



Low-level radon detection in the SuperNEMO detector

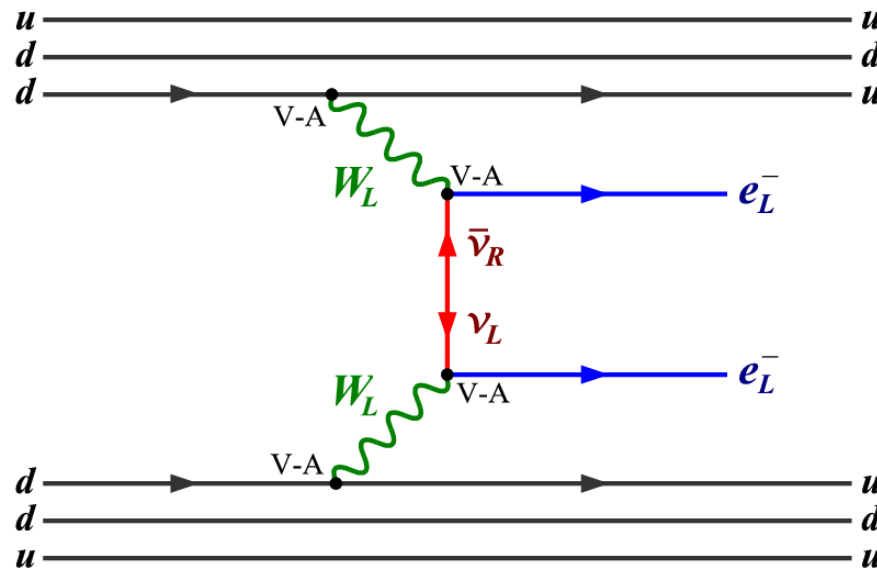
James Mott

IoP Joint HEPP and APP Meeting 2012

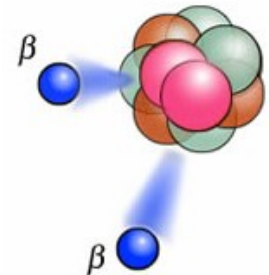
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Neutrinoless Double Beta Decay ($0\nu\beta\beta$)

- $0\nu\beta\beta$ is a lepton-number violating transition:



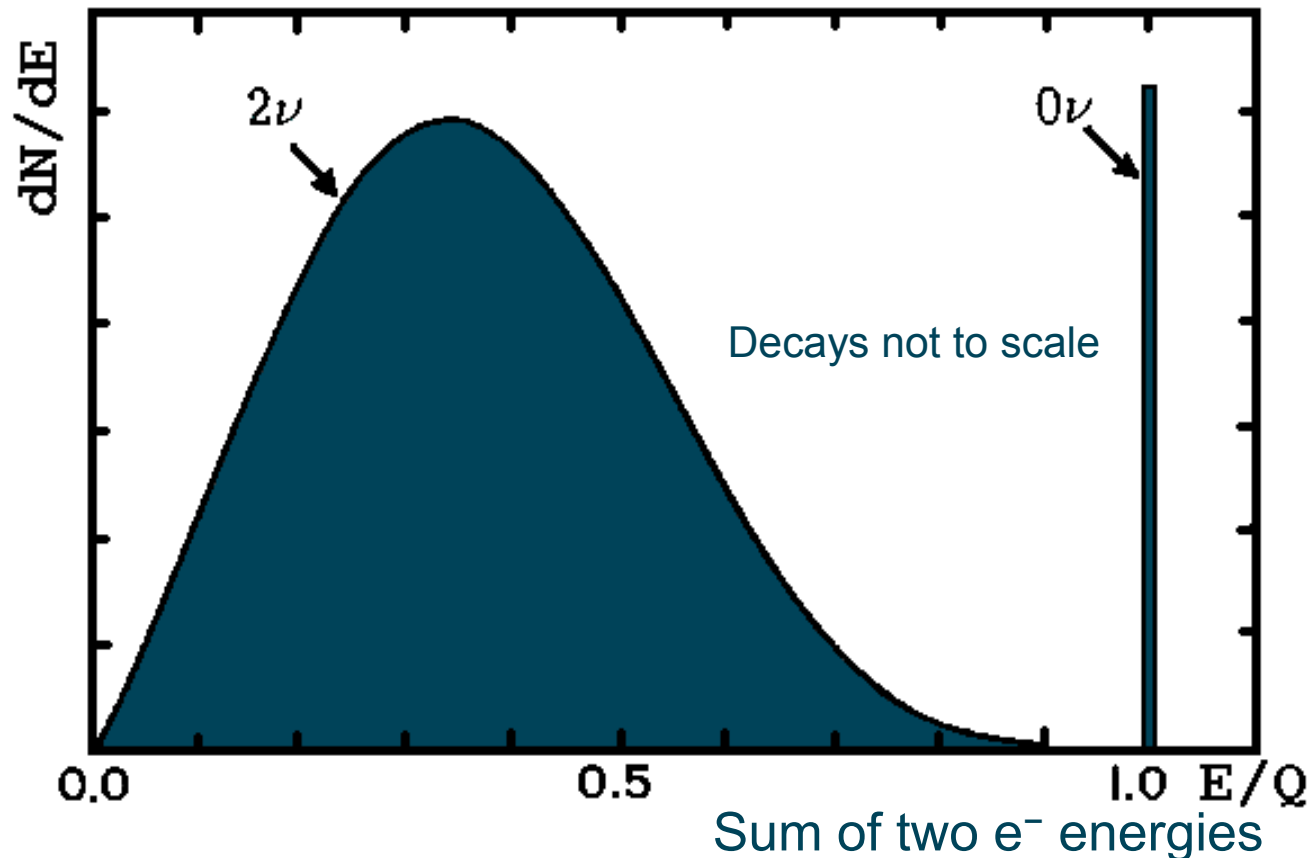
One possible mechanism: Light Majorana neutrino exchange



- $0\nu\beta\beta$ is one of the most promising ways to answer:
 - Is the neutrino a Majorana or Dirac particle?
 - What is the absolute neutrino mass scale?

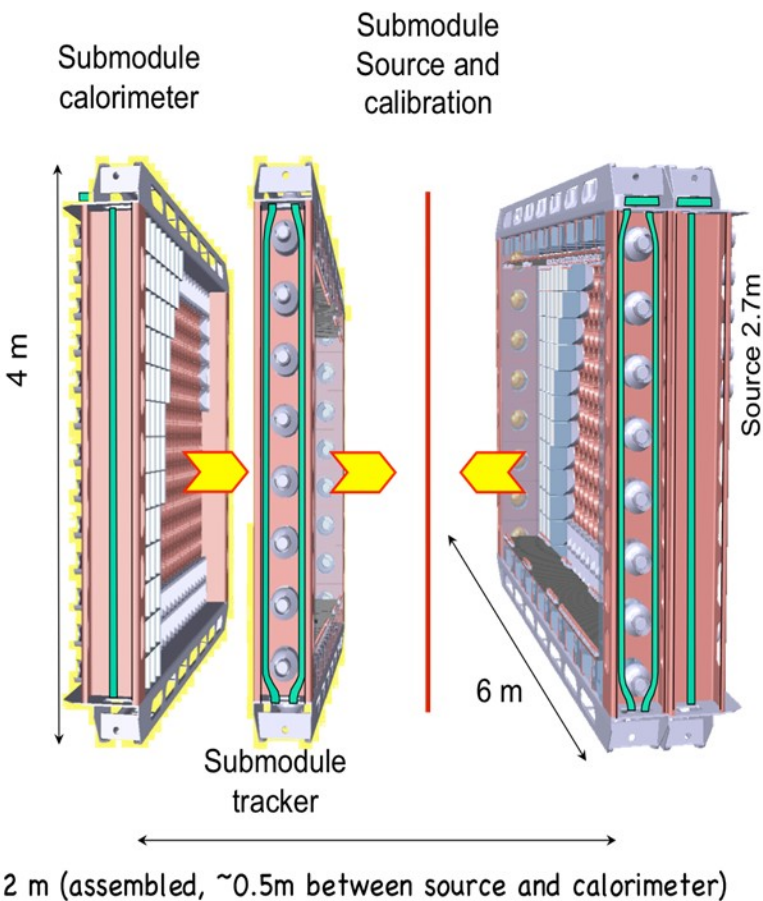
$0\nu\beta\beta$: Experimental Signature

- The majority of experiments measure the sum of the energy of the two electrons to identify the decay.
- The signal for $0\nu\beta\beta$ is a line at $Q_{\beta\beta}$ for the isotope:



The SuperNEMO Experiment

- SuperNEMO is a next-generation $0\nu\beta\beta$ experiment.
- It uses the same tracking and calorimeter techniques as NEMO-3



- It will consist of 20 identical planar modules, each containing:
 - **Source foil:** ~ 7 kg of ^{82}Se (or $^{150}\text{Nd}/^{48}\text{Ca}$)
 - **Tracker:** Drift chamber with ~ 2000 vertical cells in He, Ar and alcohol.
 - **Calorimeter:** 500 PMTs & plastic scintillator blocks
- Design allows particle ID, topology reconstruction & strong background rejection
- Aims to study half-life up to 10^{26} yr
 $\langle m_{\beta\beta} \rangle = 40 - 100$ meV

From NEMO-3 to SuperNEMO...

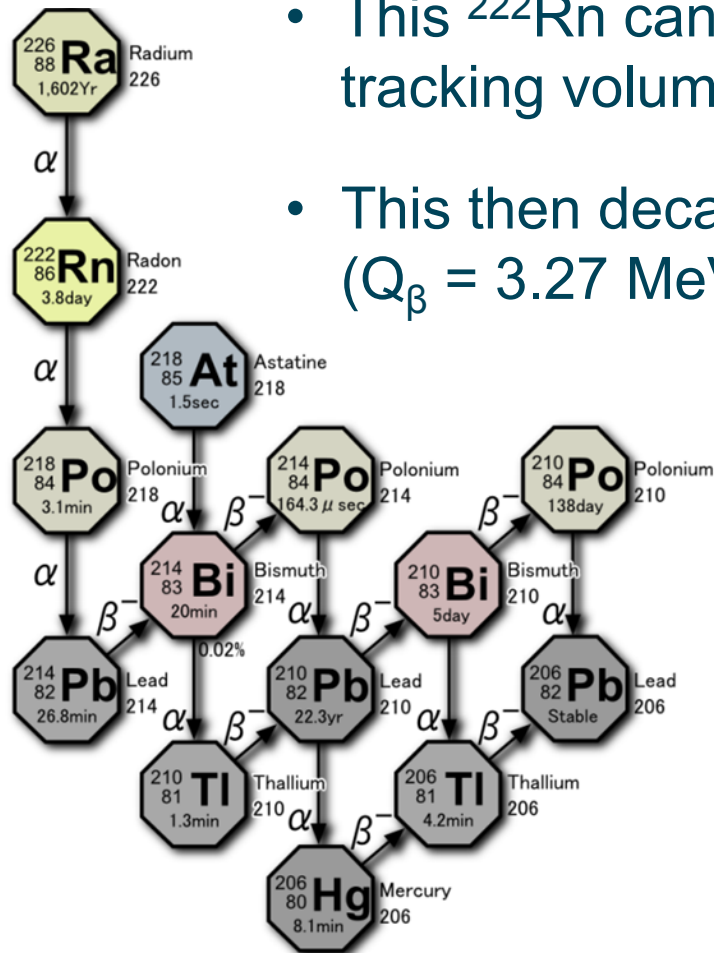
| NEMO-3 | → | SuperNEMO |
|--|------------------------------|--|
| ^{100}Mo | Isotope | ^{82}Se (or ^{150}Nd or ^{48}Ca) |
| 10 kg | Mass | 100+ kg |
| $< 5 \text{ mBq/m}^3$ | ^{222}Rn in tracker | $< 0.15 \text{ mBq/m}^3$ |
| 8% @ 3 MeV | Energy Resolution (FWHM) | 4% @ 3 MeV |
| $T_{1/2} > 10^{24} \text{ y}$ | Sensitivity | $T_{1/2} > 10^{26} \text{ y}$ |
| $\langle m_\nu \rangle < 0.3 - 0.9 \text{ eV}$ | | $\langle m_\nu \rangle < 40 - 100 \text{ meV}$ |

- 1st SuperNEMO module is being built will be installed by 2014.
- By 2016 it will be able to probe the controversial Klapdor claim to have observed $0\nu\beta\beta$ with $\langle m_\nu \rangle \sim 0.4 \text{ eV}$.
- It will be followed by 19 similar modules, to be built from 2015 and achieving full sensitivity by 2019.

Radon as a SuperNEMO background

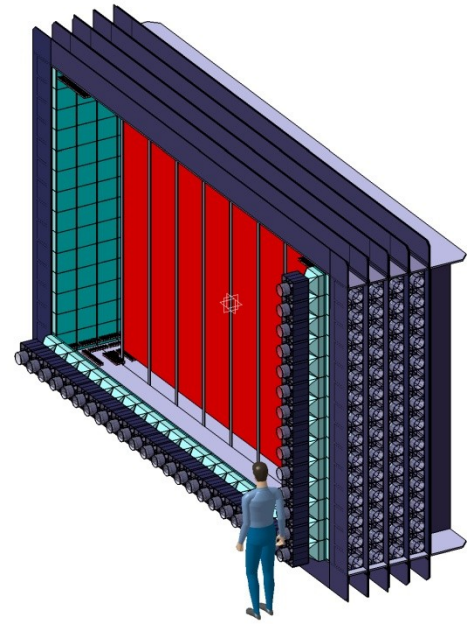
- All materials contain small traces of ^{226}Ra which decays to ^{222}Rn .

- This ^{222}Rn can diffuse out of the detector materials into the tracking volume.
- This then decays to ^{214}Bi which has a problematic β decay ($Q_\beta = 3.27 \text{ MeV}$)



- ^{222}Rn concentration in the tracker must be $< 0.15 \text{ mBq/m}^3$ to achieve the target sensitivity.
- Best radon detectors (electrostatic collection) are typically sensitive to $\sim 1 \text{ mBq/m}^3$ so to measure our target level we need a different technique.

Radon Concentration Line: Concept



SuperNEMO tracker sub-module (~ 14m³)



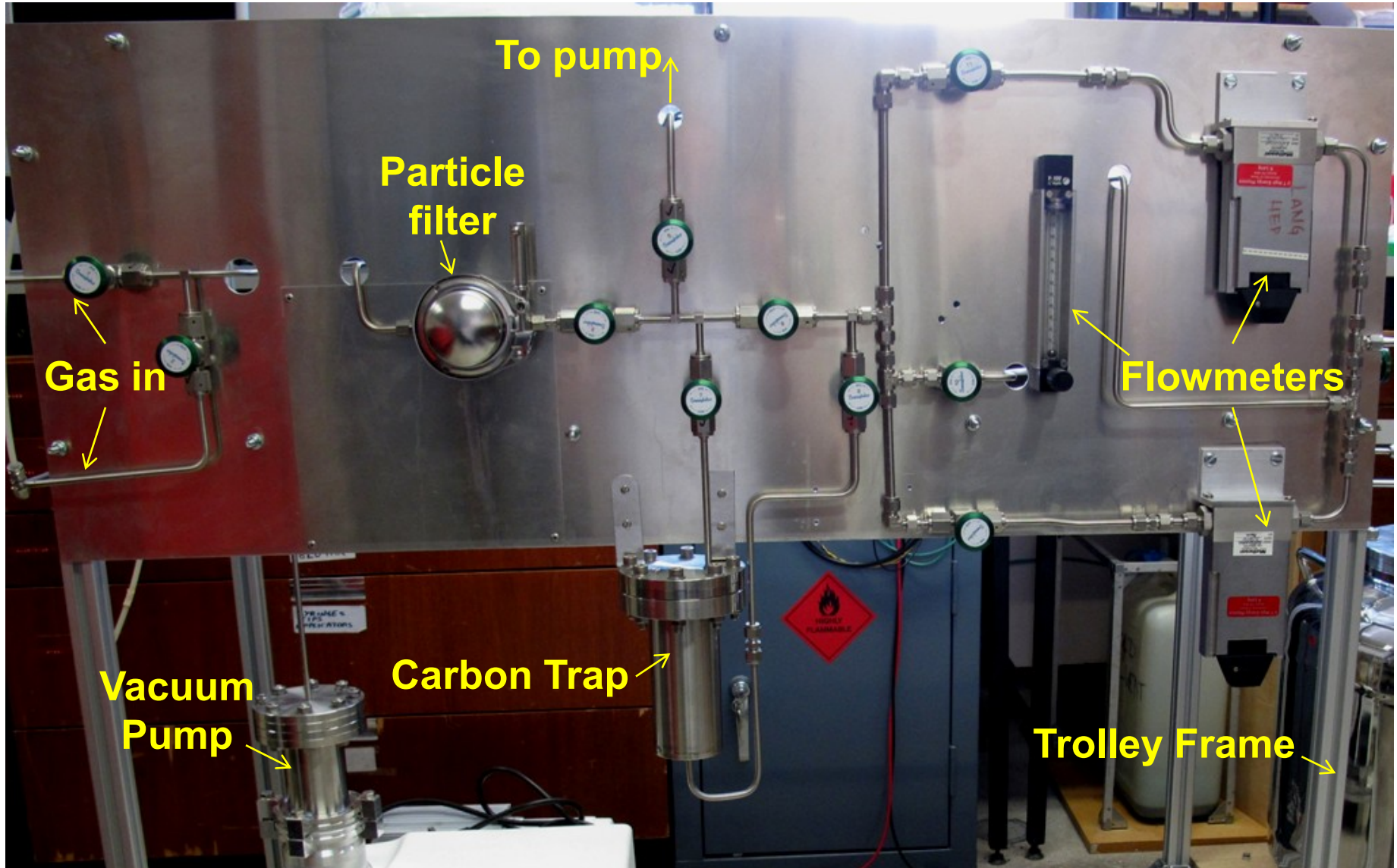
Radon concentration line



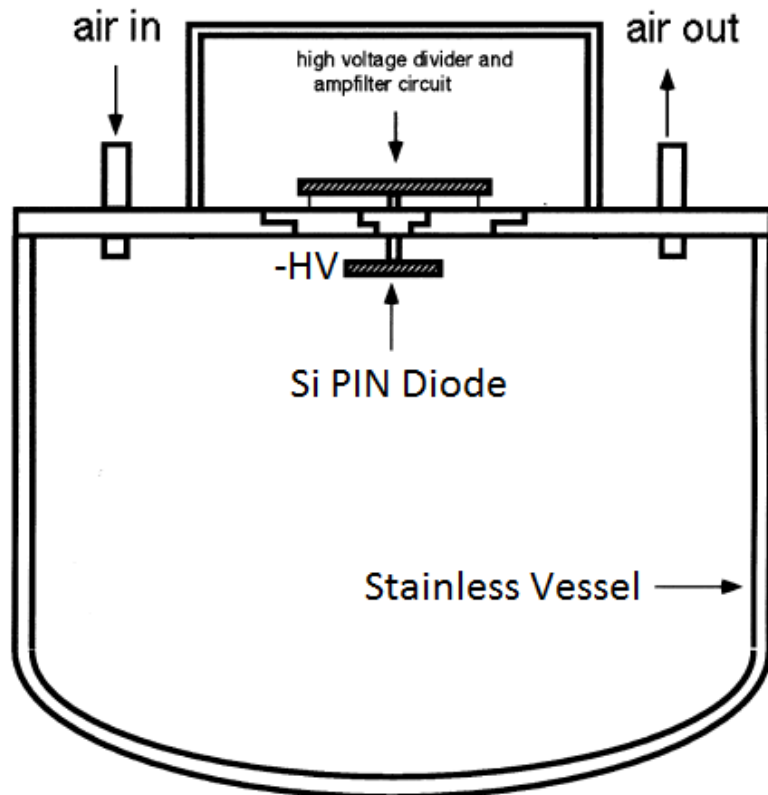
Electrostatic detector

- Gas from the tracker is pumped through a cooled ultra-pure carbon trap and the ²²²Rn in the gas is adsorbed.
- The concentrated sample is then heated and transferred to an electrostatic detector via helium purge.

Radon Concentration Line: Real Life

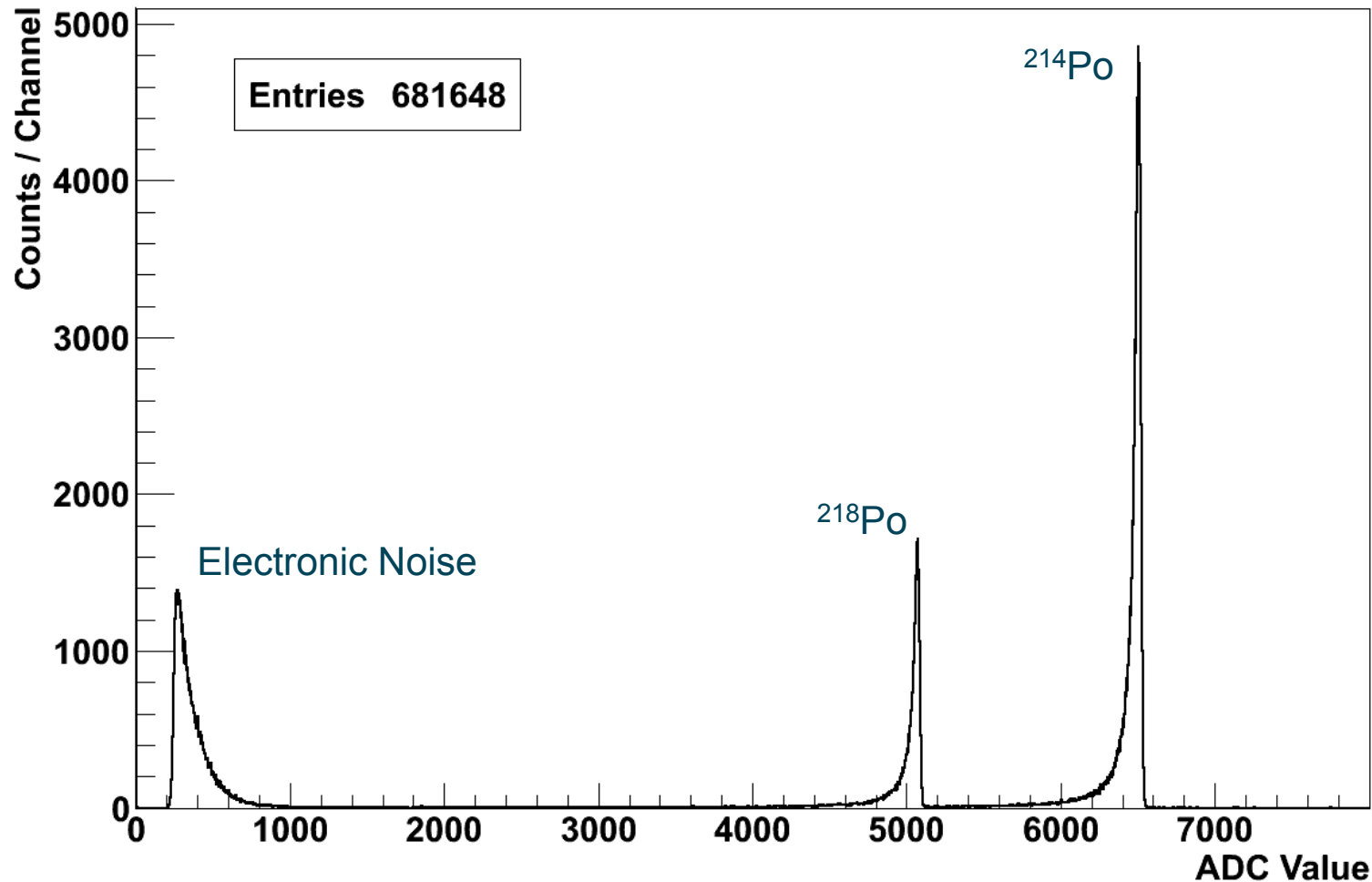


Electrostatic Detector: Principle of Operation



- ^{222}Rn is pumped into the vessel where it decays.
- Daughters of ^{222}Rn decay are mostly positive ions.
- These ions collect on the PIN diode due to the applied negative HV.
- Once on the photodiode, they decay and ^{218}Po , ^{214}Po & ^{210}Po alphas can be identified by the energy deposited.
- ^{214}Po is the most commonly used isotope for measuring ^{222}Rn .

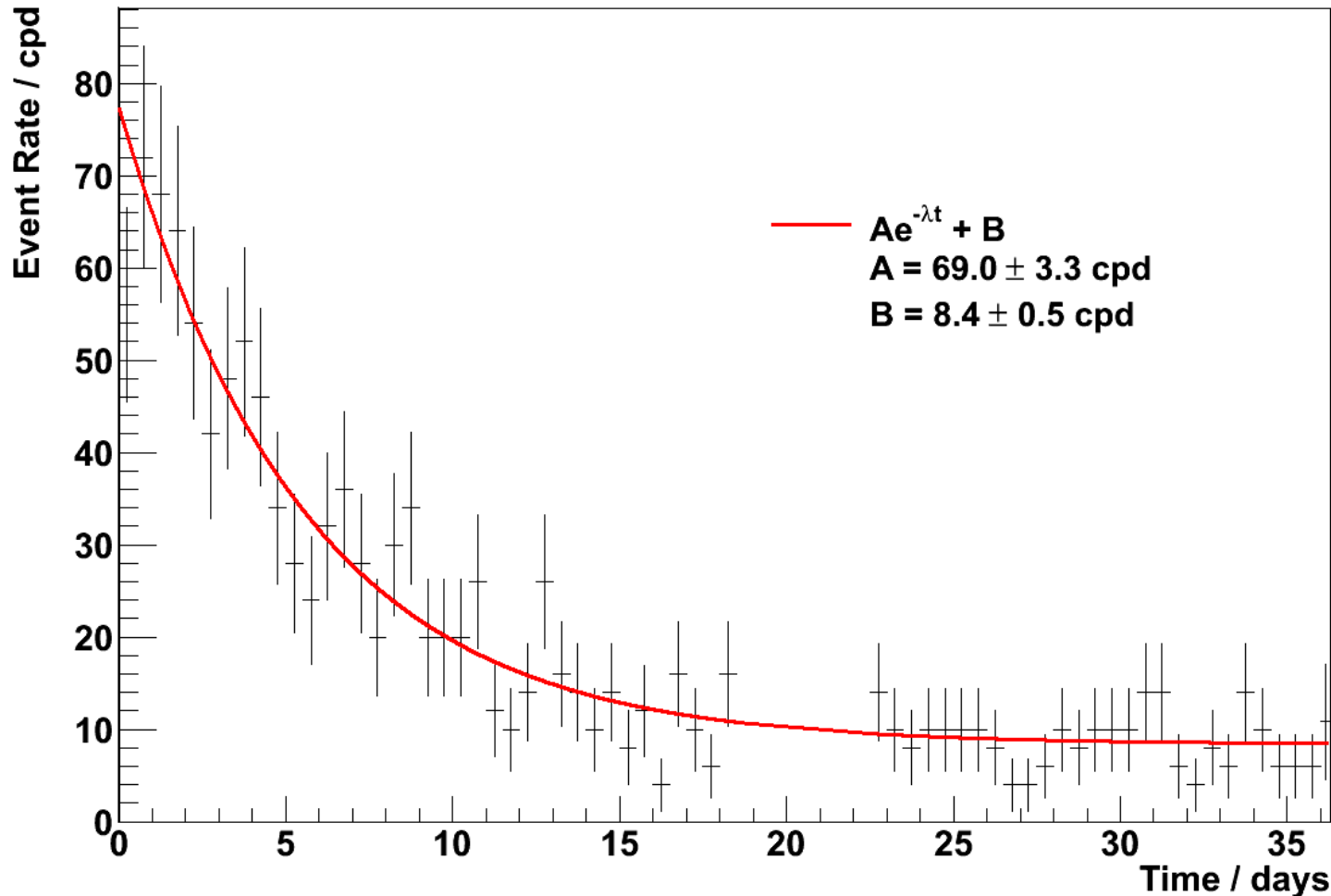
Electrostatic Detector: Energy Spectrum



- Peaks from ^{218}Po (6.1 MeV) and ^{214}Po (7.9 MeV) are visible and have excellent resolution

Electrostatic Detector: Background Rate

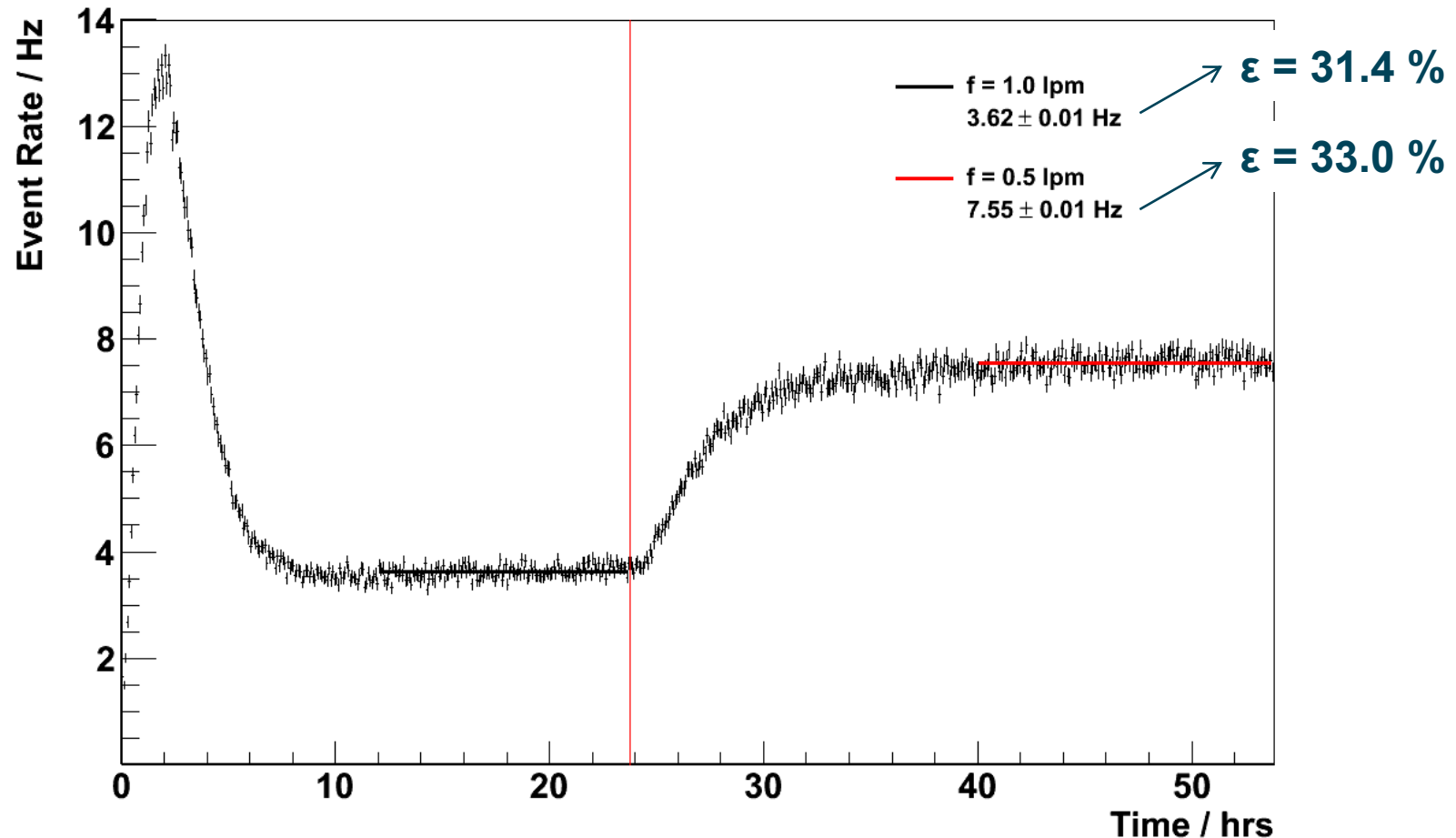
- Look at the rate of ^{214}Po events to measure detector background:



- BG Rate of detector is measured to be $8.4 \pm 0.5 \text{ cpd}$

Electrostatic Detector: Efficiency

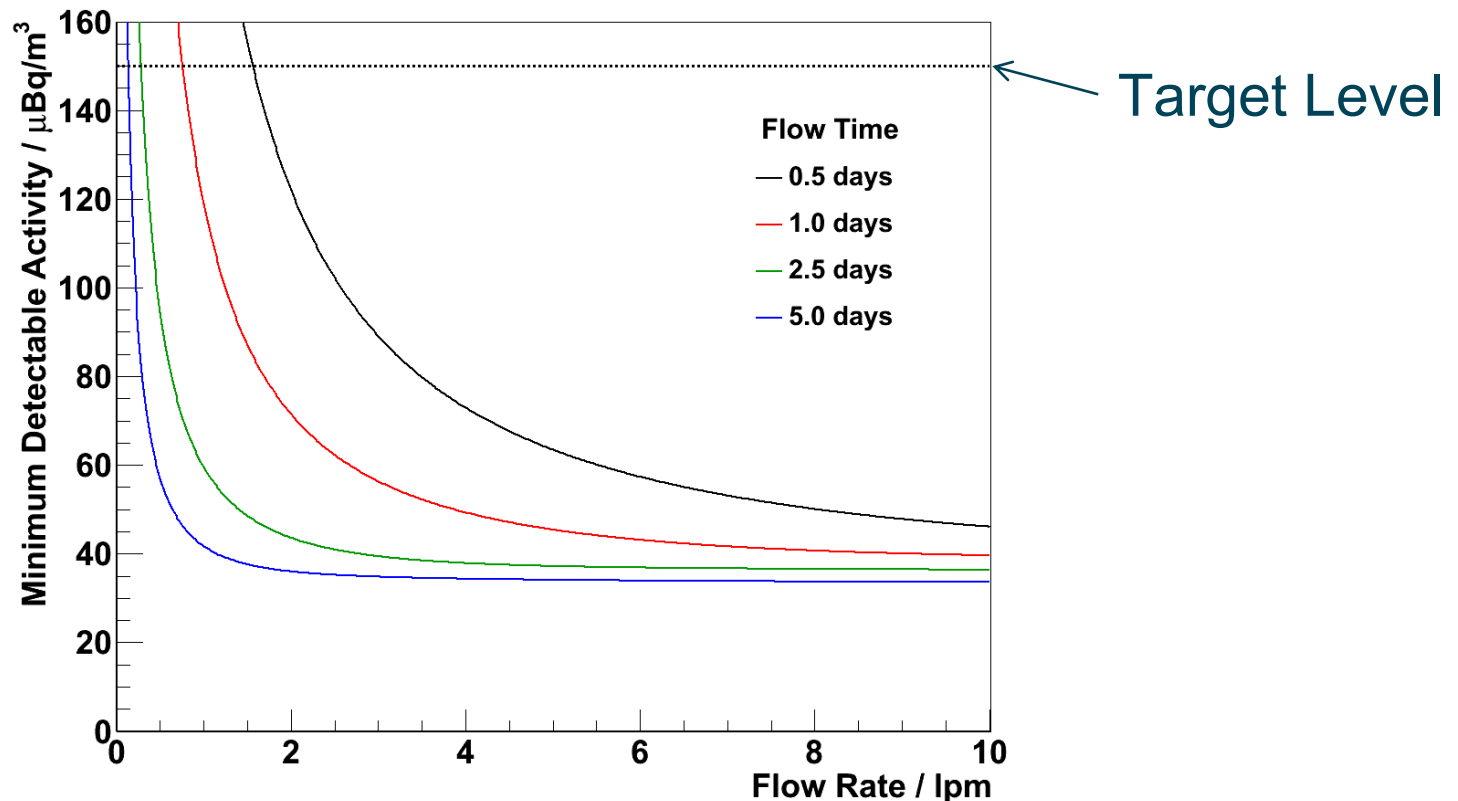
- Use a calibrated flow-through ^{222}Rn source at different flow rates to measure efficiency of ^{222}Rn detection



- Efficiency of detector is preliminarily measured to be $\sim 30\%$

RnCL & Detector: Sensitivity

- Measurement of emanation from quarter sub-section of SuperNEMO tracker:



- Measurement of gasses to select the cleanest cylinders:
 - Sensitive to $C_{Rn} \geq 20 \mu\text{Bq/m}^3$ for one cylinder

Future work

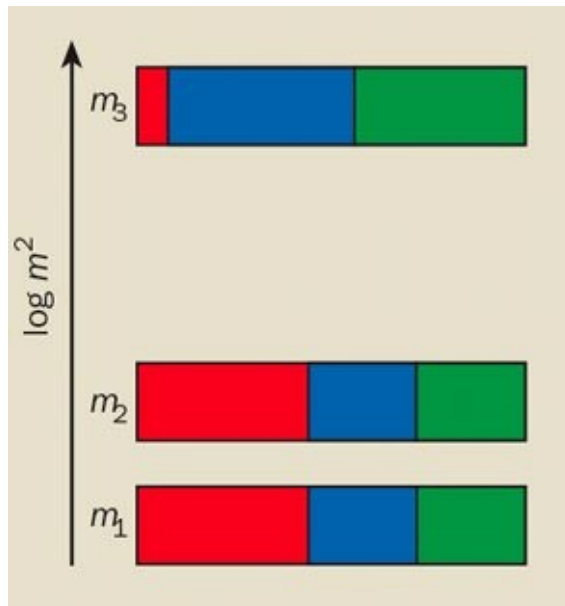
- **Soon:**
 - Finish calibration of detector and measure trapping and transfer efficiency
 - Measure N₂, He and Ar from different cylinders to find the ones most suitable for use with the SuperNEMO demonstrator
- **Longer Term:**
 - Measure the emanation of radon from detector materials in the SuperNEMO tracker before it goes underground
 - Analyse NEMO-3 data for ⁸²Se – the baseline isotope for SuperNEMO

END

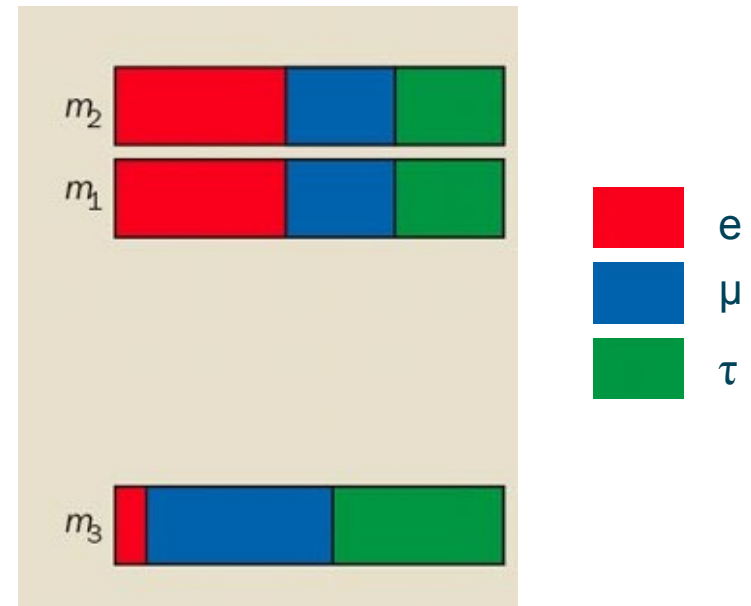
Supplementary Slides

Neutrinos: Oscillations & Mass

- Neutrino oscillation experiments have shown us that neutrinos must have non-zero mass
- Two-neutrino case :
$$P_{\alpha \rightarrow \beta, \alpha \neq \beta} = \sin^2(2\theta) \sin^2\left(\frac{\Delta m^2 L}{4E}\right)$$
- So oscillation experiments are unable to offer any information about the absolute mass scale or (with current precision) hierarchy:



OR



Neutrino Mass: Current Limits

- **$0\nu\beta\beta$:**

Heidelberg-Moscow Experiment: $\langle m_{\beta\beta} \rangle < 0.3 - 0.8 \text{eV}$

- **Tritium decay:**

Based on kinematics of end-point of decay so is model independent.

PDG : $m_{\nu_e} < 2 \text{eV}$

- **Cosmology**

Shaun Thomas: $\sum_{\nu} m_{\nu} < 0.28 \text{eV}$

0νββ vs. 2νββ: Half-life and Neutrino Mass

- The half-life for 0νββ can be related to neutrino mass by:

$$[T_{1/2}^{0\nu}]^{-1} = G^{0\nu}(Q_{\beta\beta}, Z) |M^{0\nu}|^2 \langle m_{\beta\beta} \rangle^2$$

where $G^{0\nu}$ is the (exactly calculable) phase space
 $M^{0\nu}$ is the nuclear matrix element (NME)
 $\langle m_{\beta\beta} \rangle$ is a lepton number violating parameter

- In the light-neutrino exchange mechanism, $\langle m_{\beta\beta} \rangle$ is the effective Majorana neutrino mass, given by:

$$\langle m_{\beta\beta} \rangle = \left| \sum_k m_k U_{ek}^2 \right|$$

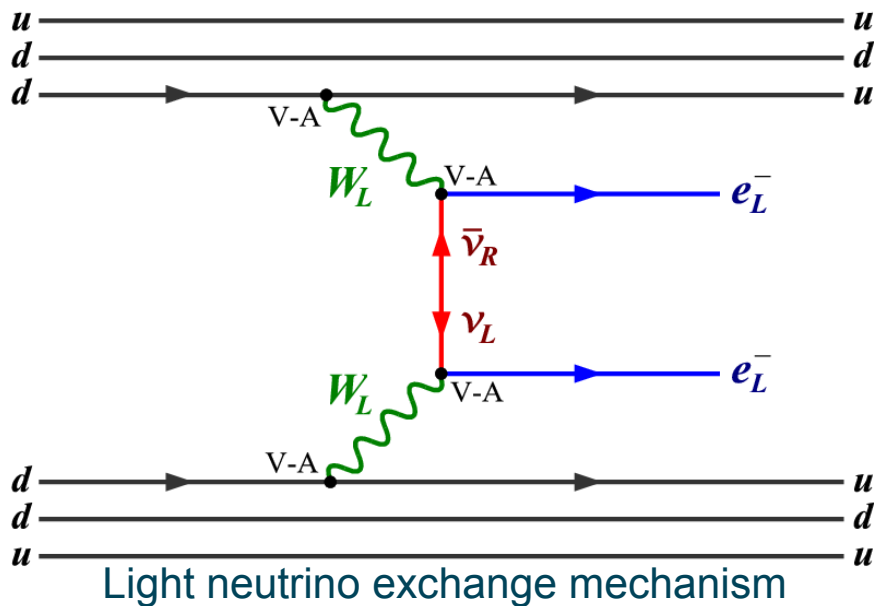
- In 2νββ, we can get no information about neutrino mass:

$$[T_{1/2}^{2\nu}]^{-1} = G^{2\nu}(Q_{\beta\beta}, Z) |M^{2\nu}|^2$$

- But NME is important for reducing the error in the 0νββ NME. 19

$0\nu\beta\beta$: Light neutrino exchange mechanism

- Can break down $0\nu\beta\beta$ into two parts:



i) $n \rightarrow p + e^- + \nu_{eR}^M$

ii) $n + \nu_{eL}^M \rightarrow p + e^-$

- Process requires that the neutrino must be its own antiparticle and must flip helicity.

- Helicity flip is possible if we consider a reference frame moving faster than the neutrino
- Implies that neutrinos must be **Majorana** and have **mass**

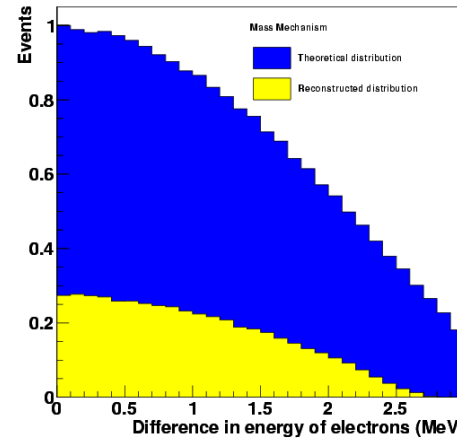
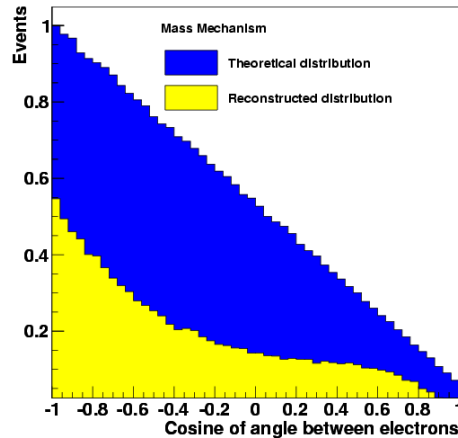
$0\nu\beta\beta$: Signatures for different mechanisms

- SuperNEMO is unique among the next generation $0\nu\beta\beta$ experiments as it may allow us to disentangle the physics mechanism, for example:

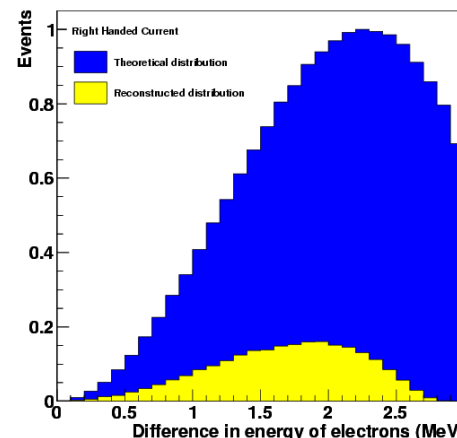
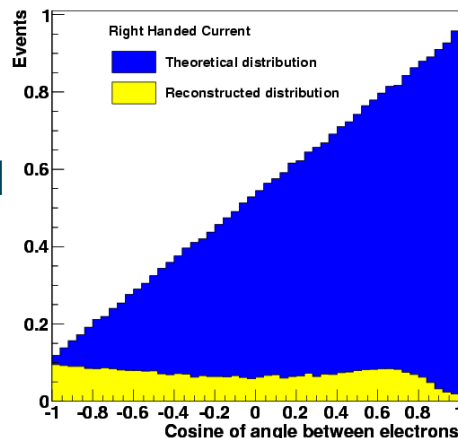
Cosine of electron opening angle

Electron energy difference

Mass Mechanism



Right Handed Current



Theoretical Distribution

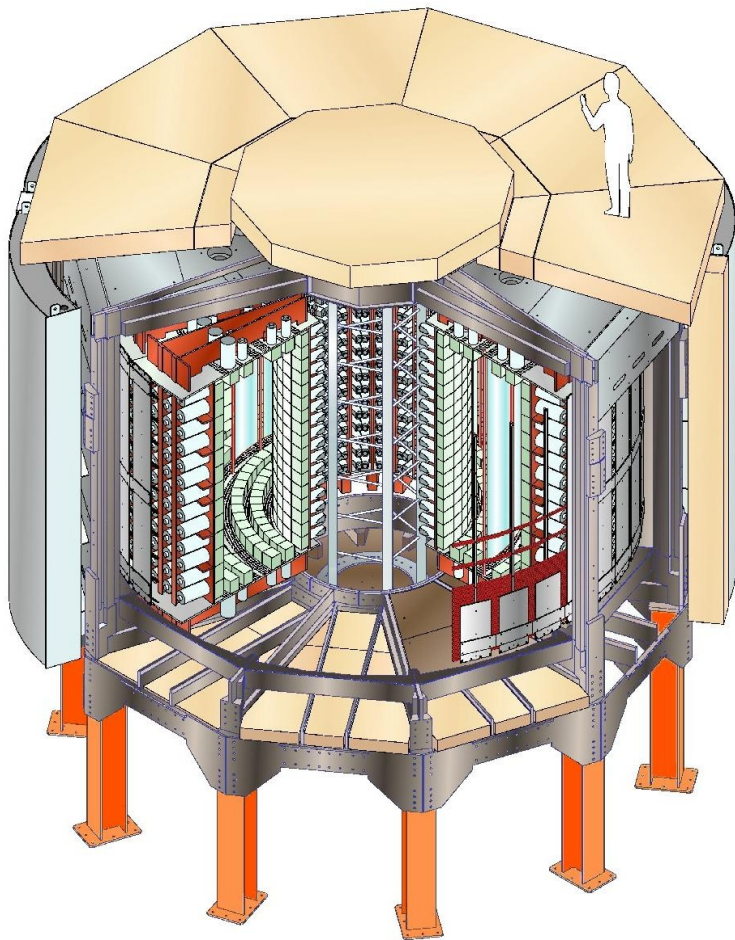
Reconstructed Distribution

$0\nu\beta\beta$: Notable Experiments

| Experiment | Isotope | Mass / kg | $\langle m_{\beta\beta} \rangle$ lower limit / meV | Current Status | Start of data taking |
|--------------------|-------------------|-----------|--|----------------|----------------------|
| SuperNEMO | ^{82}Se | 100 | 40 | R & D | ~2013 – 2015 |
| GERDA | ^{76}Ge | 40 | 70 | Running | 2011 |
| | | 1000 | 10 | R & D | ~2015 |
| CUORE | ^{130}Te | 200 | 20 | In progress | ~2013 |
| SNO+ | ^{150}Nd | 56 | 100 | In progress | ~2012 |
| | | 500 | 40 | R & D | ~2015 |
| EXO | ^{136}Xe | 200 | 100 | Running | 2011 |
| | | 1000 | 30 | R & D | ~2015 |
| KamLAND-ZEN | ^{136}Xe | 400 | 40 | In progress | ~2012 |
| | | 1000 | 25 | R & D | ~2013 – 2015 |
| MAJORANA | ^{76}Ge | 30 – 60 | 70 | In progress | ~2013 |
| | | 1000 | 10 | R & D | ~2015 |

The NEMO-3 Experiment

- NEMO-3 was the predecessor to SuperNEMO, which ran from Feb 2003 – Jan 2011.



- Cylindrical design with source foils of different $\beta\beta$ isotopes surrounded by a gas tracker and a calorimeter.
- Employed a ‘smoking-gun’ approach:
 - Particle ID, event topology reconstruction & strong background rejection
 - Compromise on energy resolution
- World’s best $T_{1/2}$ measurements of seven $2\nu\beta\beta$ isotopes (out of only 12 observed):

^{100}Mo , ^{82}Se , ^{150}Nd , ^{96}Zr , ^{48}Ca , ^{116}Cd , $^{130}_{23}\text{Te}$

The NEMO-3 Experiment

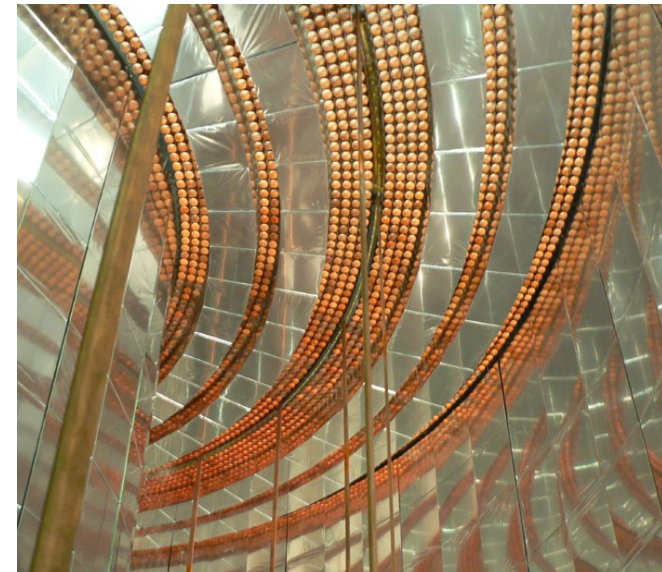


Source foil: 10kg of different $\beta\beta$ isotopes

Tracker: Drift chamber with 6180 vertical cells in He, Ar, alcohol & water.

Calorimeter: 1940 PMTs & plastic scintillator blocks

Shielding: Wood, iron & borated water to stop different external backgrounds

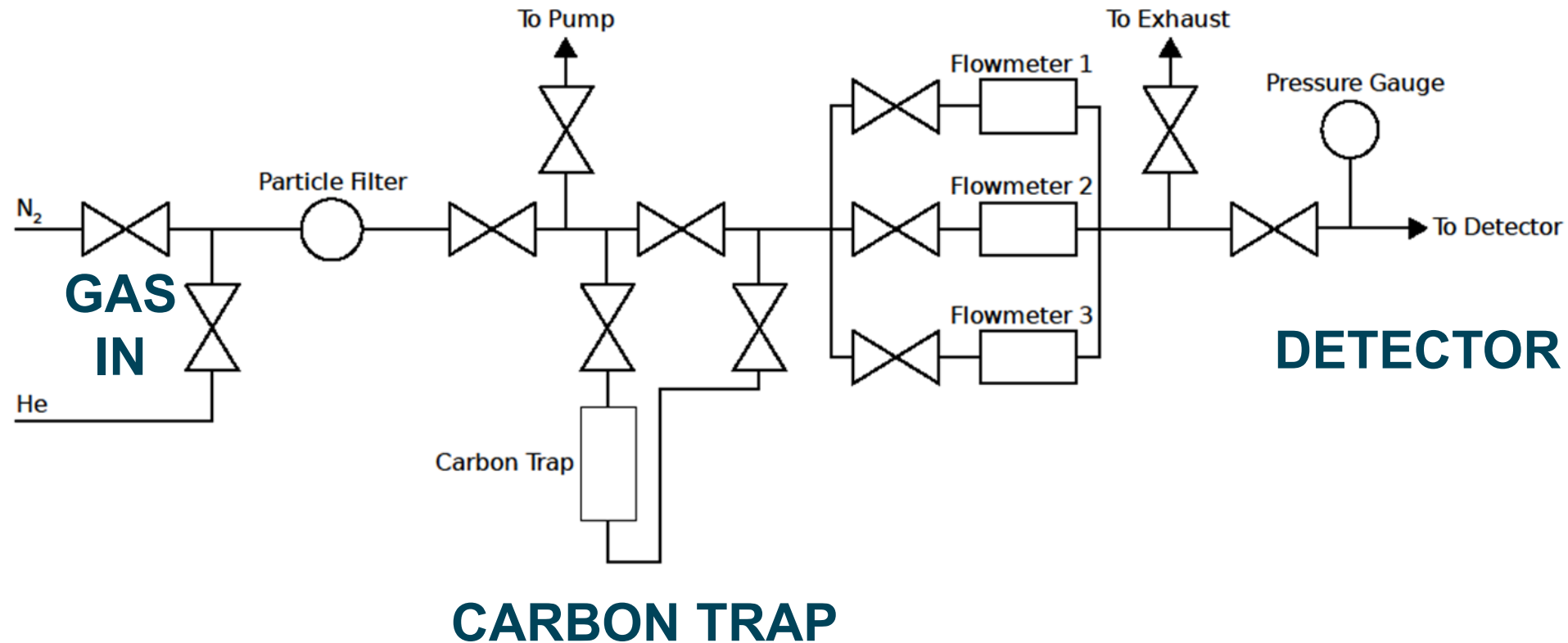


Some important measurements:

^{100}Mo : $T_{1/2}(2\nu) = [7.16 \pm 0.01(\text{stat}) \pm 0.54(\text{sys})] \times 10^{18} \text{ y}$
 $T_{1/2}(0\nu) > 1.0 \times 10^{24} \text{ y @ 90\% CL}$

^{82}Se : $T_{1/2}(2\nu) = [9.6 \pm 0.1(\text{stat}) \pm 1.0(\text{sys})] \times 10^{19} \text{ y}$
 $T_{1/2}(0\nu) > 3.2 \times 10^{23} \text{ y @ 90\% CL}$

RnCL: Schematic Diagram



RnCL: Method

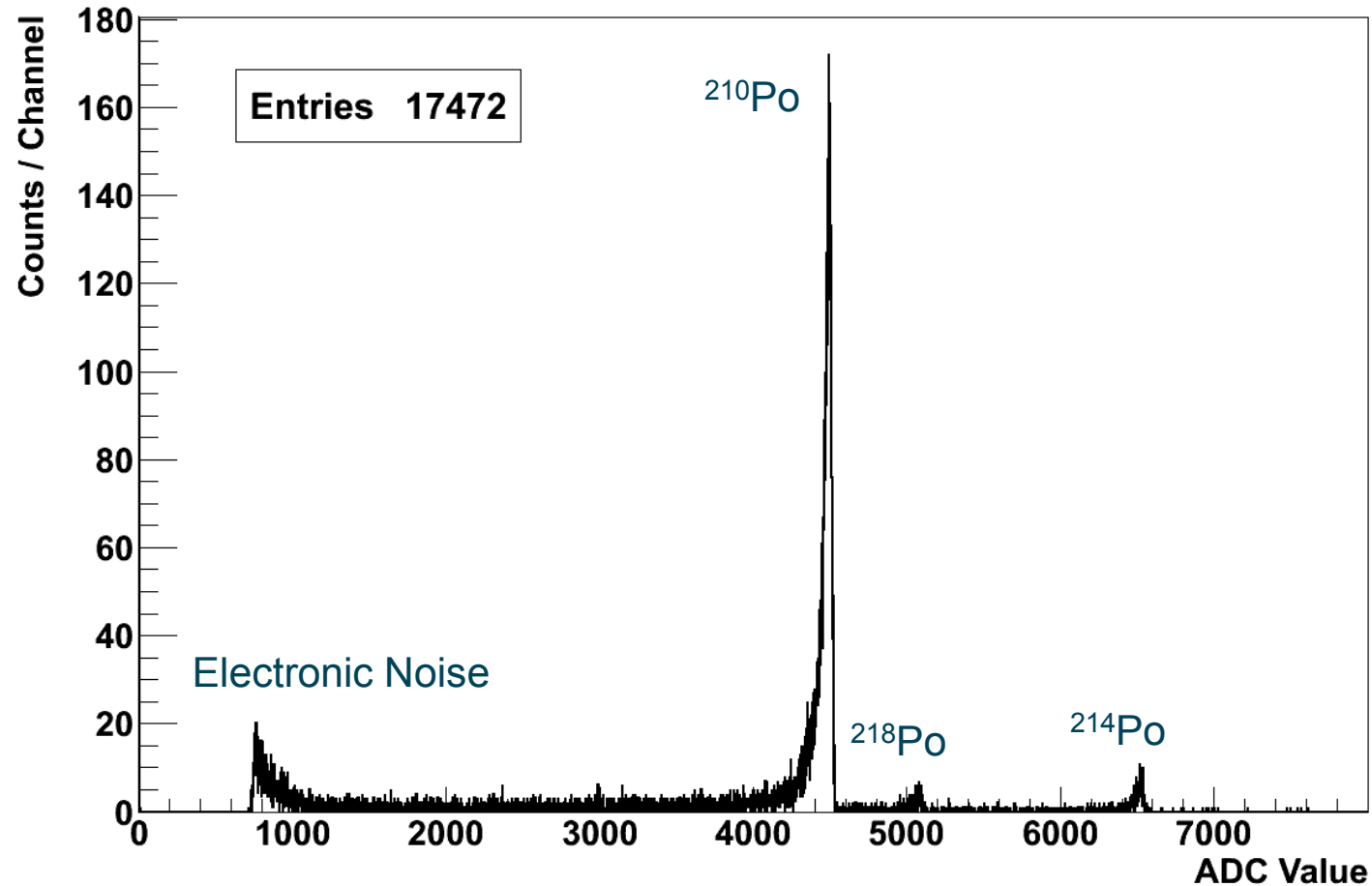
- Operation of the RnCL is as follows:
 1. Evacuate system down to 10^{-7} mbar.
 2. Cool trap down to -196 °C in Dewar containing liquid N_2 .
 3. Pump gas to be measured through carbon trap – ^{222}Rn is adsorbed.
 4. After collection of the sample, evacuate trap at -196 °C and then at -100 °C.
 5. Heat trap and transfer ^{222}Rn to detector by He purge.

RnCL: Real Life – in situ



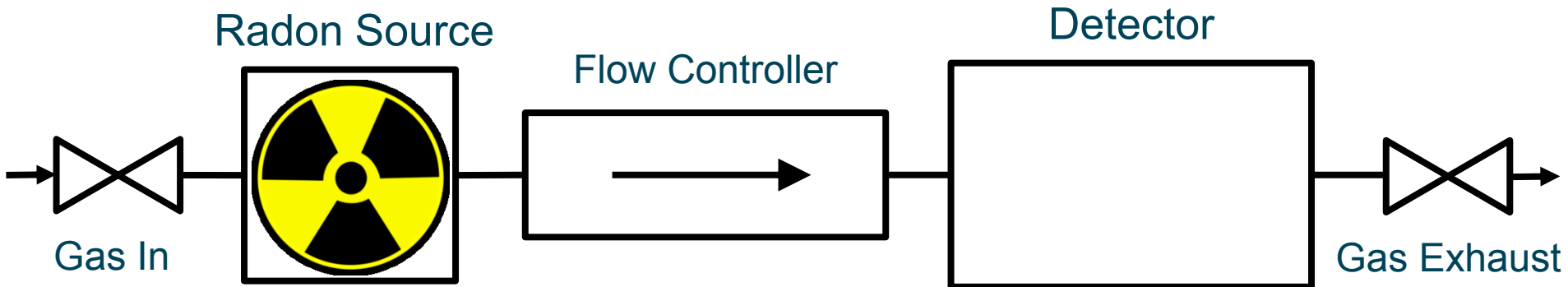
Electrostatic Detector: Background Spectrum

- Close the detector and allow ^{222}Rn to decay. Then take a long run to get background spectrum of detector:



- Peak from ^{210}Po (5.3 MeV) has now appeared.

Flow-through Calibration Method



- Flow gas at constant rate through source and detector.
- In equilibrium, activity in detector is given by:

$$A = \frac{A_0 \lambda}{\left(\lambda + \frac{f}{V_D} \right)}$$

Electrostatic Detector: Signal

- A ^{226}Ra source was attached to the detector and alpha events were seen:

