#### **Higgs Boson Mass in GMSB with Messenger-Matter Mixing**

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#### Based on:

Higgs boson of mass 125 GeV in GMSB models with messenger—matter mixing

A. Albaid and K.S. Babu, arXiv: 1207.1014 [hep-ph]

# Higgs boson mass in GMSB models

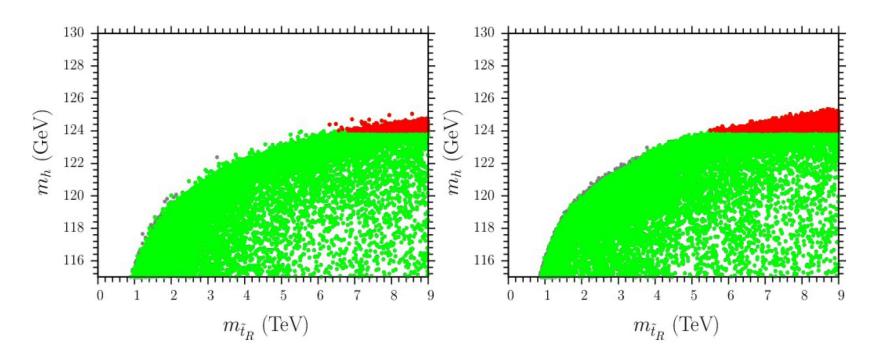
Observation of Standard Model–like Higgs particle severely strains minimal gauge mediation models

$$\begin{split} m_h^2 &= M_Z^2 \cos^2 2\beta \left( 1 - \frac{3}{8\pi^2} \frac{m_t^2}{v^2} t \right) \\ &+ \frac{3}{4\pi^2} \frac{m_t^4}{v^2} \left[ \frac{1}{2} X_t + t + \frac{1}{16\pi^2} \left( \frac{3}{2} \frac{m_t^2}{v^2} - 32\pi\alpha_3 \right) \left( X_t t + t^2 \right) \right], \\ v^2 &= v_d^2 + v_u^2, \ t = \log \left( \frac{M_s^2}{M_t^2} \right), \ X_t = \frac{2\tilde{A}_t^2}{M_s^2} \left( 1 - \frac{\tilde{A}_t^2}{12M_s^2} \right), \ \tilde{A}_t = A_t - \mu \cot \beta \end{split}$$

To obtain  $m_h \simeq 125$  GeV,  $X_t \simeq 6$  needed

GMSB models typically have  $X_t \ll 6$ , since  $A_t = 0$  at  $M_{\text{mess}}$   $m_h = 125$  GeV requires stop mass larger than 6 TeV!

## Higgs boson mass in GMSB models



Stop mass versus Higgs mass in GMSB with n=1(5) Ajaib, Gogoladze, Nasir, Shafi (2012)

If stop masses < 2 TeV,  $m_h <$  118 GeV

# **SUSY Spectrum in GMSB Models**

Messenger fields  $\Phi_i + \overline{\Phi}_i$  belonging to complete vector-like multiplets of SU(5) introduced

They couple to a singlet Z which has VEVs along scalar and F-components

$$W = \lambda_i Z \Phi_i \overline{\Phi}_i$$

$$M_a = \frac{\alpha_a}{4\pi} \wedge n_a(i) g(x_i) \quad (a = 1 - 3),$$

$$\tilde{m}^2 = 2 \wedge^2 \sum_{a=1}^3 \left(\frac{\alpha_a}{4\pi}\right)^2 C_a n_a(i) f(x_i) .$$

$$\wedge \equiv \langle F_Z \rangle / \langle Z \rangle, \ n_a(i) = 1 \text{ for } N + \overline{N}, \ C_a(N) = (N^2 - 1)/(2N)$$

$$A_f = 0 \text{ for all } f$$

$$B = 0$$

Degneracy of scalars solves SUSY flavor problem

# **Messenger-Matter Mixing**

Consider  $5 + \overline{5}$  messenger fields

$$5 = (\overline{d^c}_m + \overline{L}_m)$$
 and  $\overline{5} = (d_m^c + L_m)$ 

Normally these fileds are assumed not to mix with MSSM fields

(⇒ Messenger number conservation!)

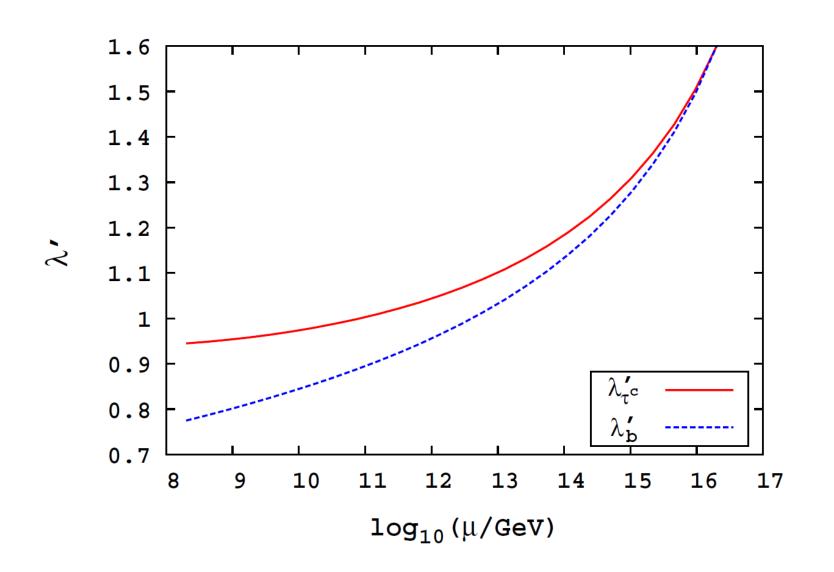
Messengers can mix with MSSM fields, however

$$W_{5+\overline{5}} = f_d \overline{d^c}_m d^c_m Z + f_e \overline{L}_m L_m Z + \lambda'_b Q_3 d^c_m H_d + \lambda'_{\tau^c} L_m e^c_3 H_d$$

Only 3rd family mixing is assumed

In 
$$SU(5)$$
,  $W = f_0 5_m \overline{5}_m Z + \lambda_0' 10_3 \overline{5}_m \overline{5}_H$   
 $(f_0, \lambda_0')$  at GUT scale

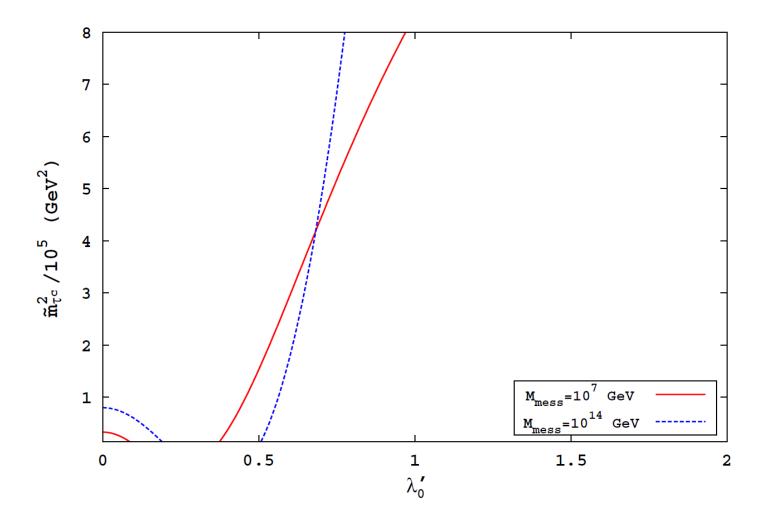
### **Evolution of Mixed Yukawas**



# **Messenger-Matter Mixing (cont.)**

New contributions to soft SUSY breaking parameters:

$$\delta \tilde{m}_{Q_{3}}^{2} = \frac{\alpha'_{b} \Lambda^{2}}{8\pi^{2}} \left( 3\alpha'_{b} + \frac{1}{2}\alpha'_{\tau^{c}} - \frac{8}{3}\alpha_{3} - \frac{3}{2}\alpha_{2} - \frac{7}{30}\alpha_{1} \right), 
\delta \tilde{m}_{\tau^{c}}^{2} = \frac{2\alpha'_{\tau^{c}} \Lambda^{2}}{8\pi^{2}} \left( 2\alpha'_{\tau^{c}} + \frac{3}{2}\alpha'_{b} - \frac{3}{2}\alpha_{2} - \frac{9}{10}\alpha_{1} \right), 
\delta \tilde{m}_{H_{d}}^{2} = \frac{\delta \tilde{m}_{\tau^{c}}^{2}}{2} + 3\delta \tilde{m}_{Q_{3}}^{2} + \frac{3\Lambda^{2}\alpha'_{b}\alpha_{t}}{16\pi^{2}} 
\delta A_{t} = -\frac{1}{4\pi}\alpha'_{b}\Lambda, 
\delta A_{b} = -\left( \frac{4\alpha'_{b} + \alpha'_{\tau^{c}}}{4\pi} \right) \Lambda, \qquad \alpha'_{b} = \frac{\lambda'_{b}^{2}}{4\pi}, \quad \alpha'_{\tau^{c}} = \frac{\lambda'_{\tau^{c}}^{2}}{4\pi} 
\delta A_{\tau} = -\left( \frac{3\alpha'_{b} + 3\alpha'_{\tau^{c}}}{4\pi} \right) \Lambda,$$



Constraints on Yukawa from positivity of stau mass-squared

# **Improved Higgs boson mass**

$\lambda'_0$	$m_h({ m GeV})$	$\Lambda(10^5 { m GeV})$	$M(10^{13}\mathrm{GeV})$	$\tilde{m}_{t_1}(\text{GeV})$	$\tilde{m}_{t_2}(\mathrm{GeV})$
0	114	2	1.78	1249	1695
0.8	116	2	10	1212	1583
1.2	119	2	10	384	2613

Messenger-matter mixing increases  $m_h$  by 5 GeV

 $m_h \simeq 121$  GeV can be realized with light SUSY spectrum

# **Messenger-Matter Mixing with 10+10\***

Consider  $10 + \overline{10}$  messenger fields

$$10 + \overline{10} = (Q_m + \overline{Q}_m) + (u_m^c + \overline{u}^c_m) + (e_m^c + \overline{e}^c_m)$$

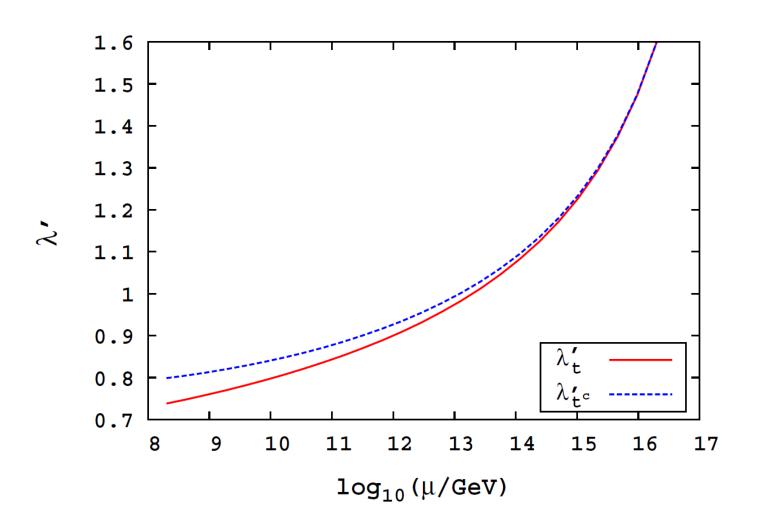
$$W_{10+\overline{10}} = \lambda'_{t^c}Q_3u_m^cH_u + \lambda'_tQ_mu_3^cH_u + \lambda'_mQ_mu_m^cH_u + f_{e^c}\overline{e^c}_me_m^cZ + f_{u^c}\overline{u^c}_mu_m^cZ + f_Q\overline{Q}_mQ_mZ.$$

Only 3rd family mixing is assumed

In 
$$SU(5)$$
,  $W \supset \lambda'_0 10_3 10_m 5_H + \lambda'_{m0} 10_m 10_m 5_H + f_0 10_m \overline{10}_m Z$ 

 $(f_0, \lambda'_0, \lambda'_{m0})$  at GUT scale

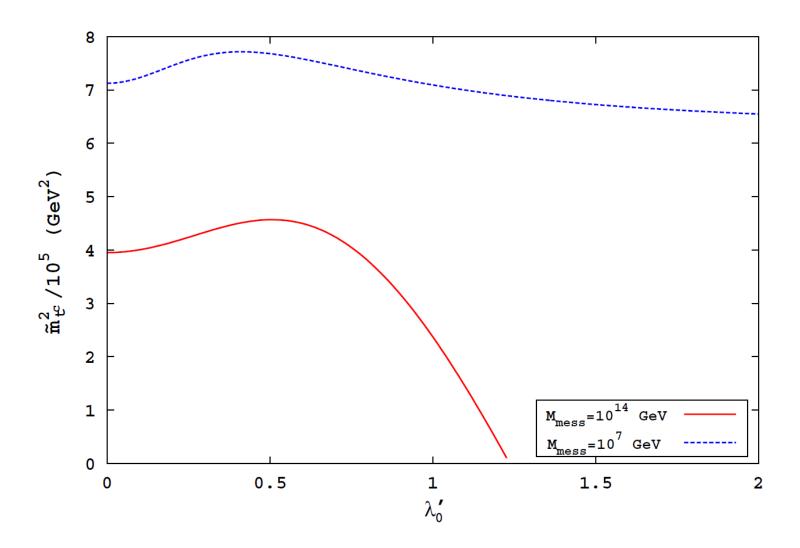
### **Evolution of mixed Yukawas in 10+10\***



# Messenger-Matter Mixing in 10+10\*

#### New contributions to soft SUSY breaking parameters:

$$\begin{split} \delta \tilde{m}_{Q_3}^2 &= \frac{\Lambda^2}{8\pi^2} \left[ \alpha'_{t^c} \left( 3\alpha'_{t^c} + \frac{3}{2}\alpha'_t + \frac{5}{2}\alpha'_m - \frac{8}{3}\alpha_3 - \frac{3}{2}\alpha_2 - \frac{13}{30}\alpha_1 \right) \right. \\ &- \alpha_t \left( \frac{5}{2}\alpha'_t + \frac{3}{2}\alpha'_m \right) \right], \\ \delta \tilde{m}_{t^c}^2 &= \frac{2\Lambda^2}{8\pi^2} \left[ \alpha'_t \left( 3\alpha'_t + \frac{3}{2}\alpha'_{t^c} + 2\alpha'_m - \frac{8}{3}\alpha_3 - \frac{3}{2}\alpha_2 - \frac{13}{30}\alpha_1 \right) \right. \\ &- \alpha_t \left( 2\alpha'_{t^c} + \frac{3}{2}\alpha'_m \right) \right], \\ \delta \tilde{m}_{H_u}^2 &= \frac{3\Lambda^2}{8\pi^2} \left[ \alpha'_{t^c} \left( 3\alpha'_{t^c} + \frac{3}{2}\alpha'_t + \frac{5}{2}\alpha'_m - \frac{8}{3}\alpha_3 - \frac{3}{2}\alpha_2 - \frac{13}{30}\alpha_1 \right) \right. \\ &+ \alpha'_t \left( 3\alpha'_t + \frac{3}{2}\alpha'_{t^c} + 2\alpha'_m - \frac{8}{3}\alpha_3 - \frac{3}{2}\alpha_2 - \frac{13}{30}\alpha_1 \right) \\ &+ \alpha'_m \left( 3\alpha'_m + 2\alpha'_t + \frac{5}{2}\alpha'_{t^c} - \frac{8}{3}\alpha_3 - \frac{3}{2}\alpha_2 - \frac{13}{30}\alpha_1 \right) \right] \\ \delta A_t &= - \left[ \frac{5\alpha'_t + 4\alpha'_{t^c} + 3\alpha'_m}{4\pi} \right] \Lambda, \\ \delta A_b &= -\frac{\alpha'_{t^c}}{4\pi} \Lambda \qquad \qquad \alpha'_{t^c} = \frac{\lambda'_{t^c}^2}{4\pi}, \quad \alpha'_t = \frac{\lambda'_t^2}{4\pi}, \quad \alpha'_m = \frac{\lambda'_t^2}{4\pi} \end{split}$$



No major constraint from positivity of stop mass-squared

# **Improved Higgs boson mass**

$\lambda'_0$	$m_h({ m GeV})$	$\Lambda(10^5{ m GeV})$	$M_{ m mess}({ m GeV})$	$\tilde{m}_{t_1}(\mathrm{GeV})$	$\tilde{m}_{t_2}(\mathrm{GeV})$	$A_t/M_s$
0	121	0.97	$2 \times 10^{13}$	928	1636	-1.8
0.4	123	0.91	$3 \times 10^{13}$	656	1612	-2.3
0.6	123	0.848	$10^{12}$	673	1512	-2.3
0.8	123	0.784	$10^{11}$	682	1509	-2.3
2	123	0.784	$10^{8}$	753	1425	-2.2

 $m_h \simeq 125$  GeV can be realized with light SUSY spectrum

# **Sample SUSY Spectrum**

Particle		$10 + \overline{10}$	$10 + \overline{10}$	$5+\overline{5}$
Inputs	$M_{ m mess}$	10 <sup>8</sup>	$4 \times 10^{5}$	108
	$N_{ m mess}$	3	3	1
	$\Lambda(10^5{ m GeV})$	0.45	0.3	1.5
	an eta	10	6.1	15.6
	$f_0$	0.25	0.25	0.25
	$\lambda_0$	1.3	1.2	1.2
Higgs:	$m_h$	122	118	114.5
	$m_H^0$	858	592	1690
	$m_A$	858	591	1690
	$m_{H^{\pm}}$	862	597	1689
Gluino:	$ ilde{m}_g$	980	667	1041
Neutralinos:	$m_{\chi_1}$	186	124	208
	$m_{\chi_2}$	346	225	408
	$m_{\chi_3}$	800	557	781
	$m_{\chi_4}$	807	569	790
Charginos:	$\chi_1^+$	347	227	409
	$\chi_2^+$	807	569	790
Squarks:	$\tilde{m}_{u_L,c_L}$	972	657	1480
	$\tilde{m}_{u_R,c_R}$	929	632	1377
	$\tilde{m}_{d_L,s_L}$	971	657	1480
	$\tilde{m}_{d_R,s_R}$	922	630	1365
	$ ilde{m}_{b_L}$	800	555	1315
	$ ilde{m}_{b_R}$	919	629	1294
	$ ilde{m}_{t_L}$	853	621	1315
	$ ilde{m}_{t_R}$	412	270	1123
Sleptons:	$\tilde{m}_{e_L,\mu_L}$	323	200	596
	$ ilde{m}_{ u_{eL}, u_{\mu_L}}$	323	200	596
	$ ilde{m}_{e_R,\mu_R}$	152	92	290
	$ ilde{m}_{ au_L}$	322	197	539
	$ ilde{m}_{ au_R}$	151	92	1543

#### **SUSY Flavor Violation**

 $10 + \overline{10}$  model embedded in a U(1) flavor symmetry:

SU(5)	$10_1$	$10_2$	$10_{3}$	$\overline{5}_1$	$\overline{5}_2,\overline{5}_3$	$5_u, \overline{5}_d$	S	$10_m$	$\overline{10}_m$	Z
U(1)	4	2	0	1+p	p	0	-1	0	$-\alpha$	$\alpha$

$$W_{10+\overline{10}} = (\lambda'_{u^{c}}\epsilon^{4}Q_{1} + \lambda'_{c^{c}}\epsilon^{2}Q_{2} + \lambda'_{t^{c}}Q_{3})u_{m}^{c}H_{u} + Q_{m}(\lambda'_{u}\epsilon^{4}u_{1}^{c} + \lambda'_{c}\epsilon^{2}u_{2}^{c} + \lambda'_{t^{c}}Q_{3})u_{m}^{c}H_{u} + \lambda'_{u}\epsilon^{4}u_{1}^{c} + \lambda'_{c}\epsilon^{2}u_{2}^{c} + \lambda'_{t^{c}}Q_{m}u_{m}^{c}H_{u} + \lambda'_{b}\epsilon^{p}Q_{m}d_{3}^{c}H_{d} + \lambda'_{\tau}\epsilon^{p}L_{3}e_{m}^{c}H_{d} + f_{e^{c}}\overline{e^{c}}_{m}e_{m}^{c}Z + f_{u^{c}}\overline{u^{c}}_{m}u_{m}^{c}Z + f_{Q}\overline{Q}_{m}Q_{m}Z.$$

With  $\epsilon \simeq 0.2$ , SUSY FCNC suppressed sufficiently

# Flavor Symmetry and Neutrino Mixing

$$M^{u} = Y^{u}v_{u} = \begin{pmatrix} y_{11}^{u} \epsilon^{8} & y_{12}^{u} \epsilon^{6} & y_{13}^{u} \epsilon^{4} \\ y_{21}^{u} \epsilon^{6} & y_{22}^{u} \epsilon^{4} & y_{23}^{u} \epsilon^{2} \\ y_{31}^{u} \epsilon^{4} & y_{32}^{u} \epsilon^{2} & y_{33}^{u} \end{pmatrix} v_{u} ,$$

$$M^{d} = Y^{d}v_{d} = \epsilon^{p} \begin{pmatrix} y_{11}^{d} \epsilon^{5} & y_{12}^{d} \epsilon^{3} & y_{13}^{d} \epsilon \\ y_{21}^{d} \epsilon^{4} & y_{22}^{d} \epsilon^{2} & y_{23}^{d} \\ y_{31}^{d} \epsilon^{4} & y_{32}^{d} \epsilon^{2} & y_{33}^{d} \end{pmatrix} v_{d} ,$$

$$M^{e} = Y^{e}v_{d} = \epsilon^{p} \begin{pmatrix} y_{11}^{e} \epsilon^{5} & y_{12}^{e} \epsilon^{4} & y_{13}^{e} \epsilon^{4} \\ y_{21}^{e} \epsilon^{3} & y_{22}^{e} \epsilon^{2} & y_{23}^{e} \epsilon^{2} \\ y_{31}^{e} \epsilon & y_{32}^{e} & y_{33}^{e} \end{pmatrix} v_{d} .$$

Lopsided mass matrices explain small quark mixings and large neutrino mixings

Babu, Barr (1996) Albright, Babu, Barr (1998) Elwood, Irges, Ramond (1998) Sato, Yanagida (1998)

# **SUSY Flavor Violation**

Mass Insertion $(\delta)$	$5+\overline{5}$	$10 + \overline{10}$	Process	Exp. Bounds
$(\delta_{12}^l)_{LL}$	-	$\epsilon^{1+2p}$		0.00028
$(\delta_{12}^l)_{RR}$	$r \epsilon^6$	-	$\mu \to e \gamma$	0.0004
$(\delta_{12}^l)_{RL,LR}$	$r \kappa_5^l(\epsilon^4, \epsilon^3)$	$\kappa_{10}^l \ (\epsilon^{4+2p}, \epsilon^{3+2p})$		$1.3 \times 10^{-6}$
$(\delta^l_{13})_{LL}$	-	$\epsilon^{1+2p}$		0.026
$(\delta^l_{13})_{RR}$	$r \epsilon^4$	-	$\tau \to e \gamma$	0.04
$(\delta^l_{13})_{RL,LR}$	$r \kappa_5^l(\epsilon^4, \epsilon^1)$	$\kappa_{10}^l(\epsilon^{4+2p}, \epsilon^{1+2p})$		0.002
$(\delta^l_{23})_{LL}$	-	$\epsilon^{2p}$		0.02
$(\delta_{23}^l)_{RR}$	$r \epsilon^2$	_	$\tau \to \mu \gamma$	0.03
$(\delta_{23}^l)_{RL,LR}$	$r \kappa_5^l(\epsilon^2, 1)$	$\kappa_{10}^l(\epsilon^{2+2p},\epsilon^{2p})$		0.0015
$\boxed{ \left( \sqrt{ \mathrm{Re}(\delta_{12}^{\mathrm{d}})_{\mathrm{LL}}^2 }, \sqrt{ \mathrm{Im}(\delta_{12}^{\mathrm{d}})_{\mathrm{LL}}^2 } \right)}$	$\epsilon^6$	$\epsilon^6$		(0.065, 0.0052)
$\left(\sqrt{ \mathrm{Re}(\delta_{12}^{\mathrm{d}})_{\mathrm{RR}}^{2} },\sqrt{ \mathrm{Im}(\delta_{12}^{\mathrm{d}})_{\mathrm{RR}}^{2} }\right)$	-	$\epsilon^{1+2p}$		(0.065, 0.0052)
$\left(\sqrt{ \mathrm{Re}(\delta_{12}^{\mathrm{d}})_{\mathrm{LR}}^{2} },\sqrt{ \mathrm{Im}(\delta_{12}^{\mathrm{d}})_{\mathrm{LR}}^{2} }\right)$	$\kappa_5^d \epsilon^3$	$\kappa_{10}^d \epsilon^3$	$K - \overline{K}$	$(0.007, 5.2 \times 10^{-5})$
$\left(\sqrt{ \mathrm{Re}(\delta_{12}^{\mathrm{d}})_{\mathrm{RL}}^{2} },\sqrt{ \mathrm{Im}(\delta_{12}^{\mathrm{d}})_{\mathrm{RL}}^{2} }\right)$	$\kappa_5^d \epsilon^4$	$\kappa_{10}^d \epsilon^4$		$(0.007, 5.2 \times 10^{-5})$
$\sqrt{ \mathrm{Re}(\delta_{12}^{\mathrm{d}})_{\mathrm{LL}}(\delta_{12}^{\mathrm{d}})_{\mathrm{RR}} }$	-	$\epsilon^{3.5+p}$		0.00453
$\sqrt{ \mathrm{Im}(\delta_{12}^{\mathrm{d}})_{\mathrm{LL}}(\delta_{12}^{\mathrm{d}})_{\mathrm{RR}} }$	-	$\epsilon^{3.5+p}$		0.00057
$(\mathrm{Re}\delta_{13}^\mathrm{d},\mathrm{Im}\delta_{13}^\mathrm{d})_\mathrm{LL}$	$\epsilon^4$	$\epsilon^4$		(0.238, 0.51)
$(\mathrm{Re}\delta_{13}^\mathrm{d},\mathrm{Im}\delta_{13}^\mathrm{d})_\mathrm{RR}$	_	$\epsilon^{1+2p}$	$B_d - \overline{B}_d$	(0.238, 0.51)
$(\mathrm{Re}\delta_{13}^\mathrm{d},\mathrm{Im}\delta_{13}^\mathrm{d})_{\mathrm{LR},\mathrm{RL}}$	$\kappa_5^d(\epsilon^4,\epsilon)$	$\kappa^d_{10}(\epsilon, \epsilon^4)$		(0.0557, 0.125)
$(\delta^d_{23})_{LL}$	$\epsilon^2$	$\epsilon^2$		1.19
$(\delta^d_{23})_{RR}$	-	$\epsilon^{2p}$	$B_s - \overline{B}_s$	1.19
$(\delta^d_{23})_{LR,RL}$	$\kappa_5^d(1,\epsilon^2)$	$\kappa^d_{10}(1,\epsilon^2)$	$b \to s \gamma$	0.04

# **Summary**

- $m_h = 125$  GeV can be naturally realized in minimal GMSB models with messenger—matter mixing
- Relatively light SUSY spectrum obtained