

**The discovery reach for  
mini-black holes with  
the ATLAS Detector at the LHC**

Michiru Kaneda

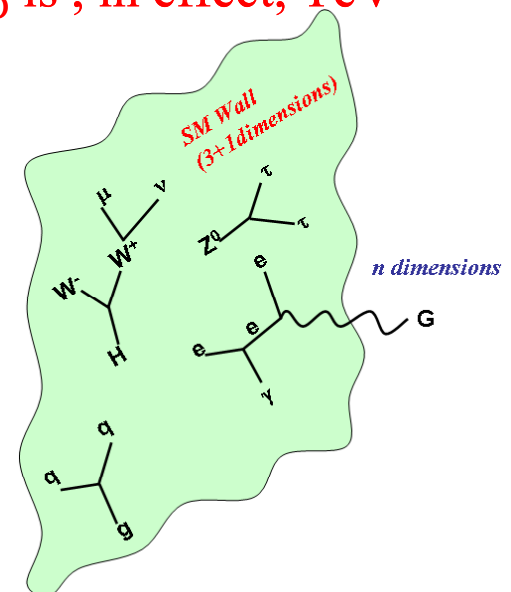
*University of Tokyo*

On behalf of the ATLAS Collaboration



# TeV-Scale Gravity

- Hierarchy problem, one of the big unsolved problems in 20 century
  - The large deviation between electroweak scale and the Planck scale:  
 $M_{Pl}(10^{19}\text{GeV}) \gg M_W(10^2\text{GeV})$
- Extra dimensions
  - One of the solutions of the hierarchy problem
  - Only gravitational field is allowed to expand into the extra dimensions
  - “The Fundamental” d-dimensional Planck scale,  $M_D$  is , in effect, TeV scale
- Some approach:
  - Additional large flat dimensions
    - Arkani-Hamed, Dimopoulos and Dvali (ADD)
  - A single warped extra dimension
    - Randall and Sundrum (RS)



# Black Hole

## Production

- If the particles (with center of mass energy =  $\hat{s}$ ) collide with the impact parameter  $b$  which is smaller than 2 times Schwarzschild radius,  $r_h \sim \hat{s}/M_D^2$ , Black Hole will be formed

➤ If  $M_D \sim \text{TeV}$ , LHC can generate TeV-scale mini-black hole!



## Decay

- Such a mini-black hole decays in  $\sim 10^{-26} \text{s}$

- There are 4 decay phase

- The balding phase: Loose the “hair” (multipole moment)
- Spin-down phase: Loose angular momentum by emitting high-spin state particles
- Schwarzschild phase: Hawking evaporation
- Planck phase:  $M_{BH} \sim M_D$ , need quantum gravity, difficult to calculate

The smaller  $M_{BH}$  becomes,  
The higher Hawking Temperature becomes

$$T_H = M_D \left[ \frac{M_D}{M_{BH}} \left( \frac{n+2}{8\Gamma((n+3)/2)} \right) \right]^{1+n}$$

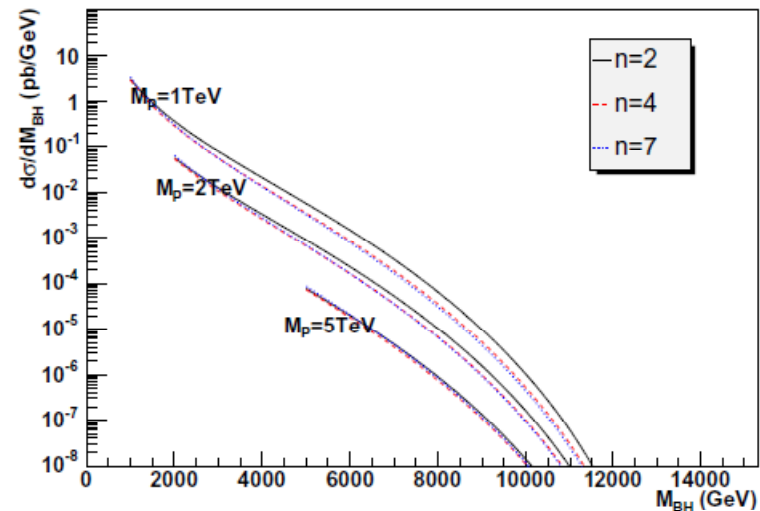
# Working Model

## Black Hole Event Generator: CHARYBDIS:

- Based on ADD model
- Parton level xsec is calculated with assumption of semi-classical model (valid only when  $M_{BH} \gg M_D$ )

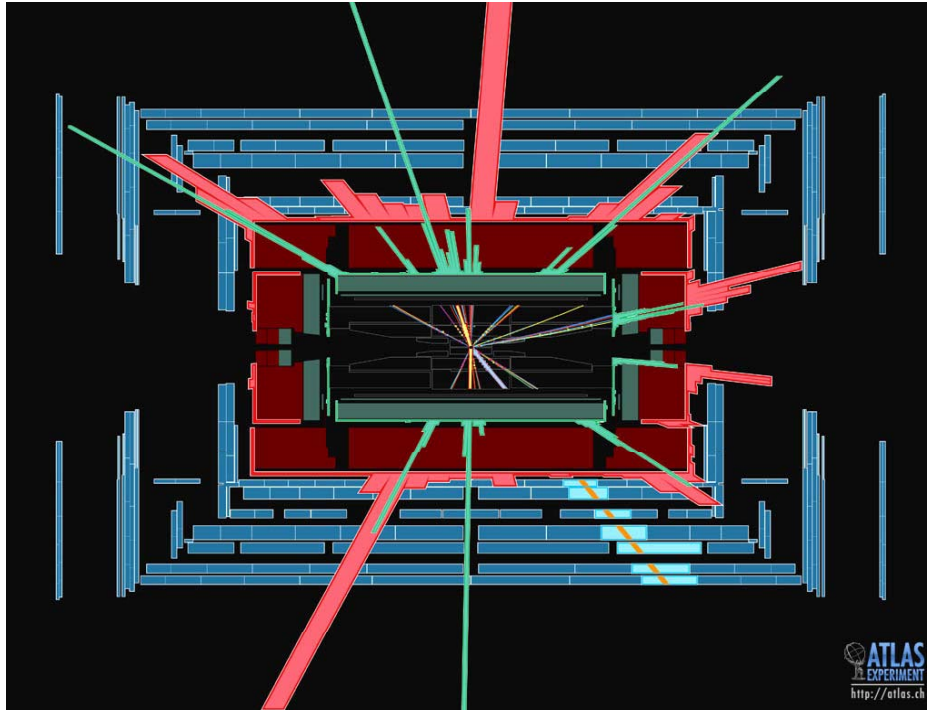
$$\hat{\sigma}_{ab \rightarrow BH} = \pi r_h^2 \quad r_h = \frac{1}{\sqrt{\pi} M_D} \left[ \frac{M_{BH}}{M_D} \left( \frac{8\Gamma((n+3)/2)}{n+2} \right) \right]^{\frac{1}{1+n}}$$

- Ignore balding and spin-down phase
- No graviton emission
- Available grey-body factor
- Just a N-body decay at Plank Phase
- Main input parameters:
  - The Fundamental Plank Scale:  $M_D$
  - Number of extra dimensions:  $n$
  - Minimum BH Mass to be produced ( $M_{BH} \gg M_D$ )



Cross section of Black Hole in 14TeV pp collision

# Black Hole Event



ATLAS Event display: Simulation of Black Hole event

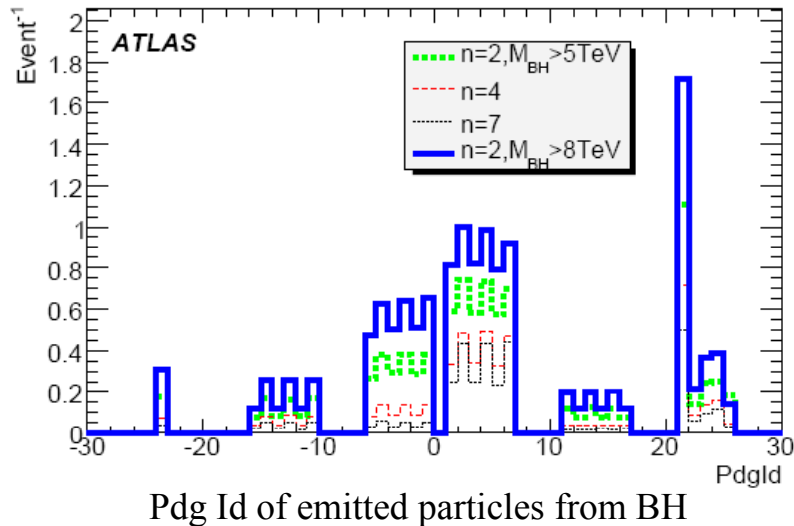
- Large cross section ( $\sim 1$  event/s for TeV BH)
- Very crowded events
- High energy particles
- Most of events have leptons
- Easy to trigger

a) CHARYBDIS:  $n = 2, m > 5$  TeV

| Trigger | L1    | L2    | EF    |
|---------|-------|-------|-------|
| j100    | 1     | 1     | 1     |
| j400    | 0.997 | 0.997 | 0.997 |
| 3j100   | 0.998 | 0.998 | 0.998 |
| 3j250   | 0.972 | 0.971 | 0.971 |
| 4j100   | 0.985 | 0.985 | 0.985 |
| 4j250   | 0.865 | 0.862 | 0.862 |

Efficiency for jet trigger

# Decay Particles from Black Hole



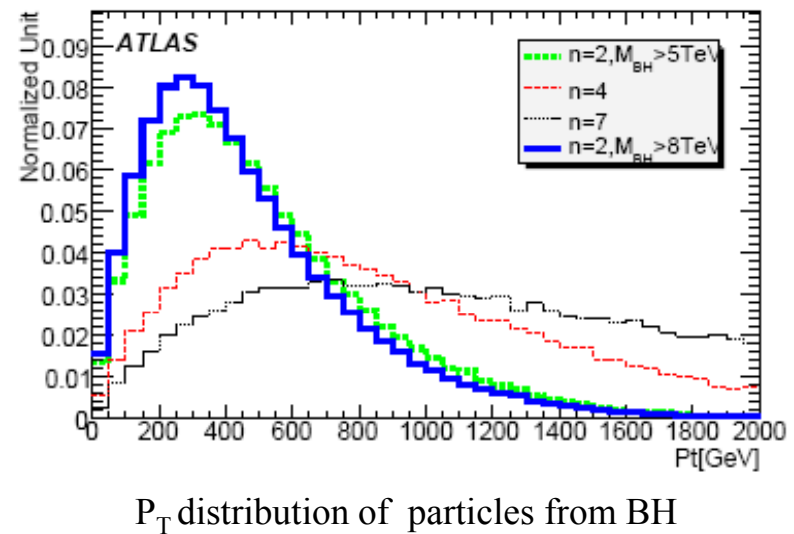
- Emit very high  $P_T$  particles
- Higher  $n$  shows higher  $P_T$  because Hawking temperature is higher

$$\frac{dN}{dE} \propto \frac{(E/T_H)^2}{\exp(E/T_H) + c}$$

$\leftarrow$  Hawking Temperature  
 $\leftarrow$  =-1 for boson(fermion)

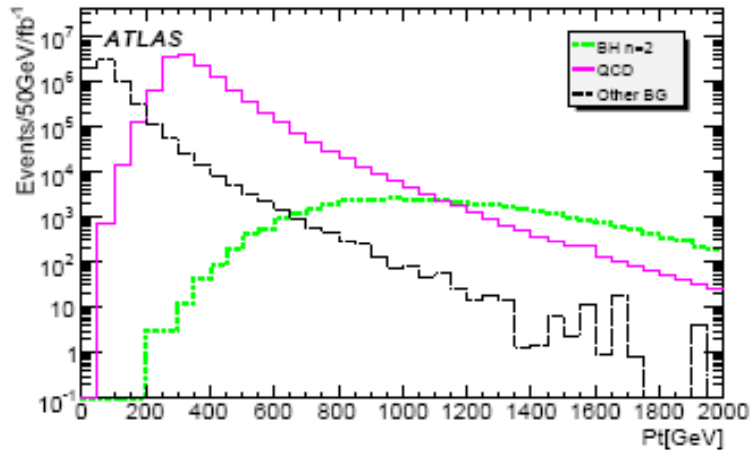
$$T_H = M_P \left[ \frac{M_D}{M_{BH}} \left( \frac{n+2}{8\Gamma((n+3)/2)} \right) \right]^{\frac{2}{1+n}}$$

- Particles are emitted by Hawking radiation
- Emission probability depends on degree of freedom of quantum variables
- A break of perfect democratic decay comes from conservation of charge, color, etc.

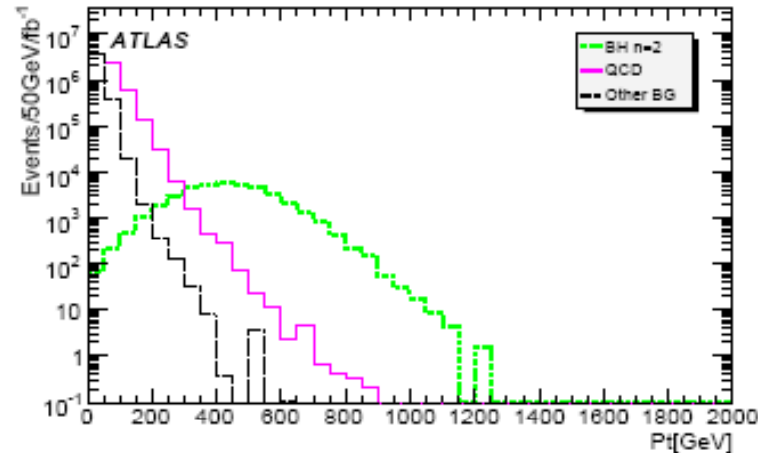


$P_T$  distribution of particles from BH

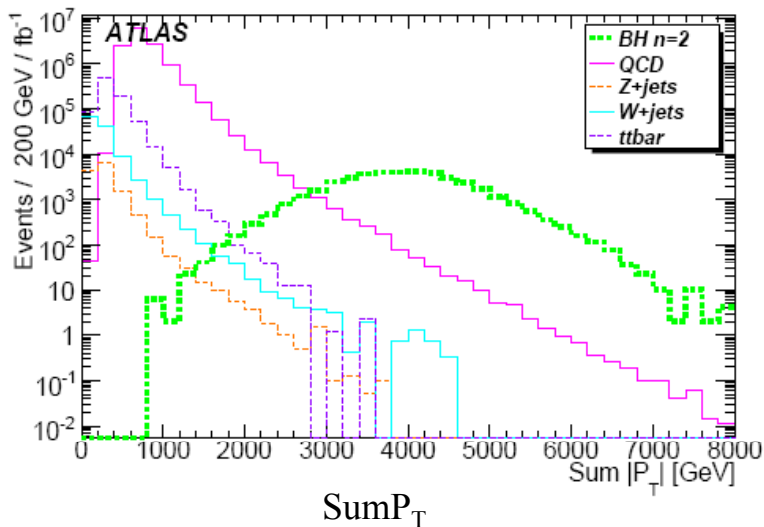
# $P_T$ of Particles observed in Events



$P_T$  distribution of leading particle

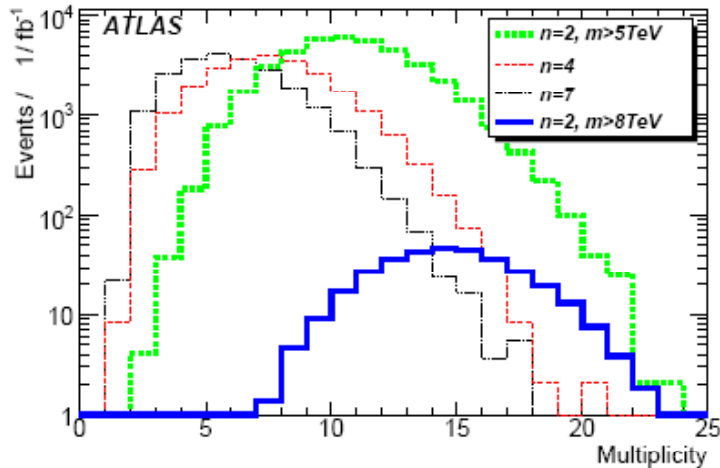


$P_T$  distribution of 4th leading particle



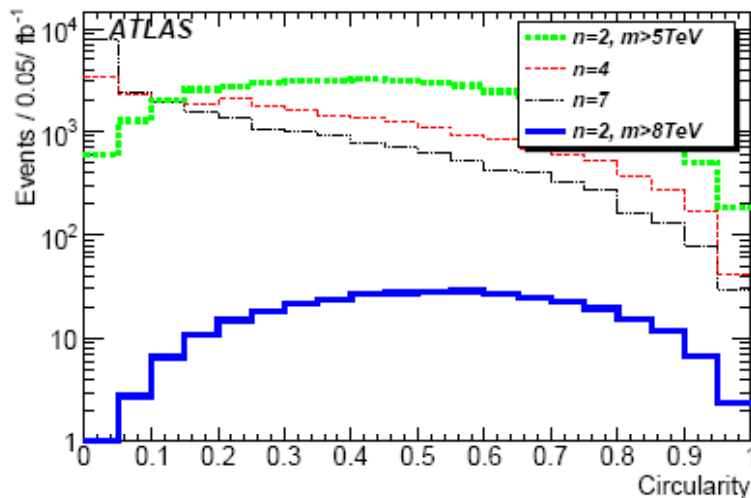
- Even 4th leading particle has very large  $P_T$
- Sum  $P_T$ :
  - A scalar sum of  $P_T$  of all particles in the event

# Multiplicity and Circularity

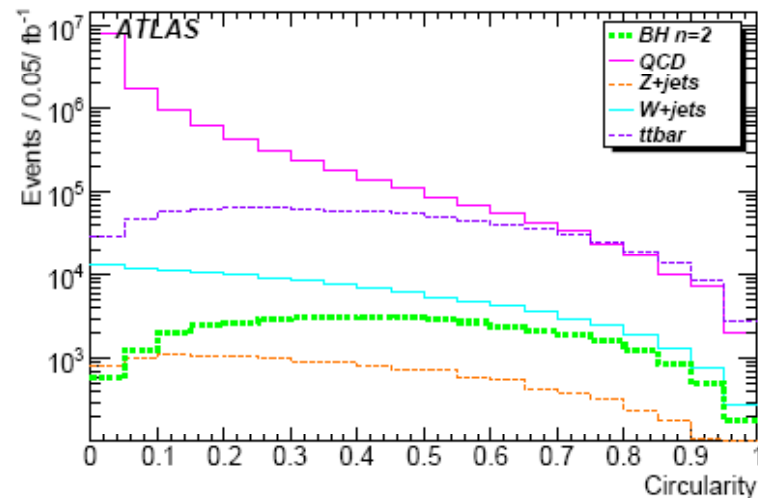


Particle Multiplicity of BH with different parameters

- Higher  $n$  shows lower multiplicity
  - Hawking temperature is higher
  - One particle carries larger energy
- Circularity, Sphericity or Thrust are also the candidates of event selections
  - But they strongly depend on multiplicity



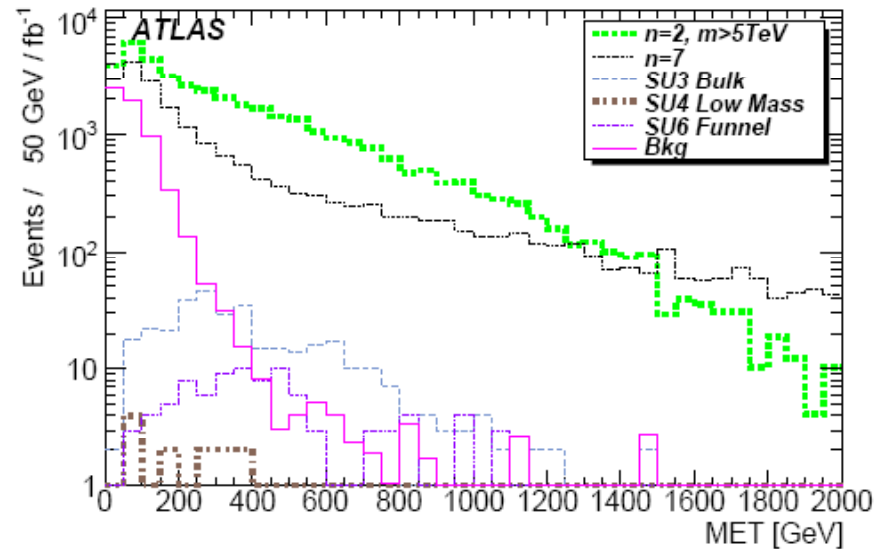
Circularity of Black Holes



Circularity, compared with backgrounds



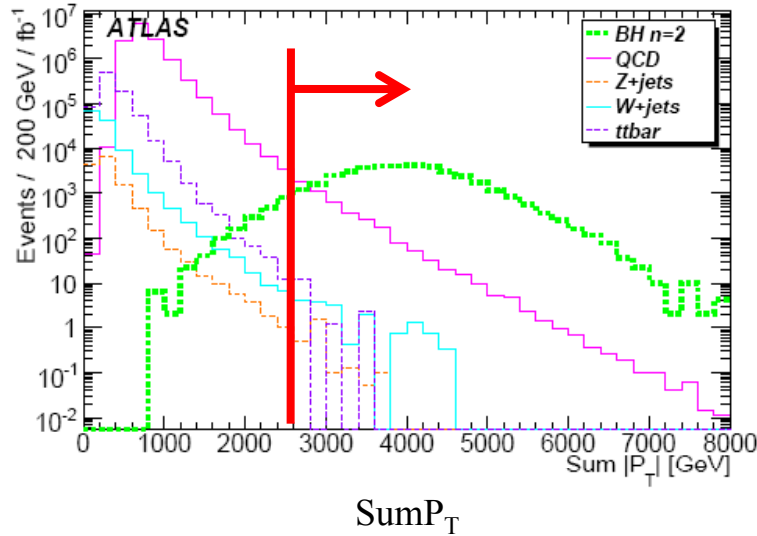
# Missing $E_T$



Missing  $E_T$  (with event selection:  $\text{Sum}P_T > 2500\text{GeV}$ )

- Black Hole emits also high  $P_T$  neutrinos
  - Large missing  $E_T$  source
- Even compared to SUSY signal, BH has a long tail in high MET region
- Charybdis can not emit graviton and not emit into bulk
  - They may be also large missing  $E_T$  source
  - New generator which can treat such features are being investigating

# Event Selection 1: Sum $P_T$

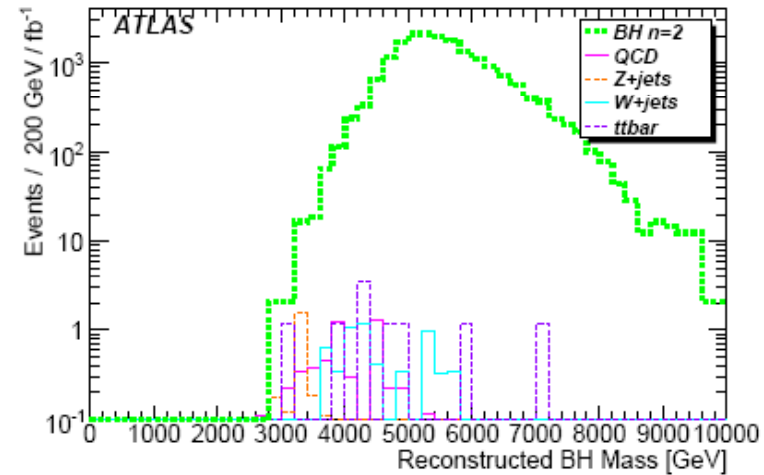


- Sum  $P_T$  selection
  - Sum  $P_T > 2.5\text{TeV}$
  - Require at least 1 of the 50GeV lepton
- **Not dependent on particle multiplicity**
- Black Hole Mass is reconstructed from particles (only high  $P_T$ ,  $>15\text{GeV}$  for  $e, \mu, \gamma$  and  $>20\text{GeV}$  for jet) in the event
- Missing ET is also included as a particle of  $P_Z=0$

Event Efficiency: Sum  $P_T$  selection

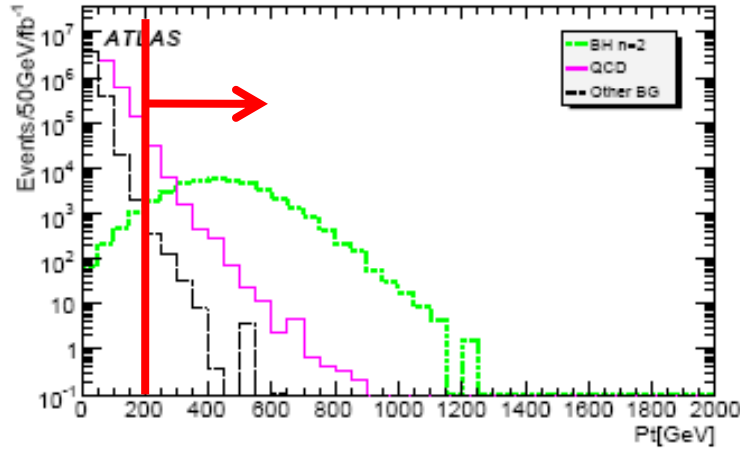
| Dataset                          | Before selection (fb)      | $\sum  p_T  > 2.5\text{TeV}$ (fb) | After requiring a lepton (fb) | acceptance           |
|----------------------------------|----------------------------|-----------------------------------|-------------------------------|----------------------|
| $n = 2, m > 5\text{ TeV}$        | $40.7 \pm 0.1 \times 10^3$ | $39.2 \pm 0.3 \times 10^3$        | $18.6 \pm 0.2 \times 10^3$    | 0.46                 |
| $n = 4, m > 5\text{ TeV}$        | $24.3 \pm 0.1 \times 10^3$ | $22.6 \pm 0.2 \times 10^3$        | $6668 \pm 83$                 | 0.27                 |
| $n = 7, m > 5\text{ TeV}$        | $22.3 \pm 0.1 \times 10^3$ | $20.1 \pm 0.2 \times 10^3$        | $3574 \pm 60$                 | 0.17                 |
| $n = 2, m > 8\text{ TeV}$        | $338.2 \pm 1$              | $338.1 \pm 2.5$                   | $212 \pm 16$                  | 0.63                 |
| $t\bar{t}$                       | $833 \pm 100 \times 10^3$  | $23.6^{+12.2}_{-6.7}$             | $8.2^{+2.43}_{-2.43}$         | $9.8 \times 10^{-6}$ |
| QCD dijets                       | $12.8 \pm 3.7 \times 10^6$ | $5899^{+1773}_{-1771}$            | $5.37^{+3.25}_{-2.02}$        | $4.3 \times 10^{-7}$ |
| $W\ell\nu + \geq 2\text{ jets}$  | $1.9 \pm 0.04 \times 10^6$ | $12.3^{+9.0}_{-1.8}$              | $4.67^{+8.75}_{-0.93}$        | $2.4 \times 10^{-6}$ |
| $Z\ell\ell + \geq 3\text{ jets}$ | $51.8 \pm 1 \times 10^3$   | $2.75^{+2.02}_{-2.01}$            | $2.57^{+0.95}_{-0.64}$        | $5.0 \times 10^{-5}$ |

Table 6: Acceptance for each signal and background dataset in fb after requiring  $\sum |p_T| > 2.5\text{TeV}$ , and a lepton with  $p_T > 50\text{GeV}$ .



Reconstructed Black Hole Mass

# Event Selection 2: Multi Object



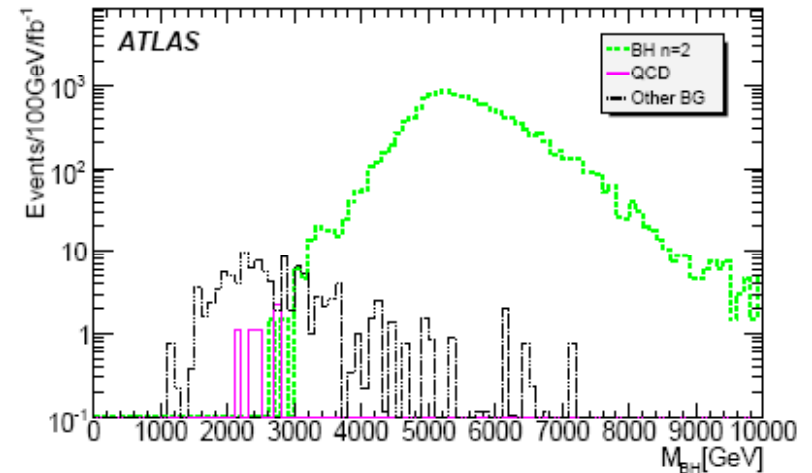
$P_T$  distribution of 4th leading particle

- Multi Object selection
  - Require 4 Objects:  $P_T > 200 \text{ GeV}$
  - Including at least 1 lepton
- Assume high multiplicity
- Less efficiency for high  $n$

Event Efficiency: Multi Object selection

| Dataset                    | Before selection<br>(fb) | After multi-object<br>requirement (fb) | After lepton requirement<br>(fb) | Acceptance           |
|----------------------------|--------------------------|--|----------------------------------|----------------------|
| $n = 2, m > 5 \text{ TeV}$ | $40.7 \times 10^3$       | $38.9 \pm 0.4 \times 10^3$             | $14.0 \pm 0.2 \times 10^3$       | 0.34                 |
| $n = 4, m > 5 \text{ TeV}$ | $24.3 \times 10^3$       | $17.9 \pm 0.3 \times 10^3$             | $4521 \pm 126$                   | 0.19                 |
| $n = 7, m > 5 \text{ TeV}$ | $22.3 \times 10^3$       | $9953 \pm 185$                         | $1956 \pm 82$                    | 0.087                |
| $n = 2, m > 8 \text{ TeV}$ | 338                      | $338 \pm 4$                            | $164 \pm 3$                      | 0.49                 |
| $t\bar{t}$                 | $833 \times 10^3$        | $129 \pm 27$                           | $36^{+12}_{-9}$                  | $4.3 \times 10^{-5}$ |
| QCD dijets                 | $12.8 \times 10^6$       | $38.9 \pm 1.9 \times 10^3$             | $6^{+107}_{-3}$                  | $5.6 \times 10^{-7}$ |
| W+jets                     | $560 \times 10^3$        | $99^{+28}_{-22}$                       | $56^{+24}_{-13}$                 | $1 \times 10^{-3}$   |
| Z+jets                     | $51.8 \times 10^3$       | $29^{+90}_{-4}$                        | $19^{+90}_{-3}$                  | $4 \times 10^{-4}$   |
| $\gamma(\gamma)$ +jets     | $5.1 \times 10^6$        | $285^{+87}_{-76}$                      | $0^{+40}_{-0}$                   | $< 10^{-5}$          |

Table 7: Acceptance of the 4-object requirements for each dataset in fb. 90% confidence limits are used when no events passed the requirements.

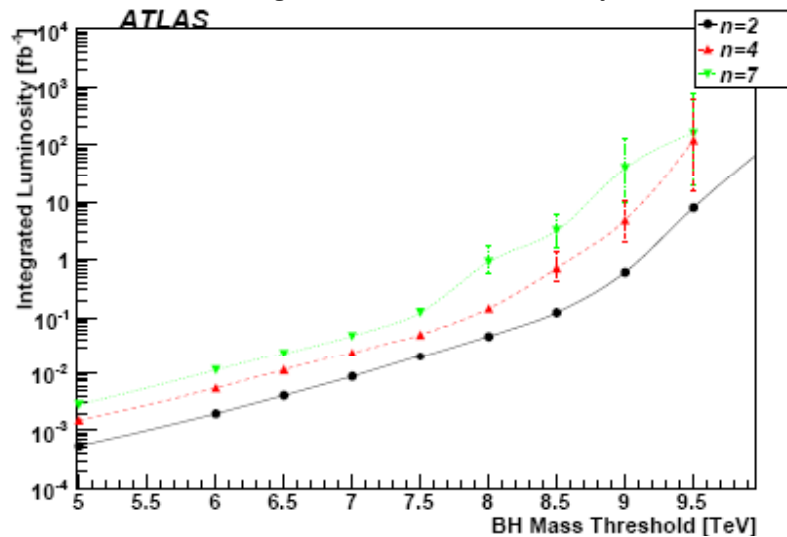


Reconstructed Black Hole Mass

# Discovery Potential for $M_p=1\text{TeV}$ Black Hole

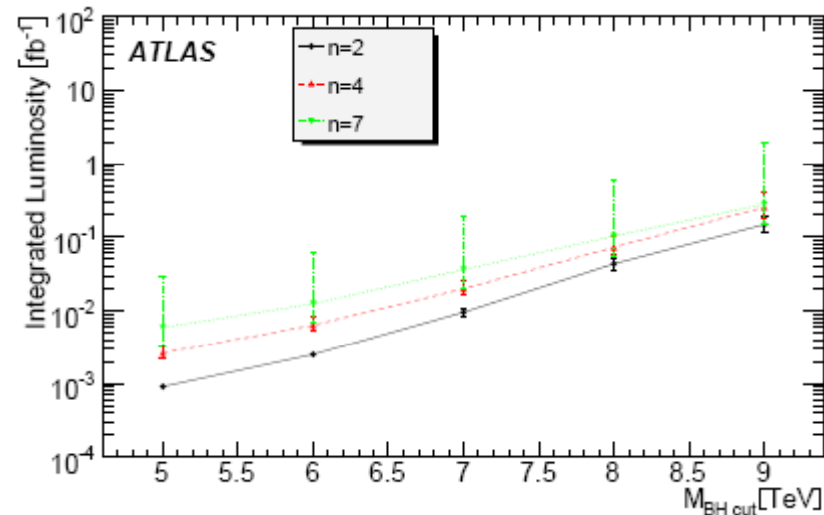
Discovery Potential:

▣ Integrated Luminosity for  $S/\sqrt{B} > 5$  &&  $S > 10$



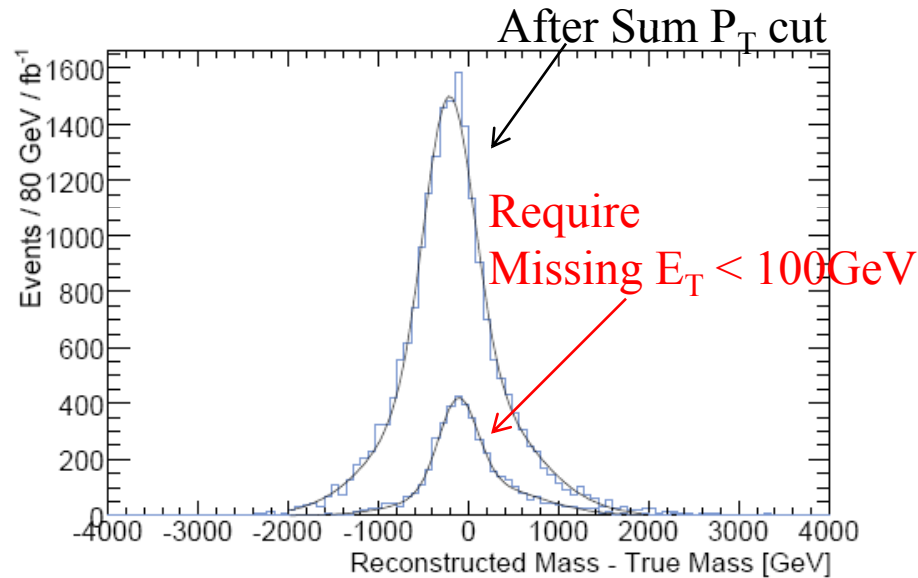
▣ Discovery Potential with Sum $P_T$  selection  
 ▣ Horizontal axis shows production threshold of  $M_{BH}$

- ▣ There is large uncertainty when  $M_{BH}$  is close to  $M_p$ 
  - ▣ Our assumption (semi-classical calculation) is not reliable
- ▣ We set minimum  $M_{BH}$  at 5TeV, and above two methods have been studied for calculation discovery potential
- ▣ **Only a few  $\text{pb}^{-1}$  is needed for 5TeV discovery**



▣ Discovery Potential with Multi Object Selection  
 ▣ Horizontal axis shows additional cut on reconstructed  $M_{BH}$

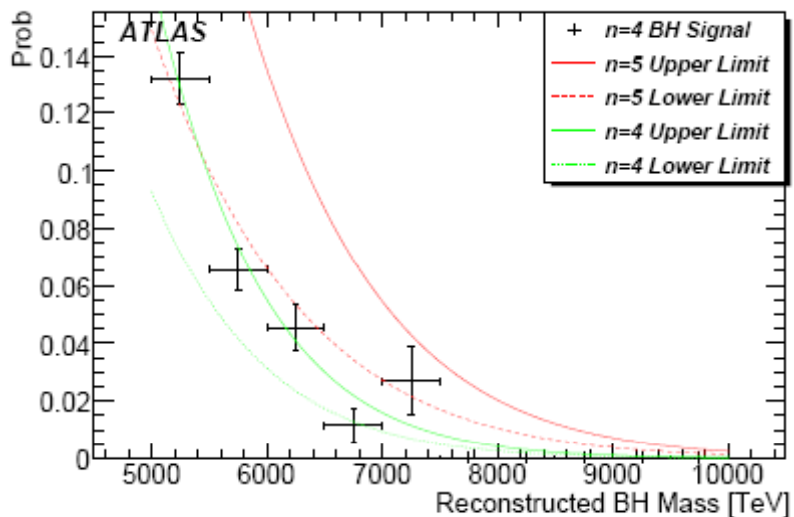
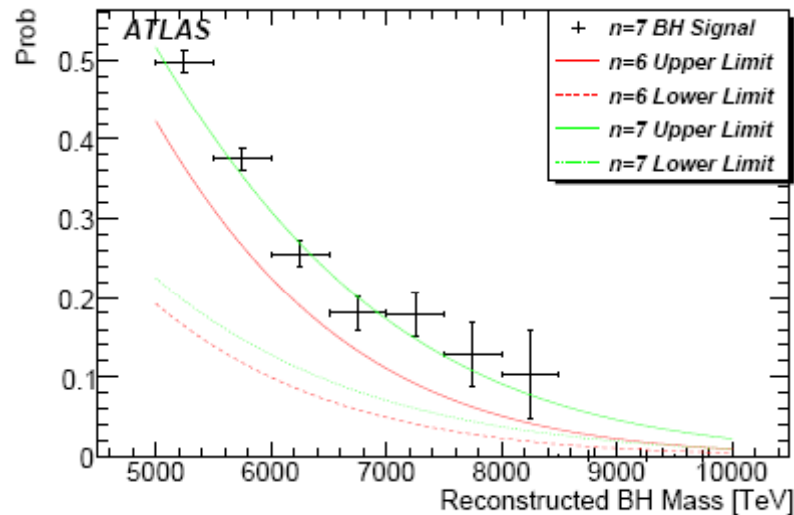
# Mass Reconstruction



- Mass information is important for more study: cross section, Planck Scale and number of extra dimensions
- A part of mass information is missed as a momentum of an undetected particle
  - A requirement of small missing  $E_T$  can improve the center value and also the resolution

|                           |        | Normalisation | Mean (GeV)   | Resolution (GeV) |
|---------------------------|--------|---------------|--------------|------------------|
| Without $E_T$ requirement | Narrow | $1018 \pm 26$ | $-217 \pm 5$ | $276 \pm 9$      |
|                           | Wide   | $276 \pm 30$  | $-148 \pm 9$ | $722 \pm 13$     |
| With $E_T$ requirement    | Narrow | $318 \pm 12$  | $-116 \pm 8$ | $215 \pm 9$      |
|                           | Wide   | $108 \pm 7$   | $118 \pm 18$ | $635 \pm 16$     |

# Measurement of Features of TeV-Scale Gravity



➤ A attempt to estimate the features using energy spectrum of Hawking radiation had been studied

➤ But non-Hawking radiation effects (such a grey-body factor) and detector effects make it difficult

➤ One of the possibility is extract from  $x_{sec}$

➤  $x_{sec}$  strongly depends on  $M_p$

➤  $n$  dependence is not so strong

➤ Left figures show one of the other methods

➤ Using emission probability of high energy particles ( $E \sim M_{BH}/2$ )

➤ Such particles should be generated at fist of BH decay

➤ Then, they should be radiated by Hawking radiation with generated  $M_{BH}$

# Summary

- Black hole production is one of the helpful signal for TeV scale gravity
  - Which can solve one of the big homework in 20century
- LHC is the first experiment which can produce TeV energy objects directly
  - **LHC has a potential of generating mini(TeV scale)-black holes**
- The ATLAS experiment has an enough potential for detection of black holes
  - Most of events passed trigger, separation from backgrounds is easy
  - Only a few  $\text{pb}^{-1}$  is needed for 5TeV discovery
    - **Black Hole may be discovered in a few days**
  - With  $100\text{pb}^{-1}$ , the discovery potential reaches to 8TeV black holes
- Some methods to estimate the parameters of TeV-scale gravity have been studied
- There are uncertainties which are inherent in the model especially around  $M_{\text{BH}} \sim M_{\text{D}}$ 
  - **But if the semi-classical estimation is valid even only in  $M_{\text{BH}} \gg M_{\text{D}}$ , we will see the mini-black holes in an early stage of the ATLAS experiment**