## Design Principles for Muon Colliders

J. Scott Berg Brookhaven National Laboratory Topical Meeting on the Neutrino Factory and the Muon Collider 22 October 2007



#### **Muon Collider Goals**

- OReach high energy
- Achieve high luminosity
- Avoid excessive neutrino radiation
  - Neutrinos create showers in massive objects near site
  - Highly concentrated neutrino beam
     Increases rapidly with energy









- Beam isn't stored long (decays)
  - Muon production rate determines luminosity
     Similar to linear collider
- Beams can collide multiple times
  - Not arbitrarily large number (decays)
     Here we beat linear colliders
  - Advantages to throwing out your beam
- Long-distance neutrino radiation





### Muon Collider Challenges

- Muons produced with large emittances
   Requires massive amounts of cooling
   Muons decay
  - Preserving beam all the way to collider
     Cooling and acceleration must be extremely rapid







# **Muon Collider Components**

- High-power proton driver
- Target
- Make the beam sane
- Capture, phase rotation, bunching
   Could be something else...
   Cooling





## **Muon Collider Components**

Rebunching: several to one
During cooling
During acceleration
Not at all...
Acceleration

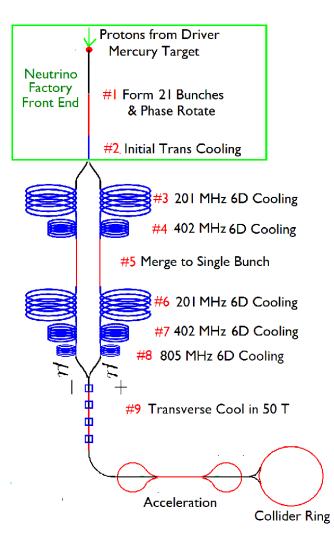
○ Collider ring







## **Muon Collider System**







## Luminosity Luminosity for Decaying Beams



 $\odot$  Round Gaussian beams collide, all  $\mu$ s decay

$$\mathcal{L} = \frac{\gamma \tau_{\mu}}{2T_{\text{rev}}} \frac{N^2 n_B f_P}{4\pi\sigma^2} = \frac{\tau_{\mu} e B_{\text{avg}} N^2 n_B f_P}{4\pi\sigma^2}$$

 $\circ N$ : particles per bunch  $\circ \gamma$ : energy/(rest mass)

- $\circ \tau_{\mu}$ : muon rest lifetime
- $\circ T_{rev}$ : revolution time
- $\circ f_P$ : driver rep. rate
- e: electron charge

- $\circ \sigma$ : RMS size
- $\circ n_B$ : no. of bunches
- $\circ B_{avg}$ : avg. ring field
- $\circ m_{\mu}$ : Muon mass



## Luminosity Multiple-Crossing Factor



$$\mathcal{L} = \frac{\gamma \tau_{\mu}}{2T_{\text{rev}}} \frac{N^2 n_B f_P}{4\pi\sigma^2} = \frac{\tau_{\mu} e B_{\text{avg}} N^2 n_B f_P}{4\pi m_{\mu}} \frac{1}{4\pi\sigma^2}$$

First factor is the average number of crossings

- Always choose the highest possible magnet fields in the collider ring
  - Everything else is much harder







#### **Beam-Beam Tune Shift**

- $\odot$  Maximize luminosity by running at maximum beam-beam tune shift ( $\Delta \nu$ )
- $\odot\, {\rm We}$  only have around 1000 turns, potentially allowing large  $\Delta \nu$

$$\Delta \nu = \frac{\beta^* N r_{\mu}}{4\pi\sigma^2\gamma} = \frac{N r_{\mu}}{4\pi\epsilon_n}$$

 $\circ \beta^*$ : C-S beta at IP  $\circ r_{\mu}$ : classical  $\mu$  radius

 $\circ \epsilon_n$ : normalized transverse emittance







#### **Proton Driver Power**

Power is product of
Particles per bunch
Bunches per cycle
Cycle frequency
Energy of beam

 $\odot$  Energy transferred to muons with efficiency  $\eta_{cap}$ 

 $\circ$  Average muon energy at capture:  $E_{cap}$ 





#### **Proton Driver Power**

$$P_P = \frac{E_{cap} N n_B f_P}{\eta_{cap} \ \eta_{trans}}$$

- Fraction of captured muons making it to collider:  $\eta_{\text{trans}}$
- $\circ \eta_{cap} \eta_{trans} / E_{cap}$  is the physical quantity  $\circ \eta_{trans}$  depends on where "capture" ends  $\circ \eta_{cap} / E_{cap}$  depends only on target/capture system





#### **Neutrino Radiation**

- Increases strongly with energy
- Proportional to muons per second times average turns

$$C_{\mathsf{rad}} = rac{ au_{\mu} e B_{\mathsf{avg}}}{2\pi m_{\mu}} N n_b f_P$$

Make ring deeper to reduce radiation
 Reduction factor proportional to ring depth
 Straights create strong radiation



## Luminosity in Terms of Other Quantities



OUse beam-beam tune shift and proton power:

$$\mathcal{L} = \frac{\tau_{\mu} e B_{\text{avg}} \eta_{\text{trans}} P_P \Delta \nu}{2\pi m_{\mu}} \frac{\eta_{\text{cap}} \gamma}{\beta^*} \frac{\beta_{\text{cap}} \gamma}{E_{\text{cap}} r_{\mu}}$$

Improve luminosity with

Higher B<sub>avg</sub>
 Reducing losses getting to collider
 Increasing proton driver power
 Running with larger beam-beam tune shift
 Lower beta at IP





# Improving Luminosity

$$\mathcal{L} = \frac{\tau_{\mu} e B_{\text{avg}} \eta_{\text{trans}} P_P \Delta \nu}{2\pi m_{\mu}} \frac{\eta_{\text{cap}} \gamma}{\beta^*} \frac{1}{E_{\text{cap}} r_{\mu}}$$

What doesn't directly appear

Beam emittance

 $\Box$  Bunch structure (fewer bunches for same  $P_P$ )







# Improving Luminosity

 $\mathcal{L} = \frac{\tau_{\mu} e B_{\text{avg}} \eta_{\text{trans}} P_P \Delta \nu}{2\pi m_{\mu}} \frac{\eta_{\text{cap}} \gamma}{\beta^*} \frac{\eta_{\text{cap}} \gamma}{E_{\text{cap}} r_{\mu}}$  $\circ$  Lower emittance may allow lower  $\beta^*$  $\Box$  Smaller  $\sigma_z$ , energy spread (longitudinal)  $\Box$  Smaller  $\beta^*$  has smaller dynamic aperture  $\Box$  Improve  $\beta^*$  faster than  $\eta_{\text{trans}}$  reduction  $\odot$  Dependence of  $\beta^*$  on emittances not obvious  $\Box$  Except  $\sigma_z$ , may be weak



## Luminosity Constrained by Radiation



What if we are constrained by radiation
 Maybe we can't go below a certain depth
 Assuming a given energy

$$\mathcal{L} = C_{\mathsf{rad}} rac{\Delta 
u \gamma}{eta^* r_\mu}$$

• Only improve with • Larger  $\Delta \nu$ • Smaller  $\beta^*$ 





#### **Beam-Beam Tune Shift**

Reach some beam-beam tune shift

$$\Delta \nu = \frac{N r_{\mu}}{4 \pi \epsilon_n} \qquad P_P = \frac{E_{\text{cap}} N n_B f_P}{\eta_{\text{cap}} \eta_{\text{trans}}}$$

 $\circ$  Lower N in proportion to  $\epsilon_n$ 

○ Lower N permits increasing n<sub>B</sub>f<sub>P</sub>
 □ Increased f<sub>P</sub> reduces p driver space charge
 □ ε<sub>n</sub> sufficiently low for p driver
 □ Increased n<sub>B</sub>, interesting...





## Making the Beam Sane

- Immediately at target
  - Energy spread too large
  - Angular spread large
- Tapered solenoid to reduce angular spread
- Reducing energy spread: phase rotation
   Drift to introduce time spread
   Time dep. voltage reduces energy spread
   Bunch at 200 MHz for cooling



### Making Beam Sane Neuffer Phase Rotation



Imagine beam already bunched

- Higher frequency RF than bunches
  - □ Early (high energy) bunches decelerated
  - Late (low energy) bunches accelerated
- Bunch and phase rotate together
- Avoids low frequency RF and/or induction linac

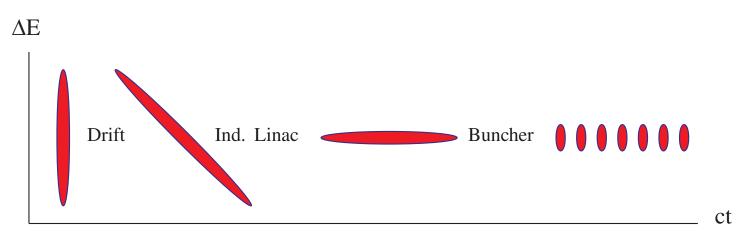


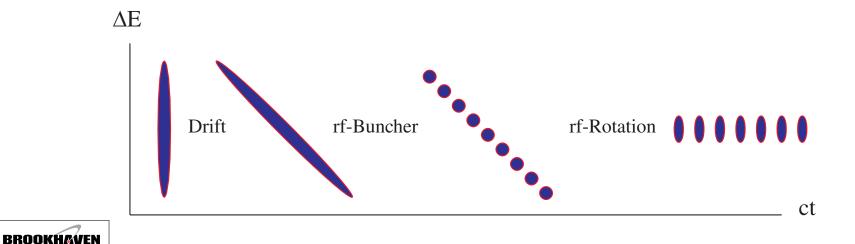




#### **Phase Rotation Scenarios**

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## Making Beam Sane Low Frequency Cooling



Low frequency ionization cooling ring/spiral
 Reduces longitudinal emittance rapidly
 Simulated rings (Balbekov) worked well
 Maybe not realistic...
 Would avoid bunch train





## Ionization Cooling Challenges



- Technology for reaching sufficiently low emittance
- High field magnets
   High magnetic fields on cavities
   Minimizing losses in system
   Cost of system





## Ionization Cooling Amount of RF Required



Energy losses in absorber restored in RF
 Amount of RF largely determines system cost

 $a \rho a$ 

$$\Delta E = \sum_{k} E_k \cos \phi_k = \frac{p \rho c}{\Lambda \eta_{\text{cool}}} \ln \frac{\epsilon_{6i}}{\epsilon_{6f}}$$
  
 $\circ E_k$ : RF energy gain  $k$   $\circ \phi_k$ : RF phase  $k$   
 $\circ p$ : momentum  $\circ \beta c$ : velocity  
 $\circ \Lambda$ : partition no. sum  $\circ \eta_{\text{cool}}$ : efficiency  
 $\circ \epsilon_{6i}$ : initial 6-D emit.  $\circ \epsilon_{6f}$ : final 6-D emit.





## Ionization Cooling Amount of RF Required



$$\Delta E = \sum_{k} E_k \cos \phi_k = \frac{p\beta c}{\Lambda \eta_{\text{cool}}} \ln \frac{\epsilon_{6i}}{\epsilon_{6f}}$$

#### $\circ$ What lowers $\eta_{cool}$ ?

- Mismatch when entering new system
- Approaching equilibrium
  - Maintain large angular, energy spread
  - $\diamond$  Taper down channel  $\beta$  function
  - Taper down bunch length (RF phase)





### Ionization Cooling Amount of RF Required



$$\Delta E = \sum_{k} E_k \cos \phi_k = \frac{p\beta c}{\Lambda \eta_{\text{cool}}} \ln \frac{\epsilon_{6i}}{\epsilon_{6f}}$$

Less voltage with smaller RF phase
 Lower longitudinal acceptance
 On crest, no longitudinal focusing





#### Ionization Cooling Losses



$$Q \equiv \frac{d\epsilon_6/\epsilon_6}{dN/N} \frac{N_f}{N_i} = \left(\frac{\epsilon_{6f}}{\epsilon_{6i}}\right)^{(1/Q)} \quad Q = \frac{\eta_Q \Lambda \tau_\mu dE}{\beta m_\mu c \, ds}$$

#### $\bigcirc$ Maximize Q to minimize losses

Keep η<sub>Q</sub> large: same problems as η<sub>cool</sub>
 Keep average gradient large

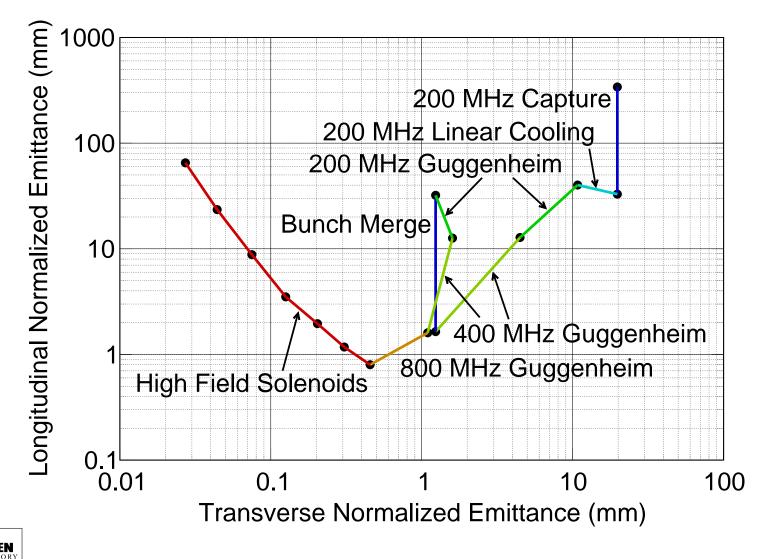
 Running closer to crest
 Densely packed lattice







## **Cooling Scheme**





## Cooling Guggenheims



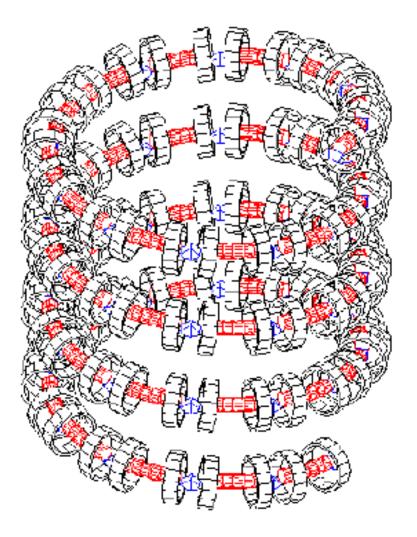
- Straight cooling lattice only cools transversely
- Add bend, wedge absorbers to couple to longitudinal
- Avoid injection/extraction
- O Long bunch trains won't fit in ring
- Taper the channel: avoid equilibrium
- One for each sign
- Increase frequency: maintain energy spread





### Cooling Guggenheim









## Cooling Bunch Merge



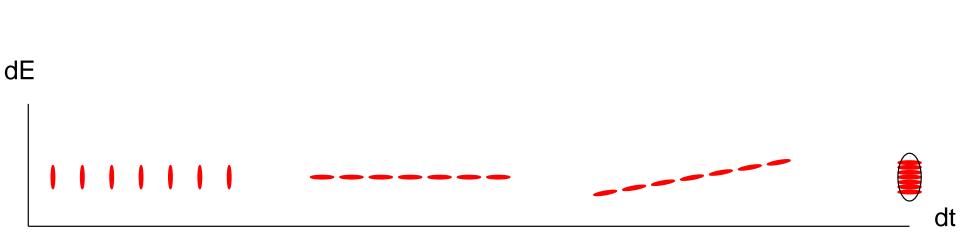
• Prefer everything in one bunch

- Needed to achieve beam-beam tune shift
   Avoid if final emittance small enough
- Induce energy spread: low frequency RF
- Drift until coincide in time
  - Accelerate drift with wiggler?
- Capture in single bucket
- O Significant losses: factor of 3!



#### Cooling Bunch Merge









## Cooling Post Merge Guggenheims



- Longitudinal emittance now large
- Reduce longitudinal emittance
- Reduce transverse emittance also
- Space charge becomes significant
- Avoided if no bunch merge



## Cooling High Field Solenoids



Need to get smaller transverse emittances
 Need large angular spreads in beam
 Use high-field (50 T) solenoids
 Little net 6-D cooling
 Reduces transverse emittance at expense of longitudinal





#### Cooling Low Emittance Scheme: Why?



- Avoid the bunch merge
  - □ Factor of 20 cooling each transverse
  - $\Box$  Reduces losses by factor of 3 ( $\eta_{\text{trans}}$ )
- $\circ$  Allows lower  $eta^*$
- Acceleration, ring beamlines less expensive
- Potentially permit high energy bunch merge
   Maybe more efficient





#### Cooling Low Emittance Scheme: How?



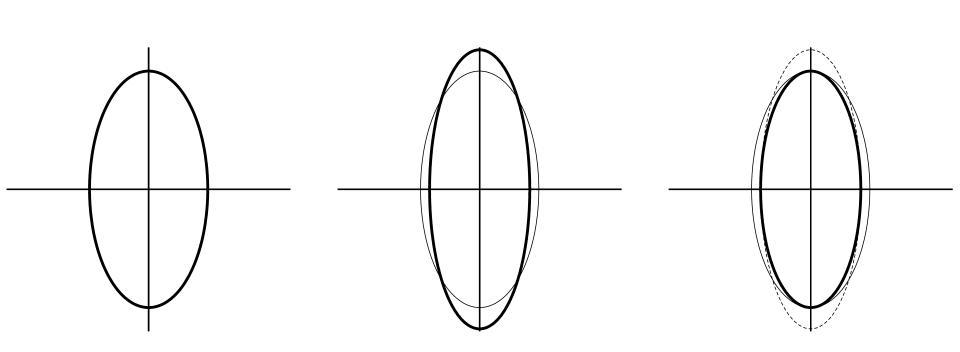
- $\circ$  Problem: getting small  $\beta$  functions
  - High magnetic fields
- OProposal "PIC"
  - Old idea of Balbekov
  - Run on linear resonance
  - Unstable direction is transverse momentum
  - Stable position is transverse position
  - Cooling reduces transverse momentum as much as instability increases





# Cooling PIC Principle









#### Cooling Low Emittance: Issues



#### ONO focusing

- System becomes sensitive to perturbations
- □ Space charge
- Chromatic aberrations
- Insuring losses don't exceed benefits





# Cooling Other Systems



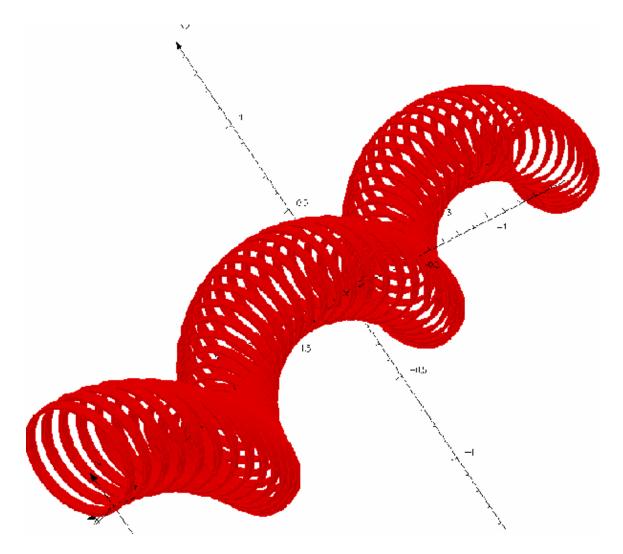
- Helical cooling channel
- Hoped to be more efficient than Guggenheim
- Still under development, various issues
   Getting cavities in channel





## Cooling Helical Cooling Channel











#### Acceleration

 $\circ$  Get to high energy without significant losses  $\circ$  Constant gradient V,

$$\frac{N_f}{N_i} = \left(\frac{E_f + p_f c}{E_i + p_i c}\right)^{-\frac{m_\mu c}{\tau_\mu V}}$$

Losses very modest if average gradient high



# Acceleration Techniques



Higher energies, more time for

□ Ramping magnets

- Varying RF frequency
- Ocan use higher frequency RF

○ RLAs

FFAGs, maybe adjusting RF frequency
 More passes through RF

Fast ramping synchrotron





# Acceleration Beam Loading



Possibly more passes through RF

- Ramping synchrotrons may use many passes
- FFAGs can adjust RF frequency
- ORF frequencies are higher

Less stored energy

OPower supplied to replace lost energy?





# **Collider Ring**



- $\odot$  Achieving sufficient dynamic aperture for small  $\beta^*$
- $\odot$  Achieving larger  $\Delta \nu$
- Question: to what extent do we get help from
  - Smaller longitudinal emittance
  - Smaller transverse emittance
- Highest field possible for luminosity







## **Sample Parameters**

CoM energy (TeV) 1.5 4 8  $\mathcal{L}$  (10<sup>34</sup> cm<sup>2</sup>s<sup>-1</sup>) 1 4 8 5.2 5.2 10.4  $B_{avg}$  (T)  $\beta^*$  (mm) 3 10 3 dp/p rms (%) 0.09 0.12 0.06 Ring depth (m) 13 135 540 0.07 0.07 0.07  $\eta$ trans  $f_P$  (Hz) 13 6 3  $P_P$  (MW) 1.8 0.8 4 0.1 N (10<sup>12</sup>)  $\Delta v$  $\epsilon_{\perp}$  (mm·mrad) 25  $|\epsilon_{\parallel}$  (mm·rad) 72







# **Final Remarks**

Getting luminosity means

Increasing proton driver power

- Increasing transmission
- □ Increasing beam-beam tune shift □ Lowering  $\beta^*$
- Increasing collider bending field
- OAII these increase radiation, except

Beam-beam tune shift







## **Final Remarks**

 Most transmission loss before acceleration Improving transmission requires Higher average RF gradients Reducing inefficiencies Addresses Matching Approaching equilibrium
 Best hope for improvement is eliminating rebunching







# **Final Remarks**

 Lowering emittance only helps indirectly Allows higher proton driver rep rate Easier on proton driver But no power reduction Potentially eliminate rebunching  $\odot$  Maybe improve  $\beta^*$ • Comes with a transmission cost



