

Anomalous-Magnetic-Moment Effects in the Equation of State of a Magnetized Fermion System

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Introduction

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Dirac Theory

$$\boldsymbol{\mu} = g\mu_B\boldsymbol{s}$$

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But refined experiments give $g \simeq 2.00232$

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$$(\Pi_\mu \gamma^\mu - m + \kappa \mu_B B \Sigma_3) \psi = 0 \quad (1)$$

where κ is the AMM of the fermion, $\Sigma_3 = i\gamma_1\gamma_2$ and $\Pi_\mu = i\partial_\mu - eA_\mu$.

[‡]J. Schwinger, *Phys. Rev.* **73**, 416 (1947).

[§]B. Jancovici, *Phys. Rev.* **187**, 2275 (1969).

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The corresponding relativistic particle energy is

$$E_{l,\sigma}^2 = [(m^2 + 2eBl)^{1/2} - \sigma \kappa \mu_B B]^2 + p_3^2, \quad l = 0, 1, 2, \dots, \quad \sigma = \pm 1 \quad (2)$$

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$$\epsilon = m - \mu_B \kappa B$$

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$$\epsilon = m - \mu_B \kappa B \quad \xrightarrow{\text{if}} \quad B = \frac{m}{\mu_B \kappa} \quad \Rightarrow \quad \epsilon = 0$$

The last conclusion could lead to several problems[§], so we can not extrapolate Schwinger results to high-field values

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Introduction

The AMM term shown before is the first term in a series expansion in B^\ddagger :

$$M^{(0)} = \frac{\alpha}{2\pi} m \left[-\frac{eB}{2m^2} + \left(\frac{eB}{m^2} \right)^2 \left(\frac{4}{3} \ln \frac{m^2}{eB} - \frac{13}{18} \right) + \dots \right] \quad (3)$$

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Can we use the linear approximation for high values of the magnetic field?

fermion	p	e	u	d	s
mass(MeV)	938	0.5	5	5	150
AMM	$2.79 \mu_N$	$0.00116 \mu_B$	$1.85 \mu_N$	$-0.79 \mu_N$	$-0.85 \mu_N$

$$\mu_N = \frac{e}{2m_p}, \quad B_c = m^2/e$$

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Introduction

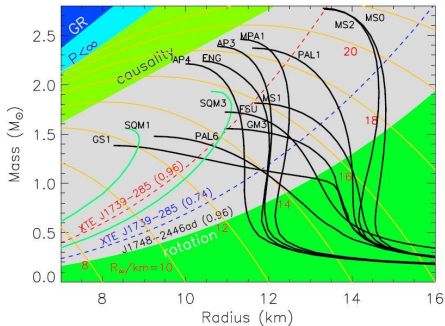
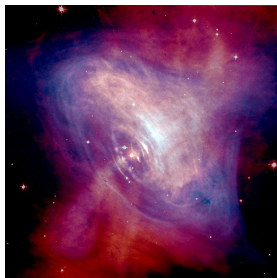
Motivation

It is interesting to see the effects of the AMM in the thermodynamical parameters of a magnetized fermion gas...

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...and in the physics of Compact Objects.

Goals

- 1 Calculation of the fermion self-energy in a constant magnetic field (B), in a non-perturbative approach (exact in B).
- 2 Calculation of the AMM in this approximation.
- 3 Study the effects of the obtained AMM in the thermodynamical quantities of the magnetized fermion gas.

Fermion Self-energy $B \neq 0$.

- Consider a constant magnetic field in x_3 -direction.
- Use the Ritus[†] transformation to diagonalize the self energy in momentum space.
- We will calculate the self-energy in the infrared limit ($p \rightarrow 0$) for $T = 0$.

The AMM enters as a correction of the energy of particles in the calculations of the thermodynamical quantities.

[†]V. I. Ritus, *Ann.Phys.* 69 (1972) 555; *Sov. Phys. JETP* **48** (1978) 788 [*Zh. Eksp. Teor. Fiz.* **75** (1978) 1560].

Fermion Self-energy $B \neq 0$.

Schwinger-Dyson equation for the fermion self-energy in the presence of a constant magnetic field gives us the self-energy:

$$\Sigma(x, x') = -ie^2 \gamma^\mu G(x, x') \gamma^\nu D_{\mu\nu}(x - x'). \quad (4)$$

where the bare photon propagator is (Feynman gauge):

$$D_{\mu\nu}(x - x') = \int \frac{d^4q}{(2\pi)^4} \frac{e^{-iq \cdot (x - x')}}{q^2 - i\epsilon} g_{\mu\nu}, \quad (5)$$

and the full fermion propagator obeys the equation

$$[\Pi_\mu \gamma^\mu - \Sigma(x, x')] G(x, x') = \delta^4(x - x'). \quad (6)$$

We go to momentum space by means of Ritus transformation

$$G(x, x') = \int \frac{d^4p''}{(2\pi)^4} E_{p''}^{l''}(x) \Pi(l'') \tilde{G}^{l''}(\bar{p}'') \bar{E}_{p''}^{l''}(x') \quad (7)$$

Fermion Self-energy $B \neq 0$.

The $E_p^l(x)$ functions are used for the transformation to momentum space and are defined as:

$$E_p^l(x) = E_p^+(x)\Delta(+) + E_p^-(x)\Delta(-) \quad (8)$$

where

$$\Delta(\pm) = \frac{I \pm i\gamma^1\gamma^2}{2}, \quad (9)$$

are the spin up (+) and down (-) projectors, and $E_p^\pm(x)$ are the corresponding eigenfunctions

$$\begin{aligned} E_p^+(x) &= N_l e^{-i(p_0x^0 + p_2x^2 + p_3x^3)} D_l(\rho), \\ E_p^-(x) &= N_{l-1} e^{-i(p_0x^0 + p_2x^2 + p_3x^3)} D_{l-1}(\rho) \end{aligned} \quad (10)$$

with normalization constant $N_l = (4\pi eB)^{1/4} / \sqrt{l!}$, and $D_l(\rho)$ denoting the parabolic cylinder functions of argument $\rho = \sqrt{2eB}(x_1 - p_2/eB)$, and index given by the Landau level numbers $l = 0, 1, 2, \dots$

Fermion Self-energy $B \neq 0$.

so we obtain for the self-energy in momentum space:

$$\Sigma^l(\bar{p})\Pi(l) = -ie^2(2eB)\Pi(l) \int \frac{d^4\hat{q}}{(2\pi)^4} \frac{e^{-\hat{q}_\perp^2}}{\hat{q}^2} [L_l + L_{l+1} + L_{l-1}]. \quad (11)$$

where we have used the notation $\hat{x} = x/\sqrt{2eB}$, $\Pi(l) = \Delta(+)\delta^{l0} + I(1 - \delta^{l0})$ and the three L factors on the right hand side of the equation (11) have the explicit form

$$\begin{aligned} L_l &= \gamma_\mu^\parallel G^l(\overline{p-q})\gamma_\mu^\parallel, \\ L_{l\pm 1} &= \Delta(\pm)\gamma_\mu^\perp G^{l\pm 1}(\overline{p-q})\gamma_\mu^\perp \Delta(\pm), \end{aligned}$$

where

$$G^l(\bar{p}) = -\frac{\bar{p} \cdot \gamma + m}{\bar{p}^2 + m^2}, \quad \gamma_\mu^\parallel = \gamma_0, \gamma_3, \quad \gamma_\mu^\perp = \gamma_1, \gamma_2 \quad (12)$$

and

$$(\overline{p-q})_l = (p^0 - q^0, 0, -\sqrt{2eBl}, p^3 - q^3).$$

Anomalous Magnetic Moment

The general structure of the self-energy is:

$$\tilde{\Sigma}^l(\bar{p}) = Z_{\parallel}^l \bar{p}_{\parallel}^{\mu} \gamma_{\mu}^{\parallel} + Z_{\perp}^l \bar{p}_{\perp}^{\mu} \gamma_{\mu}^{\perp} + M^l I + iT^l \gamma^1 \gamma^2 \quad (13)$$

where γ_{μ} are the Dirac matrixes, Z_{\parallel}^l , Z_{\perp}^l are the wave function's renormalization coefficients, M^l and T^l are the mass and anomalous magnetic moment terms.

Now we can compare (11) and (13) to obtain the AMM term:

$$M^0 + T^0 = \frac{e^2 m}{8\pi^2} \int d\hat{q}_{\parallel}^2 d\hat{q}_{\perp}^2 \frac{e^{-\hat{q}_{\perp}^2}}{\hat{q}^2} \left[\frac{1}{\hat{q}_{\parallel}^2 + \hat{m}^2} + \frac{1}{\hat{q}_{\parallel}^2 + 1 + \hat{m}^2} \right]^{\dagger}, \quad (14)$$

$$T^{l \neq 0} = -\frac{e^2 m}{16\pi^2} \int d\hat{q}_{\parallel}^2 d\hat{q}_{\perp}^2 \frac{e^{-\hat{q}_{\perp}^2}}{\hat{q}^2} \left[\frac{1}{\hat{q}_{\parallel}^2 + l + 1 + \hat{m}^2} - \frac{1}{\hat{q}_{\parallel}^2 + l - 1 + \hat{m}^2} \right] \quad (15)$$

These integrals must be solved numerically.

†E. Ferrer and V de la Incera, Nuclear Physics B 824, 217 (2010).

Anomalous Magnetic Moment

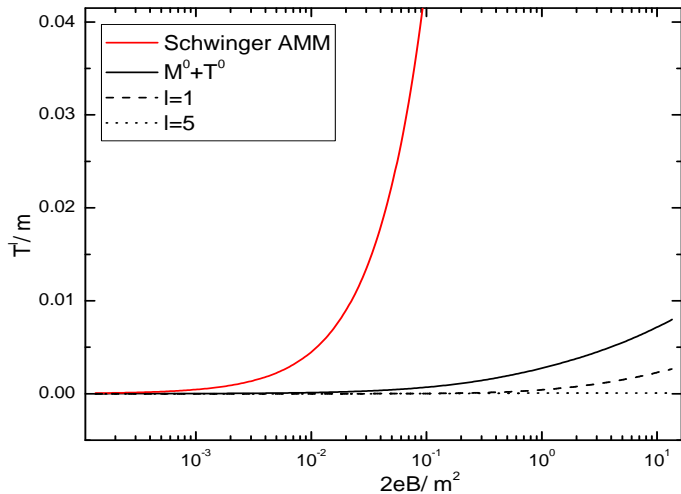


Figure: AMM for $l = 0, 1, 5$. Schwinger value is shown for comparison.

Dispersion relation of magnetized fermions

The dispersion relation is obtained from G_l^{-1} as

$$\det G_l^{-1} = \det[\bar{p} \cdot \gamma - m - M^l I - iT^l \gamma^1 \gamma^2], \quad (16)$$

to give the following particle energy spectra

$$\epsilon_{\sigma,l}^2 = p_3^2 + [\sqrt{(m + M^l)^2 + 2|eB|l} + \sigma T^l]^2, \quad \sigma = \pm 1 \quad (17)$$

and the LLL energy is

$$\epsilon_{1,0}^2 = p_3^2 + (m + M^0 + T^0)^2 \quad (18)$$

so the rest-energy is

$$\epsilon^0 = m + M^0 + T^0. \quad (19)$$

Dispersion relation of magnetized fermions

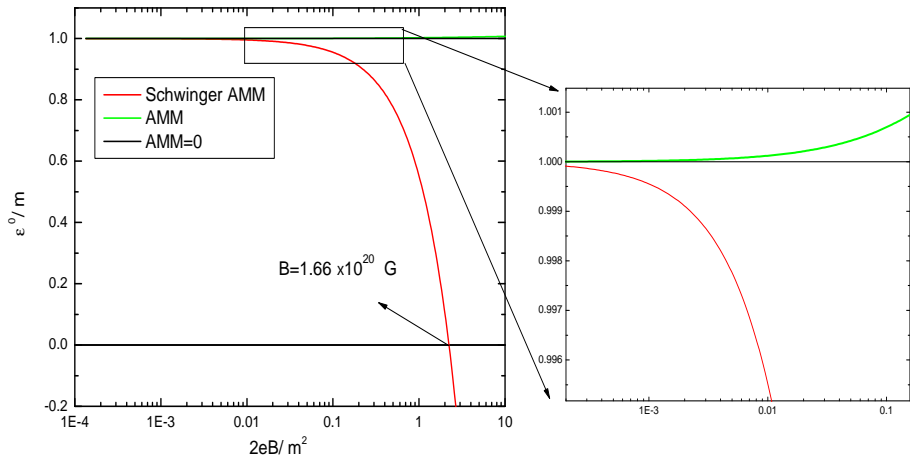


Figure: Rest energy for fermions in a constant magnetic field.

AMM effect in the thermodynamical properties

The thermodynamical potential at finite temperature and density for the magnetized system ($\Omega_f(B, \mu, T)$) including in the fermion inverse propagator the quantum contribution of the AMM is

$$\Omega_f(B, \mu, T) = -\frac{eB}{\beta} \left[\sum_{p_4} \int_{-\infty}^{\infty} \frac{dp_3}{(2\pi)^2} \ln \det \tilde{G}_0^{-1}(\bar{p}^*) + \sum_{\sigma \pm 1} \sum_{l=1}^{\infty} \sum_{p_4} \int_{-\infty}^{\infty} \frac{dp_3}{(2\pi)^2} \ln \det \tilde{G}_l^{-1}(\bar{p}^*) \right], \quad (20)$$

here $\bar{p}_\nu^* = (ip^4 - \mu, 0, \sqrt{2eBl}, p^3)$.

We can obtain the limit $T = 0...$

AMM effect in the thermodynamical properties

$$\Omega = -B \frac{em^2}{4\pi^2} \left[\Omega^0 + \sum_{l=1}^{l_{max}} \sum_{\sigma=\pm 1} \left(\tilde{\mu} \tilde{p}_F^\sigma - (\varepsilon_{\sigma,l}^0)^2 \ln \frac{\tilde{\mu} + \tilde{p}_F^\sigma}{\varepsilon_{\sigma,l}^0} \right) \right], \quad (21)$$

where

$$\Omega^0 = \left(\tilde{\mu} \tilde{p}_F^0 - (\varepsilon_{1,0}^0)^2 \ln \frac{\tilde{\mu} + \tilde{p}_F^0}{\varepsilon_{1,0}^0} \right), \quad (22)$$

$$\tilde{p}_F^\sigma = \sqrt{\tilde{\mu}^2 - (\varepsilon_{\sigma,l}^0)^2}, \quad \tilde{p}_F^0 = \sqrt{\tilde{\mu}^2 - (\varepsilon_{1,0}^0)^2}, \quad \tilde{\mu} = \frac{\mu}{m} \quad (23)$$

$$\varepsilon_{\sigma,l}^0 = \begin{cases} \sqrt{(m + M^l)^2 + 2|eB|l} + \sigma T^l, & \text{AMM,} \\ \sqrt{m^2 + 2eBl} - \sigma \kappa \mu_B B, & \text{Schwinger AMM.} \end{cases} \quad (24)$$

AMM effect in the thermodynamical properties

From Ω we can obtain the particle density $N = -(\partial\Omega/\partial\mu)$, the magnetization $M = -(\partial\Omega_f/\partial B)$, the energy density and the pressures

$$N = \frac{m^3}{2\pi^2} \frac{B}{B_c} \left(\tilde{p}_F^0 + \sum_{l=1}^{l_{max}} \sum_{\sigma=\pm 1} \tilde{p}_F^\sigma \right), \quad (25)$$

$$M = \frac{em^2}{4\pi^2} \left\{ M^0 + \sum_{l=1}^{l_{max}} \sum_{\sigma=\pm 1} \tilde{\mu} \tilde{p}_F^\sigma - [(\varepsilon_{\sigma,l}^0)^2 + 2\varepsilon_{\sigma,l}^0 \tilde{\gamma}^\sigma] \ln \frac{\tilde{\mu} + \tilde{p}_F^\sigma}{\varepsilon_{\sigma,l}^0} \right\}, \quad (26)$$

$$E = \Omega + \mu N, \quad (27)$$

$$P_{\parallel} = -\Omega, \quad (28)$$

$$P_{\perp} = -\Omega - BM, \quad (29)$$

where

$$M^0 = \tilde{\mu} \tilde{p}_F^0 - [(\varepsilon_{1,0}^0)^2 + 2\varepsilon_{1,0}^0 \tilde{\gamma}^0] \ln \frac{\tilde{\mu} + \tilde{p}_F^0}{\varepsilon_{1,0}^{(0)}},$$

$$\tilde{\gamma}^\sigma = \frac{\partial \varepsilon_{\sigma,l}^0}{\partial B}$$

AMM effect in the thermodynamical properties

Results

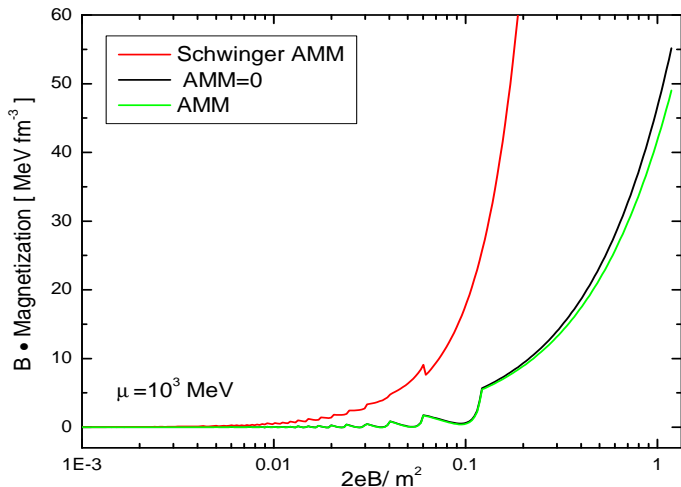


Figure: Magnetization vs B . Schwinger value and AMM=0 are shown for comparison.

AMM effect in the thermodynamical properties

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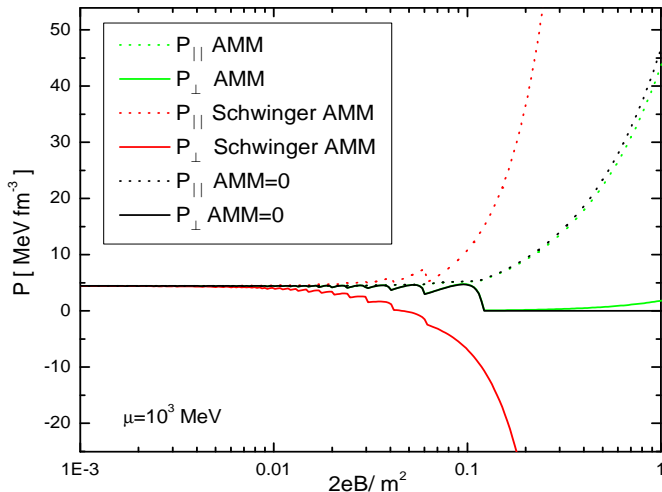


Figure: Pressures vs B . Schwinger value and AMM= 0 are shown for comparison.

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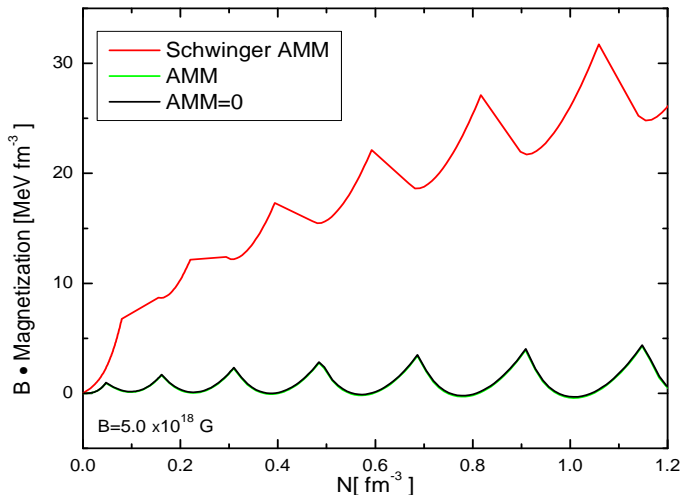


Figure: Magnetization vs density. Schwinger value and AMM= 0 are shown for comparison.

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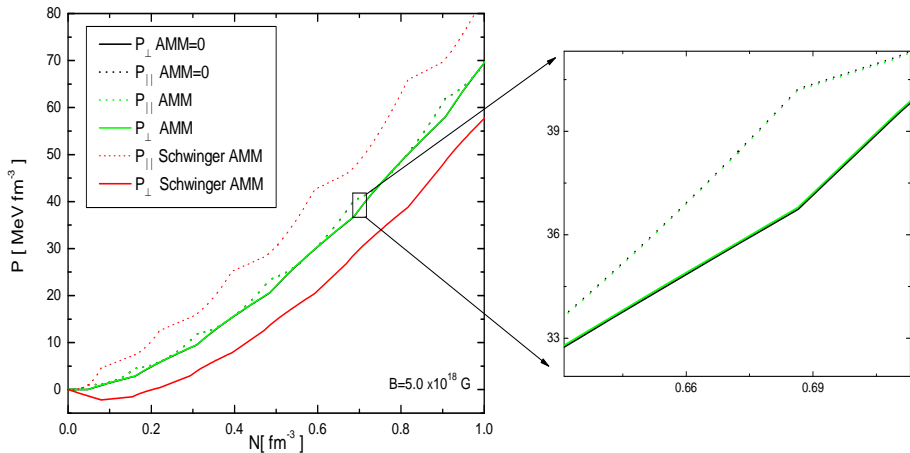


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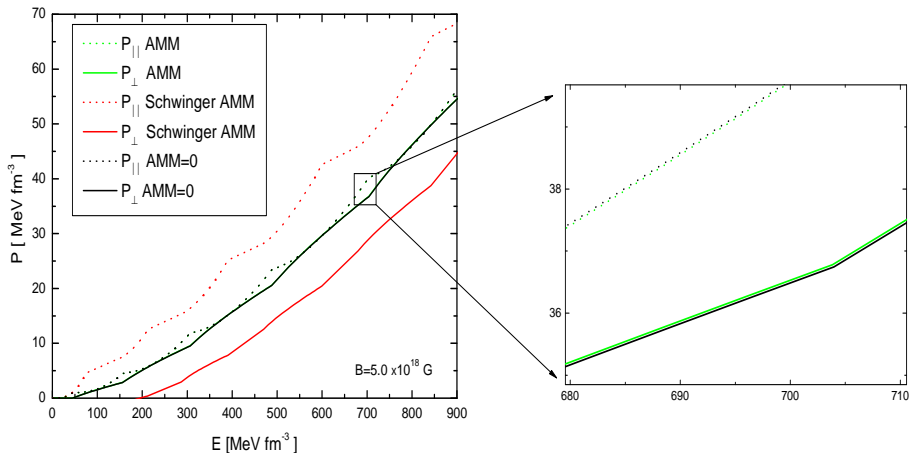


Figure: EoS, Schwinger value is shown for comparison.

Summary & Conclusions

- 1 The AMM for fermions in presence of constant magnetic field has been calculated at $T = 0$ in a consistent way (exact in B).
 - Schwinger AMM is valid if $\frac{B}{B_c} \lesssim 10^{-3}$
- 2 The effects of the AMM in the thermodynamical quantities has been studied.
 - The effects of the AMM in our approximation are negligible for $B \lesssim B_c$.
- 3 Our result is also applicable to other charged fermions with AMM.

Future work:

- Include $T \neq 0$.
- Include $\mu \neq 0$ in the AMM.
- Prove that our conclusions are true under stellar equilibrium condition.

Thanks
Gracias