

NLO Assistance to LHC Searches with Complex Final States using BlackHat and Sherpa



Lance Dixon (SLAC)

for the **BlackHat** collaboration

Z. Bern, LD, G. Diana, F. Febres Cordero,
S. Höche, H. Ita, D. Kosower, D. Maître, K. Ozeren

ICHEP

Melbourne, Australia

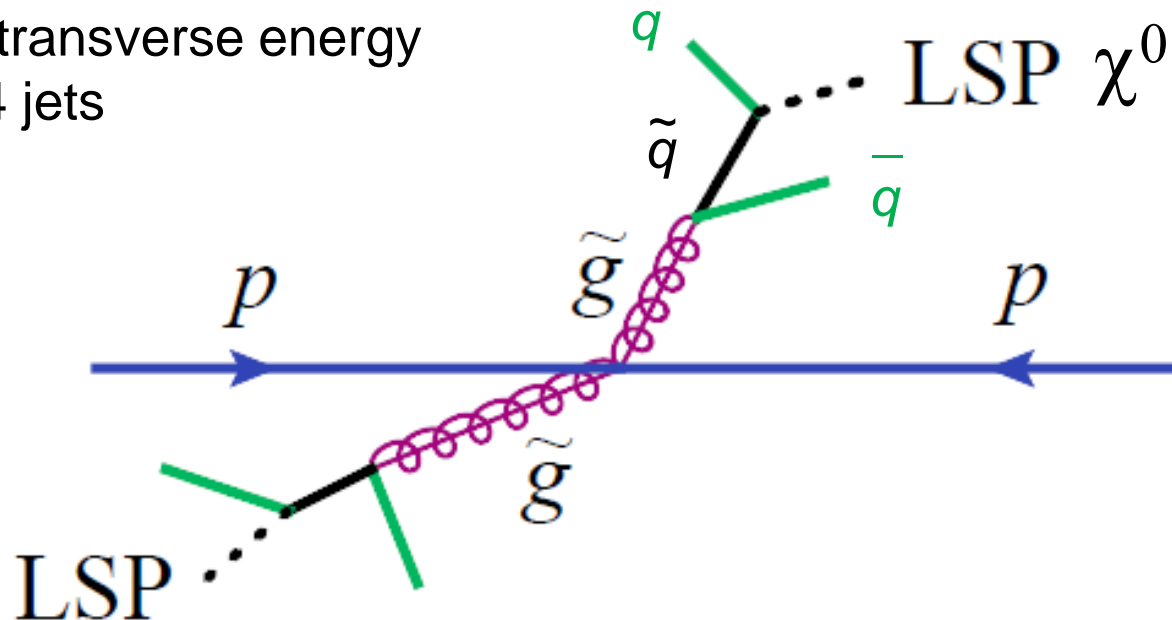
July 5, 2012

Classic SUSY dark matter signature

→ Multiple jets + missing energy (+ lepton(s)?)

In models such as supersymmetry, heavy produced particles (colored) decay rapidly to stable Weakly Interacting Massive Particle (WIMP) plus jets

→ Missing transverse energy
MET + 4 jets



Irreducible Standard Model Background

- **MET + 4 jets from**

$$pp \rightarrow Z + 4 \text{ jets},$$

$$Z \rightarrow \nu \nu$$

- Neutrinos also weakly interacting, escape detector.

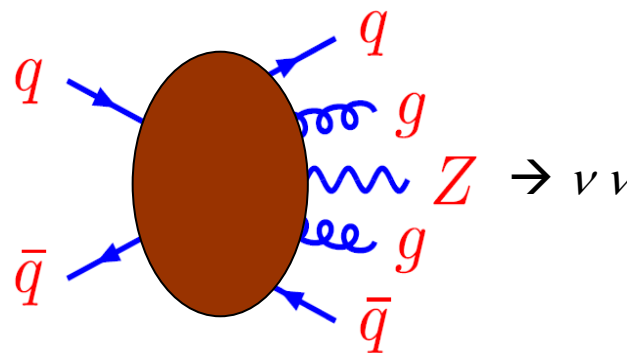
- Also large background from

$$pp \rightarrow W + 4 \text{ jets},$$

$$W \rightarrow l \nu$$

($\sim 10\times Z \rightarrow \nu \nu$ rate)

- if you lose the charged lepton
(- or if you **want** a lepton)



- Motivates theoretical and experimental study of $V + n \text{ jets}$ at Tevatron and LHC.

- Talks in this session by Strauss, Mesropian, Beauchemin, Lenzi, Ganguli, Kosower, Schönherr

Recent progress on $V + \text{jets}$ at NLO

MCFM: $V + 0,1,2 \text{ jets}$ Campbell, Ellis, hep-ph/0202176

Rocket: $W + 3 \text{ jets}$ Ellis, Melnikov, Zanderighi, 0901.4101, 0906.1445

Blackhat+Sherpa: Berger, Bern, LD, Diana, Febres Cordero, Forde, Gleisberg, Höche, Ita, Kosower, Maître, Ozeren

$W + 3 \text{ jets}$ 0902.2760, 0907.1984

$Z + 3 \text{ jets}$ 1004.1659

$W + 4 \text{ jets}$ 1009.2338

$Z + 4 \text{ jets}$ 1108.2229

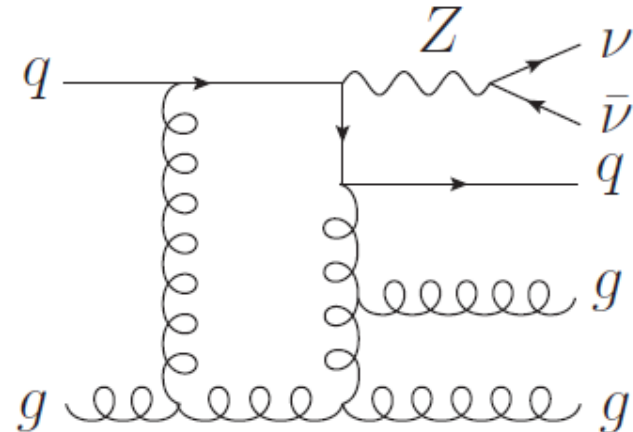
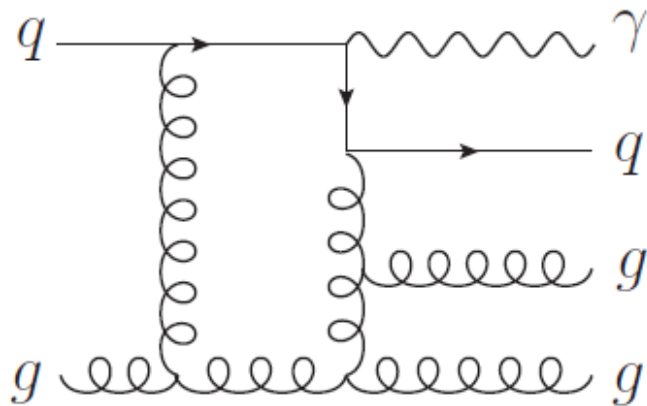
$W + 5 \text{ jets}$ 12mm.nnnn

- Could try to use such predictions **directly** for backgrounds to experimental searches.
- However, it is generally safer to use **data-driven techniques**

Data Driven Techniques

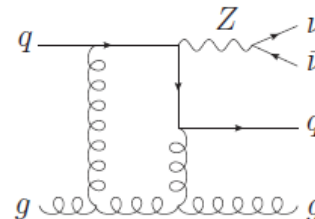
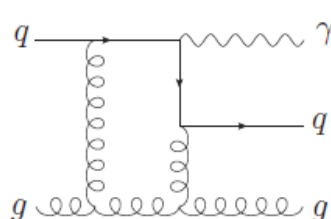
- Measure process “close” to the one you want to estimate. (Possibly the same process in a different kinematic region.)
- Rely on theory only for **ratio** of desired process to measured one.
- **Ratios** can be considerably **less sensitive** to:
 - perturbative uncertainties
 - shower + nonperturbative effects
 - jet energy scale
 - pdf uncertainties
- Nevertheless, useful to have at NLO **as well as** LO+shower.
- Examples of $V + \text{jets}$ ratios:
 - $[W + n \text{ jets}]/[Z + n \text{ jets}]$
 - $[W^+ + n \text{ jets}]/[W^- + n \text{ jets}]$
 - $[\gamma + n \text{ jets}]/[Z + n \text{ jets}]$
 - W polarization fractions
 - $[V + n \text{ jets}]/[V + (n-1) \text{ jets}]$

$\gamma + \text{jets}$ for $Z(\rightarrow \nu\nu) + n \text{ jets}$



- CMS [CMS PAS SUS-08-002, SUS-10-005, 1106.4503] and ATLAS [1107.2803, 1109.6572] both use $\gamma + \text{jets}$ to “calibrate” $Z(\rightarrow \nu\nu) + \text{jets}$ SUSY background.
- High rate compared to $Z(\rightarrow l^+l^-)$, relatively clean.
- But: How much does a γ behave like a Z?
- E.g., photon-quark collinear pole is cut off by Z mass in the Z case. Does this make much difference?

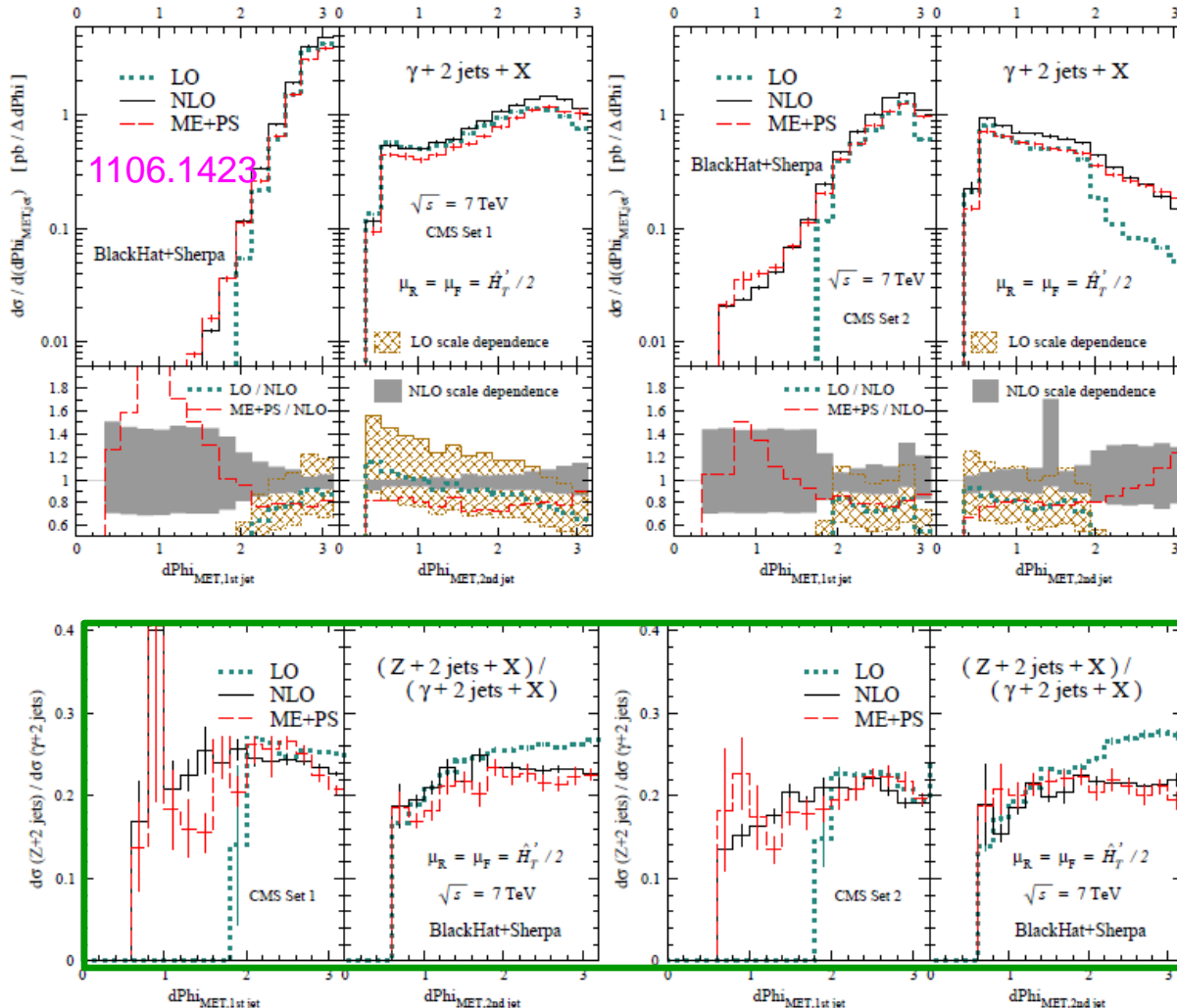
NLO ($Z + 2 \text{ jets}$)/ ($\gamma + 2 \text{ jets}$)



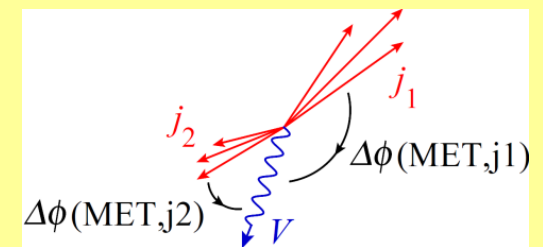
1106.1423

- Computed ($Z + 2 \text{ jets}$)/ ($\gamma + 2 \text{ jets}$) as a function of various kinematic variables, 3 different ways:
- LO (just for reference)
- NLO (probably the most reliable)
- LO+shower (ME+PS) – to estimate NLO error, and because it is similar to what CMS/ATLAS rely on.
- Traditional method of varying renormalization and factorization scales does not provide useful uncertainty estimate for ratios of similar quantities
- We used a “Frixione” photon isolation to simplify the NLO theory, but checked that it’s within $\sim 1\%$ of CMS’s isolation cone

$(Z + 2 \text{ jets})/(\gamma + 2 \text{ jets})$ distributions



- Azimuthal angle distribution, between MET vector and p_T vector of 1st, 2nd jets



- LO distribution wrong – kinematics too restrictive.
NLO and ME+PS agree to within about 10% in Z/γ ratio.

NLO ($Z + 3$ jets) / ($\gamma + 3$ jets)

- Most events in CMS samples have at least 3 jets
- For 2011 data, new (tighter) kinematic cuts

2010 data
1106.4503

Set 1: $H_T^{\text{jet}} > 300$ GeV, $|\text{MET}| > 250$ GeV

Set 2: $H_T^{\text{jet}} > 500$ GeV, $|\text{MET}| > 150$ GeV

Set 3: $H_T^{\text{jet}} > 300$ GeV, $|\text{MET}| > 150$ GeV

Set 4: $H_T^{\text{jet}} > 350$ GeV, $|\text{MET}| > 200$ GeV

control
regions

2011 data

Set 5: $H_T^{\text{jet}} > 500$ GeV, $|\text{MET}| > 350$ GeV

Set 6: $H_T^{\text{jet}} > 800$ GeV, $|\text{MET}| > 200$ GeV

Set 7: $H_T^{\text{jet}} > 800$ GeV, $|\text{MET}| > 500$ GeV

NLO ($\gamma + 3$ jets)/ ($Z + 3$ jets) results

Set	Prediction	$Z + 3\text{-jet}/\gamma + 3\text{-jet}$	$Z + 2\text{-jet}/\gamma + 2\text{-jet}$	ratio
4	LO	0.215(0.001)	0.2336(0.0003)	0.922(0.003)
	ME+PS	0.194(0.003)	0.213(0.002)	0.908(0.01)
	NLO	0.209(0.003)	0.215(0.001)	0.973(0.01)
5	LO	0.245(0.001)	0.257(0.001)	0.952(0.01)
	ME+PS	0.230(0.004)	0.239(0.004)	0.961(0.02)
	NLO	0.242(0.01)	0.246(0.002)	0.981(0.02)
6	LO	0.220(0.002)	0.232(0.001)	0.948(0.01)
	ME+PS	0.218(0.004)	0.232(0.003)	0.940(0.02)
	NLO	0.222(0.01)	0.224(0.002)	0.988(0.03)
7	LO	0.257(0.003)	0.259(0.001)	0.992(0.01)
	ME+PS	0.244(0.01)	0.261(0.003)	0.935(0.02)
	NLO	0.254(0.01)	0.255(0.003)	0.993(0.03)

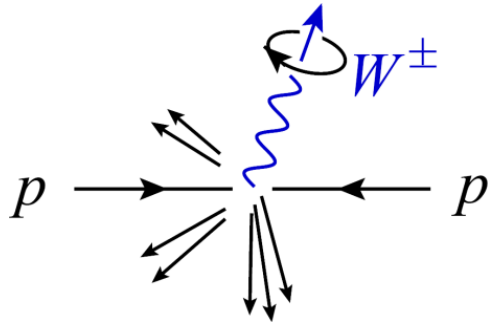
BH+S, 1206.nnnn

ME+PS, NLO
always within
10%

pdf and other
uncertainties
5% or less

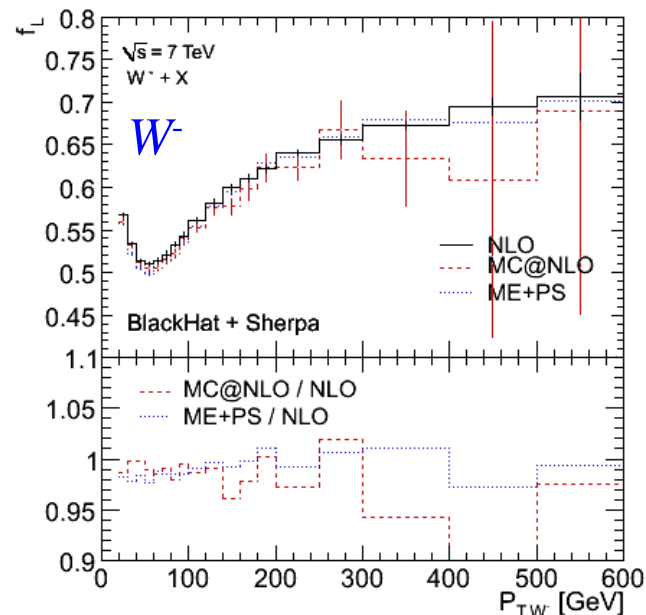
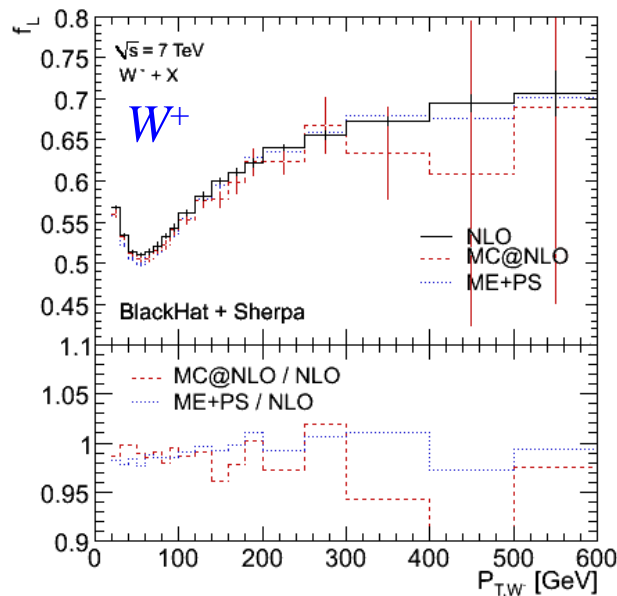
Validates this
method of
estimating
background

W^+ and W^- “differ” at LHC: polarized **same** way [left-handed]



Helicity frame:

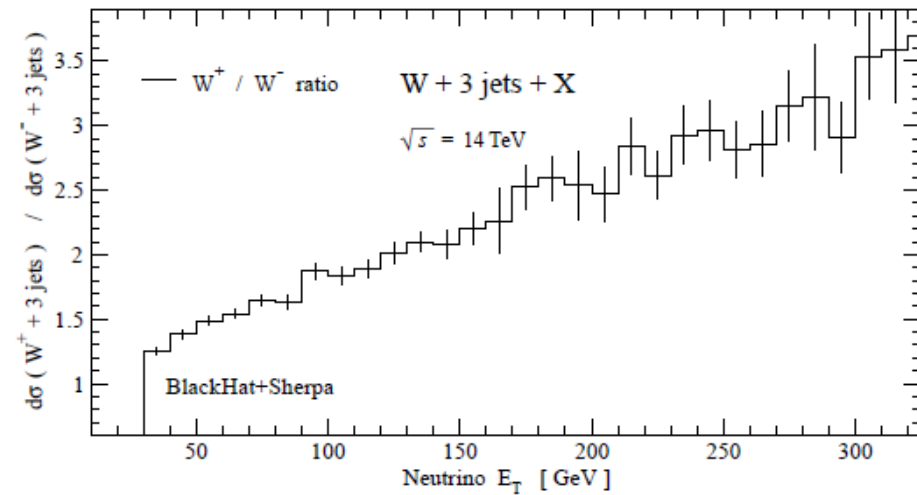
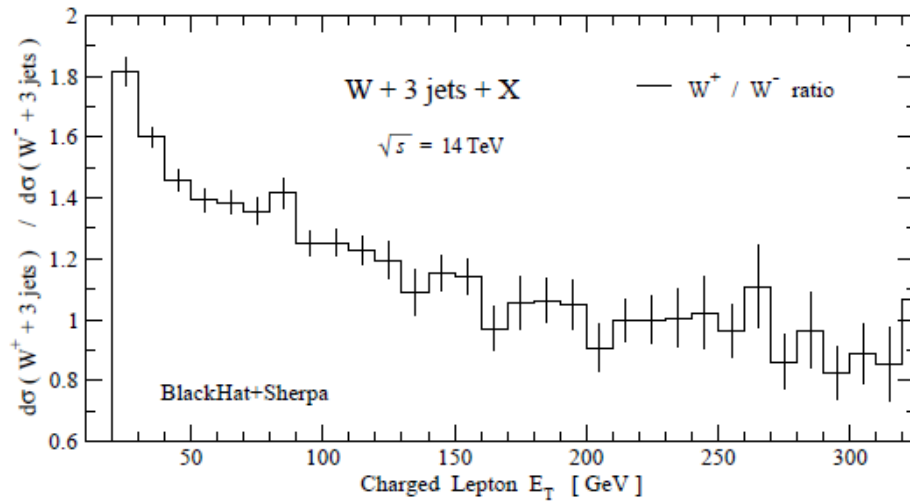
$$\frac{1}{\sigma} \frac{d\sigma}{d\cos\theta^*} = \frac{3}{8}(1 \mp \cos\theta^*)^2 f_L + \frac{3}{8}(1 \pm \cos\theta^*)^2 f_R + \frac{3}{4}\sin^2\theta^* f_0$$



1103.5445

Leptonic E_T in $W^\pm + 3$ jets

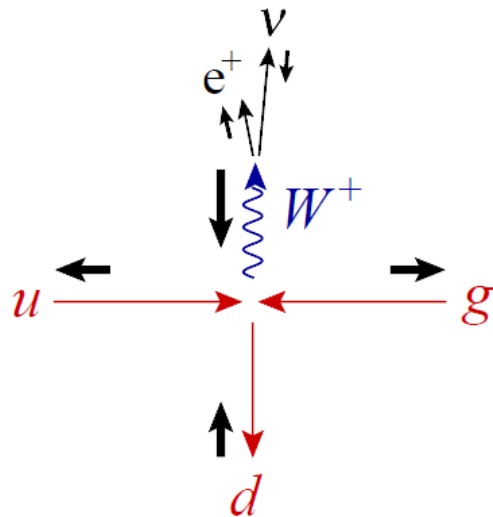
0907.1984



W^+/W^- transverse lepton ratios are **skewed** because they are analyzing a **large left-handed W** polarization at large $p_T(W)$

Origin of W polarization at LHC at large $p_T(W)$

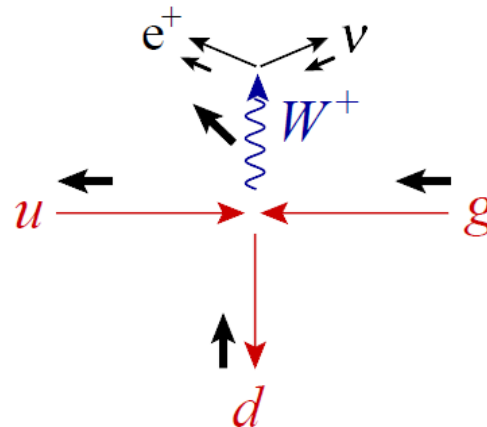
$ug \rightarrow W^+ d$ dominates due to pdfs at a pp machine.
Only 2 relevant helicity configurations:



$$A^{\text{tree}} \propto \frac{\langle d \nu \rangle^2}{\langle u g \rangle \langle g d \rangle}$$

$$d\sigma \propto (k_d \cdot k_\nu)^2$$

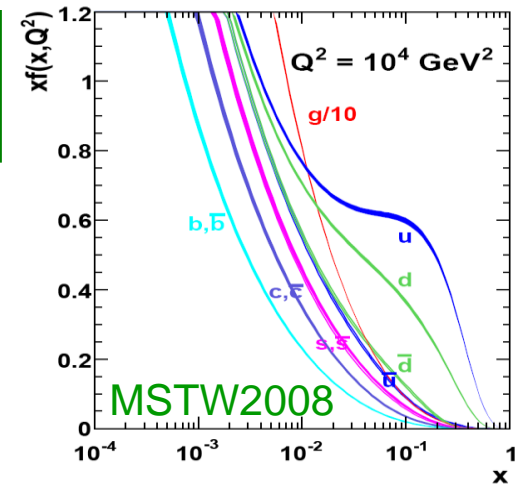
100% left-handed
(in partonic CM frame)



$$A^{\text{tree}} \propto \frac{[u e]^2}{[u g][g d]}$$

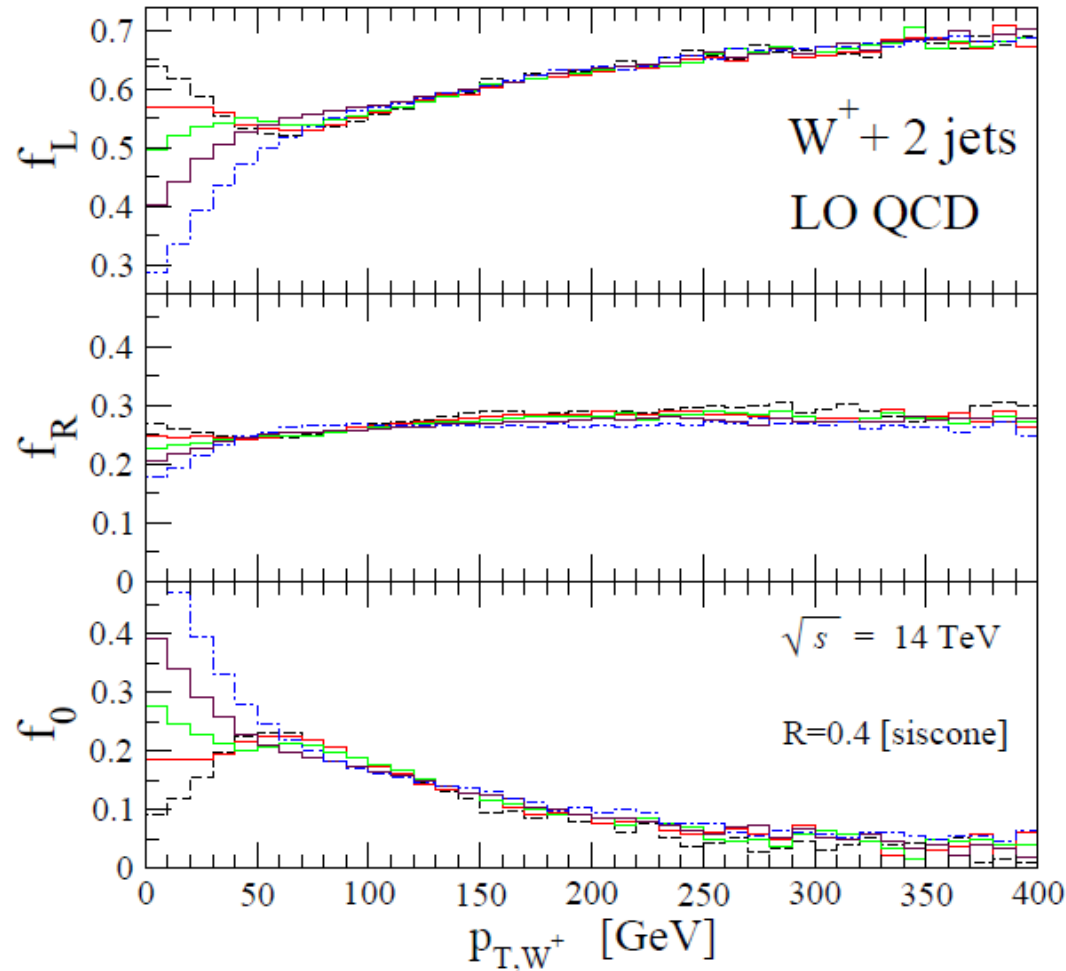
$$d\sigma \propto (k_u \cdot k_e)^2$$

Mixture of polarizations
→ 100% right-handed, but only 1/4 the size



1103.5445

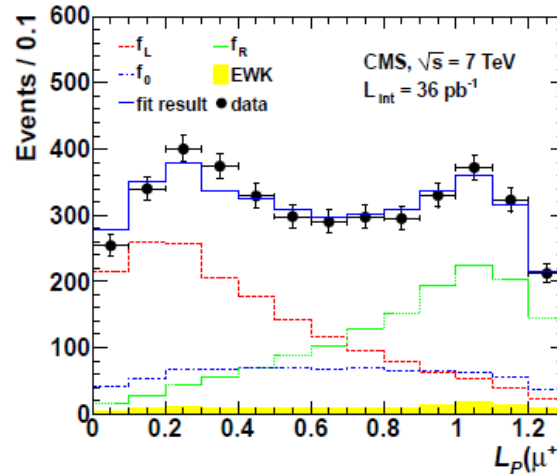
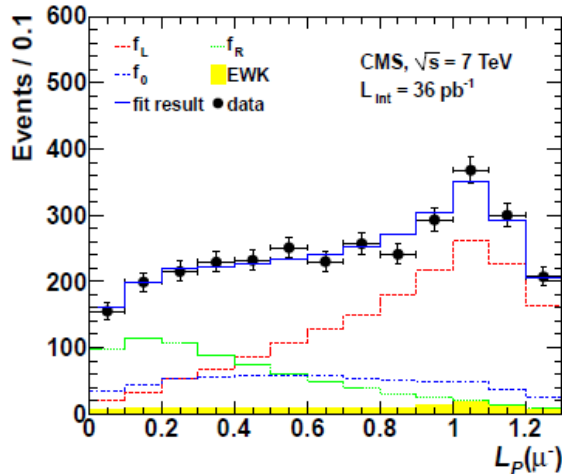
Stable W polarization: $W + 2$ jets, vs. Jet p_T cut



1103.5445

Also stable vs.
number of jets

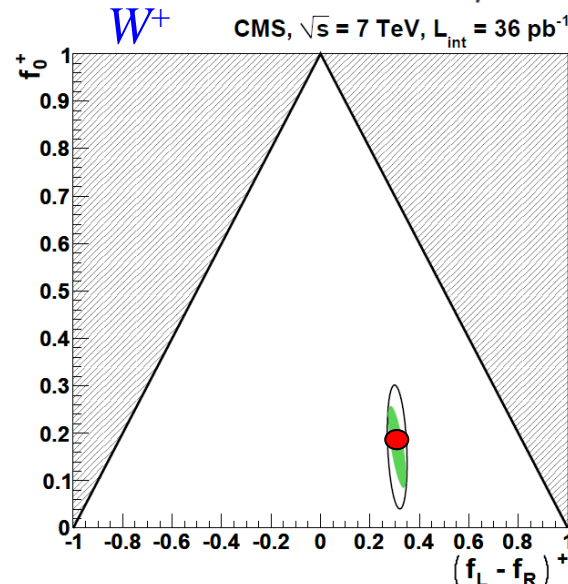
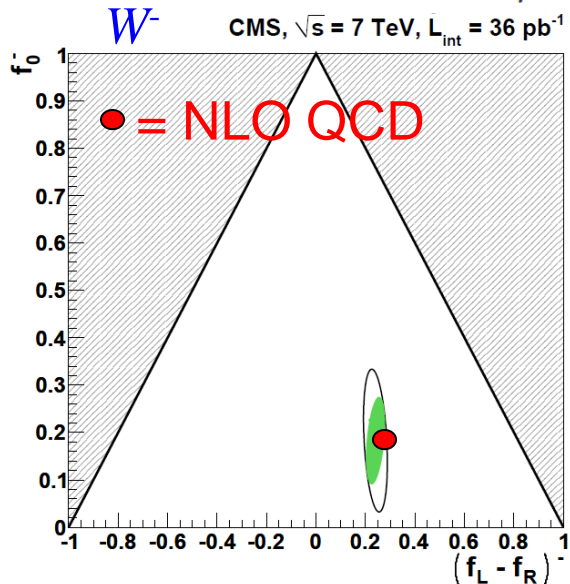
CMS measurement – no explicit jet cuts



1104.3829

$$p_T(W) > 50 \text{ GeV}$$

$$L_P = \frac{\vec{p}_T(\ell) \cdot \vec{p}_T(W)}{|\vec{p}_T(W)|^2}$$



Also ATLAS
measurement
(smaller
uncertainties)
using

$$\cos \theta_{2D} = \frac{\vec{p}_T^{\ell*} \cdot \vec{p}_T^W}{|\vec{p}_T^{\ell*}| |\vec{p}_T^W|}$$

1203.2165

Conclusions

- We compared $\gamma + 2,3 \text{ jets}$ to $Z + 2,3 \text{ jets}$ for cuts relevant for CMS SUSY searches with 2010 and 2011 data.
- We found very similar results for the ratio, between NLO and ME+PS approximations,
- This validates the data-driven method of using $\gamma + \text{jets}$ to calibrate the $Z + \text{jets}$ background to the MET + jets SUSY searches.
- Left-handed W polarization can provide another handle on $W + \text{jets}$ backgrounds, due to the charge asymmetries it induces.
- In fact, CMS [1107.1870] has used the measured lepton p_T spectrum in $W + \text{jets}$, plus the predicted W polarization to infer the MET distribution in $W + \text{jets}$ backgrounds to SUSY.
- Many other ratios out there to study and exploit!

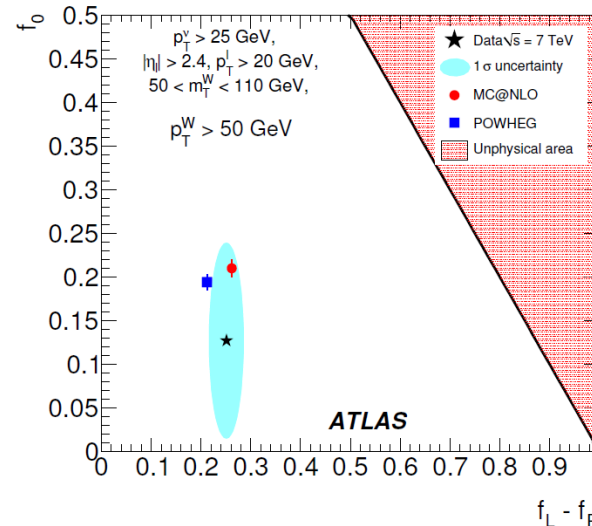
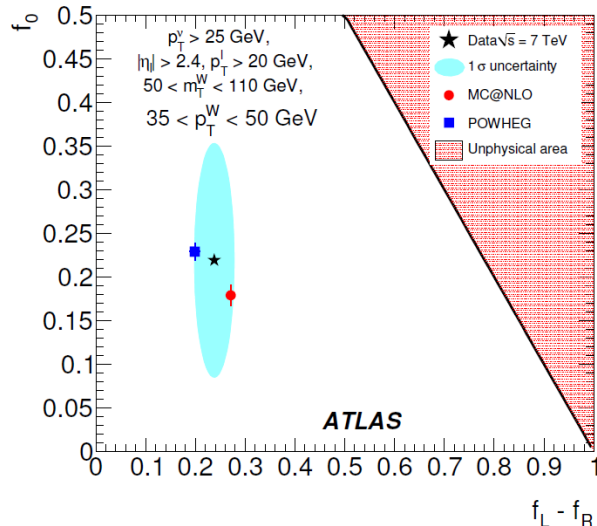
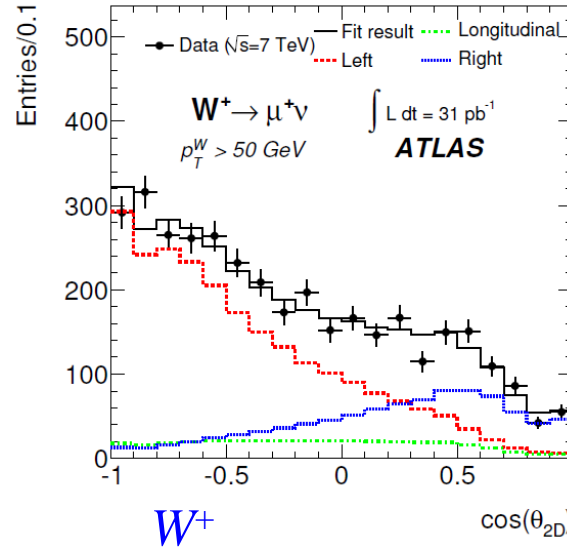
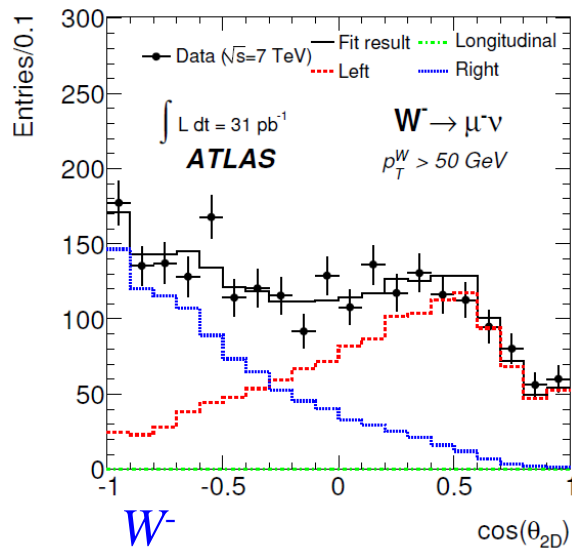
Extra slides

ATLAS measurement – no explicit jet cuts

1203.2165

$p_T(W) > 50 \text{ GeV}$

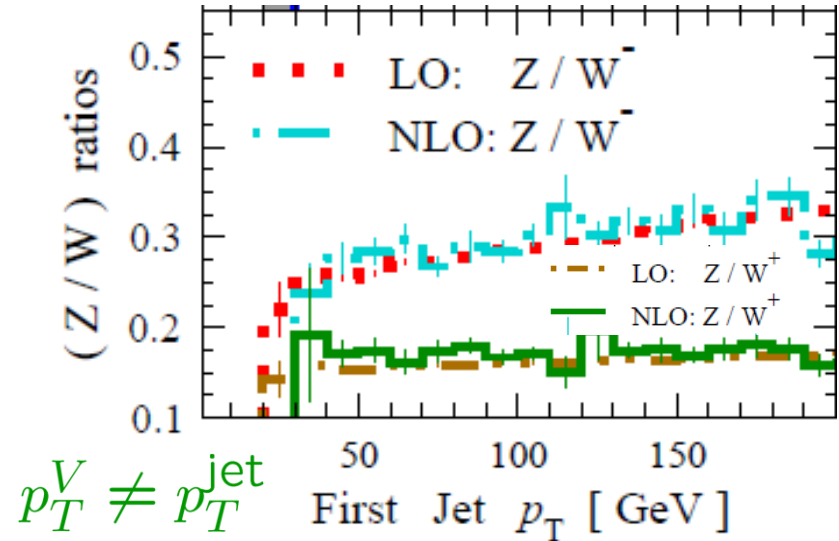
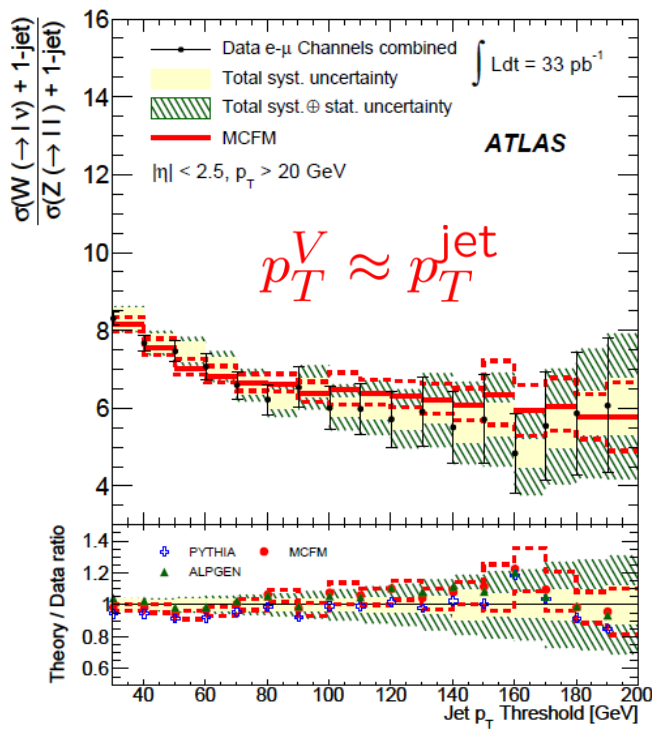
$$\cos \theta_{2D} = \frac{\vec{p}_T^{\ell*} \cdot \vec{p}_T^W}{|\vec{p}_T^{\ell*}| |\vec{p}_T^W|}$$



Different dynamics for $W/Z + \text{jets}$ ratios for 1 jet, versus more jets

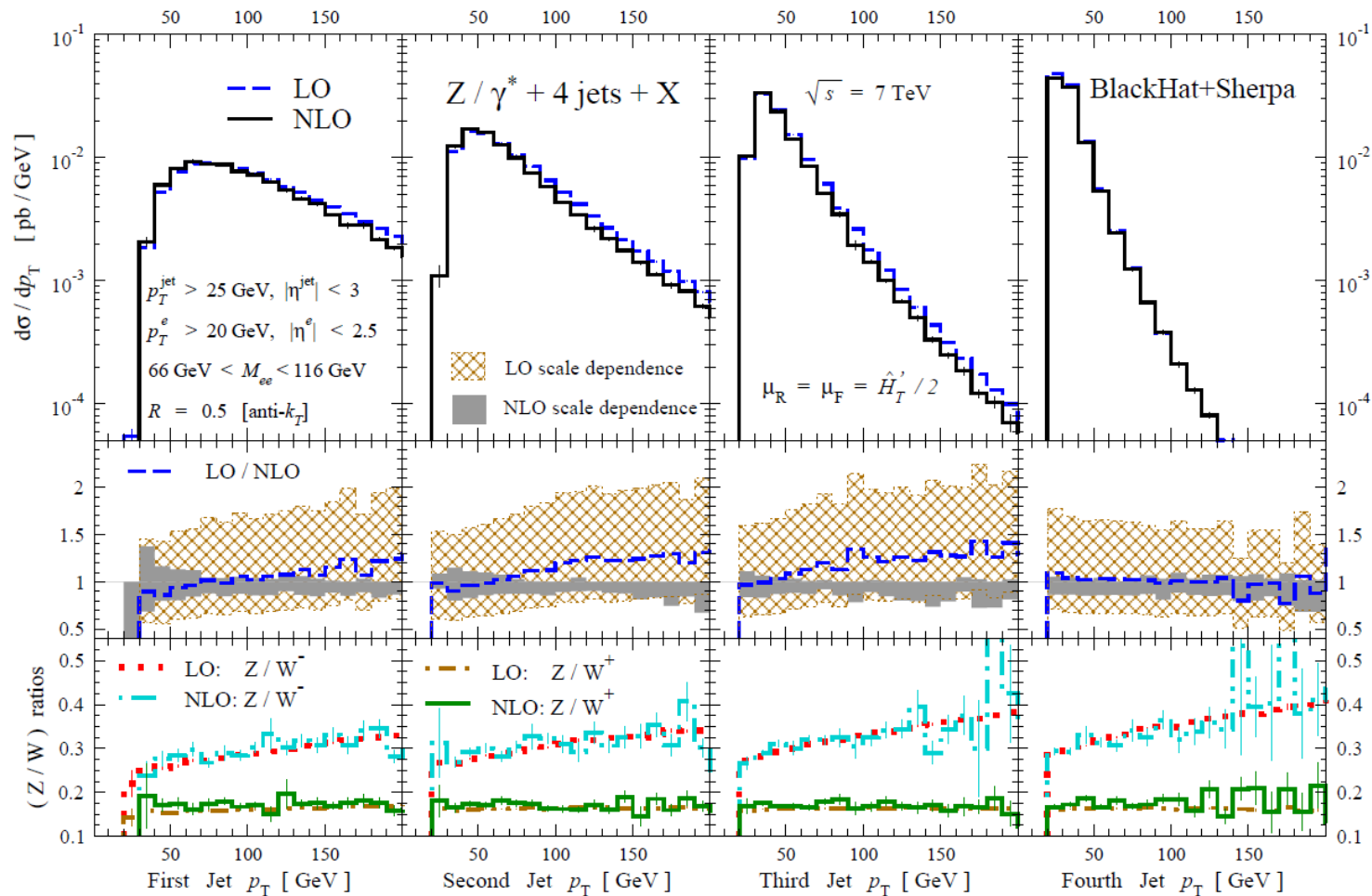
Recent ATLAS measurement of $W/Z + \text{exactly 1 jet}$ ratio 1108.4908 – strong dependence on jet p_T

First jet in $W/Z + 4 \text{ jet}$ ratio: $\sim \text{flat}$ in jet p_T



- Would be nice to measure with 2,3,4 jets!
- Also, why not separate W^+ from W^- ?

NLO $pp \rightarrow Z + 4 \text{ jets}$, and ratio to W^\pm



Ita et al.
1108.2229

Ratio of W^+ to W^- rates with jets

Kom, Stirling, 1004.3404

$$R^\pm(n) \equiv \frac{\sigma(W^\pm + n \text{ jets})}{\sigma(W^\mp + n \text{ jets})}$$

- Very small experimental systematics
- NLO QCD corrections quite small, 2% or less
- → Intrinsic theoretical uncertainty very small.
- PDF uncertainty also ~1-2%. Driven by PDF ratio

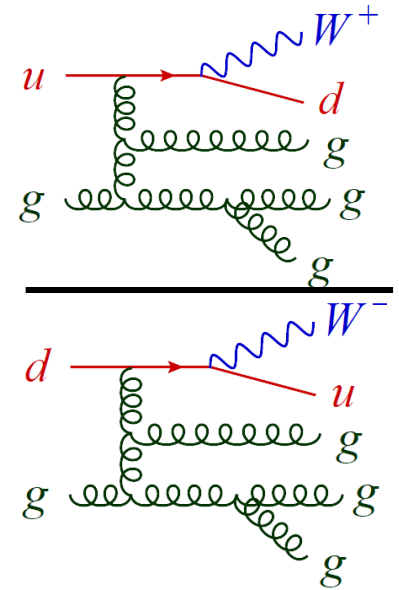
$$u(x)/d(x)$$

in well-measured valence region of moderate x .

- Sensitive to **new physics** (or Higgs, or top quark pairs) that produces W^\pm **symmetrically**
- Fraction of **new physics** in sample is:

$$f_{\text{NP}} = \frac{2(R_{\text{SM}}^\pm - R_{\text{exp.}}^\pm)}{(R_{\text{SM}}^\pm + 1)(R_{\text{exp.}}^\pm - 1)}$$

n	QQ	Qg	gg
0	100	0	0
1	18	82	0
2	21	73	6
3	23	70	7
4	25	67	8



W^+ to W^- ratios at NLO

BH+S, 1009.2338

no. jets	W^- LO	W^- NLO	W^+/W^- LO	W^+/W^- NLO
0	$1614.0(0.5)^{+208.5}_{-235.2}$	$2077(2)^{+40}_{-31}$	$1.656(0.001)$	$1.580(0.004)$
1	$264.4(0.2)^{+22.6}_{-21.4}$	$331(1)^{+15}_{-12}$	$1.507(0.002)$	$1.498(0.009)$
2	$73.14(0.09)^{+20.81}_{-14.92}$	$78.1(0.5)^{+1.5}_{-4.1}$	$1.596(0.003)$	$1.57(0.02)$
3	$17.22(0.03)^{+8.07}_{-4.95}$	$16.9(0.1)^{+0.2}_{-1.3}$	$1.694(0.005)$	$1.66(0.02)$
4	$3.81(0.01)^{+2.44}_{-1.34}$	$3.55(0.04)^{+0.08}_{-0.30}$	$1.812(0.001)$	$1.73(0.03)$

$$p_T^{\text{jet}} > 25 \text{ GeV}, |\eta^{\text{jet}}| < 3$$

$$E_T^e > 20 \text{ GeV}, |\eta^e| < 2.5$$

$$E_T^{\nu} > 20 \text{ GeV}, M_T^W > 20 \text{ GeV}$$

$$R = 0.5 \text{ [anti-}k_T\text{]}$$

- Huge scale dependence at LO
cancels in ratio
- Small corrections from LO \rightarrow NLO
- Increases with n due to increasing x

