Implications of a 125 GeV Higgs for the MSSM

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The MSSM Higgs sector
 Implications for the pMSSM
 Implications for the cMSSM

4. Conclusion

Mainly based on Arbey+Battaglia+Mahmoudi+Quevillon+AD, arXiv:1112.3028 and work in preparation.

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In MSSM with two Higgs doublets: $H_1=inom{H_1^0}{H_1^-}$ and $H_2=inom{H_2^+}{H_2^0}$,

- $\ensuremath{\bullet}$ to cancel the chiral anomalies introduced by the new h field,
- give separately masses to d and u fermions in SUSY invariant way.

After EWSB (which can be made radiative: more elegant than in SM): three dof to make $W_L^{\pm}, Z_L \Rightarrow$ 5 physical states left out: h, H, A, H^{\pm} Only two free parameters at the tree level: $tan\beta, M_A$; others are:

$$\begin{split} \mathbf{M_{h,H}^2} = \frac{1}{2} \begin{bmatrix} \mathbf{M_A^2} + \mathbf{M_Z^2} \mp \sqrt{(\mathbf{M_A^2} + \mathbf{M_Z^2})^2 - 4\mathbf{M_A^2}\mathbf{M_Z^2}\mathbf{cos^2}2\beta} \\ \mathbf{M_{H^{\pm}}^2} = \mathbf{M_A^2} + \mathbf{M_W^2} \\ \mathbf{tan2}\alpha = \mathbf{tan2}\beta \left(\mathbf{M_A^2} + \mathbf{M_Z^2}\right) / (\mathbf{M_A^2} - \mathbf{M_Z^2}) \end{split}$$

Radiative corrections very important in the MSSM Higgs sector.

a huge effort from early 1990s up to now to calculate them...

Dominant corrections are due to top (s)quark at the one-loop level

$$\mathrm{M_h} \stackrel{\mathrm{M_A} \gg \mathrm{M_Z}}{
ightarrow} \mathrm{M_Z} |\mathrm{cos}2eta| + rac{3 ar{\mathrm{m}_t^4}}{2 \pi^2 \mathrm{v}^2 \sin^2 eta} \left| \ \log rac{\mathrm{M_S^2}}{ar{\mathrm{m}_t^2}} + rac{\mathrm{X_t^2}}{2 \mathrm{M_S^2}} igg(1 - rac{\mathrm{X_t^2}}{6 \mathrm{M_S^2}}igg)
ight|$$

Okada+Yamaguchi+Yanagida, Ellis+Ridolfi+Zwirner, Haber+Hempfling (1991)

depending on
$$tan\beta$$
, $M_S \equiv \sqrt{m_{\tilde{t}_1}m_{\tilde{t}_1}}$, $X_t = A_t - \mu/tan\beta$:
 $M_h^{max} \rightarrow M_Z + 30 - 50$ GeV...

• Full one-loop including all contributions $\mathbf{ ilde{t}}, \mathbf{ ilde{b}}, \mathbf{ ilde{q}}, \mathbf{\Phi}, \mathbf{ ilde{\ell}}, \chi, \mathbf{etc..}$

Brignole, Chankowski+Rosiek+Pokorski, Dabelstein, Pierce+Bagger+Matchev+Zhang (92-96)

RGE improved one–loop corrections

Carena+Espinosa+Quiros+Wagner, Haber+Hempfling+Hoang (95–96)

- Dominant two–loop corrections: $\mathcal{O}(\alpha_{\mathbf{t}}\alpha_{\mathbf{s}}), \mathcal{O}(\alpha_{\mathbf{b}}\alpha_{\mathbf{s}}), \mathcal{O}(\alpha_{\mathbf{t}}^{2}), \mathcal{O}(\alpha_{\mathbf{b}}^{2})$ Heinemeyer+Hollik+Weiglein, Brignole+Degrassi+Slavich+Zwirner (98–02)
- Dominant three–loop corrections: $\mathcal{O}(\alpha_t \alpha_s^2)$ contributes \approx 0.5 GeV Harlander+Kant+Mihaila+Steinhauser (2010)

Impact of missing corrections estimated below 1 GeV (HKMS)!

Radiative corrections implemented in two different ways in general:

- On–shell scheme (OS) as in the MSSM Higgs code FeynHiggs Heinemeyer+Hollik+Weiglein+Han+....
- DR scheme à la BDSZ as in RGE codes Softsusy, Spheno, Suspect Slavich, Allanach, Porod, Kneur+Moultaka+AD

Difference between the two approaches: $\Delta M_{
m h} pprox 2$ GeV in general, assumed to be the theoretical+"experimental" uncertainty on M_h no–mixing case: $m M_{H}\!\lesssim\!120$ GeV; max–mixing case: $m M_{H}\!\lesssim\!135$ GeV



In the following, I assume that a 125 ± 2 GeV Higgs has been observed, (no choice anyway as only $122.5\!\le\!M_{
m h}\!\le\!127.5$ GeV is now allowed...)



and that it is the one of the MSSM... I will ask the following questions:

- what are the implications in unconstrained and constrained MSSMs?
- what happens to MSSM Higgs sector if one includes other constraints?
- could one increase the rate for the ${f h} o \gamma\gamma$ signal?
- what are the implications for sparticle searches (mainly stops)?

The mass value 125 GeV is rather large for the MSSM h boson, \Rightarrow one needs from the very beginning to almost maximize it... Maximizing M_h is maximizing the radiative corrections; at 1-loop:

$$\mathrm{M_h} \stackrel{\mathrm{M_A} \gg \mathrm{M_Z}}{
ightarrow} \mathrm{M_Z} |\mathrm{cos}2eta| + rac{3 ar{\mathrm{m}_t^4}}{2 \pi^2 \mathrm{v}^2 \mathrm{sin}^2 \, eta} \left| \ \log rac{\mathrm{M_S^2}}{ar{\mathrm{m}_t^2}} + rac{\mathrm{X_t^2}}{2 \mathrm{M_S^2}} \left(1 - rac{\mathrm{X_t^2}}{6 \mathrm{M_S^2}}
ight)
ight|$$

- decoupling regime with $\mathbf{M}_{\mathbf{A}}\!\sim\!\mathcal{O}$ (TeV);
- \bullet large values of $\tan\!\beta\gtrsim 10$ to maximize tree-level value;
- ullet maximal mixing scenario: $\mathbf{X_t} = \sqrt{6} \mathbf{M_S}$;
- \bullet heavy stops, i.e. large $M_{\mathbf{S}}\!=\!\sqrt{m_{\tilde{t}_1}m_{\tilde{t}_2}};$

we choose at maximum $M_{
m S}\!\lesssim\!3$ TeV, not to have too much fine-tuning....

Do the complete job as in real life:

- small contributions of entire SUSY spectrum: $\Phi, \chi^{\pm}_i, \chi^0_i, \tilde{q}_i, \tilde{l}_i, \tilde{g}_{...}$
- complete radiative corrections up to two–loops

We use the RGE codes Suspect Kneur+Moultaka+AD and Softsusy Allanach which implement the known radiative corrections in the \overline{DR} scheme.

To evaluate M_h , perform a full scan of the MSSM parameter space; too complicated in the general MSSM as there are 105 free parameters

 \Rightarrow work in the phenomenological MSSM or pMSSM:

- no CP or flavor-violation: no new phase and diagonal ${f { extbf{m}}}, A$ matrices,
- universal first and second generation sfermions to cope with flavor.

Only 22 free parameters: $tan\beta$, M_A , μ , $M_{1,2,3}$, $m_{\tilde{f}_L}$, $m_{\tilde{f}_R}$, A_f and only a few of them will play and important role in the Higgs sector...

Perform a full and fine scan of the pMSSM parameter space:

 $\mathbf{1} \leq \tan\beta \leq \mathbf{60}\,,\; \mathbf{50}\,\mathrm{GeV} \leq \mathbf{M_A} \leq \mathbf{3}\,\mathrm{TeV},\; -\mathbf{9}\,\mathrm{TeV} \leq \mathbf{A_f} \leq \mathbf{9}\,\mathrm{TeV},$

 $50\;{\rm GeV}\!\le\!m_{{\bf \tilde{f}_L}},m_{{\bf \tilde{f}_R}},M_3\le 3\;{\rm TeV},50\;{\rm GeV}\!\le\!M_1,M_2,|\mu|\!\le\!1.5\;{\rm TeV}$

• determine the regions of parameter space where $123 \le M_h \le 127$ GeV (2 GeV uncertainty includes both "experimental" and "theoretical" error)

• require h to be SM–like: $\sigma(h) \times BR(h \! \rightarrow \! VV) \! \gtrsim 0.9 H_{\rm SM}$

(we will also consider the possibility that H is the $H_{\rm SM}$, see later).

Main results:

- \bullet Large $M_{\mathbf{S}}$ values needed:
- $M_{\mathbf{S}} pprox 1$ TeV: only maximal mixing
- $M_{\rm S}\approx 3$ TeV: only typical mixing.
- Large tan β values favored but tan $\beta\!\approx\!3$ possible if $M_{\rm S}\!\approx\!3\text{TeV}$
- What about other benchmarks?

Carena+Heinemeyer+Wagner+Weiglein

- small $\alpha_{\rm eff}$ scenario with $g_{\rm hbb}\approx$ 0: ruled out by LHC/Tevatron data.
- gluophobic h with $g_{hgg} \ll g_{H_{\rm SM}gg}$ ruled out by $4\ell^+, \gamma\gamma$ signals at LHC (difficult to achieve as \tilde{t}_1 heavy..).
- no SUSY regime with light sparticles: BR(h $\rightarrow \chi_1^0 \chi_1^0)$ should be small...
- max and no-mix need to be updated!



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update of $[M_{\mathbf{A}},tan\beta]$ propaganda plot is desperately needed!

Besides LEP2 and $A/H/h \rightarrow \tau \tau$ searches, one must now include:

- ullet combined ATLAS+CMS of au au and $t o bH^+$ searches at low $M_{f A}$
- \bullet the limit $122.5\!\leq\!M_h\!\leq\!127.5$ GeV from $H_{\rm SM}$ searches
- \bullet constraints from flavor: at least (direct!) limits from $B_{\mathbf{s}}\!\rightarrow\!\mu\mu$...



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Are we really in decoupling regime?

- \bullet are small values of M_A allowed? \bullet can H be the SM-like Higgs boson? YES!, if no other constraints than:
- $M_{H}\approx 125\pm 2~\text{GeV}$
- $g_{HVV} \approx g_{H_{SM}VV}$

Heinemeyer+Stal+Weiglein

$$\begin{split} \mathbf{M_A} &\approx & \mathbf{100} \ \mathbf{GeV}, \mathbf{tan}\beta \approx \mathbf{6-10}, \\ \mathbf{M_S} &\approx & \mu \approx & \mathbf{1} \ \mathbf{TeV}, \mathbf{X_t} \approx & \sqrt{\mathbf{6}}\mathbf{M_S}, \\ &\Rightarrow \mathbf{M_H} \approx & \mathbf{125} \ \mathbf{GeV} \text{ ; } \mathbf{M_h} \approx & \mathbf{98} \ \mathbf{GeV!} \end{split}$$

[in ABDMQ scan, only very few points (20 out of 10^6 valid) satisfy conditions but they are all ruled out by $b\to s\gamma$

\Rightarrow only h SM–like is likely...

maybe needs more detailed studies?]





Can one change the h prod rates?

• suppress g_{hWW} or g_{hbb} couplings \Rightarrow loose Wh $\!\!\!\!\!\rightarrow \ell \nu b \bar{b}$ Tevatron signal

• suppress g_{hZZ} or g_{htt} (incr. g_{hbb}) \Rightarrow loose $h \rightarrow ZZ \rightarrow 4\ell$ ATLAS signal hard to change tree-level couplings

Only change is the $h\gamma\gamma$ coupling: increase to explain $\gamma\gamma$ LHC excess?

- light stau's and large $\mu an\!eta$
 - Carena+Gori+Shah+Wagner
- light $\chi_{\mathbf{1}}^{\pm}$ in non-univ MSSM

Driesen+Illana+Hollik+AD

- possibility of light \widetilde{t} :
- \Rightarrow max-mixing: $\sigma(\mathbf{gg} \rightarrow \mathbf{h})$ suppressed.
- \Rightarrow no mixing: yes, but stops too heavy.

Arvanitaki+Villadoro,AD

BMSSM? Ellwanger etal, King etal.,
 Kraml+Jiang+Gunion · · · · see J. Gunion's talk

TLAS Preliminary 2011 Data CMS Preliminary. √s = 7 Te\ Combined. L = $4.6-4.7 \text{ fb}^{-1}$ Observed · Expected _dt = 1.0-4.9 fb⁻ ±1σ s = 7 TeV +2 σ ATLAS 300 100 110 120 130 140 150 400 500 6 M, [GeV] n" (GeV/c²) Higgs bosor $\tan\beta = 60$ 1400 1200 μ [GeV] 1000 800 600 300 350 400 450 500 m_{L3} [GeV] $\sigma(gg \to \gamma\gamma)|_{\rm MSSM}$ $\sigma(gg \to \gamma \gamma)|_{\rm MSSM}$ $\tan \beta = 2.5$ $\tan\beta = 50$ $M_A = 1 \text{ TeV}$ $M_A = 1 \text{ TeV}$ 1.20.4 $m_{i_1} = 200 \text{ GeV}$ 0.21000 15002000 $1\,000$ 5001500 $X_t \, [\text{GeV}]$ $-\mu$ [GeV]

Fevatron Run II Preliminarv. L \leq 10.0 fb⁻¹

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How light superparticles can be in pMSSM with a 125 GeV Higgs?

- ullet non-universal gaugino masses and μ parameter unconstrained,
- non-universal sfermions masses: decouple sleptons from squarks do not affect $M_h \Rightarrow$ light $\chi^{\pm}_{1,2}, \chi^0_{1...4}, \tilde{\ell}^{\pm}, \tilde{\nu}$ beyond LEP2 possible!
- first/second gen. squarks as well as gluinos can be very heavy...

But not main player stop! How light or heavy can the stops be?



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2. Implications in pMSSM: high scale SUSY

The scale $M_{
m S}$ seems to be large. There are two extreme possibilities

• Split SUSY: allow fine-tuning scalars (including H_2) at high scale gauginos-higgsinos at weak scale (unification+DM solutions still OK) $M_h \propto \log(M_S/m_t) \rightarrow \text{large}$ Arkani-Hamed+DimopoulosGiudice, Romanino

• SUSY broken at the GUT scale...

give up fine-tuning and everything else still, $\lambda\!\propto\!M_{H}^{2}$ related to gauge cplgs $\lambda(\tilde{m})\!=\!\frac{\mathbf{g_{1}^{2}}(\tilde{m})\!+\!\mathbf{g_{2}^{2}}(\tilde{m})}{8}(1+\delta_{\tilde{m}})$... leading to $M_{H}\!=\!\!120\!-\!140$ GeV ... Hall+Nomura, Giudice+Strumia Bernal+Slavich+AD

In both cases small $an\!eta$ needed...



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Constrained MSSMs are interesting from model building point of view:

- provide concrete schemes for supersymmetry breaking
- solve some problems of unconstrained MSSM: flavor, CPV, universality,
- reduce number of input parameters and are thus more predictive

Prototype model: the minimal supergravity model (mSUGRA).

- Underlying assumption: SUSY–breaking occurs in a hidden sector communicating with visible sector through gravitational interactions,
- parameters obey a set of boundary conditions at $M_{
 m GUT}\!pprox\!10^{16}\,$ GeV
- universal soft terms emerge if the interactions are "flavor-blind"

 $\Rightarrow \text{ only 4.5 inputs: } \tan\beta \ , \ m_{1/2} \ , \ m_0 \ , \ A_0 \ , \ \ \text{sign}(\mu)$ In GMSB, SSB transmitted to MSSM fields via SM gauge interactions. Minimal inputs: $\tan\beta \ , \ \text{sign}(\mu) \ , \ M_{mes} \ , \ \Lambda_{SSB} \ , \ N_{mess \ fields}$ In AMSB, SSB in hidden sector transmitted via (super-Weyl) anomalies. Minimal inputs: $m_0 \ , \ m_{3/2} \ , \ \tan\beta \ , \ \text{sign}(\mu)$ Using Suspect+Softsusy, perform scans of the models parameter space and confront them with LHC constraint $123 \ GeV \leq M_h \leq 127 \ GeV$

The following ranges are considered for the model input parameters besides $1 \le aneta \le 60$ and sign(μ) = ± 1 that are common to all:

 $\begin{array}{ll} \mbox{mSUGRA:} & \mbox{50GeV} \leq m_0 \leq \mbox{2TeV}, & \mbox{50GeV} \leq m_{1/2} \leq \mbox{3TeV}, \ |A_0| \leq \mbox{9TeV}; \\ \mbox{mGMSB:} & \mbox{10TeV} \leq \Lambda \leq \mbox{1000 TeV}, \ \mbox{1} \leq M_{\rm mes} / \Lambda \ \leq \ \mbox{10}^{11}, \ \ \mbox{N}_{\rm mess} = \mbox{1;} \\ \mbox{mAMSB:} & \mbox{1 TeV} \leq m_{\frac{3}{2}} \leq \mbox{100TeV}, \mbox{50 GeV} \leq m_0 \leq \mbox{2 TeV}. \end{array}$

In mSUGRA we further consider the following (over-constrained) cases:

• no–scale:
$$\mathbf{m_0} = \mathbf{A_0} = \mathbf{0}$$

• cNMSSM:
$$\mathrm{m_0}=0, \mathrm{A_0}=-rac{1}{4}\mathrm{m_{1/2}}$$

• vcMSSM: $\mathbf{m_0} = \mathbf{A_0}$

as well as as the less constrained non-universal Higgs mass model:

$$\bullet$$
 NUHM: $m_{1/2}, m_0, A_0$ and m_{H_u}, m_{H_d}

In mSUGRA case and its variants, we impose in addition bounds from:

- correct relic density of DM neutralino as measured by WMAP,
- constraints from flavor physics: ${f b}
 ightarrow {f s} \gamma, {f B}_{f s}
 ightarrow \mu \mu$,
- constraints from heavy MSSM Higgs production at the LHC.

Less freedom for $A_t \Rightarrow M_h$ is much more constraining!

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also: Buchmuller etal, Drapper etal., Baer etal., Raidal etal., Li etal, Roszkowski etal... and in other (many!!) BMSSM including NMSSM scenarios, talk of Jack Gunion....

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4. Conclusions

A 125 GeV Higgs provides information on BSM and SUSY in particular:

- $\bullet~M_{H}\,{=}\,119$ GeV would have been a boring value: everybody OK
- $\bullet~M_{H}\!=\!145$ GeV would be a devastating value: everybody dead
- $M_{
 m H}\!pprox\!125$ GeV is Darwinian: (natural) selection among models..

Many questions remain:

- is the 125 GeV Higgs really there? any wrong cable connection?
- if yes, is it really SM–like? What about the $\gamma\gamma, 4\ell^{\pm}, bar{b}$ rates?
- if yes, SUSY spectrum heavy; except maybe for weakly interacting sparticles and also stops \Rightarrow more focus on them in SUSY searches!

Some answers in July or December. More complete answer later! My personal feeling or bet:

- a ($\mathbf{7.3}\sigma)$ Higgs in 2012, Higgstoric year!
- a stop and a chargino in 2015: my favorite/best-guess SUSY signal:

$$pp \rightarrow \tilde{t}_1 \tilde{t}_1 \rightarrow b \chi_1^+ \bar{b} \chi_1^- \rightarrow b \bar{b} e \mu + E_T$$

– following years, search for $gg \to \widetilde{t}_1 \widetilde{t}_1$ and measurement of $A_t...$

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