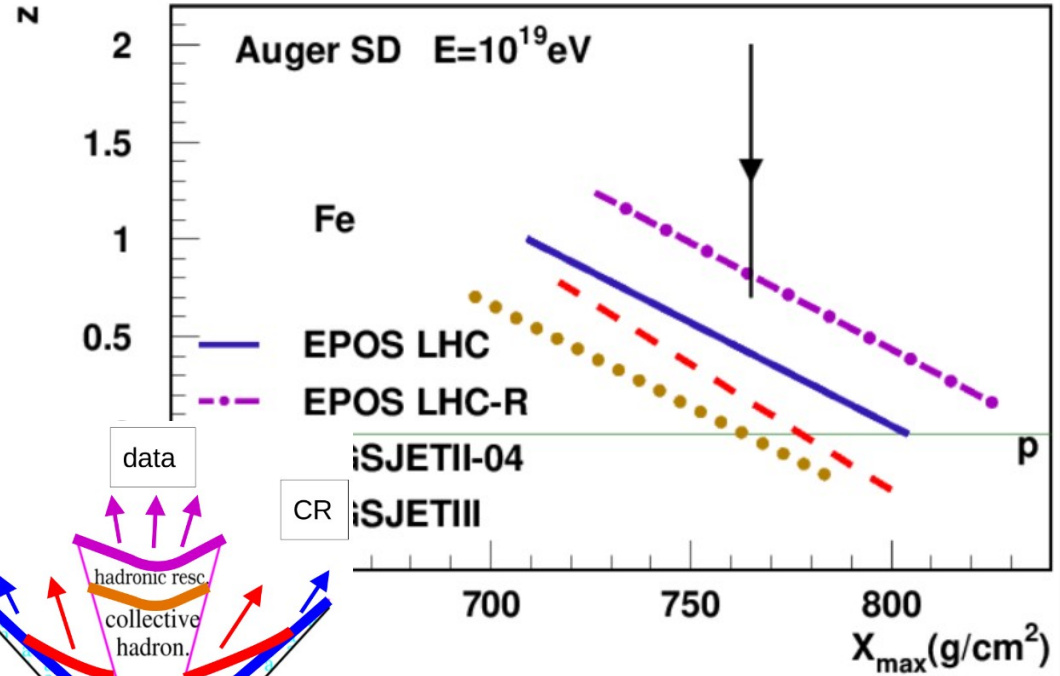
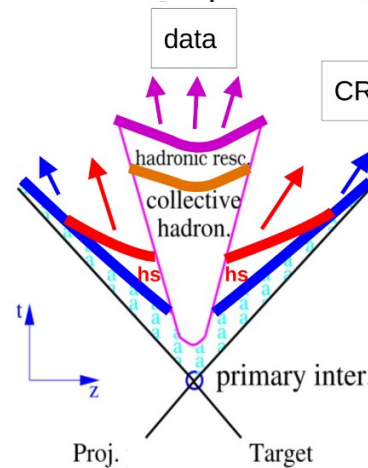
Changes arise from re-tuning with all effects (QGP, had. rescatter) included when tuning (e+e-,hp,pp,pA,AA)



No collective effects



Hadronic reScattering (hs)

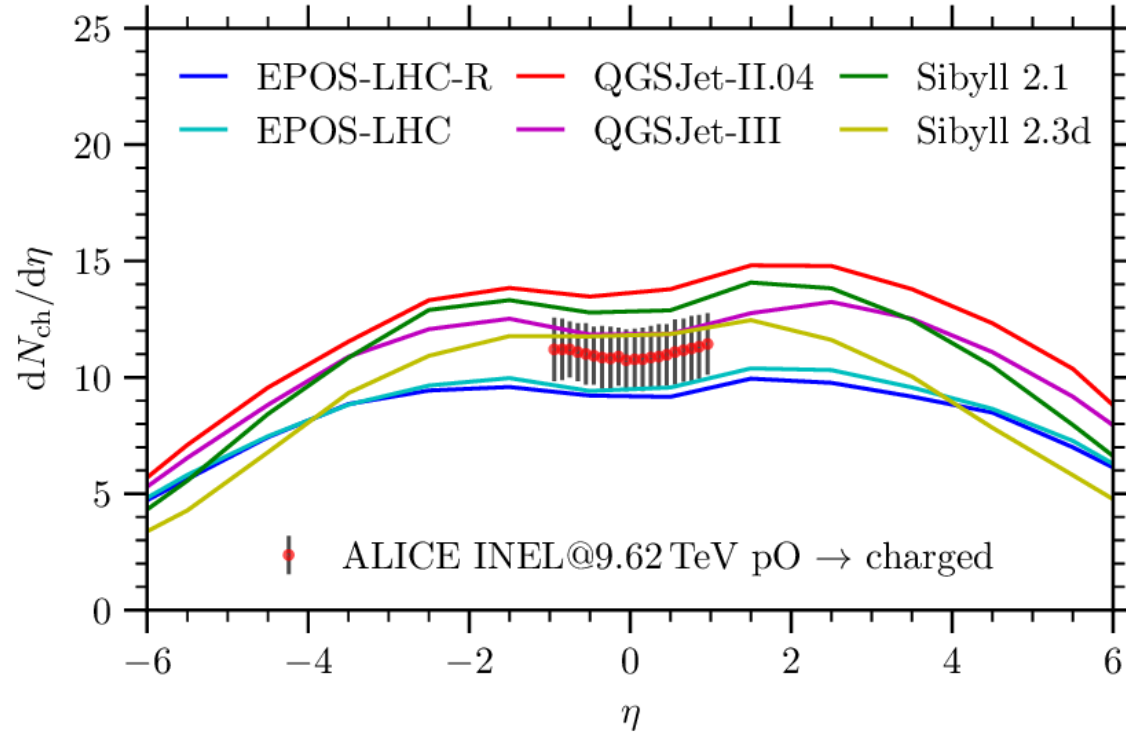


(T. Pierog ICRC 2025)

# Expected impact of light-ion collisions at the LHC on models of cosmic showers:

*Insights on which model approach is correct*

# First pO results



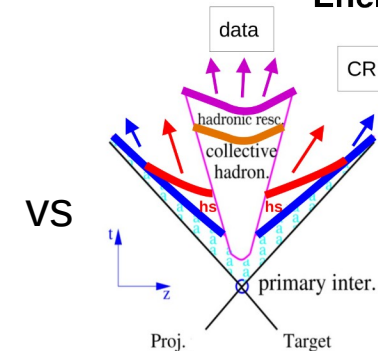
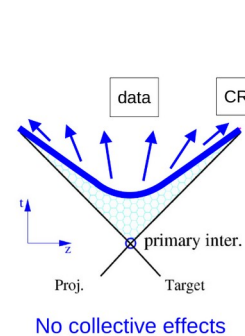
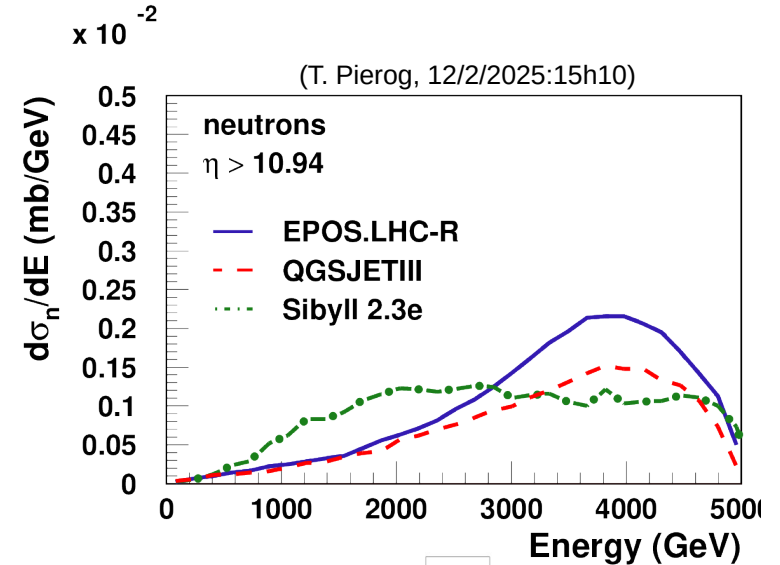
\* **all** models make reasonable predictions for central particle production

(ALICE, initial stages 2025)

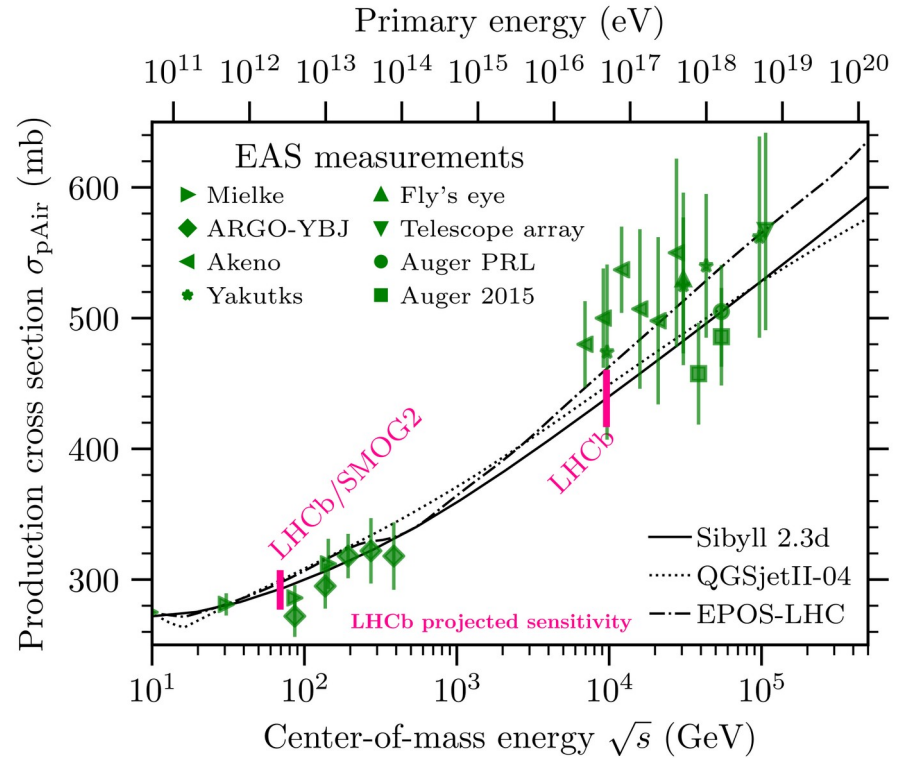
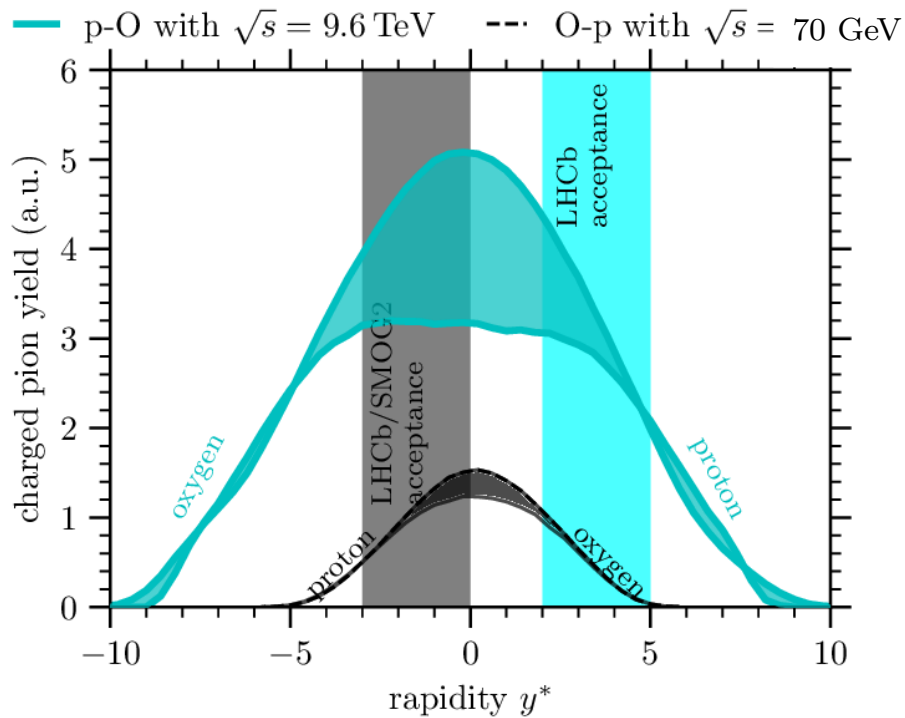
# Measurement wish-list for light-ion

*In absence of identified particle spectra at  $\eta > 7$*

- LHCf neutrons in pO  
→ constrain elasticity
- central baryon & strangeness production
- measurements of central activity vs forward energy (ATLAS+LHCf, ALICE w ZDC?, CMS?)



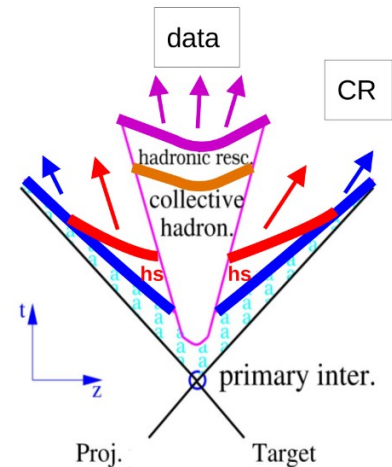
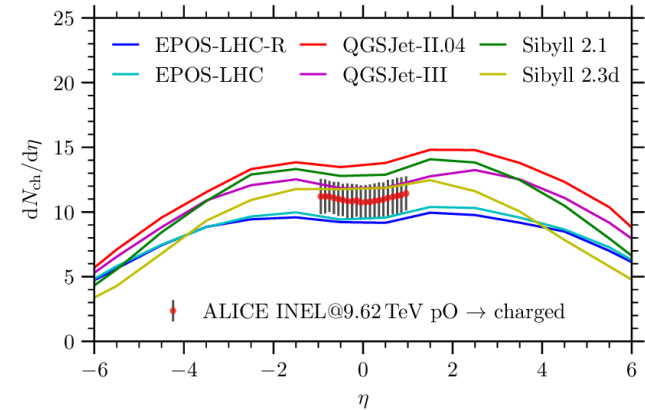
# Future measurements with LHCb



+ scan A dependence for light-ions w SMOG2!

# Summary

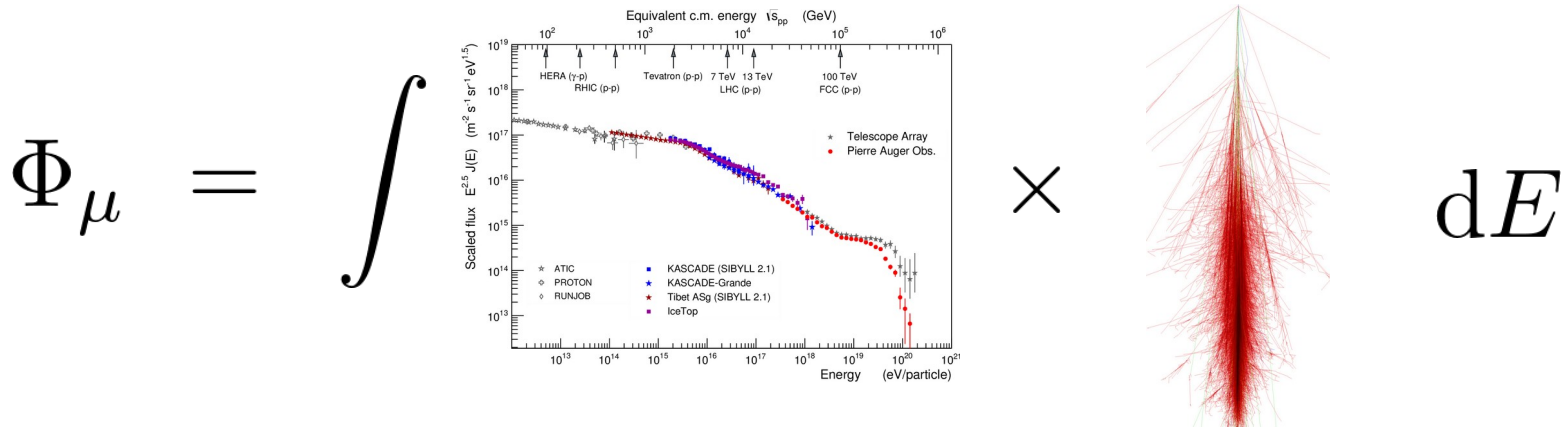
- \* cosmic showers sensitive to very forward particle production
- \* first pO results of  $dN_{ch}/d\eta$  in central region indicate all CR models close to the measurements
- \* QGP and HS have indirect effect on the cosmic shower predictions
- \* awaiting decisive measurements
  - LHCf neutrons
  - baryon & strangeness production
  - central & forward correlation





# The problem: air showers not correctly modeled

## Problem II: modeling all air showers



Flux of muons and neutrinos in atmosphere due to meson decay

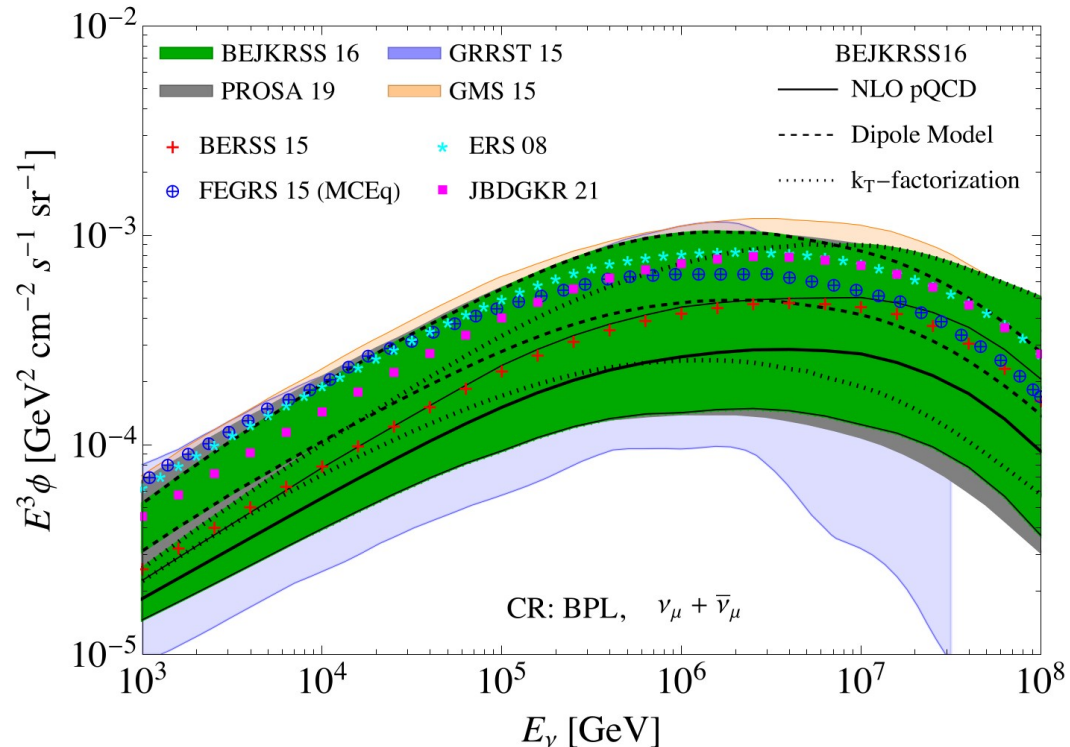
==> compounded problem. Particle production spectra in EAS are weighted by primary spectrum  $E^{**2...3}$   
 + sensitive to rare particles

# The problem: air showers not correctly modeled

## Problem II: modeling all air showers

Broad range of predictions.

Large uncertainty of atmospheric flux at high energies



# CR models

SIBYLL  
h+A(<18)

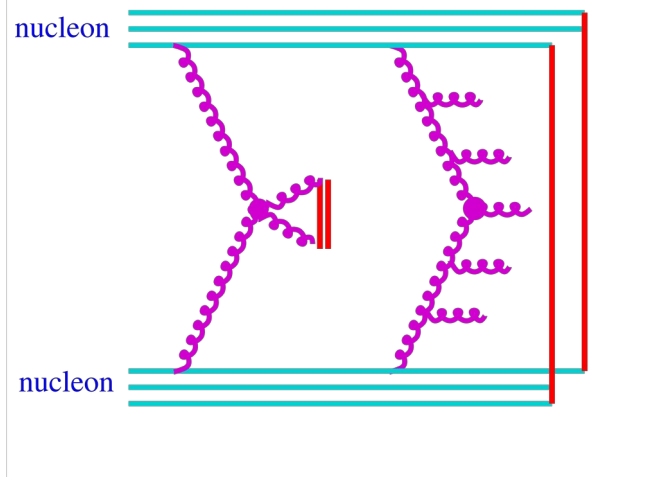
QGSJET II (h+A,A+A)

EPOS (e+e-, h+A, A+A)



Complexity, runtime, #parameters, #data

SIBYLL



EPOS, QGSJET II

