The angular analysis of $B^0 \rightarrow K^{*0} \mu^+ \mu^-$ at LHCb

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 $B^0 \rightarrow K^{*0} \mu^+ \mu^-$

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Outline

Introduction



- 3 Acceptance Correction
 - Result extraction
- 5 The zero crossing point of $A_{\rm FB}$
- 6 Systematics



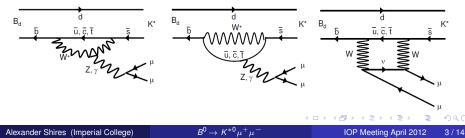
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$B^0 ightarrow K^{*0} \mu^+ \mu^-$: basics

- $B^0 \rightarrow K^{*0} \mu^+ \mu^-$ is a flavour-changing neutral current decay
- The $b \rightarrow s$ transition proceeds via a loop diagram.
- New physics can enter inside the loop.
- Branching fraction is well measured at 1.05^{+0.16}_{-0.13}x10⁻⁶, close to the SM value.
- However, smaller effects subtly affect the angular distribution of the decay products.



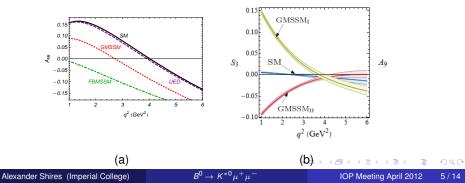


$B^0 \rightarrow K^{*0} \mu^+ \mu^-$: angular distribution

- The decay $B^0 \to K^{*0} \mu^+ \mu^-$ can be described by four kinematic variables: θ_ℓ , θ_K , ϕ and q^2 .
- The angles are defined with respect to the daughter kinematics in the *B*⁰ rest frame and form the basis for the angular distribution.
- The angular distribution can expressed in terms of theoretically clean observables (for example *A*_{FB}, *F*_L, *S*₃).
- These observables consist of combinations of the transversity (or heliticty) amplitudes which are sums of the Wilson coefficients $C_{7.}^{(eff)} C_{9}^{(eff)}$ and $C_{10.}$
- These Wilson coefficients parameterise the contributions from the penguin diagrams.

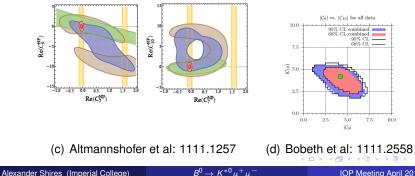
$B^0 ightarrow K^{*0} \mu^+ \mu^-$: Motivation

- b → s transitions have been explored to determine the effects of new physics models.
- Contributions from the effects of new physics can be encoded in the Wilson coefficients (Altmannshofer et al: 0811.1214)
- This leads to calculations made of the observables for different scenarios:



Current theoretical constraints

- The LHCb result in 2011 along with other b → s transition measurements have been used to provide experimental input on model independant fits to the Wilson coefficients: (a) & (b).
- The results from (a) give a 95% best fit compatibility with the standard model and the subsequent constraints give the scale of new physics in this system (Λ_{NP}) from 5 to 40 TeV.



$B^0 ightarrow K^{*0} \mu^+ \mu^-$ event selection

- Cut based preselection with an MVA (BDT) offline selection.
- Select on daughter impact parameter, B vertex, particle identification, track quality
- There are minimal cuts on p_T to retain angular sensitivity.
- Packing backgrounds and mis-identified candidates vetoed.

•
$$B_s^0 \rightarrow \phi \mu^+ \mu^-$$
,
• $B_s^0 \rightarrow \overline{K}^{*0} \ \mu^+ \mu^-$,
• $\pi \rightarrow \mu$ swaps etc

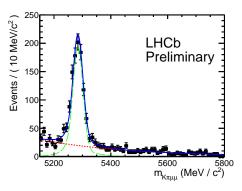


Figure: 900 \pm 34 $B^0 \rightarrow K^{*0} \mu^+ \mu^-$ candidates.

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- The effects of the detector acceptance can bias the angular distributions.
- Correct on an event by event basis.
- Efficiency calculated using phase space simulated events
 model independant
- Simulation corrected for known differences with data
- Assume efficiency factorises in phase space : $\epsilon(\cos \theta_l, \cos \theta_K, \phi, q^2) =$ $\epsilon_{\cos \theta_l}(q^2)\epsilon_{\cos \theta_K}(q^2)\epsilon_{\phi}(q^2)$

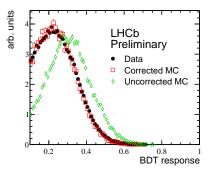


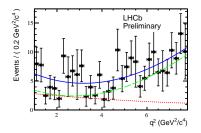
Figure: MVA distribution of $B^0 \rightarrow J/\psi$ K^{*0} events showing the impact of data/MC corrections. • Fit angular distribution :

$$\frac{1}{\Gamma} \frac{\mathrm{d}^4 \Gamma}{\mathrm{d}\cos\theta_\ell \,\mathrm{d}\cos\theta_K \,\mathrm{d}\hat{\phi} \,\mathrm{d}q^2} = \frac{9}{16\pi} \left[F_L \cos^2\theta_K + \frac{3}{4}(1 - F_L)(1 - \cos^2\theta_K) + F_L \cos^2\theta_K (2\cos^2\theta_\ell - 1) + \frac{1}{4}(1 - F_L)(1 - \cos^2\theta_K)(2\cos^2\theta_\ell - 1) + \frac{3}{4}(1 - F_L)(1 - \cos^2\theta_K)(2\cos^2\theta_\ell - 1) + \frac{3}{4}(1 - \cos^2\theta_K)(1 - \cos^2\theta_\ell)\cos2\hat{\phi} + \frac{4}{3}A_{FB}(1 - \cos^2\theta_K)\cos\theta_\ell + \frac{4}{3}A_{FB}(1 - \cos^2\theta_K)(1 - \cos^2\theta_\ell)\sin2\hat{\phi} \right]$$

- This is a distribution is simplified using the symmetry : $\phi \rightarrow \phi + \pi$
- 3D phase space ($\cos \theta_K, \cos \theta_K, \phi$) with 4 free parameters (A_{FB}, F_L, S_3, A_{Im})
- Observables are correlated non-trivially via the K^{*0} polarisation amplitudes

The zero crossing point of $A_{\rm FB}$

- A_{FB} changes sign at a well defined point in the SM.
- Uncertainty on form factors roughly cancels when $A_{FB}(q^2) = 0$ and the SM prediction is between 4 and 4.3.
- Zero-crossing point extracted using 'unbinned-counting' technique
- Fit 'forward-going' and 'backward-going' events seperately,



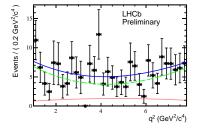


Figure: Forward-going events

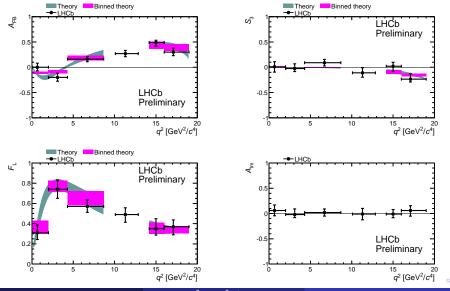
Figure: Backward-going events

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For the determination of $A_{\rm FB}$, $F_{\rm L}$, S_3 and $A_{\rm Im}$ the dominant sources of systematic uncertainty are

- uncertainty on the acceptance correction
- contains uncertainties from the data-simulation corrections: tracking, trigger, PID correction
 - This is explored by varying event weights coherently and refitting the angular distribution
- Second main source of systematic is dependance on the signal and background model
- Additional systematic uncertainties:
 - S-wave : Fit with values taken from from $B^0 \rightarrow J/\psi \ K^{*0}$ at BABAR and LHCb (8%)
 - B⁰ mis ID fraction.

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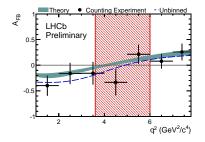
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Results

- The zero crossing point is measured to be $q_0^2 = 4.9^{+1.1}_{-1.3} \text{ GeV}^2$
- The measured range of the crossing point is shown in red below.



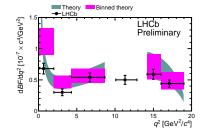


Figure: The zero crossing point showing a comparison of a counting experiment and the unbinned analysis

Figure: The differential branching ratio of ${\cal B}^0 \! \to {\cal K}^{*0} \mu^+ \mu^-$

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- $B^0 \rightarrow K^{*0} \mu^+ \mu^-$ is a promising channel to discover new physics in the flavour sector.
- However, the analysis of 341 pb⁻¹ of LHCb data in 2011 set significant constraints on the scale of new physics
- Angular analysis of 1 fb⁻¹ of LHCb data is the most precise measurement of A_{FB} , F_L , A_{Im} and S_3 yet.
- First measurement of the zero-crossing point of A_{FB}.
- LHCB-CONF-2012-008