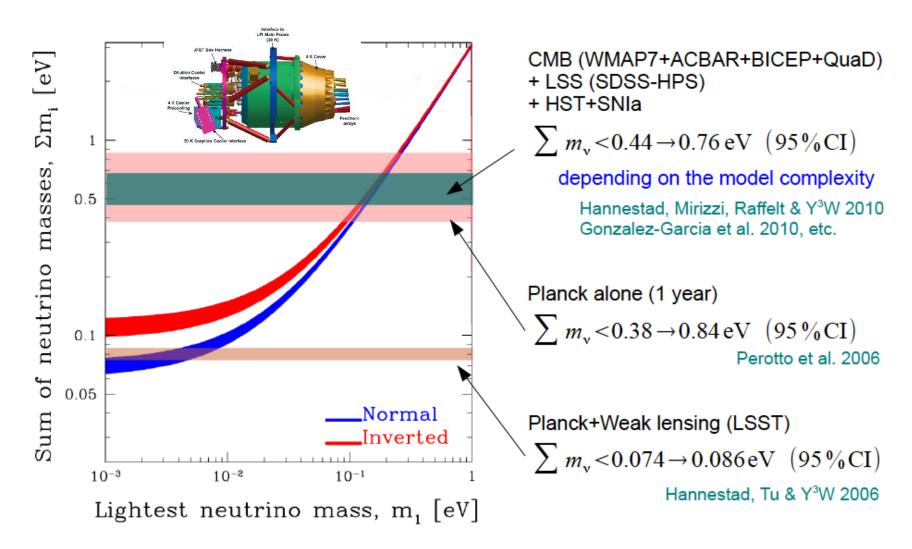


Particle Physics Cosmology What is the neutrino mass scale?

Hamish Robertson, ESPP12, Kracow, Sept. 11, 2012

Present constraints and future sensitivities...



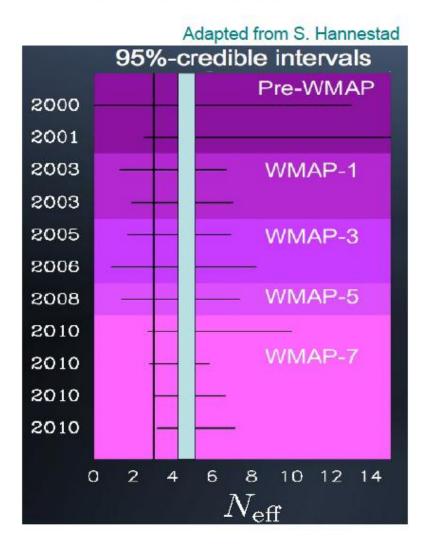
Y.Y.Y. Wong

Recent hint for N_{eff} > 3 from precision cosmology...

 Parameterise excess relativistic energy density in terms of extra species of massless neutrinos.

$$\rho_{\nu} + \rho_{X} = N_{\text{eff}} \left(\frac{7}{8} \frac{\pi^{2}}{15} T_{\nu}^{4} \right)$$
$$= (3.046 + \Delta N_{\text{eff}}) \left(\frac{7}{8} \frac{\pi^{2}}{15} T_{\nu}^{4} \right)$$

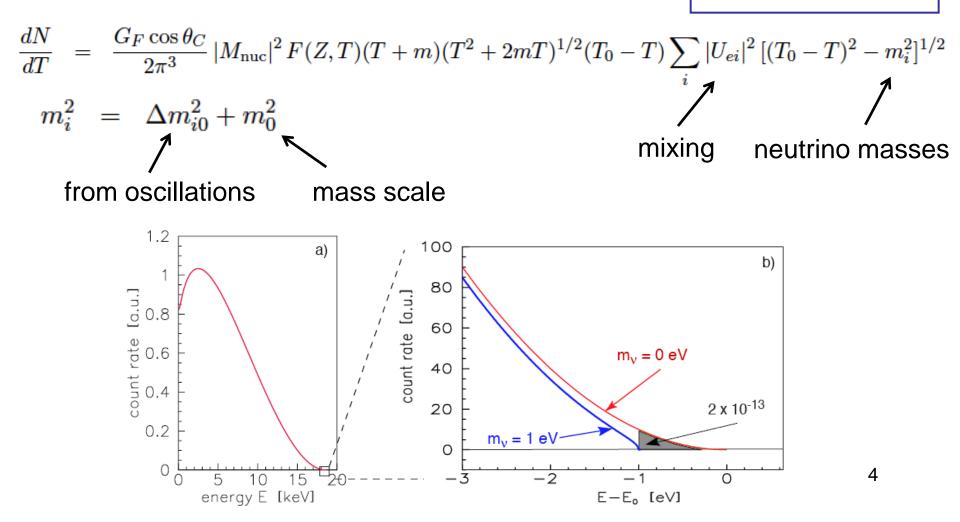
- Evidence for N_{eff} > 3:
 - @ 98.4% (CMB+LSS) Hou et al. 2011
 - @ 99.5% (CMB+LSS+BBN) Hamann et al. 2011



Y.Y.Y. Wong

Neutrino mass from Beta Spectra

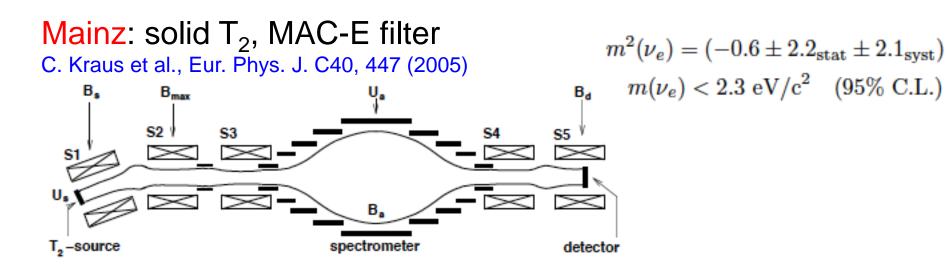
With flavor mixing:



ЗH

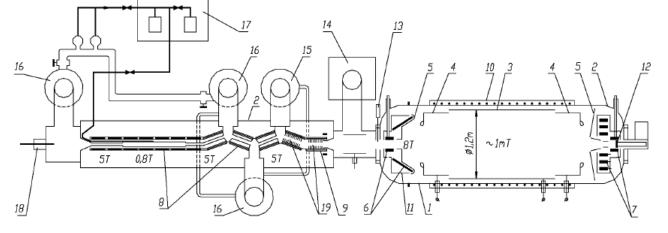
³He

Current status of direct mass measurement



Troitsk: gaseousT₂, MAC-E filter V. Aseev et al., PRD in press (2011)

 $m_{\nu}^2 = -0.67 \pm 1.89_{stat} \pm 1.68_{syst}$ $m_{\nu} < 2.05 \ eV, \ 95\% \ C. L.$



Together:... m_v < 1.8 eV (95% CL)

KATRIN

At Karlsruhe Institute of Technology unique facility for closed T₂ cycle: Tritium Laboratory Karlsruhe



<u>Ieľ</u>



Size of experiment now: Diameter 10 m.

$$\tau(m_{\nu}^2) = k \frac{b^{1/6}}{r^{2/3} t^{1/2}},$$

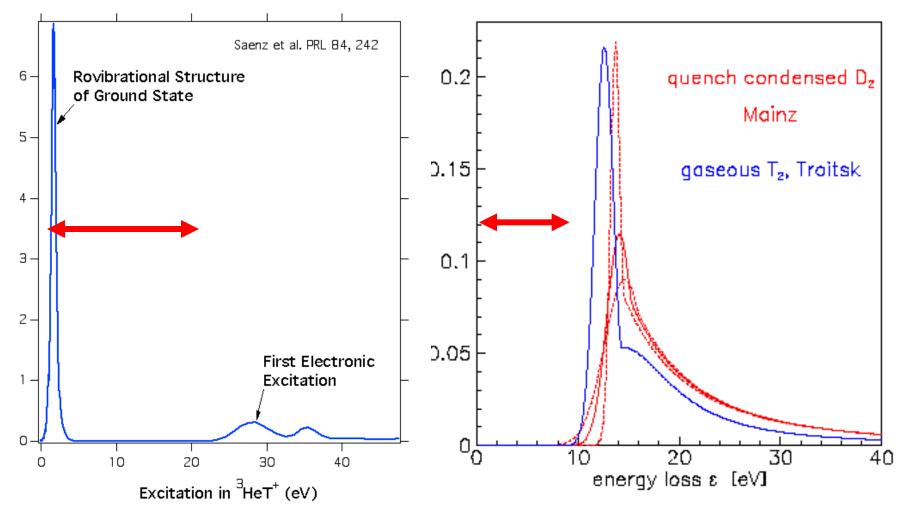
Next diameter: 300 m!

~ 75 m long with 40 s.c. solenoid

A window to work in

Molecular Excitations

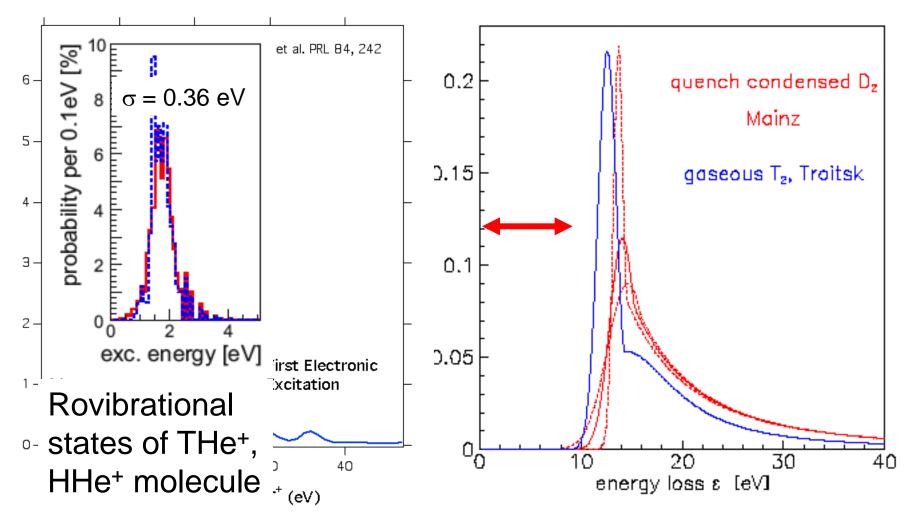
Energy loss function



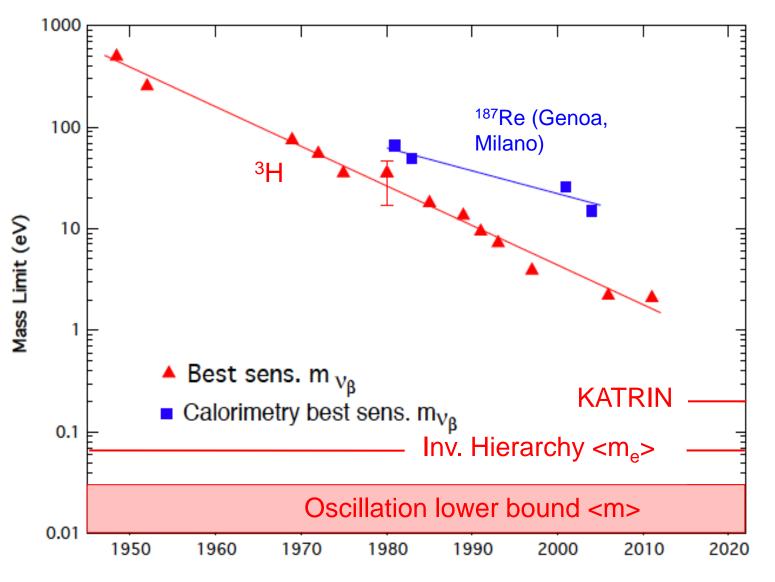
A window to work in

Molecular Excitations

Energy loss function



Neutrino Mass Limits from β decay

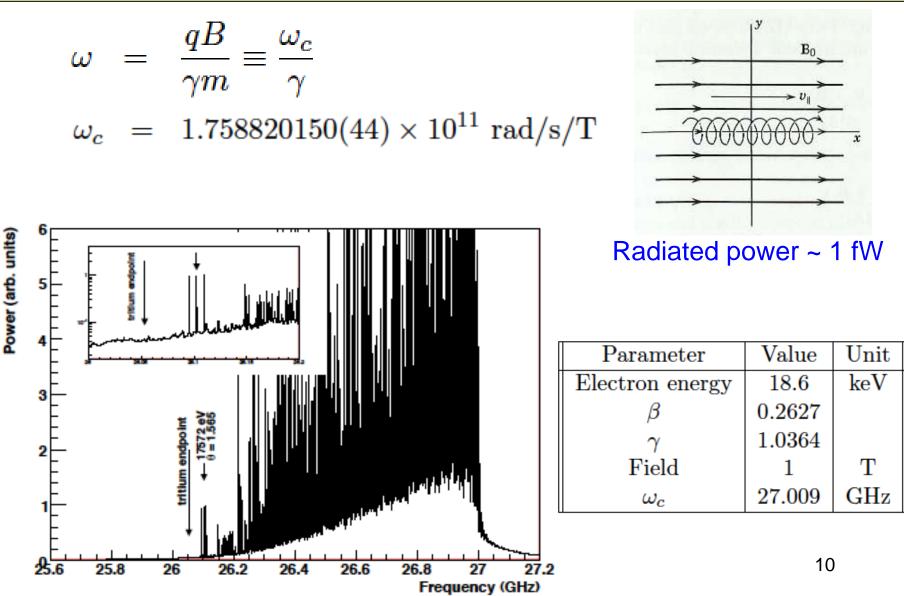


J.F. Wilkerson & HR

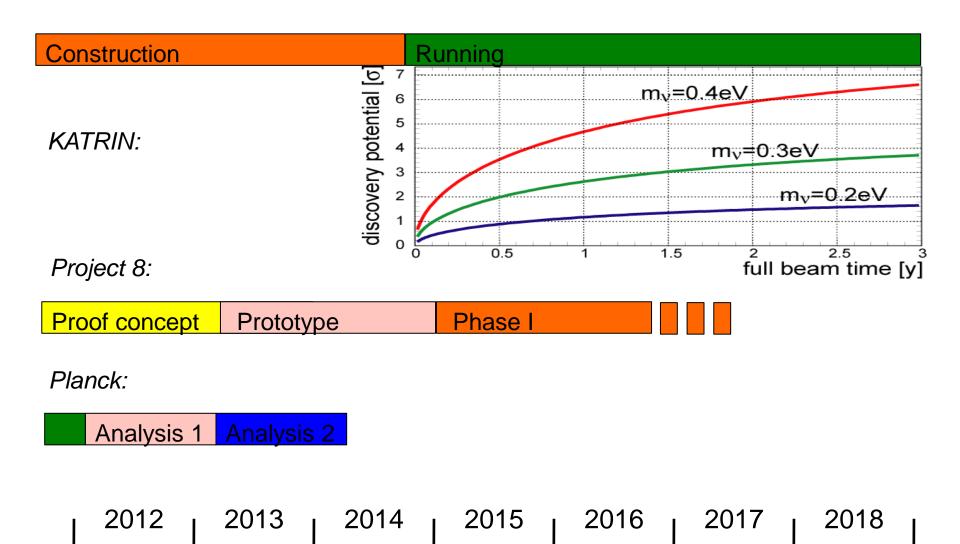


Cyclotron radiation from tritium beta decay

(B. Monreal and J. Formaggio, PRD 80:051301, 2009)



Neutrino mass: some milestones



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