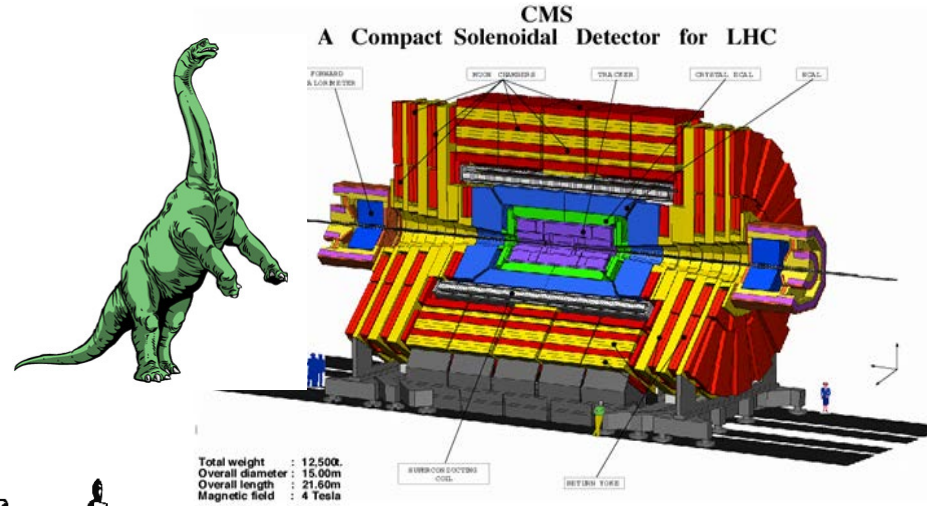
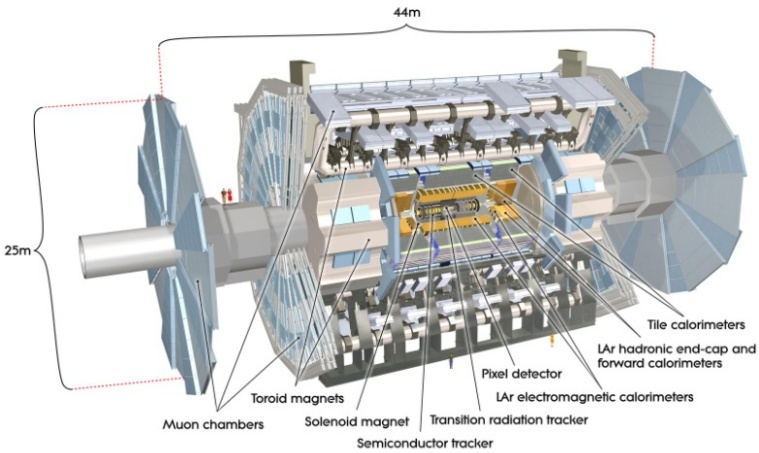


SUSY w/o Prejudice @ LHC-7 & -8



3/29/12

Searches for SUSY @ the LHC keep **going and going** but have not found any signals (**yet**)...

However they **ARE** eating into **a lot** of the **model parameter space**...

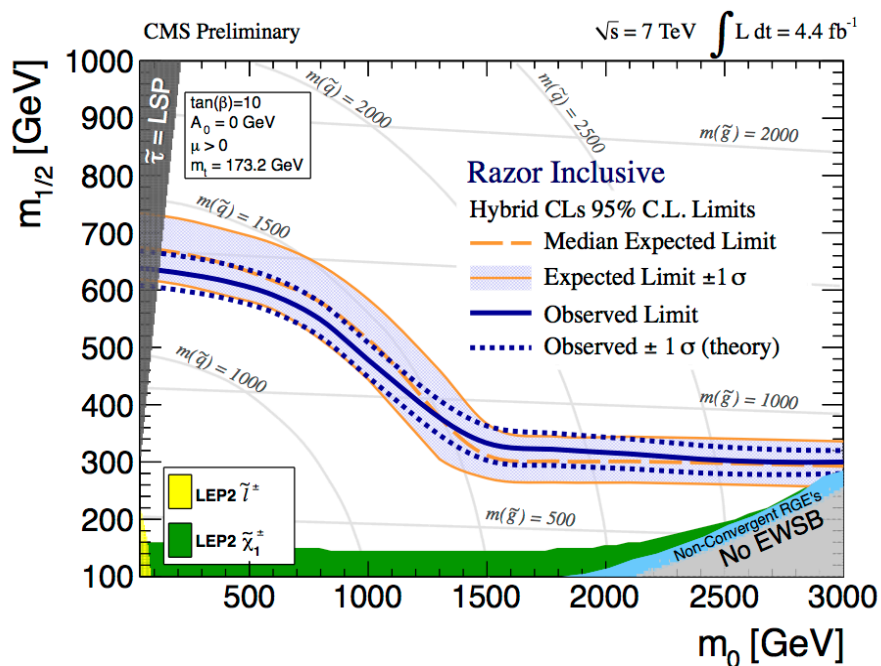
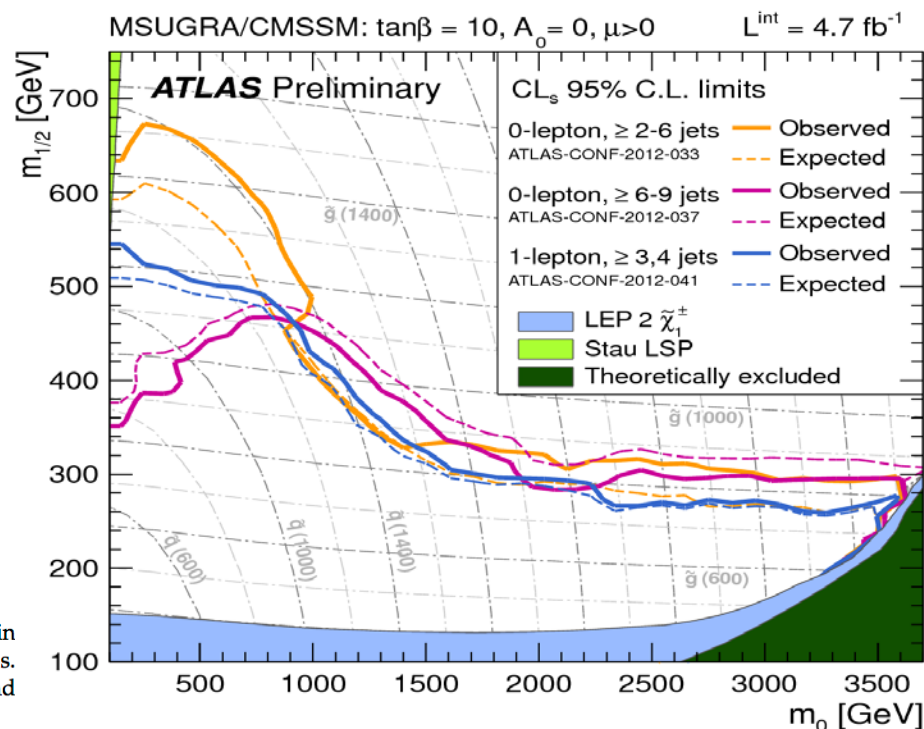


Figure 10: Observed (solid curve) and median expected (dot-dashed curve) 95% CL limits in the $(m_0, m_{1/2})$ CMSSM plane with $\tan \beta = 10$, $A_0 = 0$, $\text{sgn}(\mu) = +1$ from the razor analysis. The \pm one standard deviation equivalent variations in the uncertainties are shown as a band around the median expected limit.



Two New pMSSM Scans: Neutralino & Gravitino LSPs

(via SOFTSUSY
+SuSpect + FeynHiggs)

$$100 \text{ GeV} \leq m_{\text{L}_{1,2,3}} \leq 4 \text{ TeV}$$

$$400 \text{ GeV} \leq m_{\text{Q}_{ud1,2}} \leq 4 \text{ TeV} \quad 200 \text{ GeV} \leq m_{\text{Q}_{ud3}} \leq 4 \text{ TeV}$$

$$50 \text{ GeV} \leq |M_1| \leq 4 \text{ TeV} \quad 100 \text{ GeV} \leq |M_2, \mu| \leq 4 \text{ TeV}$$

$$400 \text{ GeV} \leq M_3 \leq 4 \text{ TeV} \quad |A_{t,b,\tau}| \leq 4 \text{ TeV}$$

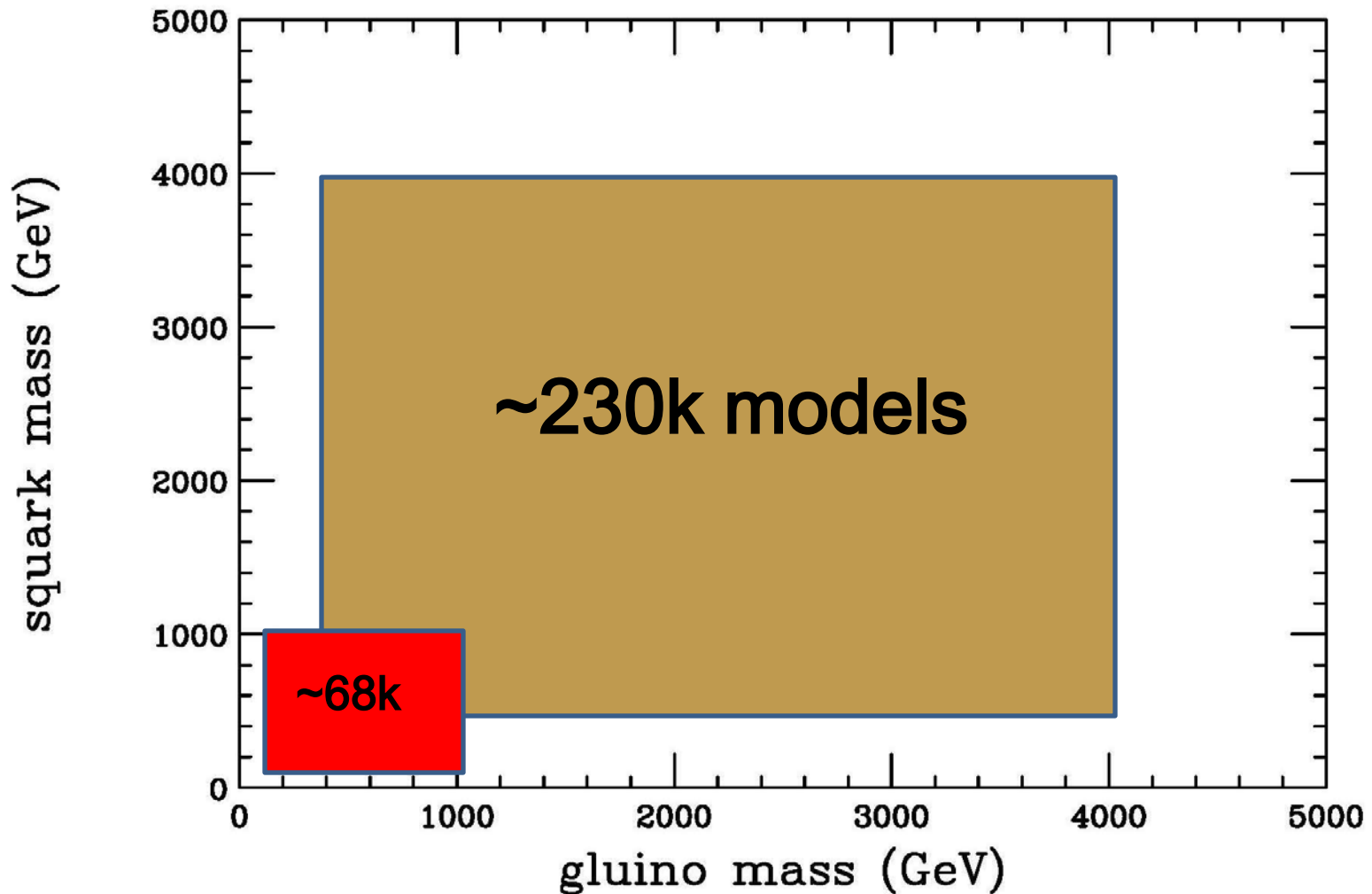
$$100 \text{ GeV} \leq M_A \leq 4 \text{ TeV}$$

$$1 \leq \tan\beta \leq 60$$

→→ For the gravitino LSP: $1 \text{ eV} \leq m_G \leq 1 \text{ TeV}$ (log scan)

- Apply all the usual non-LHC + all LHC non-MET constraints (as of 12/1/2011). Additional complexities occur, eg, BBN constraints for the **gravitino** LSP case

A much larger volume needs to be explored...



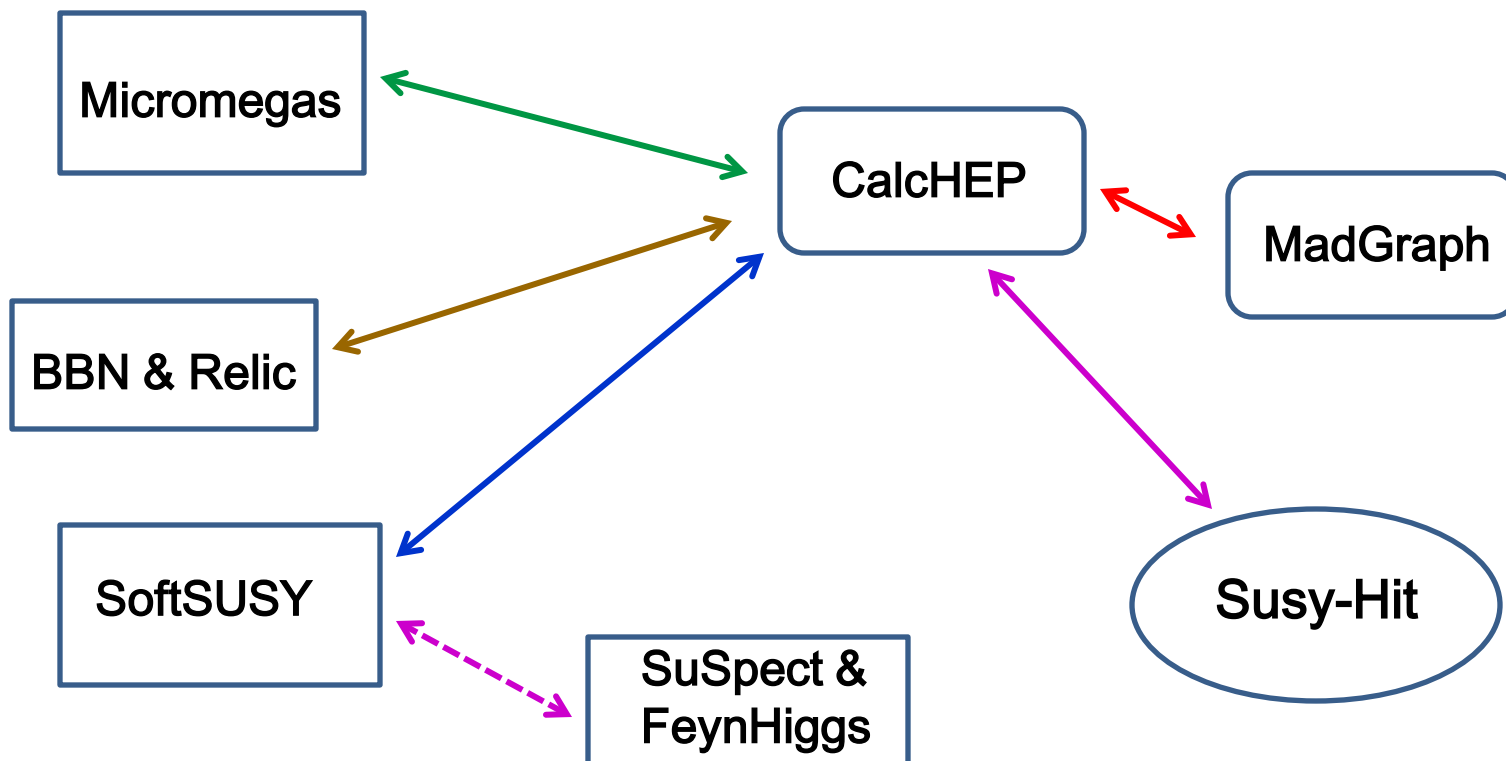
.... for both gravitino and neutralino LSPs

Let's investigate the other side of life: gravitino LSPs

- This is NOT generalized **GMSB**.. We make NO assumptions except that the gravitino is the LSP..it could be light ~ 1 ev or it could be heavy ~ 1 TeV ! Anybody can be the NLSP.
- A big issue is **BBN**... NLSPs in this scenario tend to be long lived & their decays will inject hadronic &/or EM energy into the early universe, possibly **disrupting BBN**
- Lots of **NEW** code needed, e.g., generalize all NLSP/NNLSP decays to the case of **arbitrary gravitino mass** .. **Existing codes inadequate !**
- **Reminder:** No DD or ID constraints for G LSP models

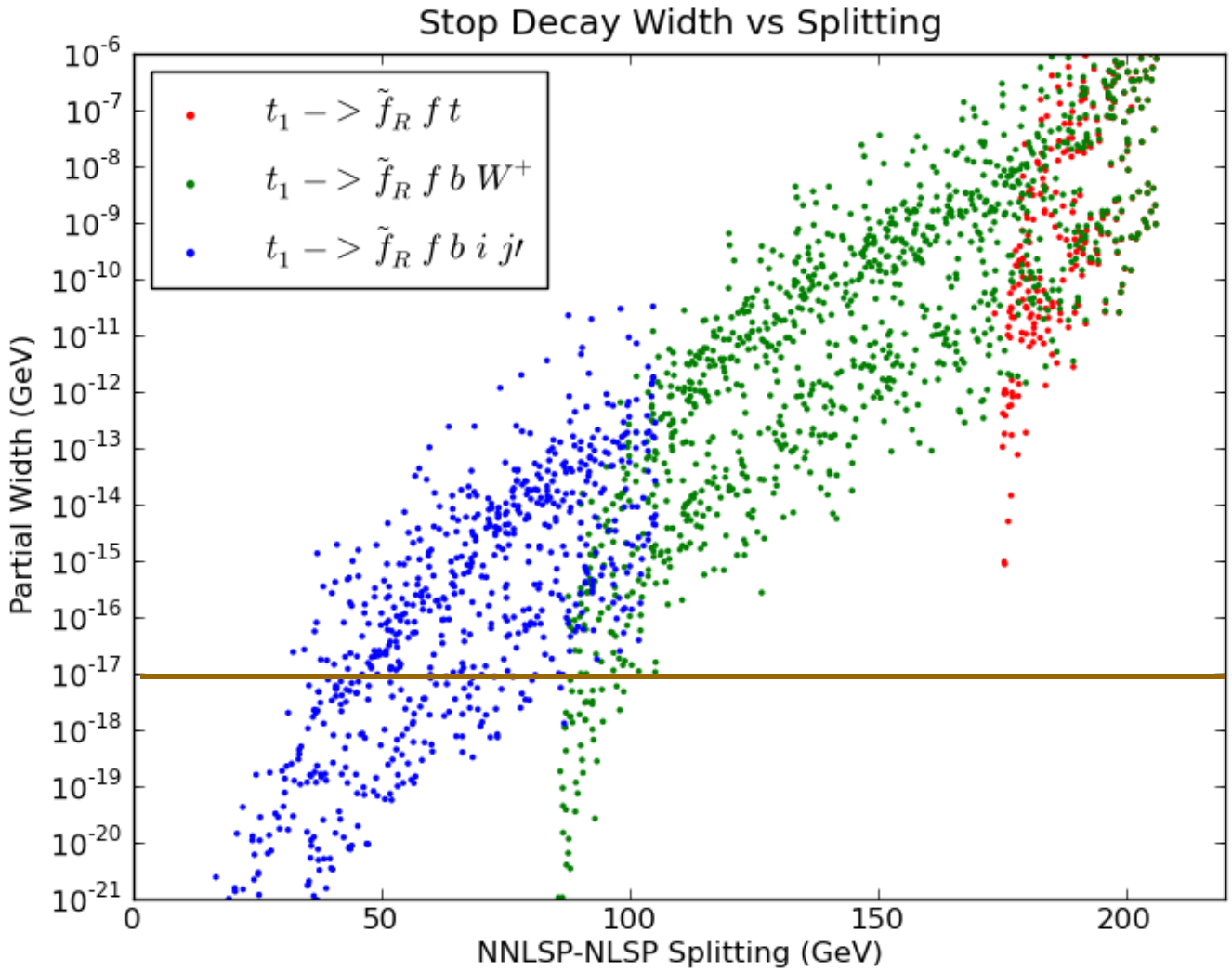
Some of the (MANY) Changes & Additions

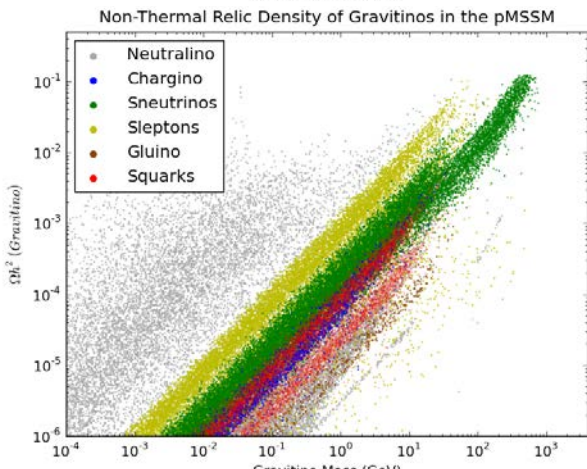
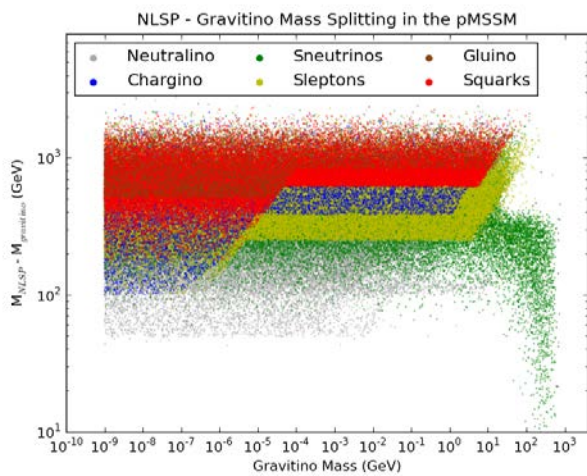
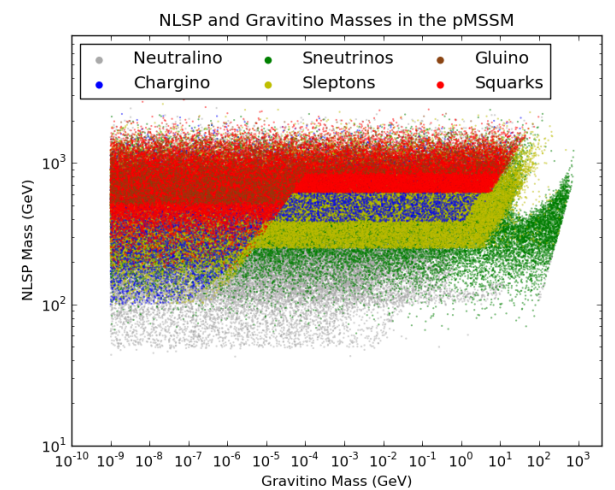
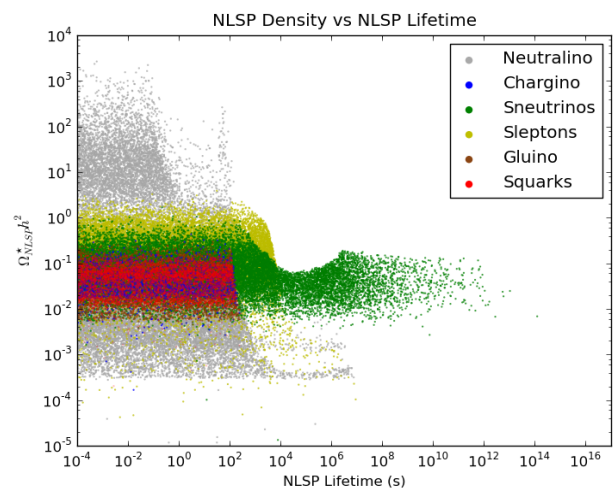
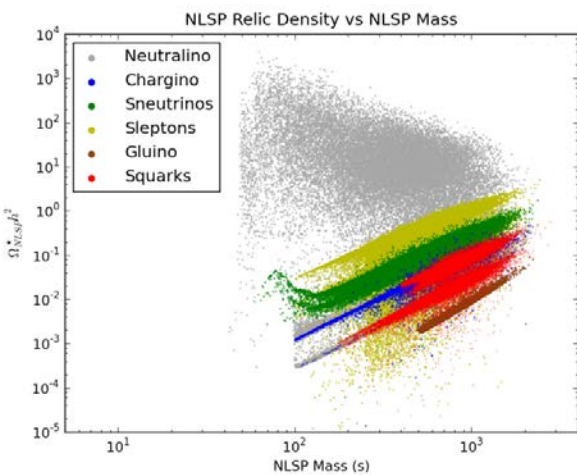
- Decays to G's can no longer be performed in the Goldstino limit so all possible NLSP decay modes involving G's need to be recalculated ...



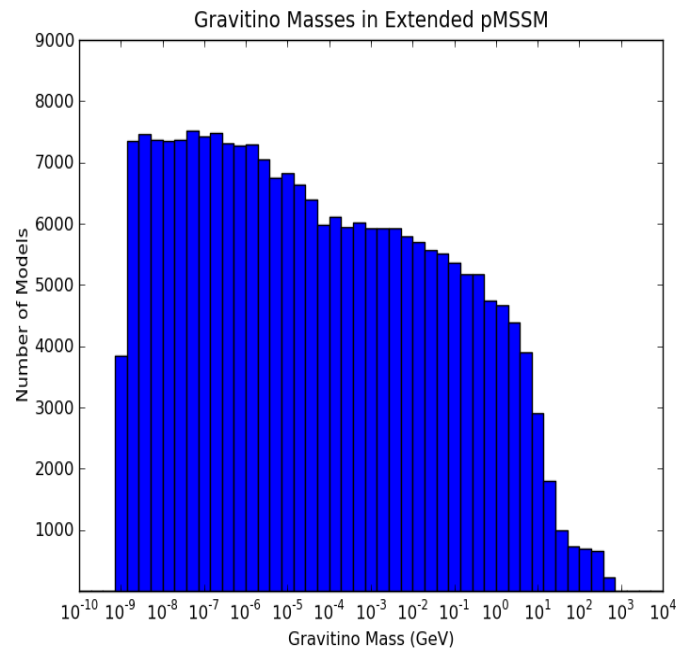
- **For non-G decays** (e.g., for the **NNLSP** \rightarrow **NLSP**) add all **3-body sparticle decays** not in SUSY-Hit via CalcHEP
 - Add relevant **4-body decays** for **gluinos**, t_1 & χ_{1^\pm}
 - Add **5-body decays** of t_1 via RH-sfermions
- \rightarrow All sparticles w/ masses larger than $m_{\text{bottom}} + m_{\text{NLSP}}$ now have complete decay tables for collider & BBN studies . NNLSPs can be detector stable
- **For NLSP decays to G**, add all 3- & 4-body modes w/ BBN relevant lifetimes ($\sim 10^{-4}$ to 10^{14} sec) via MadGraph
 - Calculate NLSP density using Micromegas & rescale to the gravitino mass
 - Use lifetime & BF info for NLSPs from modified SUSY-Hit & check the constraints on EM or hadronic energy deposition during BBN
 - Add constraints from the cosmo relic ν & diffuse photon fluxes

E.g., even if t_1 is the NNLSP it may **STILL** be **detector stable**

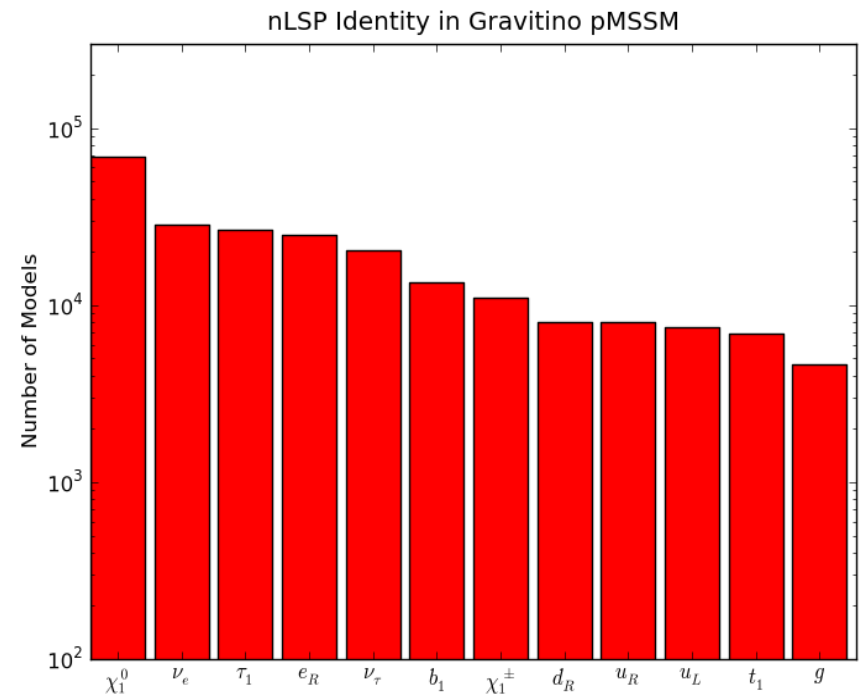
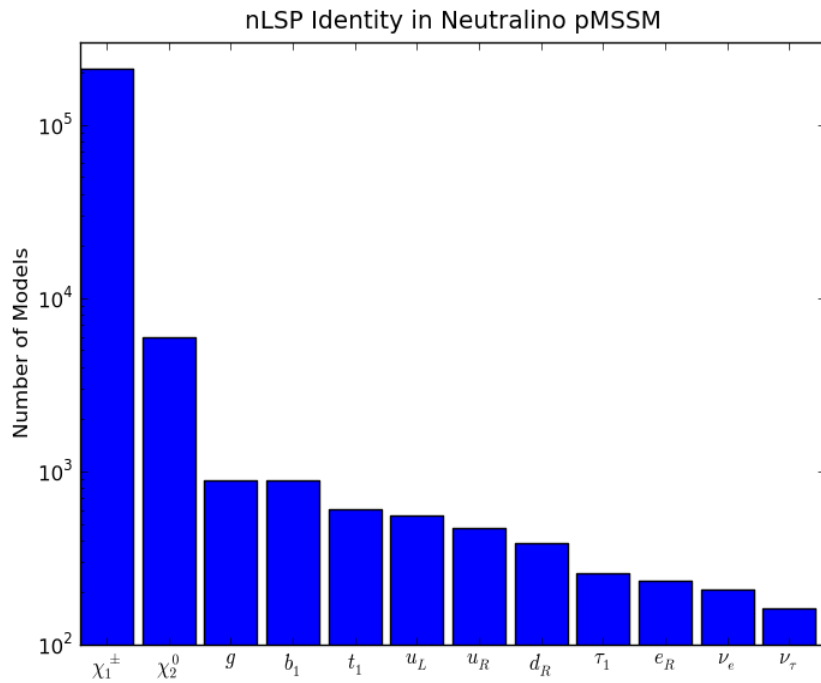




Some properties of the gravitino LSP models

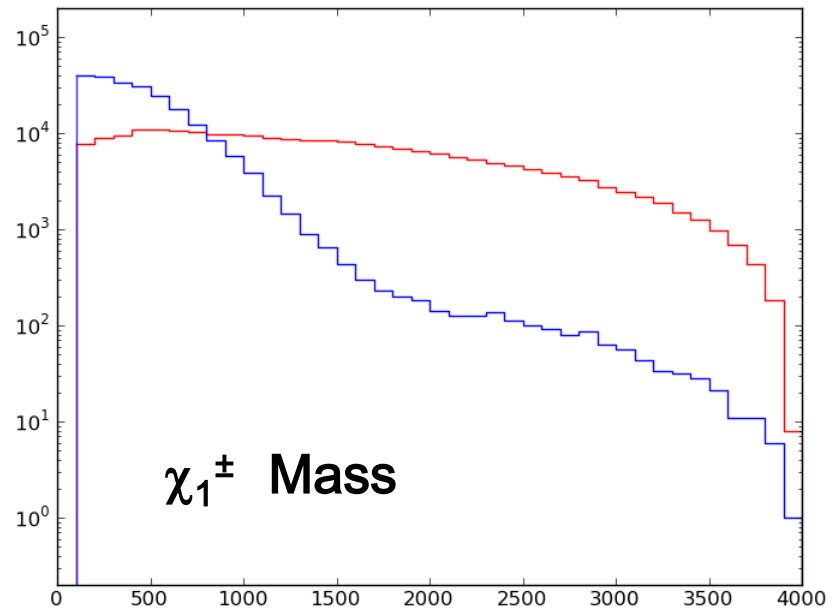
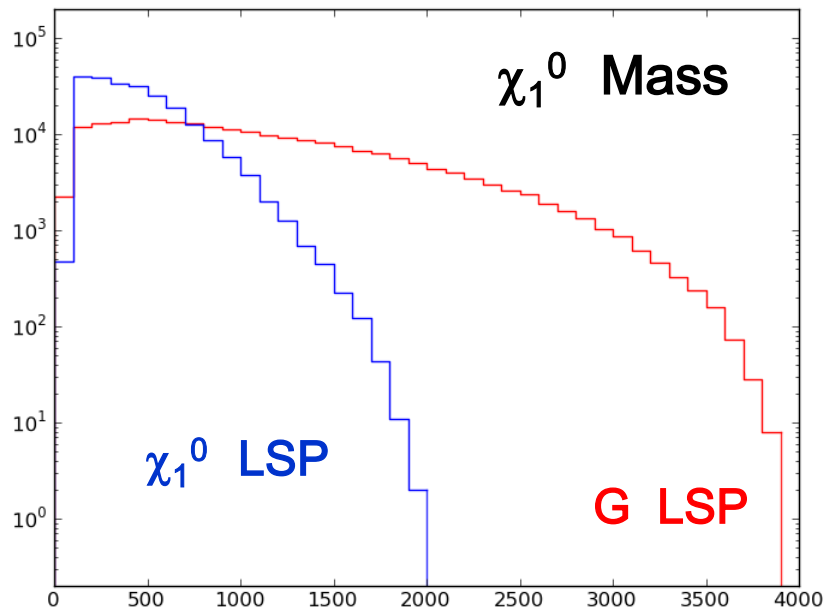


At first glance gravitino LSP models appear to be a bit different that the neutralino LSP ones... A comparison is quite interesting.



- The likelihood of various NLSP identities is **very strongly dependent** on the **LSP choice**
- This can have a **potentially large influence** on LHC SUSY searches (apart from, e.g., additional cascades)

- The mass spectra of the MSSM fields are (indirectly) influenced by the nature of the LSP, i.e., the fact that **G** can be **VERY light** whereas χ_1^0 must be $> \sim 10$'s of GeV in the scan..
- E.g., since the lightest neutralino is **at best** the NLSP in the **G** scan, its mass distribution must now **extend to larger values**
- Other sparticle masses are **less influenced** due to scan ranges



ATLAS MET Analyses @ 7 & 8 TeV

- The first step in exploring the parameter space of either pMSSM model set is to apply the SUSY MET searches
- As is our tradition, we follow the ATLAS analyses as closely as possible & we began w/ the χ model set
- At $\sim 1 \text{ fb}^{-1}$ this is 'relatively straightforward' as all the data & numerous benchmark model results exist that we can test/validate against. Only partial $\sim 5 \text{ fb}^{-1}$ results available.
- We combine the various analyses signal regions (as ATLAS does) into : nj0l, multi-j, nj1l, nj2l (+ multi-l & HF) and we quote the coverage for each as well as the combined result.. approach is CPU intensive

% models
excluded

7 TeV $\sim 1 \text{ fb}^{-1}$

7 TeV $\sim 5 \text{ fb}^{-1}$

nj0l [5/11]	6.68%	23.23%
multi-j [5/6]	0.36%	1.61%
nj1l [8/3]	0.79%	2.64%
nj2l [5]	0.15%	0.20% ^{***}
flavor	(in progress)	(ditto)
(sub)total	6.73%	23.27%

→ nj0l is by far dominant in these searches

*** In this case, we extrapolate to $\sim 5 \text{ fb}^{-1}$, since results have not yet been released. We assumed that the number of events observed equals the expected backgrounds & that the analysis cuts are exactly the same as at $\sim 1 \text{ fb}^{-1}$

- Our analyses can be updated when more data is available³

(Preliminary) Extrapolation to $\sqrt{s} = 8$ TeV

- The extrapolation here is greater than for $\sim 1 \rightarrow \sim 5 \text{ fb}^{-1}$ @ 7 TeV
- First pass: assume the cuts & analyses are as for 7 TeV & the number of observed events equals the expected backgrounds in each SR.
- However, we need to know the backgrounds for 8 TeV !
- Rescale ATLAS 7 TeV backgrounds? How? Use MC to determine the **RATIO** of the expected backgrounds in each signal region at 7 & 8 TeV as transfer factors
- When low statistics becomes an issue we closely follow ATLAS' approach using the sideband 'ABCD' method & then rescale the control regions
- Of course we still need to generate the relevant SM MC backgrounds

SM Background Generation @ $\sqrt{s}=7$ & 8 TeV

- $Z/W^\pm + (0-4)j$
- $WW/ZZ + (0-2)j$
- $t\bar{t} + (0-2)j$
- $\text{single } t + (0-2)j$
- QCD up to 6 jets

\leftrightarrow ME + PS, weighted evts

~ 1 TB

w/ Sherpa

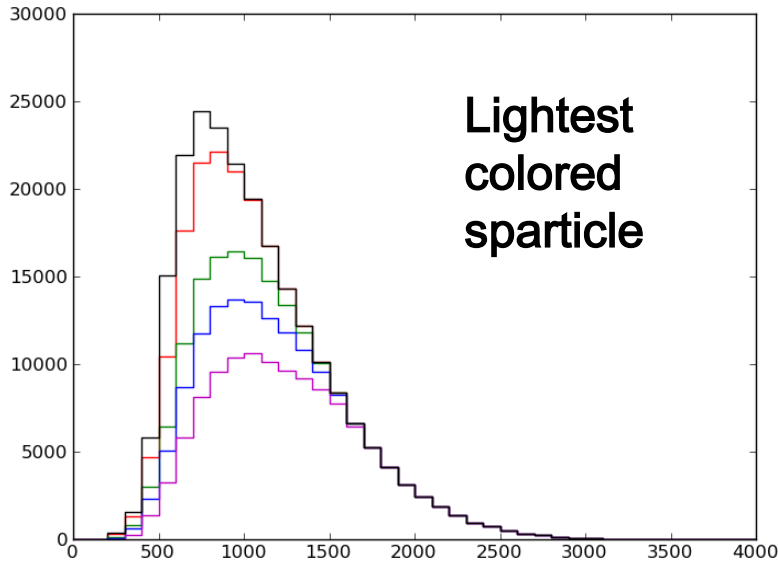
- **Not too surprisingly**, the gain in pMSSM coverage going to 8 TeV is **substantial due to the increases in σ 's**. **nj0l** continues to dominate :

	<u>8 TeV 5 fb⁻¹</u>	<u>8 TeV 20 fb⁻¹</u>
nj0l**	32.70%	45.11%
multi-j**	6.26%	7.35%
nj1l**	1.41%	1.53%
nj2l**	0.34%	0.37%
flavor	(in progress)	(ditto)
(sub)total	32.75%	45.13%

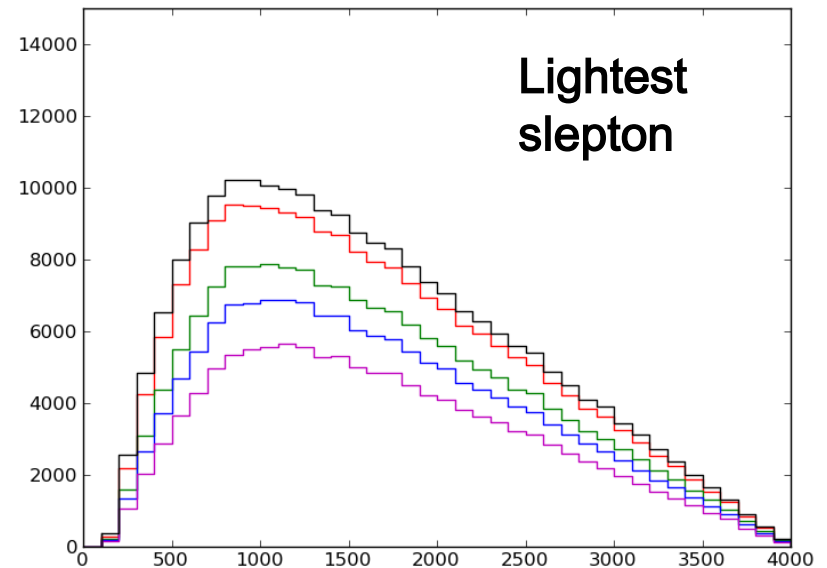
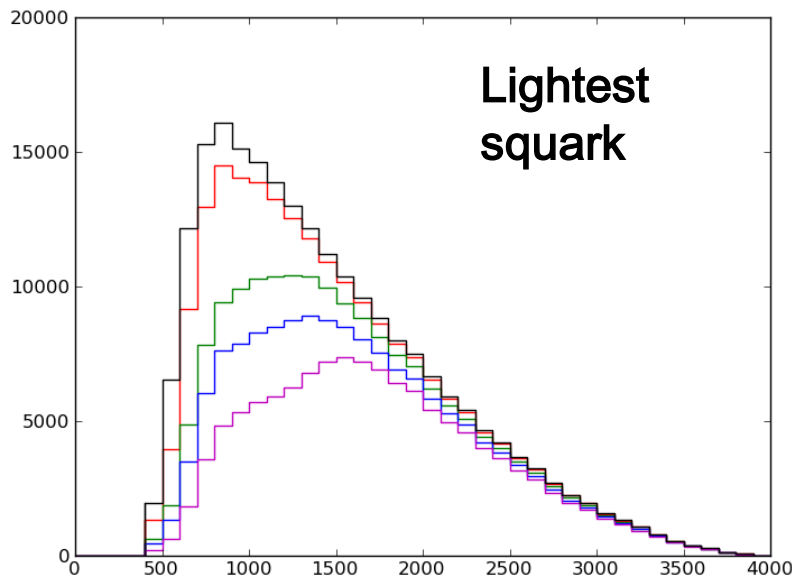
** extrapolated from $\sim 5 \text{ fb}^{-1}$ analysis ++ extrapolated from $\sim 1 \text{ fb}^{-1}$ analysis

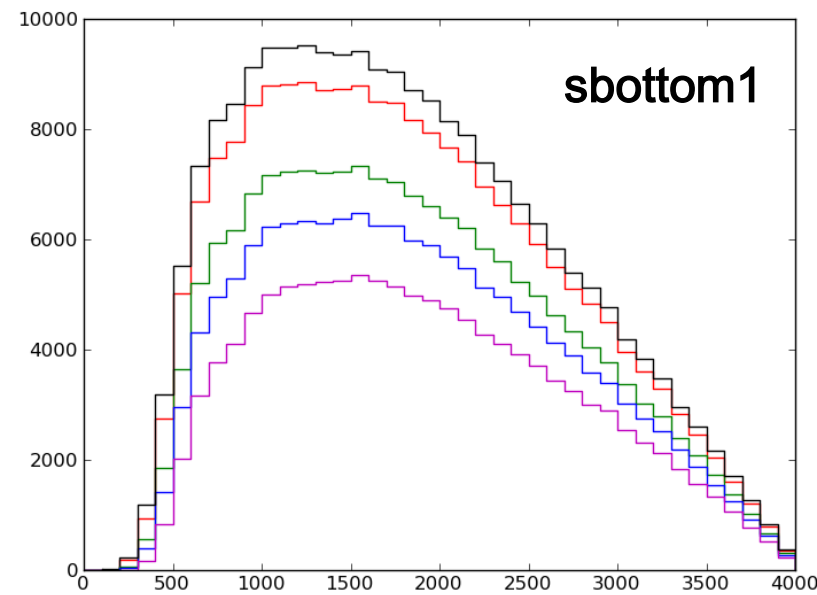
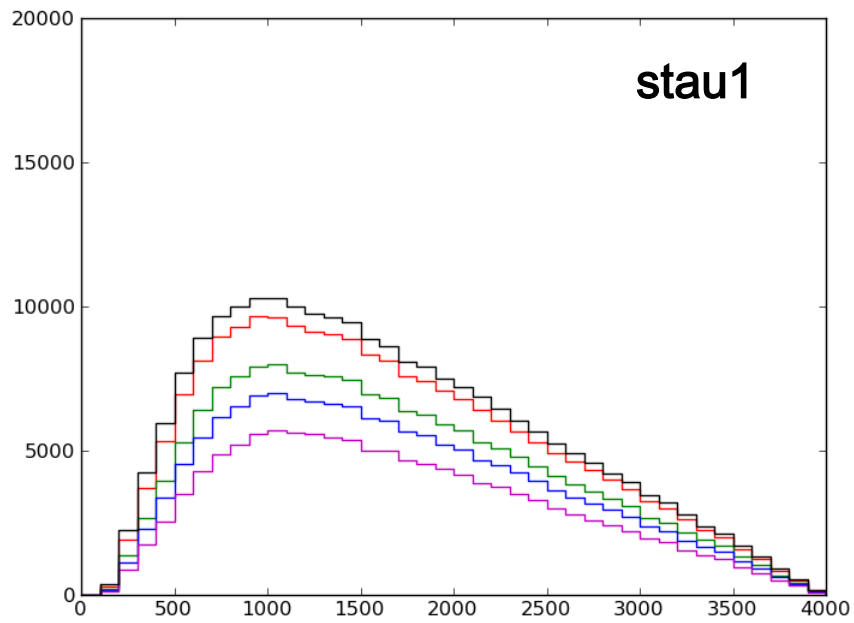
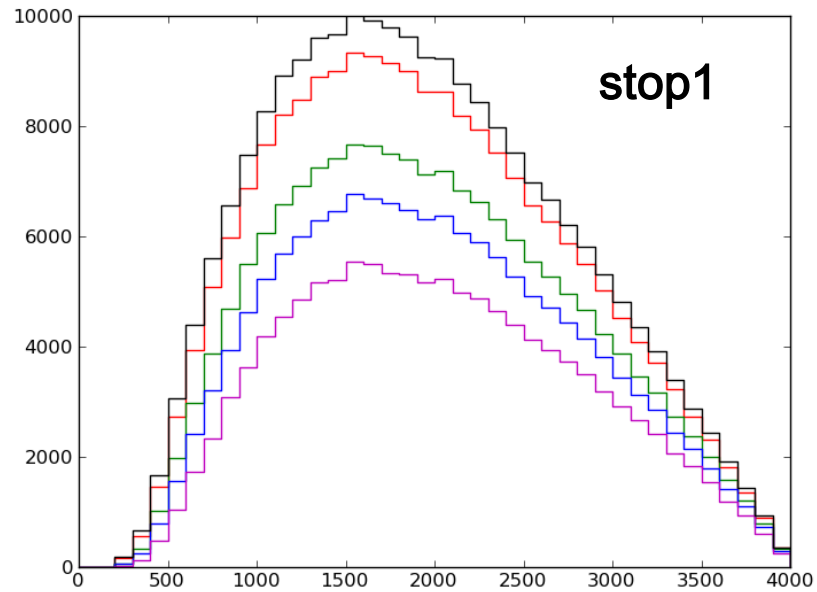
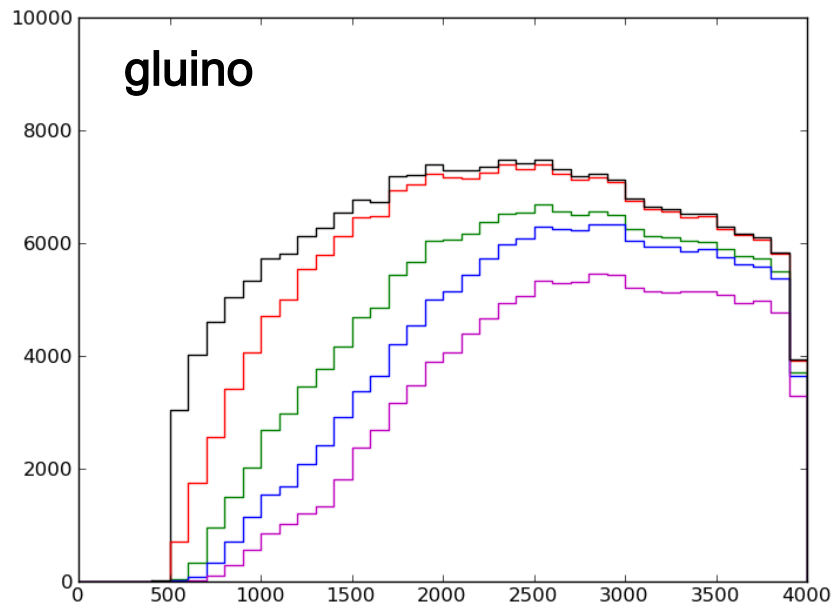
- **$\sqrt{s}=13-14\text{TeV}$ is needed for more complete coverage**

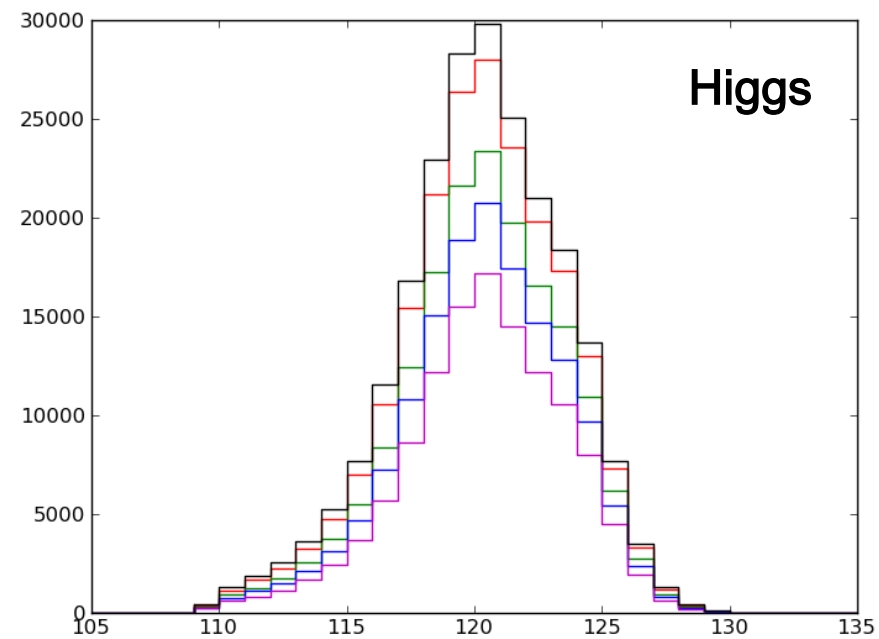
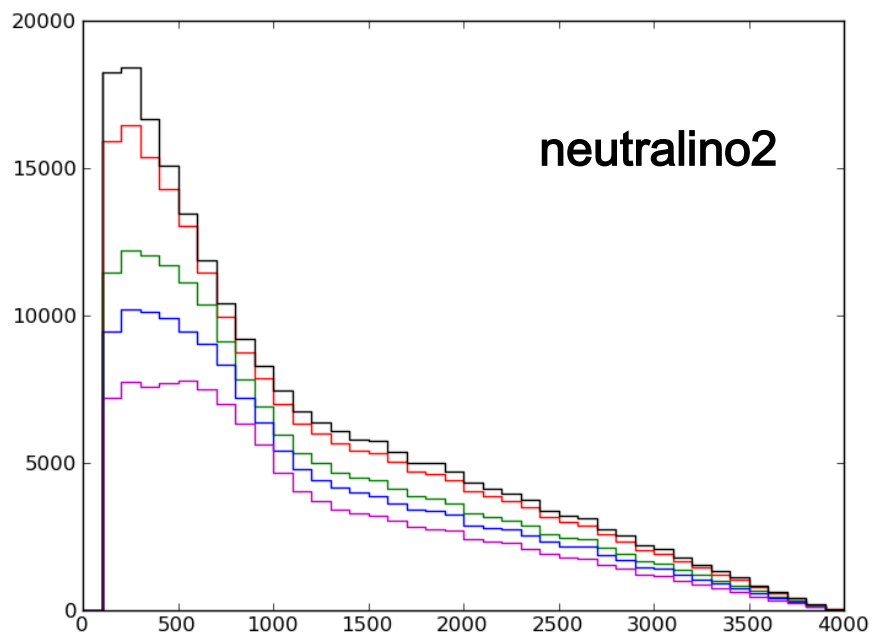
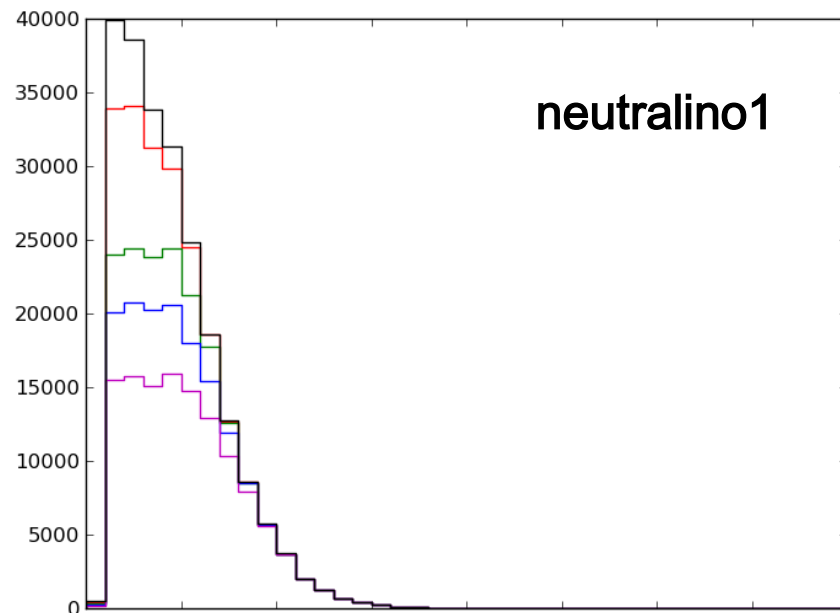
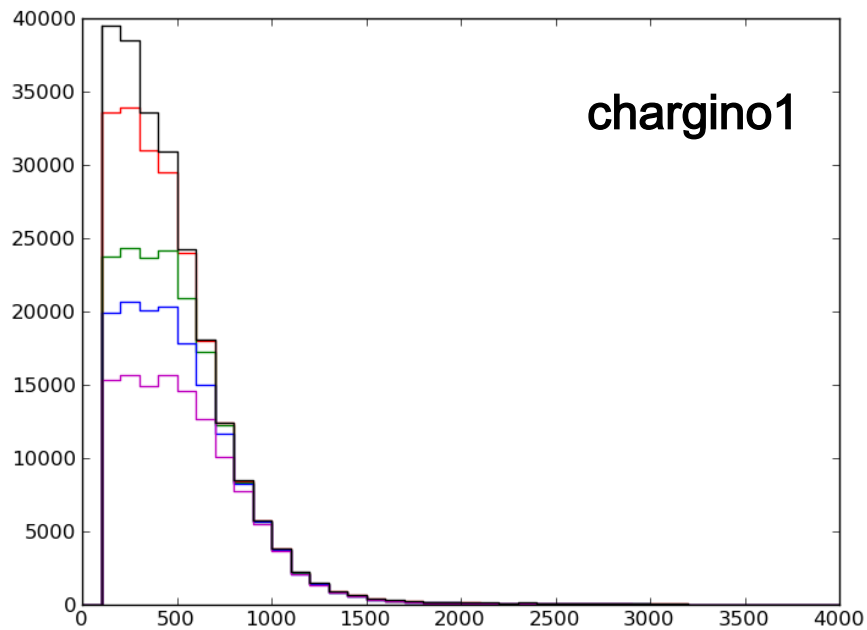
How does the pMSSM respond to negative searches ?

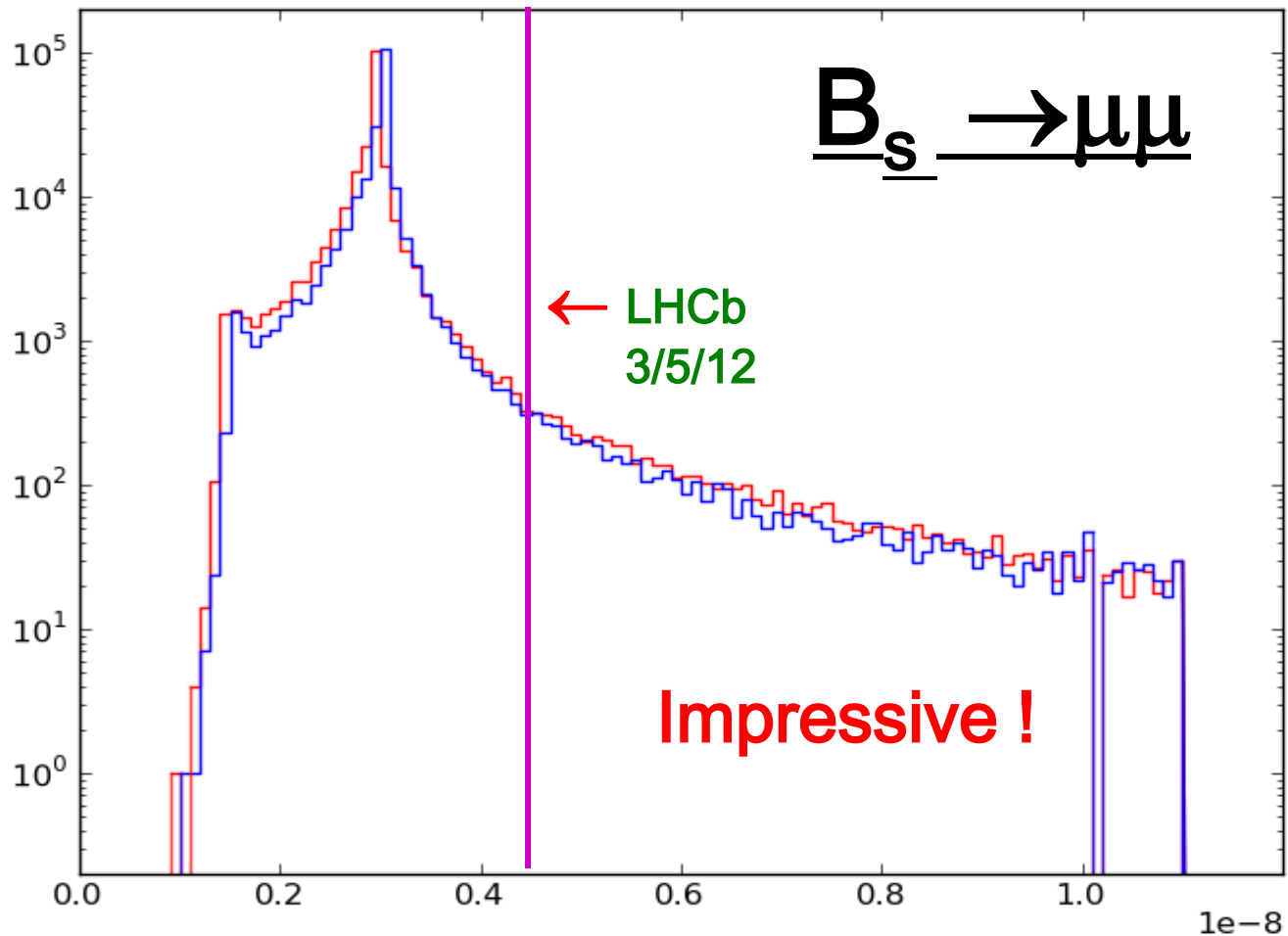


Note that **colored** sparticles get **heavier**, i.e., the distributions peak at **higher masses** as the searches progress but color singlets distributions are just seen to be **rescaled** downward



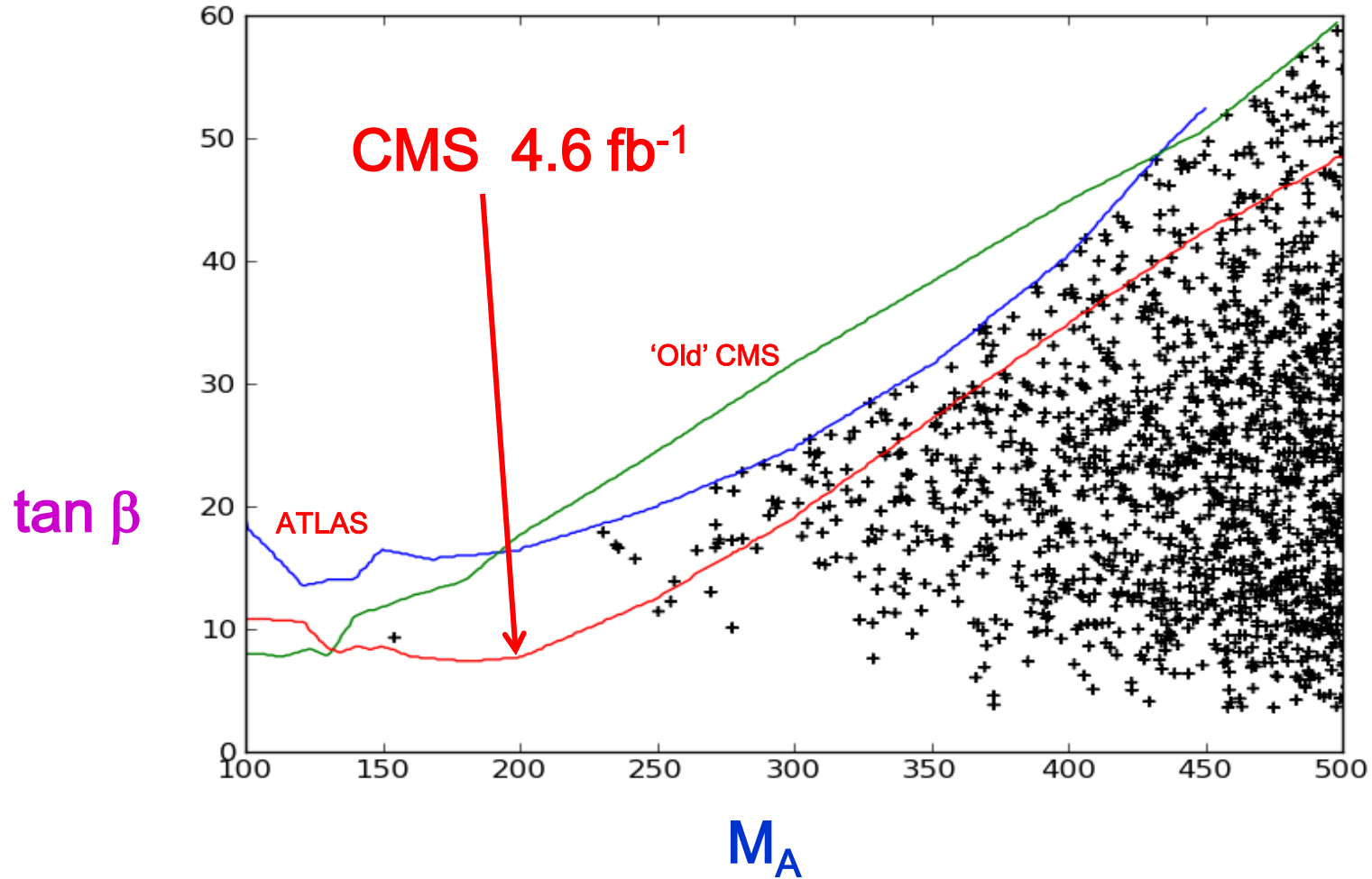






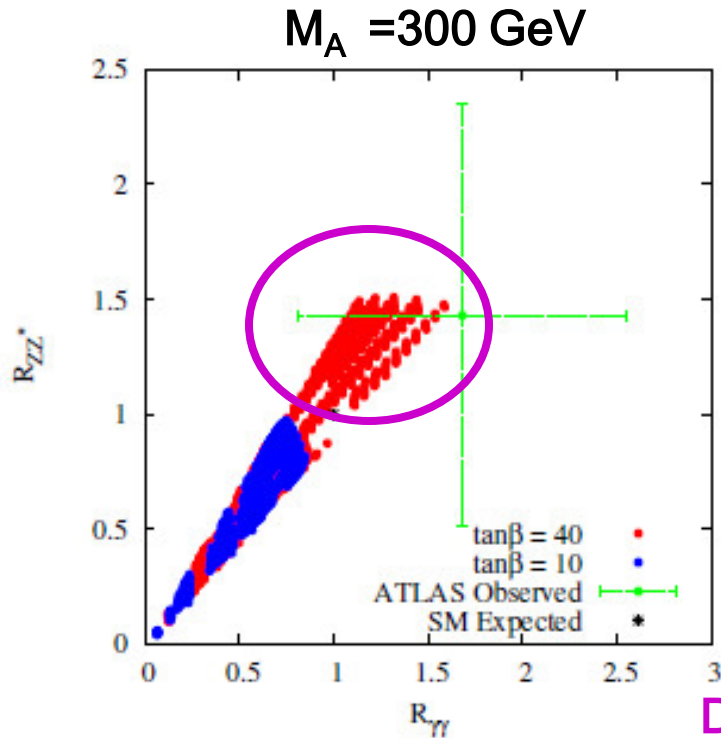
- **LHCb** result removes a total of **4899** (~~5884~~) models in the **neutralino (G)** LSP model set ...
- **non-MET** searches REALLY ARE important !

Impact of $A, H \rightarrow \tau\tau$ Searches

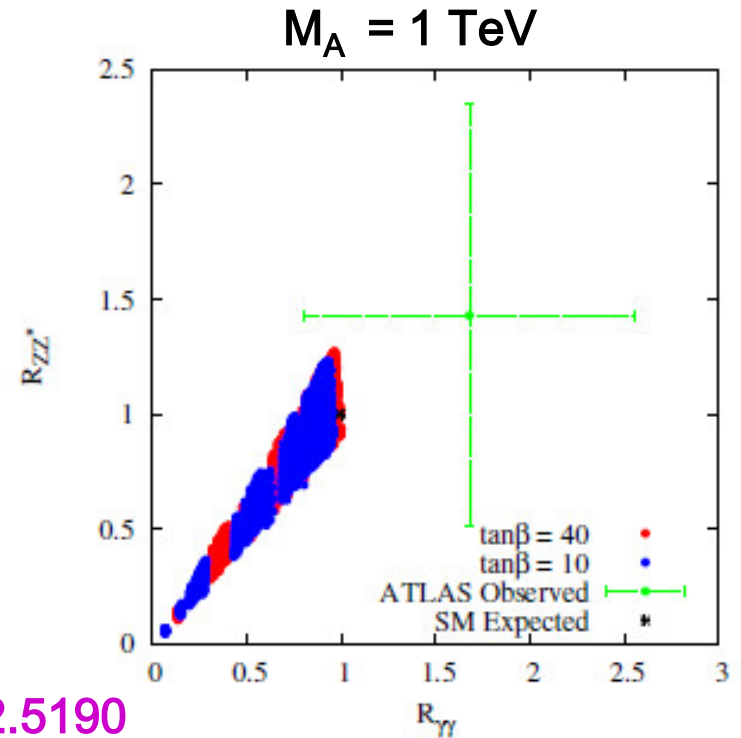


As in the case of $B_s \rightarrow \mu\mu$, improvement in non-MET searches impact the pMSSM analyses... **160 models** removed from the **neutralino LSP** set...²¹

...for example Higgs properties in the MSSM...

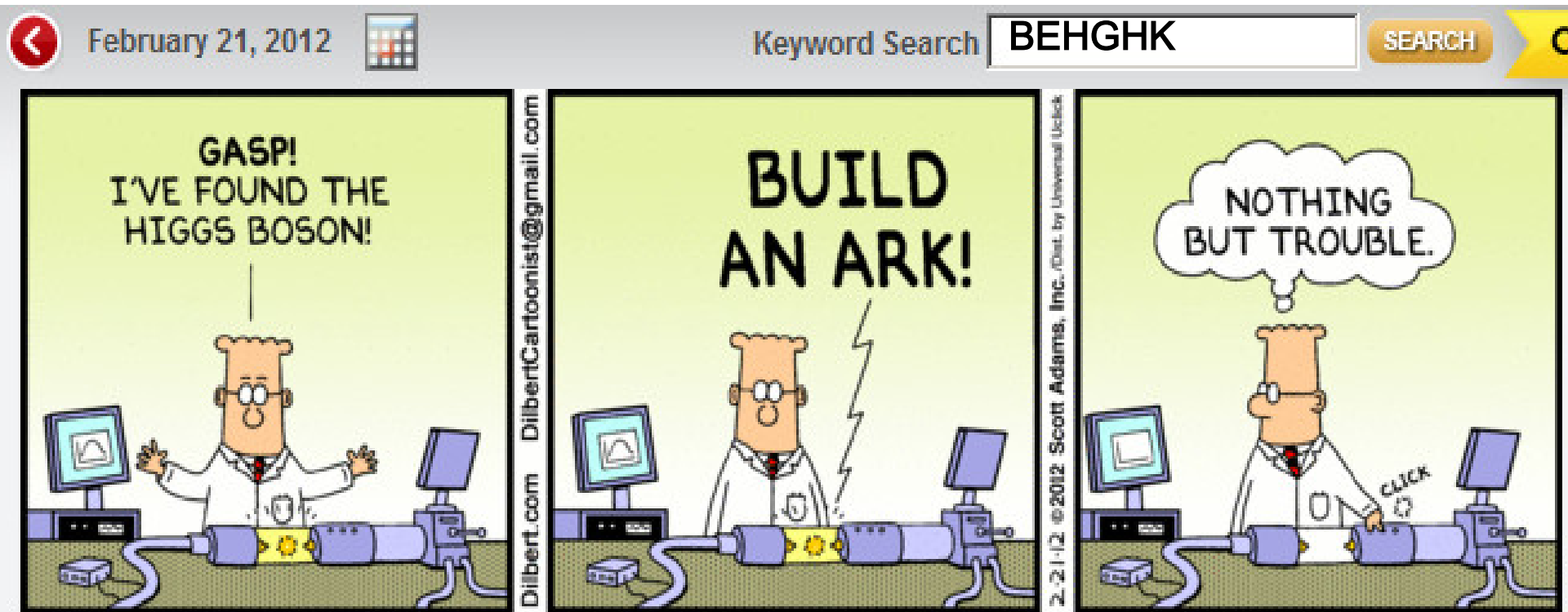


Desai et al. 1202.5190

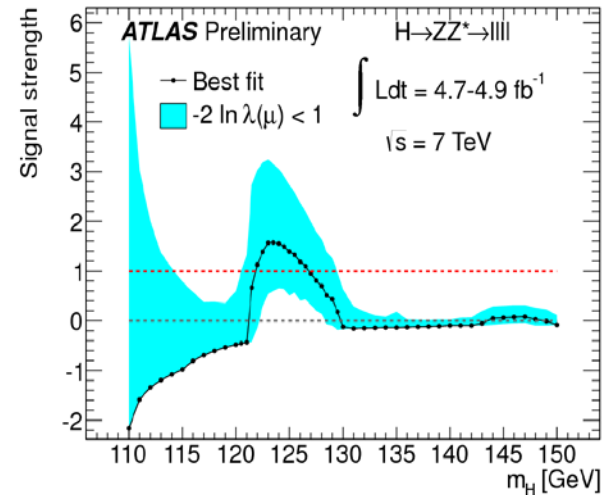
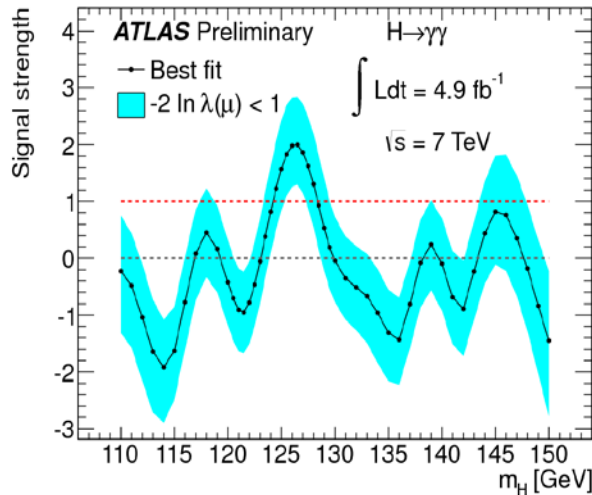
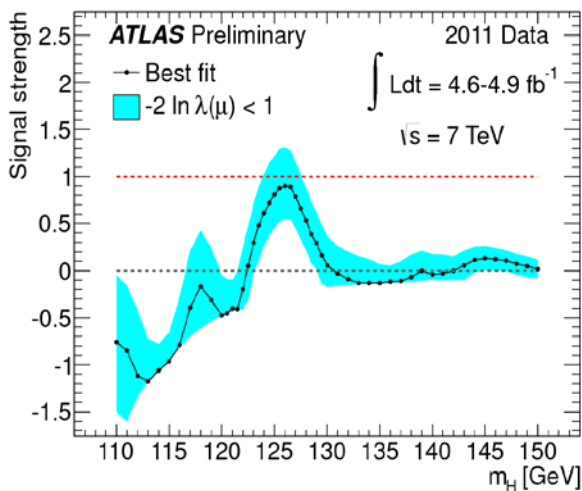


Low M_A & large $\tan\beta$ can enhance h signal rates...however this parameter range is excluded by b-physics & $H, A \rightarrow \tau\tau$ searches !

Impact of LHC SM Higgs Searches



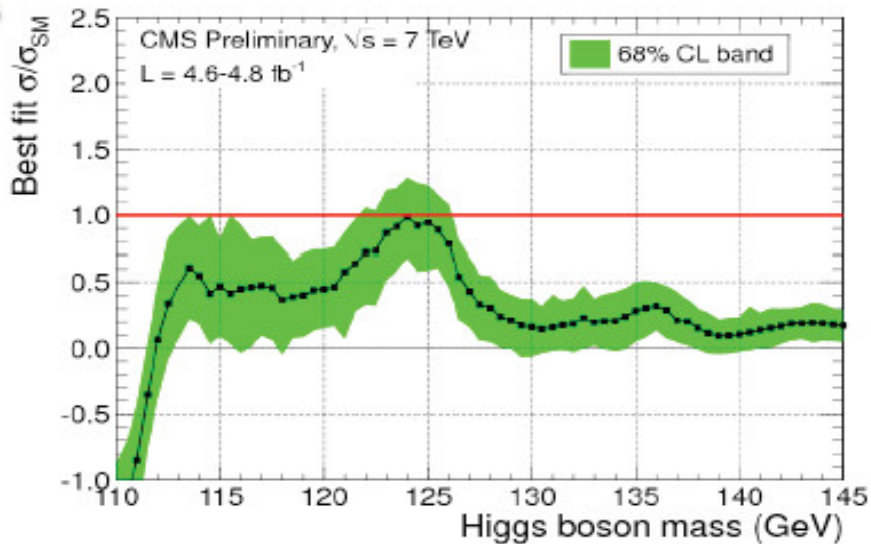
..or what will a Higgs at ~ 125 GeV tell us



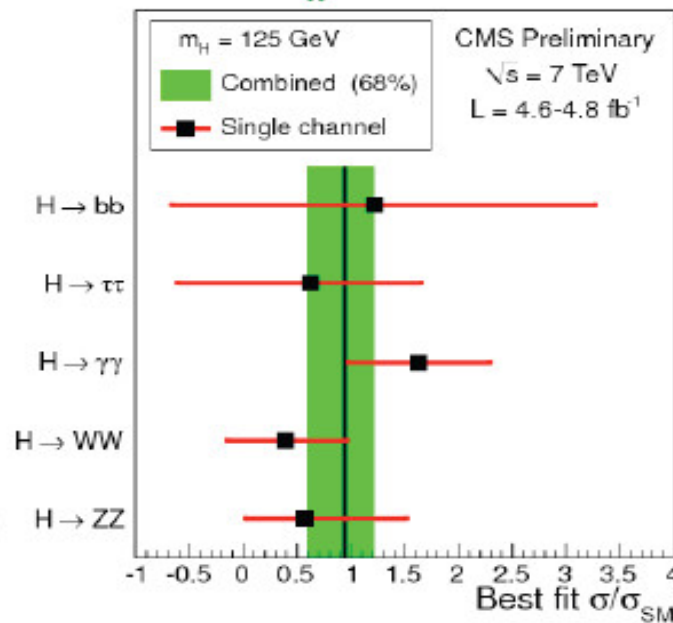
“Generally” living up to ~SM expectations... **



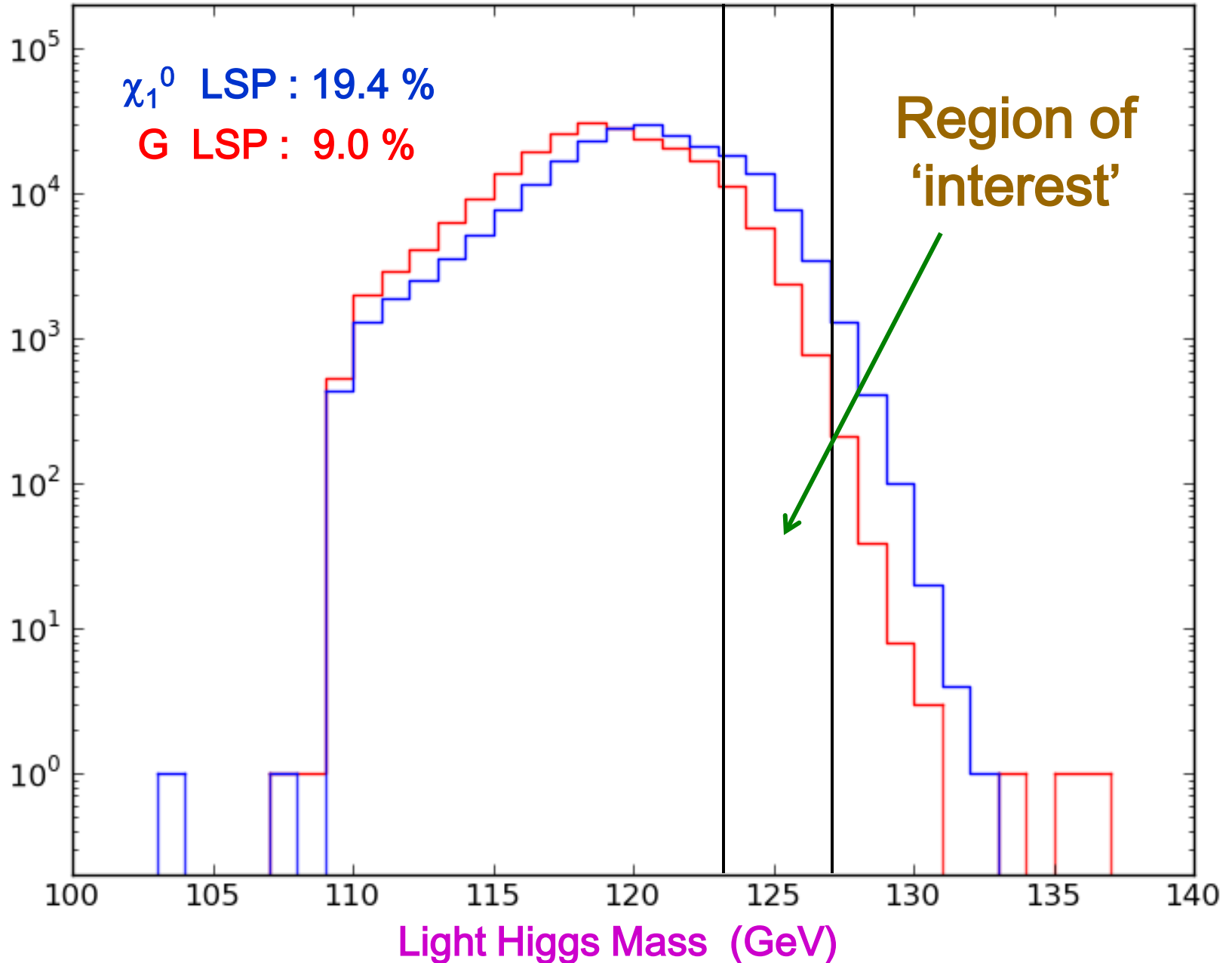
** see however, e.g., 1203.4254



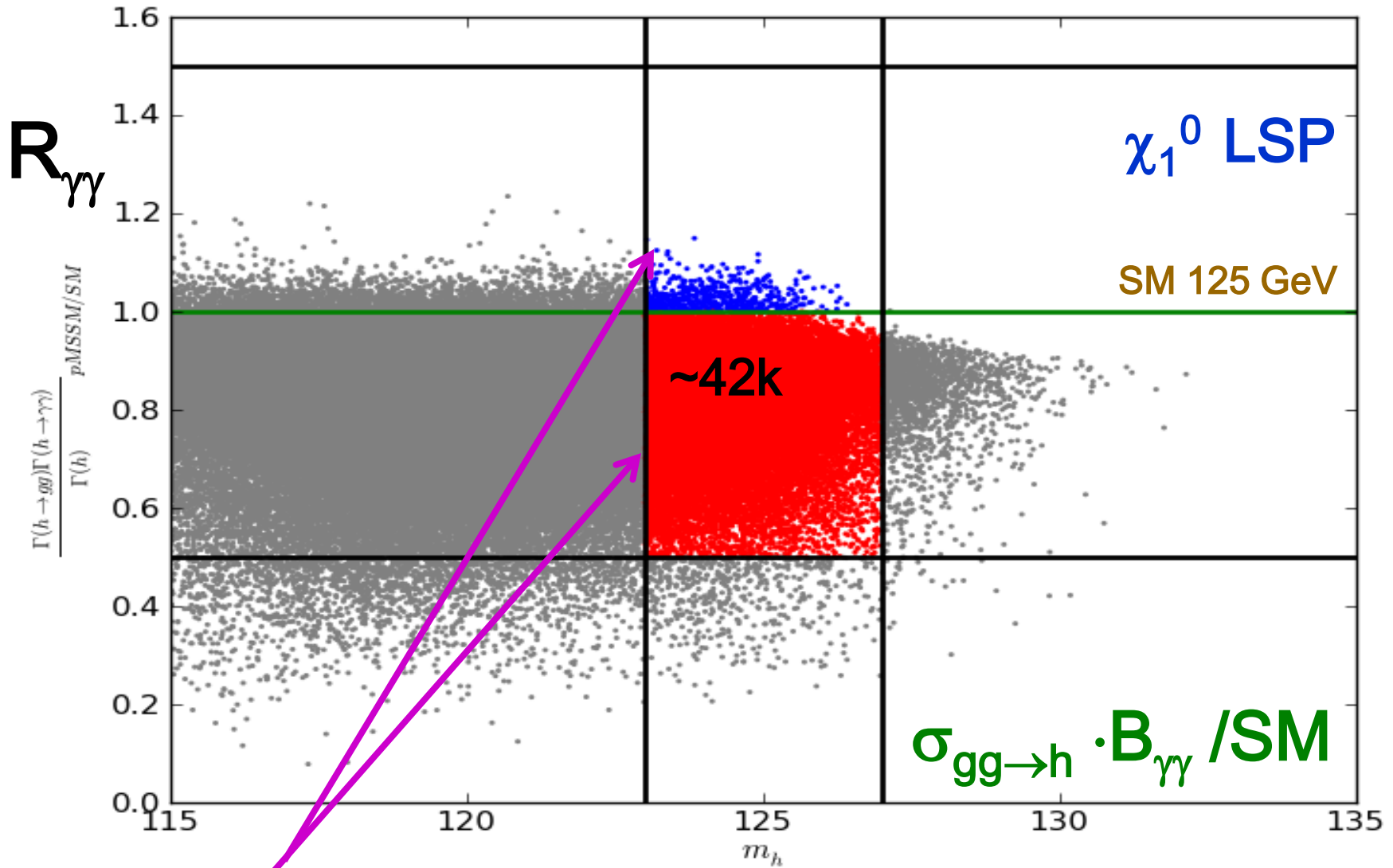
Comparison of channels for $M_H=125$ GeV



Distribution of Predicted Higgs Masses

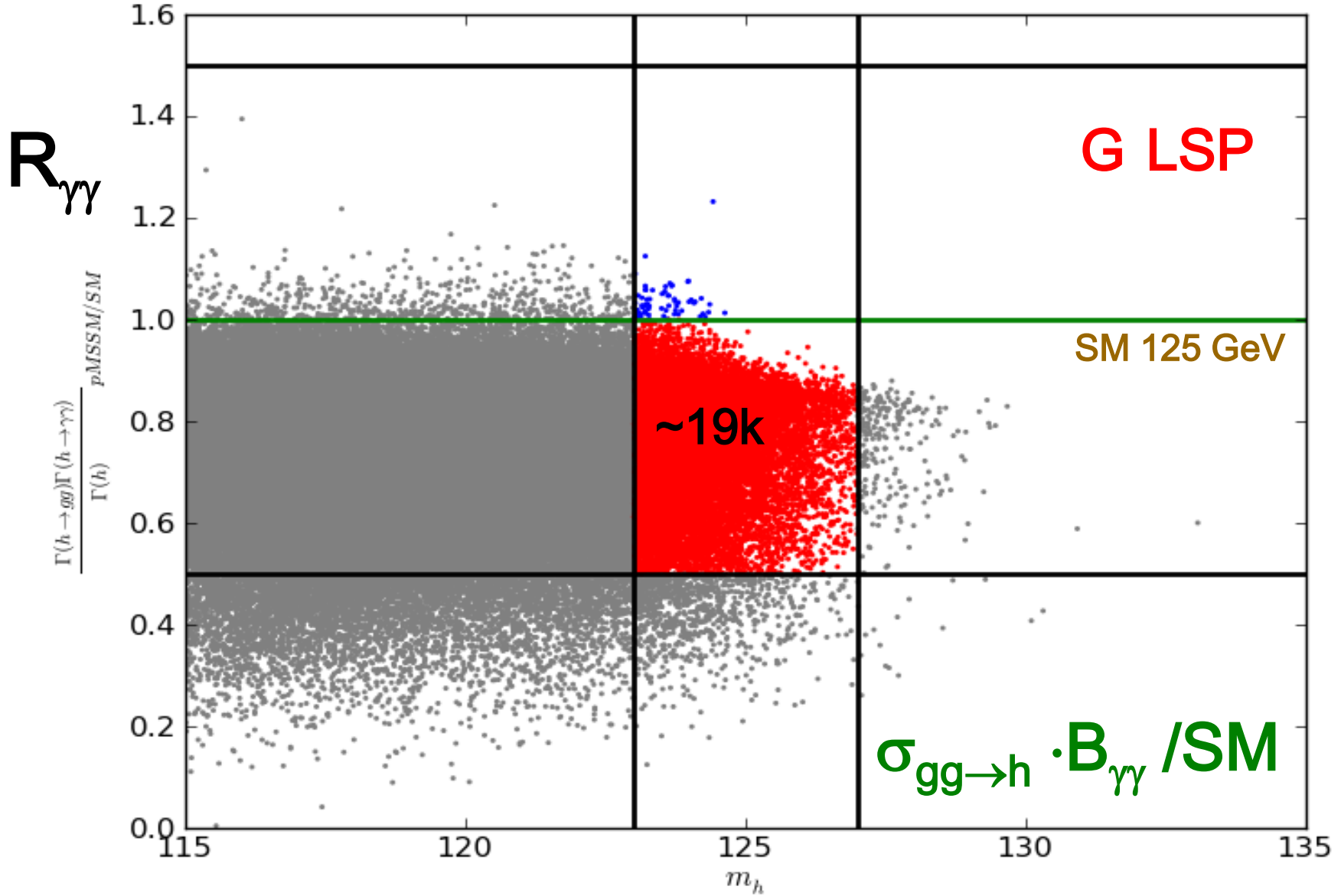


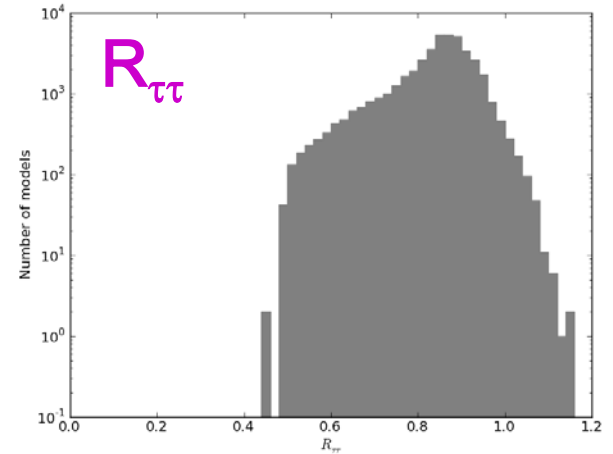
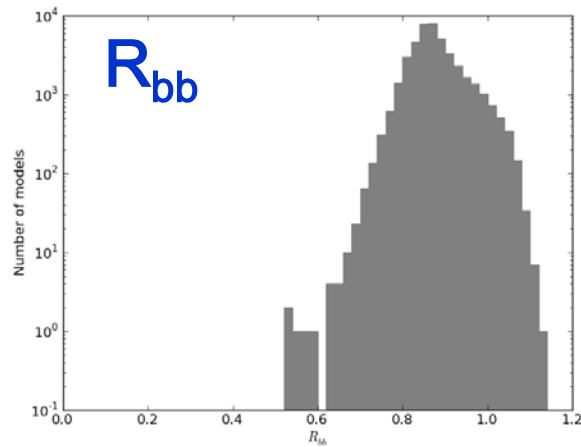
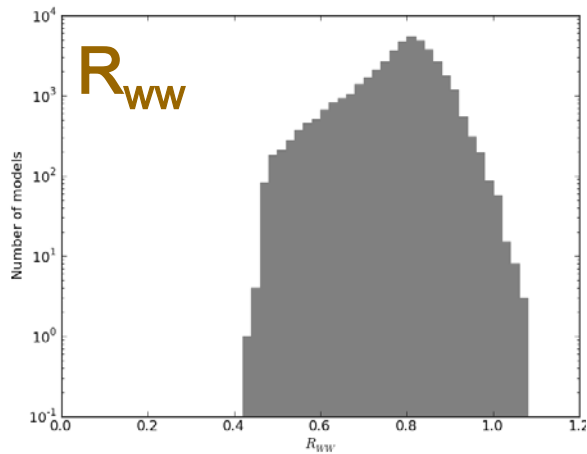
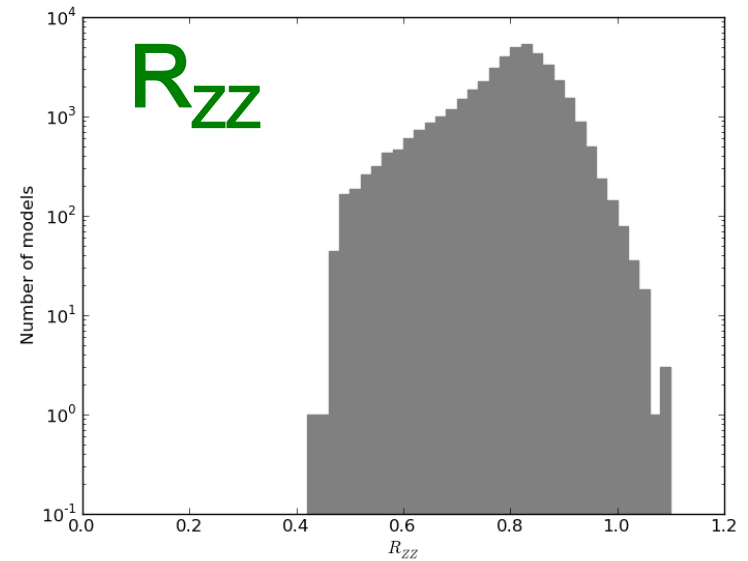
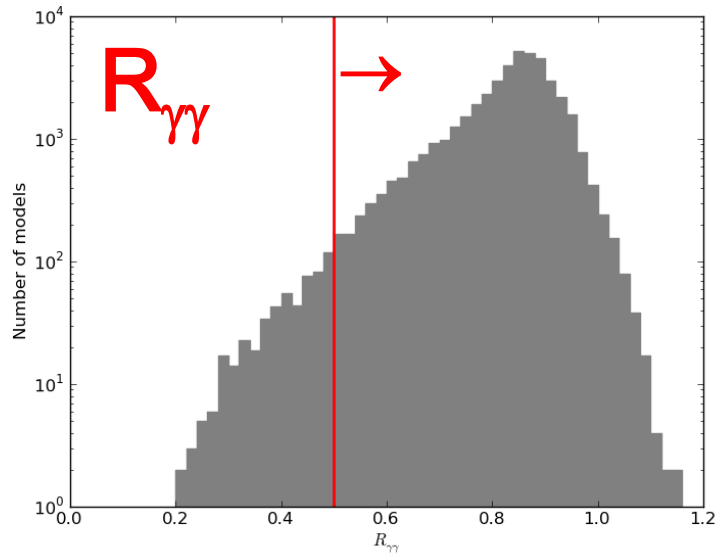
Impact of LHC SM Higgs Searches



Regions of 'Interest' → Few models larger than the SM

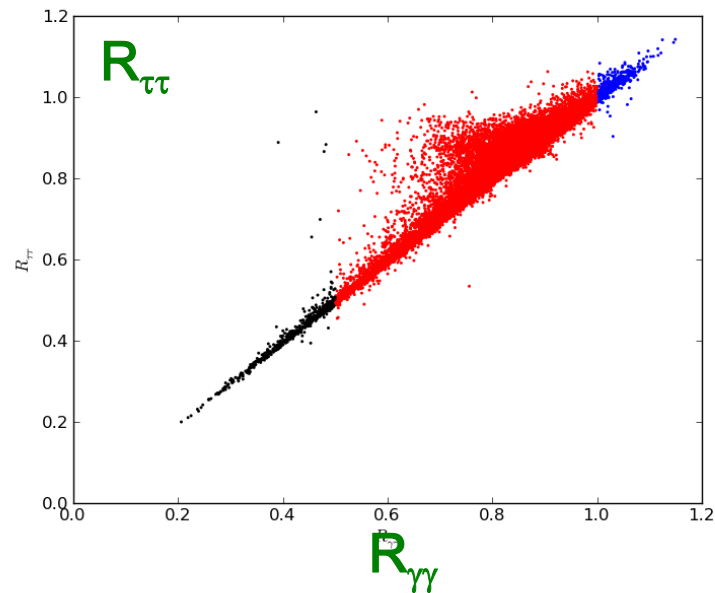
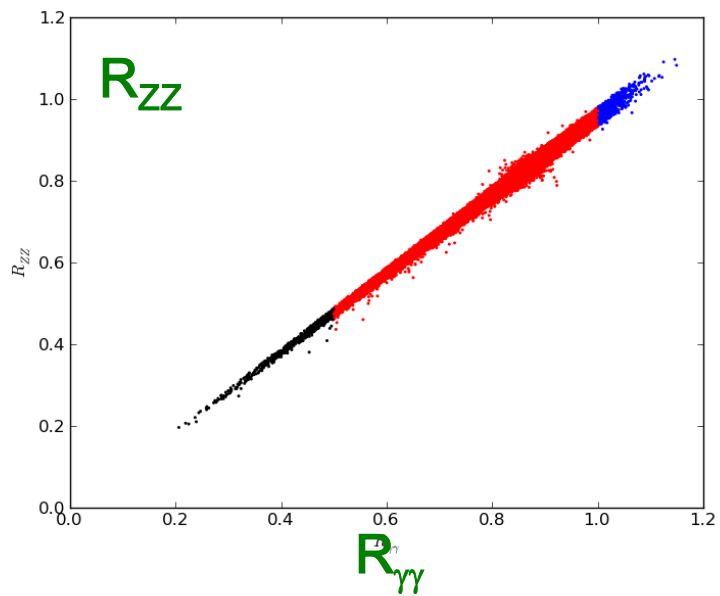
Impact of LHC SM Higgs Searches



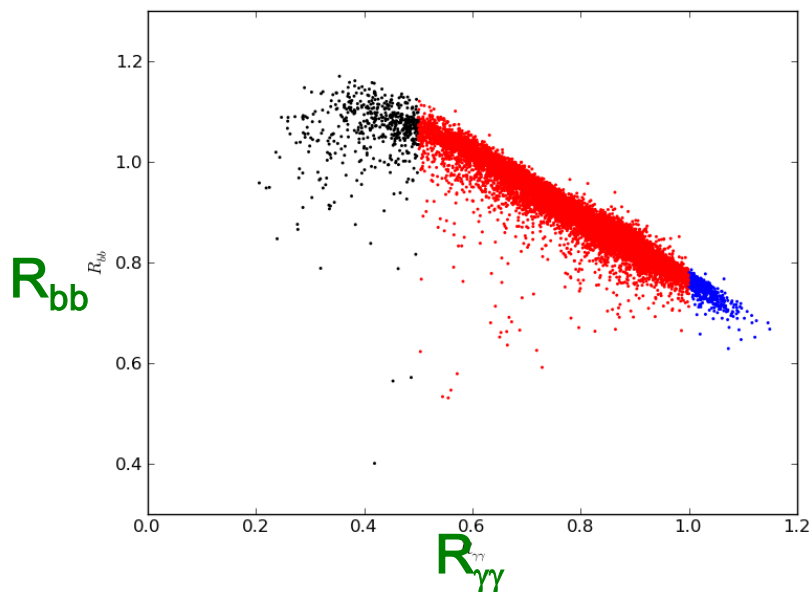
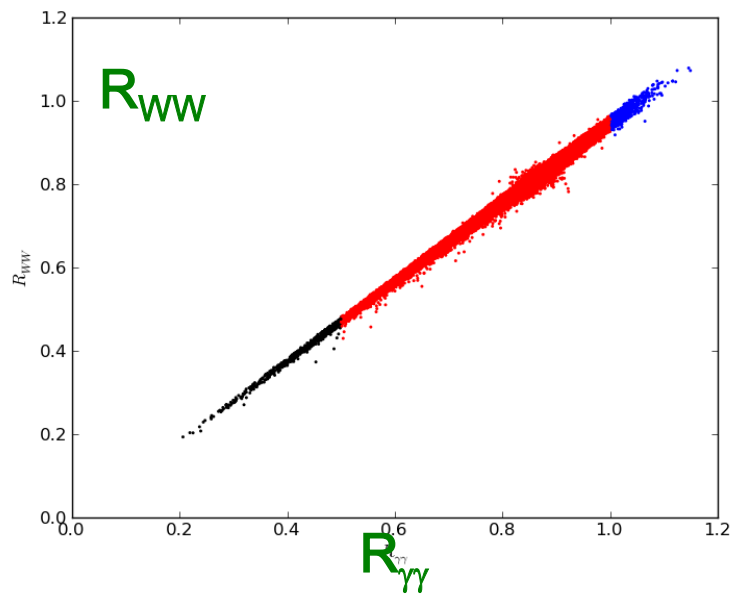


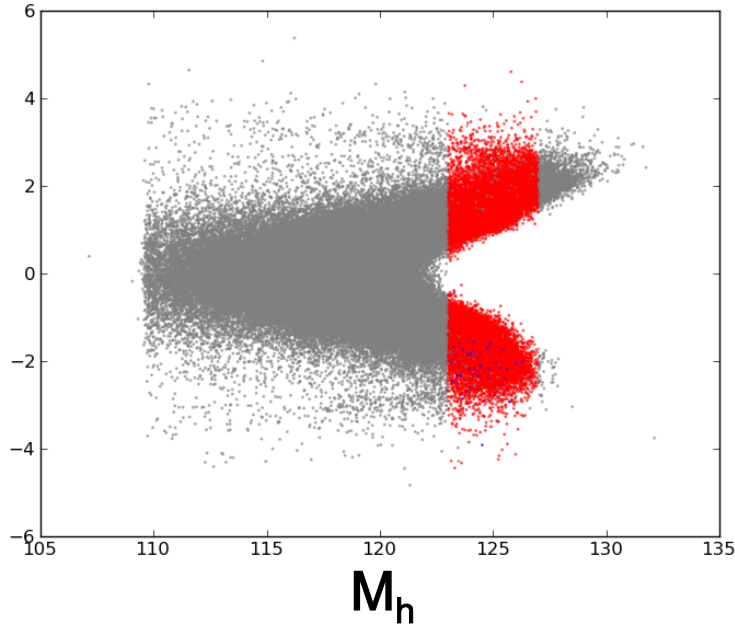
R_{XX} distributions for $m_h = 125 \pm 2$ GeV

χ_1^0 LSP



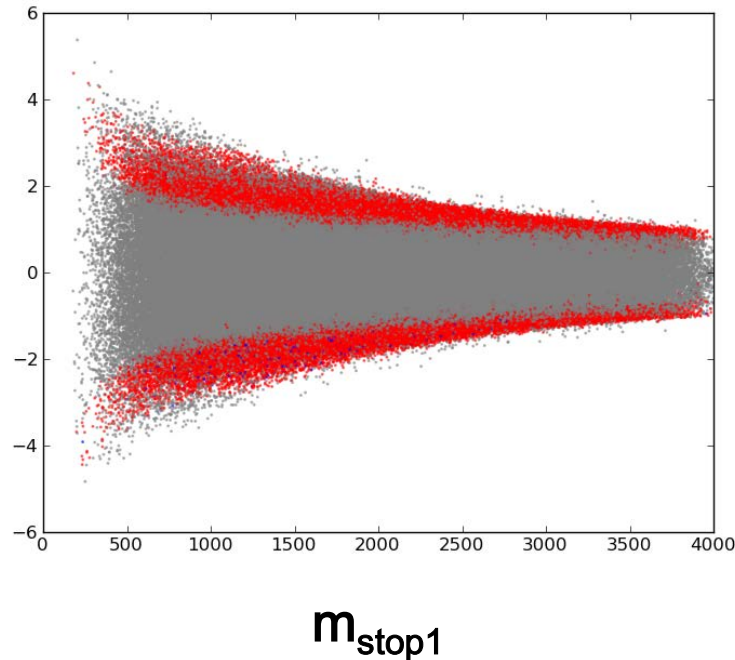
$R_{\gamma\gamma}$ vs R_{XX} for $m_h = 125 \pm 2$ GeV χ_1^0 LSP



X_t / M_S 

$$X_t = A_t - \mu/t_\beta$$

$$M_S^2 = m_{\text{stop1}} \cdot m_{\text{stop2}}$$

 X_t / M_S 

The blue points seem to prefer negative values of X_t . Stop_1 masses down to ~ 300 GeV or even less are still found for large X_t / M_S

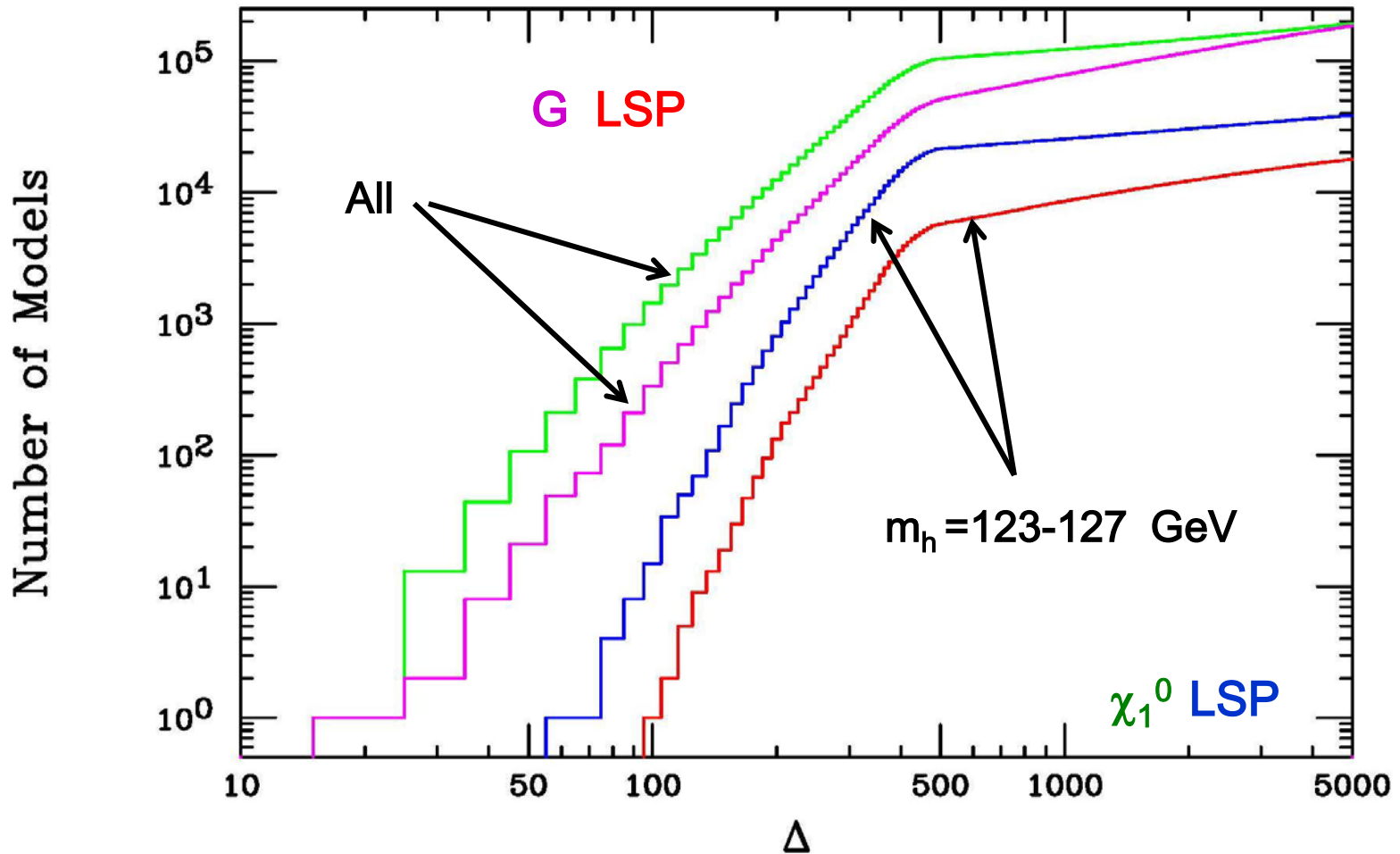
Fine-tuning in the pMSSM

- $m_h = 123-127$ GeV in the MSSM requires large stop masses and/or mixings which then \rightarrow **significant FT expected**
- To quantify FT we ask how the value of M_Z depends upon **any of the 19 parameters**, $\{p_i\}$, up to (in some cases) the **2-loop, NLL level** (c/o **Martin & Vaughn**). We follow the FT approach of **Ellis et.al.** + **Barbieri & Giudice** :

$$A_i = \left| \partial \ln M_Z^2 / \partial \ln p_i \right|, \quad \Delta = \max \{A_i\}$$

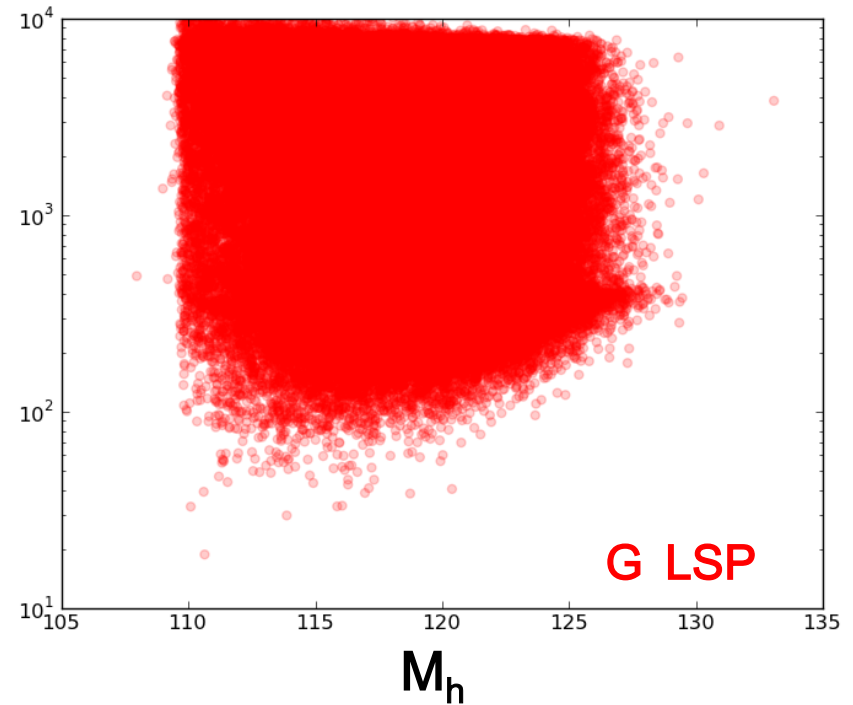
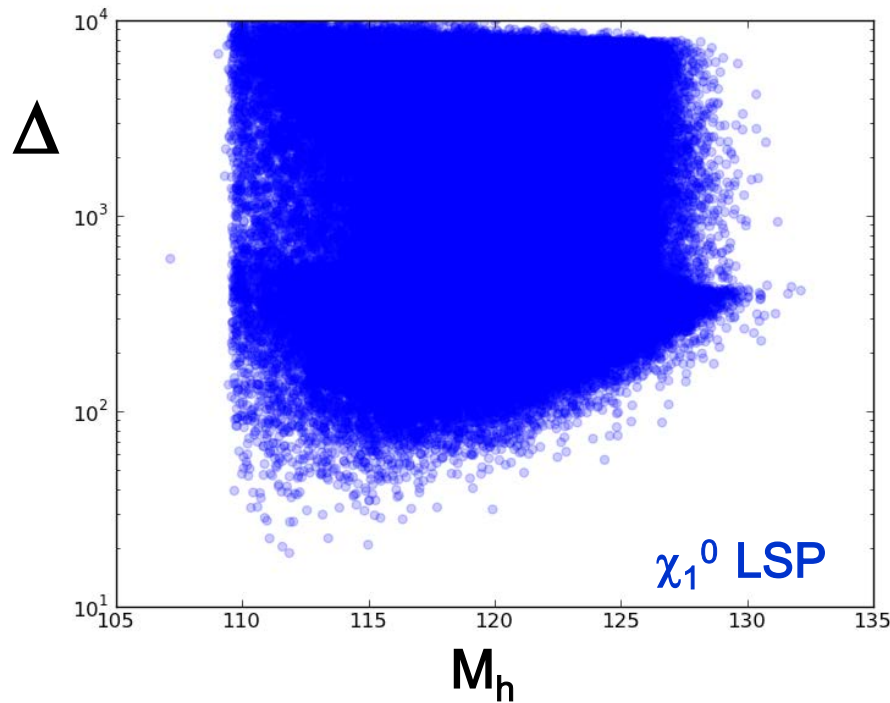
- **Specifically** we ask for the **number of models** with Δ less than a specific value...

Fine-tuning in the pMSSM



- Hence, as expected, the large Higgs mass 'cut' removes many of the models with the lowest FT values

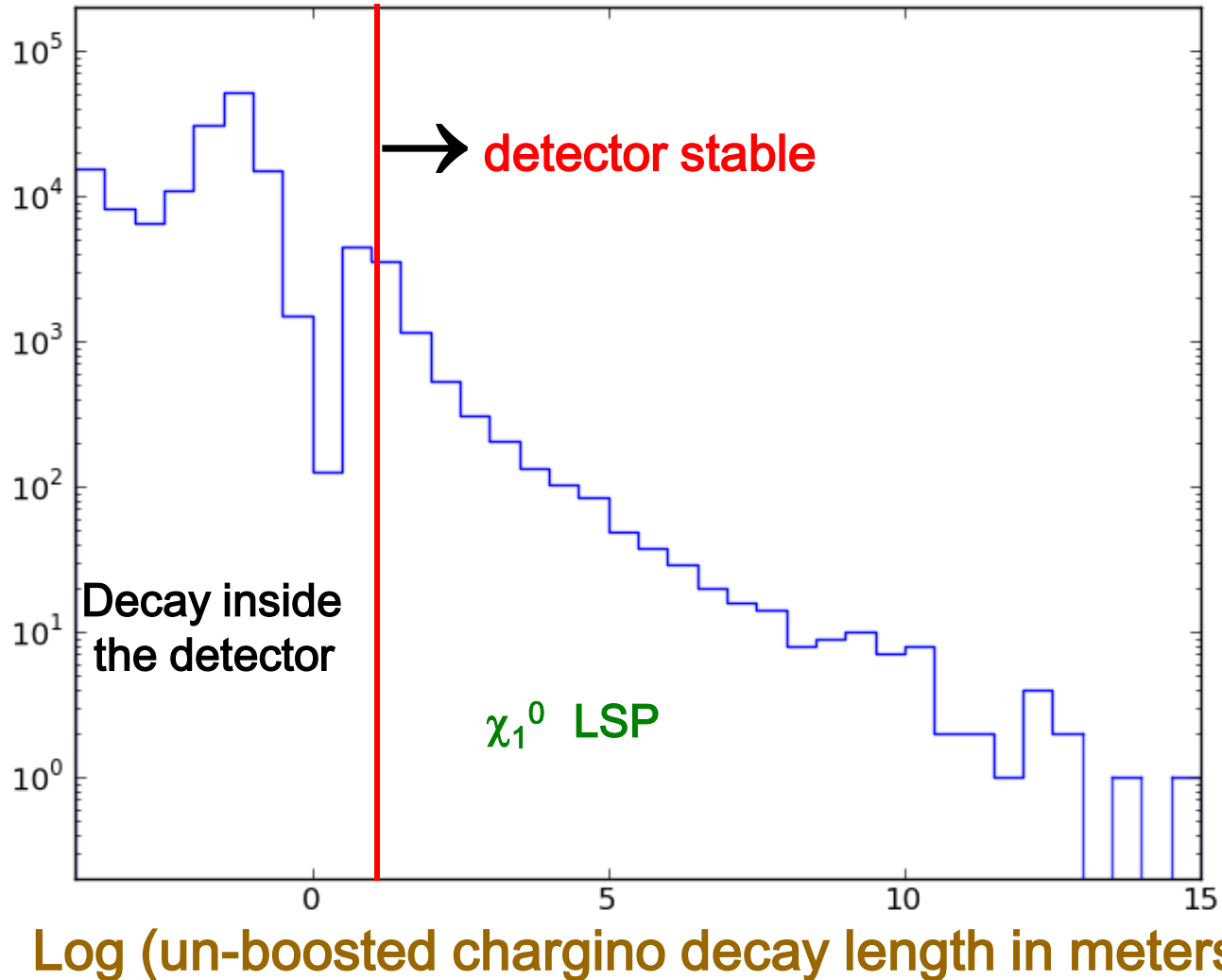
FT vs. Higgs mass distributions for both model sets



As is well-known, FT prefers lighter Higgs masses. Overall the G LSP models, on average, have slightly more FT than do χ LSP models.

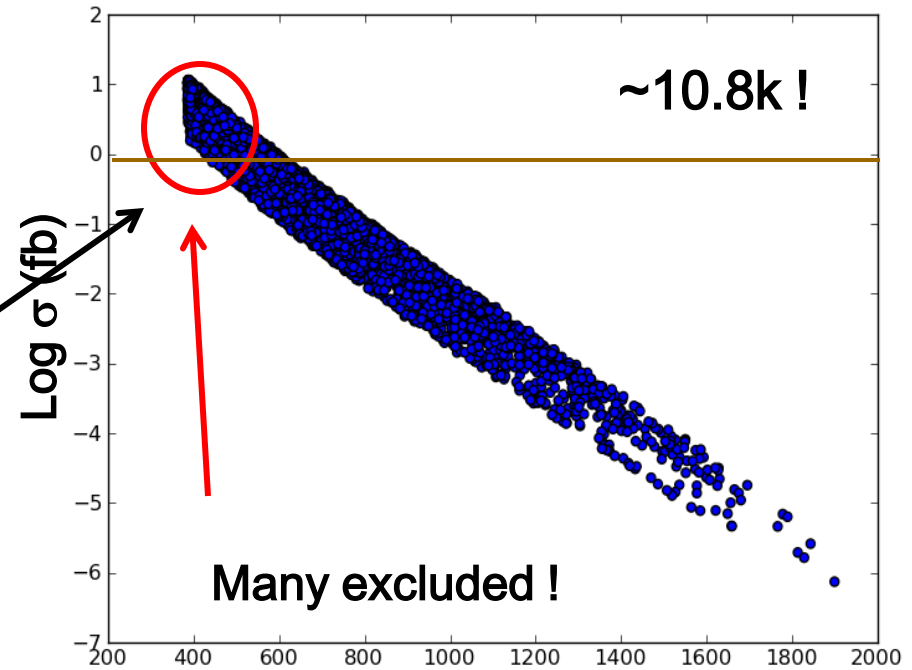
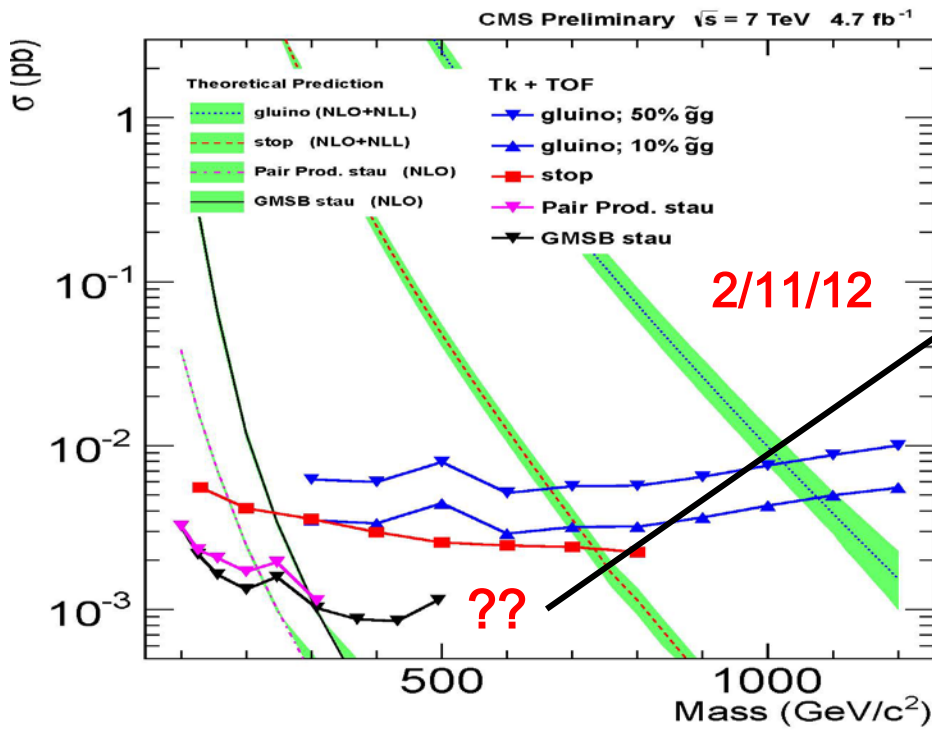
Long-Lived Sparticles: The Chargino Example

Most LSPs are nearly pure wino or Higgsino

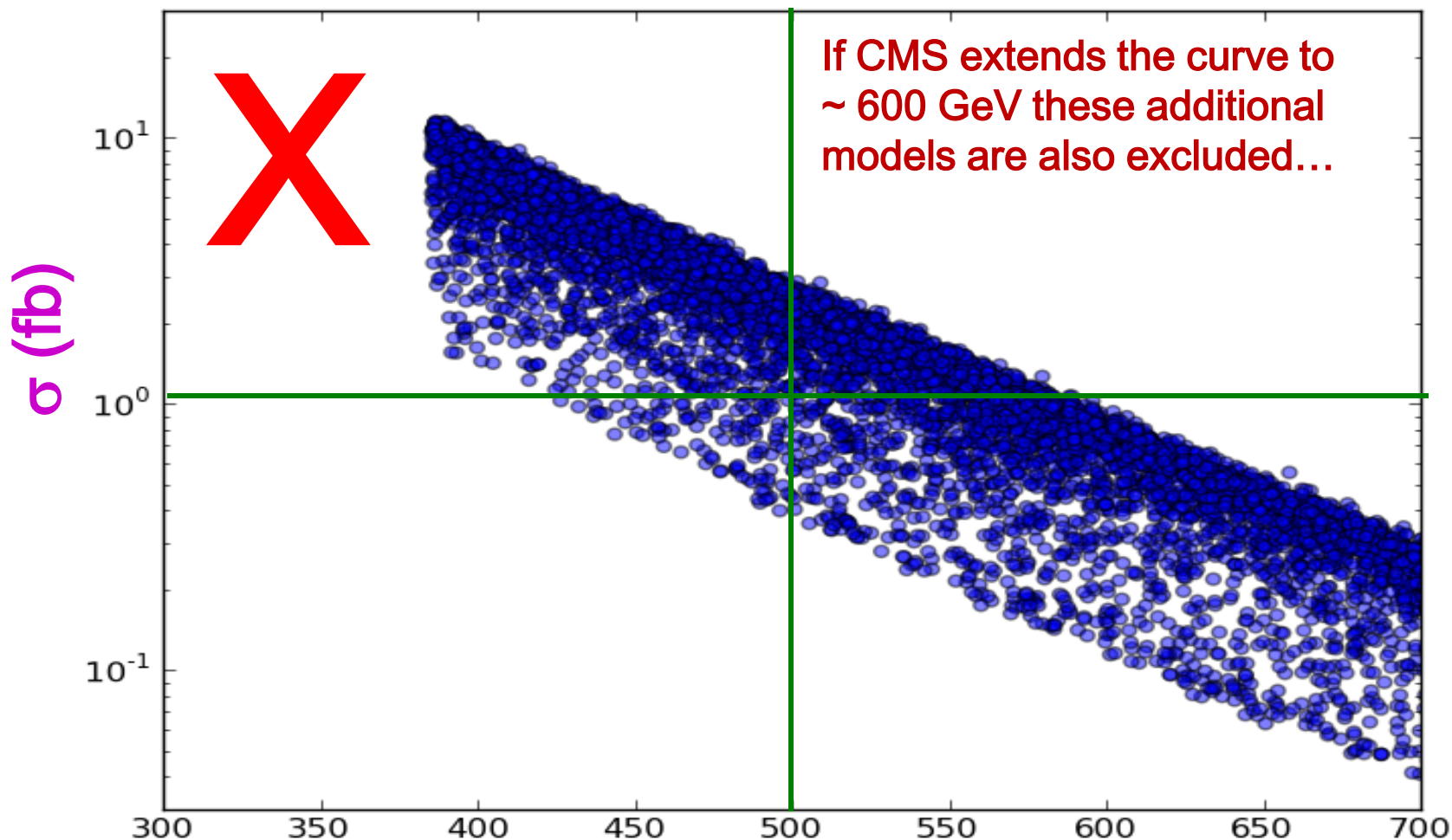


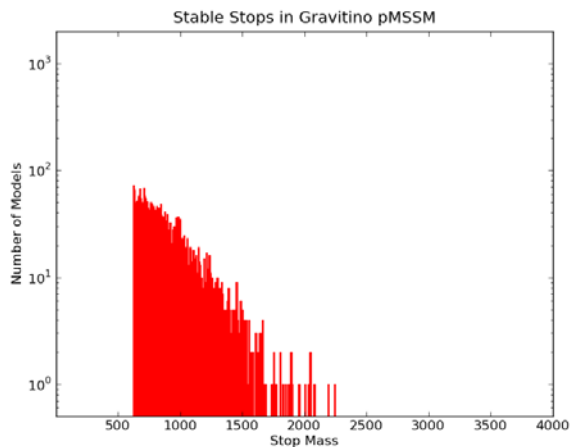
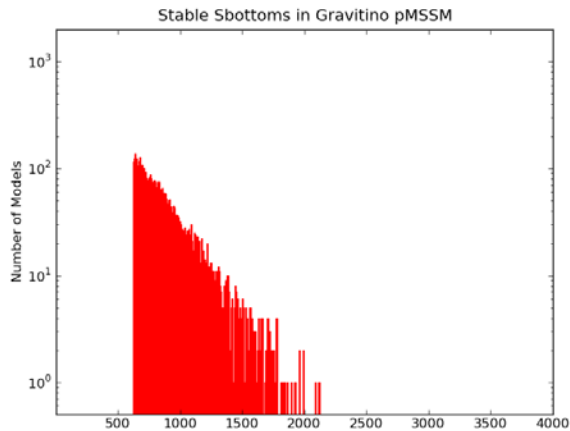
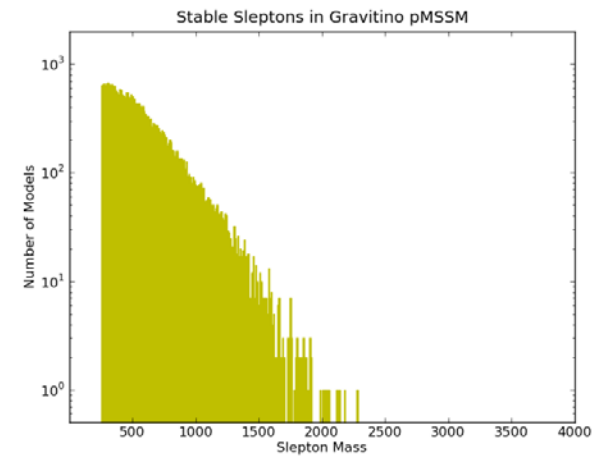
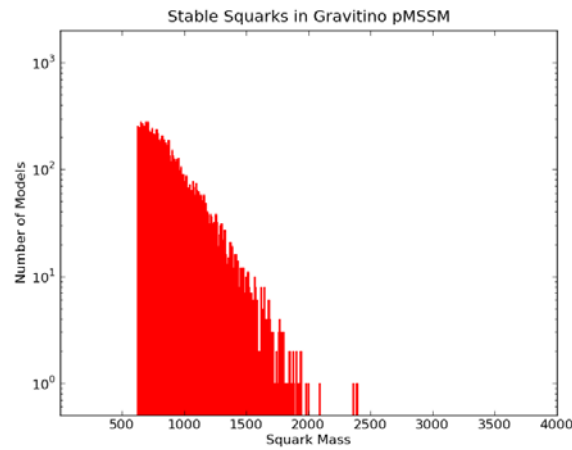
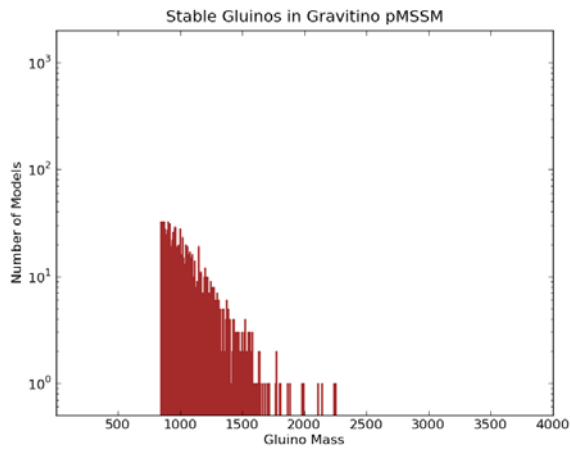
Detector Stable Charginos

- Searches for stable and/or long-lived sparticles can be quite powerful for both χ_1^0 or G LSP sets
- E.g., detector-stable charginos are quite common in χ_1^0 LSP models & extend out to large masses :



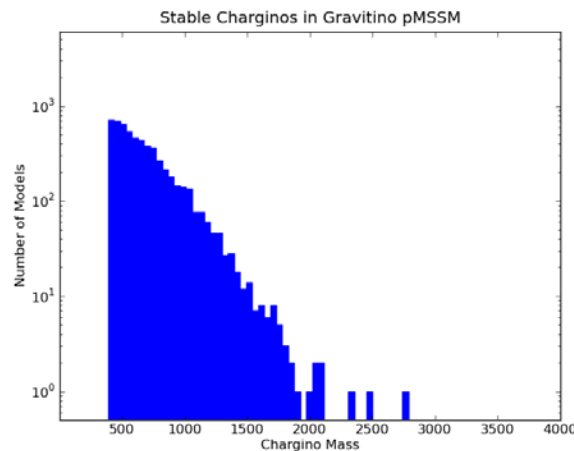
- **3581 (!!)** models (conservatively) are removed by stable particle searches w/ $\sim 5 \text{ fb}^{-1}$ @ 7 TeV





Gravitino LSP scenarios produce **many models** with **detector-stable** charged/colored sparticles over a very wide range of masses & species.

Specialized searches are required in some cases & to cover **decays inside the detector** (not shown here). This is work now in progress.



Summary & Conclusions

- The pMSSM with either neutralino or gravitino LSP shows a wide range of very interesting properties. The gravitino case has not been well explored until now & may yield some unexpected results
- LHC searches, both with & w/o MET, are cutting into these two model parameter spaces
- Going to 8 TeV will be a significant step in model coverage
- Higgs results will play a critical role in all future studies
- We look forward to the 8 TeV results in July -- Down Under!



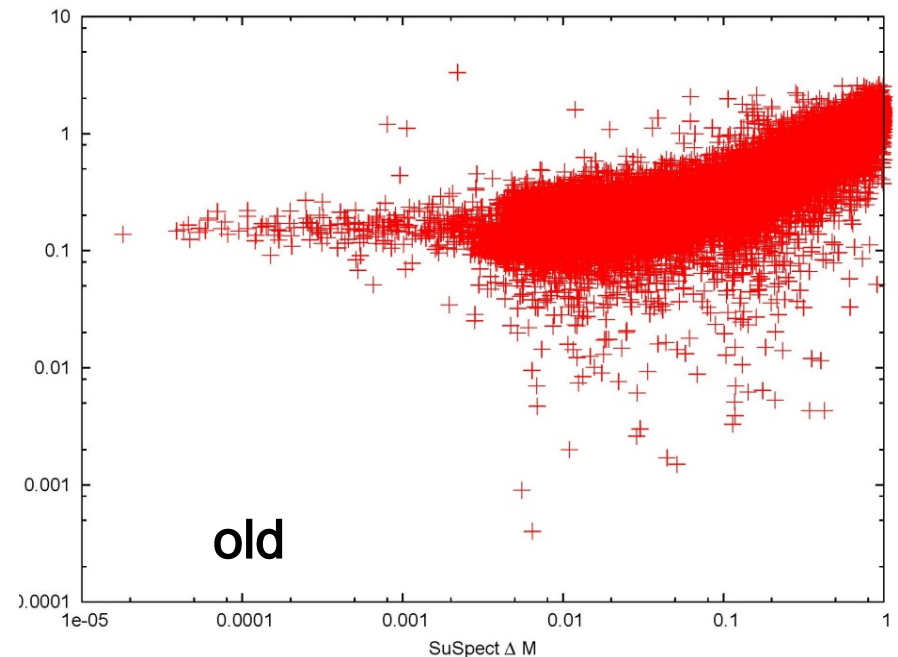
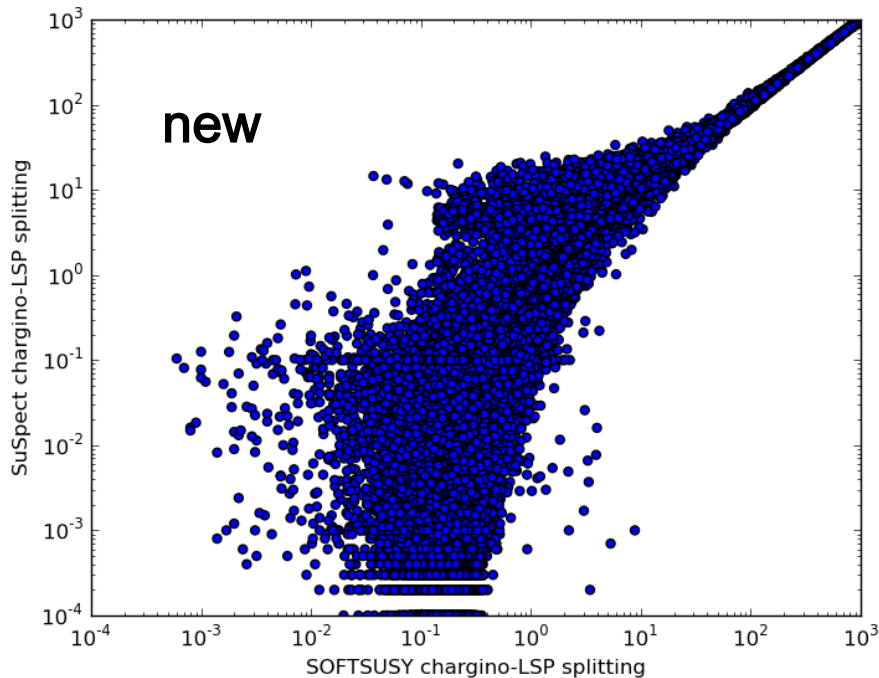
"Take a look at this everyone - it just could be the signature we've been looking for!"

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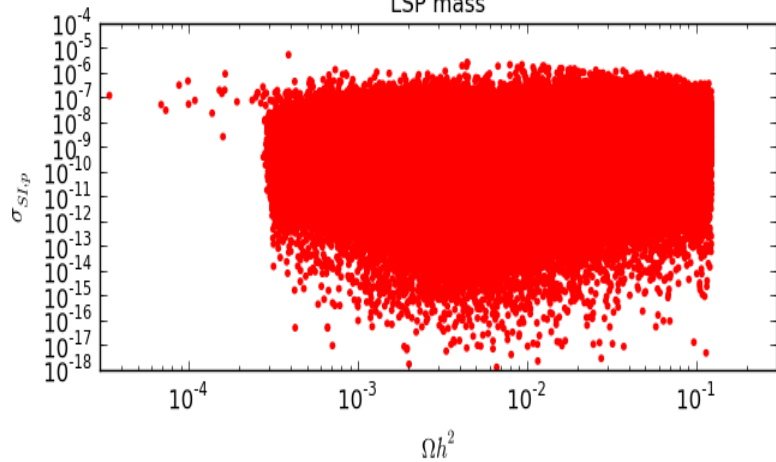
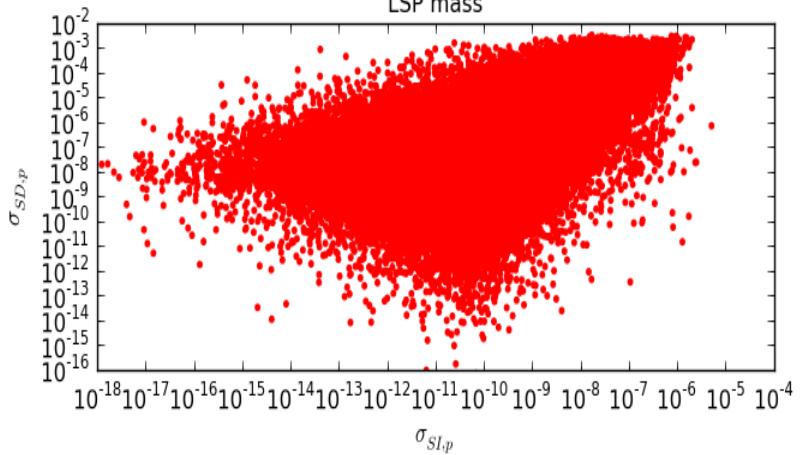
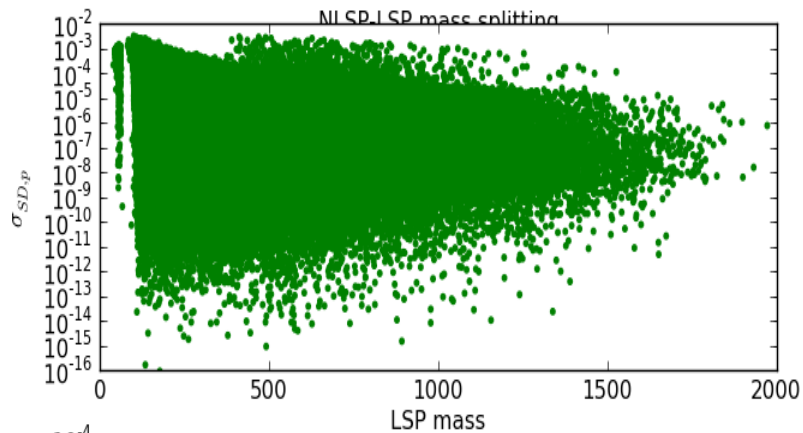
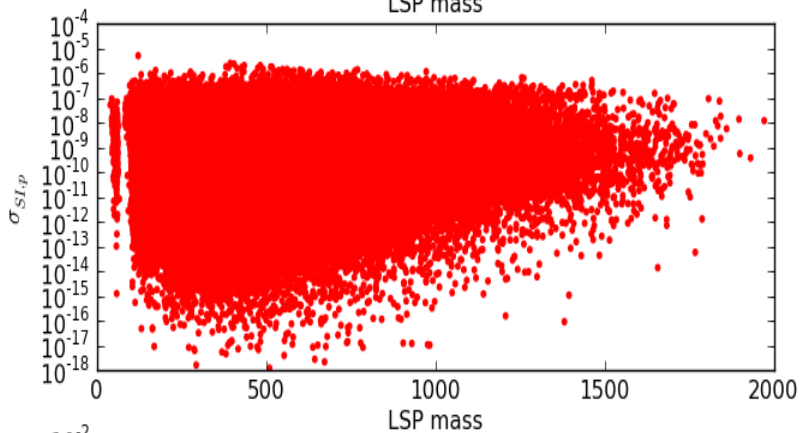
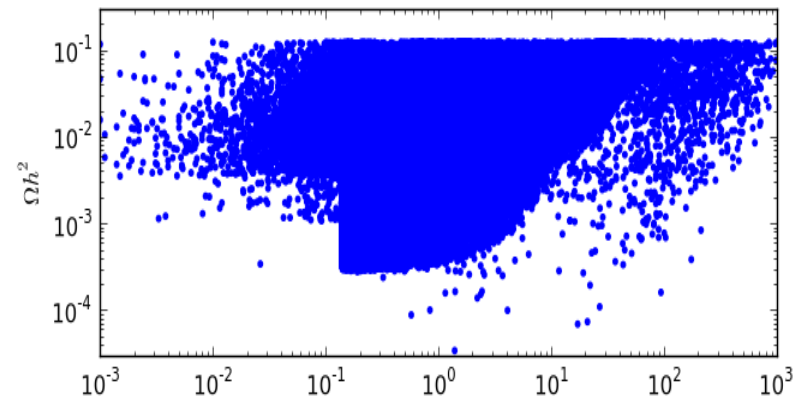
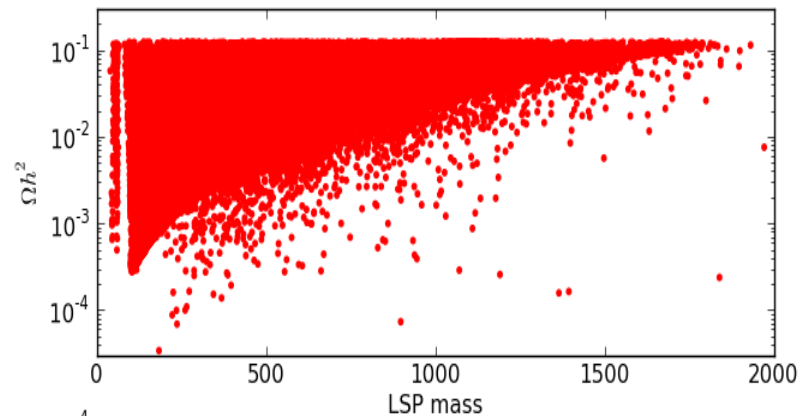
BACKUPS

Stable Chargino Spectrum Generator Issues

- **The issue still remains** as to just how many pMSSM models will lead to **long lived charginos**. Although the number is always large **SuSpect** predicts more **degeneracies** than does **SOFTSUSY**, i.e., it depends how & which RCs are included



χ_1^0 LSP DM Observables



The 19(20) Parameter pMSSM

10 sfermion masses: $m_{Q_1}, m_{Q_3}, m_{u_1}, m_{d_1}, m_{u_3}, m_{d_3}, m_{L_1},$
 $m_{L_3}, m_{e_1}, m_{e_3}$

3 gaugino masses: M_1, M_2, M_3

3 tri-linear couplings: A_b, A_t, A_τ

3 Higgs/Higgsino: $\mu, M_A, \tan\beta$

→→ (1 gravitino mass : $m_{3/2}$)

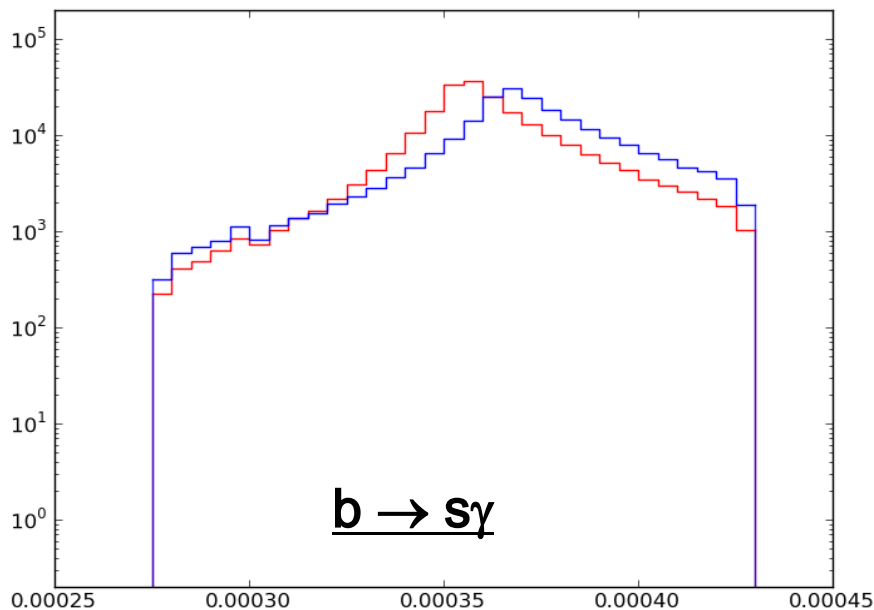
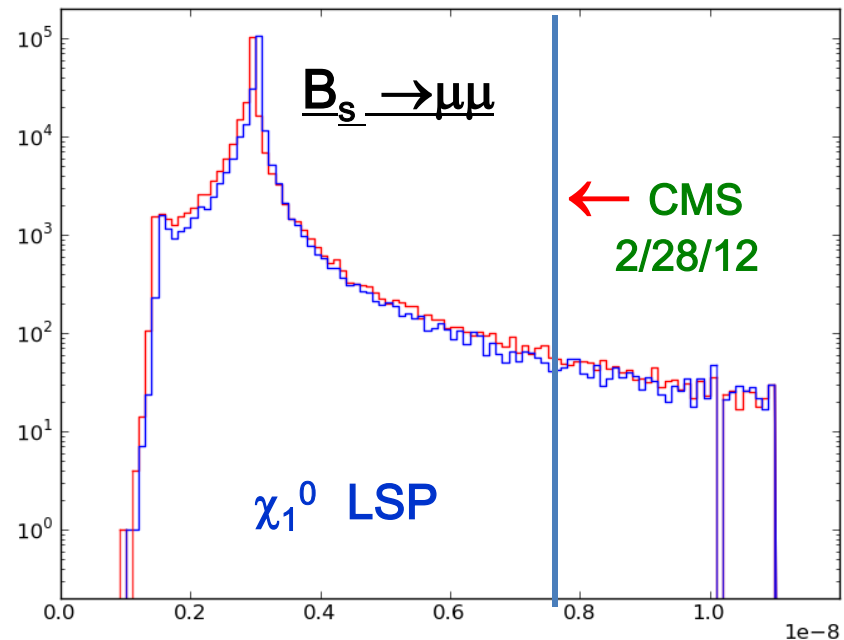
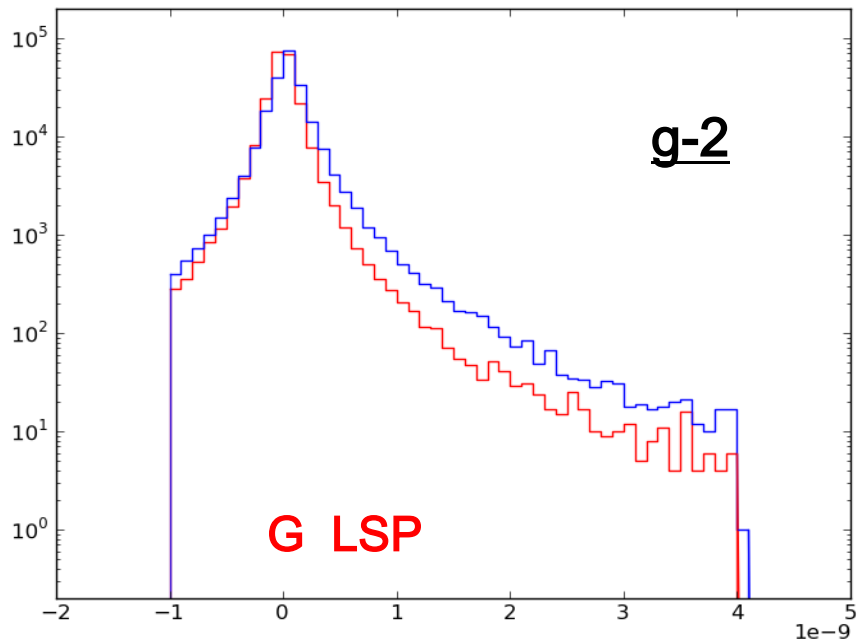
Choose the ranges of these parameters & how they're selected

Scan: look for points in this space satisfying all existing data & then study their signatures @ the LHC & elsewhere.. NO FITS!

Electroweak Content of χ_1^0

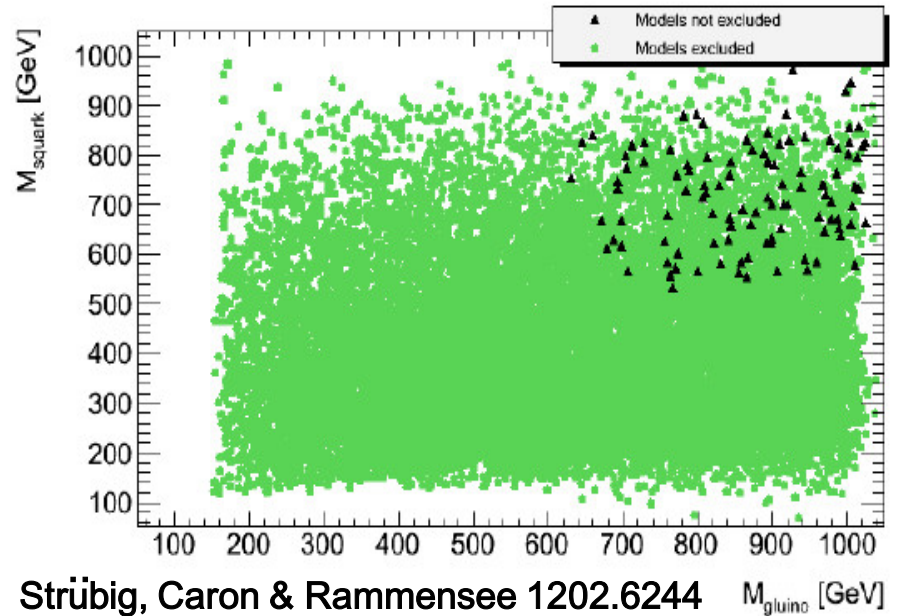
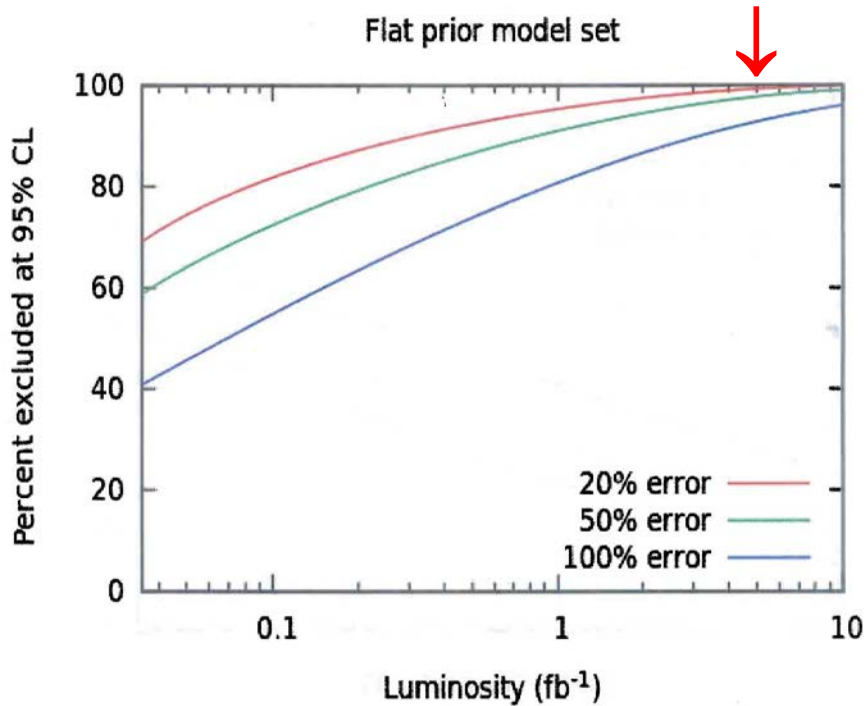
Lightest Neutralino	Definition	Neutralino LSP	Gravitino LSP
Bino	$ N_{11} ^2 > 0.95$	0.024	0.313
Mostly Bino	$0.80 < N_{11} ^2 < 0.95$	0.002	0.012
Wino	$ N_{12} ^2 > 0.95$	0.546	0.296
Mostly Wino	$0.80 < N_{12} ^2 < 0.95$	0.022	0.019
Higgsino	$ N_{13} ^2 + N_{14} ^2 > 0.95$	0.340	0.296
Mostly Higgsino	$0.80 < N_{13} ^2 + N_{14} ^2 < 0.95$	0.029	0.029
All other models	$ N_{11} ^2, N_{12} ^2, N_{13} ^2 + N_{14} ^2 < 0.80$	0.036	0.035

With most of the neutralino parameters ~ 1 TeV the mass & electroweak eigenstates are generally quite close !



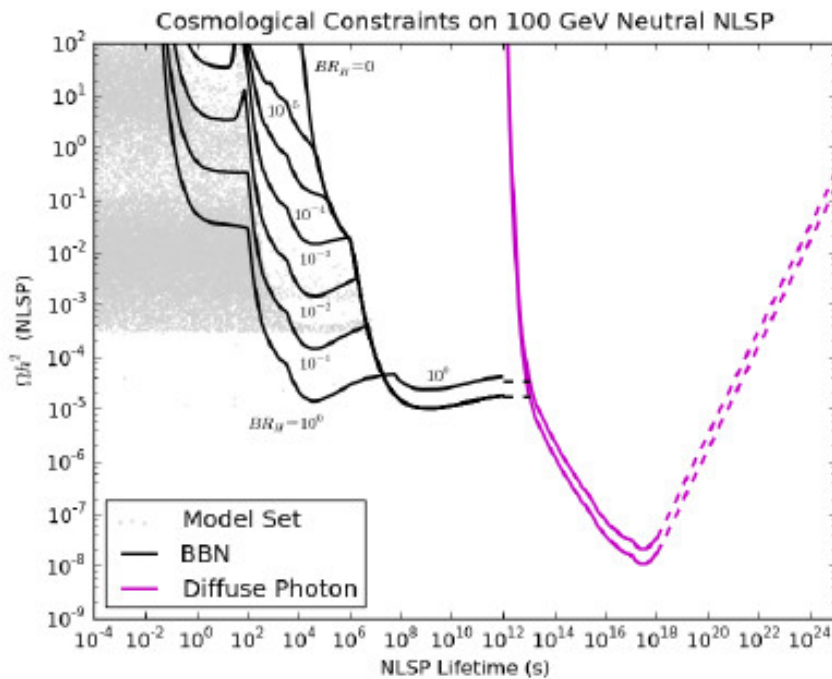
- Overall, these are quite minor differences . The important differences are the LSP itself & the NLSP
- CMS update kills 1041(1143) models in the neutralino (G) LSP set

- **ATLAS & CMS have ‘done a number’ on squark & gluinos below $\sim 0.8\text{-}1\text{ TeV}$ - - even in the pMSSM scenario :**



- **At most, only $\sim 1\%$ of our old $\sim 68\text{k}$ pMSSM model set w/ sparticle masses below 1 TeV now survive. They’re ‘dead but still walking’ ...**
→ **New pMSSM models needed!**

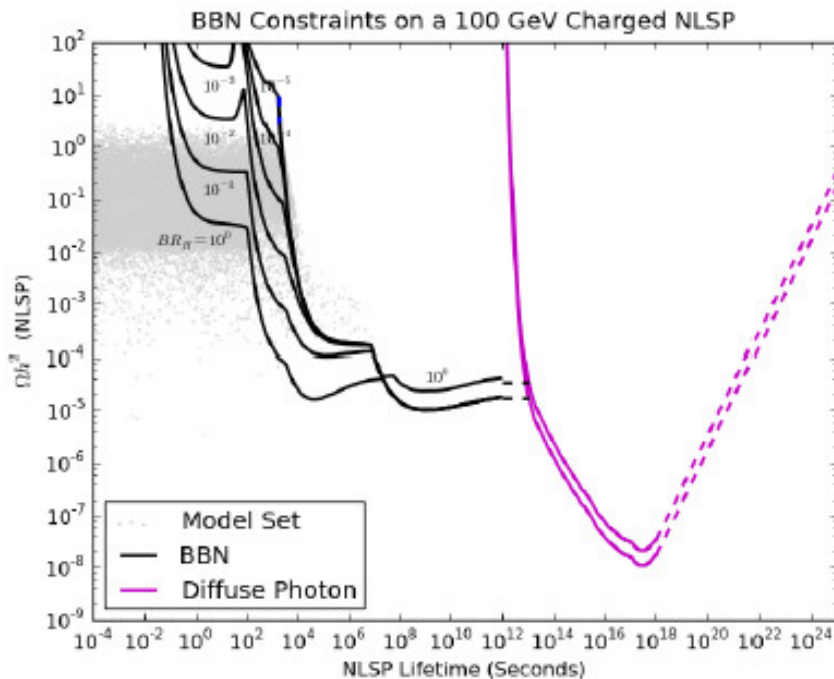




Sample constraints from
BBN and diffuse γ 's
for different hadronic
branching fractions
of the NLSP

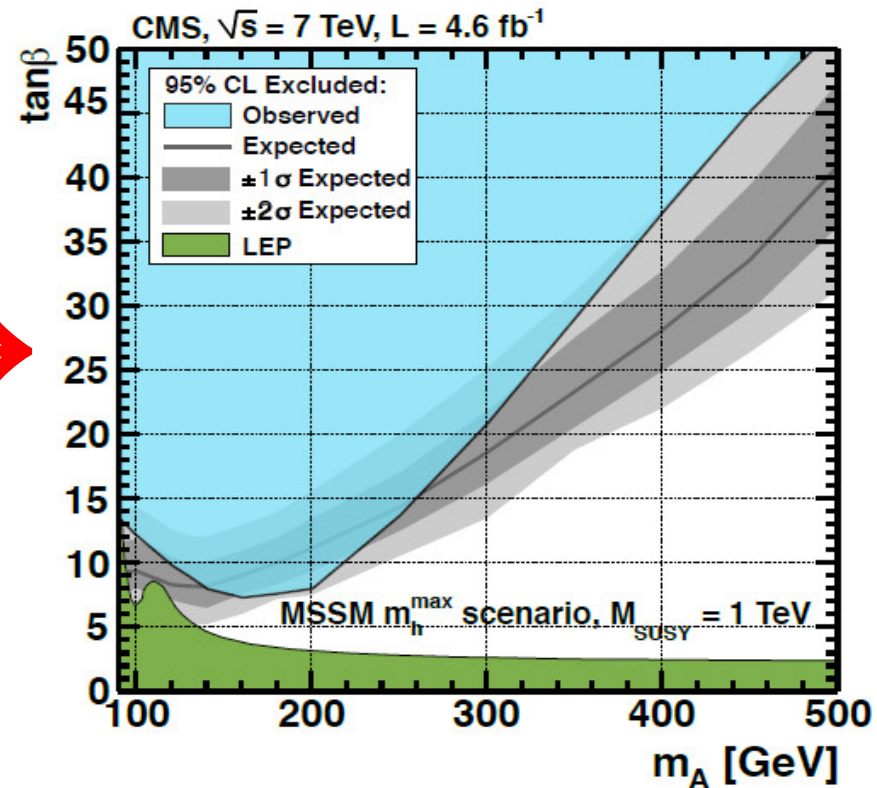
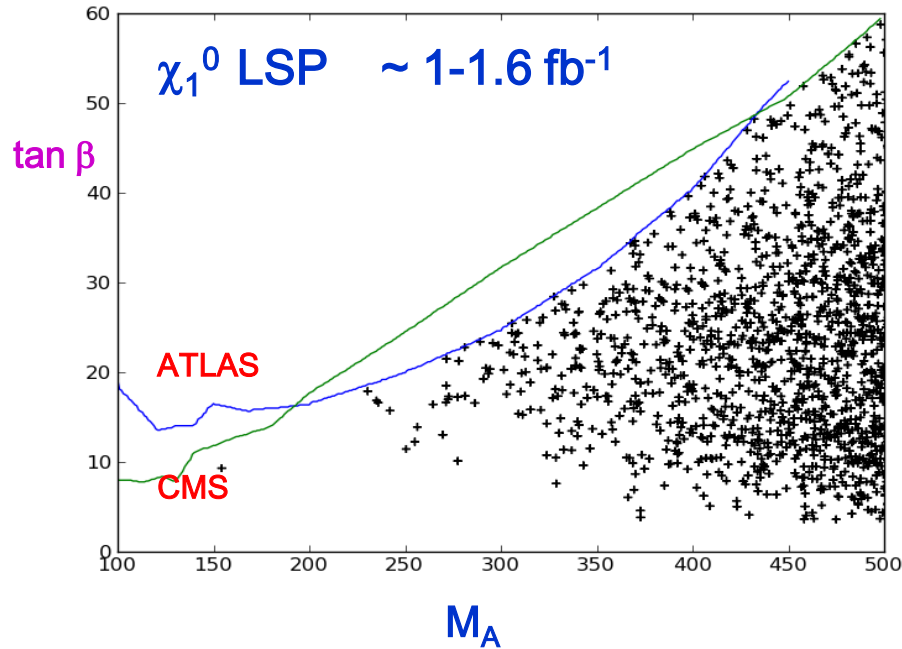
Shaded areas show where
our gravitino models live

We follow :



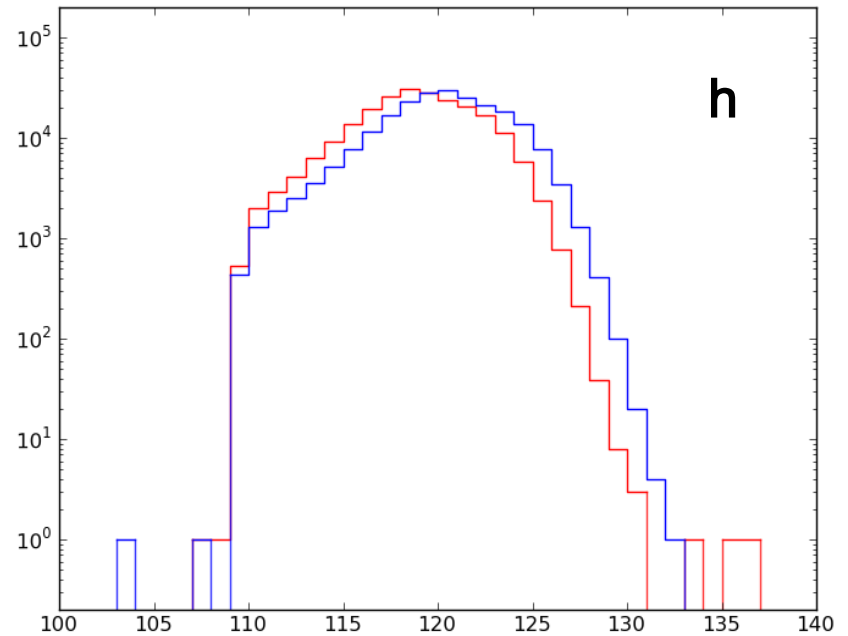
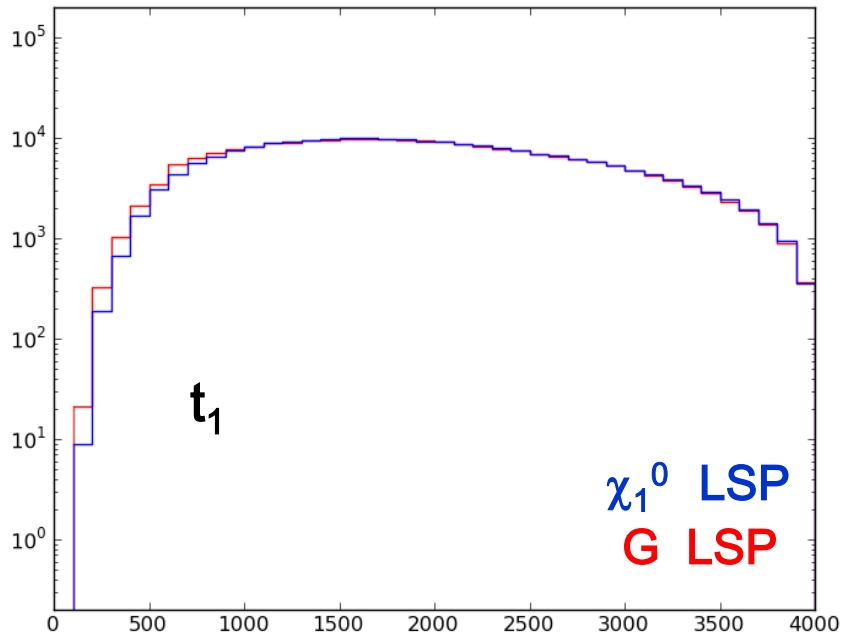
Jedamzik;
Kusakabe et al.;
Kanazki et al.;
Kribs and Rothstein

Impact of $A, H \rightarrow \tau\tau$ Searches



Increased lumi enhances coverage of the M_A - $\tan \beta$ plane

→ These searches have **important impact** other parts of the pMSSM..



- Although **the h mass** itself is **not part of the scan**, the slightly **different mass spectra in the G scan** contribute, via loops, to a somewhat **lighter Higgs**
- Some other observables show **similar size** effects but are **not really very different** in the two model sets

nLSP-LSP Mass Splitting

χ_1^0 LSP

