

SuperB Background Picture



*E. Paoloni (INFN & Università di Pisa)
for the SuperB Collaboration*

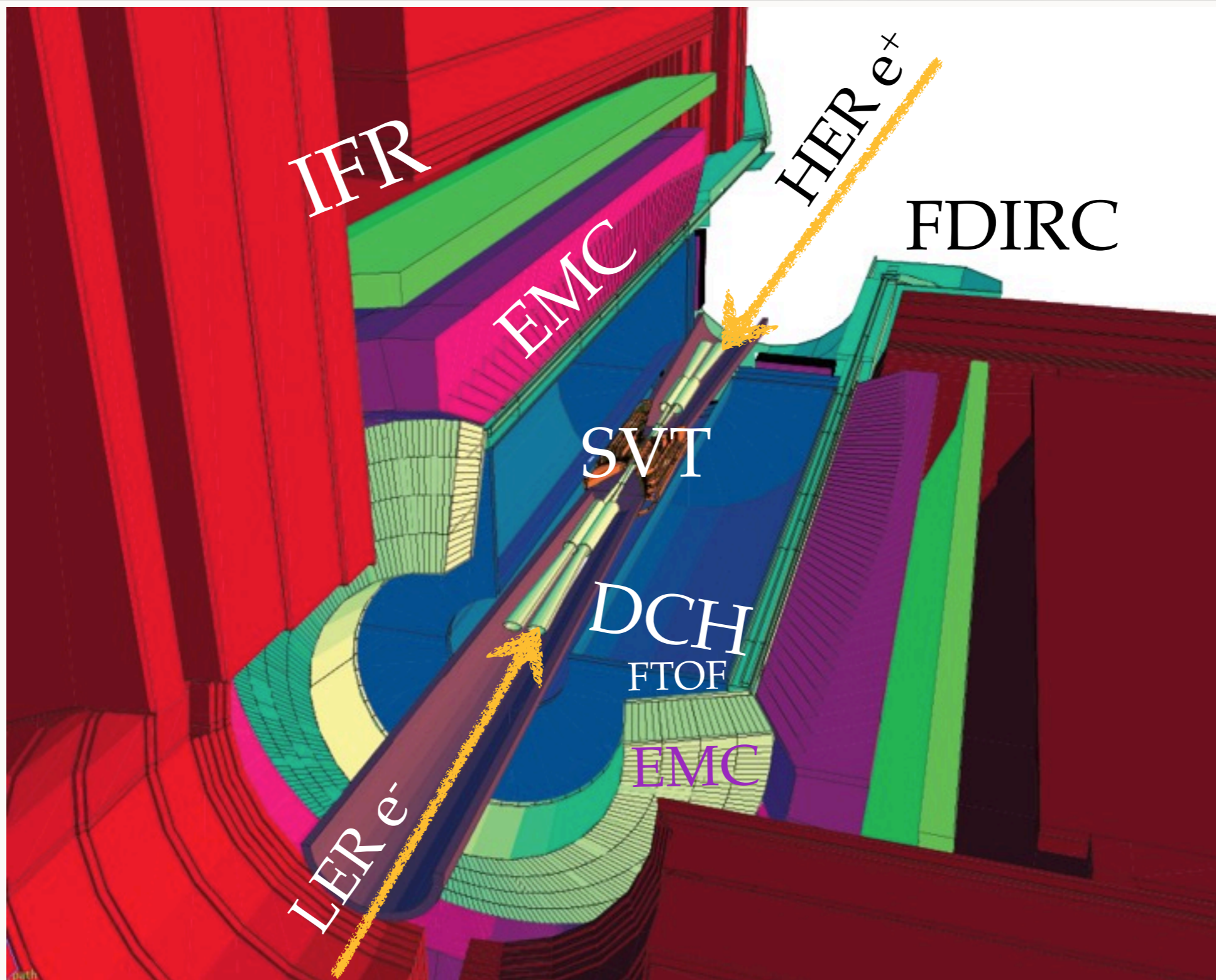
TALK OUTLINE

- The background model in SuperB:
 - Our Geant 4 simulation tool: Bruno
 - Luminosity scaling backgrounds
 - Radiative Bhabha (beam strahlung) B³Rem
 - Pairs production (2 photons) : Diag36 + Bruno
 - Intensity scaling backgrounds
 - Touschek scattering: Star + Bruno
 - Beam gas scattering: Star + Bruno
 - Synchrotron radiation

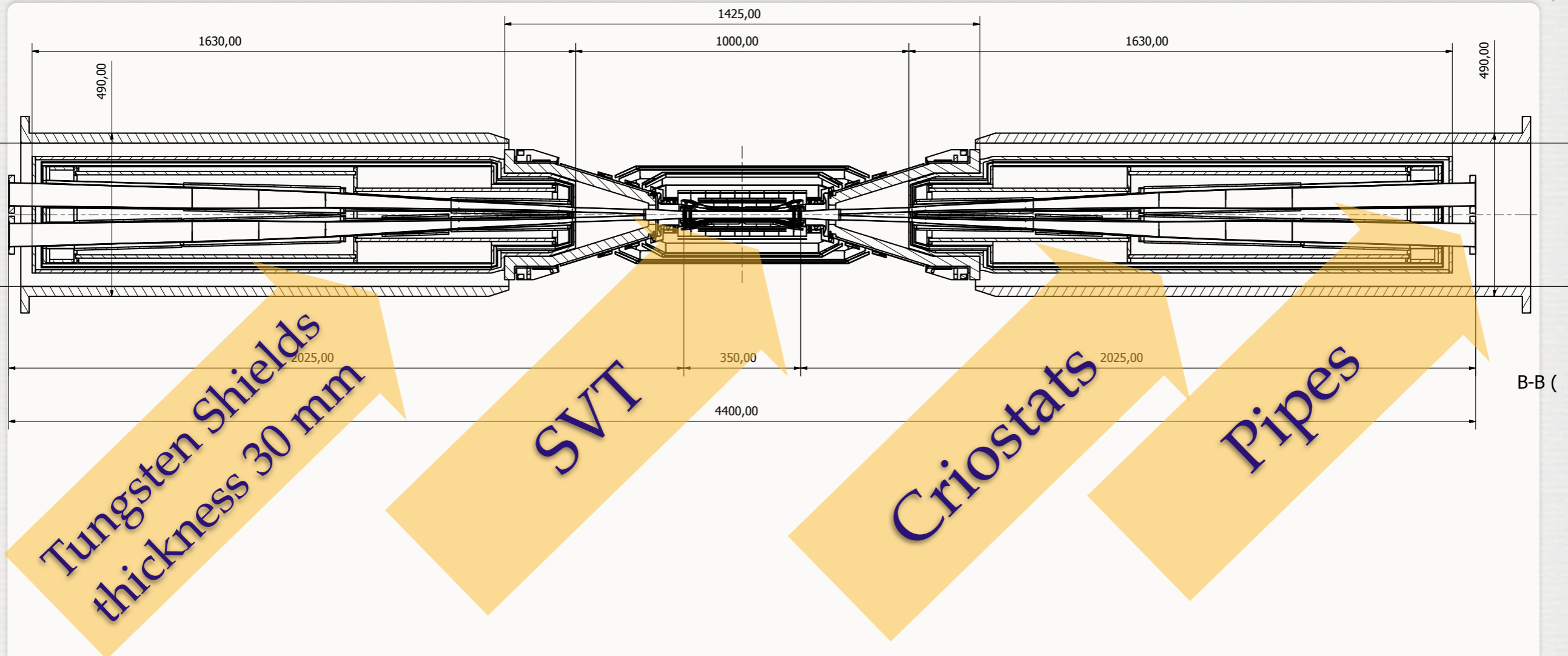
PRIMARIES PROPAGATION

- A Geant4 based program (code name Bruno, from Bruno Touschek) was created to simulate:
 - the transport of charged particles in the magnetic field of the final focus by numerical integration of the equation of motion
 - the magnetic field of the final focus is specified as a set of cylindrical regions in which the user prescribes the dipolar and quadrupolar components of the field
 - the passage of particles through matter (machine elements)
 - the effect of the secondaries in the detector (energy releases, doses)
 - at present Bruno is not able to *reconstruct* the event
 - each subsystem performs a post processing that “digitize” the energy releases: rates evaluation, impact on detector performances, physics reach...

BRUNO DETECTOR MODEL



BRUNO: MACHINE MATERIAL



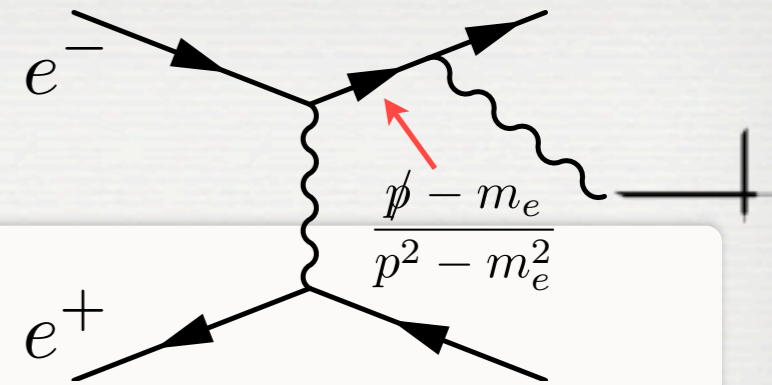
- Detailed model of the beam pipe sizes, cryostats and cold mass
- A 30 mm thick tungsten shield is put around the cryostat to protect the detectors from Radiative Bhabhas and Touschek

BACKGROUND CROSS SECTIONS

	Scattering Cross section	#Evt / crossing	Scattering Rate	
Beam Strahlung	~340 mbarn ($E_\gamma/E_{\text{beam}} > 1\%$)	~1400	0.34 THz	Luminosity lifetime driving term
Beam Strahlung	~150 mbarn ($E_\gamma/E_{\text{beam}} > 10\%$)	~630	0.15 THz	Losses “near” the IP
e^+e^- production	~7.3 mbarn	~31	7.3 GHz	
e^+e^- production (seen by L0 @ 1.4 cm coverage 300 mRad)	~ 80 μ barn	~0.34	80 MHz	Main SVT L0 Background
Elastic Bhabha	$O(10^{-4})$ mbarn (Det. acceptance)	~420/Million	100 KHz	~L1 Trigger rate
Υ (4S)	$O(10^{-6})$ mbarn	~4.2/Million	1 KHz	Physics

RADIATIVE BHABHA

$$e^+ e^- \rightarrow e^+ e^- \gamma \quad (\gamma \sim \parallel e^-)$$



- Quasi elastic Bhabha of the electron on the positron associated with the emission of a photon
- The virtual photon and the virtual electron are almost on mass shell:
 - the amplitude pinches both poles of the propagator
- The particles in the final states escapes throughs the detector acceptance holes till
 - magnetic field deflect the lepton that radiated the photon
 - the photon or the deflected lepton hit the beam pipe
 - the debris of the electromagnetic shower hit the detector

INCOMING LEPTON GENERATION

- The angular divergence of the beam @ IP $\sim 1/\gamma$ and cannot be neglected
- The momenta of the incoming particle and the position of the scattering vertex is generated from first principles
- All the gaussian features @ IP are modeled + Crab Waist as a bonus

$$\rho_{\text{ph.sp.}} \propto e^{-\frac{1}{2} \left[\left(\frac{x}{\sigma_x} \right)^2 + \left(\frac{x'}{\sigma_{x'}} \right)^2 + \left(\frac{y-y' - x \chi/\vartheta}{\sigma_y} \right)^2 + \left(\frac{y'}{\sigma_{y'}} \right)^2 + \left(\frac{s}{\sigma_z} \right)^2 + \left(\frac{\epsilon}{\sigma_\epsilon} \right)^2 \right]}$$

Radial size @ IP

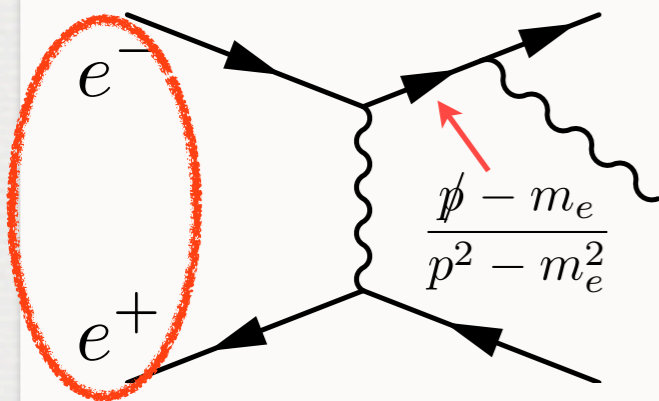
Radial div. @ IP

Vertical size @ IP
(with C.W.)

Vertical div. @ IP

Bunch length @ IP

Energy spread @ IP



PRIMARYIES GENERATION

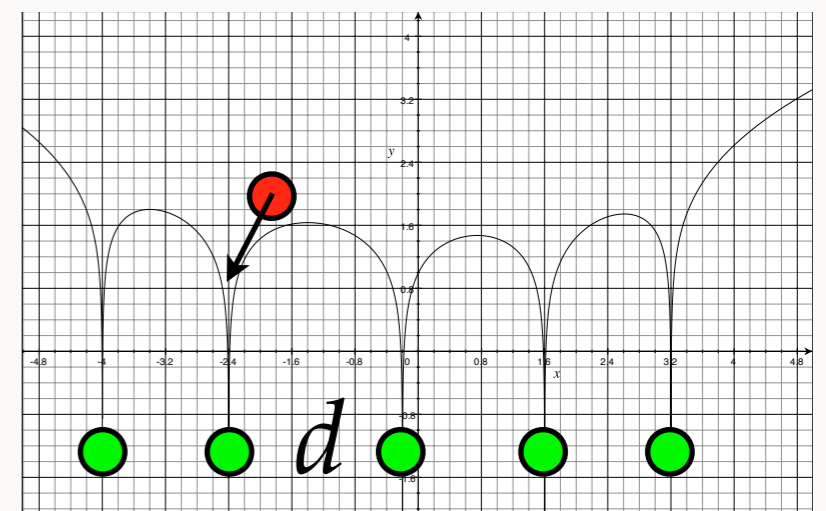
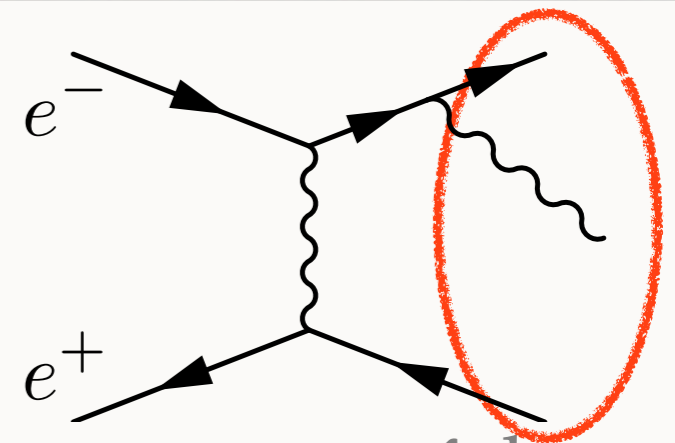
- BBRem (R. Kleiss, H. Burkhardt)
arXiv:hep-ph/9401333

- Only two Feynman diagrams are taken into account out of the eight tree level diagrams

- The electron mass is not “neglected with impunity” , i.e. the angular deflection of the lepton is properly simulated

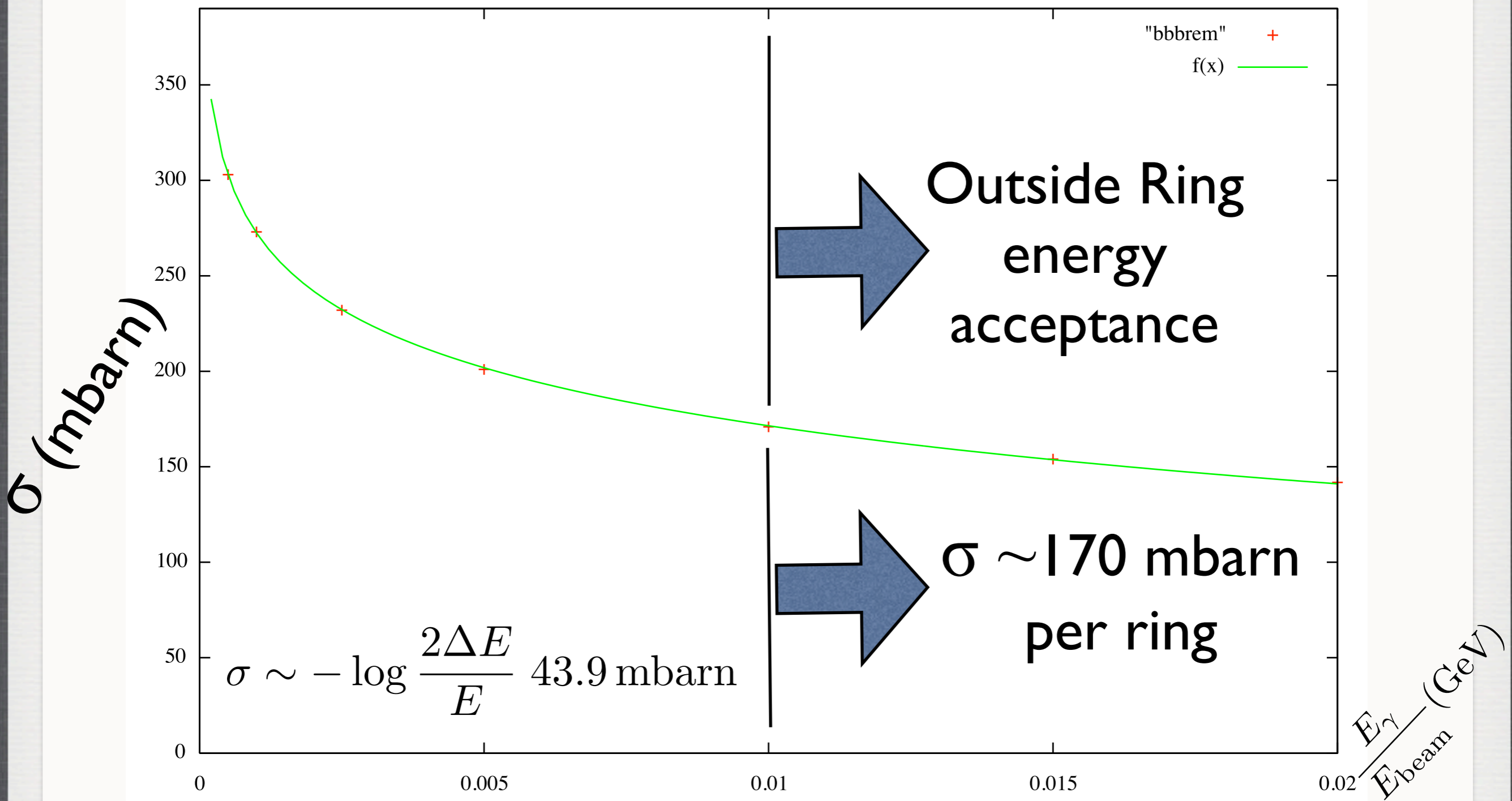
- Finite density of the bunch is taken into account by an infrared cut-off parameter that fix the minimum momentum transfer

$$p_{min} \sim \hbar c / d$$

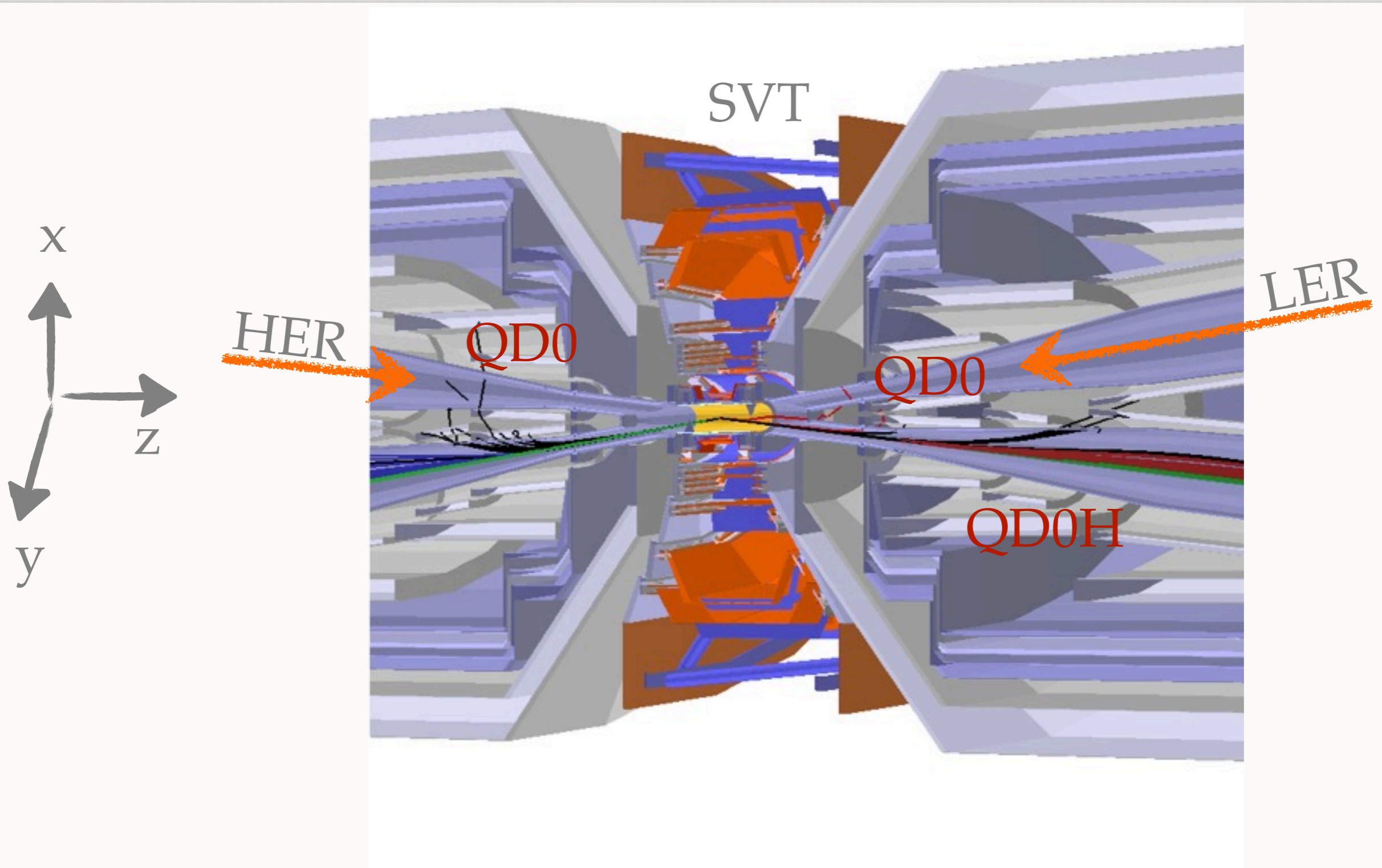


RADIATIVE BHABHA CROSS SECTION

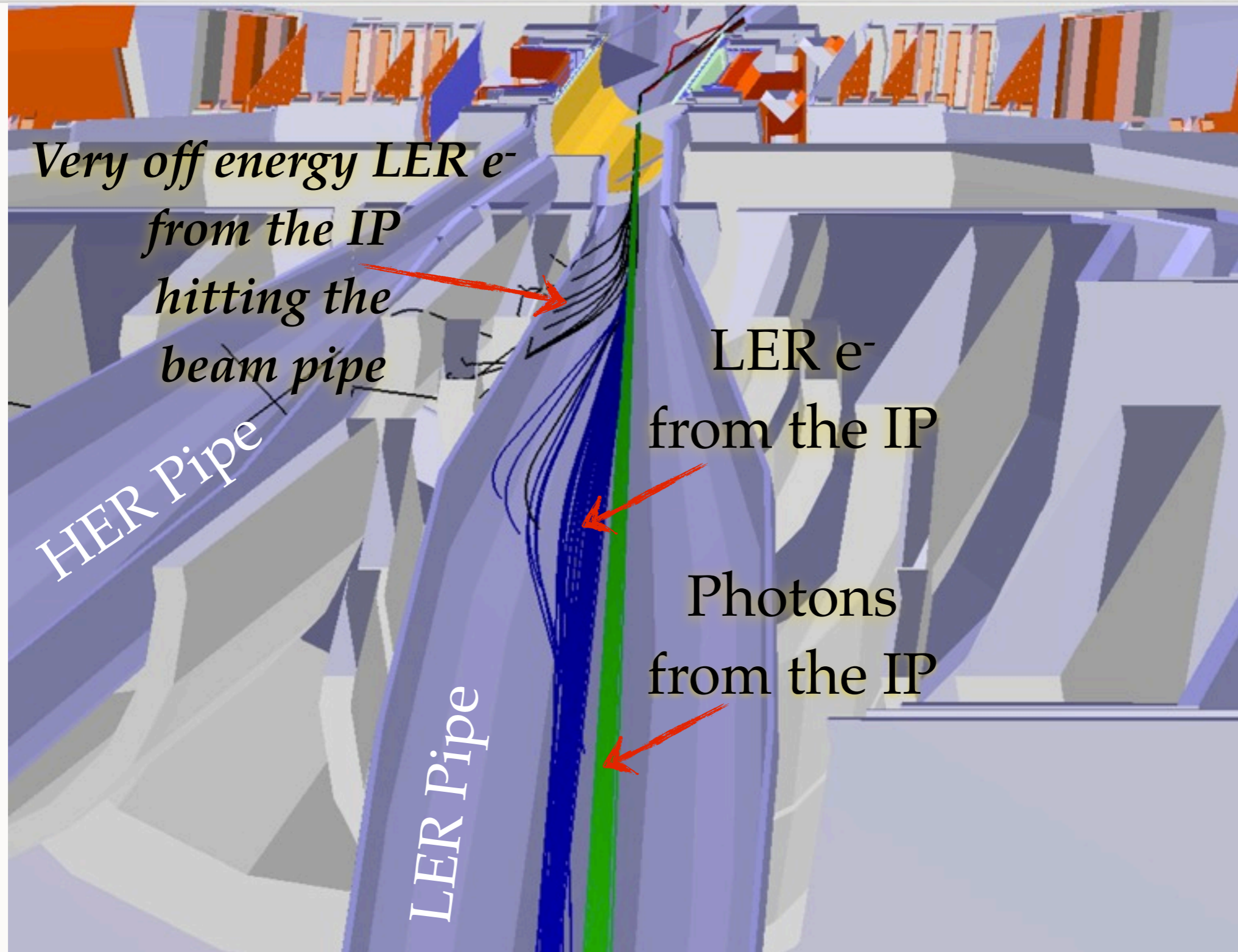
Rad. Bhabha Cross Section (mbarn) vs. Delta E / E



A SINGLE BUNCH CROSSING

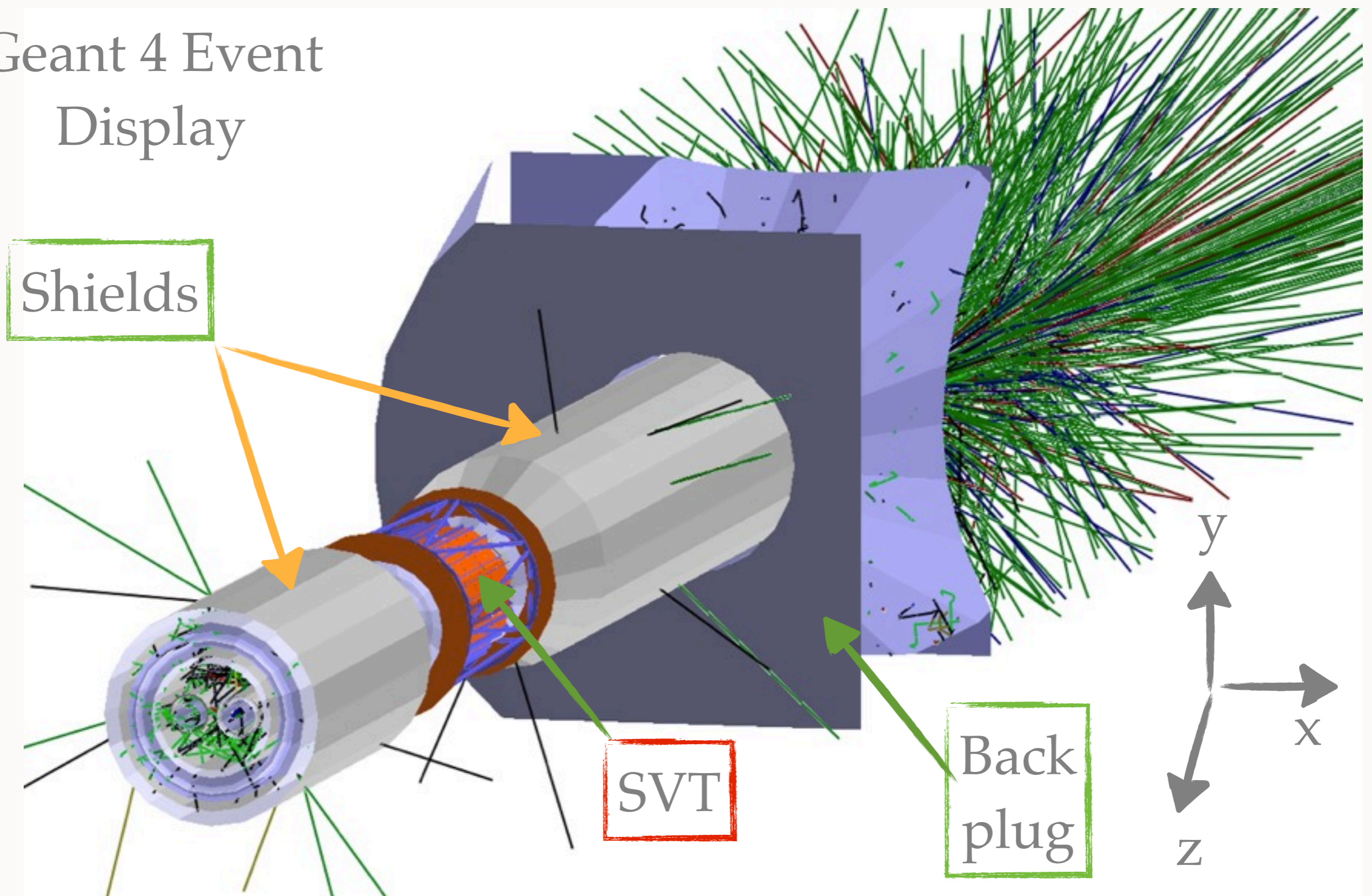


RADIATIVE BHABHA: 1 BUNCH XING



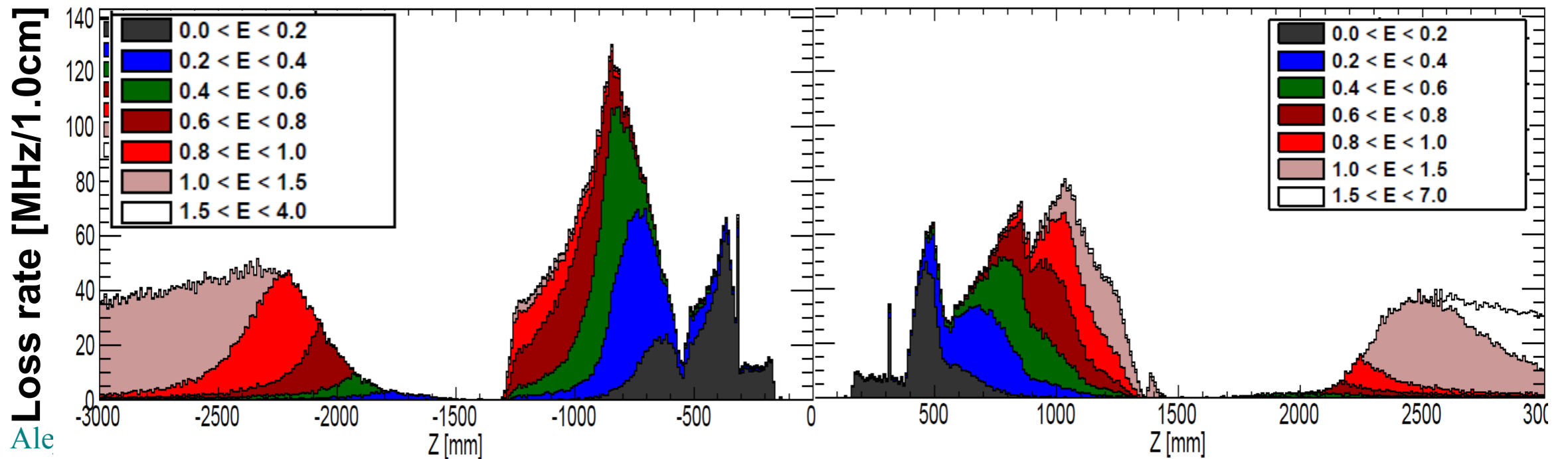
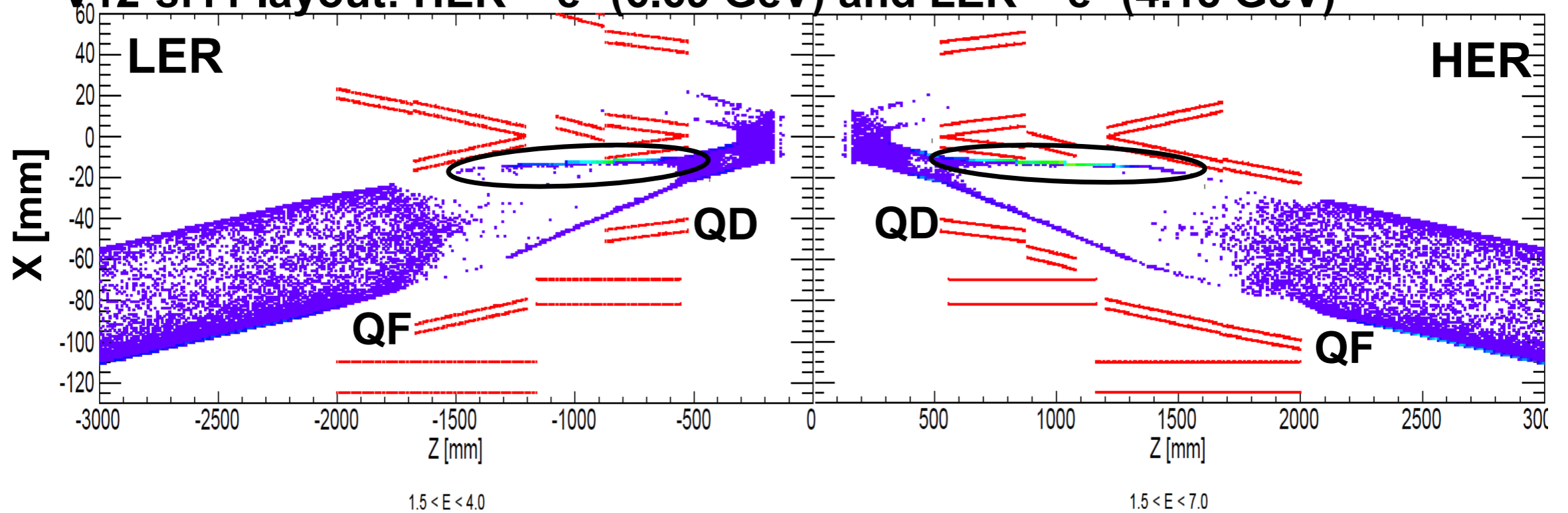
1 BUNCH CROSSING: SECONDARIES

Geant 4 Event
Display



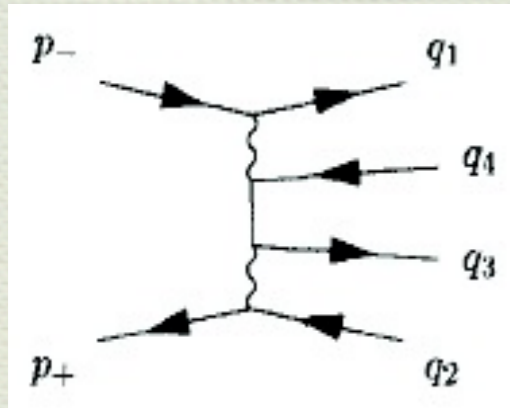
PRIMARYIES LOSS RATE

V12-sf11 layout: HER = e^+ (6.69 GeV) and LER = e^- (4.18 GeV)



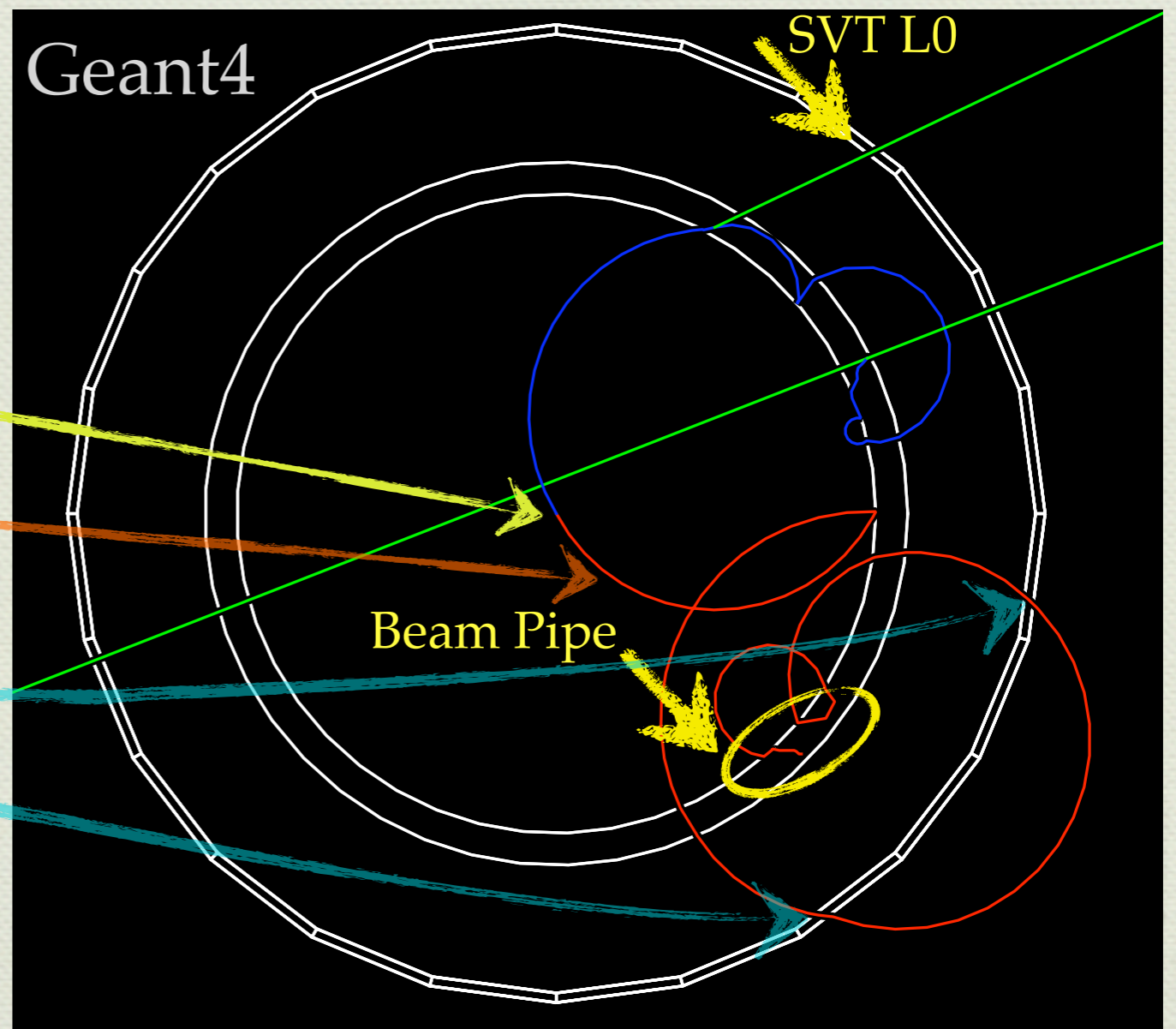
L0 main bkg.: pairs production

$$\sigma \sim \frac{\alpha^2 r_e^2}{\pi} \left(\frac{28}{27} \ln^3 \frac{s}{m^2} - 6.59 \ln^2 \frac{s}{m^2} - 11.8 \ln \frac{s}{m^2} + 104 \right)$$



$$e^+ e^- \rightarrow e^+ e^- e^+ e^-$$

- ◆ In this scattering event there is:
- ◆ 1 track that hits the L0
- ◆ this track fires 2 clusters in the L0
- ◆ each cluster will be composed by one or more hits



Definitions

Event rate:

$$\mathcal{R} = \mathcal{L} \sigma$$

track rate:

$$\mathcal{R}_{\text{trk}} = \mathcal{R} \langle \#trk \rangle_{\text{evt}}$$

cluster rate:

$$\mathcal{R}_{\text{clus}} = \mathcal{R}_{\text{trk}} \langle \#clus \rangle_{\text{trk}}$$

hit rate:

$$\mathcal{R}_{\text{hits}} = \mathcal{R}_{\text{clus}} \langle \#hits \rangle_{\text{clus}}$$

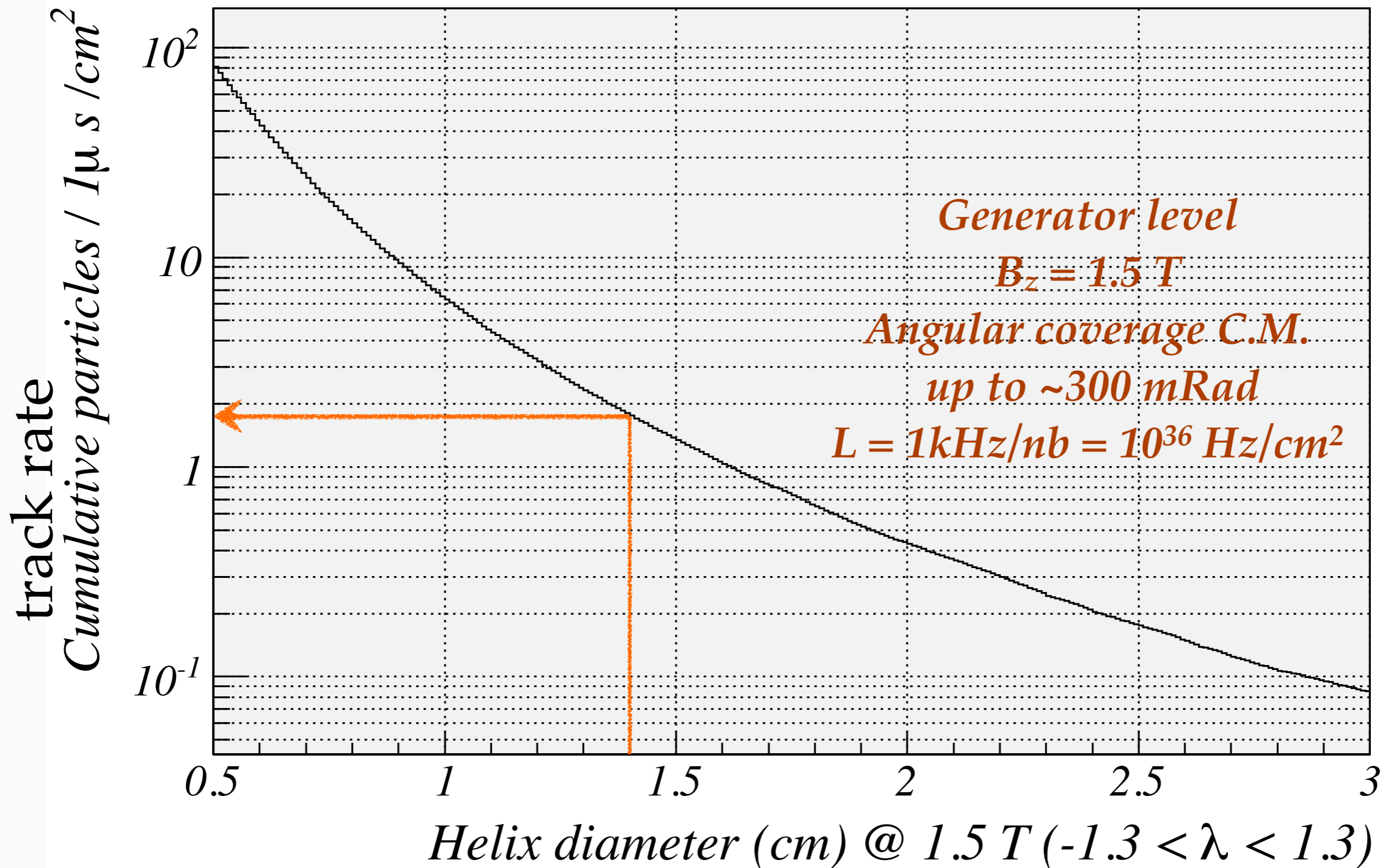
SIMULATION STRATEGY

- Primaries generated with DIAG36 then
- Bruno simulates the effects of primaries on detector

DIAG36

```
C-----ALL THE FEYNMAN DIAGRAMS CONTRIBUTING IN LOWEST ORDER-----
C-----(IPROC=1 6 DIAGRAMS, IPROC=2 12 DIAGRAMS, IPROC=3 12DIAGRAMS,--
C----- IPROC=4 12 DIAGRAMS, IPROC=5 36 DIAGRAMS)-----
C-----ARE TAKEN INTO ACCOUNT-----
C-----THE KINEMATICS IS TREATED EXACTLY-----
C-----THE PROGRAMS GENERATES EVENTS EFFICIENTLY UNDER-----
C-----NO- OR SMALL ANGLE TAGGING CONDITIONS-----
C
C*****
C*
C*   AUTHORS : F.A. BERENDS, P.H. DAVERVELDT, R. KLEISS   *
C*
C*   UNIVERSITY OF LEIDEN   *
C*   INSTITUUT-LORENTZ VOOR THEORETISCHE NATUURKUNDE   *
C*   NIEUWSTEEG 18   *
C*   2311 SB LEIDEN   *
C*   THE NETHERLANDS   *
C*
C*
C*   INSTALLATION DATE :   *
C*   LAST UPDATE       : 8 FEBRUAR 1985   *
C*
C*****
C
C FOR DETAILED INFORMATION ON THE CALCULATION OF THE MATRIX ELEMENT
C SQUARED AND ON THE PROCEDURE USED FOR THE EVENT GENERATION WE REFER
C TO THE PAPER :
C "COMPLETE LOWEST ORDER CALCULATIONS FOR FOUR-LEPTON PROCESSES IN E+
C E- COLLISIONS"
C NUCL. PHYS. B253 (1985) 441
C
C THIS PROGRAM RUNS IN DOUBLE PRECISION FORTRAN H-EXT.
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DIAG36 TRACK RATE EVALUATION



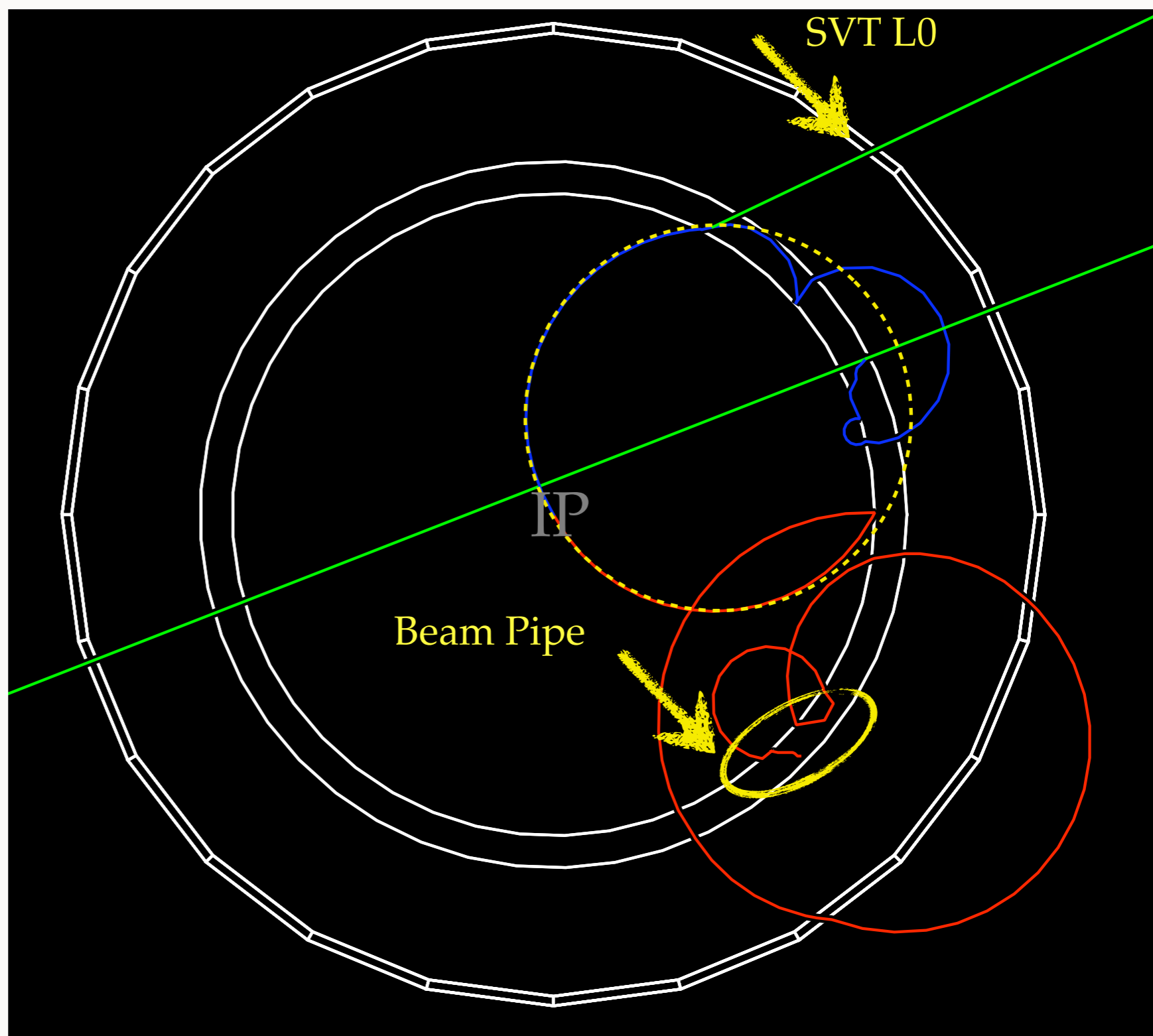
WHY WE NEED BRUNO?

Primaries are very soft

~ few MeV

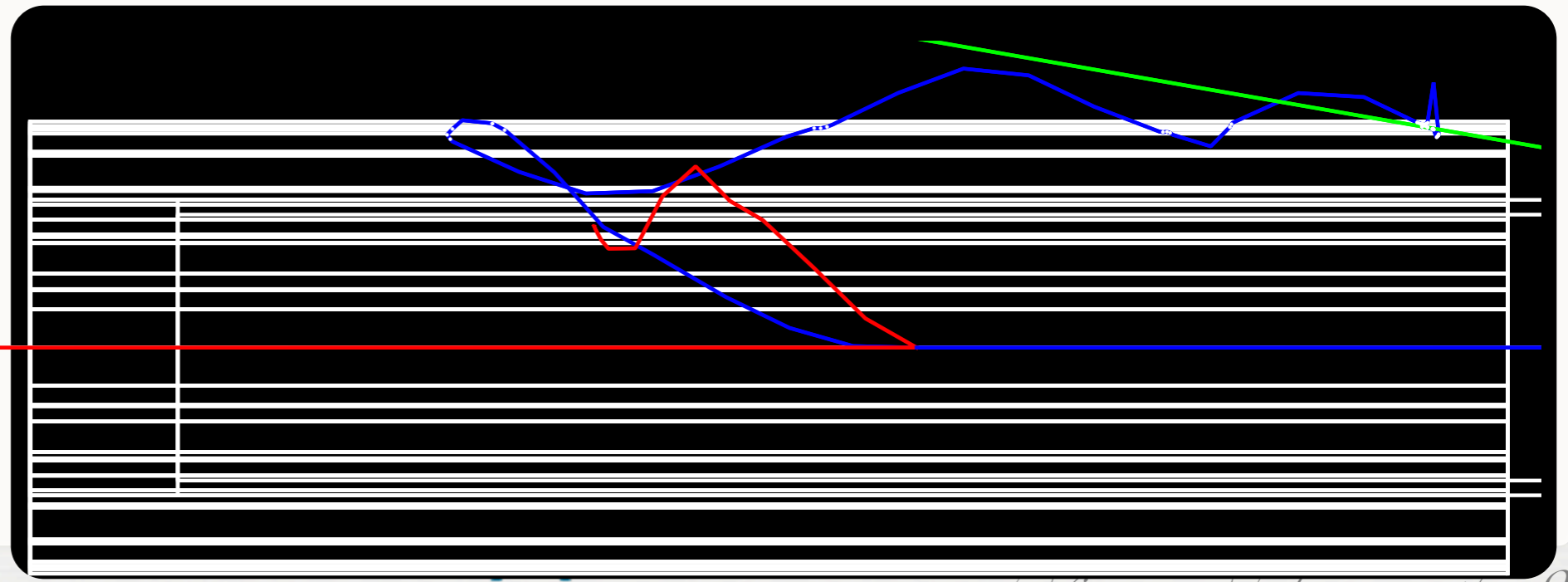
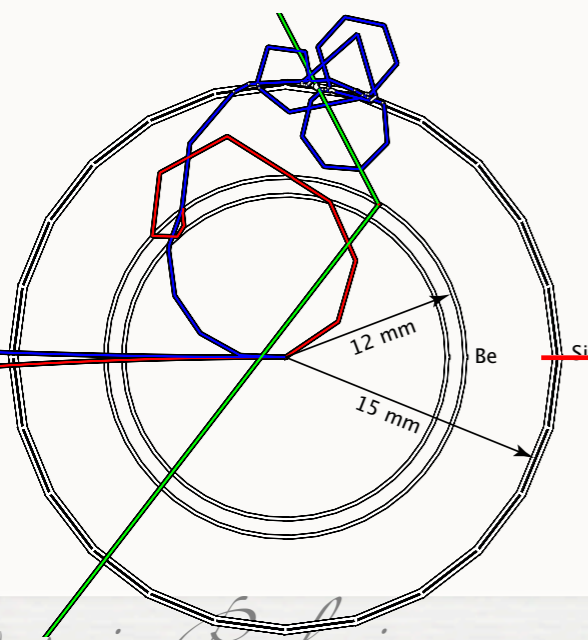
Multiple Coulomb scattering and dE/dx are not negligible

Particles with $p_t < 3.5 \text{ MeV}/c$ can still hit the pipe and the Coulomb scattering can increase p_t at expense of the long. momentum

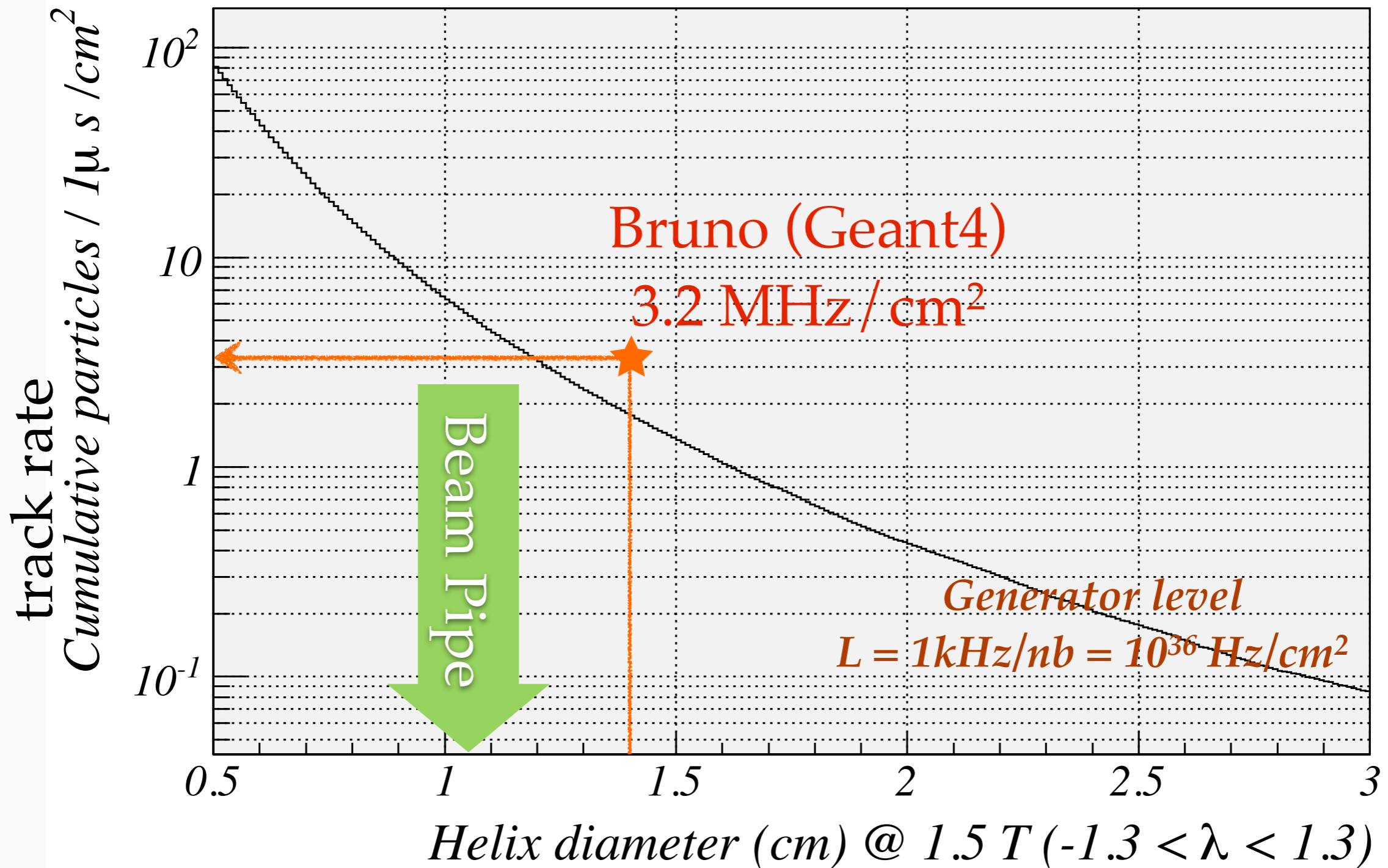


PAIR PRODUCTION

- Particles from this process can have enough transverse momentum to enter directly into the detector acceptance
 - The detector solenoidal field is the main trap for particles with small p_t
 - Non trivial effects from the interactions with the beam pipe



BRUNO EVALUATION EXAMPLE



CLUSTER MULTIPLICITY

- In our model 1mm thick Be Beam Pipe + 3 μm Au @ $r = 10$ mm
- SVT L0 mean radius 14mm
 - track rate = 3.2 MHz / cm^2
 - Average cluster multiplicity per track 2.5
 - Cluster Rate 8.2 MHz / cm^2
 - Average hit multiplicity per cluster 4 - 5 depending on the view

INTENSITY DEPENDENT BKG.

- Touschek scattering and beam-gas scattering generates non gaussian tails in the bunch transverse profile
- The finite aperture of the beam pipe and the large β functions at the final focus doublet conspire to make these particles in the tails impinge on the beam pipe
- The non gaussian tails generation and transport through the lattice is simulated with STAR (M.Boscolo)
- Bruno simulates the interaction of these particles crossing the beam pipe near the IP to predict the effects on the detector

PRIMARY LOSS RATES

Modif. V12 Conclusions (1): Lifetime summary

	HER	LER
Touschek lifetime	τ_{TOU} (min)	τ_{TOU} (min)
No collimators, nominal ϵ_x (no IBS)	26.3	7.4
No collimators, ϵ_x with IBS	26	10.2
With Collimators, ϵ_x with IBS	22	7.9
Coulomb	76min	39 min
Bremsstrahlung	72 hrs	77 hrs



M. Boscolo, December 14th 2011

Eugenio Laoloni



Vienna, Feb. 2012 the 9th

LOSS RATE AT THE IP

Modif. V12 Conclusions (2): IR rates summary

$|s| < 2$ m

Touschek	HER	LER
No collimators, ϵ_x with IBS	2.5 GHz	17 GHz
With Collimators, ϵ_x with IBS	7 MHz	100 MHz

Coulomb No collimators, ϵ_x with IBS	11 GHz	25 GHz
Coulomb with collimators, ϵ_x with IBS	11 MHz	36 MHz
Bremsstrahlung with coll	130 KHz	450 KHz



M. Boscolo, December 14th 2011

TOUSCHEK VS RAD. BHABHA

- LER Touschek IR loss rate 100 MHz vs Rad Bhabha 10 GHz
- But: the Touschek losses are fairly energetic ~ 4 GeV while the radiative Bhabha are quite soft
- The energy spectrum and the angular distributions of the secondaries are quite different.
- The total rate of the secondaries from Touschek is smaller
- The energy spectrum of the secondaries

Rad-Bhabha Losses at the Beam-pipe

Total rates around the IP (-3 to 3 mts)

HER positron rates		LER electron rates	
E range (GeV)	Rate (GHz)	E range (GeV)	Rate (GHz)
0.0 – 1.0	4.735	0.0 – 1.0	7.863
1.0 – 2.0	2.789	1.0 – 1.5	2.289
2.0 – 3.0	0.025	1.5 – 2.0	0.031
3.0 – 4.0	0.003	2.0 – 2.5	0.007
4.0 – 5.0	0.003	2.5 – 3.0	0.004
5.0 – 6.0	0.003	3.0 – 3.5	0.005
6.0 – 7.0	0.005	3.5 – 4.2	0.003
0.0 – 7.0	7.563	0.0 – 4.2	10.202

Touschek	HER	LER
No collimators, ϵ_x with IBS	2.5 GHz	17 GHz
With Collimators, ϵ_x with IBS	7 MHz	100 MHz

Coulomb No collimators, ϵ_x with IBS	11 GHz	25 GHz
Coulomb with collimators, ϵ_x with IBS	11MHz	36 MHz
Bremsstrahlung with coll	130KHz	450KHz

CONCLUSIONS

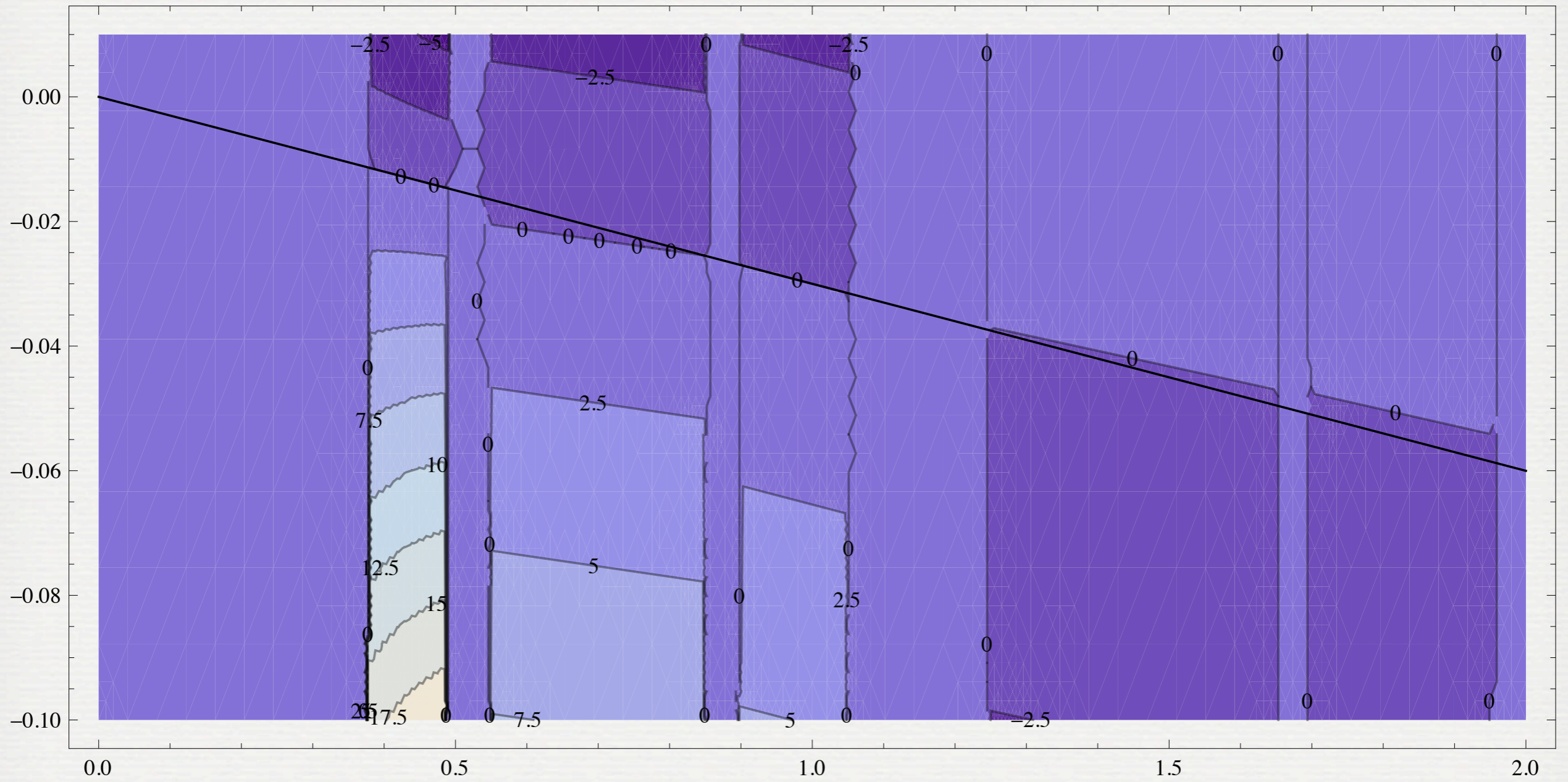
- SuperB developed (in my humble opinion) a fairly good set of tools to understand and predict the backgrounds features
- The discrepancy in the pairs background rate is probably a byproduct of lack of communication between the 2 collaborations
- The beam pipe material play a significant role in the pairs background

Thank you

For your Attention

HER B_y (T) (Mathematica)

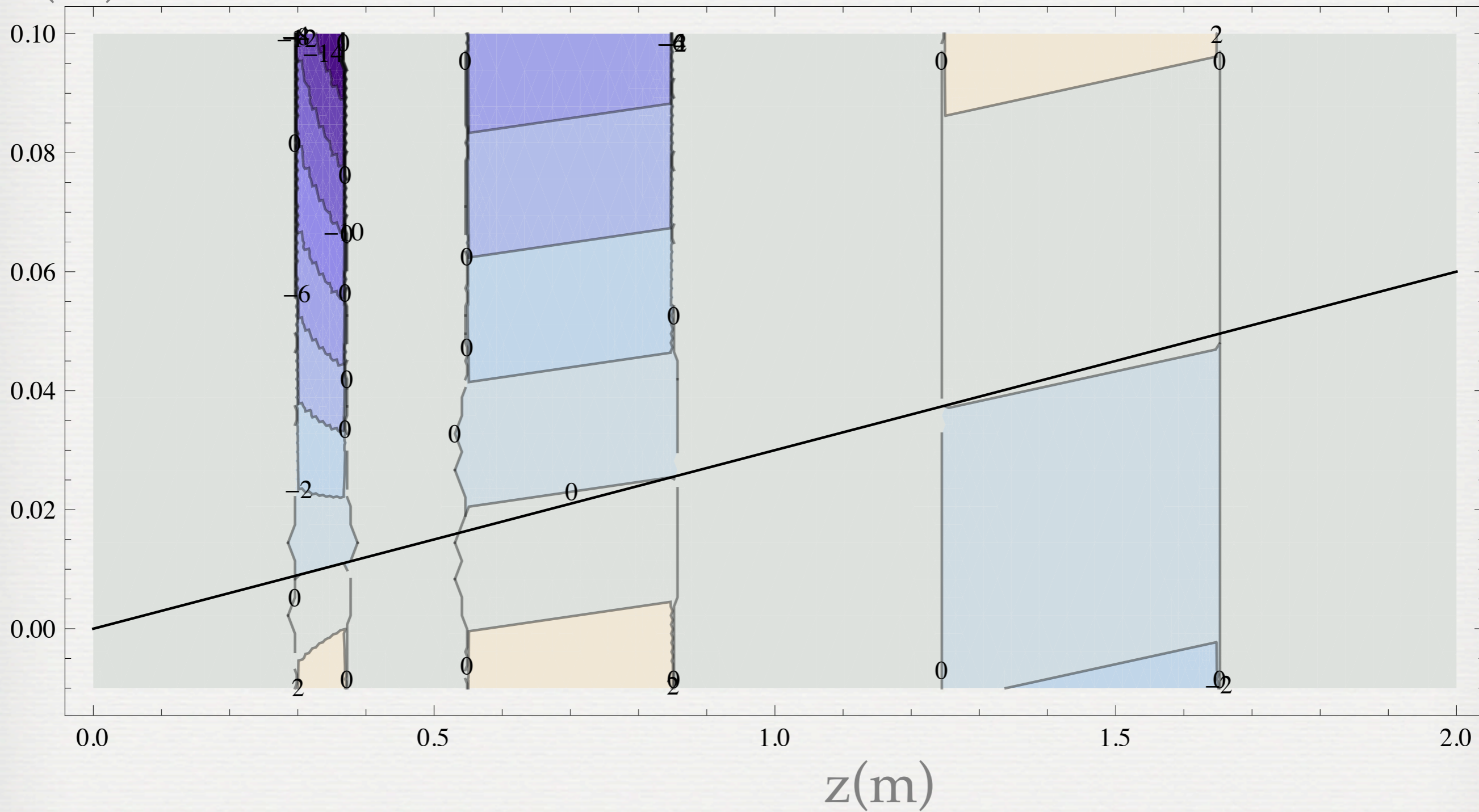
x(m)



z(m)

HER B_y (T) (Mathematica)

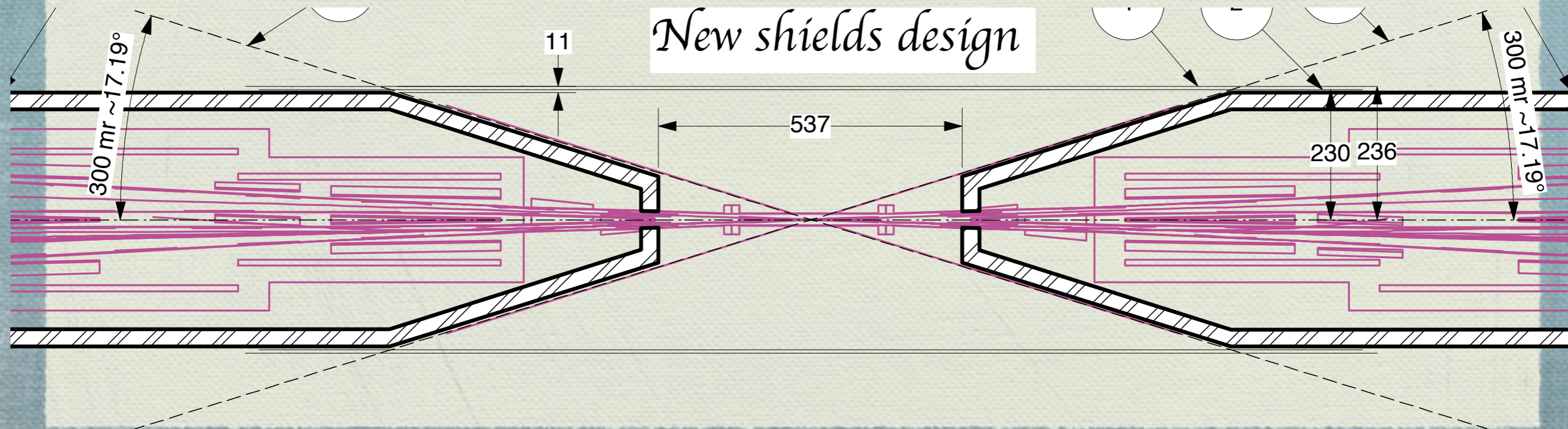
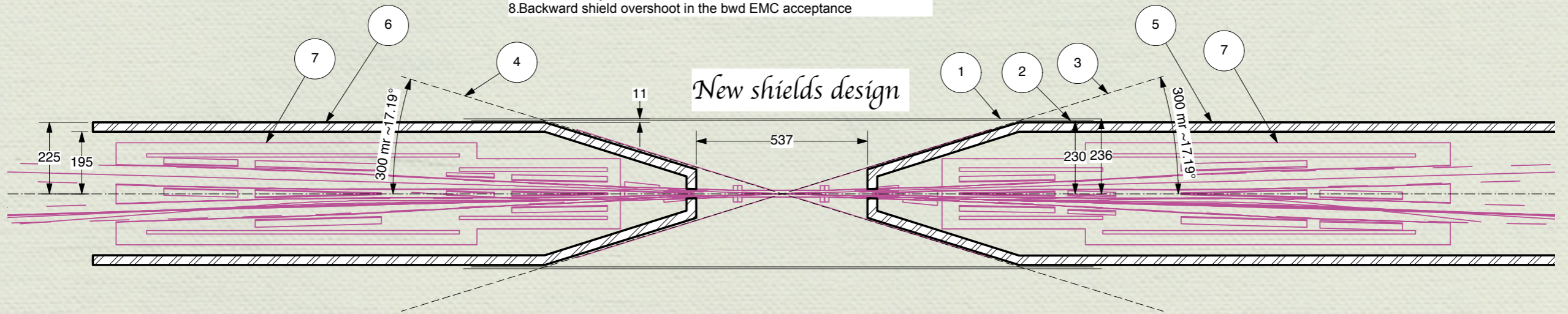
$x(m)$



Mechanical interface: boundaries

General Notes

- 1.DCH internal radius
- 2.Geant4 DCH boundary envelope
- 3.Forward acceptance limit
- 4.Backward acceptance limit
- 5.Forward shield
- 6.Backward shield
- 7.Final focus cryostat
- 8.Backward shield overshoot in the bwd EMC acceptance



GENERATOR LEVEL COMPARISON

