

*GEM and other charge multipliers  
with VLSI pixel read-out*

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**Self-assembly of icosahedral**

**X-ray astronomy Polarimetry sees the light**

new on the market  
 liquid handling

**Table 1 Physical characteristics and performances of the micro-pattern detector**

	Present prototype (2–10 keV)	Improved configuration (3.5–10 keV)
Drift/absorption gap	6 mm	30 mm
Drift field	3,000 V cm <sup>-1</sup>	1,500 V cm <sup>-1</sup>
Gas filling and pressure	(Ne 80%–DME 20%); 1 atm	(Ne 40%–DME 60%); 4 atm
Gas grain	5,000	2,500
Transverse diffusion in drift	80 μm	<100 μm
GEM thickness		kapton foil
GEM hole geometry		1-μm pitch
GEM voltage		
Detection efficiency at 5.4 keV		
Read-out pixel size		
Number of pixels		

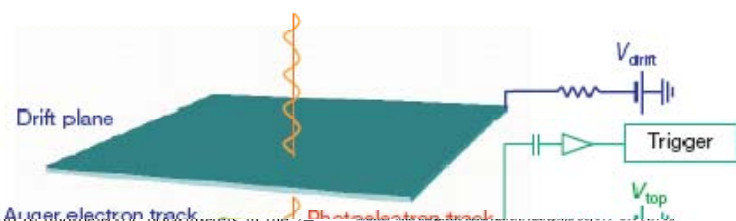
**letters to nature**

**Read-out plane technology**

- Track length/pixel size (6 keV)
- Sensitivity to Her X1
- Sensitivity to 3C-273
- Sensitivity to MCG-6-30-15
- Gain in the integration time over SXRP (strong sources)
- Gain in the integration time over SXRP (faint sources)

We show also the observing time needed to measure at 99% micro-pattern gas chamber (MPGC) is placed at the focus of the use only photons above 3.5 keV to compute the sensitivity. I unity of a new generation of photoelectric polarimeters in the 2–10 keV band. The device can simultaneously also produce good images (50–100 μm), moderately good spectroscopy (16% full-width at half-maximum, at 5.4 keV), and fast, high-rate timing down to 150 eV. Moreover, being truly two-dimensional, it is non-dispersive and does not require rotation.

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= 10%  
 JP = 1%  
 JP = 1%  
 C (over Bragg)  
 100 (over Bragg)  
 sical sources if the  
 proved design we  
 antly reduces the

We also tested our capability to model the polarization detection processes. As absorption, slowing down, scattering and transverse diffusion of electrons in the drift are well known quantities, we may reliably predict the performance of another detector configuration that would be better optimized for astrophysical applications. It is based on an existing<sup>30</sup> VLSI (Very Large Scale Integration) readout chip combined with other well established detector technology. We can derive the polarimetric sensitivity of such detectors when installed at the focus of a real X-ray telescope. In Table 1 we compare the sensitivity of the present and final configuration of the MPGC with SXRP.

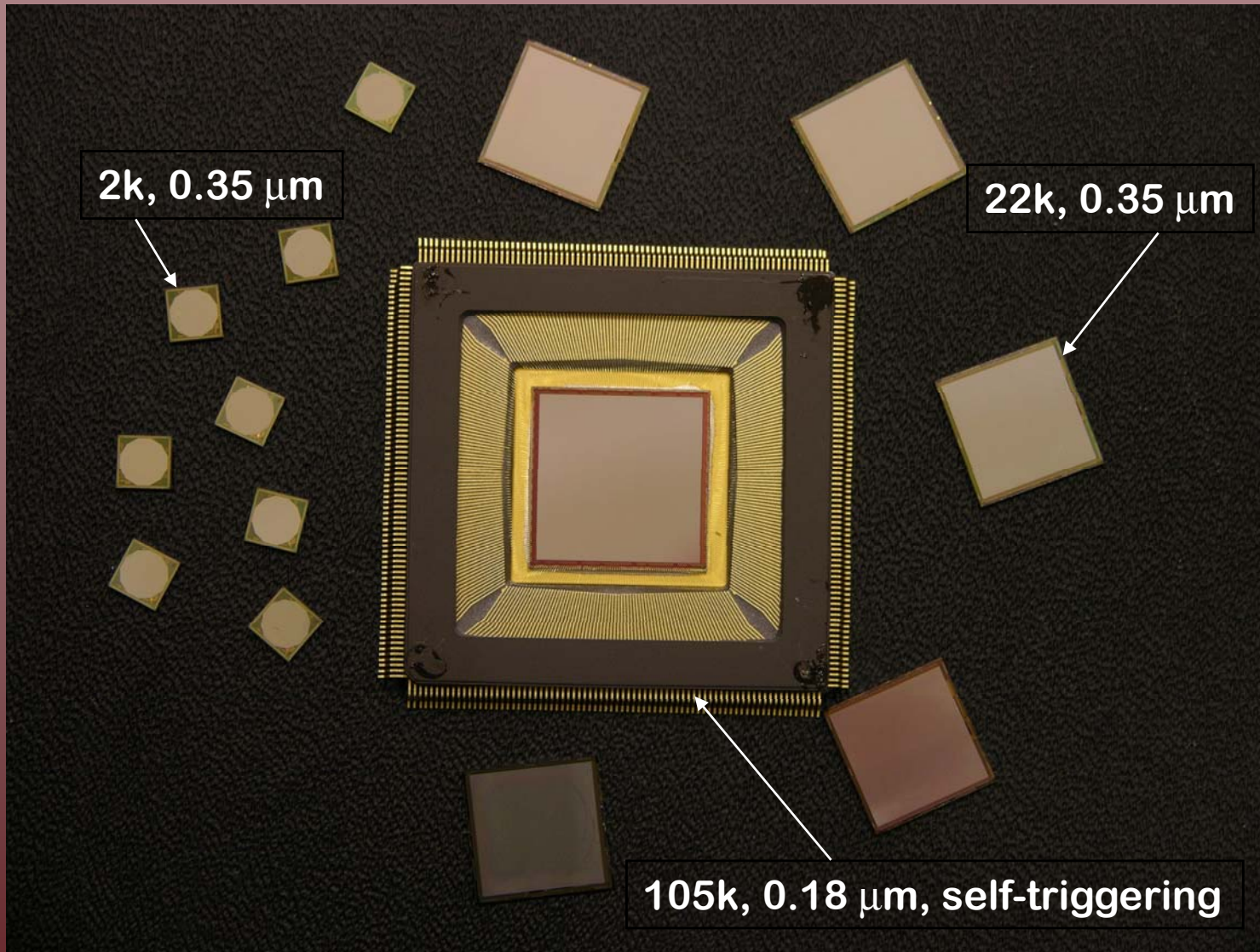
The MPGC requires integration periods that are about 100 times shorter than those of the SXRP to detect the same polarization in bright sources. With integrations of the order of one day we could perform polarimetry of active galactic nuclei at the 1% level, a breakthrough in this fascinating window of high-energy astrophysics.

- Conners, P. A. & Stark, R. E. Observable gravitational effects on polarized radiation coming from near a black hole. *Nature* 269, 128–129 (1977).
- Stark, R. E. & Conners, P. A. Observational test for the existence of a rotating black hole in Cyg X-1. *Nature* 266, 429–430 (1977).
- Conners, P. A., Piran, T. & Stark, R. E. Polarization features of X-ray radiation emitted near a black hole. *Astrophys. J.* 235, 224–244 (1980).
- Bao, G., Wita, P. & Hadrava, P. Energy-dependent polarization variability as a black hole signature. *Phys. Rev. Lett.* 77, 12–15 (1996).
- Calzetti, A. & Matt, G. Polarization properties of synchrotron self-Compton emission. *Mon. Not. R. Astron. Soc.* 268, 451–458 (1994).
- Postman, J. Relativistic jets in blazars polarization of radiation. *Astrophys. J. Suppl. Ser.* 92, 607–609 (1994).
- Tanaka, Y. et al. Gravitationally redshifted emission implying an accretion disk and massive black hole in the active galaxy MCG-6-30-15. *Nature* 375, 659–661 (1995).
- Okura, J., Nishiyori, O. & Kajino, Y. Profiles and polarization properties of emission lines from relativistic disks. *Publ. Astron. Soc. Jpn.* 32, 841–845 (2000).
- Matt, G., Fabian, A. C. & Ross, R. R. X-ray photoionized accretion discs: UV and X-ray continuum spectra and polarization. *Mon. Not. R. Astron. Soc.* 264, 839–852 (1993).
- Nowick, R., Weisskopf, M. C., Berthelndorf, R., Links, R. & Wolff, R. S. Detection of X-ray polarization of the Crab Nebula. *Astrophys. J.* 174, L1–L8 (1972).
- Weisskopf, M. C., Silver, E. H., Kestenbaum, H. L., Long, K. S. & Novick, R. A precision measurement of the X-ray polarization of the Crab Nebula without polar contamination. *Astrophys. J.* 220, L117–L122 (1978).
- Kaaret, P. et al. SXRP: a focal plane stellar X-ray polarimeter for the Spectrum-X-Gamma mission. *Opt. Eng.* 29, 773–783 (1990).
- Schnopper, H. W. & Kaku, K. Polarimeter for celestial X-rays. *Astron. J.* 74, 854–858 (1969).
- Nowick, R. In *Planets, Stars and Nebulae Studied with Photopolarimetry* (ed. Gehrels, T.) 262–317 (Univ. Arizona Press, Tucson, 1972).

**An efficient photoelectric X-ray polarimeter for the study of black holes and neutron stars**

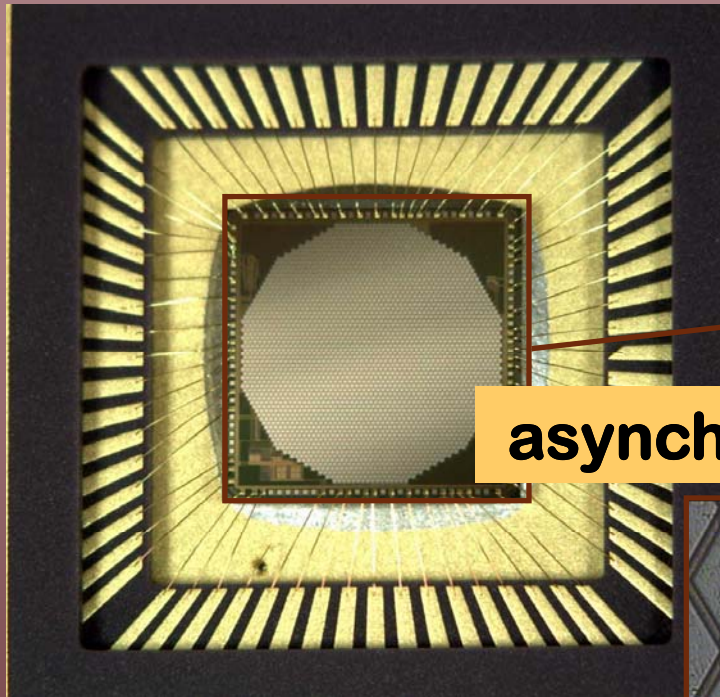
E. Costa, P. Soffitta, R. Bellazzini, A. Brez, N. Lumb, G. Spandre  
*Nature*, Vol. 411 (2001) 662.

- Sauli, F. GEM: a new concept for electron amplification in gas detectors. *Nucl. Instrum. Methods A* 386, 531–534 (1997).
- Campbell, M. et al. A pixel readout chip for 10–30 Mrad in standard 0.25 μm CMOS. *IEEE Trans. Nucl. Sci.* 46, 156–160 (1999).
- Bavdas, M. et al. Status of the X-ray evolving universe spectroscopy mission (XEUS). *Proc. SPIE* 4138, 69–78 (2000).
- Christensen, E. E. et al. X-ray calibration of the SODART flight telescope. *Proc. SPIE* 3113, 69–78 (1997).

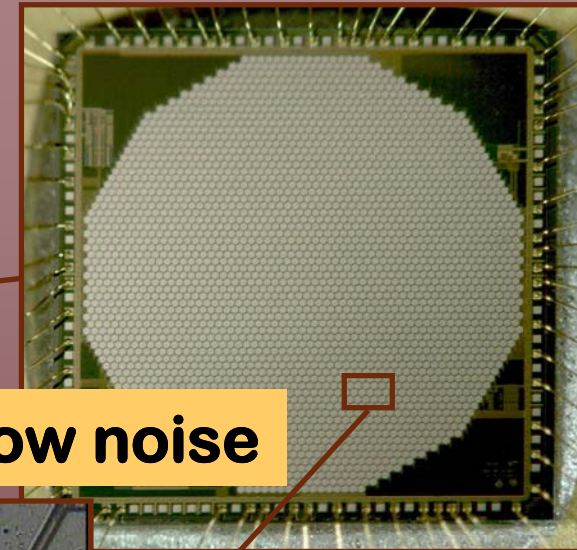


**Three ASIC generations of increasing size, reduced pitch and improved functionality have been realized**

# The collecting anode/read-out VLSI chip

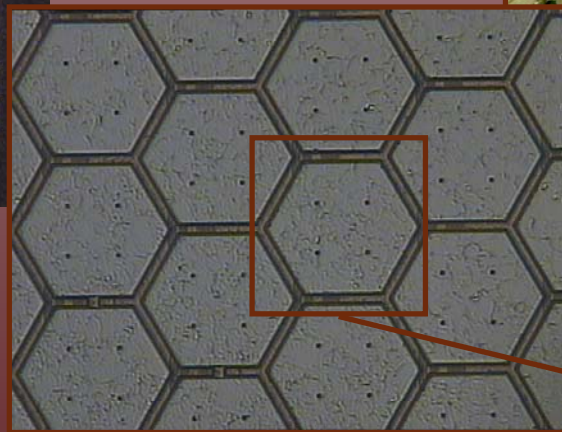


First ASIC prototype

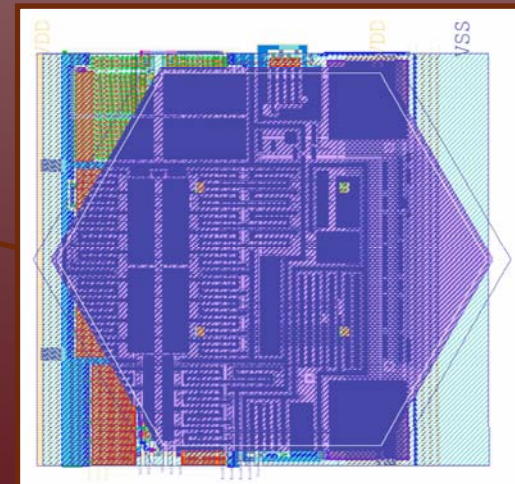


asynchronous, fast, low noise

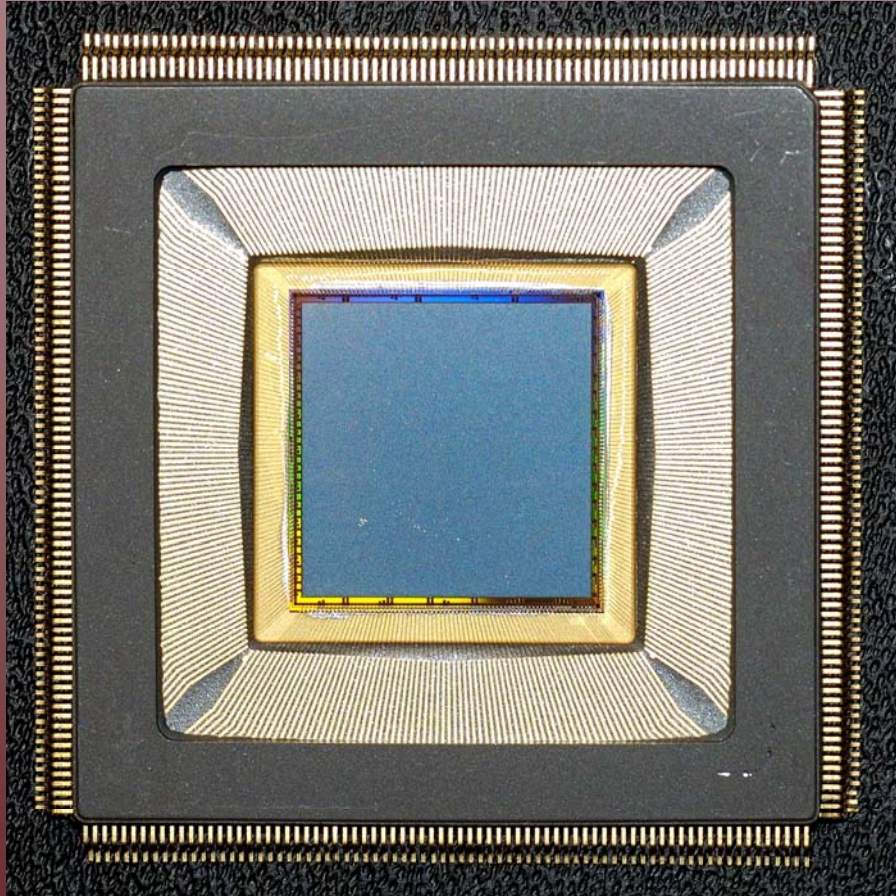
pixel electronics dimension:  
 $80\ \mu\text{m} \times 80\ \mu\text{m}$  in an  
hexagonal array,  
comprehensive of  
preamplifier/shaper, S/H and  
routing (serial read-out) for  
each pixel  
number of pixels: 2101



$\sim 3.5\ \mu\text{s}$  shaping time  
100 e<sup>-</sup> ENC  
100 mV/fC input sensitivity  
20 fC dynamic range



# Last technological step: a 0.18 $\mu\text{m}$ CMOS VLSI



The chip integrates more than 16.5 million transistors. It has a 15mm x 15mm active area of 105'600 pixels organized in a honeycomb matrix

470 pixels/mm<sup>2</sup>

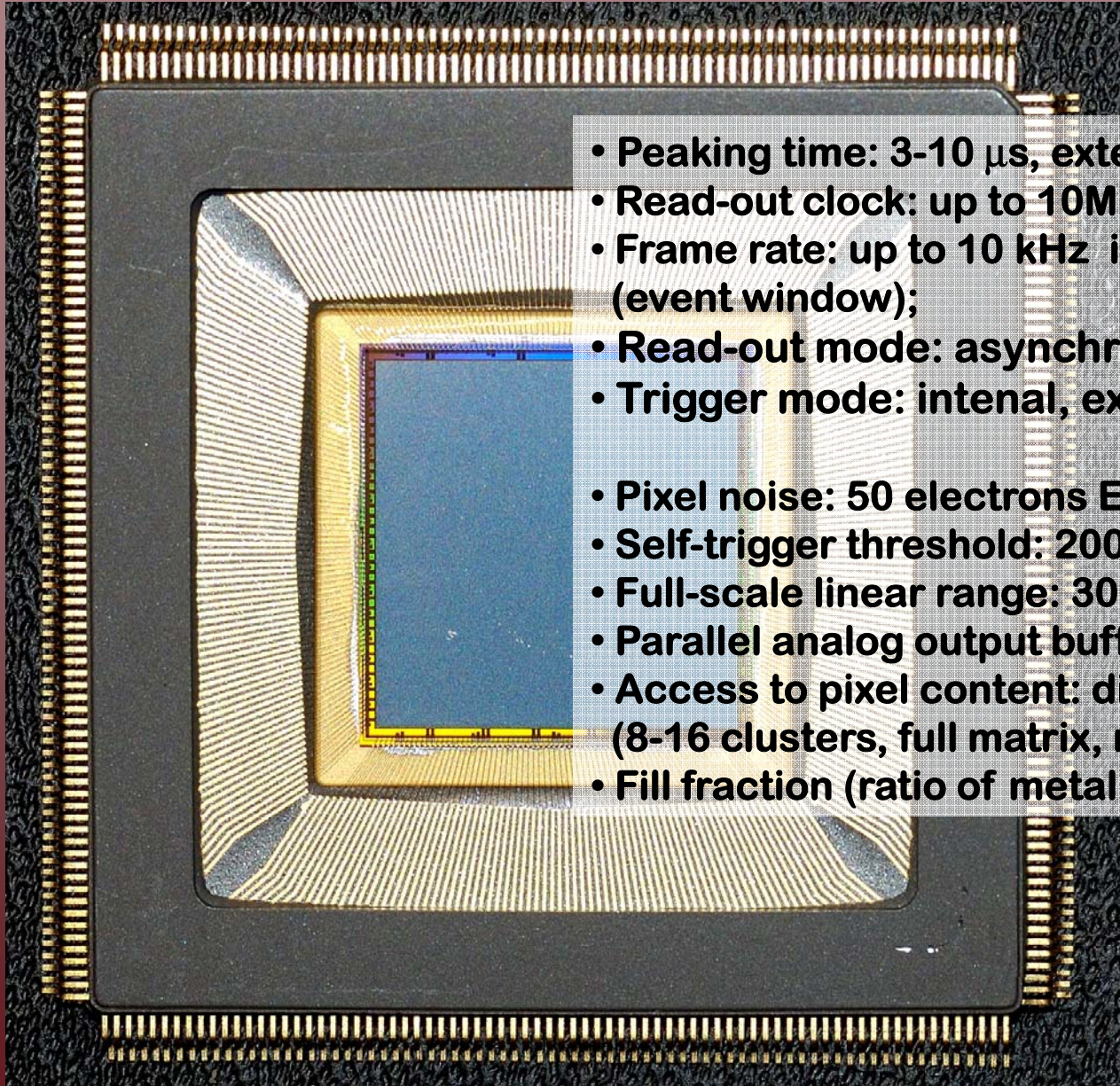
## Matrix organization

300 (width=300x50 $\mu\text{m}$ =15mm) x 352 (height=352x43.3 $\mu\text{m}$ =15.24mm) pixels

16 clusters of 300 x 22 = 6600 pixels each or

8 clusters of 300 x 44 = 13200 pixels each

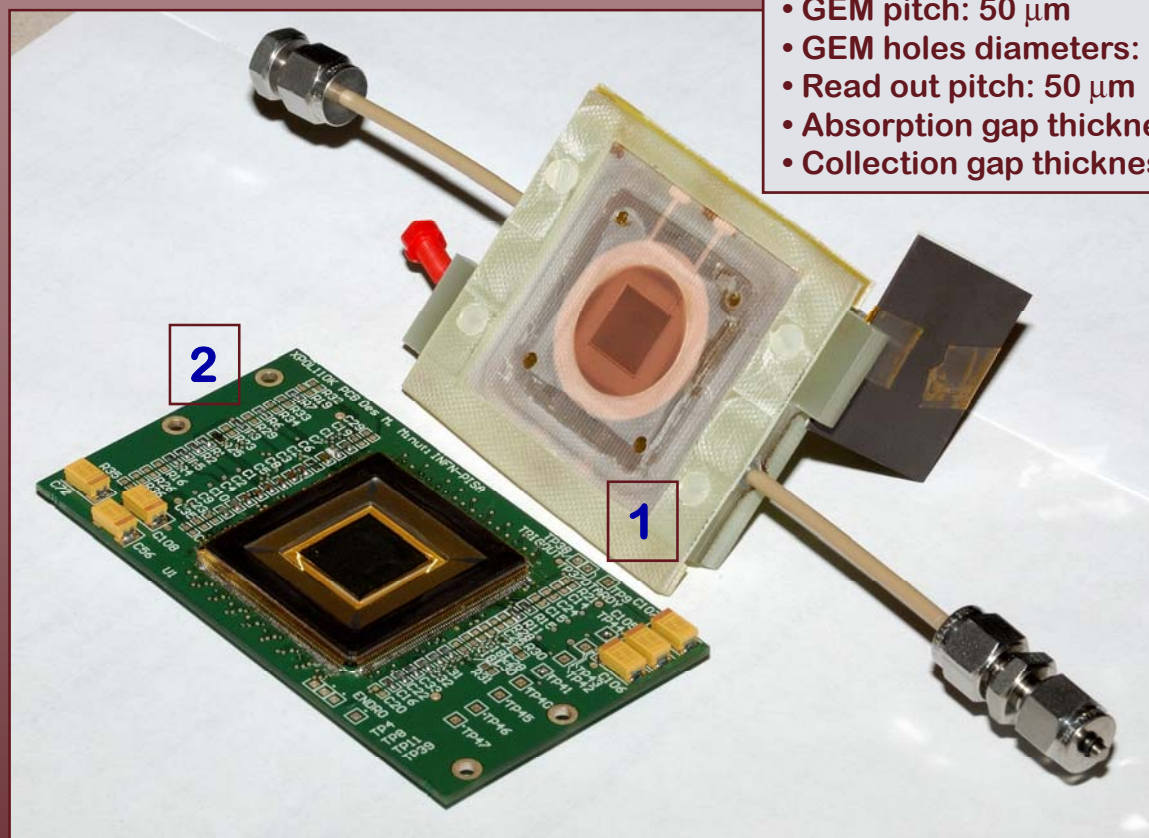
# 0.18 $\mu\text{m}$ ASIC features



- Peaking time: 3-10  $\mu\text{s}$ , externally adjustable;
- Read-out clock: up to 10MHz;
- Frame rate: up to 10 kHz in self-trigger mode (event window);
- Read-out mode: asynchronous or synchronous;
- Trigger mode: internal, external or self-trigger;
  
- Pixel noise: 50 electrons ENC;
- Self-trigger threshold: 2000 electrons;
- Full-scale linear range: 30000 electrons;
- Parallel analog output buffers: 1, 8 or 16;
- Access to pixel content: direct (single pixel) or serial (8-16 clusters, full matrix, region of interest);
- Fill fraction (ratio of metal area to active area): 92%

**Total power dissipation  
~ 0.5 Watt**

# Detector assembly



- GEM pitch: 50  $\mu\text{m}$
- GEM holes diameters: 33  $\mu\text{m}$ , 15  $\mu\text{m}$
- Read out pitch: 50  $\mu\text{m}$
- Absorption gap thickness: 10 mm
- Collection gap thickness: 1 mm

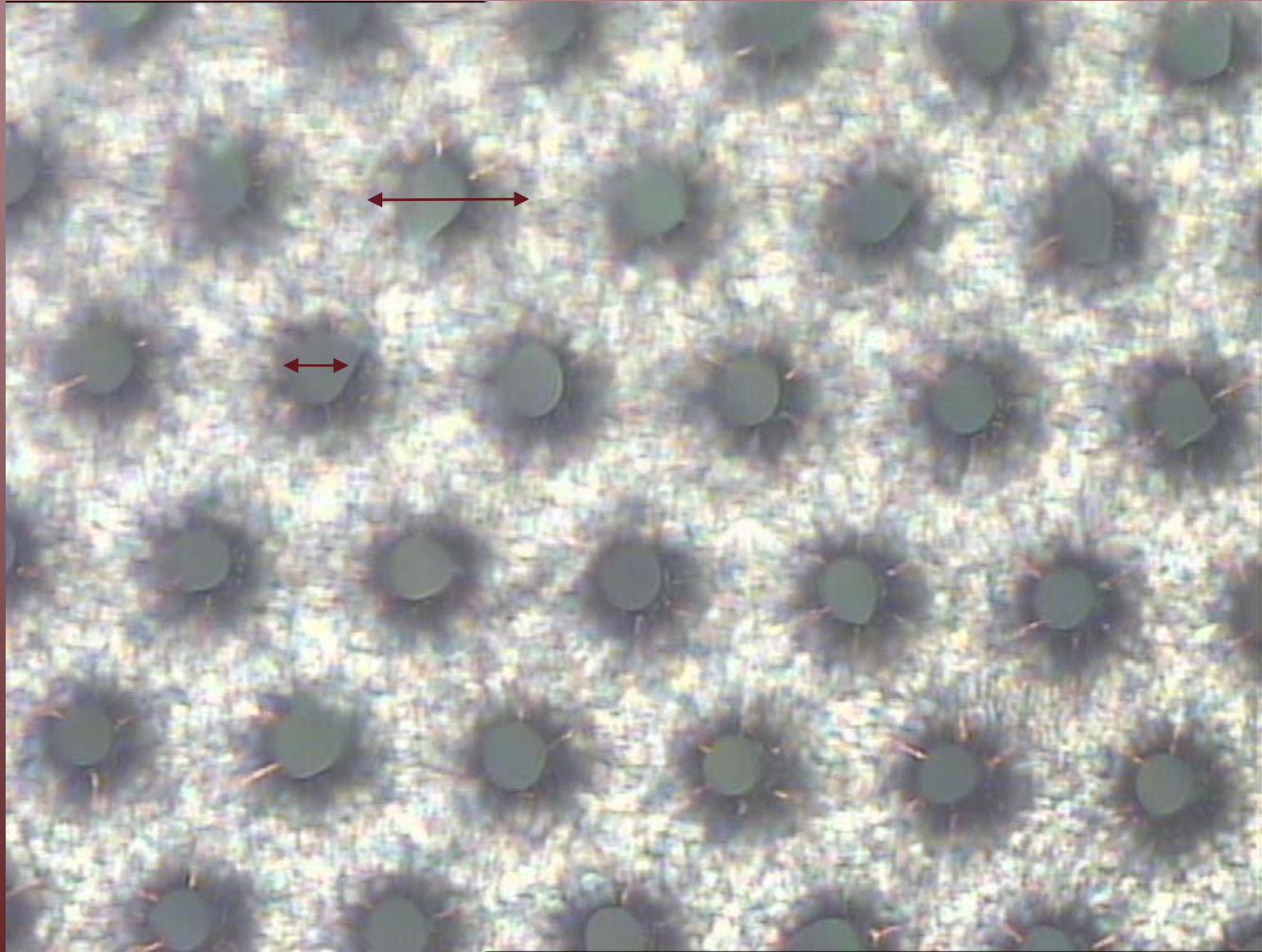
- 1 - The GEM glued to the bottom of the gas-tight enclosure
- 2 - The large area ASIC mounted on the control motherboard

Large effective gas gain around 1000 @450V in Ne(50%)-DME(50%)  
(at least 70 V less than in our standard 90  $\mu\text{m}$  pitch GEM)

# GEM specs

pitch: 50  $\mu\text{m}$

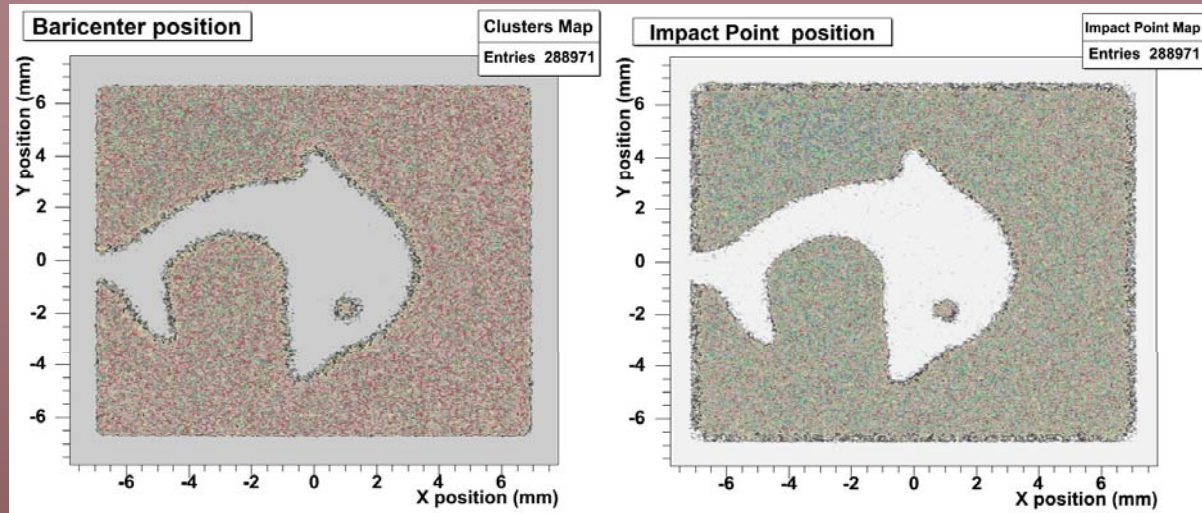
holes inner  $\varnothing$ : 33  $\mu\text{m}$



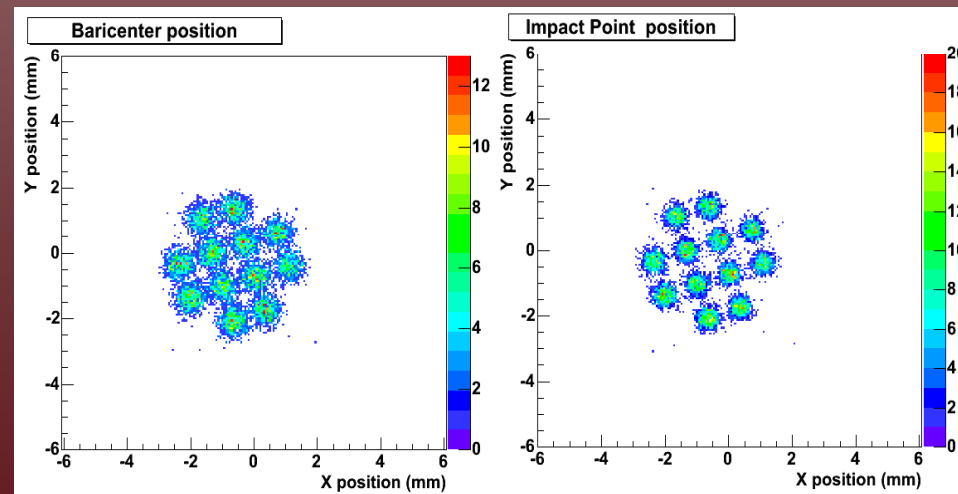
The matching of readout and gas amplification (GEM) pitch allows getting optimal results and to fully exploit the very high granularity of the device



# Imaging capability



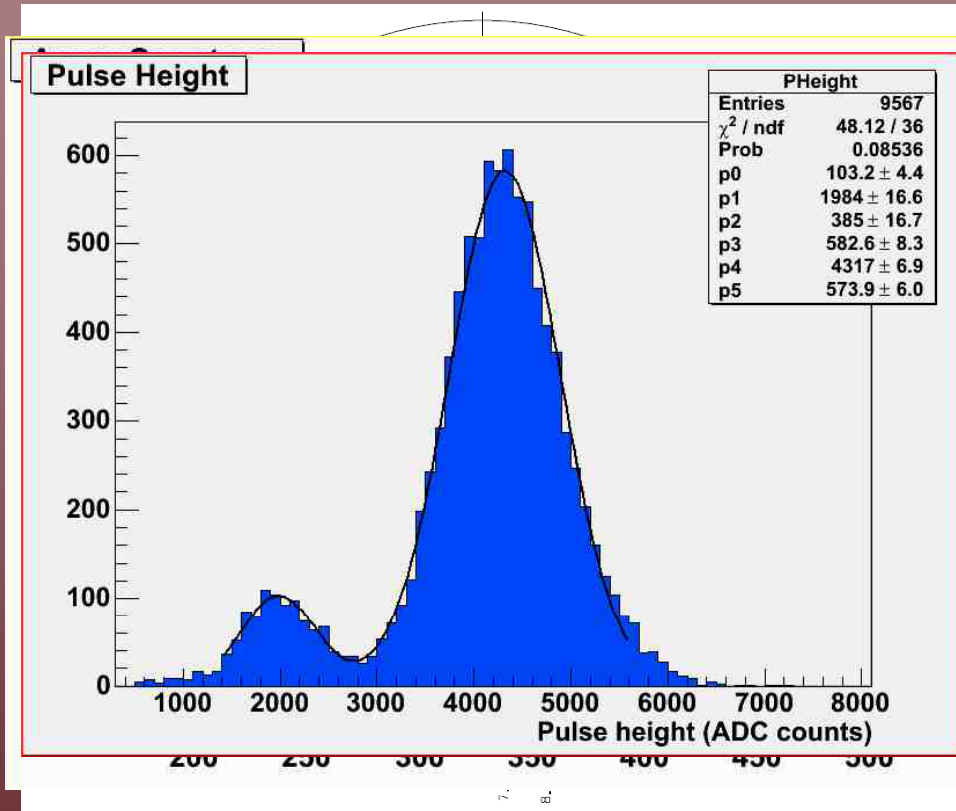
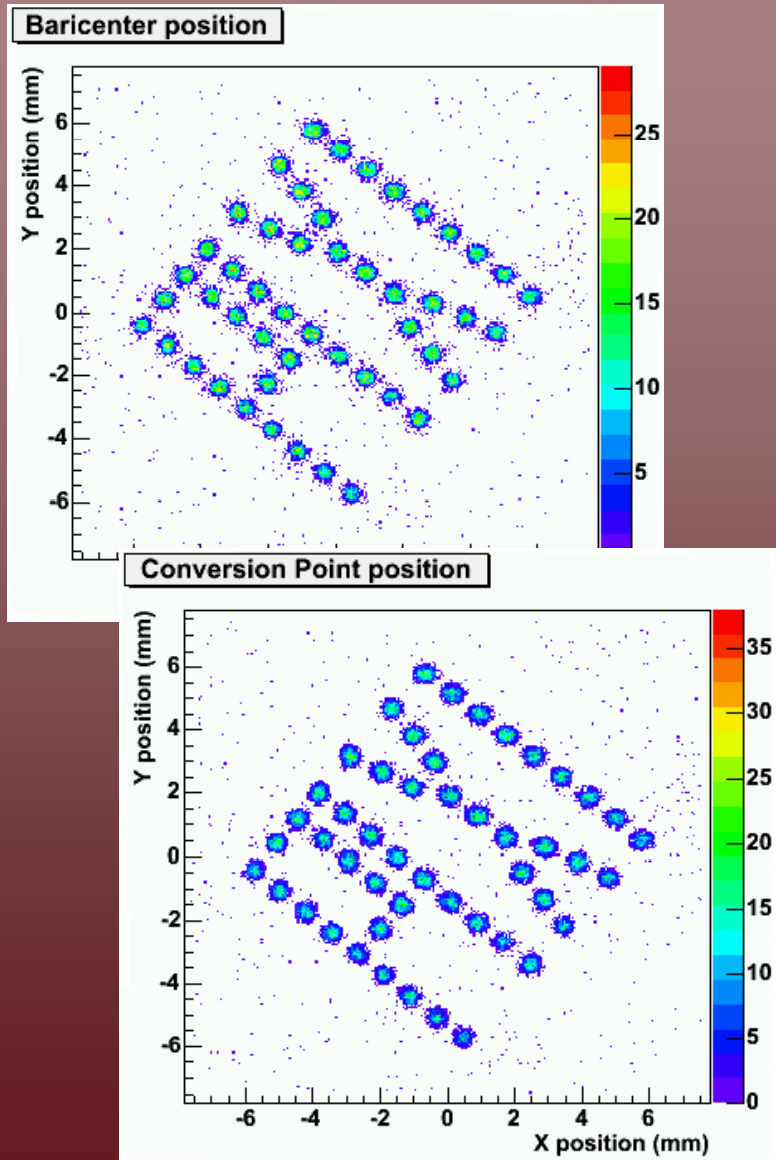
$^{55}\text{Fe}$  source Ne(50%)-DME(50%)



Holes: 0.6 mm diameter, 2 mm apart.

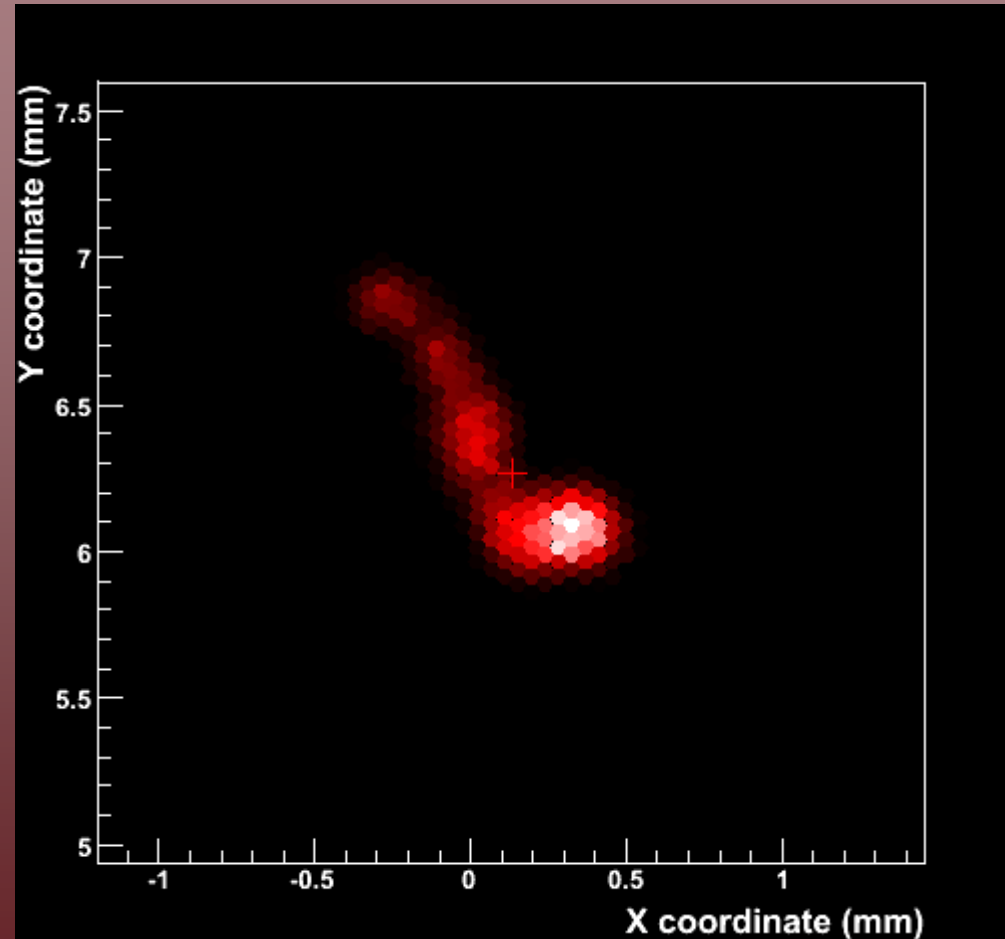
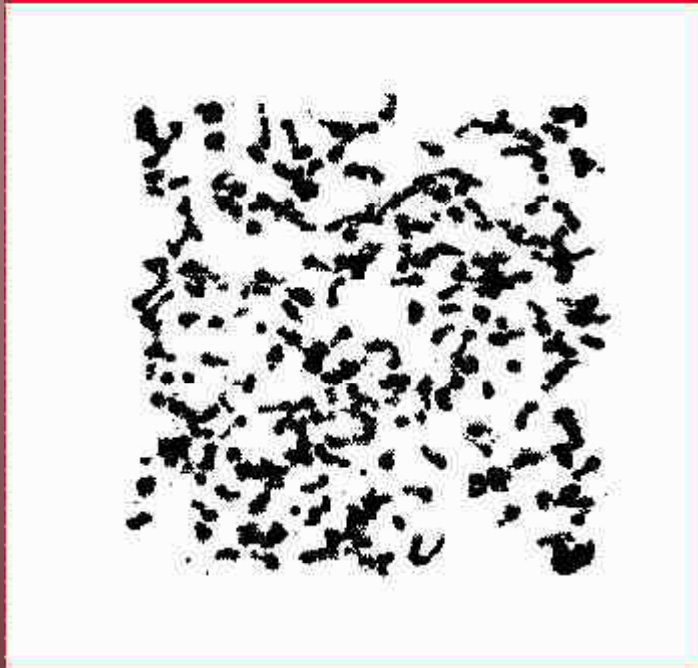
# Imaging and spectroscopic capability

Argon (50%)-DME(50%)

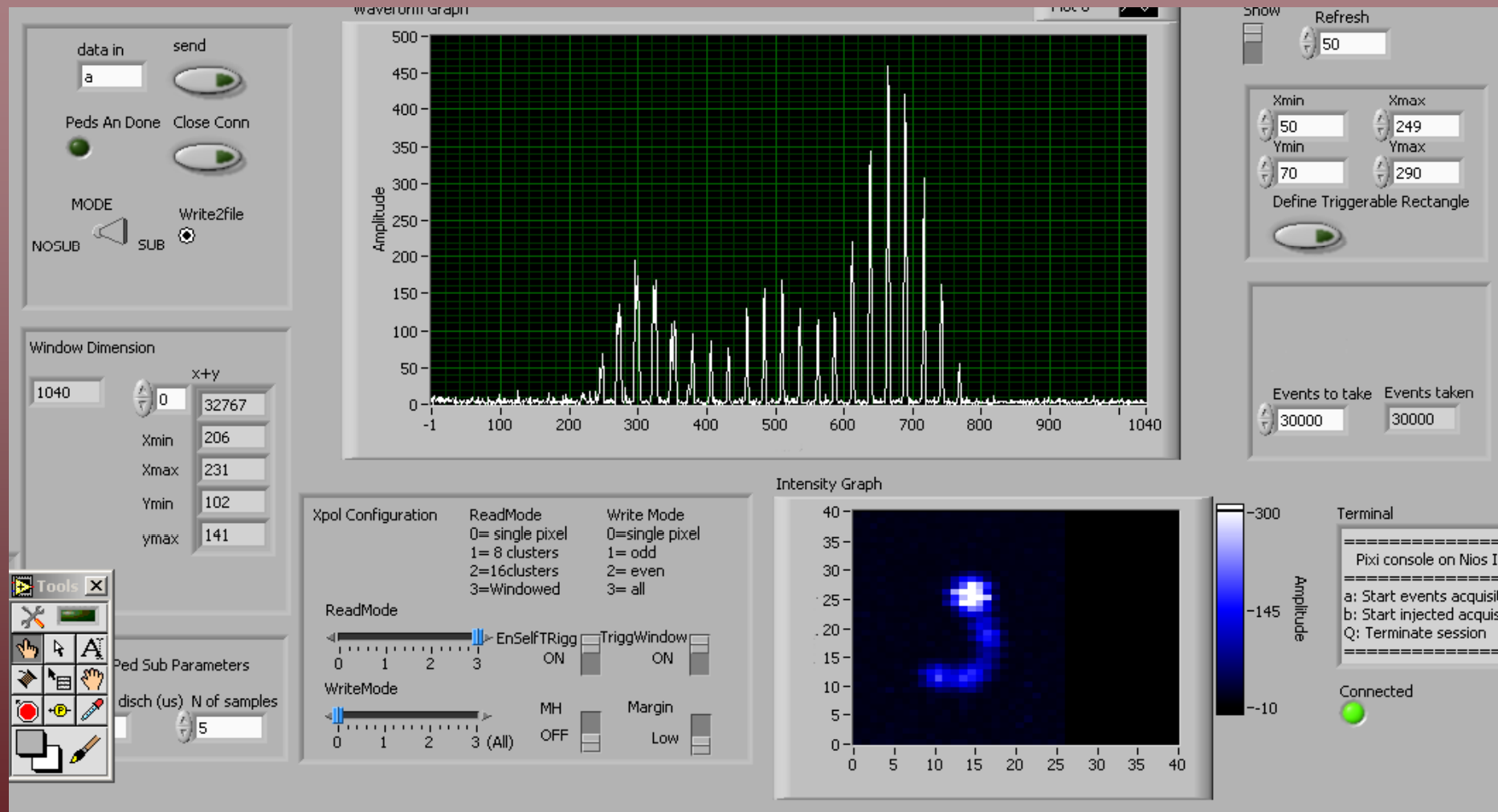


Holes:  
Ø 0.5 mm  
pitch 1 mm

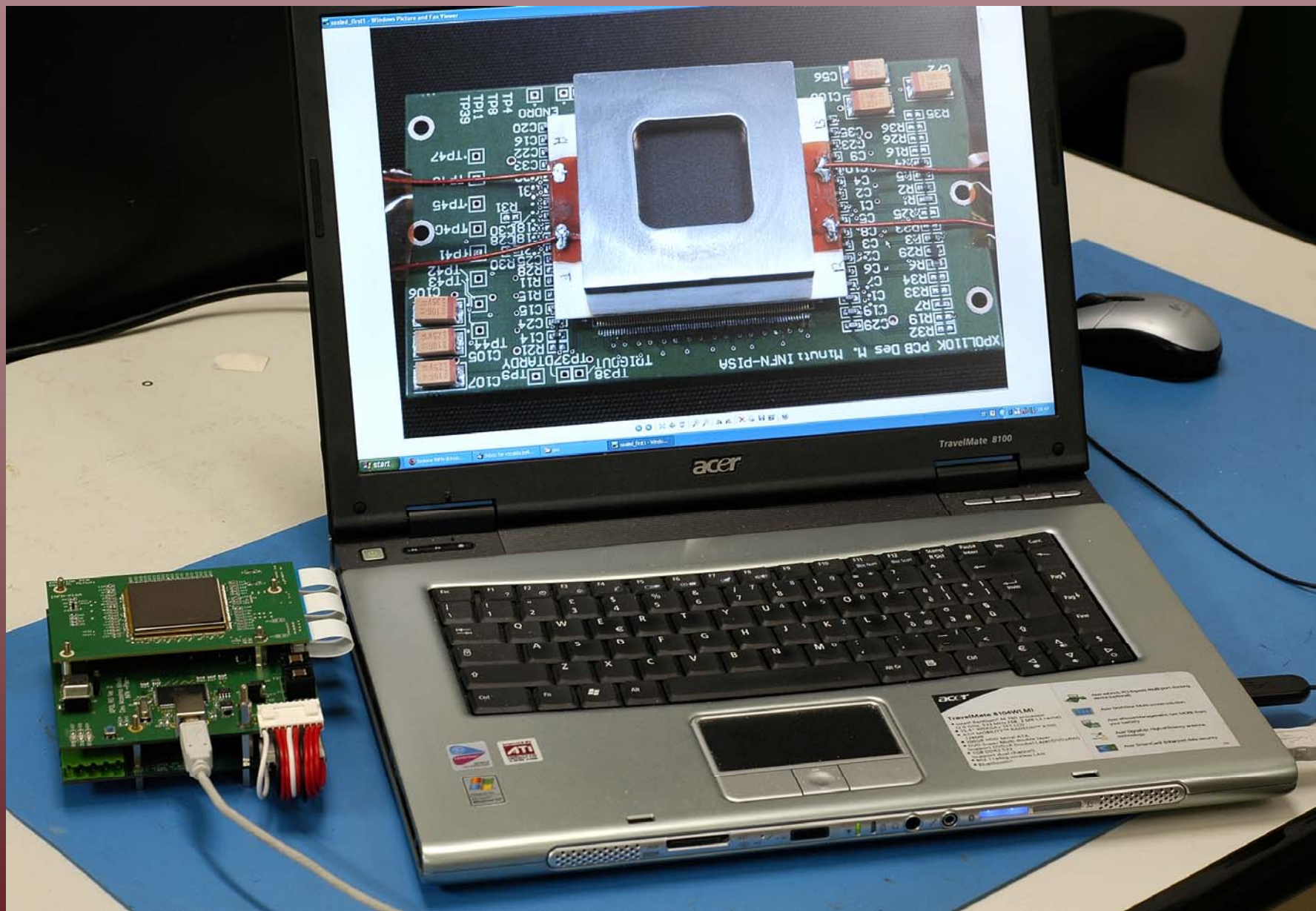
# Track morphology and angle reconstruction



# On-line monitoring



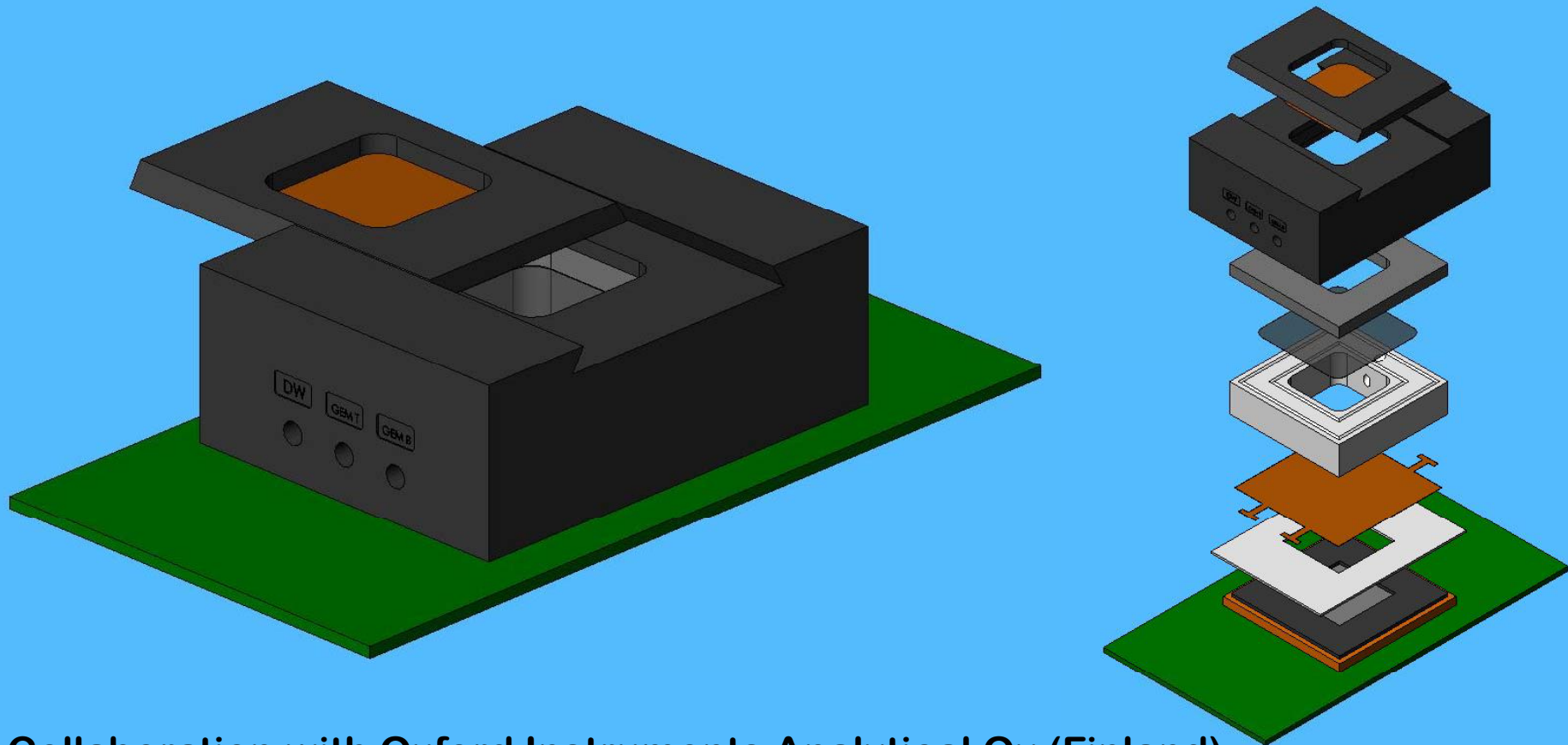
Real time pedestal subtraction



The level of integration, compactness and operational simplicity of this device is comparable to solid state detectors

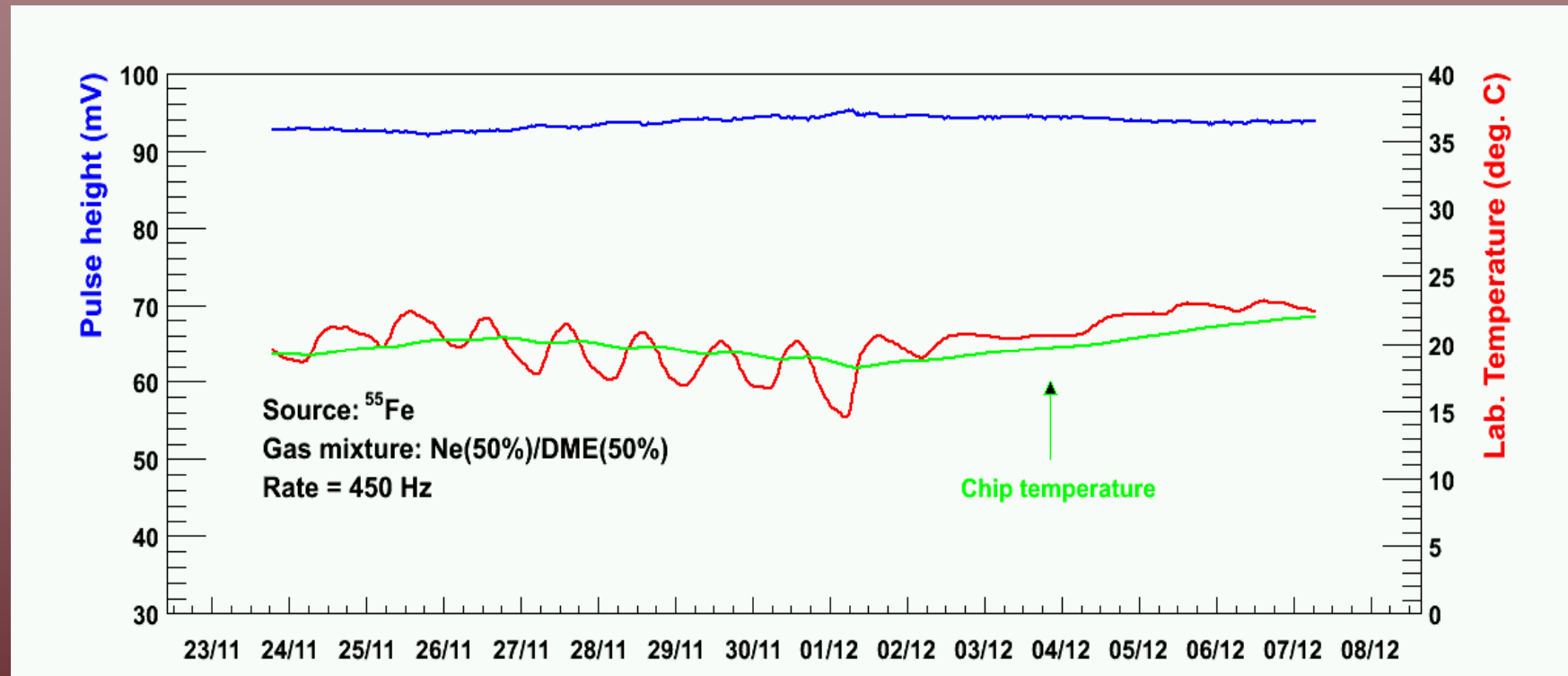
# Sealed device

(only clean materials, baking & outgassing)



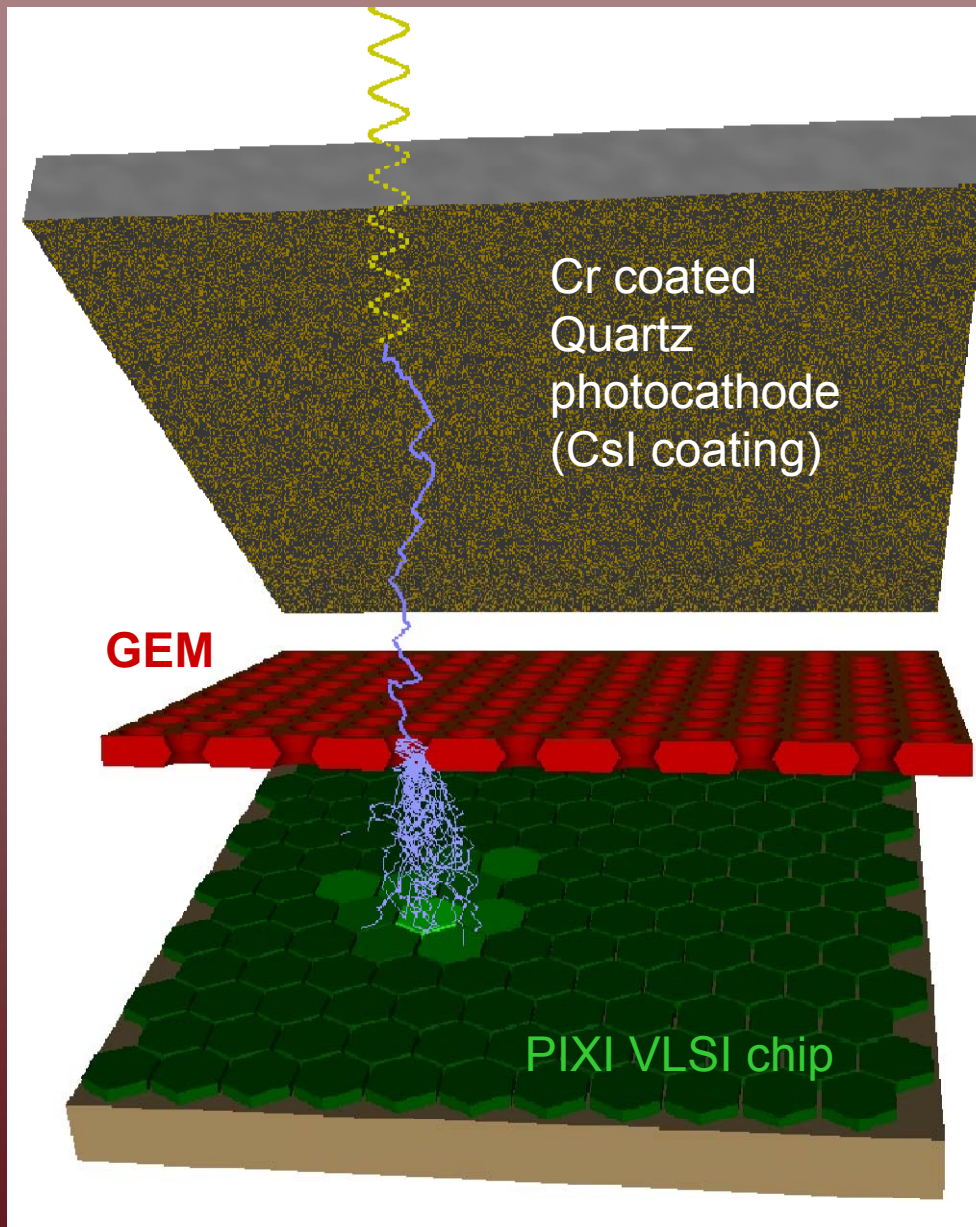
Collaboration with Oxford Instruments Analytical Oy (Finland)

# Long-term gain stability



Irradiation with  $^{55}\text{Fe}$  has started on November 20<sup>th</sup>  
Pulse height registration on 23<sup>rd</sup> at 7:30pm.

# Semitransparent Photocathode



Drift gap = 1 mm  
Transfer gap = 1mm  
GEM thickness = 50  $\mu\text{m}$   
GEM pitch = 50  $\mu\text{m}$

## Pros:

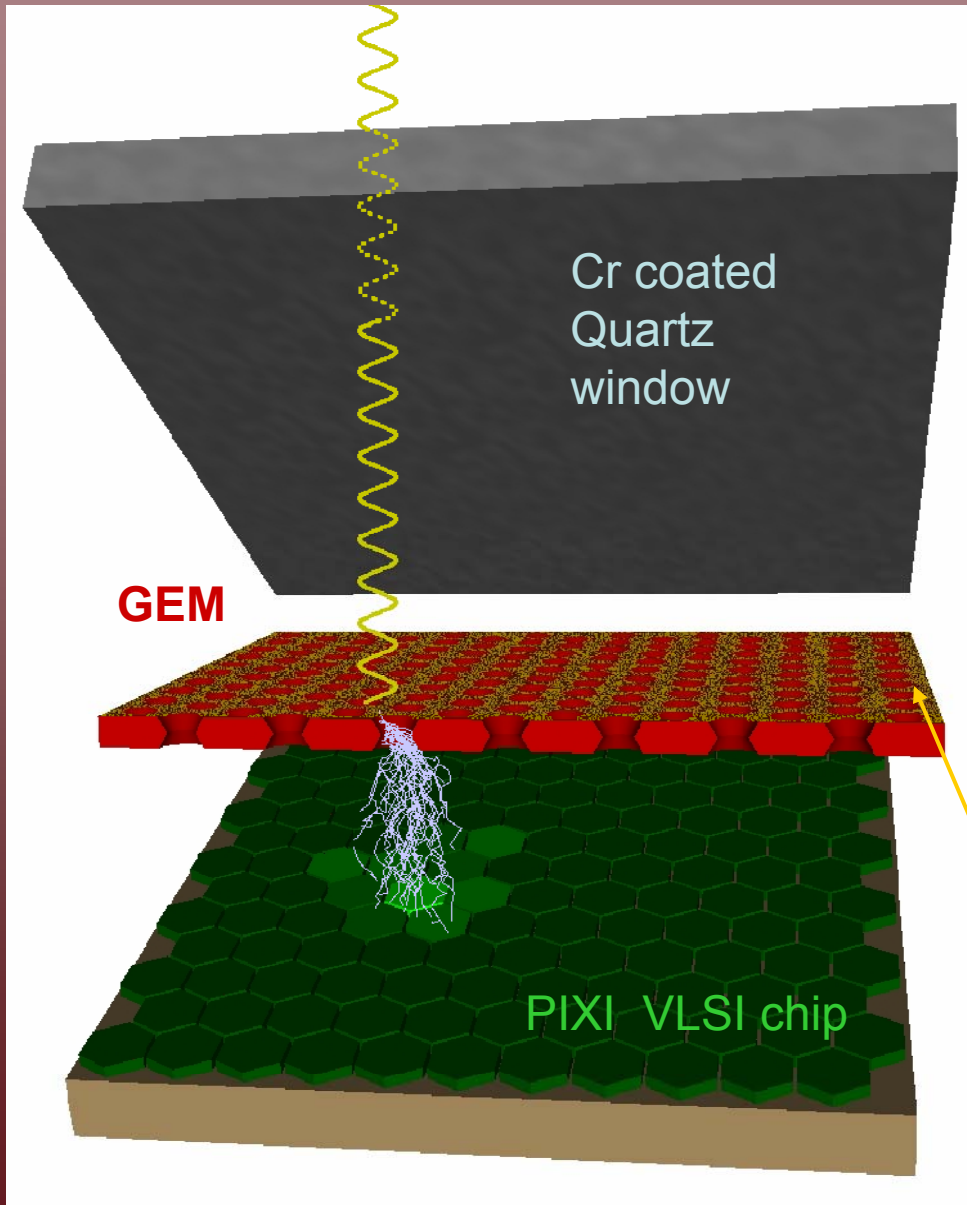
- Simple
- High gain
- High geometrical efficiency

## Cons:

- Low thickness  $\rightarrow$  Low Q.E.
- Extra diffusion in the gas layer above the GEM



# Reflective Photocathode



Drift gap = 1 mm  
Transfer gap = 1mm  
GEM thickness =  $50 \mu\text{m}$   
GEM pitch =  $50 \mu\text{m}$

## Pros:

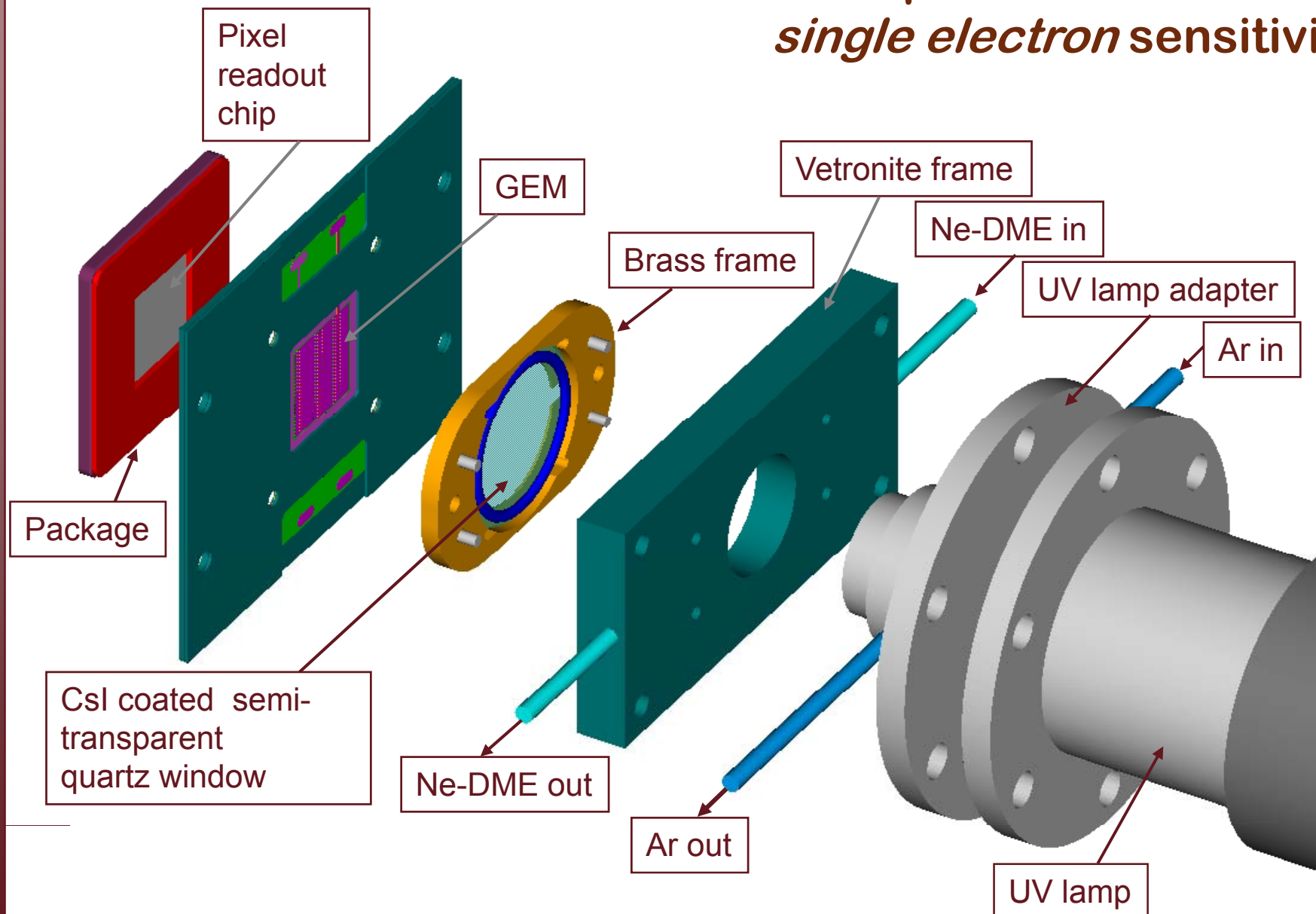
- Thick film  $\rightarrow$  high Q.E. (10-20%)

## Cons:

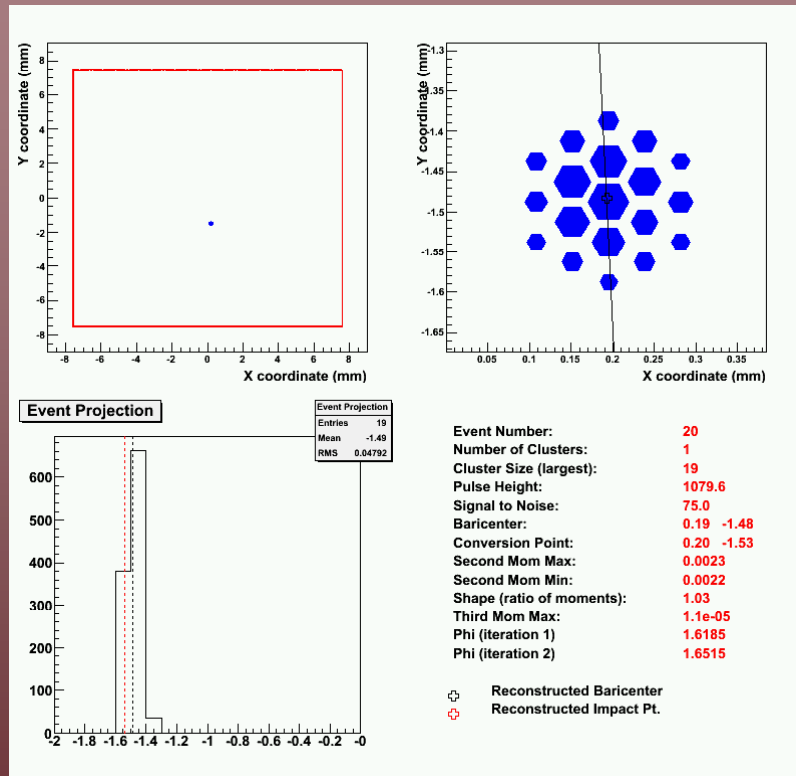
- More complicated to build
- Special gold coating on the GEM
- Low geometrical efficiency (in our case 50%)
- Lower gas gain

**CsI photocathode**

# UV photo-detector with *single electron sensitivity*

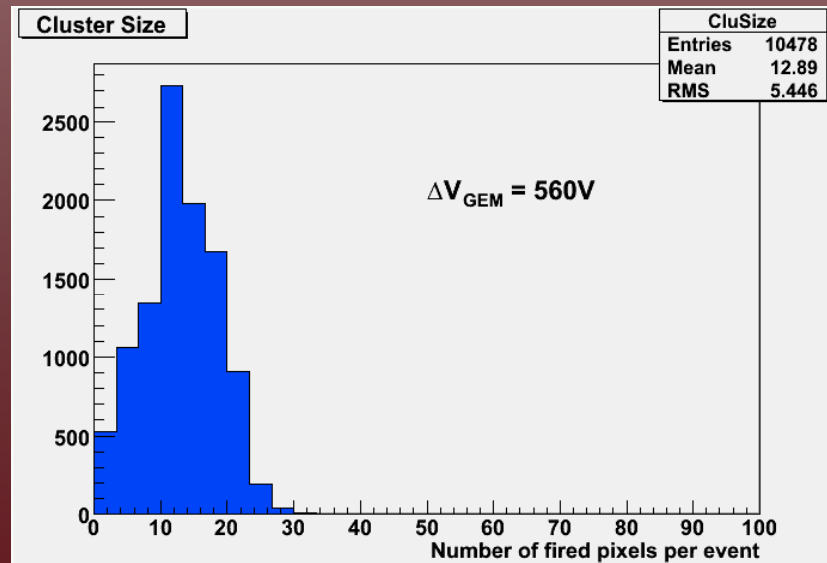
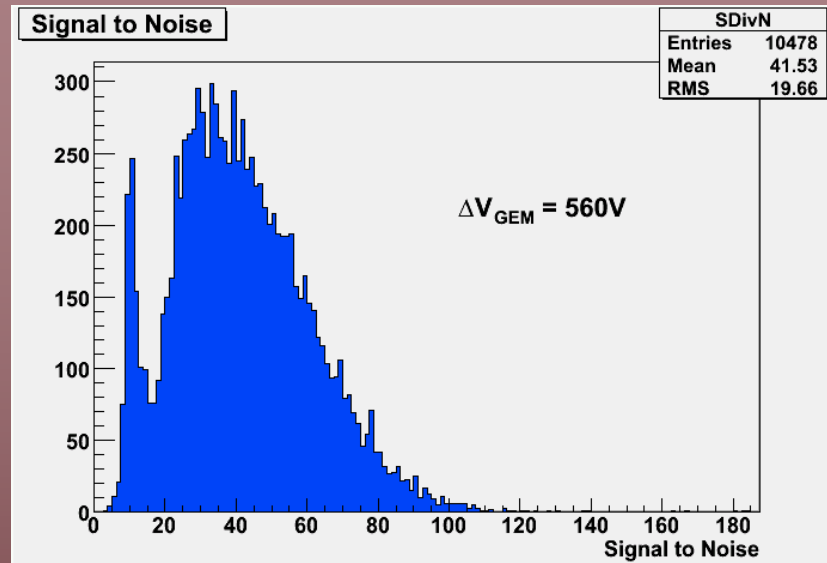


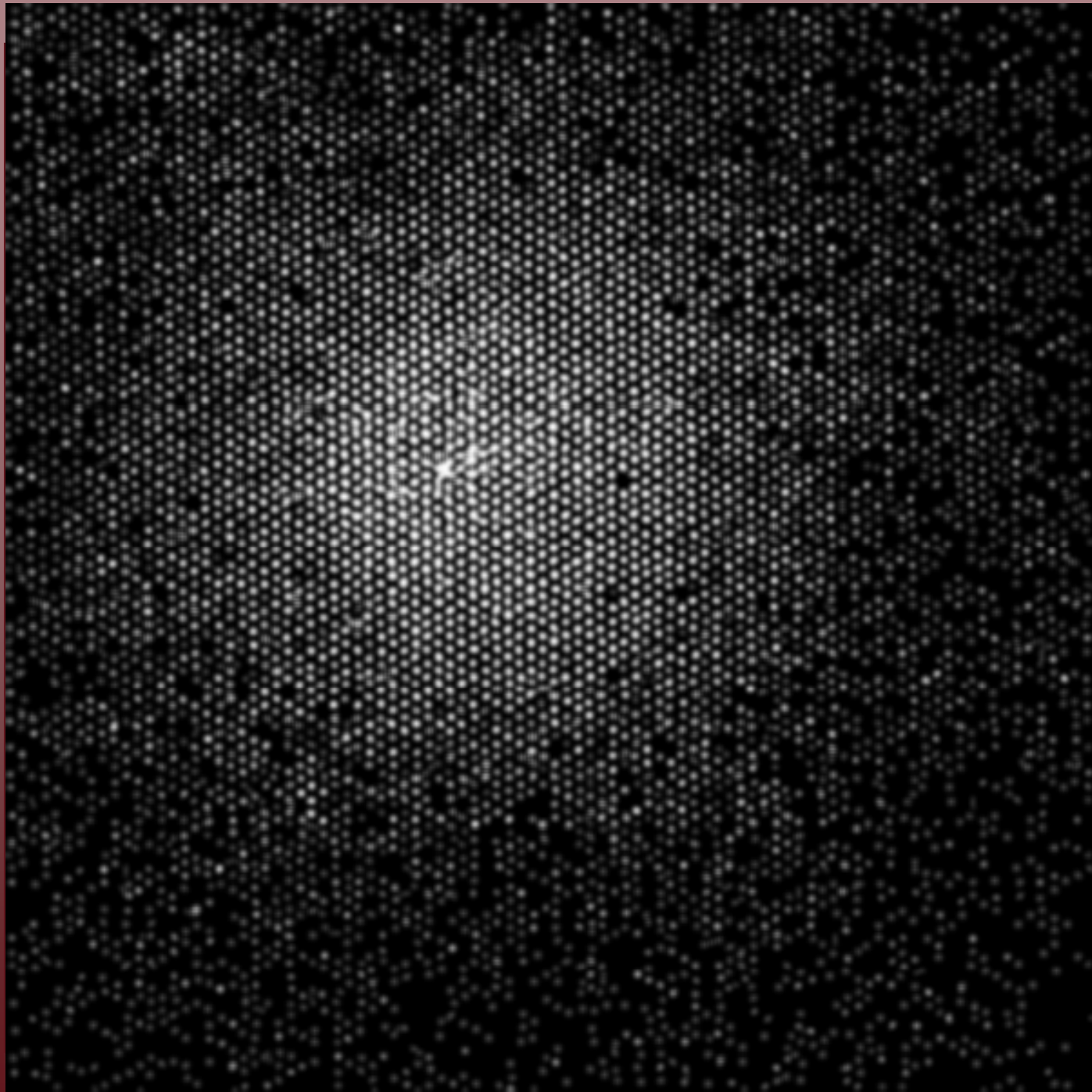
# Single photon operation

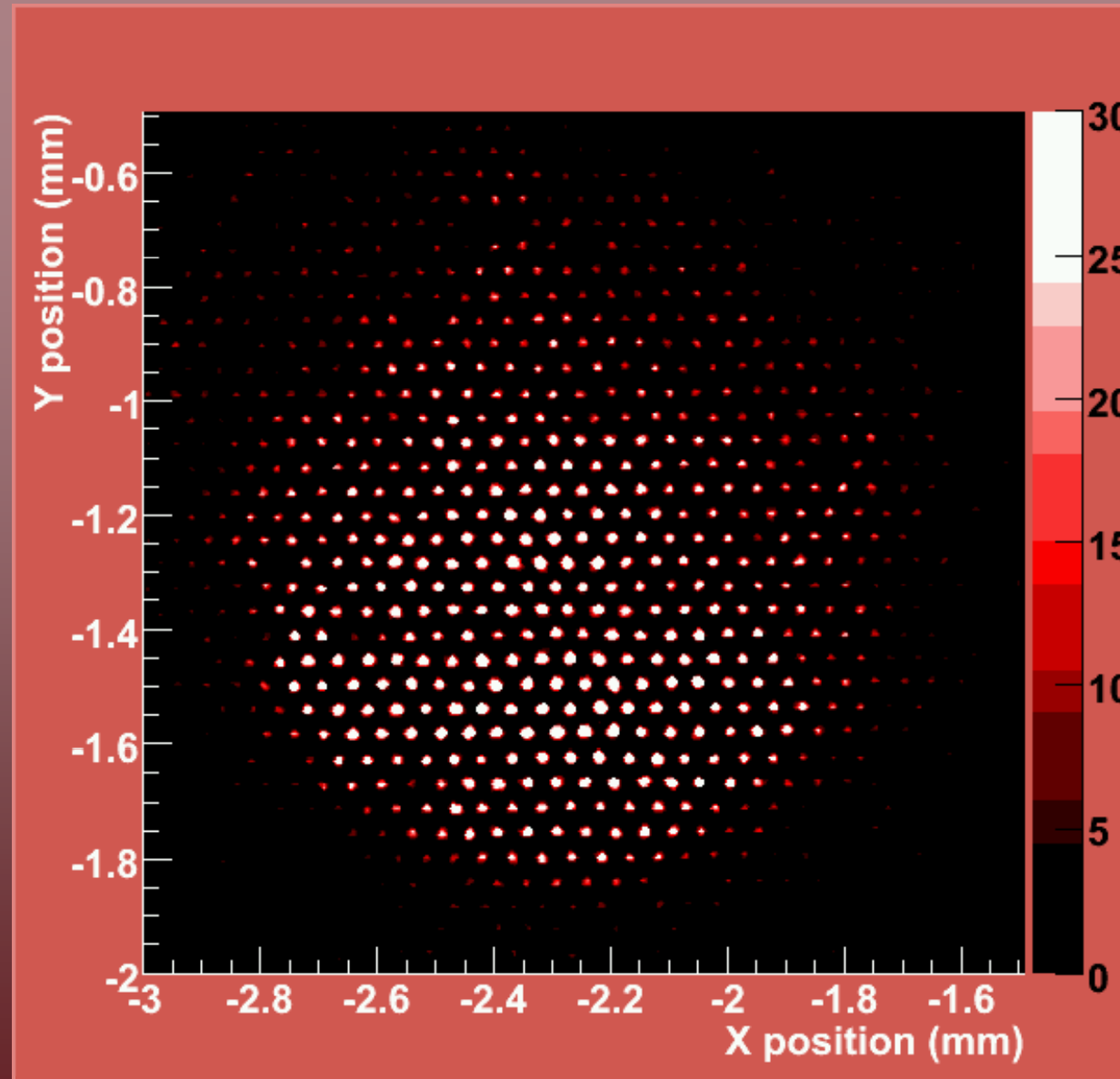


Single photon event topology

Semitransparent Photocathode

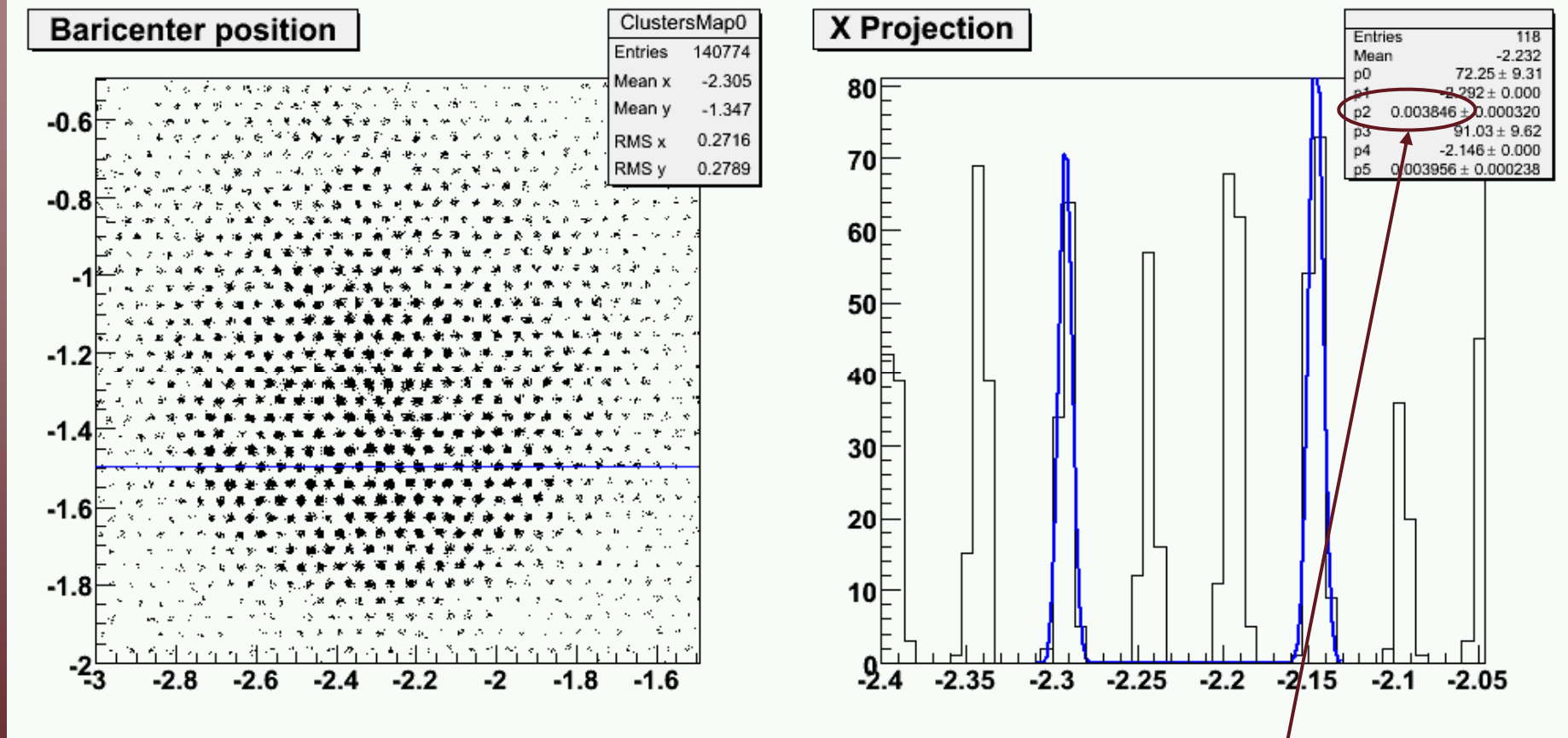






*“Self-portrait”* of the GEM amplification structure

# Intrinsic resolution of the read-out system



1 electron primary charge

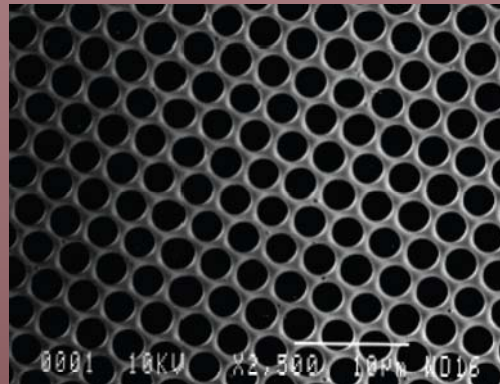
Semitransparent Photocathode



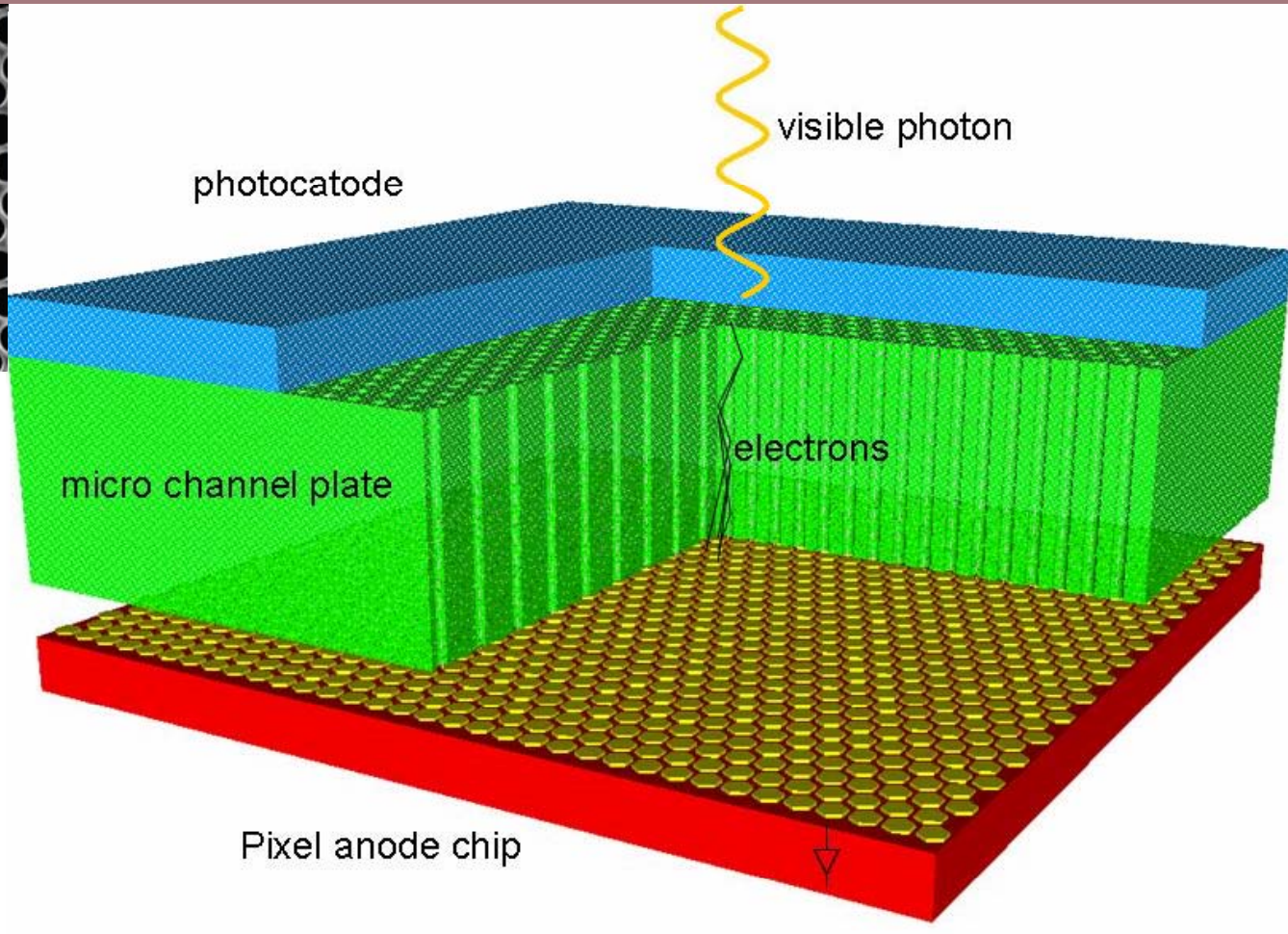
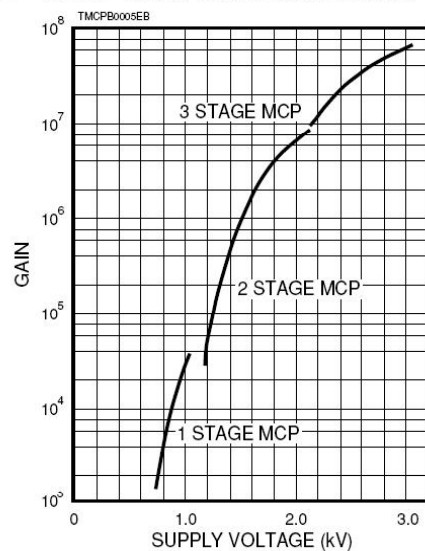
# Reflective Photocathode

# CIP

The use of Xpol as readout plane of a Micro Channel Plate coupled to a suitable photocathode in vacuum

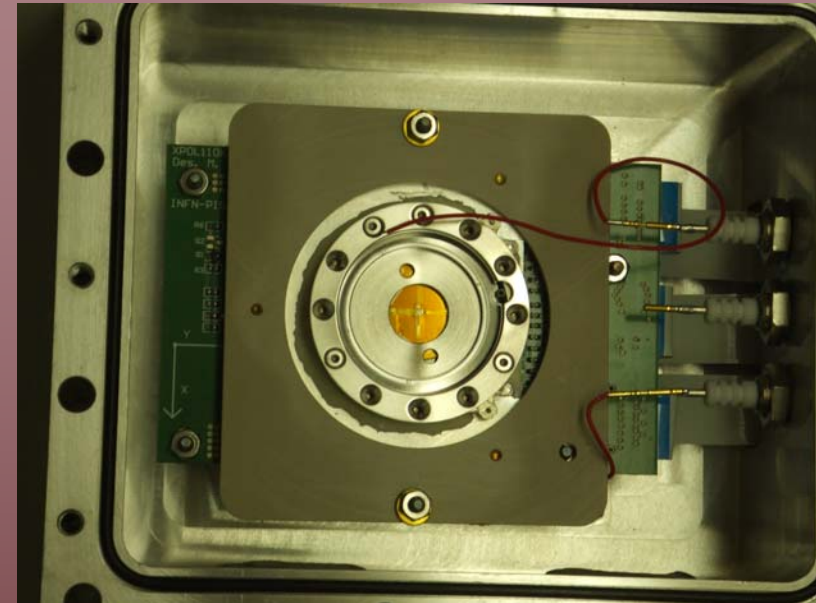
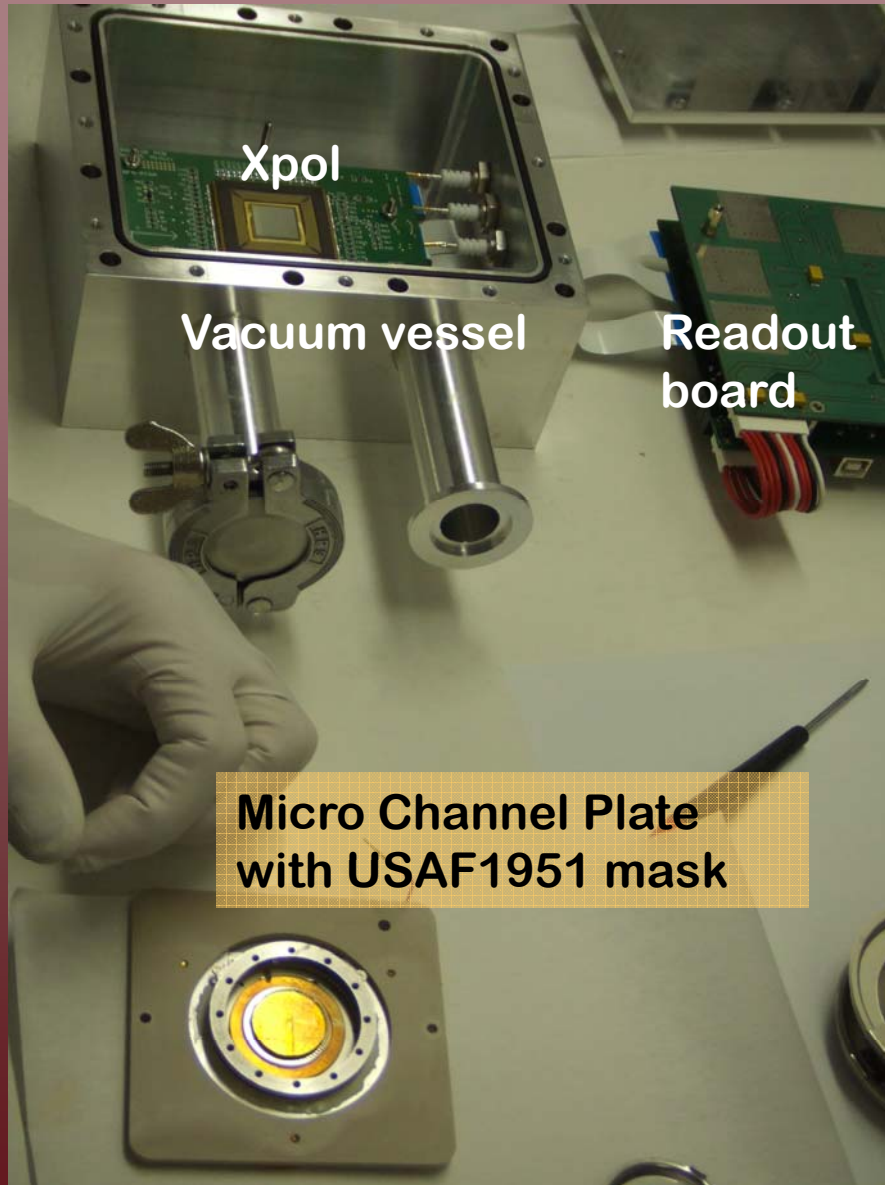


■ MCP Gain Characteristics

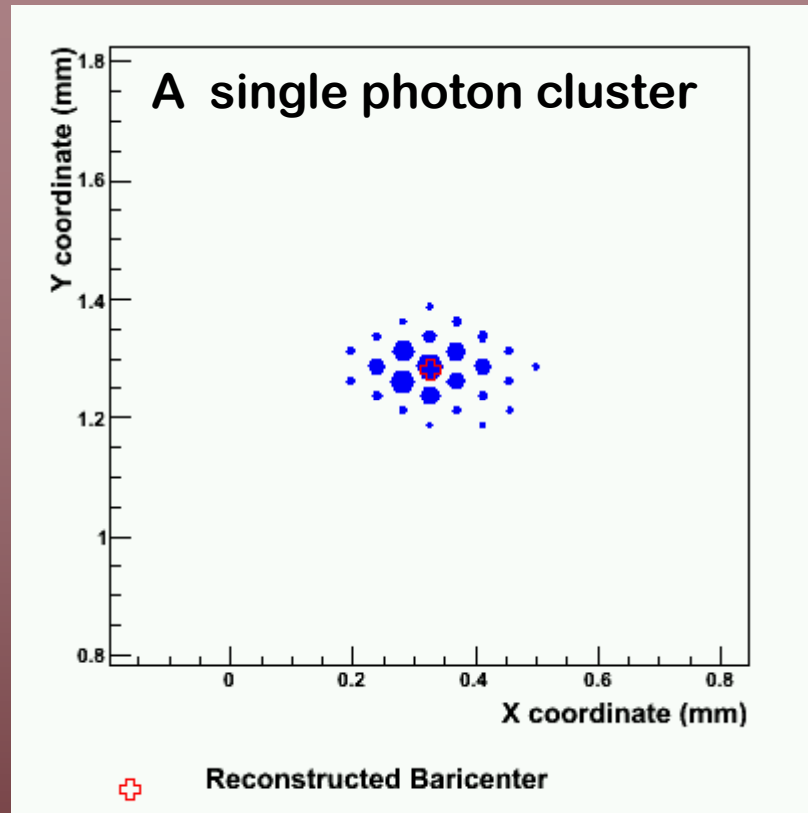




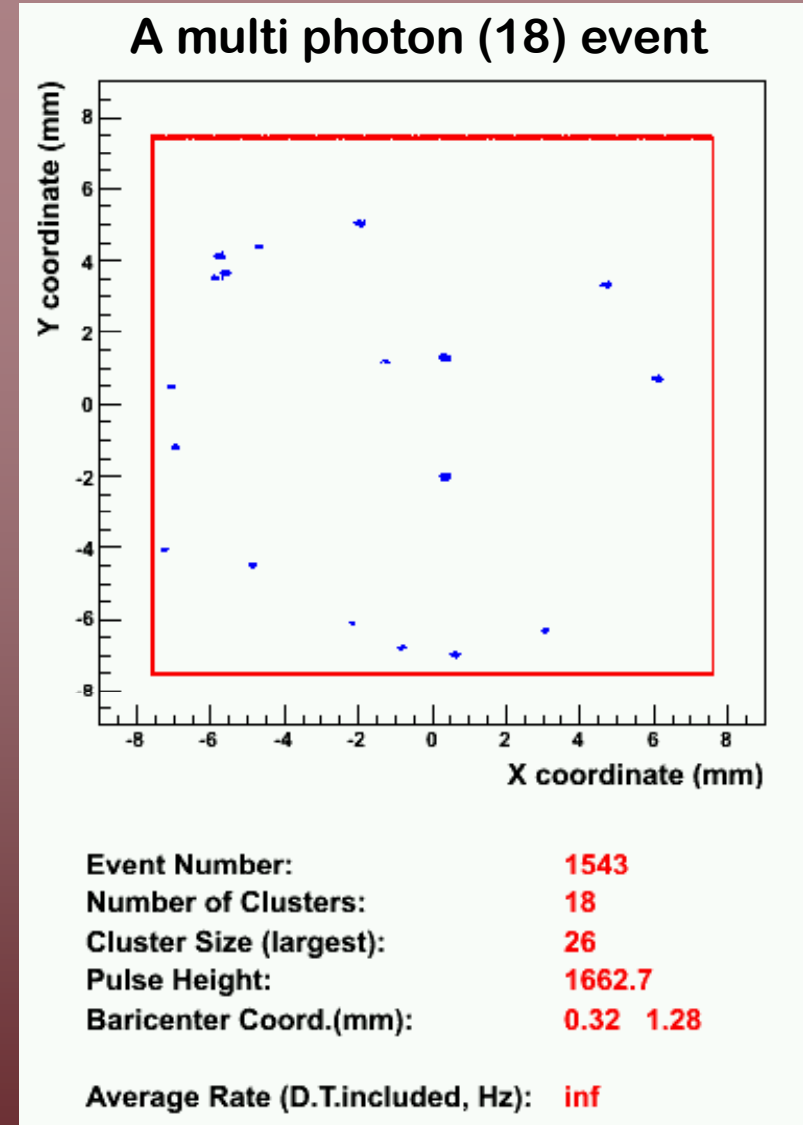
# First test in collaboration with Space Science Laboratory, Berkeley



# The events

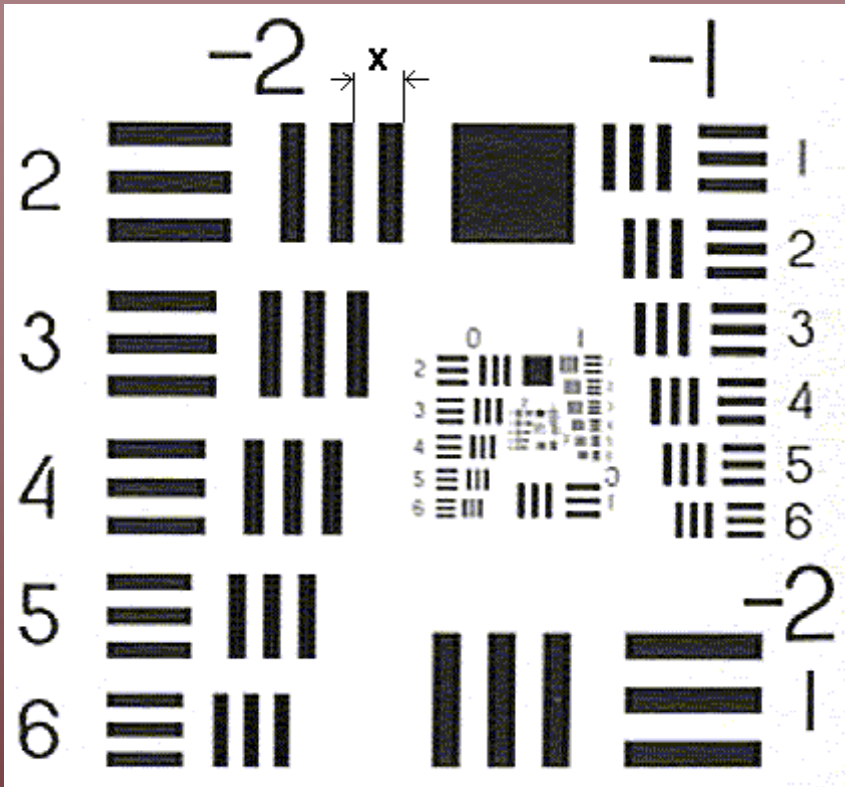


At low rate the detector records single photon events in window mode (<1K channels/frame, frame rate up to 10KHz).



Thanks to the high granularity, at high rate the detector can resolve up to few hundreds of photons/frame in full frame readout mode (1KHz frame rate, up to ~200KHz photon rate)

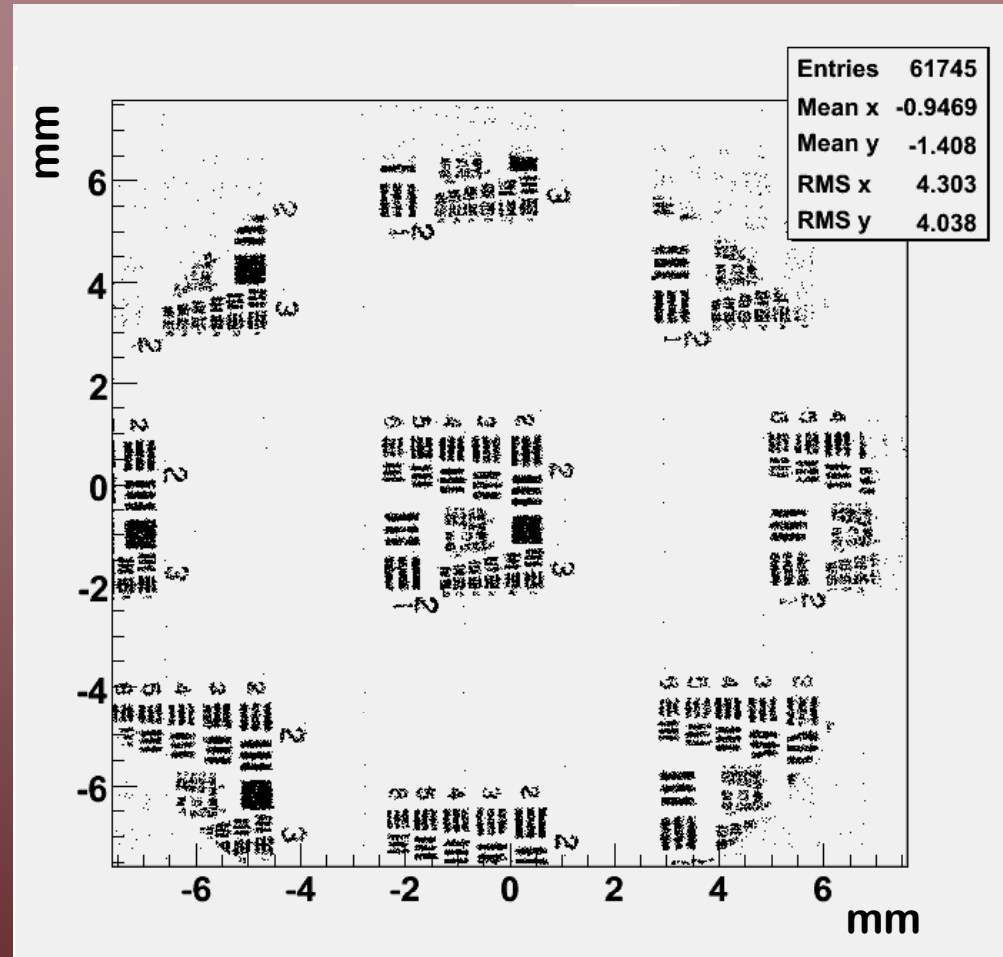
# The USAF1951 3-Bar Resolving Power Test Chart



Line pairs/mm =  $2^{n+(m-1)/6}$

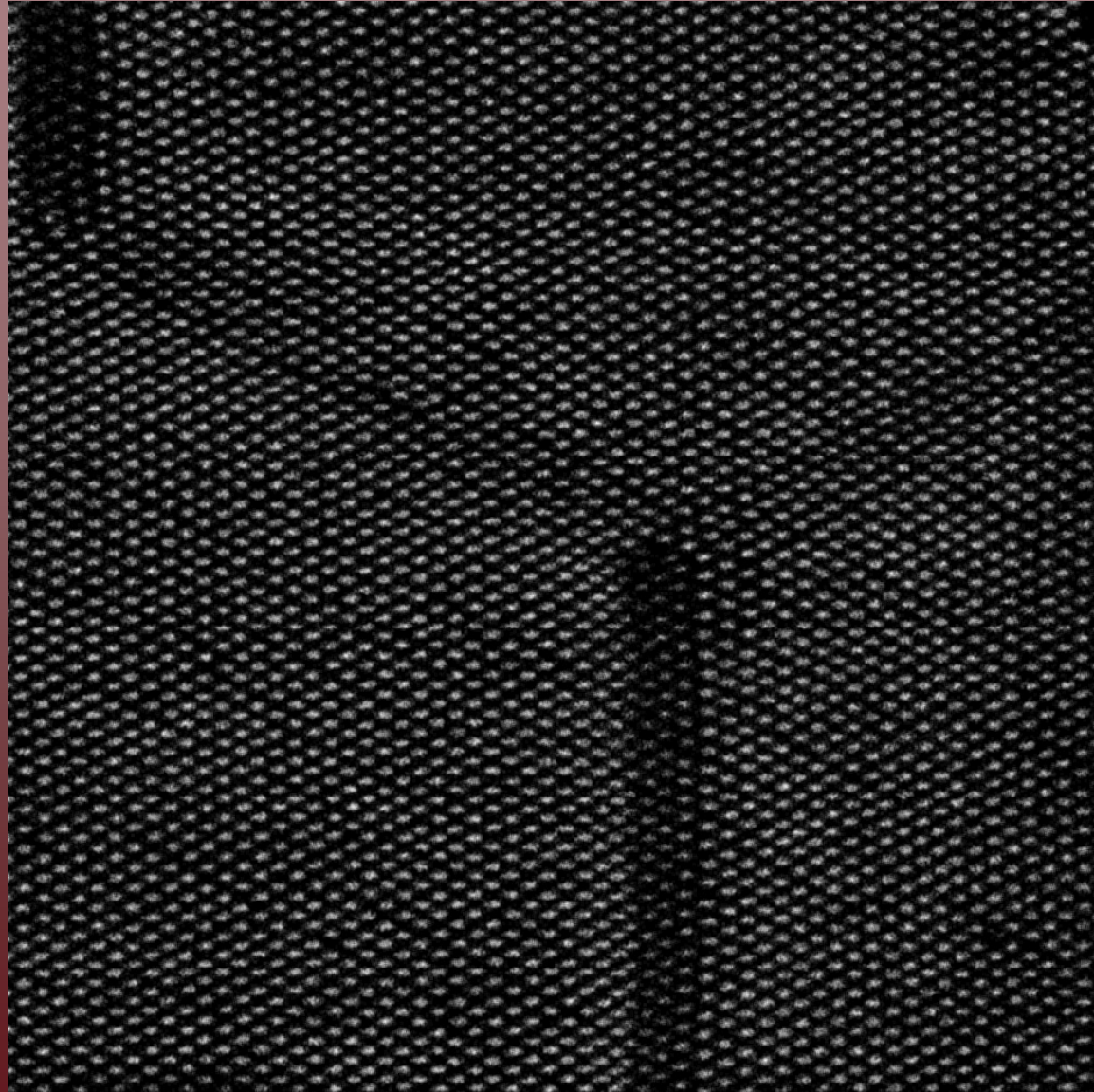
X (mm) =  $2^{-n-(m-1)/6}$

- Es. n=6, m=1, X=2<sup>-6</sup>=1/64=15.6μm
- n=6, m=2, X=2<sup>-(6+1/6)</sup>=1/72=14μm
- n=6, m=6, X=2<sup>-(6+5/6)</sup>=1/114=8.7μm
- n=7, m=1, X=2<sup>-7</sup>=1/128=7.8μm



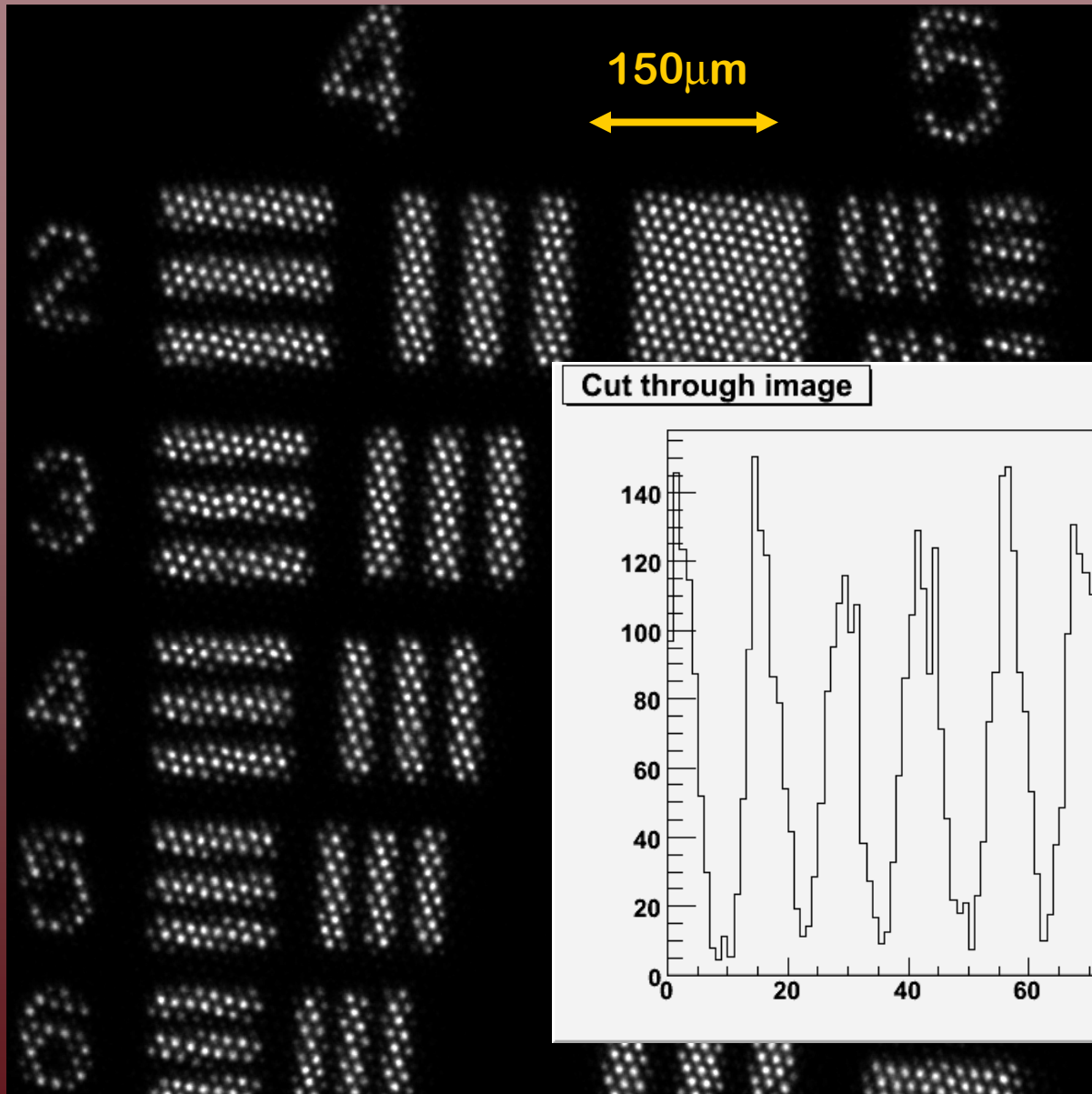
Full detector image

# MCP 12.5mm aperture 15mm pitch



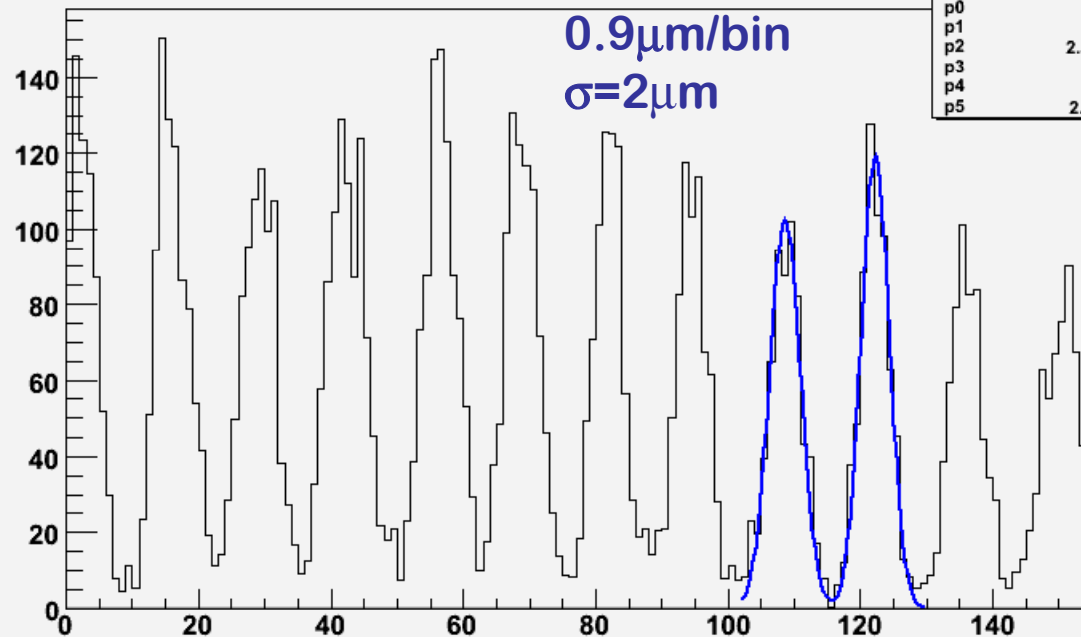
1mm<sup>2</sup> flat field image

# MCP $10\mu\text{m}$ aperture $12\mu\text{m}$ pitch



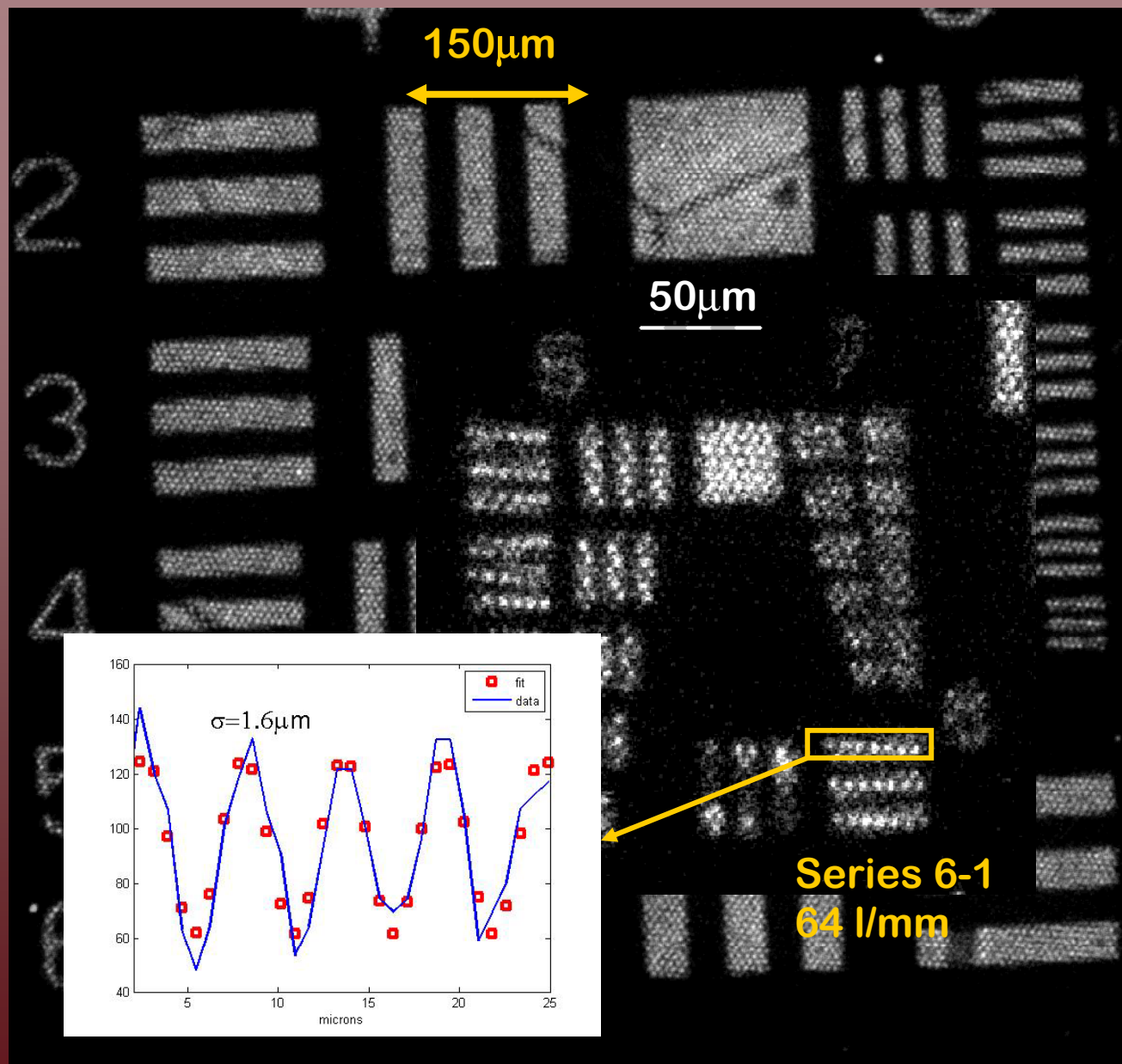
In a  $140\mu\text{m}$  square easily resolved 150 spots.  
The peak width includes the electron cloud width.

Cut through image



Cut	
Entries	154
Mean	69.98
p0	$102.3 \pm 5.0$
p1	$108.6 \pm 0.1$
p2	$2.335 \pm 0.070$
p3	$119.2 \pm 5.7$
p4	$122.2 \pm 0.1$
p5	$2.132 \pm 0.061$

# MCP 4 $\mu\text{m}$ aperture 5.5 $\mu\text{m}$ pitch



# The MCP+ASIC as visible light single photon imaging photomultiplier



The existing electronic devices with the required pitch and number of pixels are the CCDs (integrating, i.e do not allow photon by photon processing).

Photon → photo-electron →  $10^4$  gain → phosphor screen → intensified light → CCD

With the CMOS pixel ASIC the chain is

Photon → photo-electron →  $10^4$  gain → ASIC

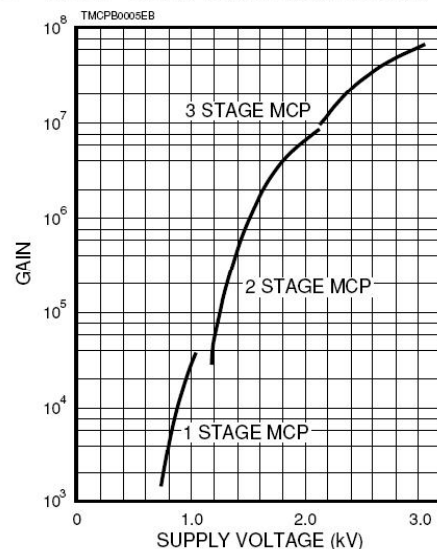
**The actual Xpol chip** is perfect for a detector with *enormous potentialities:*

- ultra high resolution ( $2\mu\text{m}$ ),
- 100ps timing capability
- reasonably fast (10KHz),
- single photon counting (noiseless)
- imaging (300 X 352 pixels, 15mmX15mm)

**Next step** → **Pixie counting chip**

- 600x800 Pixels at  $40\mu\text{m}$  pitch ( $25\text{X}30\text{mm}^2$  area)
- parallel counting 1MHz/pixel, several GHz/chip
- 5KHz Frame rate
- lower threshold

■ MCP Gain Characteristics



# applications

## Molecular Imaging

Optical spectroscopy (prismatic). Actual devices use a unimimensional delay line readout with photocathode and MCP. Xpol can improve by orders of magnitude the resolution.

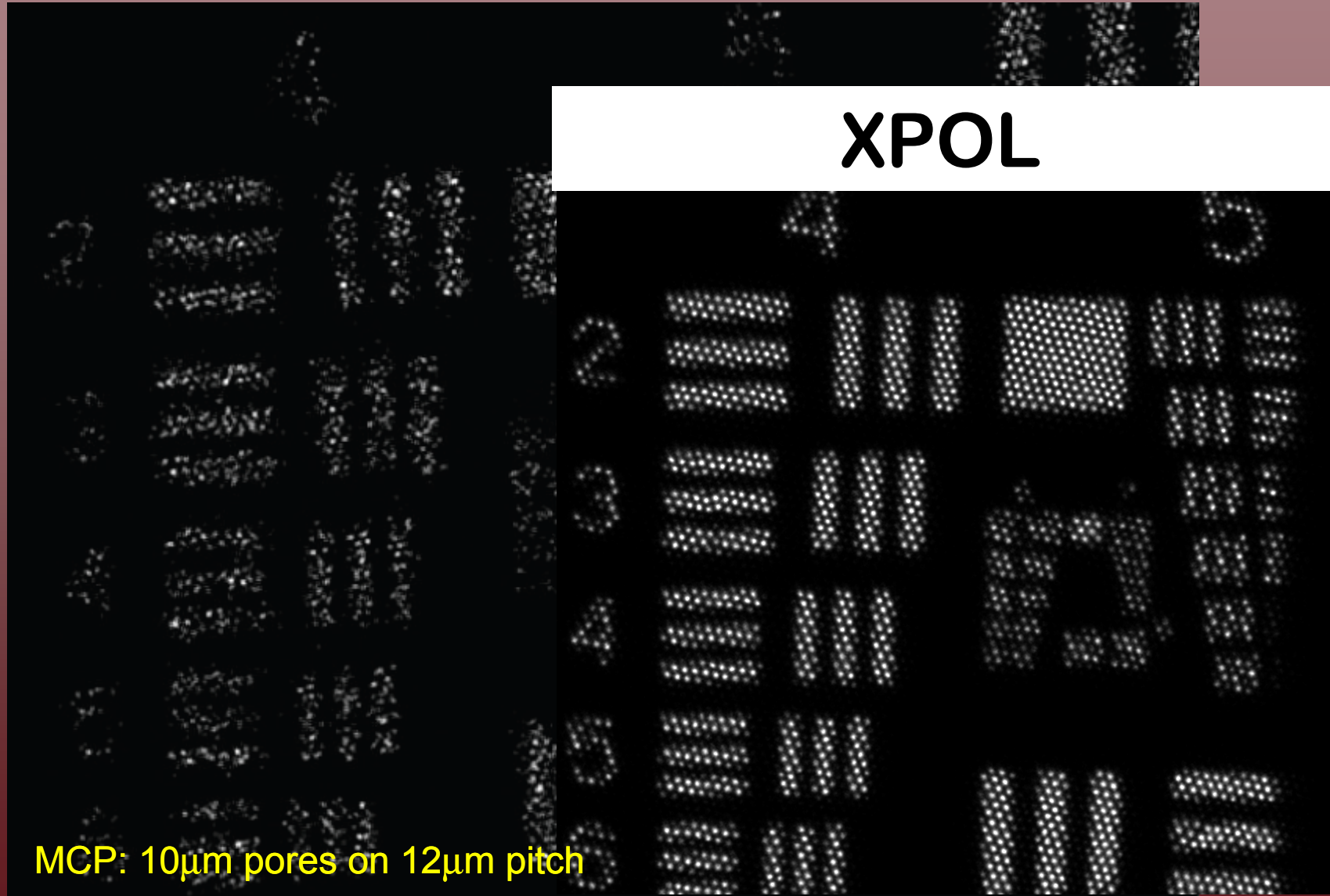
Mass spectroscopy: in this case there is no photocathode, the MCP is used directly as detector. The high Xpol spatial resolution allows very compact design.

Adaptive Optics: there is the need of a large detector (512X512 pixels) with  $\mu\text{m}$  resolution at high frame rate (>1KHz). SSL people propose 4 Medipix2 assembled together using the TOT function as ADC (pixel interpolation)



# TimePix with “analog” (TOT) interpolation

(John Vallerga – 9<sup>th</sup> IWoRiD: Erlangen July 2007)



MCP: 10 $\mu$ m pores on 12 $\mu$ m pitch

	Xpol	MEDIPIX2	PIXIE
<b>Technology:</b>	CMOS 0.18 $\mu\text{m}$	CMOS 0.25 $\mu\text{m}$	CMOS 0.18 $\mu\text{m}$
<b>Type:</b>	analog	digital (counting)	digital (counting+ ToT+Time stamp)
<b>Area:</b>	15x15=225mm <sup>2</sup> (1.1X)	14x14=196mm <sup>2</sup> (1X)	24x28=672mm <sup>2</sup> (3.4X)
<b>Pixel no.:</b>	105.600 (1.6X)	65.536 (1X)	480.000 (7.3X)
<b>Pixel density:</b>	470/mm <sup>2</sup> (1.4X)	330/mm <sup>2</sup> (1X)	720/mm <sup>2</sup> (2.2X)
<b>Pixel noise:</b>	50 electrons <i>ENC</i>	110 electrons <i>ENC</i>	50 electrons <i>ENC</i>
<b>Read-out scheme:</b>	asynchronous, synchronous	synchronous	synchronous
<b>Read-out trigger:</b>	self-trigger, internal, external	internal, external	internal, external
<b>Read-out mode:</b>	single pixel, window, full frame (8-16 nodes)	full frame (1 node)	Full frame (up to 200 nodes)
<b>Global threshold:</b>	2000 el. (unadjusted)	1000 el. (adjusted)	200 el. (auto- adjusted)
<b>Frame rate:</b>	10 kHz	1 kHz	5 kHz
<b>Event rate:</b>	$\sim 10^5/\text{s}$ ( $10^2$ ev. / frame)	$\sim 10^9/\text{s}$	$> 10^9/\text{s}$
<b>Resolution:</b>	$\sim 1\mu\text{m}$ (analog int.)	$\sim 15\mu\text{m}$ ( $55/\sqrt{12}$ )	$\sim 11\mu\text{m}$ ( $38/\sqrt{12}$ )
<b>Metal fraction:</b>	90%	13%	47%

# Conclusions

With devices like the ones presented the class of Gas Pixel Detectors has reached the level of integration, compactness and resolving power typical of solid state detectors.

Depending on type of electron multiplier, pixel and die size, electronics shaping time, analog vs. digital read-out, counting vs. integrating mode, many applications can be envisaged for this class of detectors.

- A UV photo-detector with single electron sensitivity and excellent imaging capabilities has been shown. It is based on a semitransparent or reflective CsI photocathode followed by a Gas Electron Multiplier foil and by a large area, custom, analog, VLSI ASIC. The high granularity and low noise of the read-out plane allows to reconstruct with  $4\ \mu\text{m}$  resolution the centroid of the single electron avalanche. The detector position resolution is at the moment limited by the  $50\ \mu\text{m}$  pitch of the GEM foil.

Main problem ion is back-flow control; possible solutions exist and are under development.

- Sealed operation has been demonstrated with very good gain stability after more than 40 days of irradiation. Stability of performance after 1 year.

- A Ultra High Spatial Resolution Single Photon Counting Imaging Detector based on a MCP coupled to the  $0.18\ \mu\text{m}$  CMOS Xpol ASIC (300 X 352 pixels,  $15 \times 15\ \text{mm}^2$ ).  $4\ \mu\text{m}$  pores of the MCP were resolved indicating a  $\sim 1\ \mu\text{m}$  spatial resolution capability of the device. Good uniformity, high signal/noise ratio, stable operation conditions were achieved with different MCPs. Images in single photon readout mode ( $\sim 10\ \text{KHz}$ ) and multi-photon mode (up to  $\sim 200\ \text{KHz}$ ) were acquired. Around 100 line pairs/mm can be resolved.