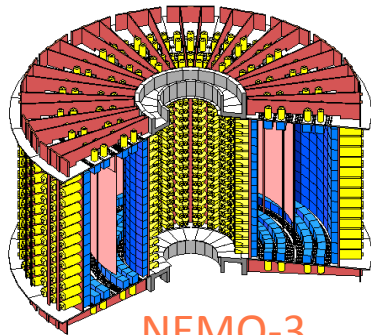


Latest Results from **NEMO-3**
and
Status of **SuperNEMO**

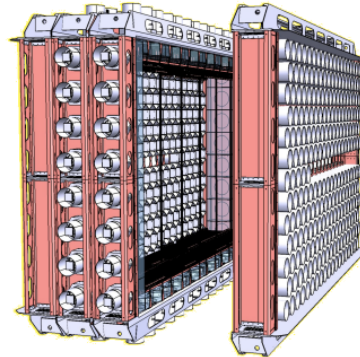
Karol Lang

University of Texas at Austin

On behalf of the NEMO Collaboration



NEMO-3



SuperNEMO

Outline:

1. $0\nu\beta\beta$ fundamentals
2. NEMO technique
3. NEMO-3 results
4. Status of SuperNEMO



36th International Conference
on High Energy Physics

4 – 11 July 2012
Melbourne Convention and Exhibition Centre

Fundamentals of $0\nu\beta\beta$ and $2\nu\beta\beta$

$$\frac{1}{T_{1/2}^{2\nu}} = G_{2\nu}(Q_{\beta\beta}^{11}, Z) \cdot |M_{2\nu}|^2$$

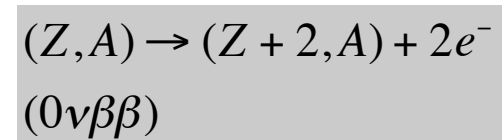
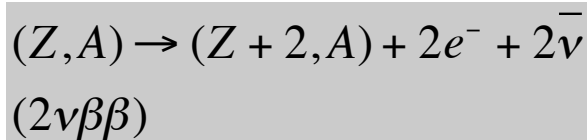
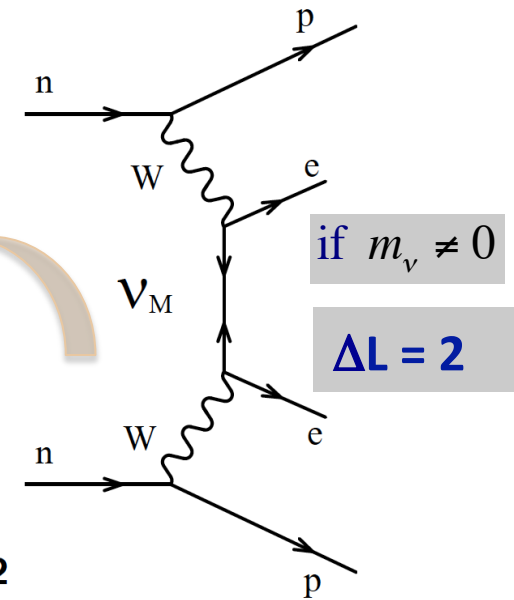
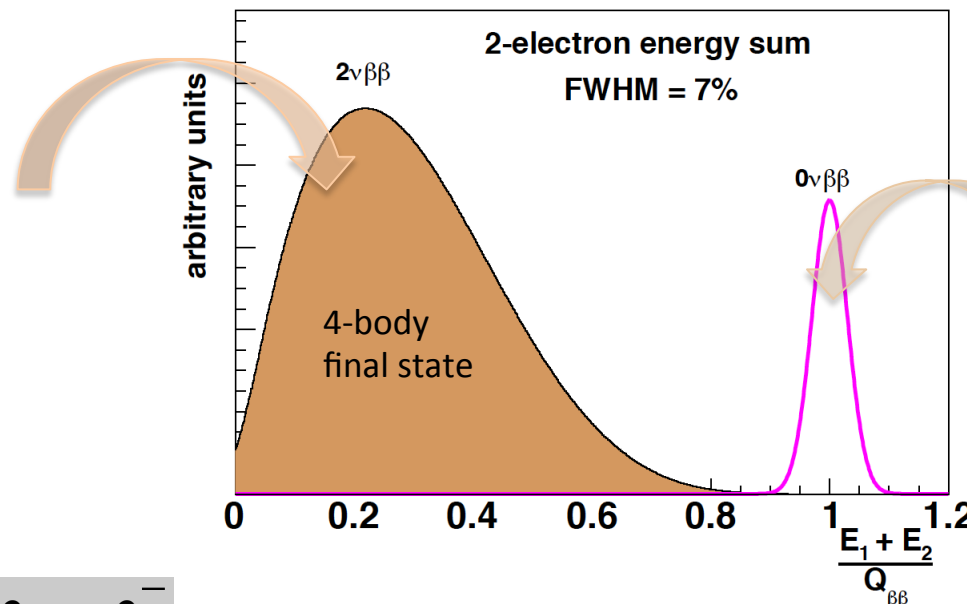
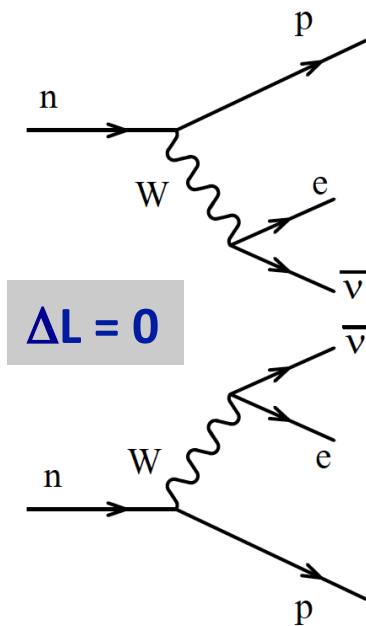
G = phase space (well known)

M = nuclear matrix element (challenging)

$$\frac{1}{T_{1/2}^{0\nu}} = G_{0\nu}(Q_{\beta\beta}^5, Z) \cdot |M_{0\nu}|^2 \cdot \langle m_{\beta\beta} \rangle^2$$

$$\langle m_{\beta\beta} \rangle \equiv \left| m_1 |U_{e1}|^2 + m_2 |U_{e2}|^2 e^{i\alpha^*} + m_3 |U_{e3}|^2 e^{i\beta^* - 2i\delta} \right|$$

α^*, β^* = linear combinations of α and β

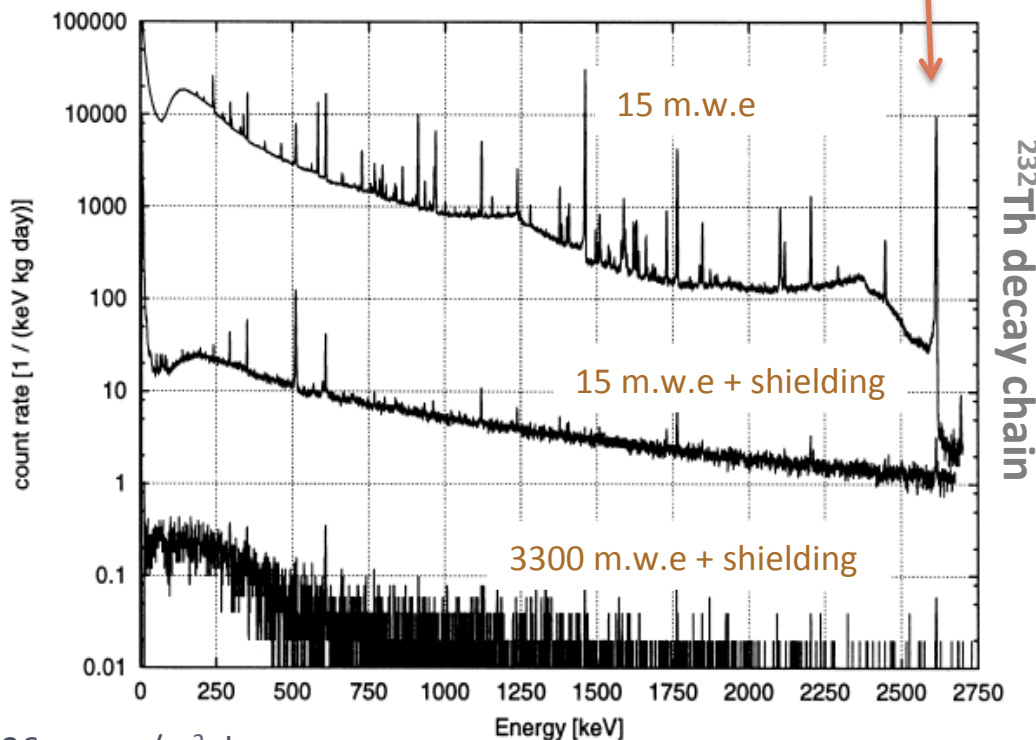




Practical fundamentals

- ◆ **Natural radioactivity and cosmic rays** dominate the backgrounds → go underground + local shielding
- ◆ **^{238}U and ^{232}Th decay chains** produce the most troubling gammas (highest energies):
 - ^{214}Bi
 - ^{208}Tl

2×10^6 muons/m² day on surface



26 muons/m² day

(Applied Rad and Isotopes 53 (2000) 191)

NEMO-3	$Q_{\beta\beta}$ (MeV)	Natural abundance (%)
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	4.272	0.187
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	3.367	5.6
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	3.350	2.8
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	3.035	9.6
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	2.995	9.2
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	2.805	7.5
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	2.529	34.5
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	2.458	8.9
$^{124}\text{Sn} \rightarrow ^{124}\text{Te}$	2.228	5.64
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	2.039	7.8
$^{110}\text{Pd} \rightarrow ^{110}\text{Cd}$	2.013	11.8

Top 11 $\beta\beta$ emitters with $Q_{\beta\beta} > 2$ MeV

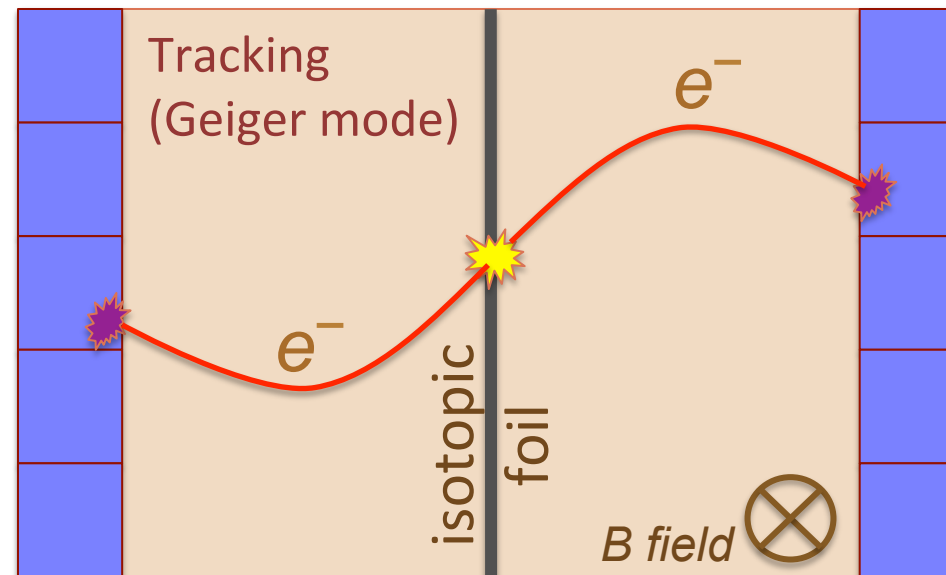
Challenge:

- ✓ suppress backgrounds
- ✓ identify the final state

The NEMO-3 Technique



The multi-observable principle:
topology, kinematics, timing



Plastic
scintillator
calorimeter

Radio-pure materials
and a multi-layer shielding

NEMO-3 detector

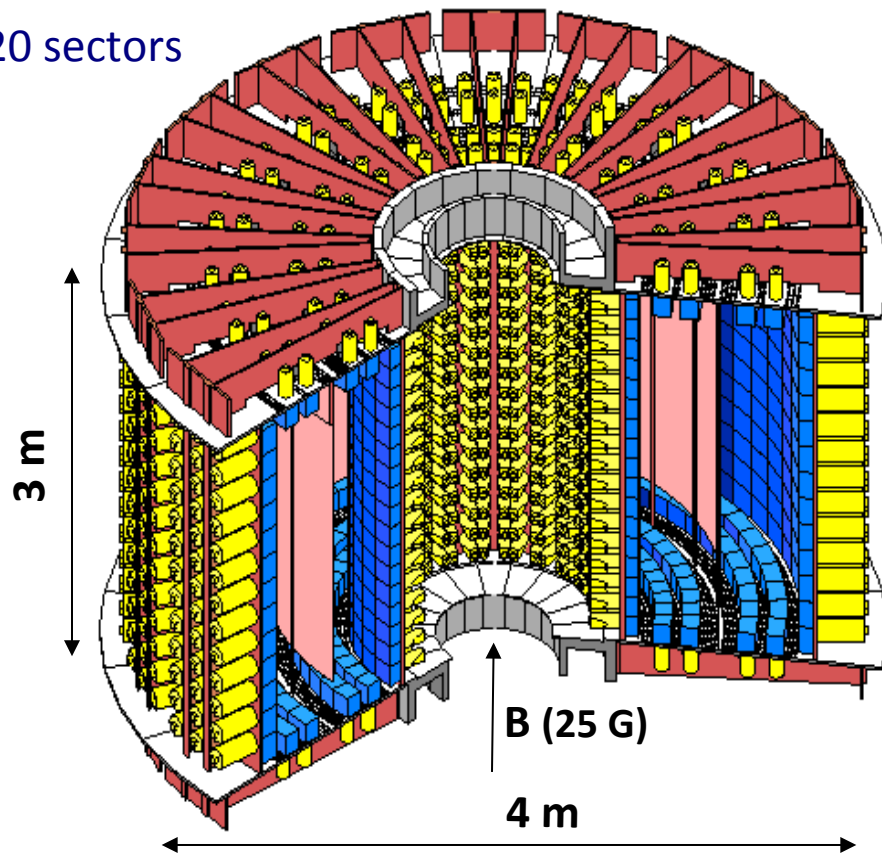


Fréjus Tunnel : 4,800 m.w.e.

Phase 1: Feb, 2003 → Sep, 2004

Phase 2: Oct, 2004 → Jan, 2011

20 sectors

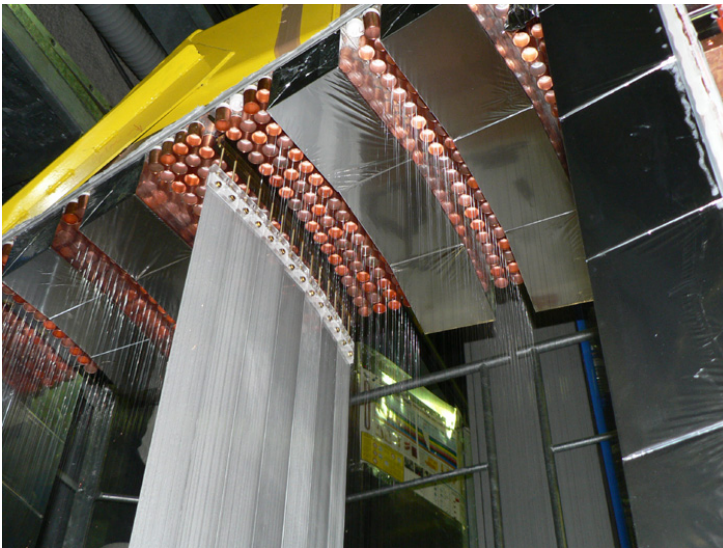
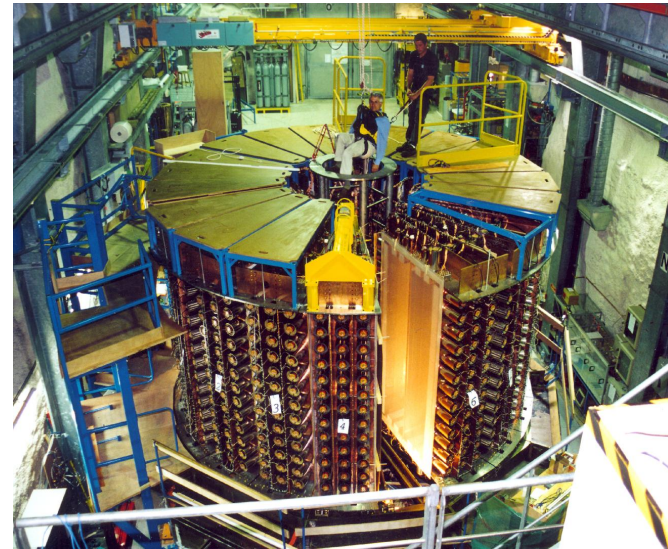
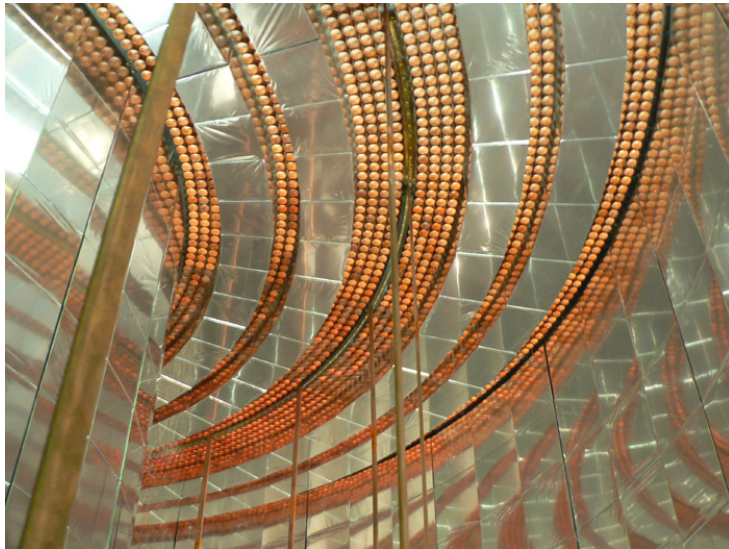


- ✓ Source: 10 kg of $\beta\beta$ isotopic foils
area = 20 m², thickness \sim 60 mg/cm²
- ✓ Tracking detector:
drift wire chamber (9 layers)
in Geiger mode (6180 cells)
Gas: He + 4% ethyl alcohol + 1% Ar + 0.1% H₂O
- ✓ Calorimeter:
1940 plastic scintillators
low radioactivity 3'' & 5'' PMTs
- ✓ B field : 25 Gauss
- ✓ Shielding:
gamma shield: pure iron (d = 18cm)
neutron shield:
30 cm water (ext. wall)
40 cm wood (top / bottom)
(since March 2004: water + boron)



Particle ID: e^- , e^+ , γ and α

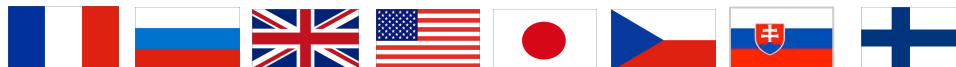
NEMO-3 data taking: 2003 - 2010



Laboratoire Souterrain de Modane (LSM) (Frejus Tunnel)

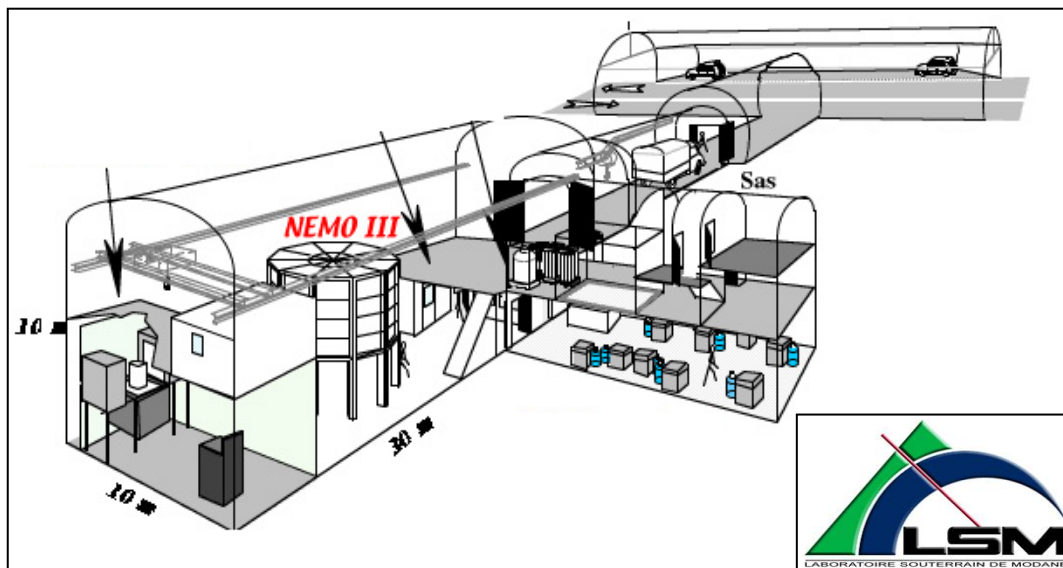
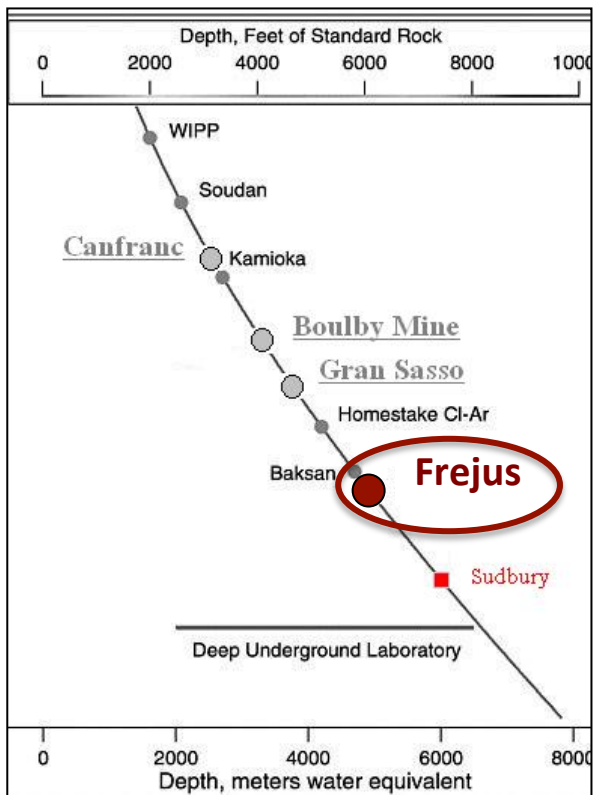
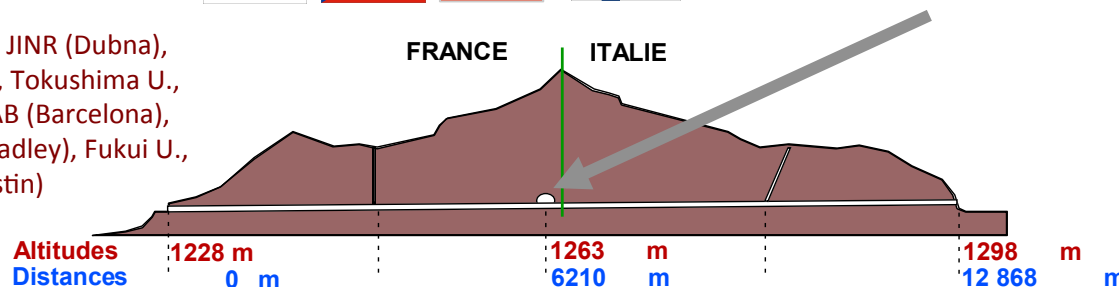


NEMO Collaboration



4,800 m.w.e

LAL (Orsay), IPHC (Strasbourg), INL (Idaho Falls), ITEP (Moscow), JINR (Dubna), LPC (Caen), CENBG (Bordeaux), UCL (London), U. of Manchester, Tokushima U., Cornelius U. (Bratislava), Osaka, IEAP & Charles U. (Prague), UAB (Barcelona), Saga U., Imperial College (London), Mount Holyoke Coll. (South Hadley), Fukui U., INR (Kiev), CPPM (Marseilles), U. Warwick, Texas (Austin)

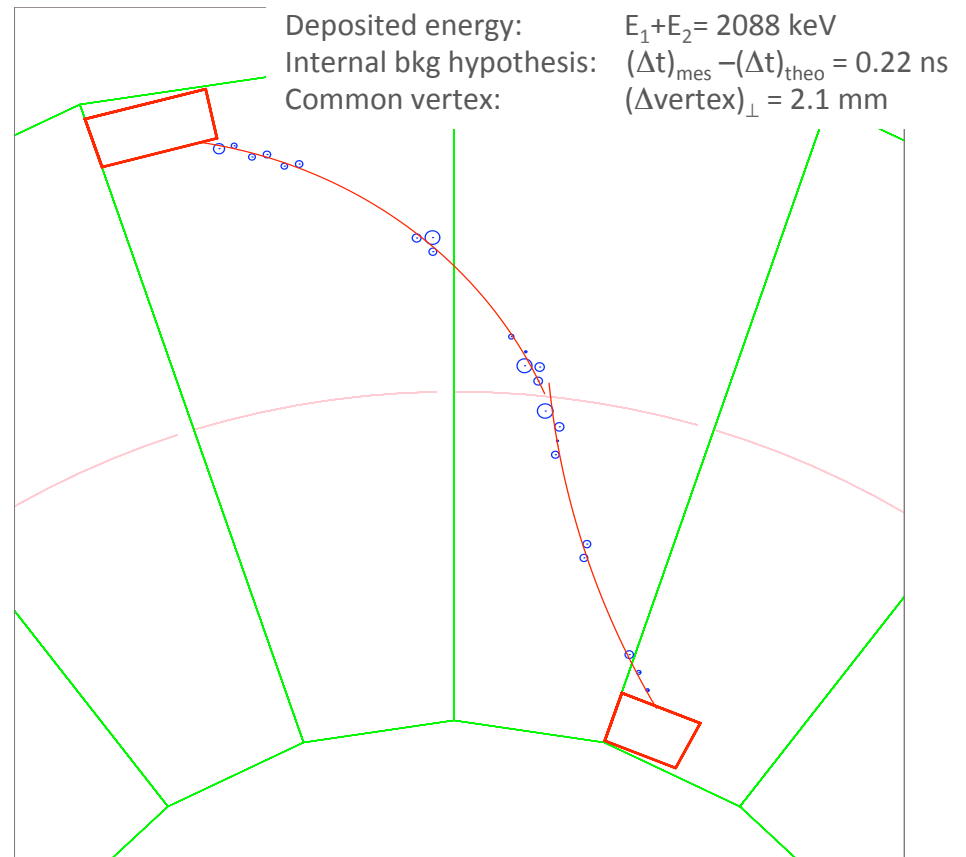
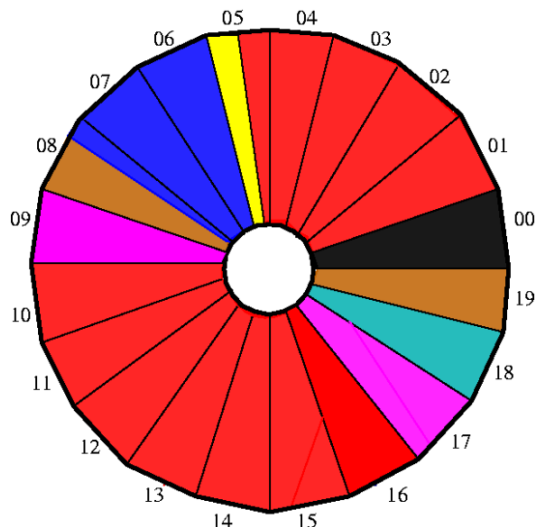


Built for τ experiment (proton decay) in 1981-1982



NEMO-3: 7 isotopes + events images

Isotope	Mass (g)	$Q_{\beta\beta}$ (keV)
^{100}Mo	6 914	3035
^{82}Se	932	2995
^{116}Cd	405	2805
^{96}Zr	9.4	3350
^{150}Nd	37	3367
^{48}Ca	7	4272
^{130}Te	454	2529
natTe	491	
natCu	621	

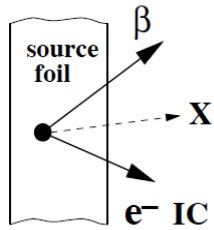


- ✓ Trigger: at least 1 PMT > 150 keV
 ≥ 3 Geiger hits (2 neighbouring layers+1)
- ✓ Trigger rate = 7 Hz
- ✓ 25 $\beta\beta$ events per hour

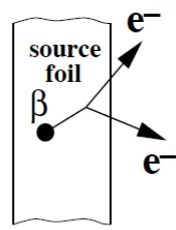
NEMO-3 backgrounds

1. Internal background (in addition to a potential $2\nu\beta\beta$ tail)

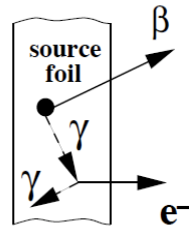
(due to ^{232}Th (^{208}Tl) and ^{238}U (^{214}Bi) radio-impurities of the isotopic source foil)



beta + IC



beta + Möller



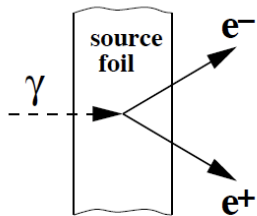
beta + Compton

(dominant)

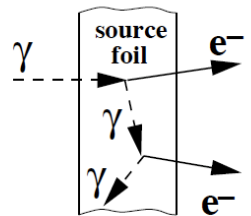
● = radioisotope β = electron from beta decay IC = internal conversion

2. External background (if the γ is not detected)

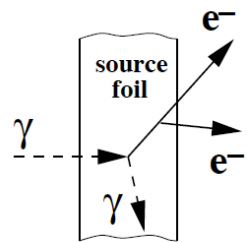
(due to radio-impurities of the detector)



pair creation



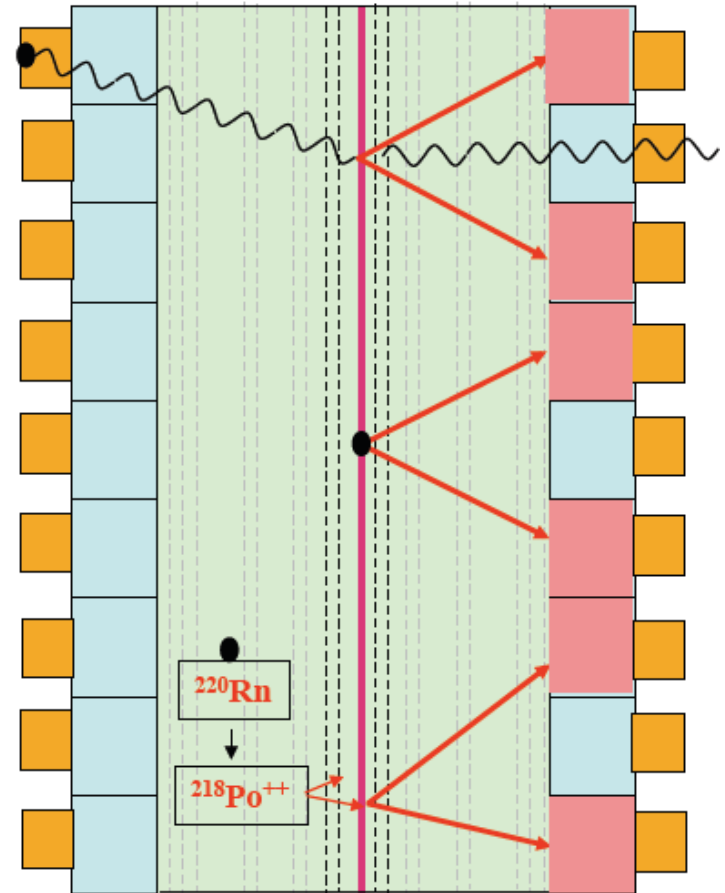
Compton + Compton



Compton + Möller

3. Radon (^{214}Bi) inside the tracking detector

- deposits on the wire near the $\beta\beta$ foil
- deposits on the surface of the $\beta\beta$ foil

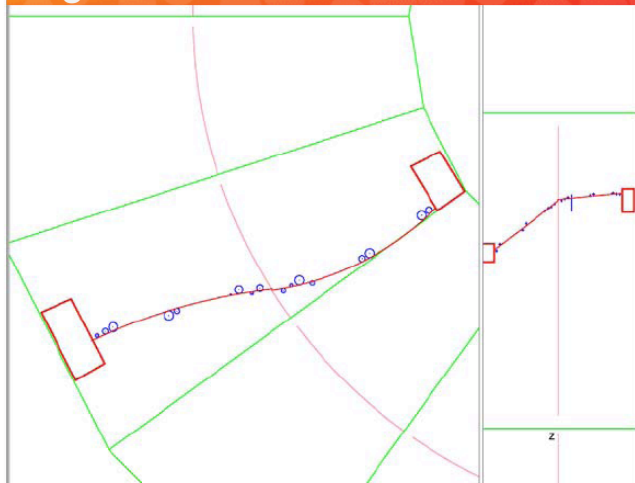


Each bkg is measured using the NEMO-3 data

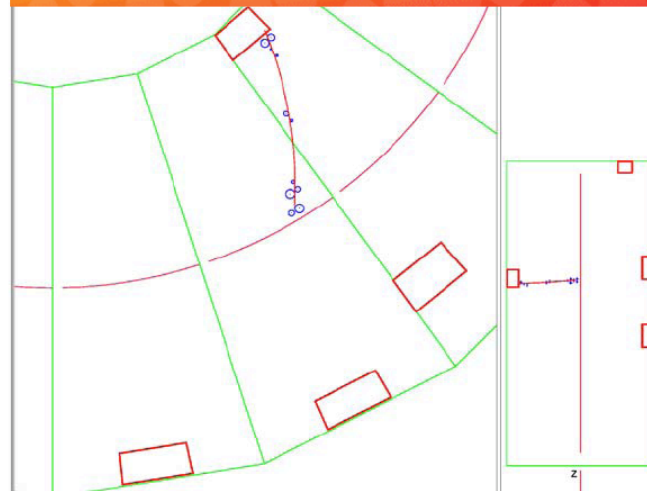
Signal and background signatures



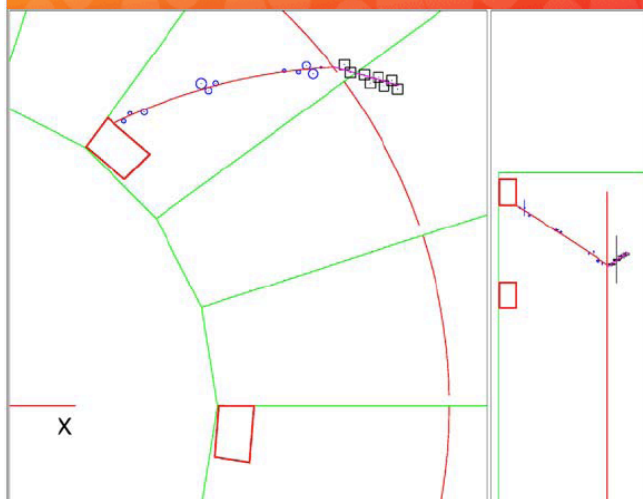
$2e^-$ event
signal



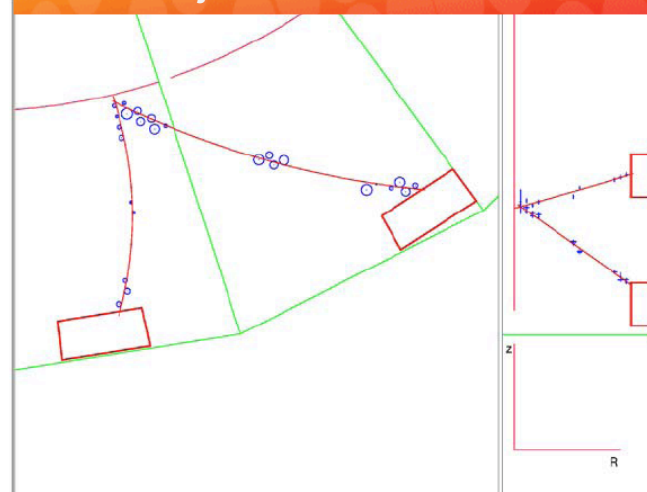
$e^- \gamma \gamma$ event
measure ^{208}Tl



$\beta - \alpha$ (delay track) event
 $^{214}\text{Bi} \rightarrow ^{214}\text{Po} \rightarrow ^{210}\text{Pb}$

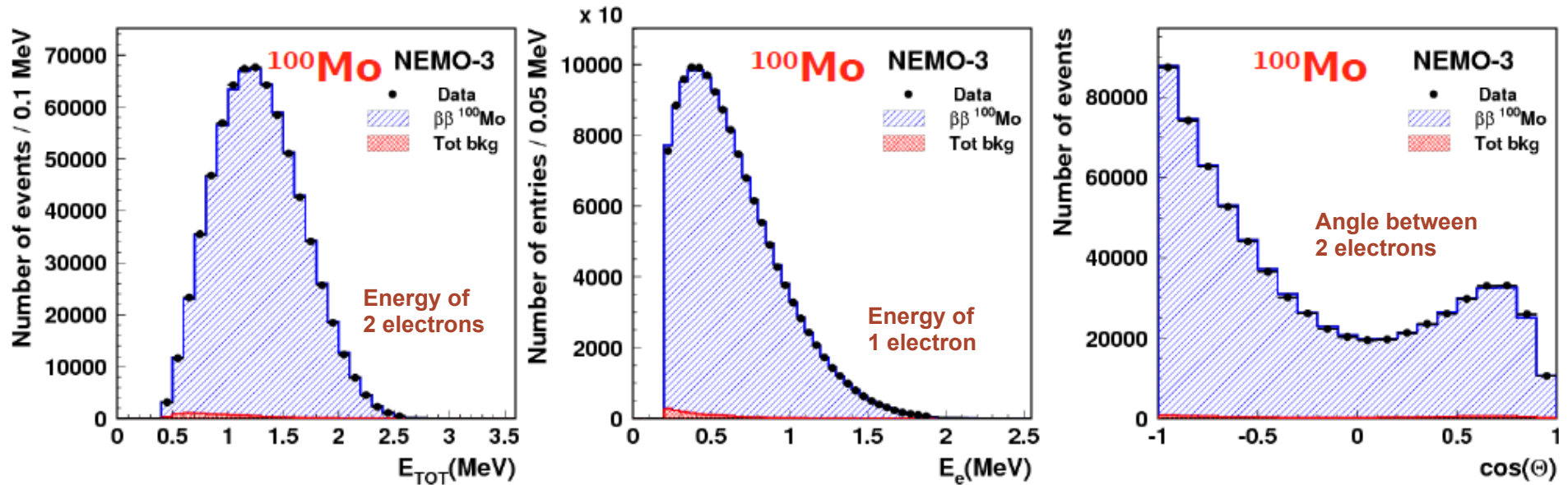


$e^+ - e^-$ pair event
B field rejection



NEMO-3 flagship measurements

$2\nu\beta\beta$ results (not final)



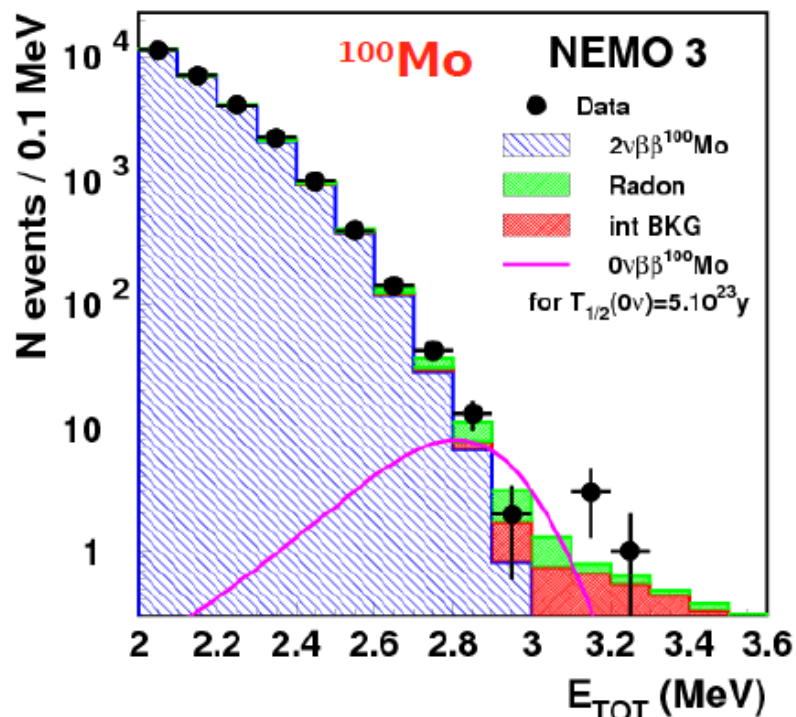
> 700 000 of 2-electron

Signal/Background : 76

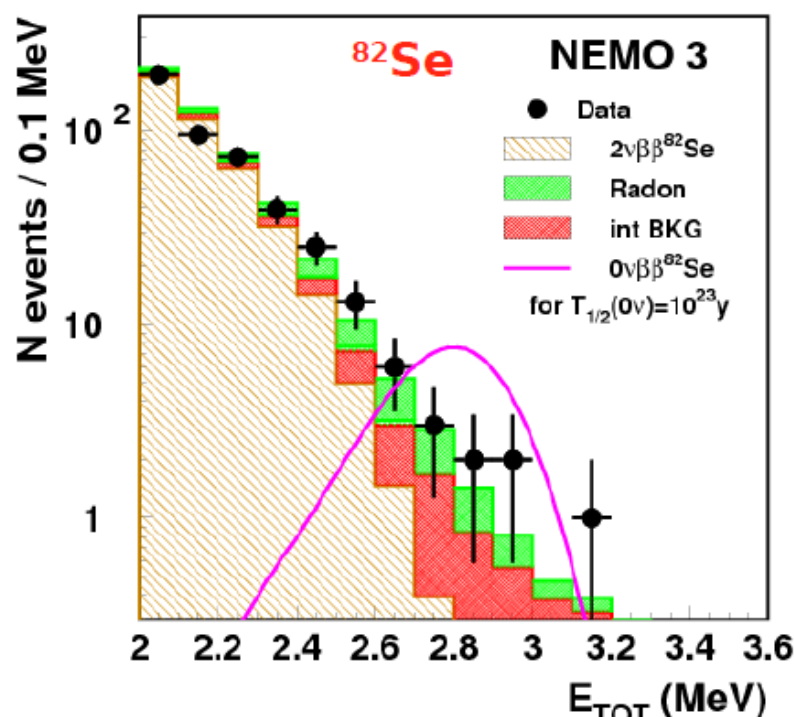
$$T_{1/2} (2\nu\beta\beta) = (7.16 \pm 0.01) \times 10^{18} \text{ y} \quad (\text{preliminary})$$

published Phase 1 $T_{1/2} = [7.11 \pm 0.02 \text{ (stat)} \pm 0.54 \text{ (sys)}] \times 10^{18} \text{ y}$

NEMO-3 flagship measurements: ^{100}Mo and ^{82}Se $0\nu\beta\beta$ results (not final)



[2.8 – 3.2] MeV 18 observed events, 16.4 ± 1.3 expected



[2.6 – 3.2] MeV 14 observed events, 11.3 ± 1.3 expected

^{100}Mo (for exposure of $31.2 \text{ kg} \cdot \text{y}$)

$$T_{1/2}(0\nu\beta\beta) > 1.0 \times 10^{24} \text{ y (90\% C.L.)}$$

$$m_{\beta\beta} < 0.31 - 0.96 \text{ eV}$$

^{82}Se (for exposure of $4.2 \text{ kg} \cdot \text{y}$)

$$T_{1/2}(0\nu\beta\beta) > 3.2 \times 10^{23} \text{ y (90\% C.L.)}$$

$$m_{\beta\beta} < 0.94 - 2.6 \text{ eV}$$

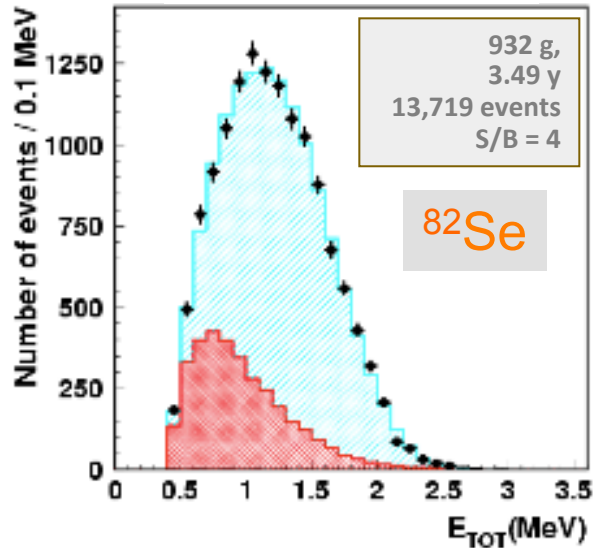
[1] QRPA M.Kortelainen and J.Suhonen, Phys.Rev. C 75 (2007) 051303(R)
[2] QRPA M.Kortelainen and J.Suhonen, Phys.Rev. C 76 (2007) 024315

[3] QRPA F.Simkovic, et al. Phys.Rev. C 77 (2008) 045503
[4] IBM2 J.Barrea and F.Iachello Phys.Rev.C 79(2009)044301

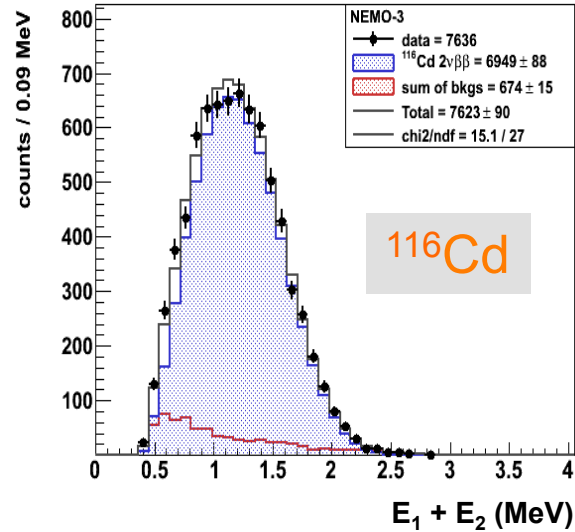
PHFB [5] P.K. Rath et al., Phys. Rev. C 82 (2010) 064310
SM [6] E.Caurrier et al. Phys.Rev.Lett 100 (2008) 052503



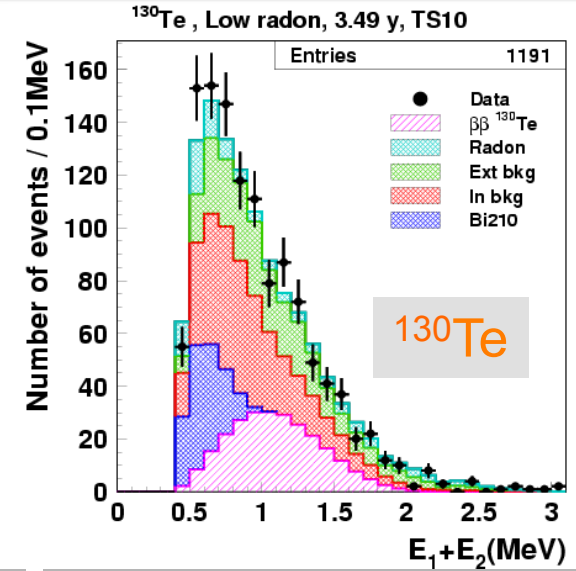
NEMO-3 $2\nu\beta\beta$ results (not final)



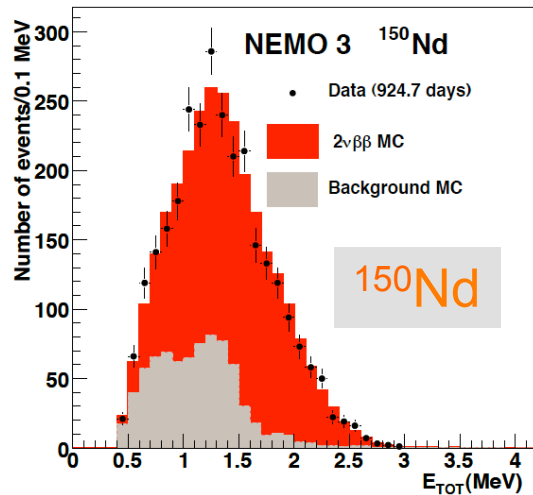
$$[9.6 \pm 0.1_{\text{(stat)}} \pm 1.0_{\text{(syst)}}] \times 10^{19} \text{ y}$$



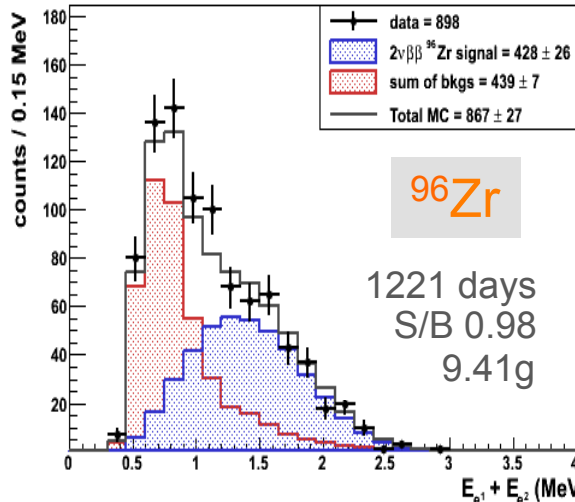
$$[2.88 \pm 0.04_{\text{(stat)}} \pm 0.16_{\text{(syst)}}] \times 10^{19} \text{ y}$$



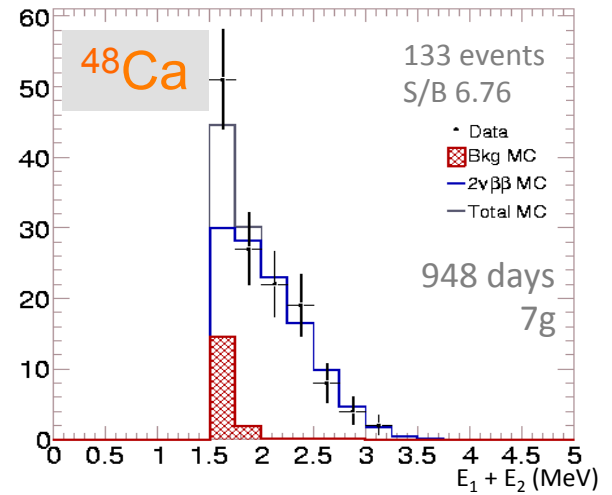
$$[7.0 \pm 0.9_{\text{(stat)}} \pm 1.1_{\text{(syst)}}] \times 10^{20} \text{ y}$$



$$[9.11^{+0.25}_{-0.22} \text{(stat)} \pm 0.63_{\text{(syst)}}] \times 10^{18} \text{ y}$$



$$[2.35 \pm 0.14_{\text{(stat)}} \pm 0.16_{\text{(syst)}}] \times 10^{19} \text{ y}$$



$$[4.4^{+0.5}_{-0.4} \text{(stat)} \pm 0.4_{\text{(syst)}}] \times 10^{19} \text{ y}$$

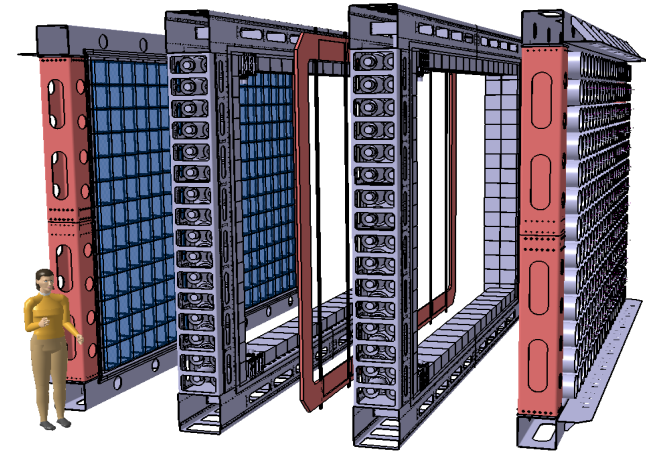
NEMO-3 → SuperNEMO



20 wedges: 10 kg of 7 isotopes



20 planar modules, each w/ 5-7 kg
Can do different isotopes & locations



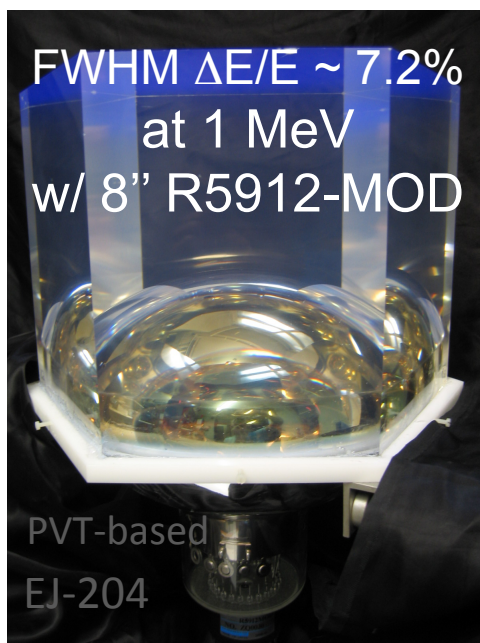
NEMO-3	→	SuperNEMO
^{100}Mo , ^{82}Se , ^{150}Nd , ^{130}Te , ^{116}Cd , ^{96}Zr , ^{48}Ca		^{82}Se , ... ^{150}Nd , ^{48}Ca
7 kg		100 - 200 kg (7 kg Demonstrator)
$T_{1/2}(0\nu\beta\beta) > 2 \times 10^{24} \text{ y}$		$T_{1/2}(0\nu\beta\beta) > 1 \times 10^{26} \text{ y}$
$m_{\beta\beta} < 0.3 - 0.8 \text{ eV}$		$m_{\beta\beta} < 40 - 100 \text{ meV}$

SuperNEMO R&D and Demonstrator construction

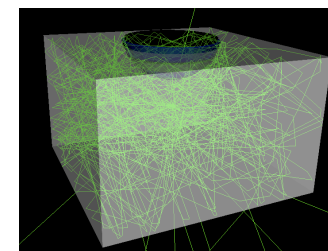


- ❑ First SuperNEMO module (Demonstrator) under construction (through 2014)
- ❑ (Most of) R&D completed; optimizing costs, some systems under full production
- ❑ Major improvements in:
 - energy resolution, source foil radio-purity measurements, radon emanation and hermeticity, source installation, ^{207}Bi calibration, tracking, software, ...
- ❑ Some highlights below:

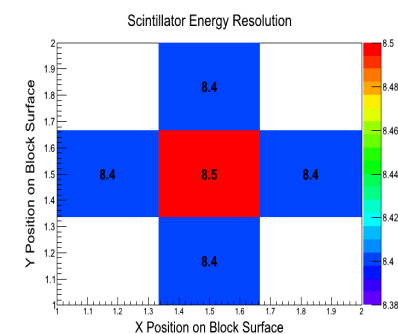
Main walls blocks



Calorimeter:
main walls
and gamma veto walls



G4 simulations → agreement
with measurements

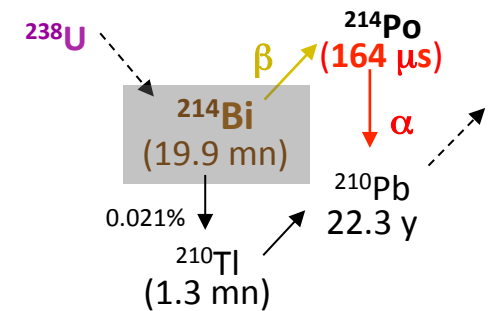
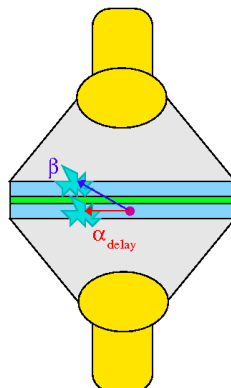
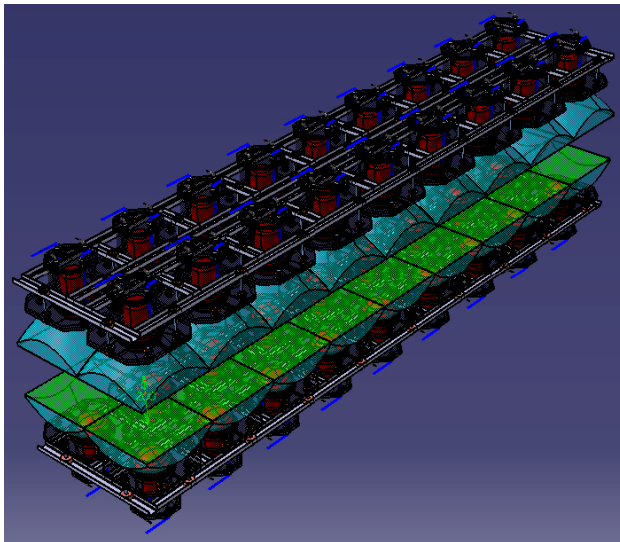
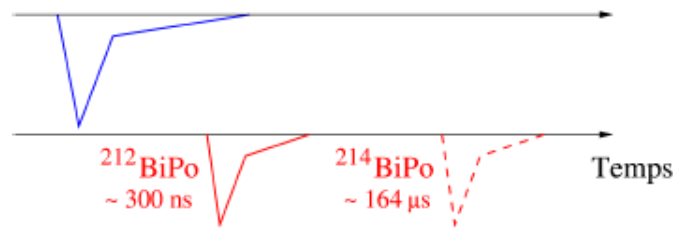
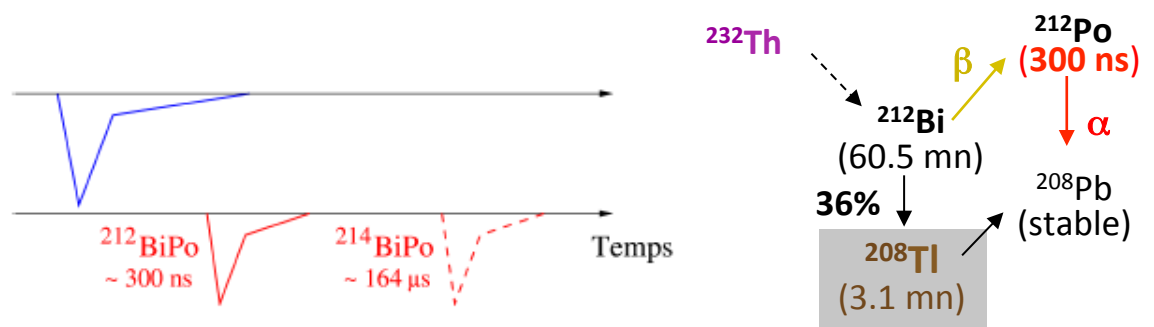


SuperNEMO Demonstrator construction (2)



- BiPo detector - to test the source foil radio-purity
- Sensitivity in 6 months
 - ^{208}Tl : $\sim 5 \mu\text{Bq} / \text{kg}$
 - ^{214}Bi : $\sim 15 \mu\text{Bq} / \text{kg}$
- Will operate at the Canfranc Lab

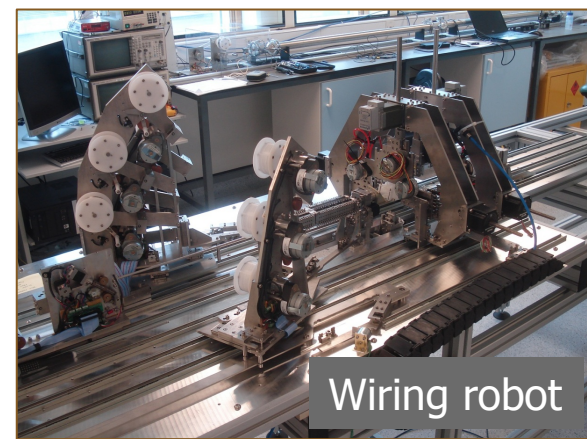
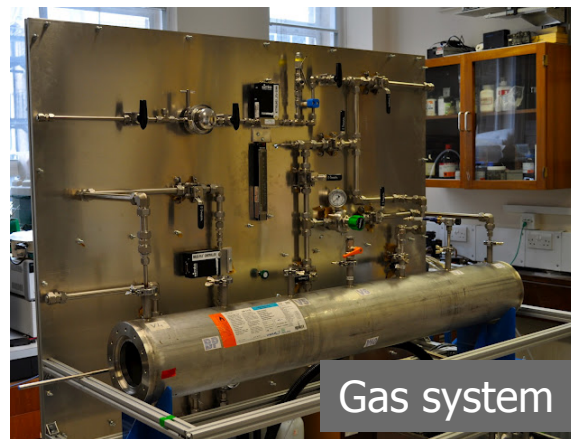
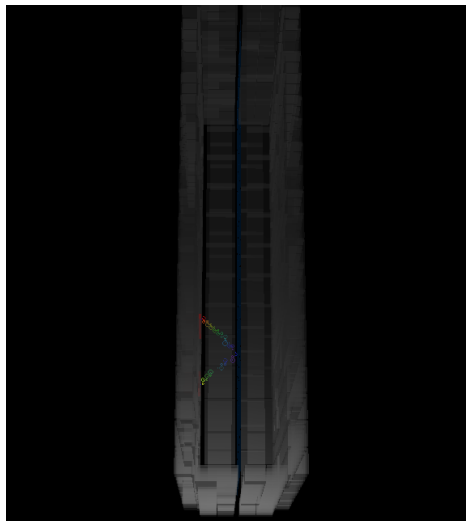
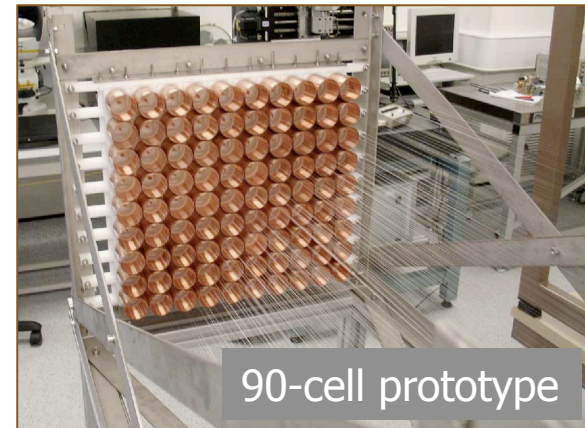
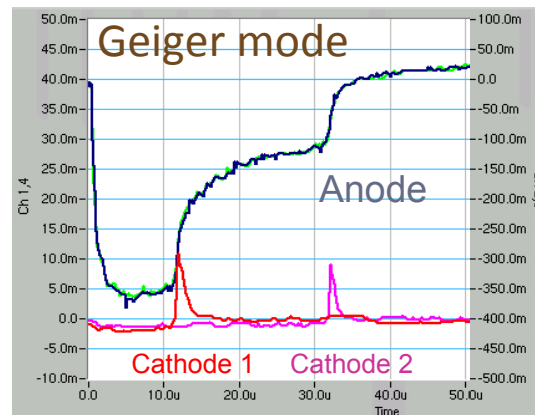
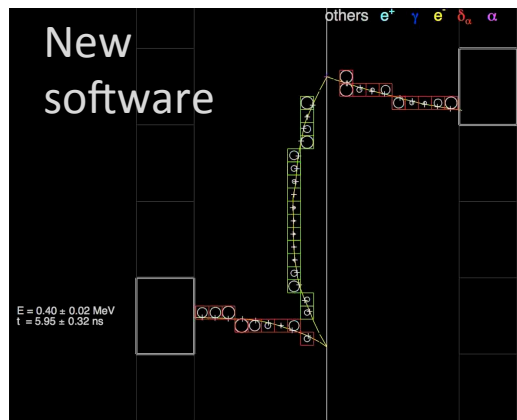
Bi \rightarrow Po \rightarrow Pb





Tracker construction (automatic wiring)

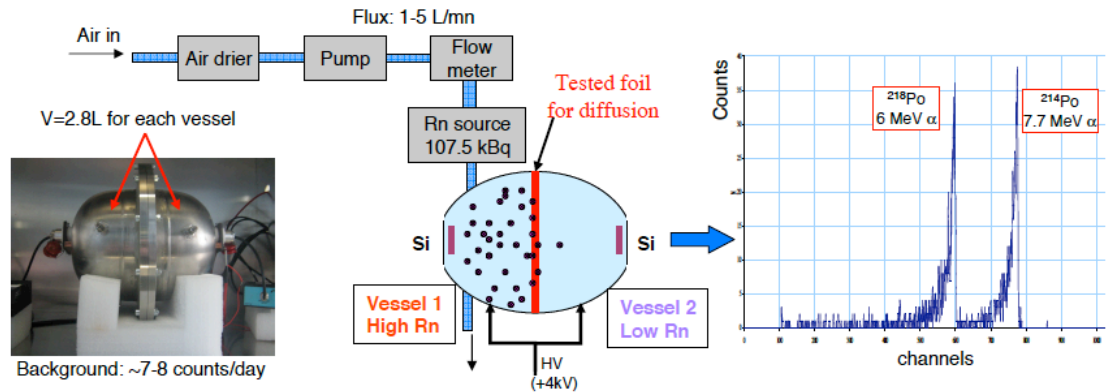
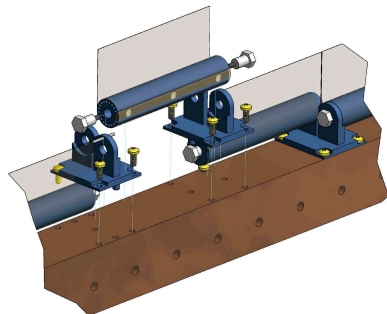
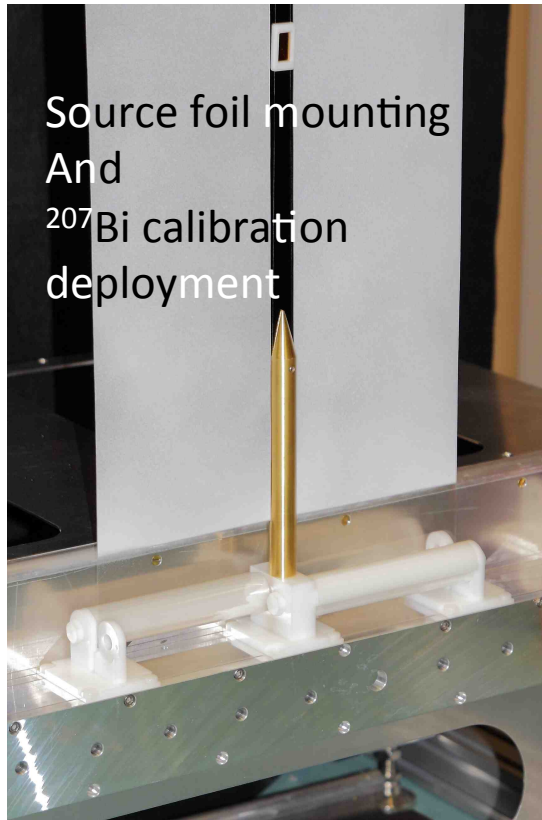
He-Ar (w/ isoprop.)
or
He-Ne (w/ isoprop.)



SuperNEMO Demonstrator construction (4)



Material radon emanation screening



Radon concentration line (sensitivity <math>< 0.15 \text{ mBq/m}^3</math>)



Summary

◆ NEMO-3 produces a unique spectrum of results on $0\nu\beta\beta$ and $2\nu\beta\beta$

$$\begin{aligned}
 {}^{100}\text{Mo}: T_{1/2}^{0\nu\beta\beta} &> 1.0 \times 10^{24} \text{ y (90\% CL)} & m_{\beta\beta} &< (0.31 - 0.96) \text{ eV} \\
 {}^{82}\text{Se}: T_{1/2}^{0\nu\beta\beta} &> 3.6 \times 10^{23} \text{ y (90\% CL)} & m_{\beta\beta} &< (0.94 - 2.6) \text{ eV}
 \end{aligned}$$

- ✓ Results for 5 other isotopes: ${}^{48}\text{Ca}$, ${}^{96}\text{Zr}$, ${}^{116}\text{Cd}$, ${}^{130}\text{Te}$, ${}^{150}\text{Nd}$
- ✓ Testing transitions models: excited states, V+A, Majorons, SSD, HSD, ...
- ✓ Stay tuned for results with final samples

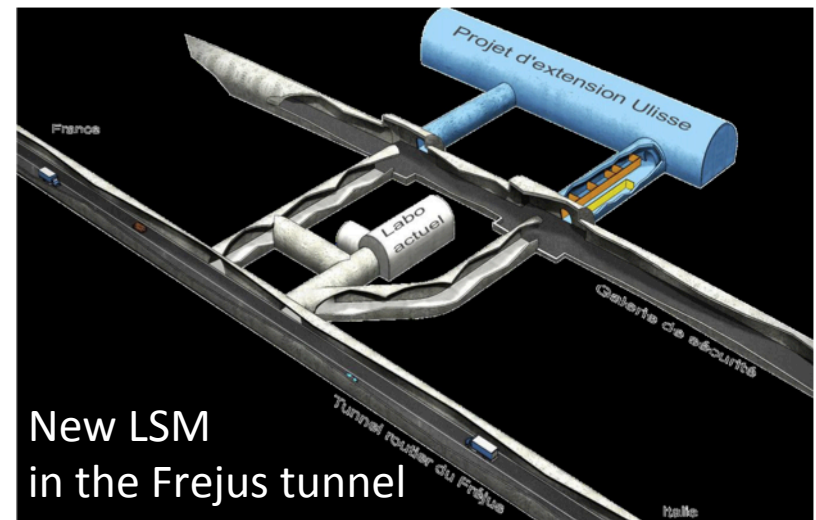
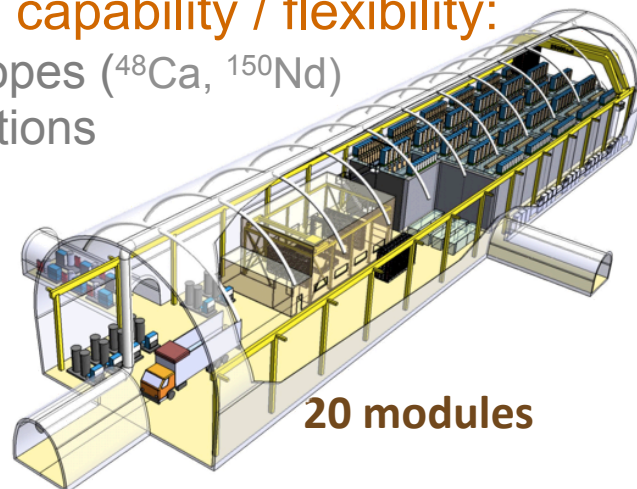
◆ SuperNEMO (Demonstrator in 2014)

- ✓ ${}^{82}\text{Se}$; sensitivity:

$$T_{1/2}(0\nu\beta\beta) = (1 - 2) \times 10^{26} \text{ y (500 kg*y exposure)} \rightarrow m_{\beta\beta} \leq 40 - 140 \text{ meV}$$

◆ SuperNEMO capability / flexibility:

- ✓ Other isotopes (${}^{48}\text{Ca}$, ${}^{150}\text{Nd}$)
- ✓ Other locations





NEMO Collaboration, Chateau d'Yquem, June'2012



BACKUP SLIDES FOR NEMO-3 & SUPERNEMO

Summary of $2\nu\beta\beta$ NEMO-3 results



Isotope	Exposure [days]	$T_{1/2}(2\nu\beta\beta)$ [years]
^{100}Mo	389 [Phase 1 ⁽¹⁾] 1,468 [Phase 2]	[7.11 ± 0.02 (stat) ± 0.54 (syst)] $\times 10^{18}$ [7.16 ± 0.02 (stat) ± 0.54 (syst)] $\times 10^{18}$ (2011)
^{82}Se	389 [Phase 1 ⁽¹⁾] 1,468 [Phase 2]	[9.6 ± 0.3 (stat) ± 1.0 (syst)] $\times 10^{18}$ [9.6 ± 0.1 (stat) ± 1.0 (syst)] $\times 10^{18}$ (2011)
^{150}Nd	925 [Phase 1 + 2 ⁽²⁾]	[$9.11^{+0.25}_{-0.22}$ (stat) ± 0.63 (syst)] $\times 10^{18}$
^{130}Te	1,275 [Phase 2 ⁽³⁾]	[7.0 ± 0.9 (stat) ± 1.1 (syst)] $\times 10^{20}$
^{116}Cd	1,471 [Phase 1 + 2]	[2.88 ± 0.04 (stat) ± 0.16 (syst)] $\times 10^{19}$
^{96}Zr	1,221 [Phase 1 + 2 ⁽⁴⁾]	[2.35 ± 0.14 (stat) ± 0.16 (syst)] $\times 10^{19}$
^{48}Ca	943 [Phase 1 + 2]	[$4.4^{+0.5}_{-0.4}$ (stat) ± 0.4 (syst)] $\times 10^{19}$

(1) R. Arnold et al., Phys. Rev. Lett. 95 182302 (2005)

(2) J. Argyriades et al., Phys. Rev. C 80, 032501R (2009)

(3) R. Arnold et al., PRL 107, 062504 (2011)

(4) J. Argyriades et al., Nucl. Phys. A 847, 168 (2010)

Phase 1: Feb, 2003 \rightarrow Sep, 2004

Phase 2: Oct, 2004 \rightarrow Jan, 2011

Summary of $0\nu\beta\beta$ NEMO-3 results



Isotope	Exposure [kg·y]	$T_{1/2}(0\nu\beta\beta)$ [years]	$\langle m_\nu \rangle$ [eV]	NME reference
^{100}Mo	31.2	$> 1.0 \cdot 10^{24}$	$< 0.31 - 0.96$	2-5,8
^{82}Se	4.2	$> 3.2 \cdot 10^{23}$	$< 0.94 - 1.6$ < 2.6	1,3-5,8 7
^{150}Nd	0.095	$> 1.8 \cdot 10^{22}$	$< 1.5 - 6.8$	4-6,8
^{130}Te	1.6	$> 1.3 \cdot 10^{23}$	$< 1.3 - 2.7$ < 3.6	2-5,8 7
^{116}Cd	1.65	$> 1.3 \cdot 10^{23}$	$< 1.3 - 3.2$	2-5,8
^{96}Zr	0.031	$> 9.2 \cdot 10^{21}$	$< 7.2 - 19.5$	2-5,8
^{48}Ca	0.017	$> 1.3 \cdot 10^{22}$	< 29.6	7

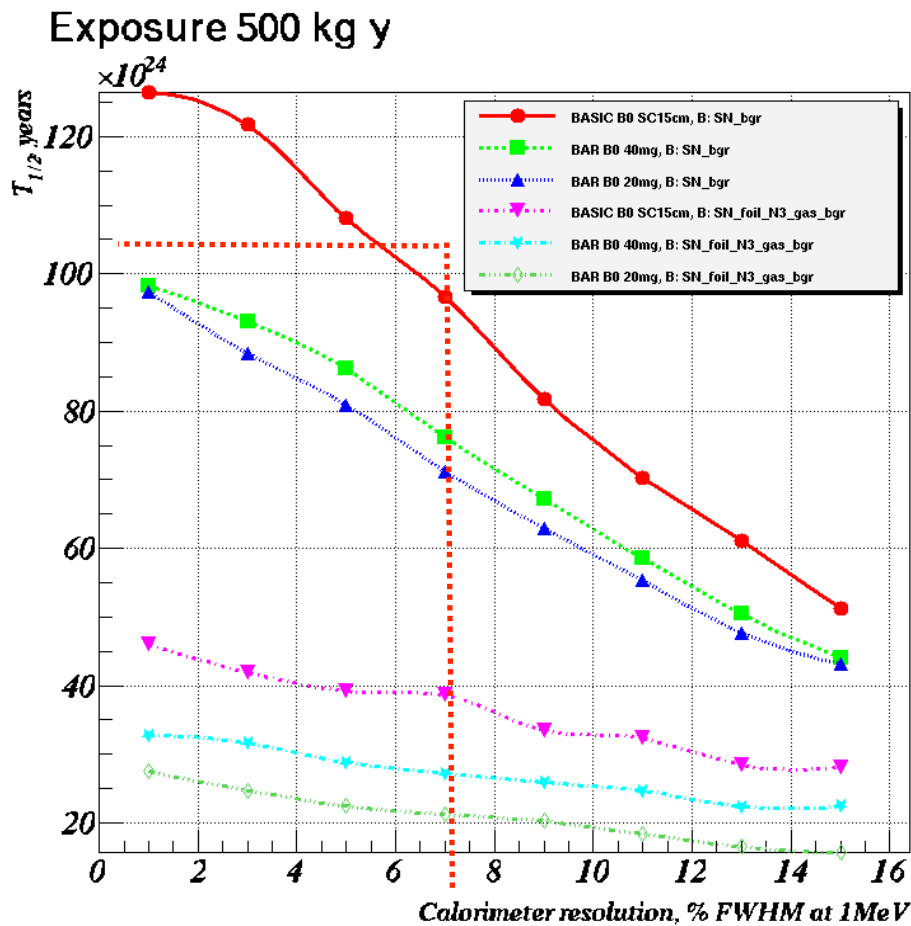
Nuclear Matrix Elements references:

[1] M.Kortelainen and J.Suhonen, *Phys.Rev. C* 75 (2007) 051303(R)
 [2] M.Kortelainen and J.Suhonen, *Phys.Rev. C* 76 (2007) 024315
 [3] F.Simkovic, et al., *Phys.Rev. C* 77 (2008) 045503
 [4] V.A. Rodin et al., *Nucl.Phys. A* 793 (2007) 213

[5] V.A. Rodin et al., *Nucl.Phys. A* 766(2006) 107
 [6] J.H.Hirsh et al., *Nucl.Phys. A* 582(1995) 124
 [7] E.Caurrier et al., *Phys.Rev.Lett* 100 (2008) 052503
 [8] P.K. Rath et al., *Phys. Rev. C* 82 (2010) 064310

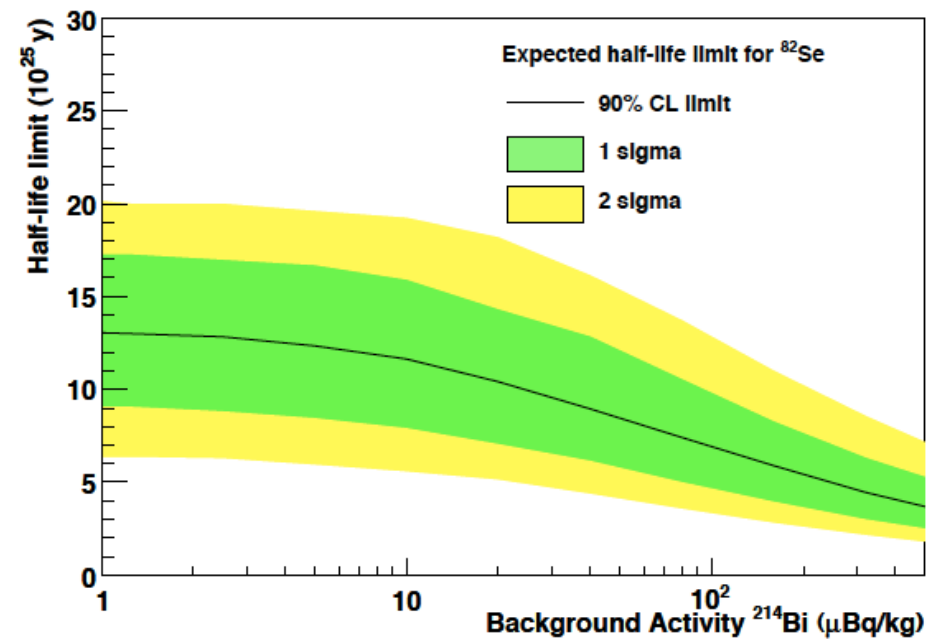


^{82}Se Sensitivity ($0\nu\beta\beta$)



Calorimeter resolution (% FWHM at 1 MeV)

GEANT-4 based model of the detector combined with NEMO-3 experience.



^{82}Se sensitivity
Exposure 500 kg * y

✓ $T_{1/2}(0\nu\beta\beta) > 2 \times 10^{26}$ y

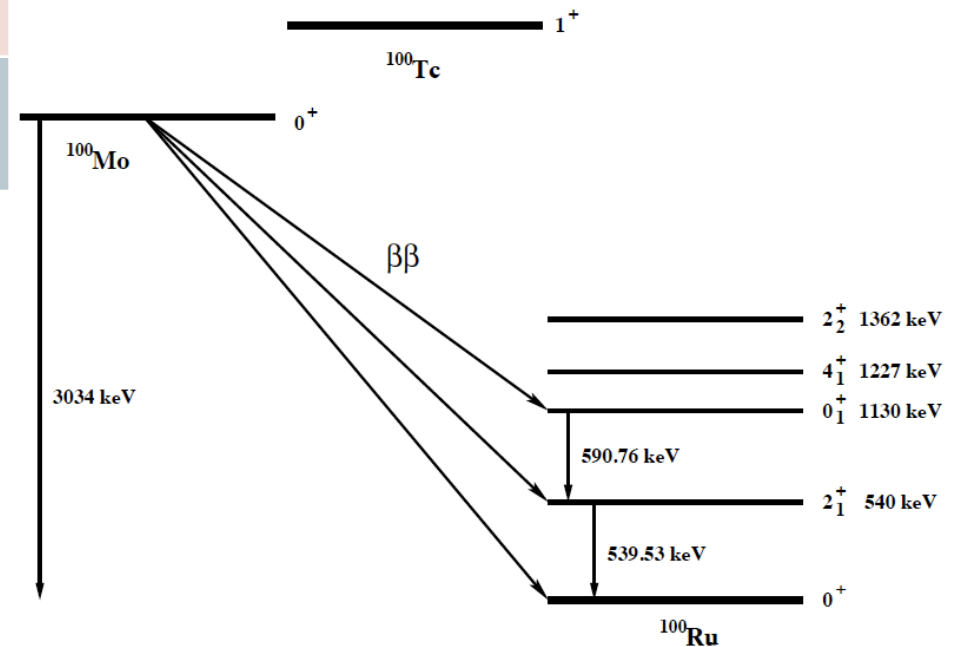
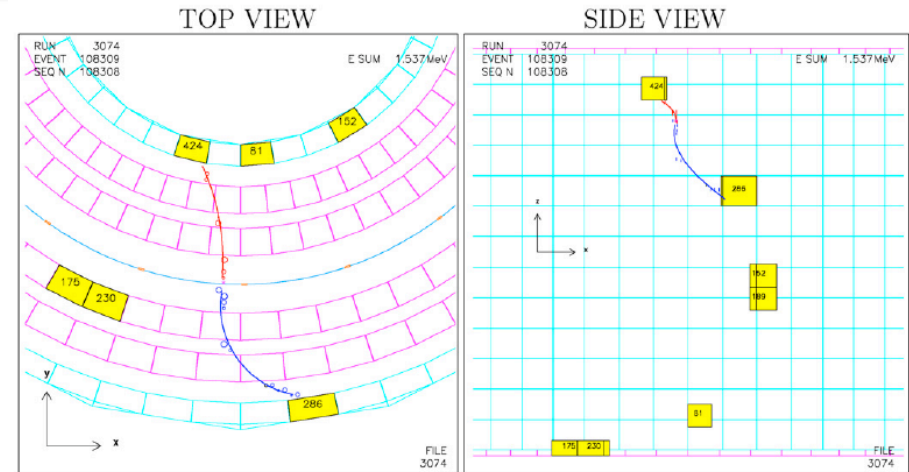
✓ $\langle m_\nu \rangle < 0.04 - 0.11$ eV

NEMO-3: $\beta\beta$ of ^{100}Mo to excited states



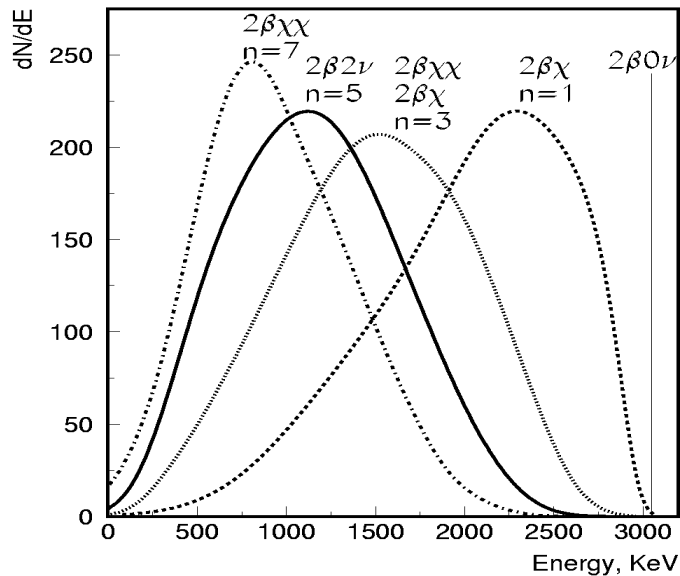
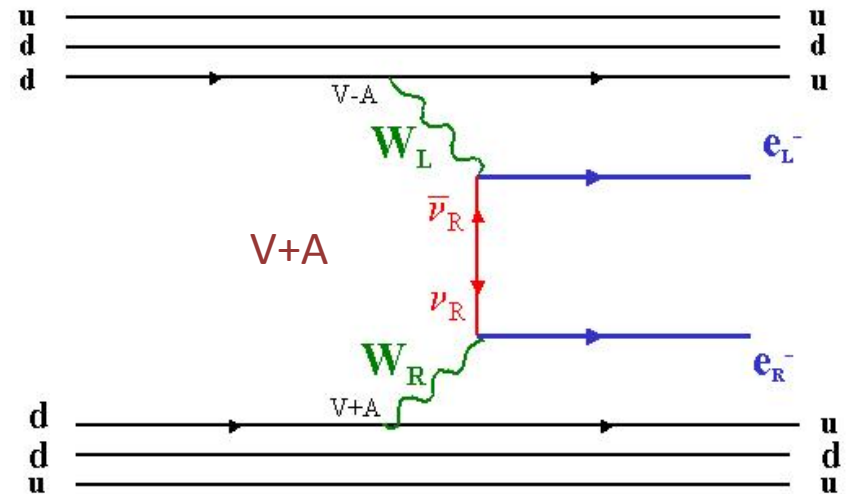
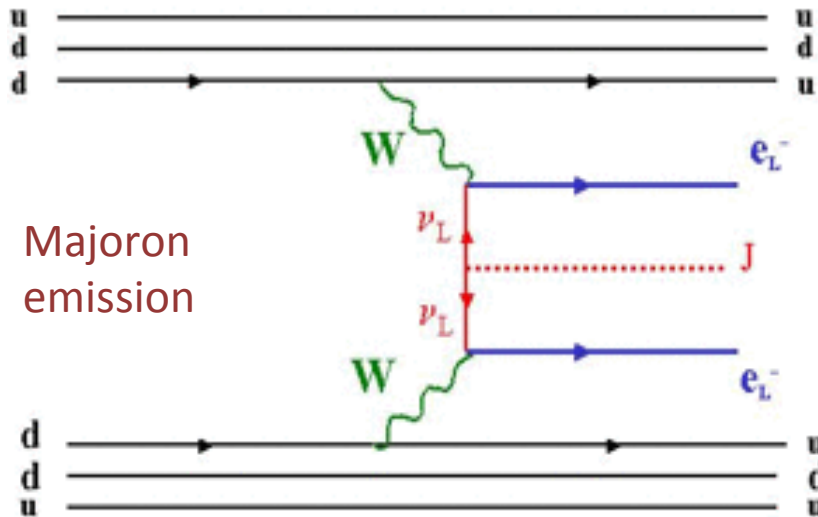
Transition	$T_{1/2}$ (y) (this work)	Theory
$0\nu\beta\beta$ $0^+ \rightarrow 2^+_1$	$> 1.6 * 10^{23}$	$6.8 * 10^{30} \langle m_\nu \rangle$ $2.1 * 10^{27} \langle \lambda \rangle$
$2\nu\beta\beta$ $0^+ \rightarrow 2^+_1$	$> 1.1 * 10^{21}$	$2.1 * 10^{21}$ $- 5.5 * 10^{25}$
$0\nu\beta\beta$ $0^+ \rightarrow 0^+_1$	$> 8.9 * 10^{22}$	$7.6 * 10^{24} \langle m_\nu \rangle$ $- 2.6 * 10^{26} \langle m_\nu \rangle$
$2\nu\beta\beta$ $0^+ \rightarrow 0^+_1$	$[5.7^{+1.3}_{-0.9}(\text{stat})$ $\pm 0.8 * 10^{20}$	$1.5 * 10^{20}$ $- 2.1 * 10^{21}$

↑
Best limits or uncertainties



NEMO Collaboration / Nuclear Physics A 781 (2007) 209–226

NEMO-3 - other physics

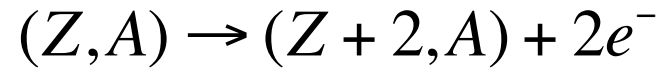


	V+A *	Majoron(s) emission (n=spectral index)**			
	$T_{1/2}(0\nu\beta\beta)$ [years]	n=1	n=2	n=3	n=7
^{100}Mo	$>5.7 \cdot 10^{23}$ $\lambda < 1.4 \cdot 10^{-6}$	$>2.7 \cdot 10^{22}$ $g_{ee} < (0.4-1.8) \cdot 10^{-4}$	$>1.7 \cdot 10^{22}$	$>1 \cdot 10^{22}$	$>7 \cdot 10^{19}$
^{82}Se	$>2.4 \cdot 10^{23}$ $\lambda < 2 \cdot 10^{-6}$	$>1.5 \cdot 10^{22}$ $g_{ee} < (0.7-1.9) \cdot 10^{-4}$	$>6 \cdot 10^{21}$	$>3.1 \cdot 10^{22}$	$>5 \cdot 10^{20}$

* Phase 1+Phase 2 data

** Phase 1 data, R. Arnold et al. Nucl. Phys. A765 (2006) 483

Neutrinoless double beta decay



Process:

- 1) Light neutrino exchange
- 2) (V+A) current
- 3) Majoron emission
- 4) SUSY

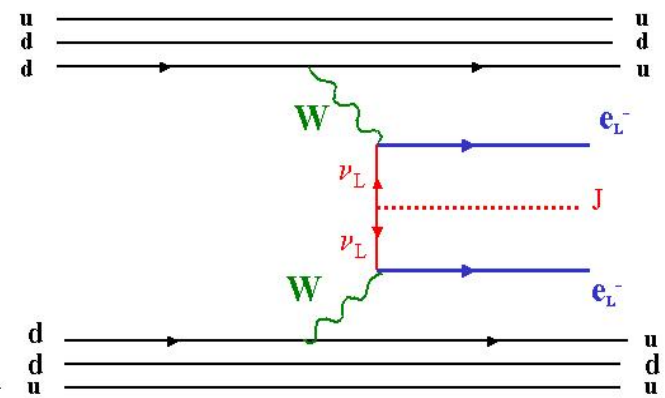
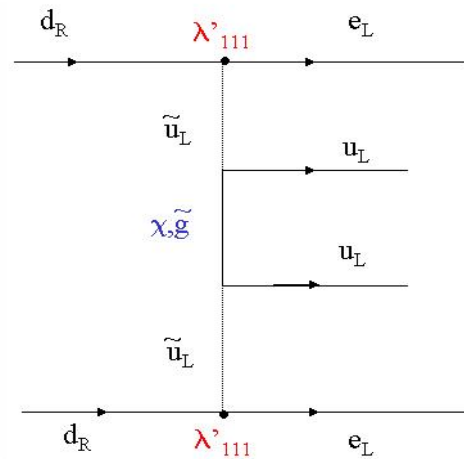
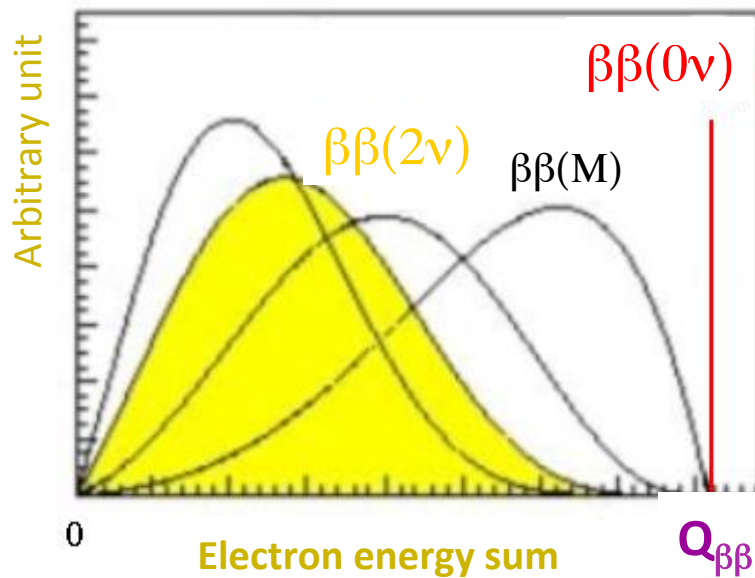
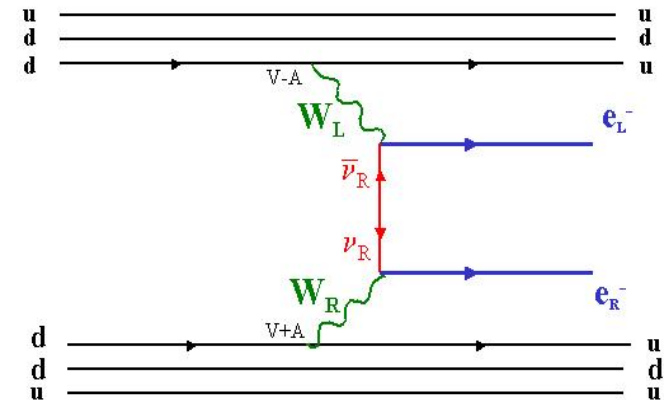
Parameters

$$\langle m_\nu \rangle$$

$$\langle m_\nu \rangle, \langle \lambda \rangle, \langle \eta \rangle$$

$$\langle g_M \rangle$$

$$\lambda'_{111}, \lambda'_{113}, \lambda'_{131}, \dots$$



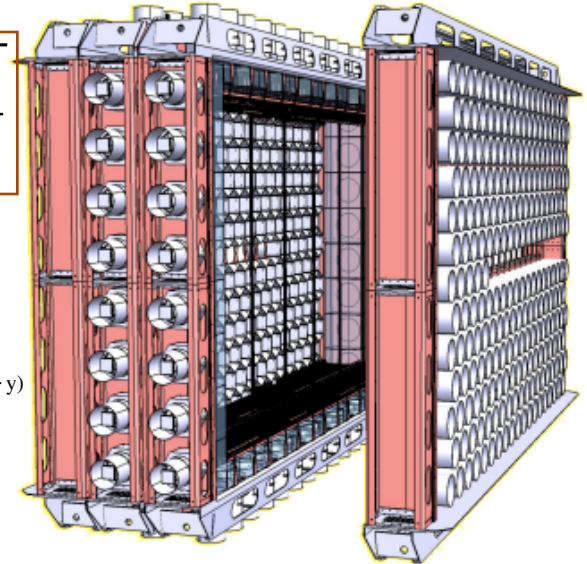
+ difference in angular distributions

NEMO-3 → SuperNEMO



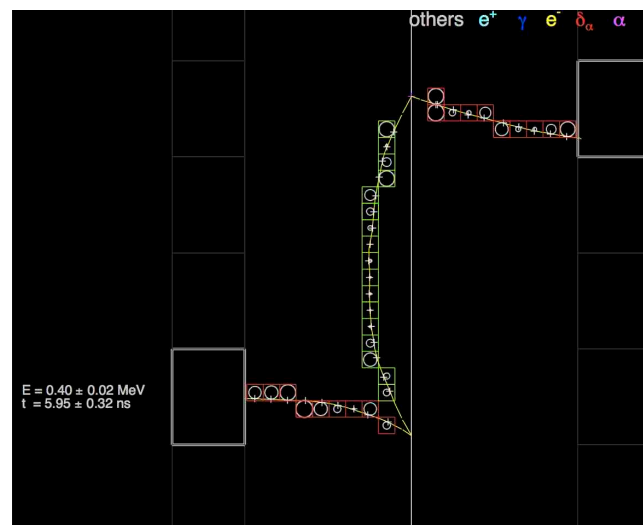
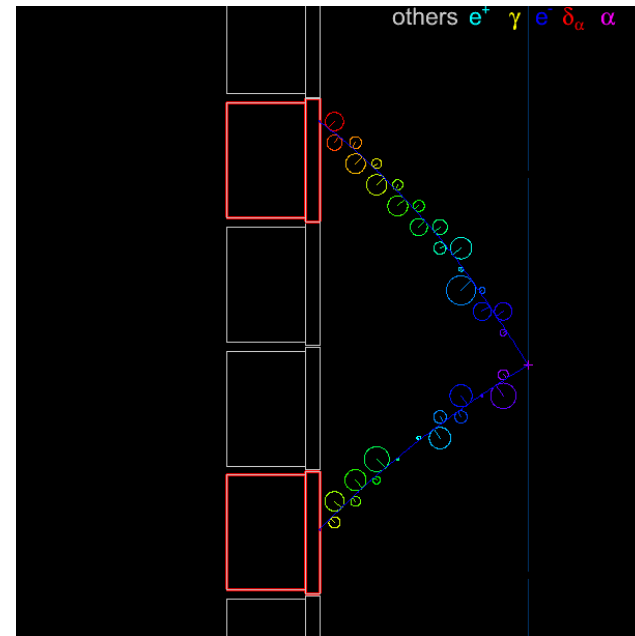
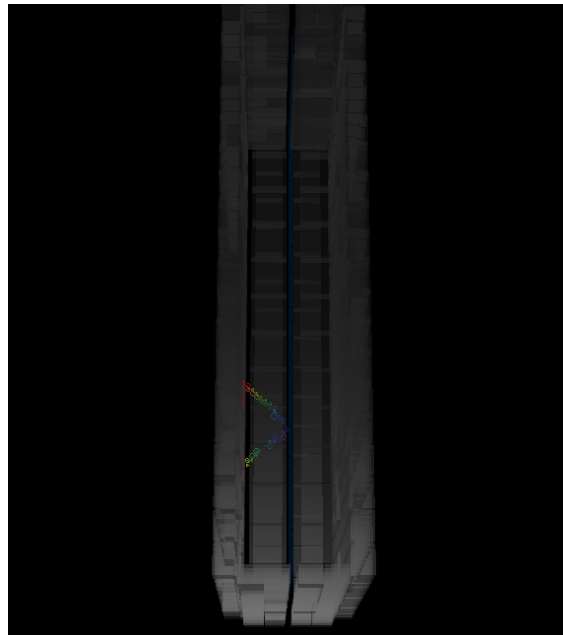
$$T_{1/2}^{0\nu}(n_\sigma) = \frac{4.16 \times 10^{26} y}{n_\sigma} \left(\frac{\epsilon a}{W} \right) \sqrt{\frac{Mt}{b\Delta E}}$$

n_σ – number of std. dev. for a given C.L. M – total mass of the source (kg)
 a – isotopic abundance t – time of data collection (y)
 ϵ – detection efficiency b – background rate in counts (keV · kg · y)
 W – molecular weight of the source ΔE – energy resolution (keV)



NEMO-3	R&D since 2005	SuperNEMO
^{100}Mo	isotope	^{82}Se (maybe also ^{150}Nd or ^{48}Ca)
7 kg	mass	100 kg
$A(^{208}\text{Tl}) < 20 \mu\text{Bq/kg}$ $A(^{214}\text{Bi}) < 300 \mu\text{Bq/kg}$ $\text{Rn} \sim 5\text{-}6 \text{ mBq/m}^3$	Radio-purity of the foil Radon in the tracker	$A(^{208}\text{Tl}) < 2 \mu\text{Bq/kg}$ $A(^{214}\text{Bi}) < 10 \mu\text{Bq/kg}$ $\text{Rn} < 0.1 \text{ mBq/m}^3$
8%	efficiency	30%
8% FWHM @ 3 MeV	Energy resolution	4% FWHM @ 3 MeV
$T_{1/2}(0\nu\beta\beta) > 2 \times 10^{24} \text{ y}$ $\langle m_n \rangle < 0.3 - 0.8 \text{ eV}$	sensitivity	$T_{1/2}(0\nu\beta\beta) > 1 \times 10^{26} \text{ y}$ $\langle m_n \rangle < 40 - 100 \text{ meV}$
1 module	modularity	>20 modules (new lab)

Improved simulations and tracking



Phenomenology of $0\nu\beta\beta$ and $2\nu\beta\beta$ (1)

$$\frac{1}{T_{1/2}^{2\nu}} = G_{2\nu}(Q_{\beta\beta}^{11}, Z) \cdot |M_{2\nu}|^2$$

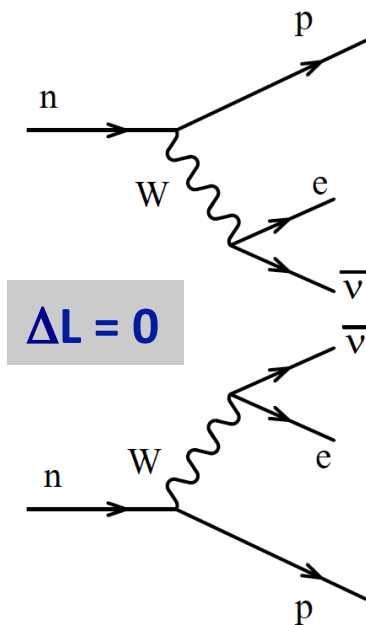
G = phase space (well known)

M = nuclear matrix element (challenging)

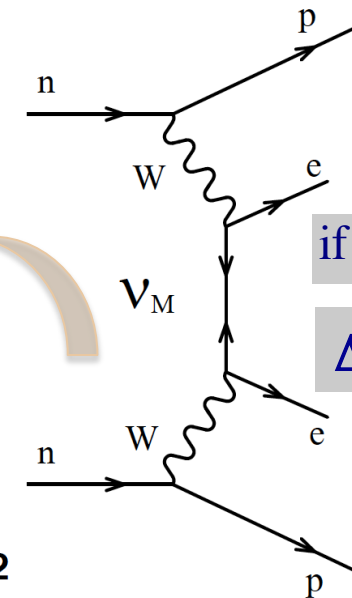
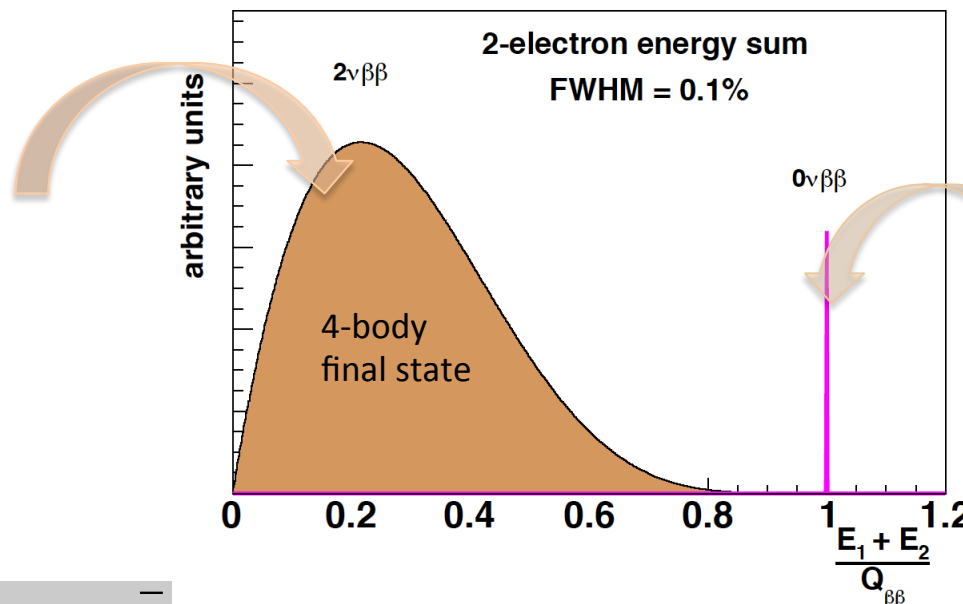
$$\frac{1}{T_{1/2}^{0\nu}} = G_{0\nu}(Q_{\beta\beta}^5, Z) \cdot |M_{0\nu}|^2 \cdot \langle m_{\beta\beta} \rangle^2$$

$$\langle m_{\beta\beta} \rangle \equiv \left| m_1 |U_{e1}|^2 + m_2 |U_{e2}|^2 e^{i\alpha^*} + m_3 |U_{e3}|^2 e^{i\beta^* - 2i\delta} \right|$$

α^*, β^* = linear combinations of α and β

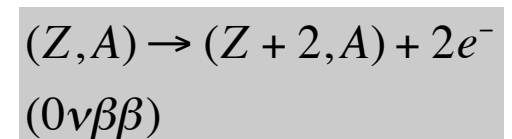
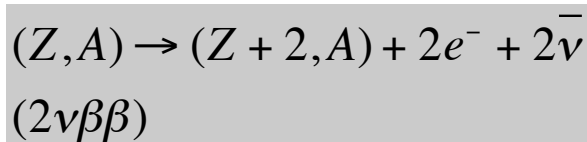


$\Delta L = 0$



if $m_\nu \neq 0$

$\Delta L = 2$



Phenomenology of $0\nu\beta\beta$ and $2\nu\beta\beta$ (1)

$$\frac{1}{T_{1/2}^{2\nu}} = G_{2\nu}(Q_{\beta\beta}^{11}, Z) \cdot |M_{2\nu}|^2$$

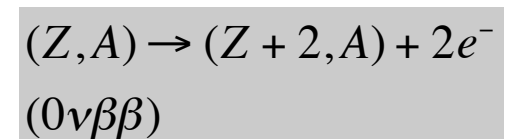
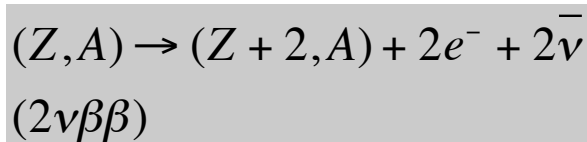
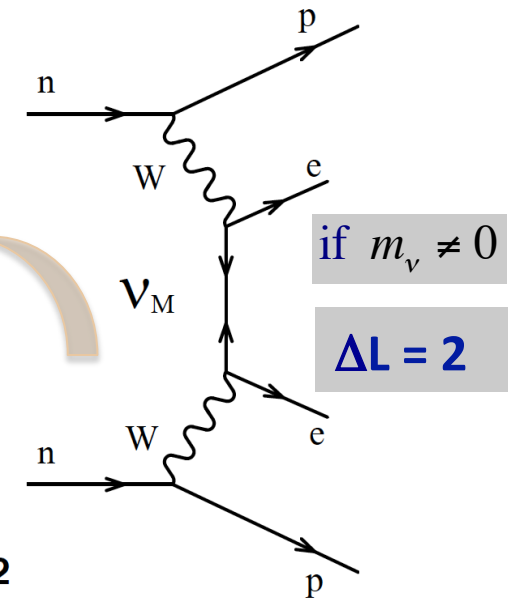
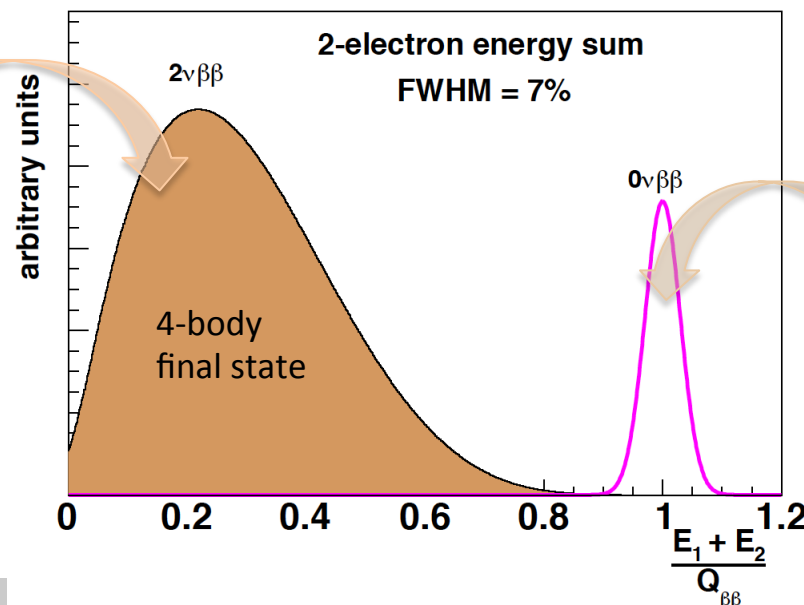
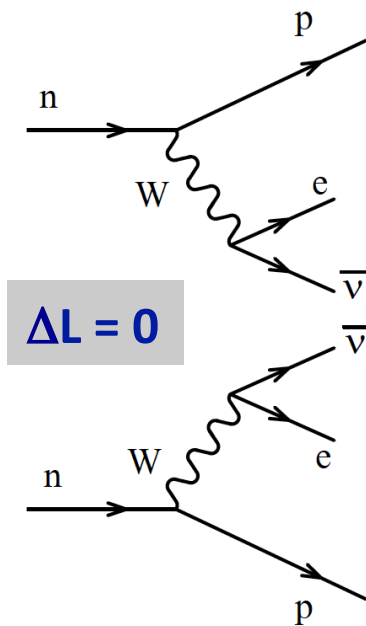
G = phase space (well known)

M = nuclear matrix element (challenging)

$$\frac{1}{T_{1/2}^{0\nu}} = G_{0\nu}(Q_{\beta\beta}^5, Z) \cdot |M_{0\nu}|^2 \cdot \langle m_{\beta\beta} \rangle^2$$

$$\langle m_{\beta\beta} \rangle \equiv \left| m_1 |U_{e1}|^2 + m_2 |U_{e2}|^2 e^{i\alpha^*} + m_3 |U_{e3}|^2 e^{i\beta^* - 2i\delta} \right|$$

α^*, β^* = linear combinations of α and β





Phenomenology of $0\nu\beta\beta$ and $2\nu\beta\beta$ (2)

$$\frac{1}{T_{1/2}^{2\nu}} = G_{2\nu}(Q_{\beta\beta}^{11}, Z) \cdot |M_{2\nu}|^2$$

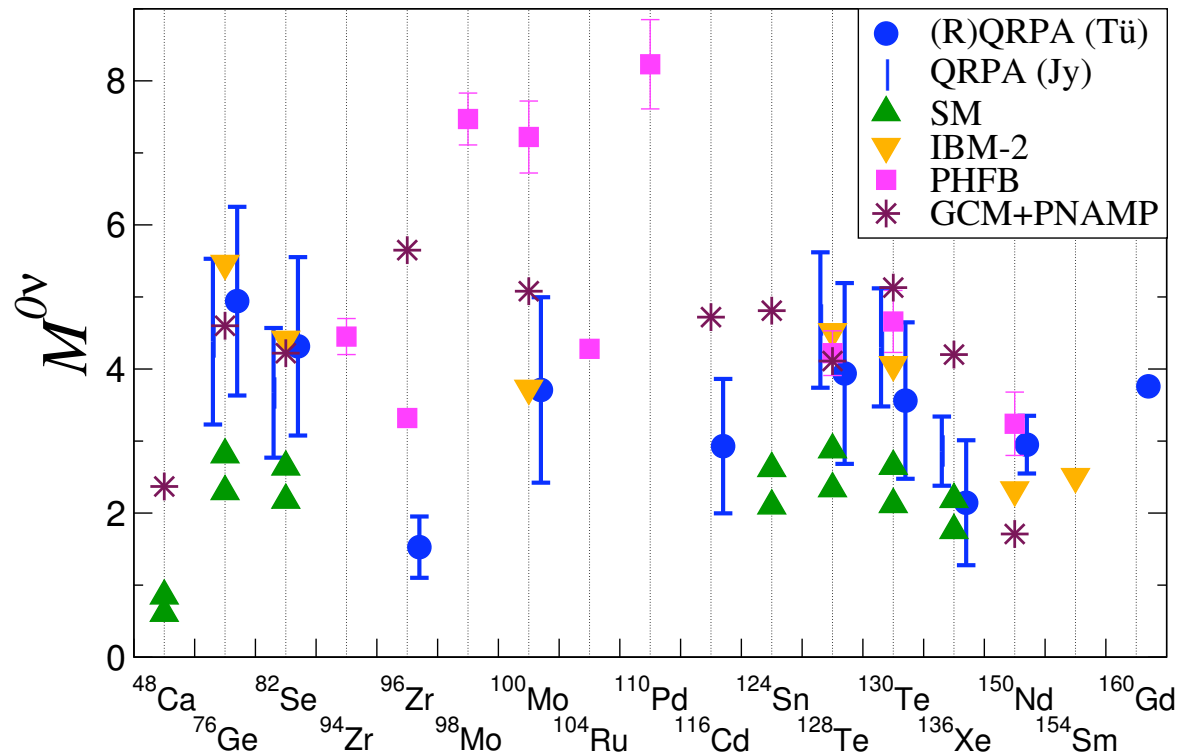
$$\frac{1}{T_{1/2}^{0\nu}} = G_{0\nu}(Q_{\beta\beta}^5, Z) \cdot |M_{0\nu}|^2 \cdot \langle m_{\beta\beta} \rangle^2$$

G = phase space (well known)

M = nuclear matrix element (challenging)

$$\langle m_{\beta\beta} \rangle \equiv \left| m_1 |U_{e1}|^2 + m_2 |U_{e2}|^2 e^{i\alpha^*} + m_3 |U_{e3}|^2 e^{i\beta^* - 2i\delta} \right|$$

α^*, β^* = linear combinations of α and β

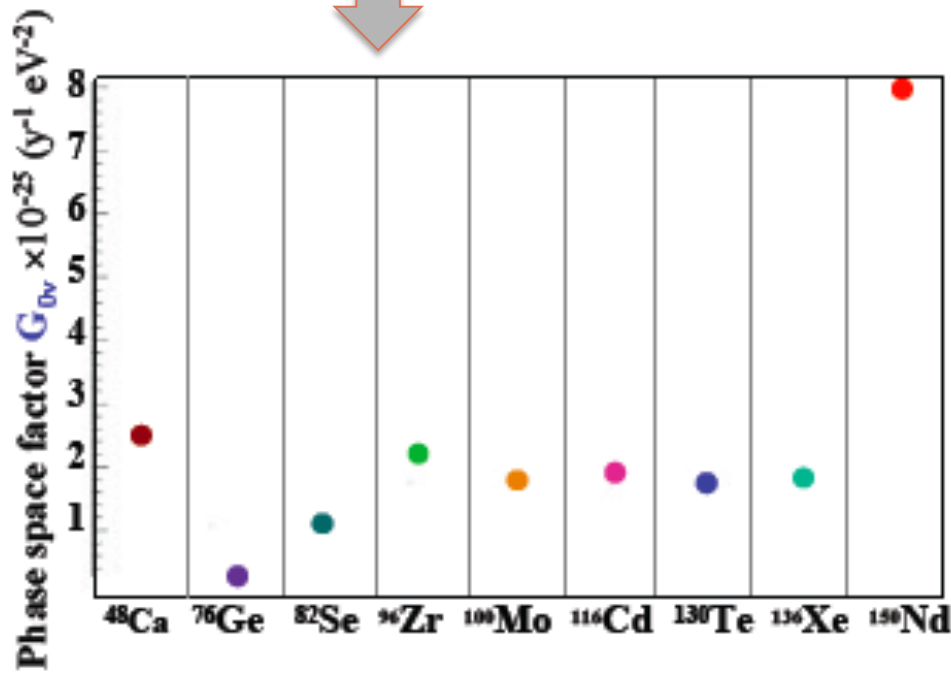
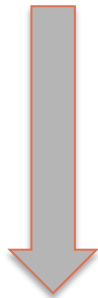


From V. Rodin,
TAUP 2011 in Munich

Practical matters



$$\frac{1}{T_{1/2}^{0\nu}} = G_{0\nu} \cdot |M_{0\nu}^{GT}|^2 \cdot \langle m_{\beta\beta} \rangle^2$$



NEMO-3	$Q_{\beta\beta}$ (MeV)	Natural abundance (%)
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	4.272	0.187
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	3.367	5.6
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	3.350	2.8
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	3.035	9.6
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	2.995	9.2
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	2.805	7.5
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	2.529	34.5
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	2.458	8.9
$^{124}\text{Sn} \rightarrow ^{124}\text{Te}$	2.228	5.64
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	2.039	7.8
$^{110}\text{Pd} \rightarrow ^{110}\text{Cd}$	2.013	11.8

Top 11 $\beta\beta$ emitters with $Q_{\beta\beta} > 2$ MeV

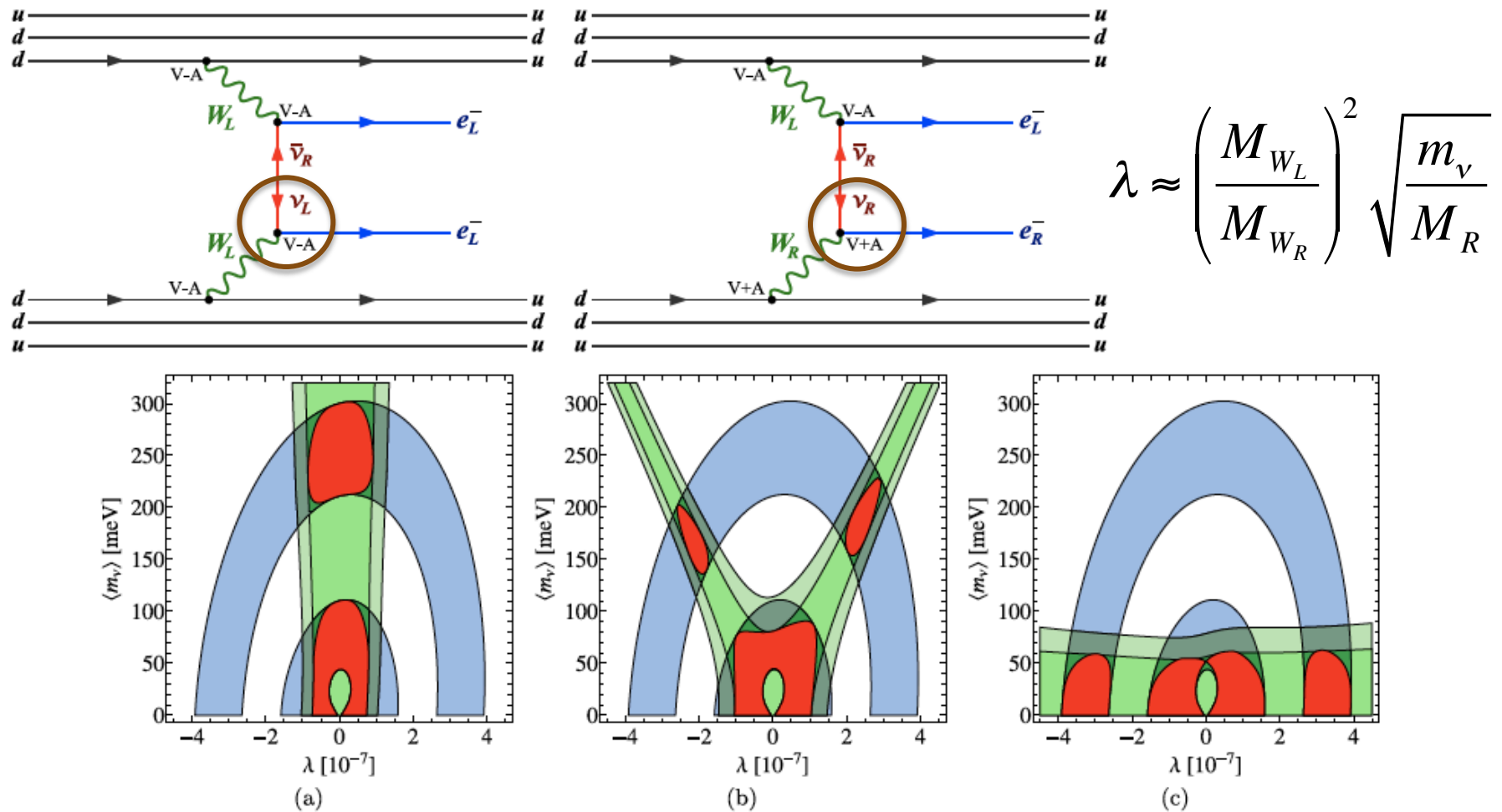


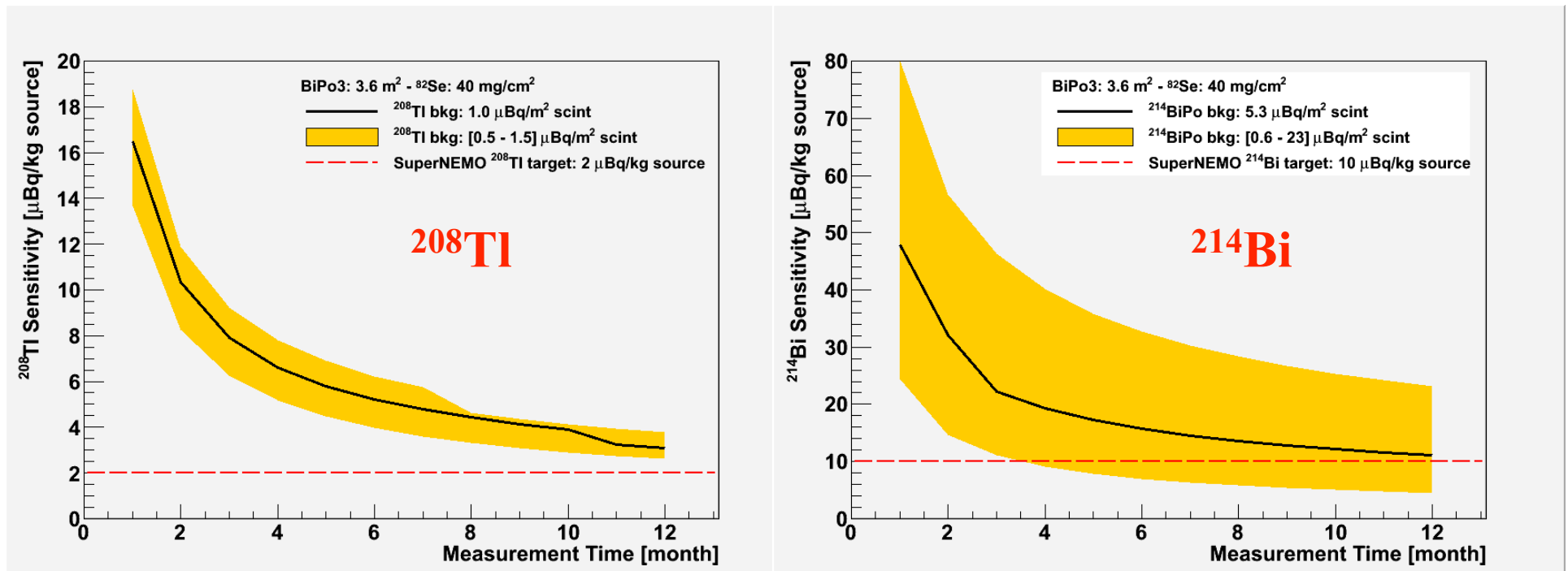
Fig. 11 (Color online) Constraints at one standard deviation on the model parameters m_ν and λ for ^{82}Se from: (1) an observation of $0\nu\beta\beta$ decay half-life at $T_{1/2} = 10^{25}$ y (outer blue elliptical contour) and 10^{26} y (inner blue elliptical contour); (2) reconstruction of the angular (outer, lighter green) and energy difference (inner, darker green) distribution shape; (3) combined analysis of (1) and (2) using decay rate and

energy distribution shape reconstruction (red contours). The admixture of the MM and RHC_λ contributions is assumed to be: a pure MM contribution; b 30% RHC_λ admixture; and c pure RHC_λ contribution. NME uncertainties are assumed to be 30% and experimental statistical uncertainties are determined from the simulation



BiPo-3 sensitivity

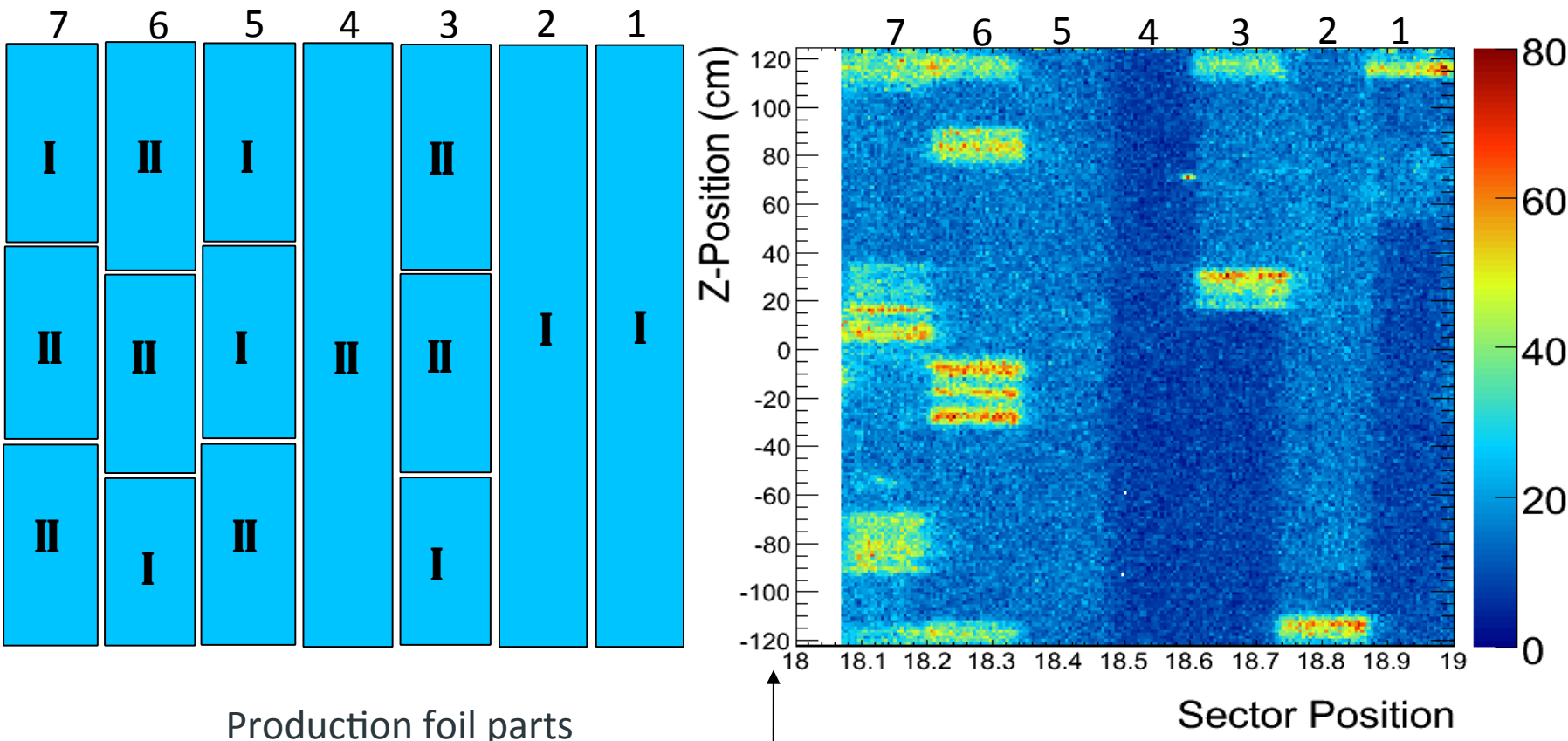
Assuming: ^{82}Se foil 40 mg/cm^2
Total surface BiPo-3 = 3.6 m^2
Energy threshold = 100 keV for prompt and delay signals ($\epsilon \sim 5\%$)



SuperNEMO demonstrator:

- Foils $270 \times 13.55 \text{ cm}^2$, 8 foils installed in BiPo-3 \Rightarrow Total surface = 2.93 m^2
efficiency reduced by $2.9/3.6 = 20\%$
- If thickness = 53 mg/cm^2
efficiency reduced by $40/53 = 25\%$

Cadmium Foil Activity and Hot Spots



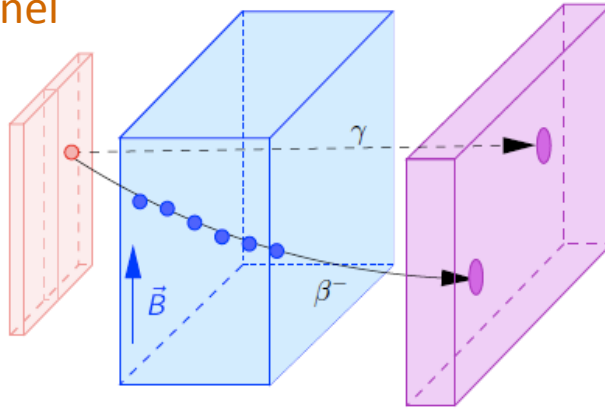
Vertex at the foil for 1 electron data

calibration tube

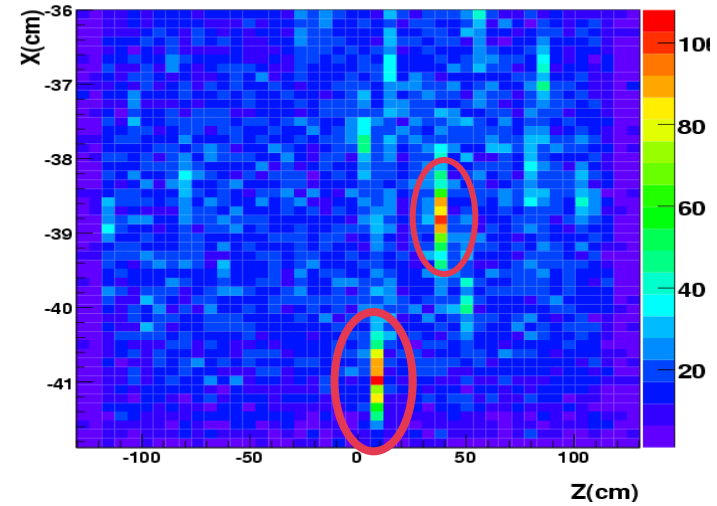
Background: control channels



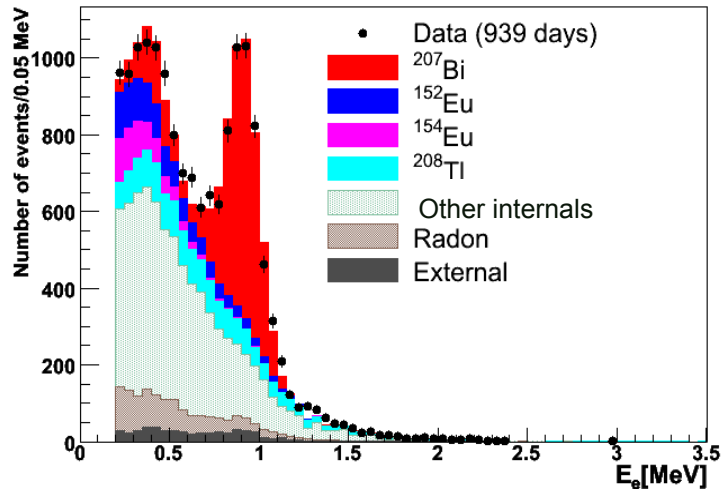
Example:
e γ control channel



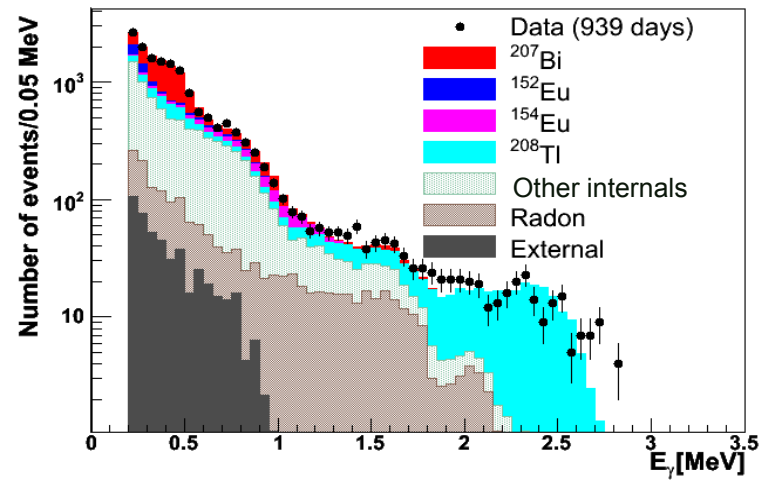
150-Nd foil



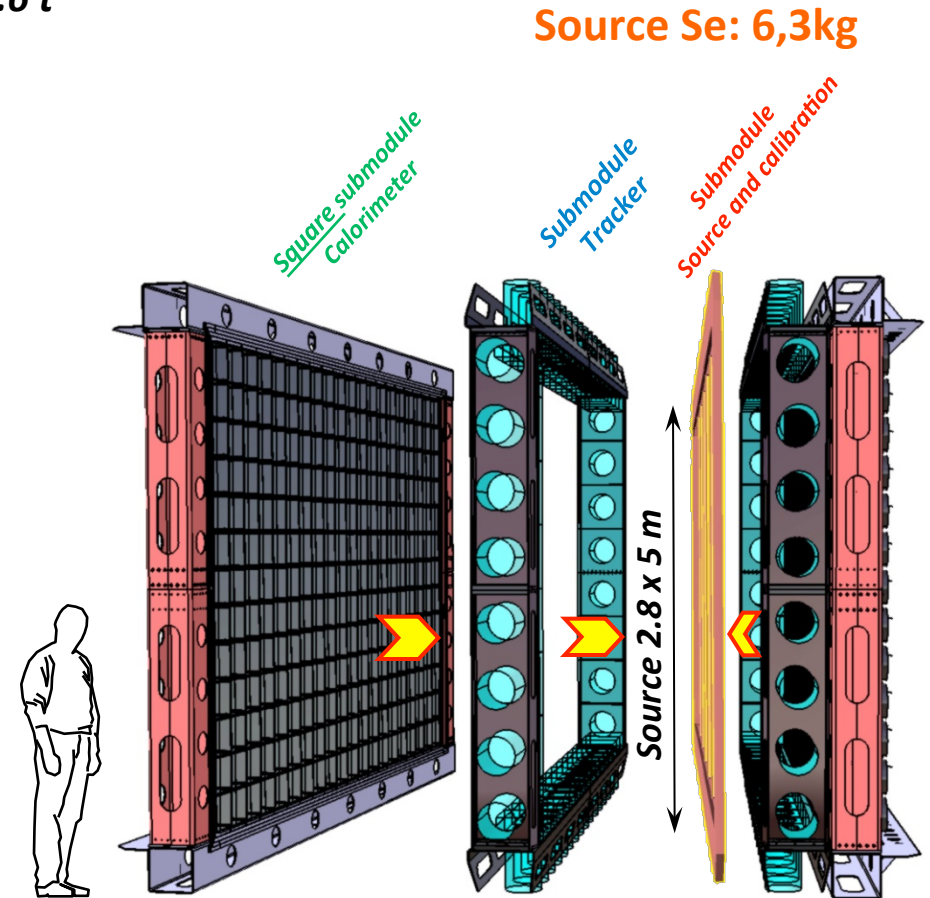
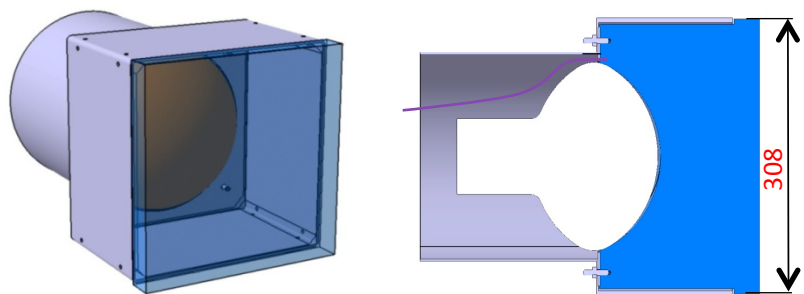
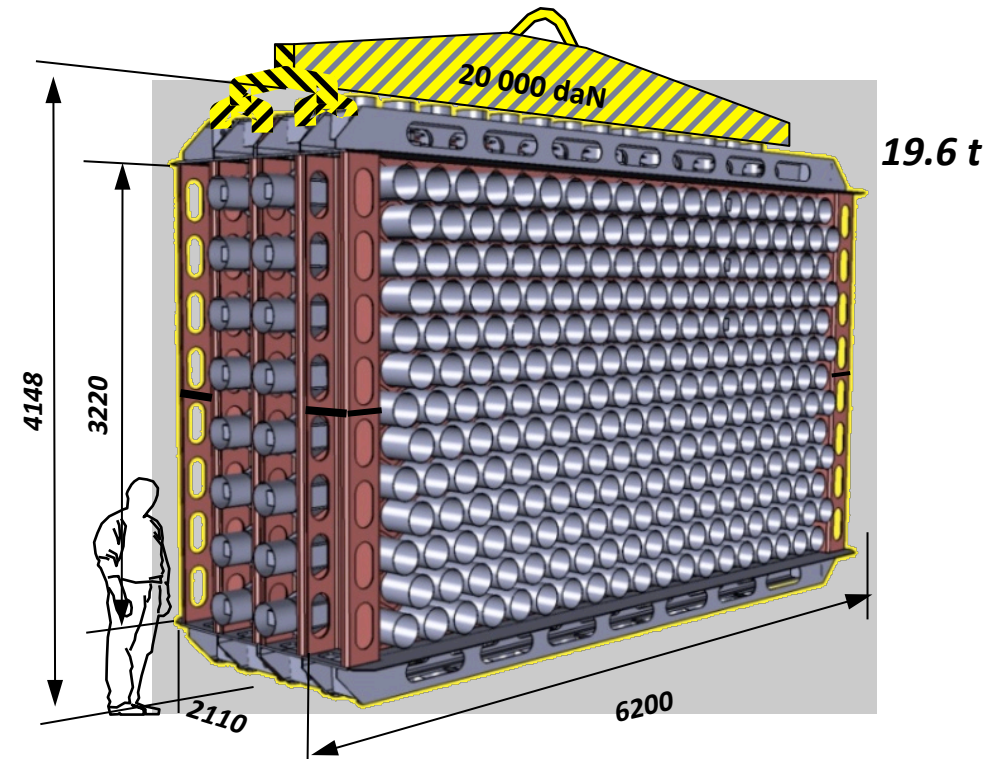
Energy of the electron



Energy of the photon



SuperNEMO Demonstrator module



« Calorimeter square design »

NEMO-3 detector



Fréjus Underground Laboratory : 4800 m.w.e.

Source: 10 kg of $\beta\beta$ isotopic foils
area = 20 m², thickness ~ 60 mg/cm²

Tracking detector:

drift wire chamber operating (9 layers)
in Geiger mode (6180 cells)

Gas: He + 4% ethyl alcohol + 1% Ar + 0.1% H₂O

Calorimeter:

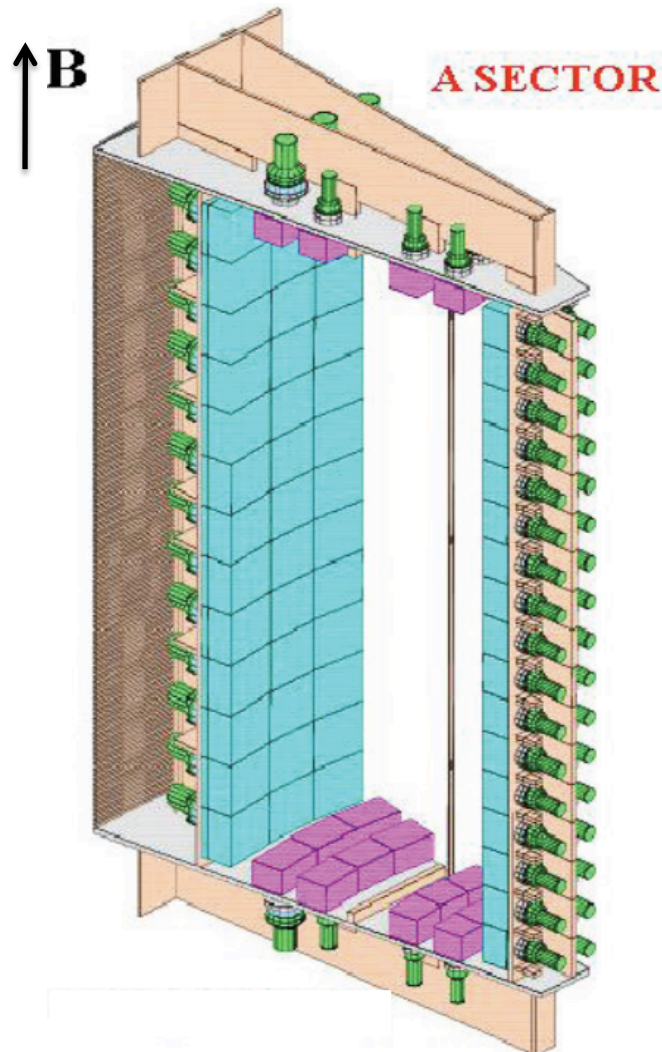
1940 plastic scintillators
coupled to low radioactivity PMTs

Magnetic field: 25 Gauss

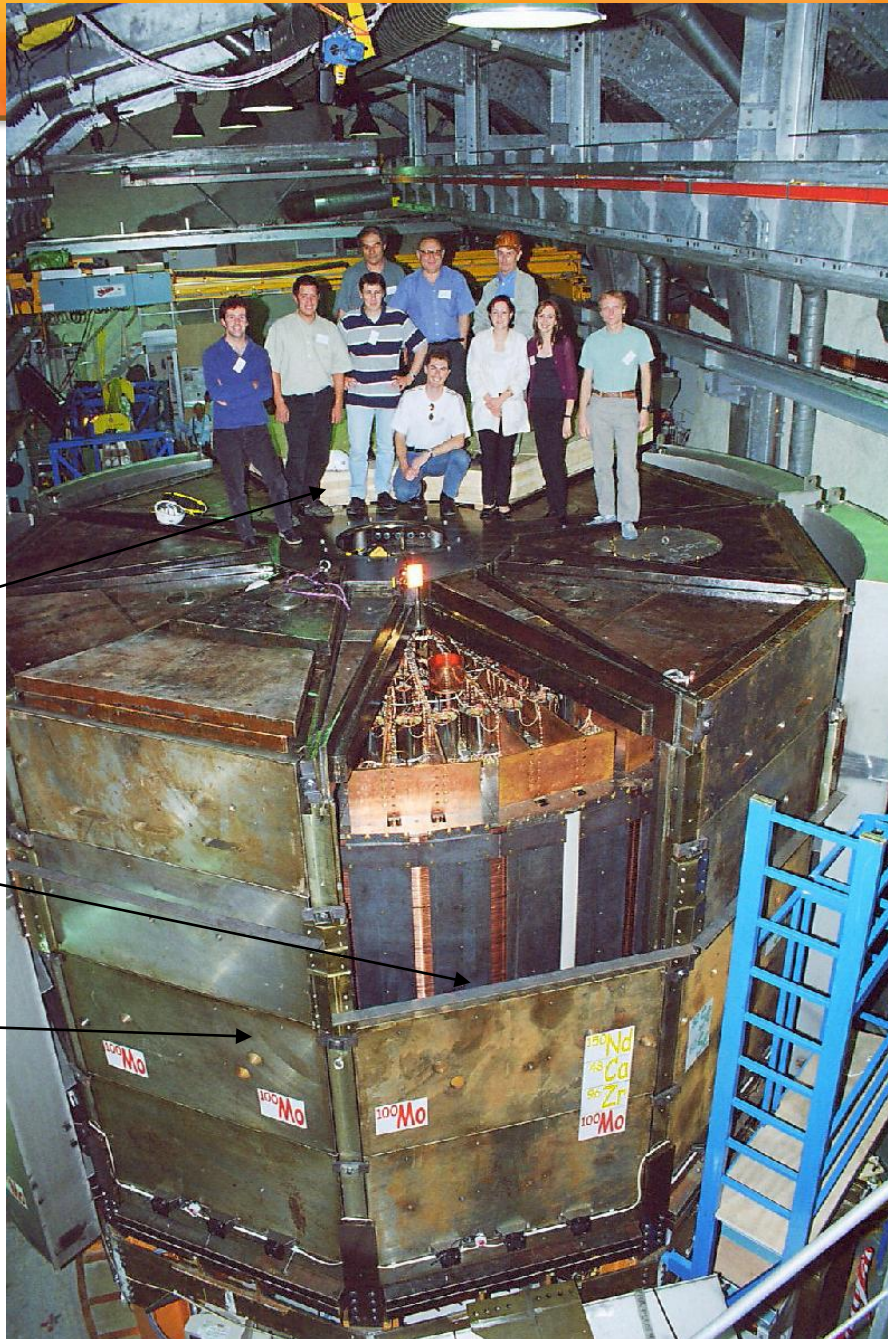
Gamma shield: pure iron (d = 18cm)

Neutron shield: 30 cm water (ext. wall)

40 cm **WOOD** (top and bottom)
(since March 2004:
water + boron)



Particle ID: e^- , e^+ , γ and α



Water tank

wood

coil

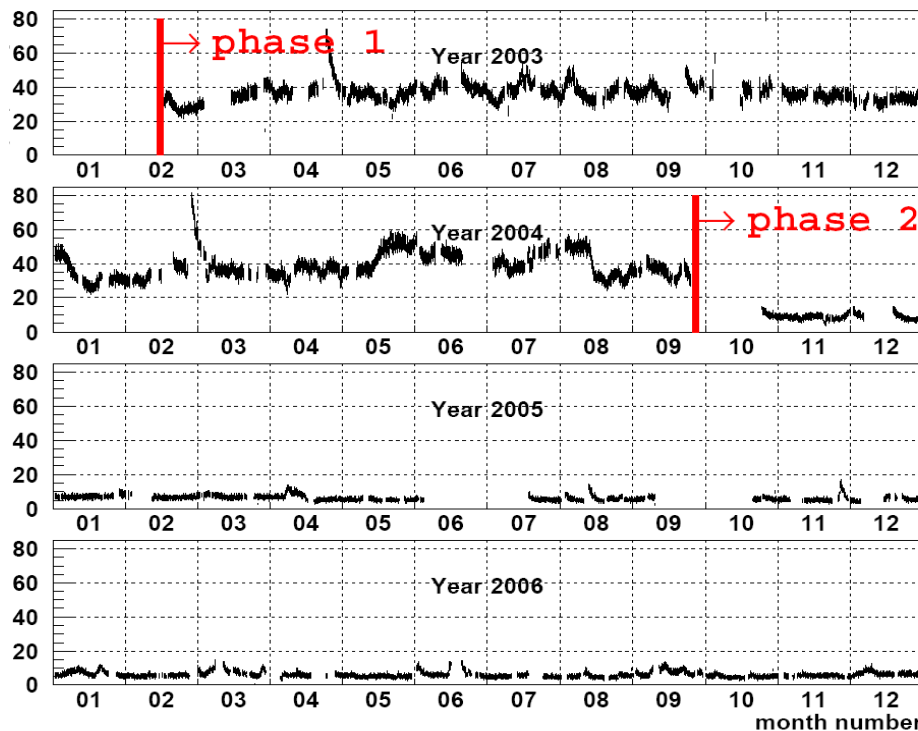
Iron shield

Radon Trapping Facility



- Radon trapping facility installed in September 2004.
- The trapping time in activated charcoal longer than ^{222}Rn half-life of 3.8 days.
- Radon level reduced by almost factor of 10 in the detector by installing radon trapping facility

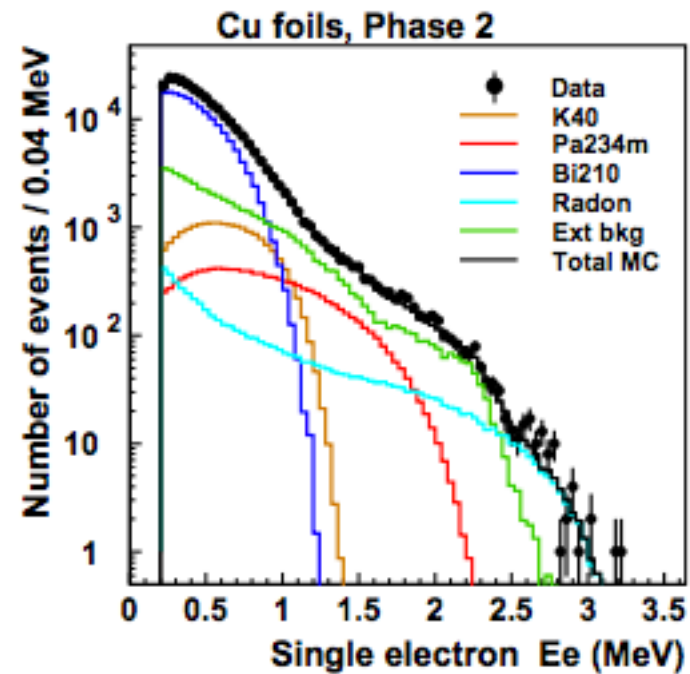
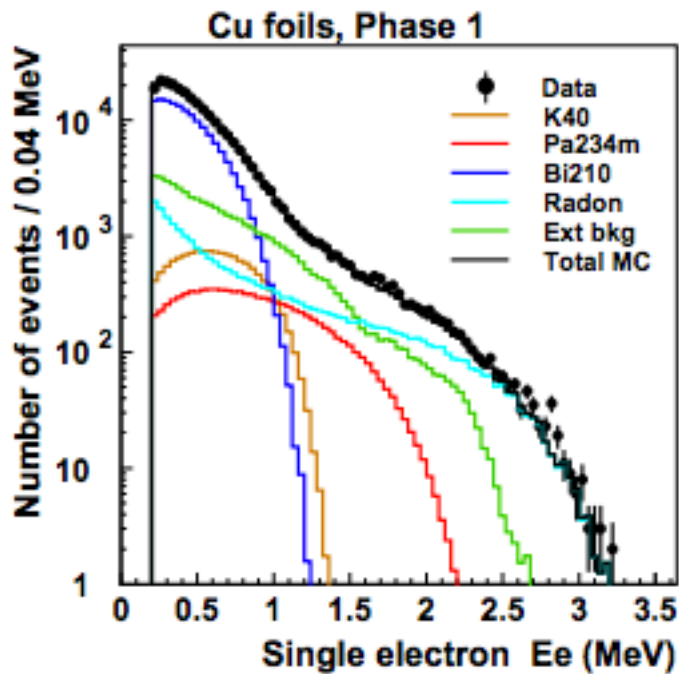
Adsorption unit @-50°C



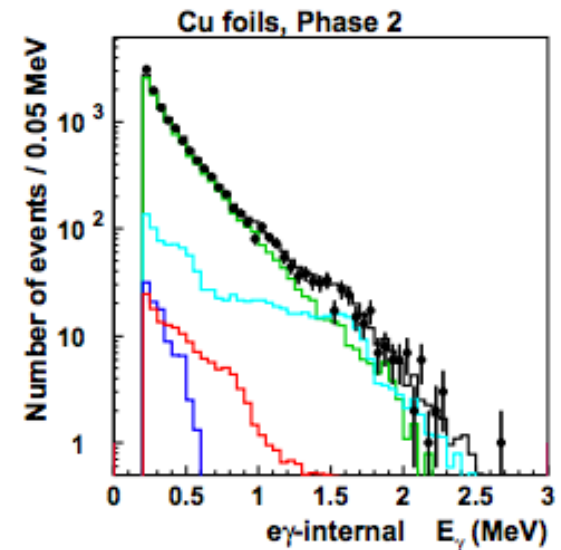
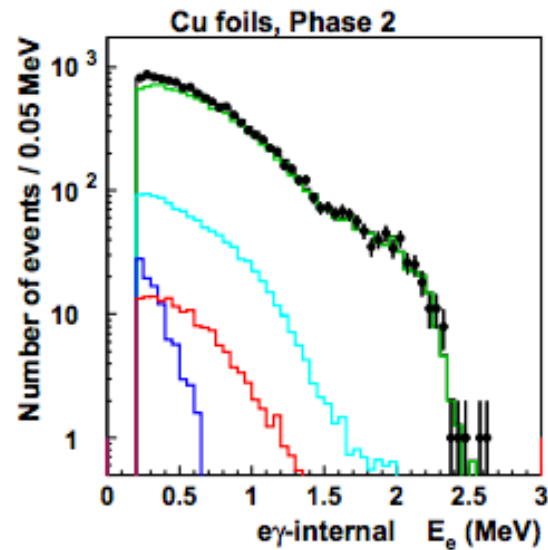
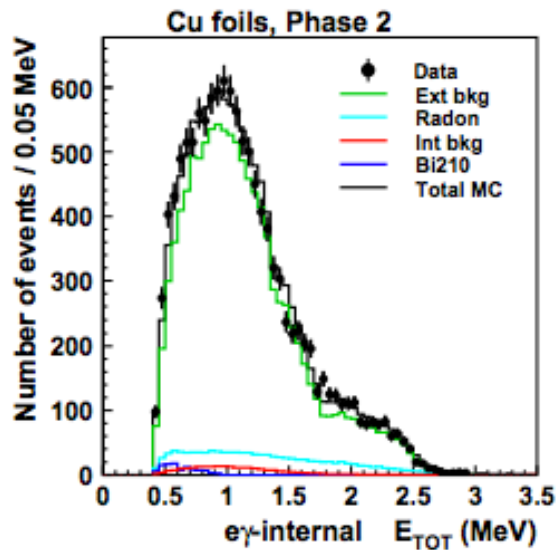
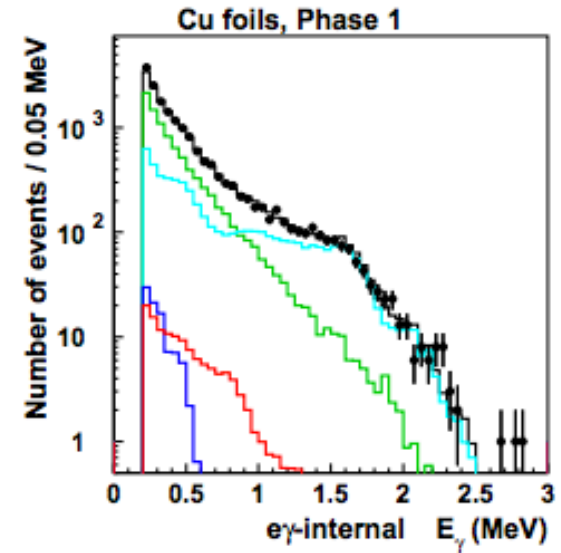
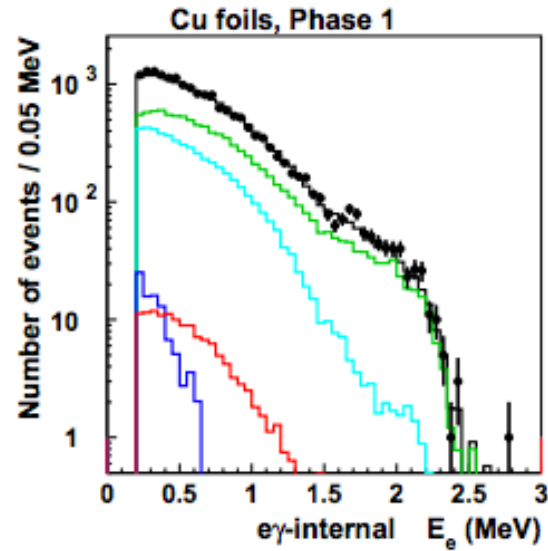
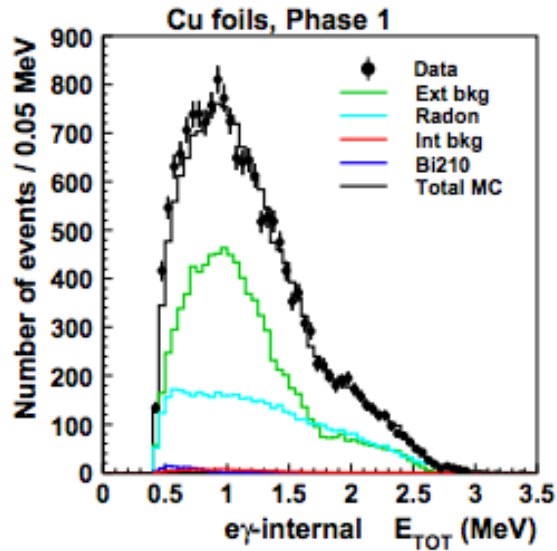
Input: $A(^{222}\text{Rn})$ 15 Bq/m^3

Output: $A(^{222}\text{Rn}) < 15 \text{ mBq/m}^3$!!
reduction factor of 1000

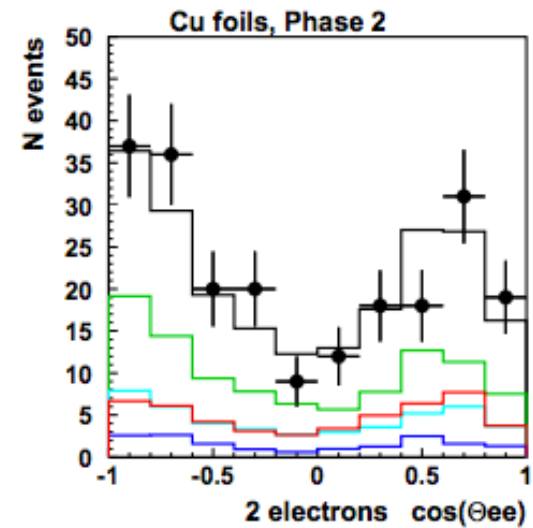
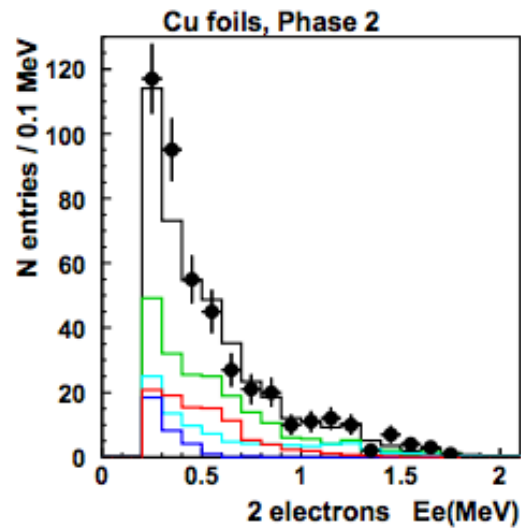
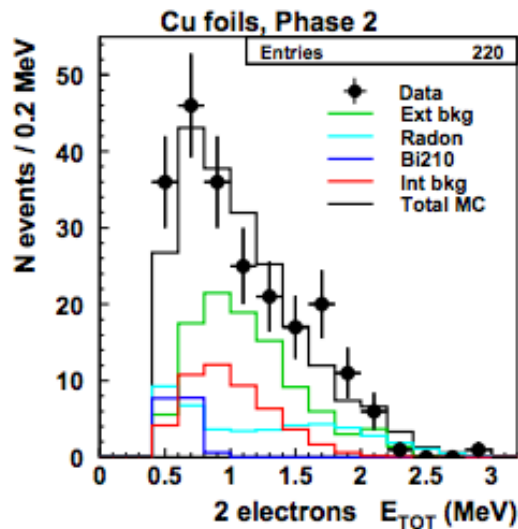
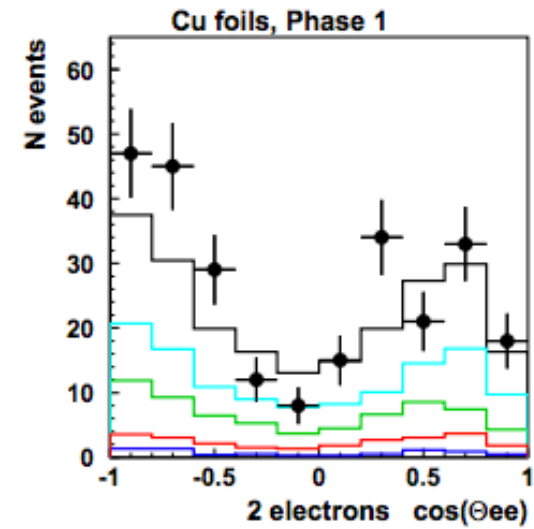
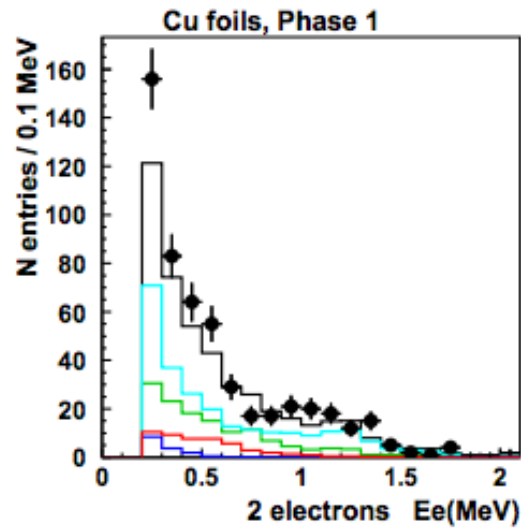
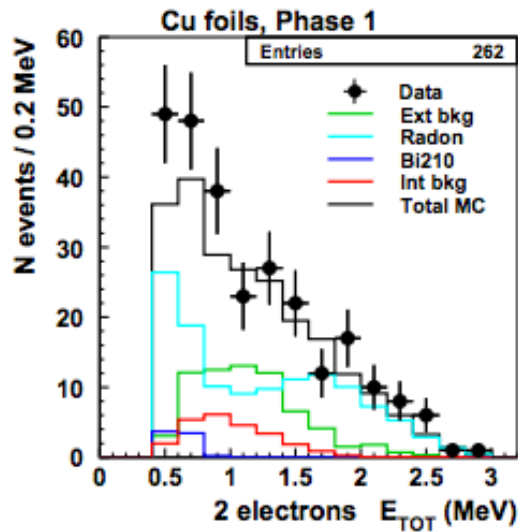
Phase 1 versus Phase 2 backgrounds



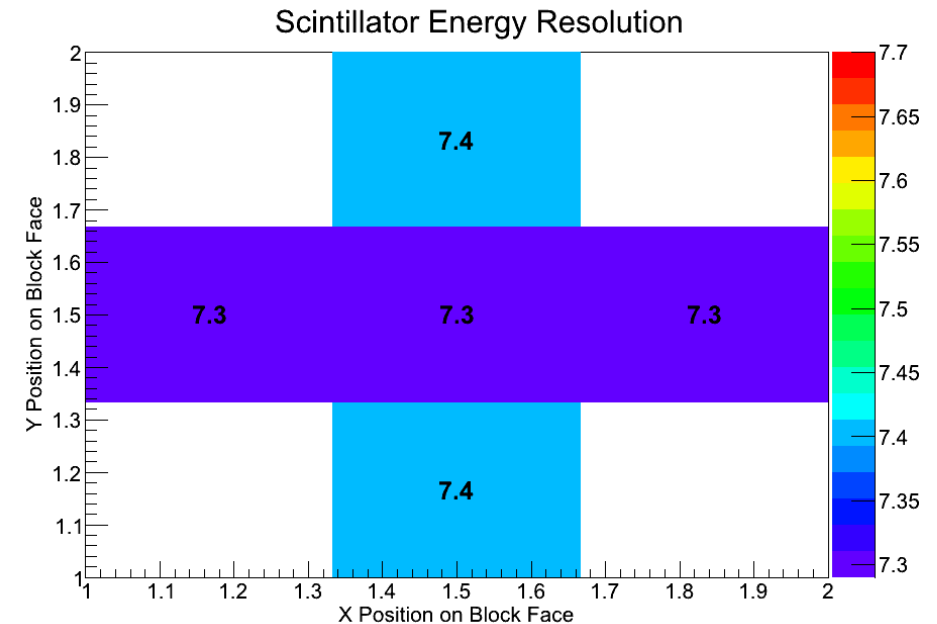
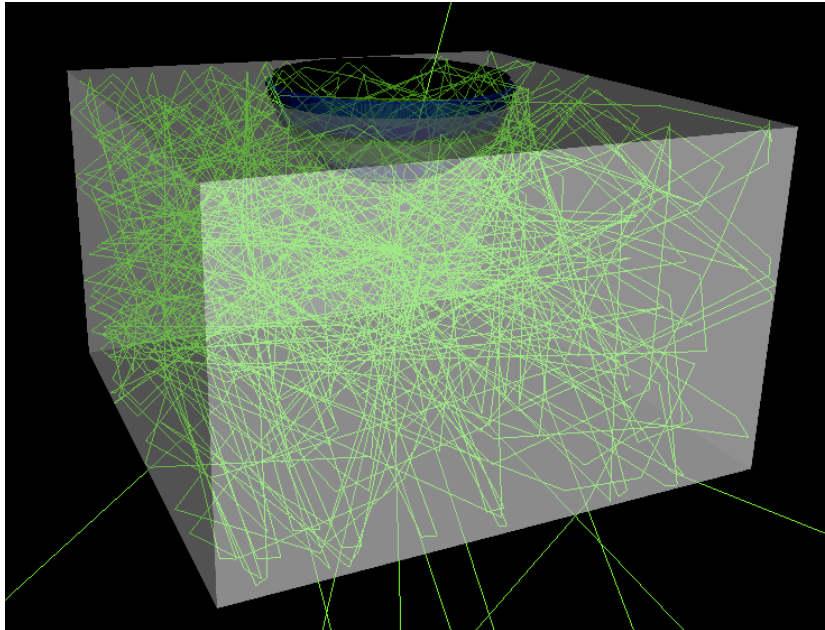
Phase 1 backgrounds



Phase 2 backgrounds



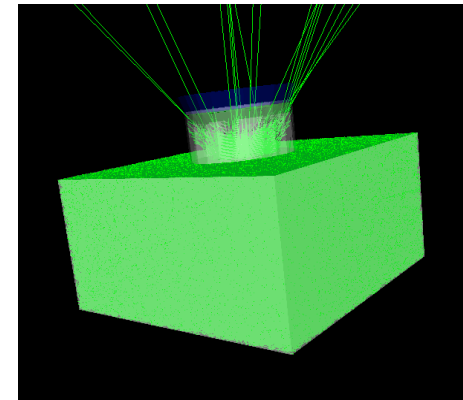
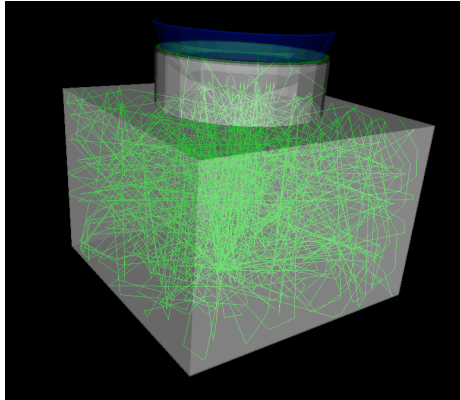
Main wall calorimeter simulations



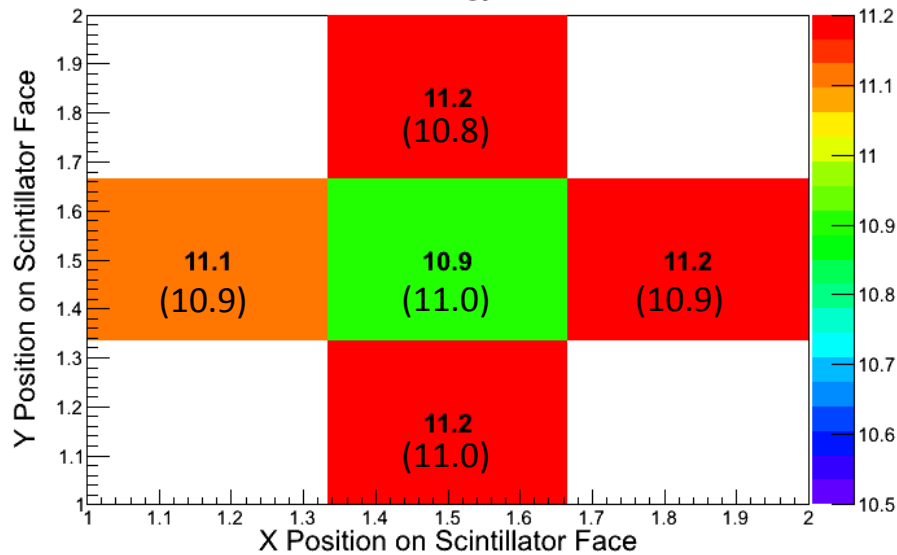
- ◆ Visualization of a GEANT-4 simulation of of a main wall calorimeter block (left), and diagram showing simulated resolution by electron entrance position on the block surface (right).
- ◆ All simulated values are in agreement with measurements.

X-Wall

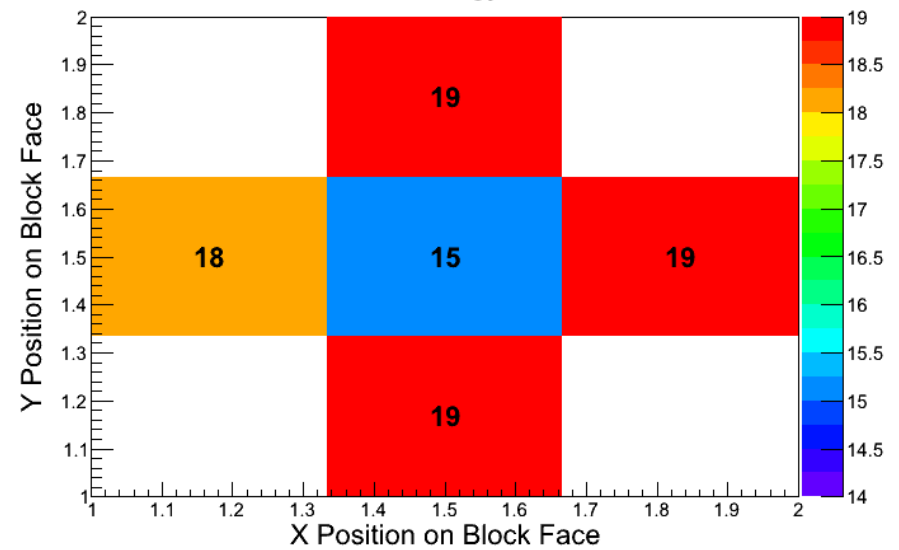
V-Wall



Scintillator Energy Resolution

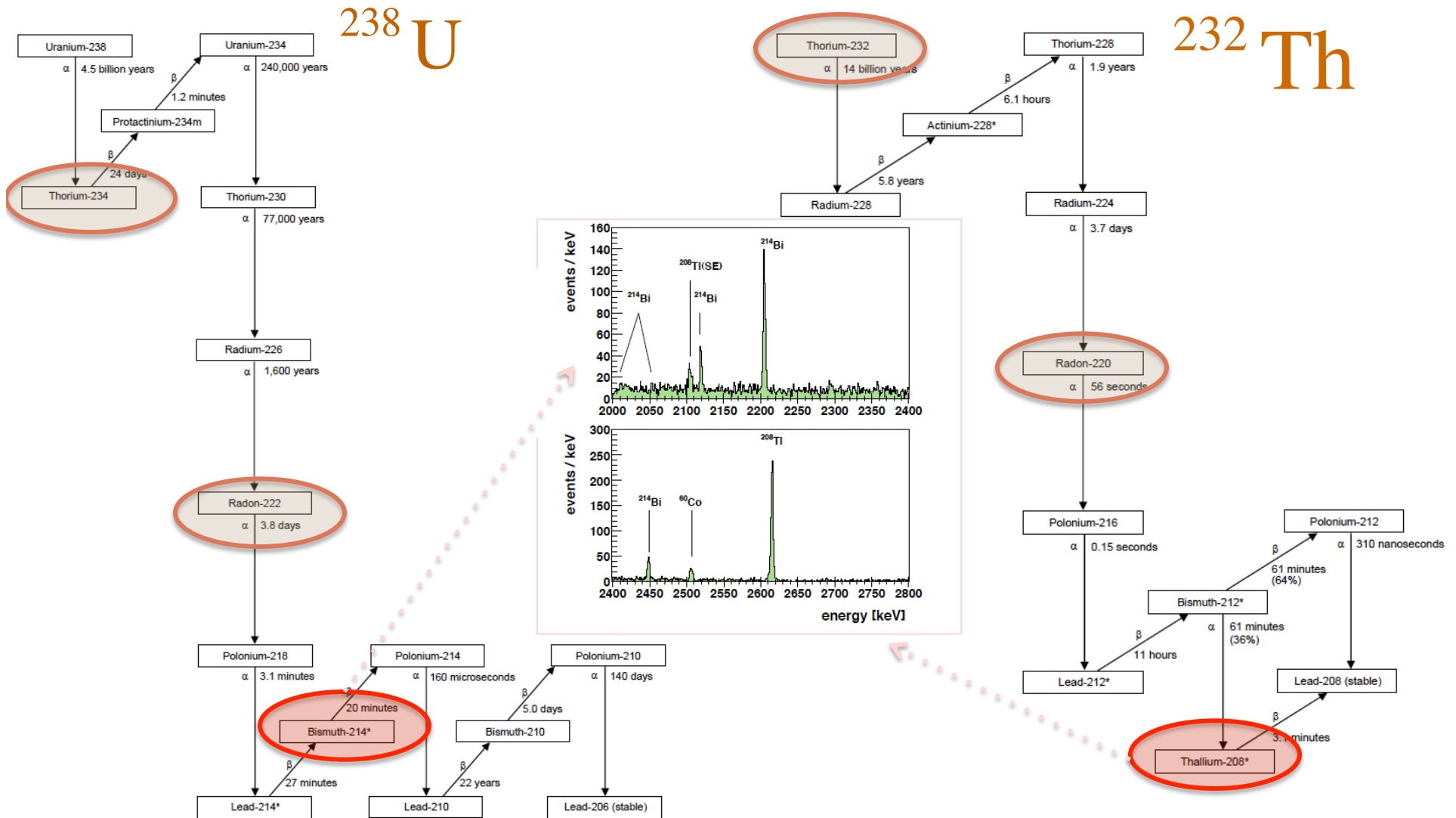


Scintillator Energy Resolution



- ◆ Visualization and resolutions by electron entry position for X-wall (left) and V-wall (right) calorimeter blocks.
- ◆ Measured positional values are in parenthesis; no such values are available for the V-wall blocks.

Perennial problem – natural radioactivity



Thorium and radon are diffusive radioactive isotopes out-gassed into the air from the rock.