

A Search for Direct Detection of Dark Contents of the Vacuum and Dark Energy

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On Behalf of the Collaboration



UNIVERSITY OF
LIVERPOOL

Martin Perl et al.
<http://arxiv.org/abs/1101.5626>



The Cockcroft Institute
of Accelerator Science and Technology

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Introduction

- Motivate the search to detect dark energy terrestrially.
- What is the technology and how does it work?

Dark Energy



The Nobel Prize in Physics 2011
Saul Perlmutter, Brian P. Schmidt, Adam G. Riess



Photo: U. Montan

Saul Perlmutter



Photo: U. Montan

Brian P. Schmidt



Photo: U. Montan

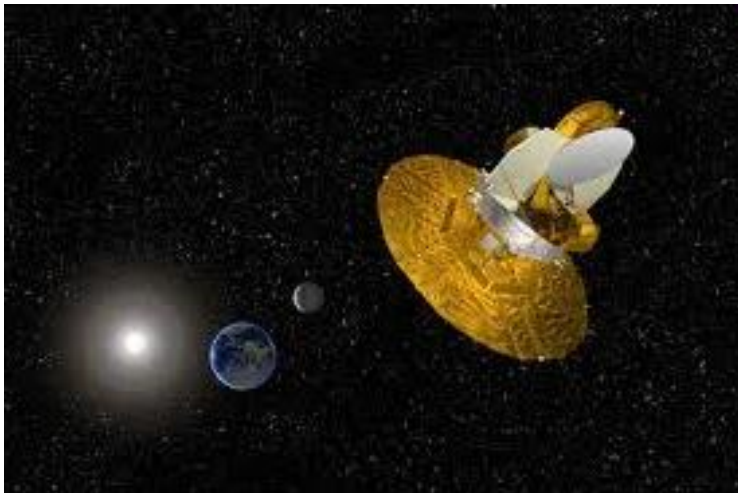
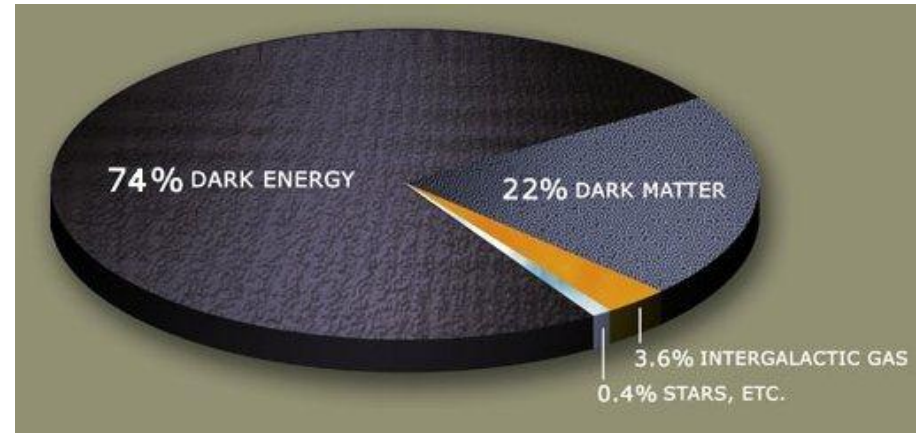
Adam G. Riess

The Nobel Prize in Physics 2011 was divided, one half awarded to Saul Perlmutter, the other half jointly to Brian P. Schmidt and Adam G. Riess *"for the discovery of the accelerating expansion of the Universe through observations of distant supernovae"*.

Photos: Copyright © The Nobel Foundation

Dark Energy

- 74% of the universe is dark energy.
- Causes the expansion of the universe to accelerate.



- Observations from Super Novae, CMB and large scale structures all agree with the Λ -CDM cosmological model of the universe.
- Can dark energy be explored terrestrially?

Terrestrial search

- Average energy density of dark energy is $\rho_{\text{DE}} = 6.3 \times 10^{-10} \text{ J m}^{-3}$.
- Same energy density as a 12 Vm^{-1} electric field!
- 12 Vm^{-1} electric field easily detectable in lab.
- Dark energy is small but non zero, may be detectable.

a SQUID in a Undergrad Lab...



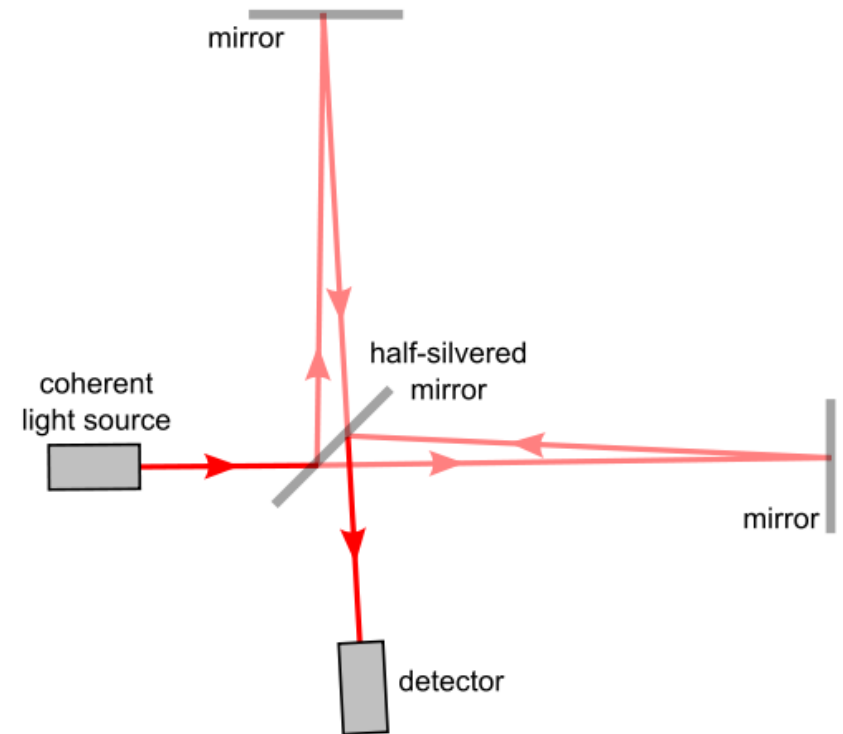
Conditions on detection

Dark energy can be detected **if**;

1. Dark energy is spatially inhomogeneous on the lab scale.
2. Dark energy interacts with atoms in a non gravitational way.

How does interferometry work?

- Split light into two beams.
- Beams travels two different paths.
- Recombine the beams into one.
- Measure the phase difference between the two paths.
- Phase difference related to difference in paths.



Why Atoms?

- Smaller wavelengths lead to a higher accuracy.
- Atoms have a wavelength $\lambda = h/p$

Laser Cooling

Laser cooling techniques are used to achieve the required velocity (wavelength) control for the atom source.

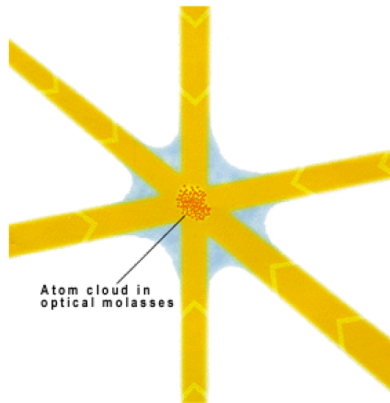


Image source: www.nobel.se/physics

Laser cooling:
Laser light is used to cool atomic vapors to temperatures of $\sim 10^{-6}$ deg K.



The Nobel Prize in Physics 1997

"for development of methods to cool and trap atoms with laser light"



Steven Chu



USA

Stanford University
Stanford, CA, USA

1948 -



Claude Cohen-Tannoudji



France

Collège de France
Paris, France
and École Normale Supérieure
Paris, France

1933 -



William D. Phillips



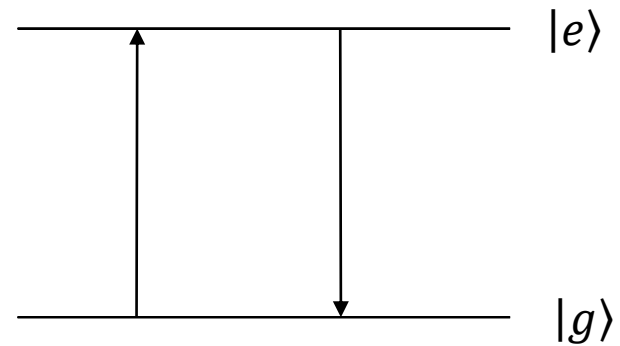
USA

National Institute of Standards and Technology
Gaithersburg, Maryland, USA

1948 -

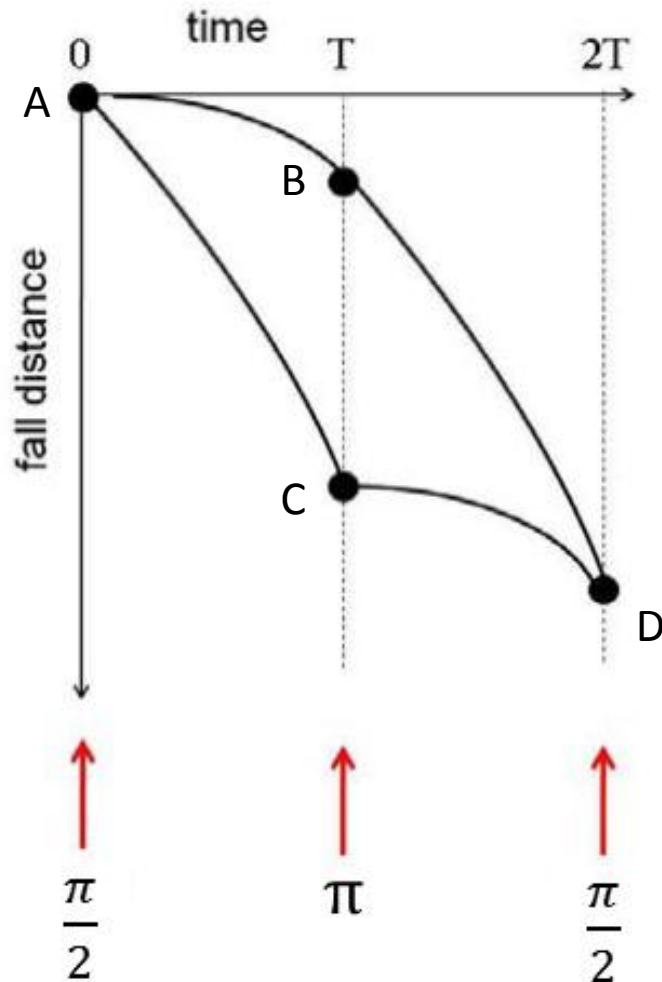
Atomic mirrors and beam splitters

- A Rabi cycle $|g\rangle \rightarrow |e\rangle \rightarrow |g\rangle$ under stimulated emission/absorption.
- Half an cycle is a π pulse.
- π pulses switches $|g\rangle$ for $|e\rangle$ and visa verse.
- Quarter of a cycle is a $\frac{\pi}{2}$ pulse.
- A $\frac{\pi}{2}$ pulse can put the atom into a superposition of two states



$$\text{i.e. } |g\rangle \rightarrow \frac{1}{\sqrt{2}} (|g\rangle + e^{-i\omega t} |e\rangle)$$

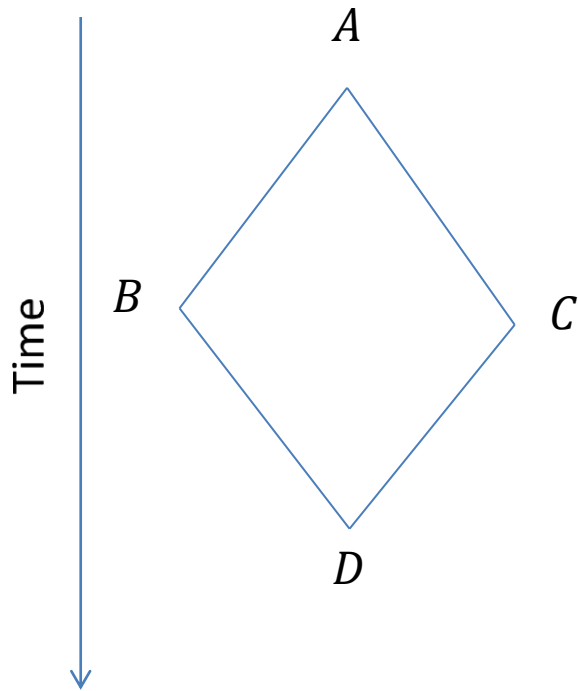
Interferometry Sequence



| | | | |
|-------------------|-----------------|--------|-----------------|
| Time | 0 | T | 2T |
| Pulse type | $\frac{\pi}{2}$ | π | $\frac{\pi}{2}$ |
| Analogy | Beam-splitter | Mirror | Beam-splitter |

- Figure shows different paths in the atom interferometer.

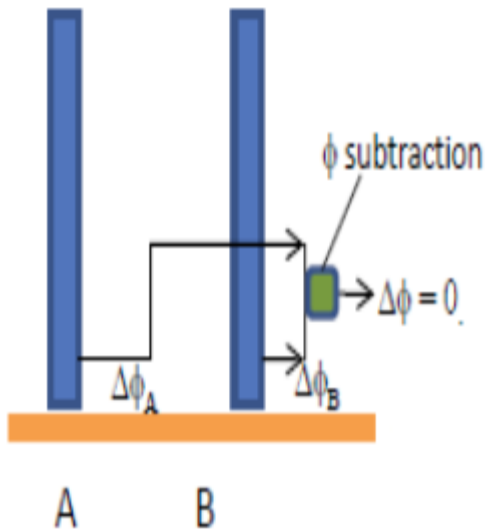
Measurement



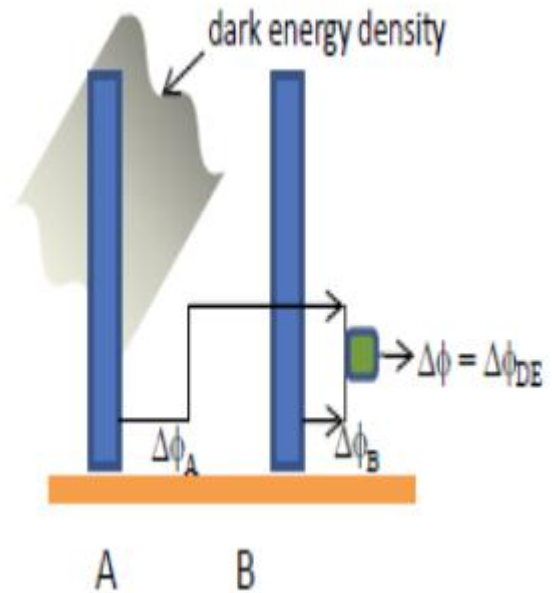
- ϕ_{ABD} is the phase from the path integral for an atom that has travelled the path ABD.
- ϕ_{ACD} is the phase from from the path integral for an atom that has travelled the path ACD.
- $\Delta\phi = \phi_{ABD} - \phi_{ACD}$ is the phase difference for the interferometry measurement.

Atom Interferometry

- An atom interferometer measures phase difference $\Delta\phi$ of the atoms.
- Two interferometers should measure the same $\Delta\phi$



- But if there is a greater dark energy density in one interferometer, it could affect the $\Delta\phi$ in one interferometer.
- This appears as a noise like signal.



Progress

- Designing and constructing our first interferometer prototype now.
- Vacuum system operating at $> 10^{-10}$ mbar.
- Waiting on last pieces of equipment to trap the atoms, should be trapping in the next few weeks.
- Expect to have a working interferometer within the next 12 months.

Summary

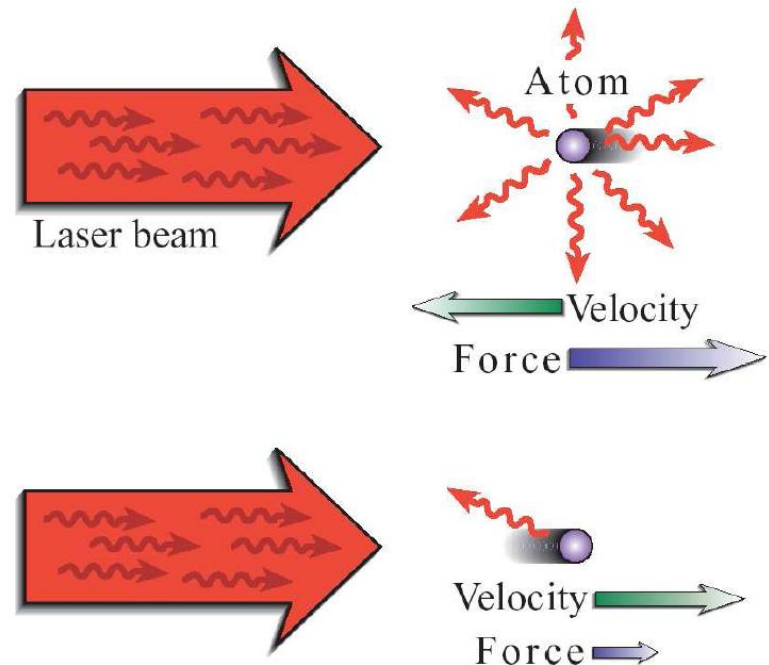
- Dark energy could be detected using atom interferometry if it is;
 - Spatially inhomogeneous on lab scale.
 - Interacts in a non gravitational way with matter.
- Explained how atom interferometry works.
- How a dark energy can produce a noise like signal between two atom interferometers.

Thanks for listening!

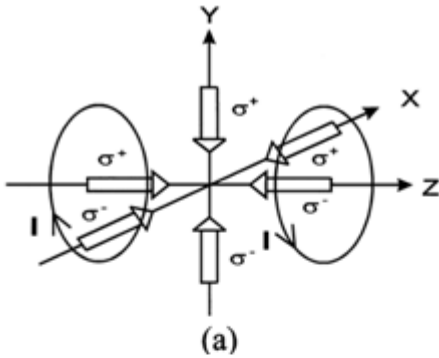
Any questions?

Atom Trapping

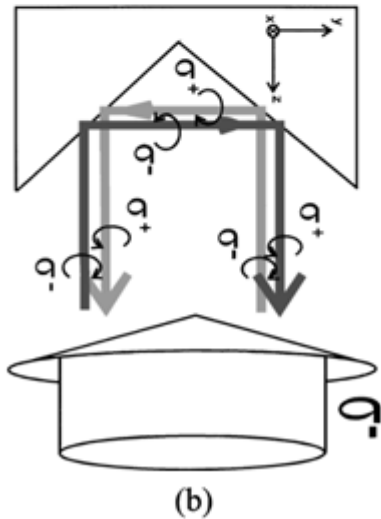
- Cooling laser frequency is just less than a transition.
- Atoms moving towards laser beam interact and slow down.
- Atoms emits photon in random direction.
- Overall beam slows down atoms.



Beam of atoms

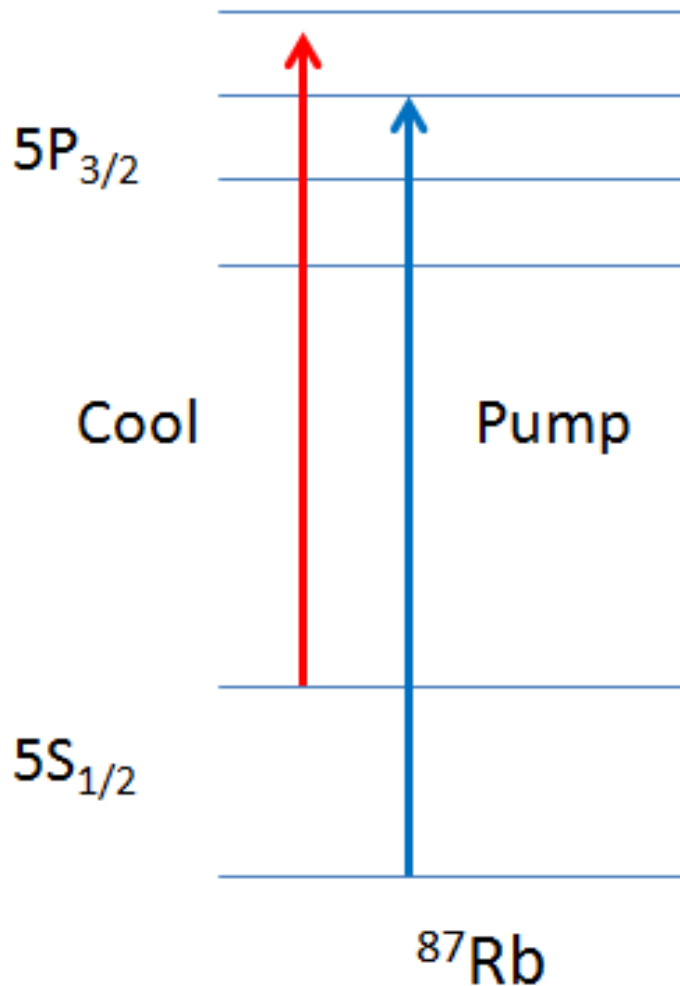


- Beam of atoms dropped from an optical trap.



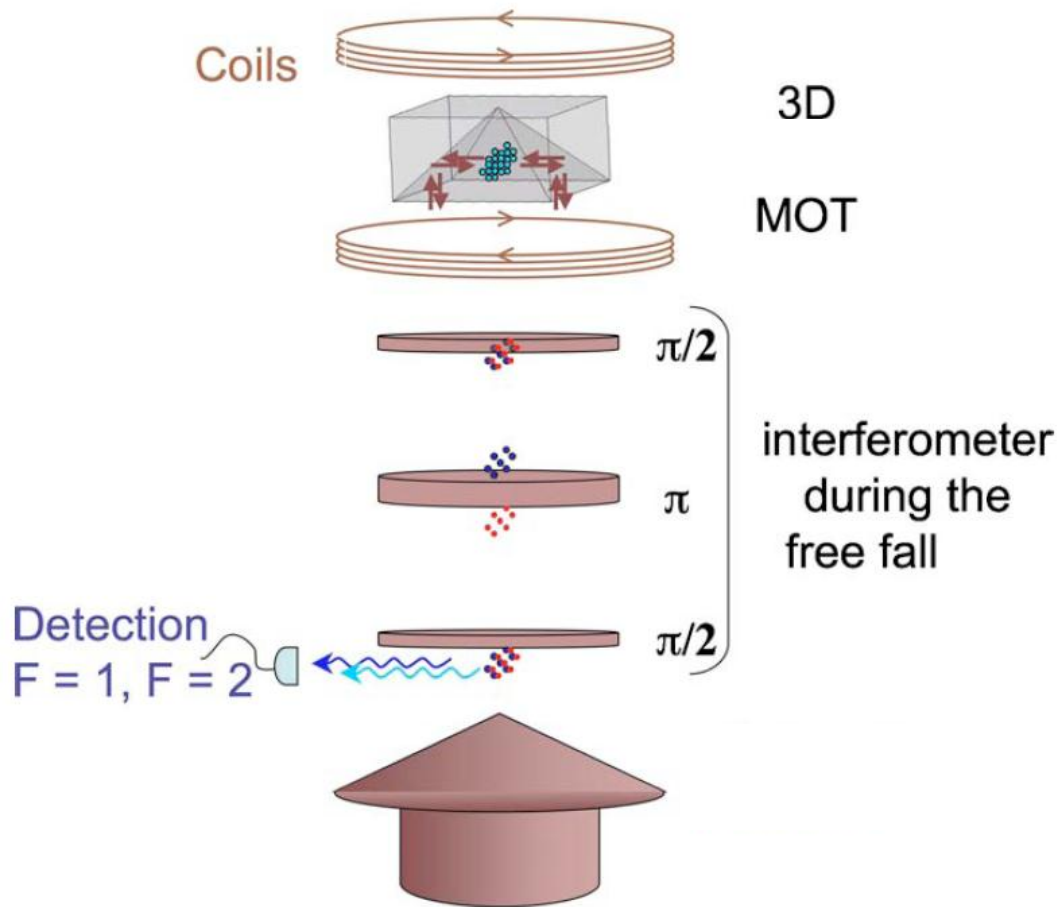
- Calibrating 6 independent beams – difficult!
- Hollow pyramidal mirror creates 6 beams when illuminated with the laser light.

Rubidium

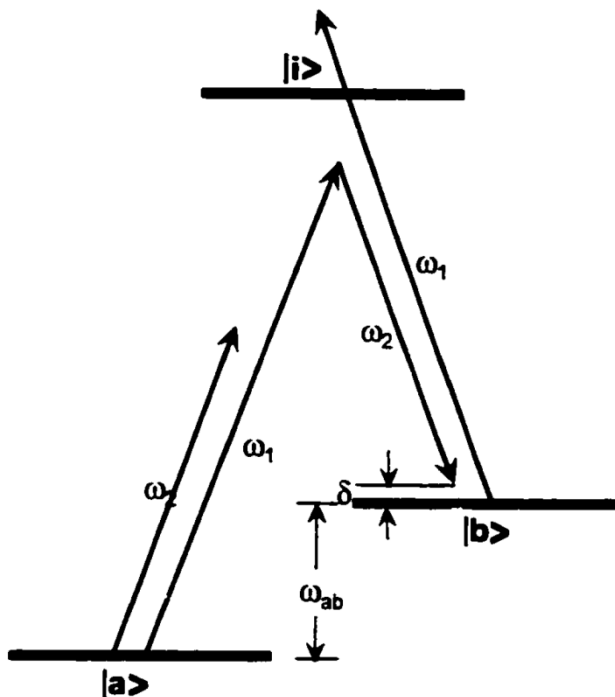


- Having just 1 electron in the outer shell makes alkali metals good for atomic cooling experiments.
- Easy to manipulate the one electron.

Interferometry Sequence

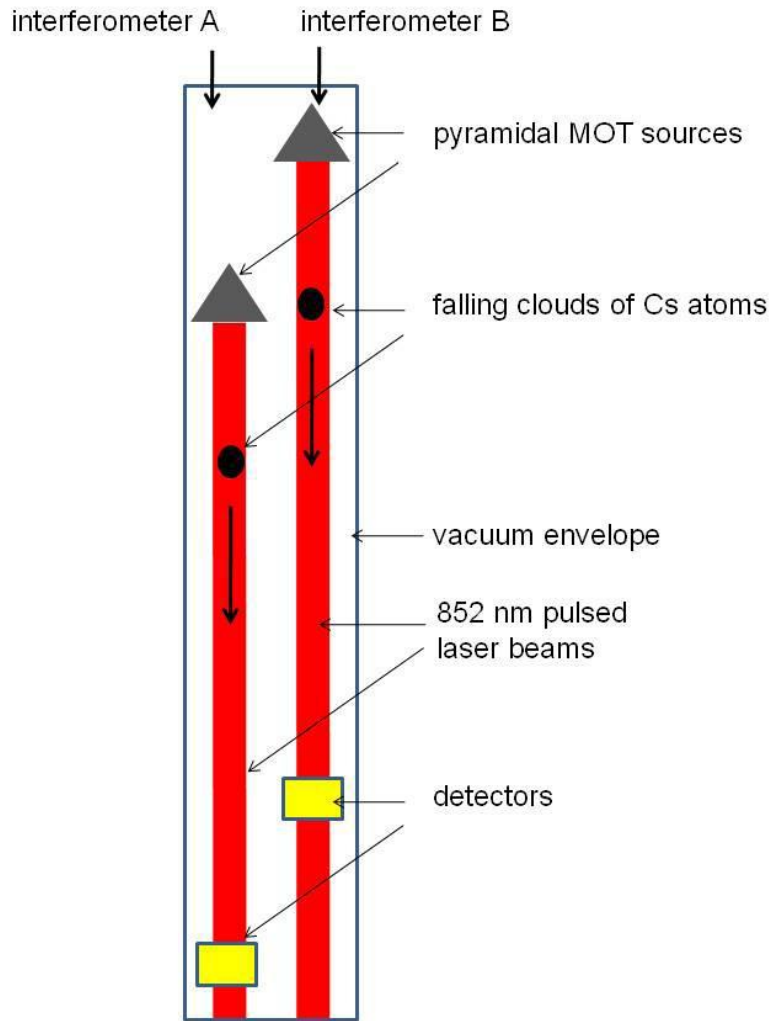


Raman Pulses



- Want the atoms in stable states – oscillate between the two spin states of ground state.
- Need laser light of two frequencies, with frequency difference equal to required transition.
- Think of 2 photon transition via virtual state.

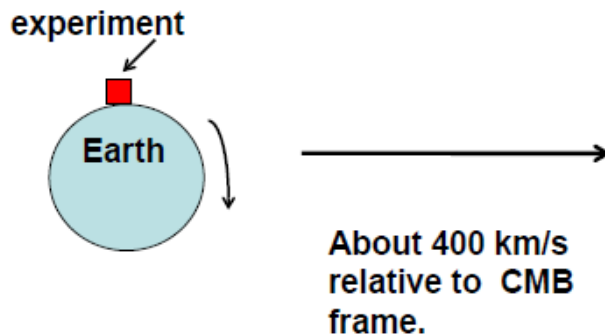
Experimental Configuration



- To cancel systematic effects:
- Incorporate two interferometers in one vacuum envelope,
 - reduce problems from common mode noises such as vibrations.
 - drop sources for simplicity.
- Sources are staggered vertically,
 - total phase change for each atom is measured during the same velocity period.

NATURE OF THE SOUGHT SIGNAL

- We do not record $\Delta\phi$ which will average to zero, we record the root mean square $\Delta\phi$ rms. We can then determine the dark energy equivalent acceleration, g_{DE} .
- We expect to be able to detect the dark energy equivalent acceleration, g_{DE} with a precision of $\rightarrow 10^{-15} \text{ m/s}^2$.



Apparatus sampling rate is of the order of Hertz.

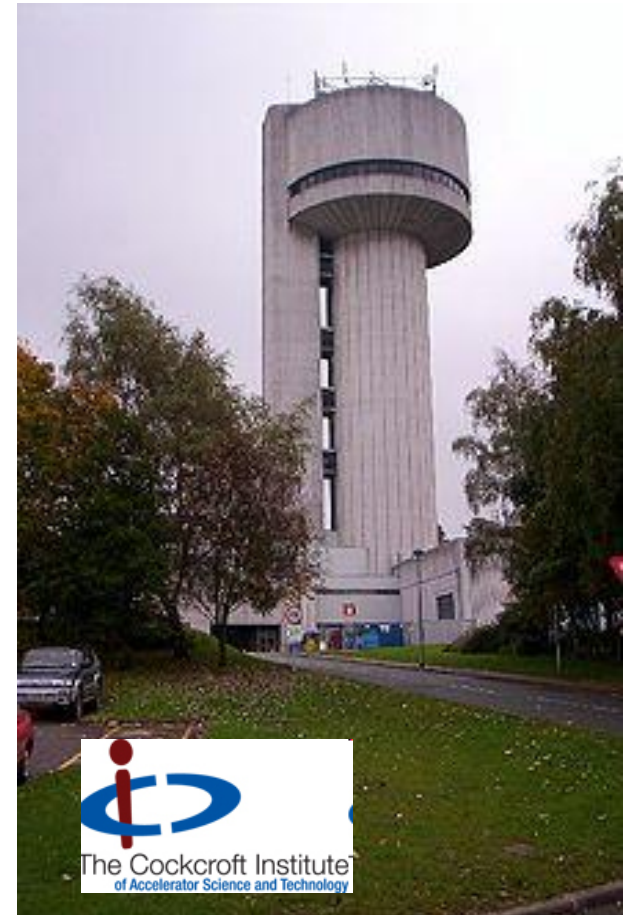
Hence the dark energy signal is a noise signal.

Improving Sensitivity

- A crucial factor in the sensitivity of the experiment depends upon having a large phase shift ϕ where

$$\phi = \text{constant } (gT^2)$$

- Here g is the acceleration of gravity and T is the time it takes for the atom cloud to fall from the source at the top to the detector at the bottom of the interferometer.
- In the apparatus under construction the fall height, h , is 1 m and f is about 107 radians.
- Since T^2 is proportional to h , f is proportional to h . if $h = 10$ m would give 10 times the present f , in principle the higher the better.
- Daresbury tower, is ~ 100 m high, benefits of exploiting this structure are obvious.





Progress

- The construction of two symbiotic experiments, began about one year ago using:
 - start-up funds from the President's Office of Stanford University.
 - Seed-corn funding and technical support from particle physics group @ U. of Liverpool, and the Cockcroft Institute, UK.

