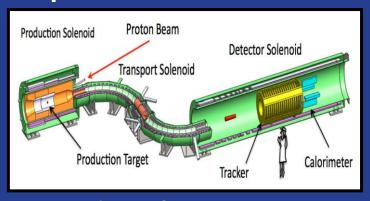


#### The Mu2e Experiment at Fermilab



Xth International Conference on Hyperons, Charm and Beauty Hadrons
Wichita, KS USA

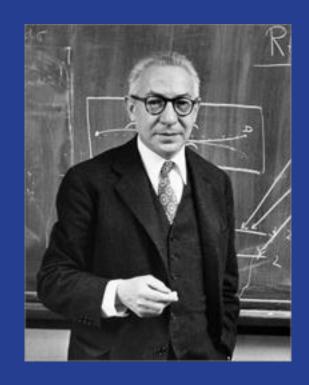
Andrew Norman, Fermilab For the Mu2e Collaboration





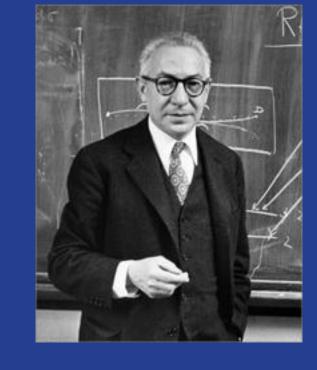
## A muon talk at a heavy flavor conference?

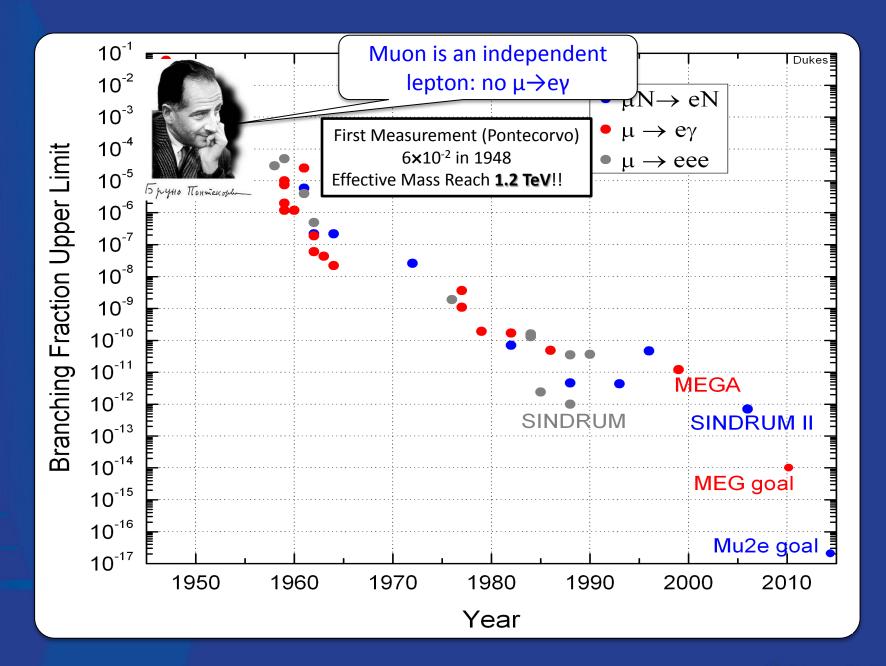
"Who ordered that?" —I.I.Rabi

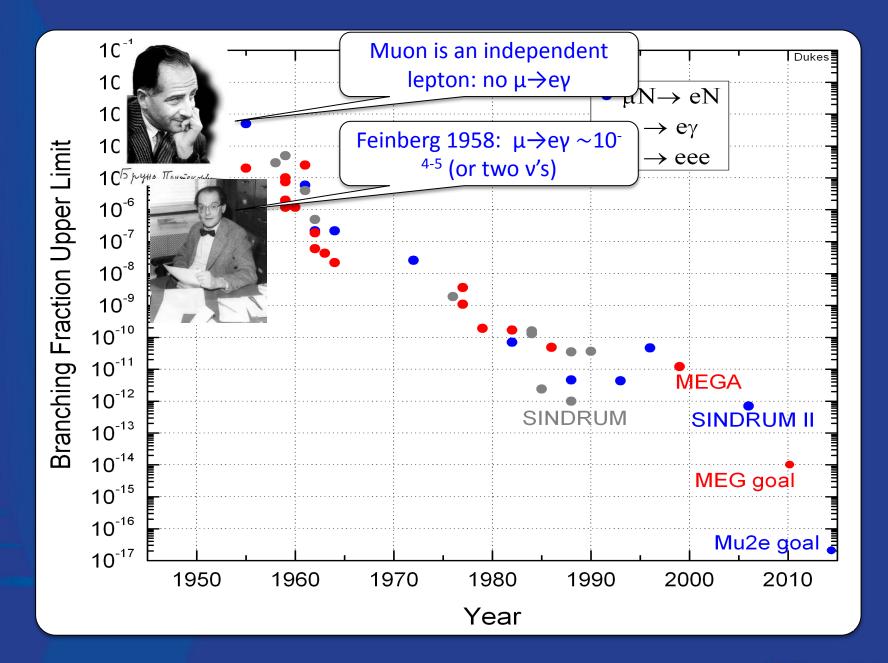


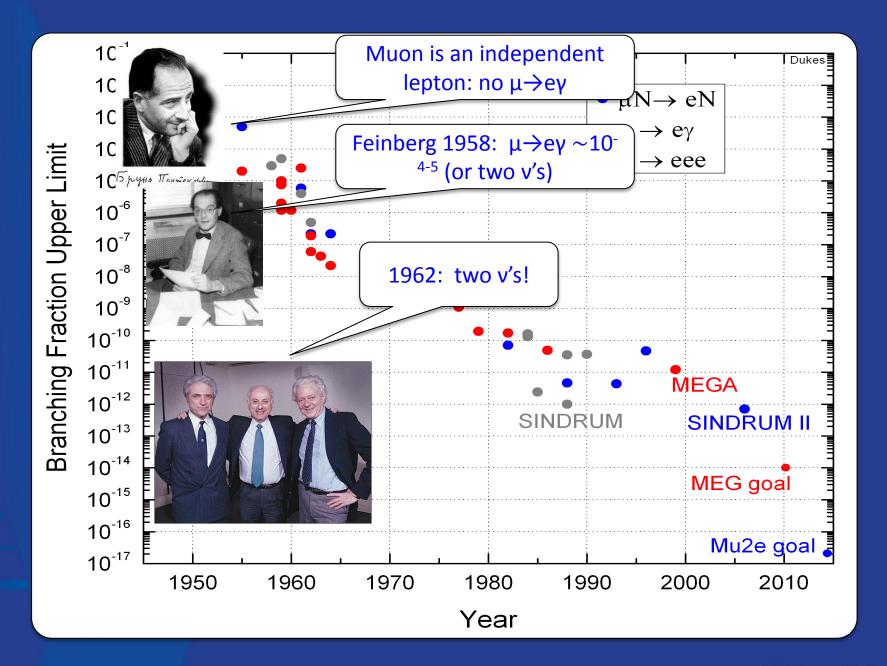
### "Who ordered that?"

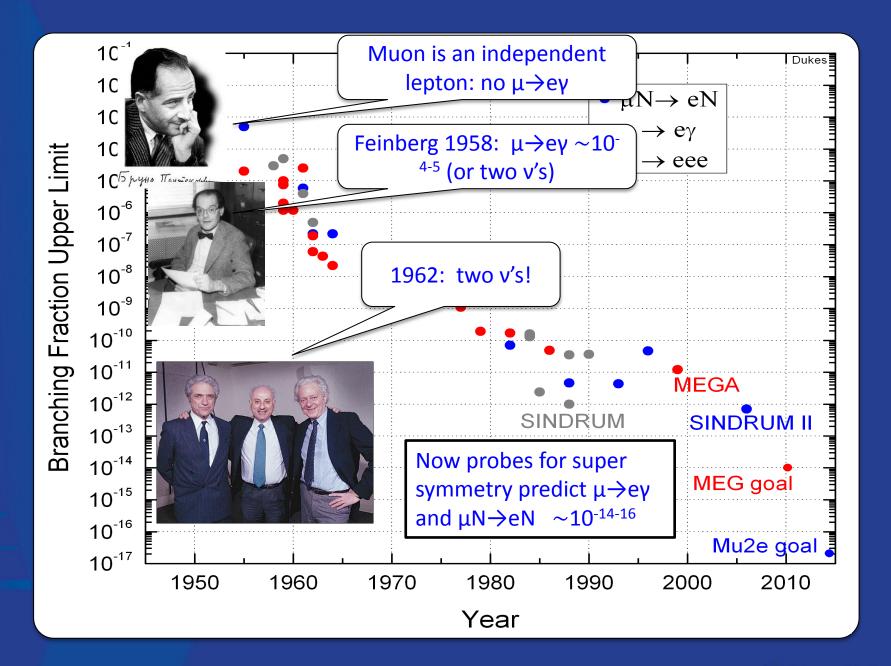
- When the μ was discovered it was logical to think of the μ as just an excited electron
  - So we would expect:
    - BR(μ→eγ)≈10<sup>-4</sup>
  - That is, unless another v, in an intermediate vector boson loop canceled it out. (Feinberg, 1958)
  - Same as GIM mechanism!
- Introduced the notion of lepton flavor





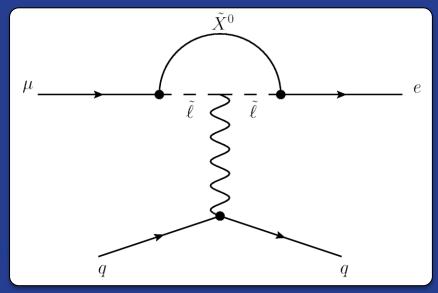






# Why Precision Measurements & Ultra-Rare Processes?

- We want to access physics beyond the standard model
  - This means access to High and Ultra-High Energy interactions
  - We get to these energies through loops
  - Getting at Loops means making precision measurements and looking for ultra-rare decays
- Ideally we start with processes that are forbidden or highly suppressed in the standard model
  - Any observation becomes proof of non-SM physics



# Why Precision Measurements & Ultra-Rare Processes?

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  - Any observation becomes proof of non-SM physics
- Flavor Changing Neutral Currents
  - FCNC in quark sector
    - $B_s \rightarrow \mu\mu$  ,  $b \rightarrow s\gamma$  ,  $K \rightarrow \pi v v$
    - Allowed but HIGHLY suppressed in Standard Model
    - Can receive LARGE enhancements in SUSY and other beyond-SM physics
  - FCNC in charged lepton sector
    - $\mu \rightarrow e\gamma$ ,  $\mu \rightarrow eee$ ,  $\mu N \rightarrow e N$  (Lepton Flavor Violating)
    - No SM amplitudes (except via v loops)
    - Permitted in beyond-SM models, and have extreme reach in energy

# Lepton Mixing in the Standard Model

We have three generations of leptons:

$$\begin{pmatrix} e \\ \nu_e \end{pmatrix} \begin{pmatrix} \mu \\ \nu_\mu \end{pmatrix} \begin{pmatrix} \tau \\ \nu_\tau \end{pmatrix}$$

 $\begin{pmatrix} e \\ \nu_e \end{pmatrix} \begin{pmatrix} \mu \\ \nu_{\mu} \end{pmatrix} \begin{pmatrix} \tau \\ \nu_{\tau} \end{pmatrix}$  No SM couplings between generation!

- In the standard model Lagrangian there is no coupling to mixing between generations
- But we have explicitly observed neutrino oscillations
- Thus charged lepton flavor is **not** conserved.
- Charged leptons must mix through neutrino loops

$$Br(\mu \to e\gamma) = \frac{3\alpha}{32\pi} \left| \sum_{\ell} V_{\mu\ell}^{\star} V_{e\ell} \frac{m_{\nu_{\ell}}^2}{M_W^2} \right|^2$$

$$< 10^{-54}$$

But the mixing is so small, it's effectively forbidden

# Charged Lepton Flavor Violation (CLFV) Processes with µ's

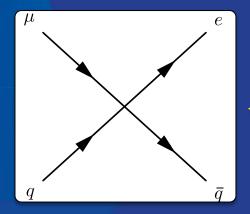
 There are three basic channels to search for μ-CLFV in:

$$\mu^{+} \to e^{+} \gamma$$

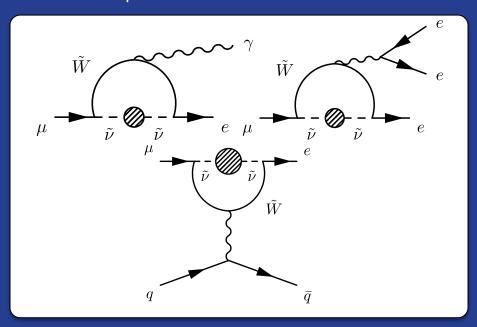
$$\mu^{+} \to e^{+} e^{+} e^{-}$$

$$\mu^{-} N \to e^{-} N$$

 If loop like interactions dominate we expect a ratio of these rates:



 New physics for these channels can come from loop level



For  $\mu N \rightarrow eN$  and  $\mu \rightarrow eee$  we also can have contact terms

If contact terms dominate then μN→eN can have rates 200 times that of μ→eγ

# Charged Lepton Flavor Violation (CLFV) Processes with µ's

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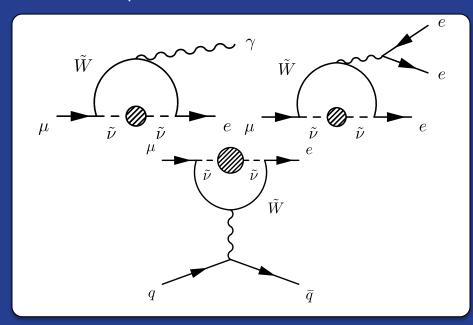
 If loop like interactions dominate we expect a ratio of these rates:

Note:  $\mu \rightarrow e \gamma$  and  $\mu \rightarrow e e e$  have experimental limitations (resolution, overlap, accidentals)

Ultimately Limits the measurement of:  $Br(\mu \rightarrow e\gamma) \approx 10^{-14}$ 

No such limits on µN→eN channel

 New physics for these channels can come from loop level

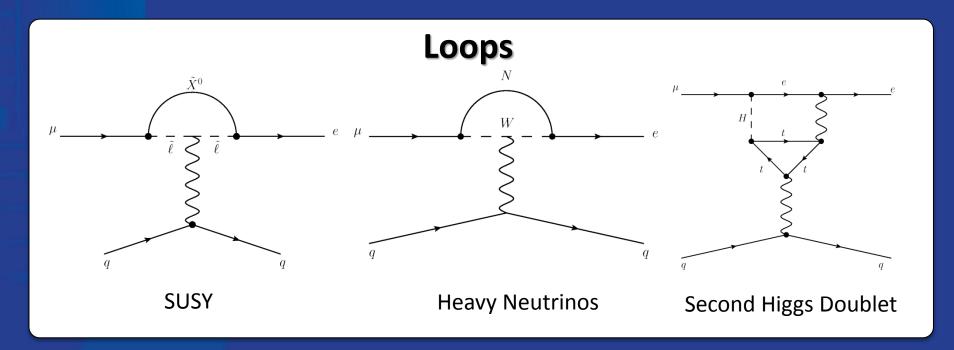


For μN→eN and μ→eee we also can have contact terms

If contact terms dominate then  $\mu N \rightarrow e N$  can have rates 200 times that of  $\mu \rightarrow e \gamma$ 

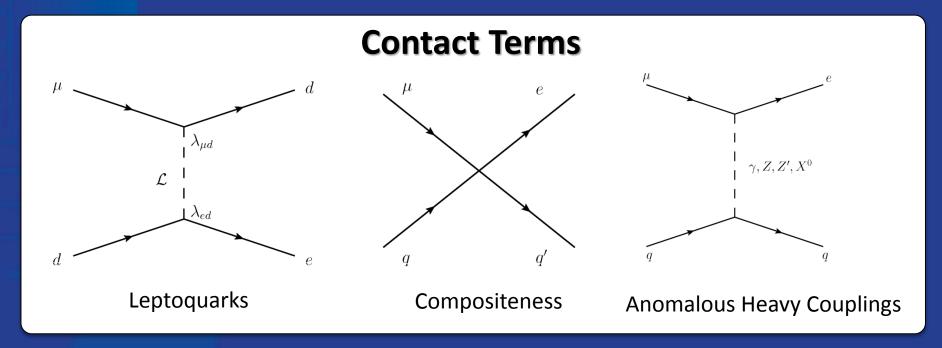
# Beyond the Standard Model

 The CLFV process can manifest in the μN→eN channel in many models with large branching fractions:



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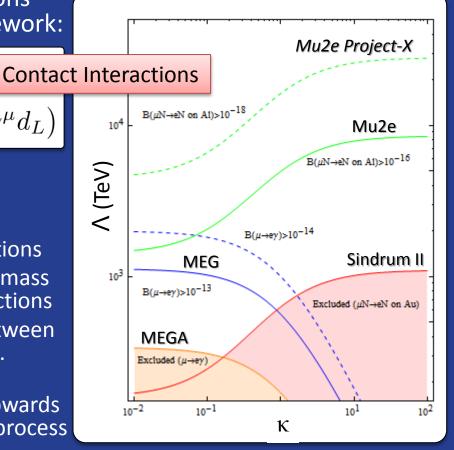


# General CLFV Lagrangian

 Recharacterize these all these interactions together in a model independent framework:

$$\mathcal{L}_{\mathcal{LFV}} = \frac{m_{\mu}}{(\kappa+1)\Lambda^{2}} \bar{\mu}_{R} \sigma_{\mu\nu} e_{L} F^{\mu\nu}$$
 Contact 
$$+ \frac{\kappa}{(1+\kappa)\Lambda^{2}} \bar{\mu}_{L} \gamma_{\mu} e_{L} \left( \bar{u}_{L} \gamma^{\mu} u_{L} + \bar{d}_{L} \gamma^{\mu} d_{L} \right)$$

- Splits CLFV sensitivity into
  - Loop terms
  - Contact terms
- Shows dipole, vector and scalar interactions
- Allows us to parameterize the effective mass scale ¤ in terms of the dominant interactions
- The balance in effective reach shifts between favoring <sup>1</sup>N!eN and <sup>1</sup>!e° measurements .
- For contact term dominated interaction (large  $\kappa$ ) the sensitivity in  $\Lambda$ , reaches upwards of 10<sup>4</sup> TeV for the coherent conversion process



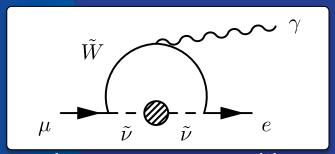
# Experimental Limits vs. SU(5) SUSY-GUT

SUSY predictions for CLFV processes are only a few orders of magnitude below current experimental limits

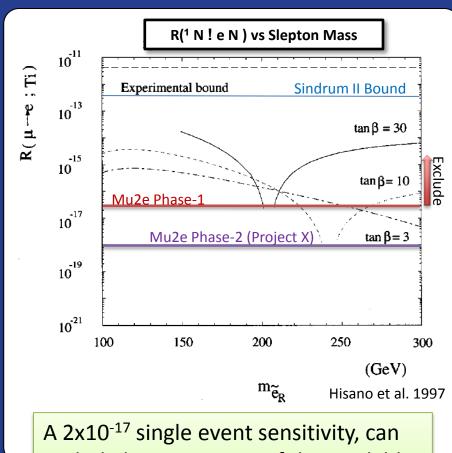
Process	Current Limit	SUSY-GUT level
$\mu \: \textbf{N} \to \textbf{e} \: \textbf{N}$	<b>7 x10<sup>-13</sup></b> W. Bertl, et al EPJ C47(06)337	<b>10</b> <sup>-16</sup>
$\mu  o$ e $\gamma$	<b>2.4 x10</b> <sup>-12</sup> J. Adam, et al PRL 107(11)171801	10 <sup>-14</sup>
$ au o \mu\gamma$	<b>4.5 x10</b> -8  K. Hayasaka, et al PL B666(08)16	10 <sup>-9</sup>

## µN→eN Sensitivity to SUSY

 Rates are not small because they are set by the SUSY mass scale



- For low energy SUSY like we would see at the LHC:
   Br(μN→eN) ~10<sup>-15</sup>
- Makes µN→eN compelling, since for Mu2e this would mean observation of ≈O(40) events [0.5 bkg]

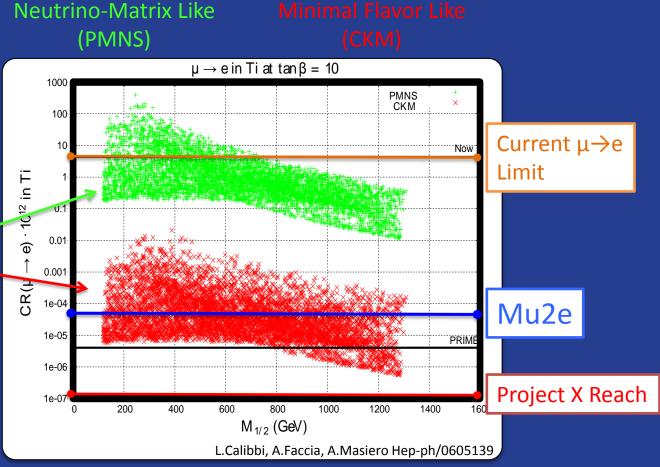


A 2x10<sup>-17</sup> single event sensitivity, can exclude large portions of the available SUSY parameter spaces

A.Norman, FNAL BEACH2012 BEACH2012

### Tests of SUSY Frameworks

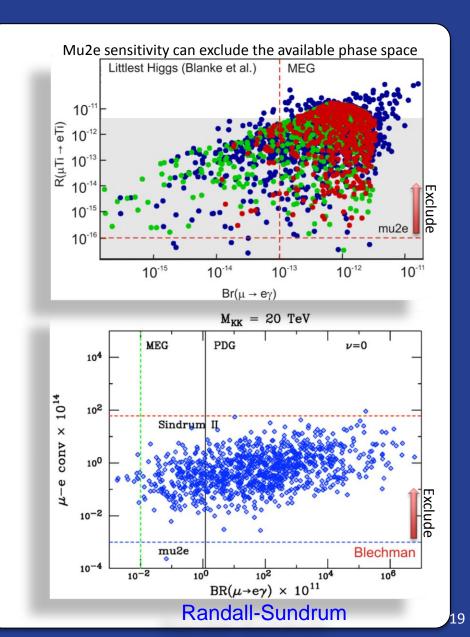
μ→e measurement can distinguish between PMNS and MFV mixing structures in SUSY frameworks



Example: neutrino masses via the seesaw mechanism, analysis is performed in an SO(10 framework). Different predictions for  $\mu e$  conversion with mixing structure.

## $\mu N \rightarrow eN$ , $\mu \rightarrow e\gamma$ , g-2 Work Together

- Knowing μN→eN , μ→eγ allow us to exclude SUSY phase space
- Also knowing the g-2 results allows us to then over constrain SUSY models
- In some cases this permits us to make strong, testable predictions for our models in terms of Br(μ→eγ) & R(μN→eN)



A.Norman, FNAL

BEACH.

## $\mu N \rightarrow eN$ , $\mu \rightarrow e\gamma$ , g-2 Work Together

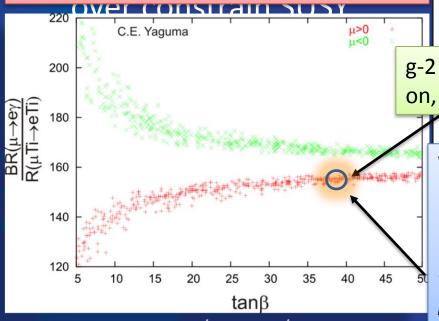
BEACH

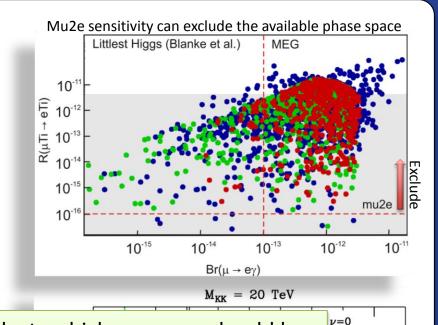
#### Example:

- From LHC we have the SUSY masses
- From g-2 we know tanβ
- From g-2 we know also know μ>0
- Combining these we get an a priori PREDICTION for:

$$\frac{Br(\mu \to e\gamma)}{R(\mu N \to eN)}$$

under MSSM/MSUGRA





g-2 selects which curve we should be on, and gives us the value of tan

We measure  $R(\mu N \rightarrow eN)$  and take the ratio to the MEG result.

Sindrum II

We use this match to prediction as a way to disentangle, or validate, or interpret manifestations of SUSY

chman

### SUSY

Many search modes have large effects for some models

#### But only:

- μ→eγ
- μe conversion



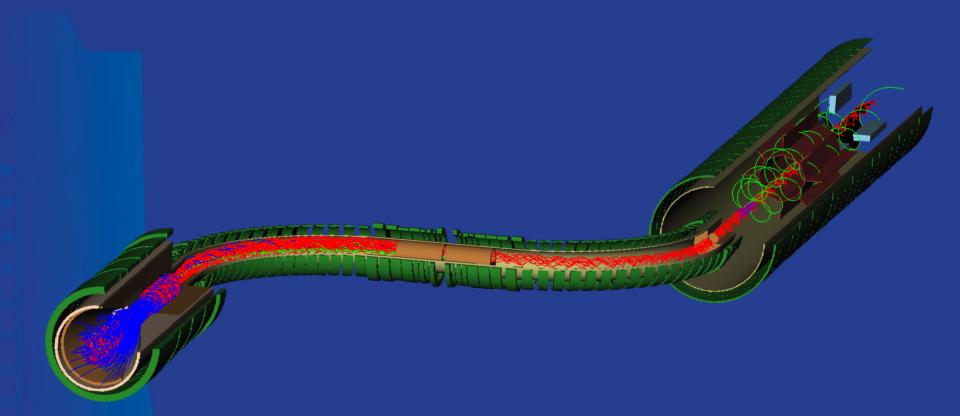




	AC	RVV2	AKM	$\delta \mathrm{LL}$	FBMSSM	LHT	RS
$D^0 - \bar{D}^0$	***	*	*	*	*	***	?
$\epsilon_K$	*	***	***	*	*	**	***
$S_{\psi\phi}$	***	***	***	*	*	***	***
$S_{\phi K_S}$	***	**	*	***	***	*	?
$A_{\rm CP}\left(B \to X_s \gamma\right)$	*	*	*	***	***	*	?
$A_{7,8}(B \to K^* \mu^+ \mu^-)$	*	*	*	***	***	**	?
$A_9(B \to K^* \mu^+ \mu^-)$	*	*	*	*	*	*	?
$B  o K^{(*)} \nu \bar{\nu}$	*	*	*	*	*	*	*
$B_s \to \mu^+ \mu^-$	***	***	***	***	***	*	*
$K^+  o \pi^+  u \bar{ u}$	*	*	*	*	*	***	***
$K_L  o \pi^0 \nu \bar{\nu}$	*	*	*	*	*	***	***
$\mu \to e \gamma$	***	***	***	***	***	***	***
$\tau \to \mu \gamma$	***	***	*	***	***	***	***
$\mu + N \rightarrow e + N$	***	***	***	***	***	***	***
$d_n$	***	***	***	**	***	*	***
$d_e$	***	***	**	*	***	*	***
$(g-2)_{\mu}$	***	***	**	***	***	*	?

Table 8: "DNA" of flavour physics effects for the most interesting observables in a selection of SUSY and non-SUSY models  $\bigstar \star \star \star$  signals large effects,  $\star \star$  visible but small effects and  $\star$  implies that the given model does not predict sizable effects in that observable.

W.Altmanshofer et al., arXiv:0909.1333v2 [hep-ph]



Ordering up  $\mu N \rightarrow eN$  at  $10^{-16}$ 

### **MAKING THE MEASUREMENT**

# The $\mu N \rightarrow eN$ measurement at Br(10<sup>-17</sup>) (in a nutshell)

- Stop  $\sim$ O(5×10<sup>10</sup>)  $\mu$  per pulse on a target (Al, Ti, Au)
- Wait 700ns (to let prompt backgrounds clear)
- Look for the coherent conversion of a muon to a monoenergetic electron:

$$E_e = M_{\mu} - N_{recoil} - (B.E.)_{\mu}^{1S}$$
  
= 104.96 MeV (on <sup>27</sup>Al)

Report the rate relative to nuclear capture

$$\mathcal{R} = \frac{\Gamma(\mu^- N \to e^- N)}{\Gamma(\mu^- N(Z) \to \nu_\mu N(Z-1)d}$$

 If we see a signal, it's compelling evidence for physics beyond the standard model!

#### µN→eN in Detail

#### **Muonic Atom**

- Start with a series of target foils
  - For Mu2E these are Al or Ti
- Bring in the low energy muon beam
  - We stop ≈ 50% of µ's
  - Stopped muons fall into the atomic potential
  - As they do they emit x-rays
- Muons fall down to the 1S state and a captured in the orbit
  - Muonic Bohr Radius

or Size 
$$\langle r_{\mu} 
angle = rac{n^2 \hbar}{m_{\mu} z e^2} pprox 19.6 \; {
m fm} \; ({
m Al})$$

Nuclear Size

$$R \approx 1.2A^{1/3} \text{fm} = 3.6 \text{ fm (Al)}$$

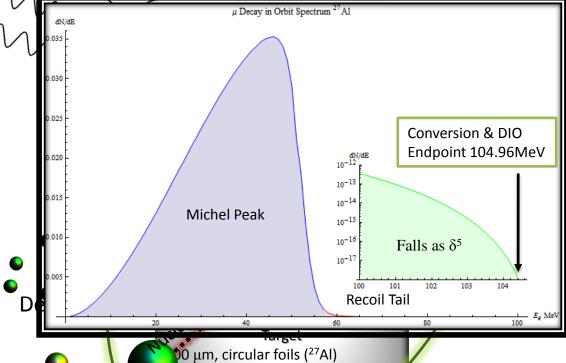
- Provides large overlap in the muon's wavefunction with the nucleous's
- For Z > 25 the muon is "inside" the nucleous
- Once captured 3 things can happen
  - Decay in Orbit:

$$\mu^- \to e^- \nu_\mu \bar{\nu}_e$$

We use the cascade of muonic x-rays and the well known spectrum to normalize the experiment.

lal time)

#### **1S Muonic Aluminum**



adius tapers from 10 cm to 6.5 cm 5cm spacing between foils

Lifetime: 864ns
DIO Fraction: 39.3%

Capture Fraction: 60.7%

#### **Muonic Atom**

- Start with a series of target foils
  - We stop ≈ 50% of µ's
- Bring in the low energy muon beam
  - We stop  $\cong$  50% of  $\mu$ 's
  - Stopped muons fall into the atomic potential
  - As they do they emit x-rays
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• Muonic Bohr Radius 
$$\langle r_{\mu} \rangle = \frac{n^2 \hbar}{m_{\mu} z e^2} pprox 19.6 \; {
m fm} \; ({
m Al})$$
 • Nuclear Size

$$R \approx 1.2 A^{1/3} \text{fm} = 3.6 \text{ fm (Al)}$$

- Provides large overlap in the muon's wavefunction with the nucleous's
- For Z > 25 the muon is "inside" the nucleous
- Once captured 3 things can happen
  - Decay in Orbit:
  - Nuclear Capture:

$$\mu^- N(Z) \to \nu N(Z-1)$$

#### **Ordinary Muon Capture (OMC)**

Problem

These protons and neutrons constitute a large source of rate in the detector ( $\approx 1.2 \text{ per } \mu$ )

The energy spectra for these ejected particles is not well known.

Capture is a contact like

iection

interaction, scales as:

 $|\phi_{\mu}(0)|^2 N_{protons} \sim Z^4$ 

Lifetime: 864ns DIO Fraction: 39.3%

**Capture Fraction: 60.7%** 

#### **Muonic Atom**

- Start with a series of target foils
  - We stop ≈ 50% of µ's
- Bring in the low energy muon beam
  - We stop ≈ 50% of μ's
  - Stopped muons fall into the atomic potential
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  - Muonic Bohr Radius

• Muonic Bohr Radius 
$$\langle r_\mu \rangle = \frac{n^2 \hbar}{m_\mu z e^2} \approx 19.6 \; {\rm fm \; (Al)}$$
 • Nuclear Size

$$R \approx 1.2 A^{1/3} \text{fm} = 3.6 \text{ fm (Al)}$$

- Provides large overlap in the muon's wavefunction with the nucleus's
- For Z > 25 the muon is "inside" the nucleus
- •Once captured 3 things can happen
  - Decay in Orbit
  - Nuclear Capture
  - New Physics! i.e.  $\mu$  N  $\rightarrow$  e N

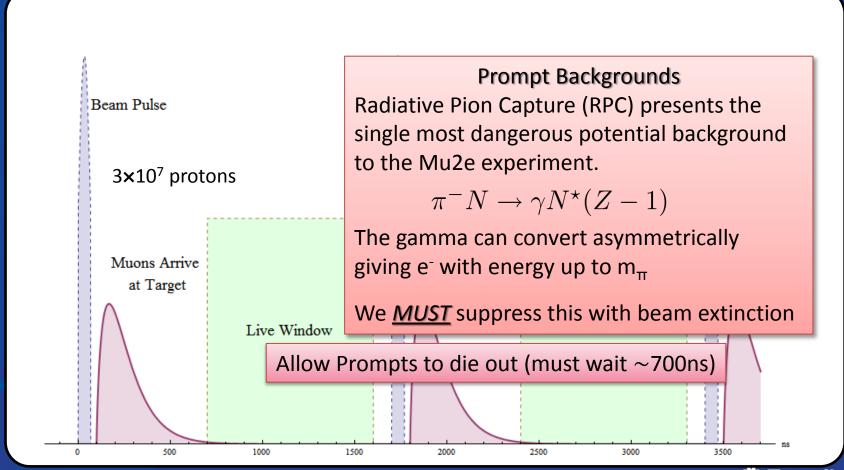
**Nucleus Is Left Unchanged** Muon Coherent Conversion to the ground state scales as  $\sim Z^5$ . Rates:  $(\mu N \rightarrow eN)/(OMC)$ rises as Z. Moving to high Z buys you sensitivity

Coherent Conversion ( $\mu \rightarrow e$ )

E<sub>a</sub> ≈ 104.96 MeV

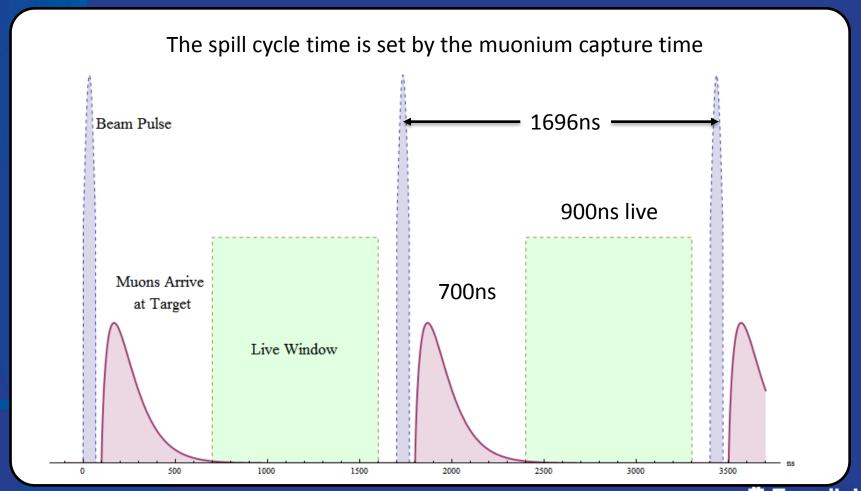
#### Beam Structure

- $\mu$ 's are accompanied by prompt e,  $\pi$ 's, ....
- These cause dangerous backgrounds (RPC)
- Must limit our beam extinction, and detector live window



#### Beam Structure

- μ's are accompanied by prompt e, π's, ....
- These cause dangerous backgrounds (RPC)
- Must limit our beam extinction, and detector live window



# **Total Backgrounds**

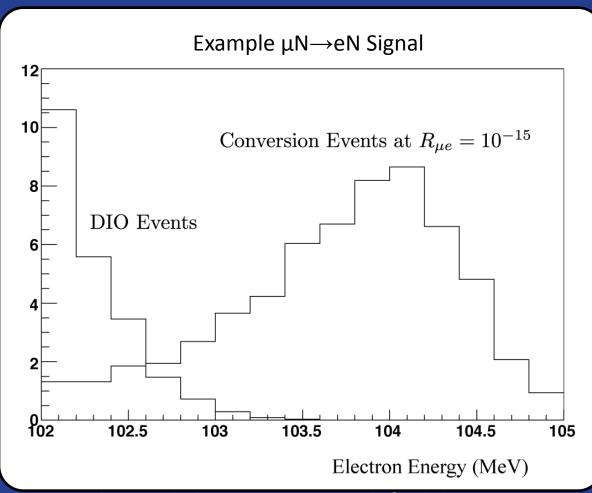
 Total expected background for SES 10<sup>-17</sup> ≈ 0.41 evts

- Largest Background
  - Decay in Orbit (DIO)

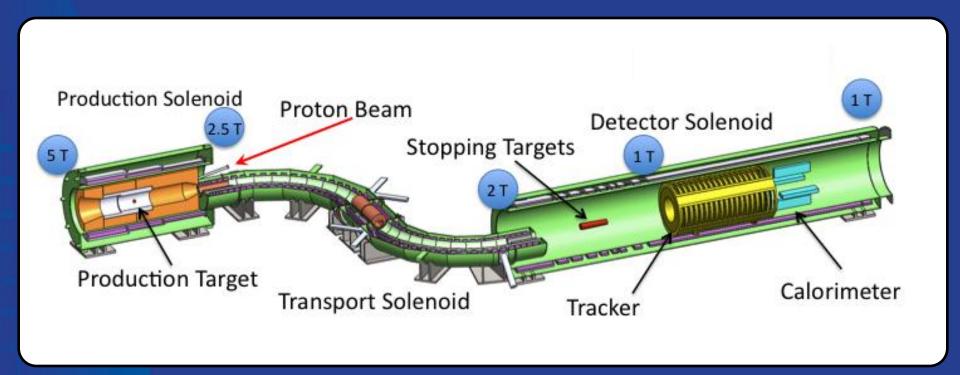
Background	Bkg Est.	Err Est.	Notes
Muon Decay-in-orbit	0.22	$\pm 0.06$	Acceptance and energy
			loss modeling, spectrum
			calculation; reco algorithm
$\bar{p}$ Induced	0.10	$\pm 0.05$	Cross-section, modeling
Cosmic Ray	0.05	$\pm 0.013$	Monte Carlo Stats.
Rad Pion Capture	0.003	$\pm 0.007$	Acceptance and energy
			loss modeling
$\mu$ decay in flight	0.01	$\pm 0.003$	
$\pi$ decay in flight	0.003	$\pm~0.0015$	
Beam electrons	0.0006	$\pm~0.0003$	
Total	0.41	$\pm 0.08$	
<b>\</b>	·	·	· · · · · · · · · · · · · · · · · · ·

# Signal Estimates

- For  $R_{\mu e} = 10^{-15}$ 40 events / 0.41 bkg (LHC SUSY)
- For R<sub>μe</sub> = 10<sup>-16</sup>
   4 events / 0.41 bkg



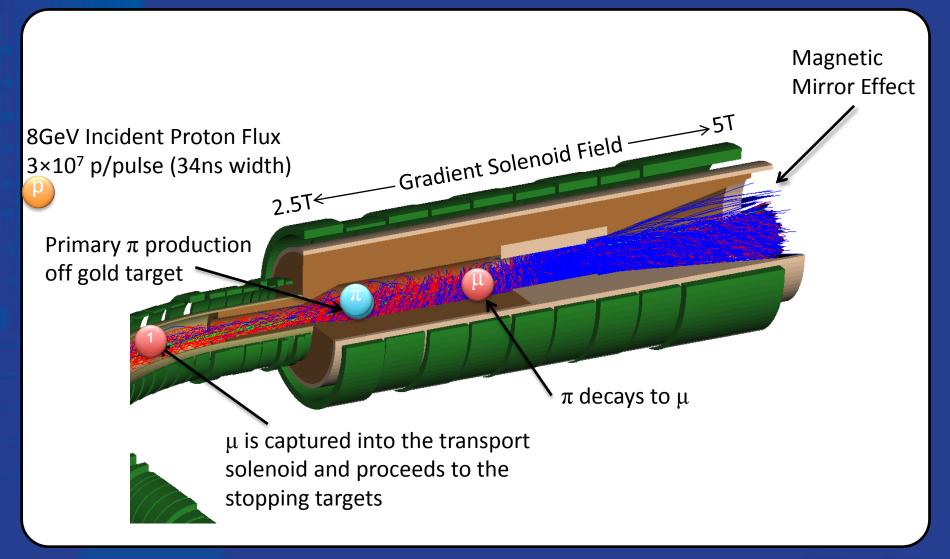
Observed electron energy is shifted down to 104 MeV due to energy loss in stopping target and smeared by detector resolution



Who ordered this?

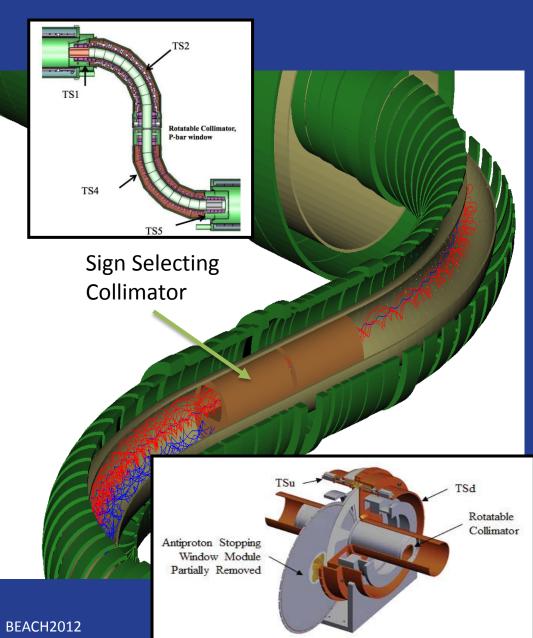
#### THE MU2E DETECTOR IN DETAIL

### **Production Solenoid**

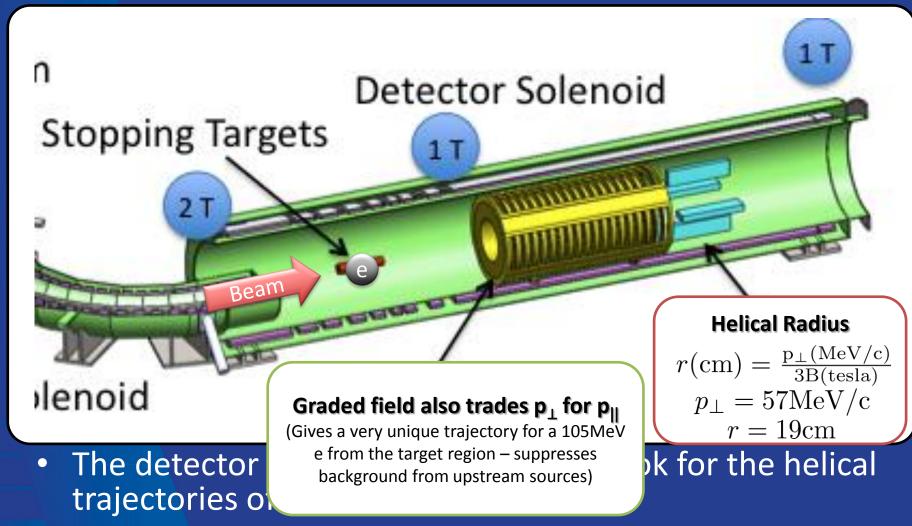


# **Transport Solenoid**

- Designed to minimize beam background rates from the production target
- Removes anti-protons from the beam line in a Be foil
- Sign selects the muon beam
  - Collimator blocks the positives after the first bend
  - Negatives are brought back on axis by the second bend
  - Allows for momentum selection of the beam



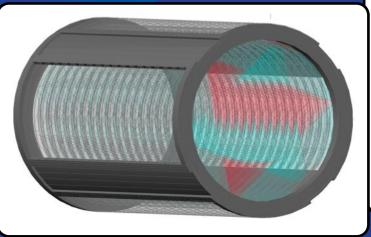
### The Detector

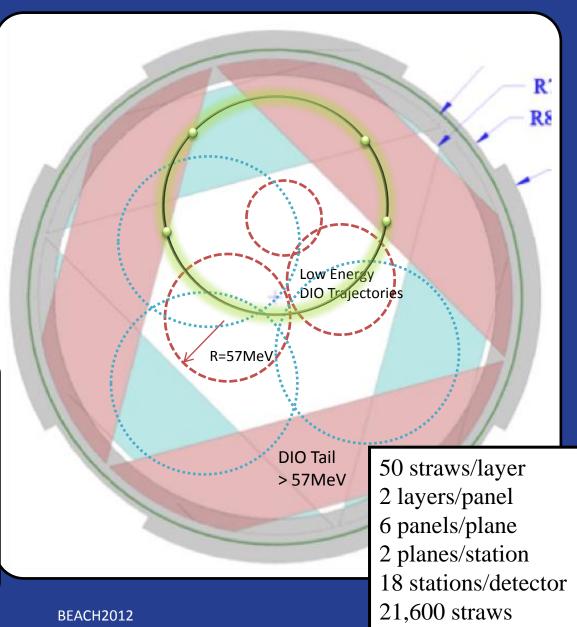


 Each component is optimized to resolve signal from the *Decay in Orbit* Backgrounds

# Straw Tracker (In Vacuum)

- Geometry is optimized for reconstruction of 105MeV helical trajectories
- Extremely low mass
- DIO tracks miss the senstive regions don't contribute to rate





#### Conclusions

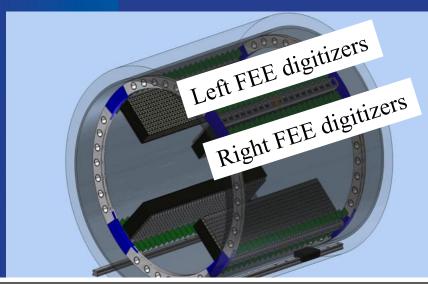
- Mu2e is unique in that it can push down the current limits on R<sub>ue</sub> by more than four orders of magnitude
- This gives the experiment real discovery potential of physics beyond the standard model
- Mu2e has the ability to complement LHC results or probe beyond the LHC to 10<sup>4</sup> TeV mass scales

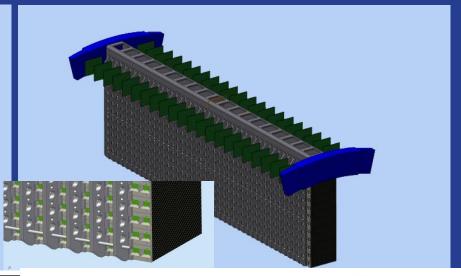
Mu2e received CD-1 on July 11th!

Project-X

### **ADDITIONAL MATERIAL**

## Calorimeter





Crystal	LYSO	PbWO <sub>4</sub>
Density (g/cm <sup>3</sup> )	7.28	8.28
Radiation length (cm) $X_0$	1.14	0.9
Molière radius (cm) R <sub>m</sub>	2.07	2.0
Interaction length (cm)	20.9	20.7
dE/dx (MeV/cm)	10.0	13.0
Refractive Index at $\lambda_{max}$	1.82	2.20
Peak luminescence (nm)	402	420
Decay time $\tau$ (ns)	40	30, 10
Light yield (compared to NaI(Tl)) (%)	85	0.3, 0.1
Light yield variation with temperature(%/°C)	-0.2	-2.5
Hygroscopicity	None	None

