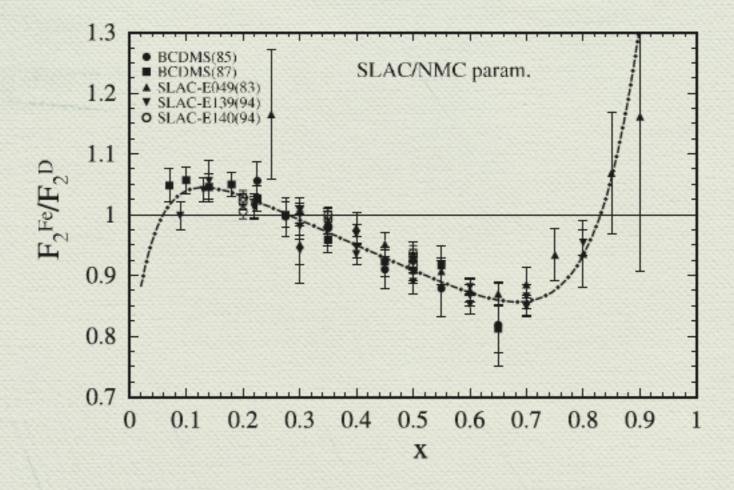
Nuclear PDFs

Pia Zurita

Universidad de Buenos Aires

CERN-ECFA-NuPECC Workshop on the LHeC, Chavannes-de-Bogis, June 2012

effects in cross-sections = effect on the PDFs?



nuclear PDFs = parton densities in bound nucleons

factorization for e-p DIS

$$\frac{d^2\sigma}{dxdQ^2} = \sum_{a} f_a \otimes d\hat{\sigma}_{al \to l'X} + \mathcal{O}\left(\frac{1}{Q^2}\right)$$

assume factorization for e-nucleus DIS

$$\frac{e\text{-A}}{\text{collisions}}$$

$$\frac{d^2\sigma}{dxdQ^2} = \sum_a f_a^A \otimes d\hat{\sigma}_{al \to l'X} + \mathcal{O}\left(\frac{1}{Q^2}\right)$$
 introduce universal nuclear PDF

assume factorization for e-nucleus DIS

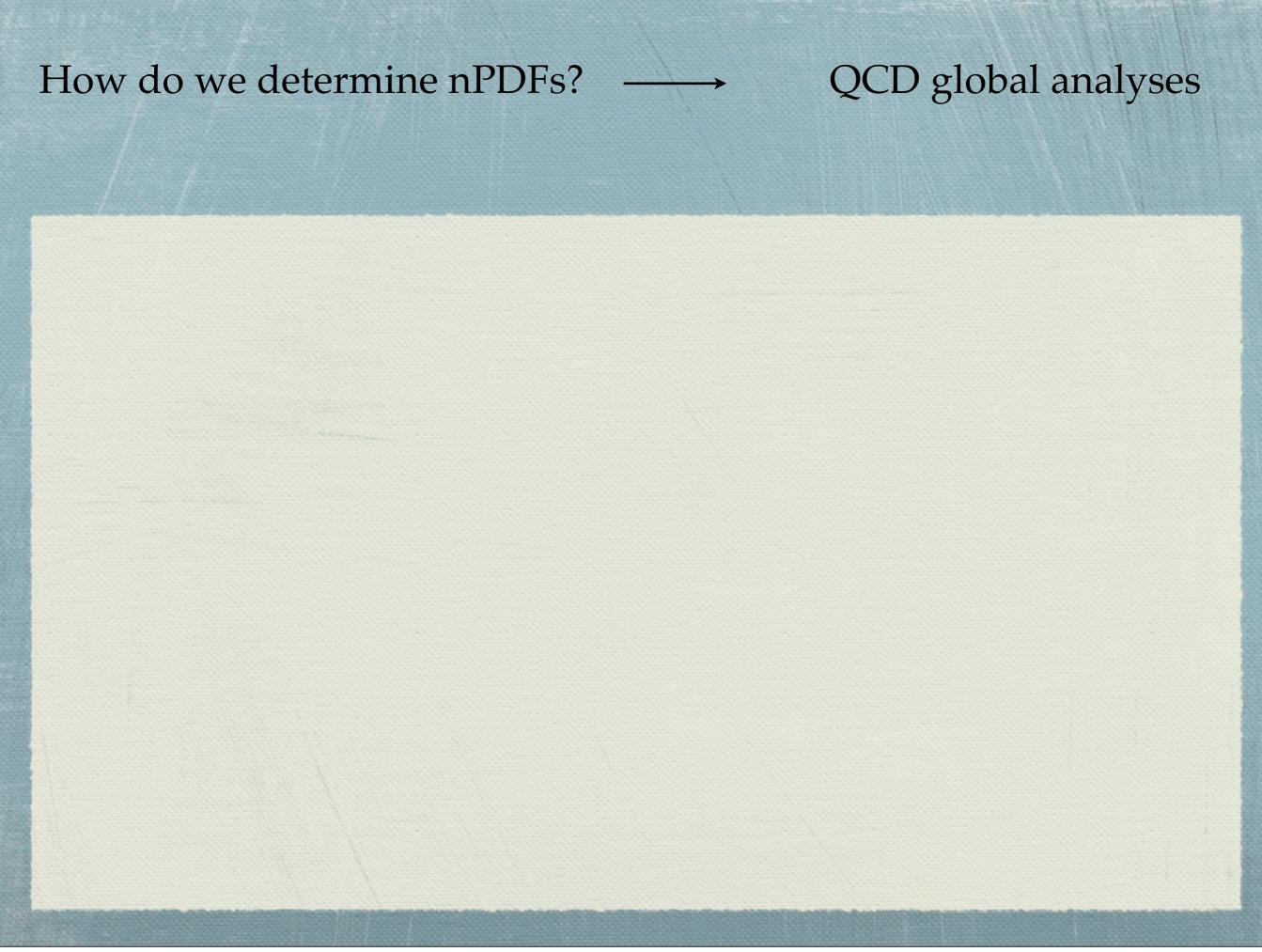
$$\frac{e\text{-A}}{\text{collisions}}$$

$$\frac{d^2\sigma}{dxdQ^2} = \sum_a f_a^A \otimes d\hat{\sigma}_{al \to l'X} + \mathcal{O}\left(\frac{1}{Q^2}\right)$$
 introduce universal nuclear PDF

- same DGLAP scale evolution
- same hard scattering cross sections

as for free proton PDFs

all nuclear effects are absorbed into non-perturbative nPDFs



QCD global analyses

different strategies for parameterization

Convolute with a weight function

$$f_i^A(x_N, Q_0^2) = \int_{x_N}^A \frac{dy}{y} W_i(y, A, Z) f_i(x_N/y, Q_0^2)$$

de Florian, Sassot [PRD69(2004)074028]

How do we determine nPDFs?

QCD global analyses

different strategies for parameterization

Convolute with a weight function

$$f_i^A(x_N, Q_0^2) = \int_{x_N}^A \frac{dy}{y} W_i(y, A, Z) f_i(x_N/y, Q_0^2)$$

de Florian, Sassot [PRD69(2004)074028]

Find them from scratch

$$f_i^A(x_N, Q_0^2) = f_i(x_N, Q_0^2, A)$$

$$f_i(x_N, Q_0^2) = f_i(x_N, Q_0^2, A = 1)$$

nCTEQ [PRD80(2009)094004]

How do we determine nPDFs? ——

QCD global analyses

different strategies for parameterization

Convolute with a weight function

$$f_i^A(x_N, Q_0^2) = \int_{x_N}^A \frac{dy}{y} W_i(y, A, Z) f_i(x_N/y, Q_0^2)$$

de Florian, Sassot [PRD69(2004)074028]

Find them from scratch

$$f_i^A(x_N, Q_0^2) = f_i(x_N, Q_0^2, A)$$

$$f_i(x_N, Q_0^2) = f_i(x_N, Q_0^2, A = 1)$$

nCTEQ [PRD80(2009)094004]

Relate to the proton PDFs by a multiplicative factor

$$f_i^A(x_N, Q_0^2) = R_i(x_N, Q_0^2, A) f_i(x_N, Q_0^2)$$

Hirai, Kumano, Nagai [PRC76(2007)065207] Eskola, Paukkunen, Salgado [JHEP0904(2009)065] de Florian, Sassot, Stratmann, Z. [PRD85(**2012**)074028]

convolution (nDS)

first NLO analysis

multiplicative factor (EPS)

first dAu

from scratch (nCTEQ)

first neutrino DIS

convolution (nDS)

multiplicative factor (EPS)

from scratch (nCTEQ)

first NLO analysis

first dAu

first neutrino DIS

DIS & DY

DIS & DY

DIS & DY

convolution (nDS)

multiplicative factor (EPS)

from scratch (nCTEQ)

first NLO analysis

first dAu

first neutrino DIS

DIS & DY

DIS & DY

DIS & DY

small effect on gluons

factorization and universality ✓



convolution (nDS)

multiplicative factor (EPS)

from scratch (nCTEQ)

first NLO analysis

first dAu

first neutrino DIS

DIS & DY

DIS & DY

DIS & DY

RHIC dAu pion production

small effect on gluons

factorization and universality ✓



convolution multiplicative factor from scratch (nDS) (EPS) (nCTEQ) first dAu first neutrino DIS first NLO analysis DIS & DY DIS & DY DIS & DY RHIC dAu pion production huge anti-shadowing/ small effect on gluons EMC effect for gluon factorization and universality ✓

convolution multiplicative factor from scratch (nDS) (EPS) (nCTEQ) first neutrino DIS first NLO analysis first dAu DIS & DY DIS & DY DIS & DY RHIC dAu pion production huge anti-shadowing/ small effect on gluons EMC effect for gluon factorization and universality ✓ factorization and universality ✓

<u>convolution</u> (nDS)	multiplicative factor (EPS)	<u>from scratch</u> (nCTEQ)
first NLO analysis	first dAu	first neutrino DIS
DIS & DY	DIS & DY	DIS & DY
	RHIC dAu pion production	neutrino DIS
small effect on gluons	huge anti-shadowing/ EMC effect for gluon	
factorization and universality ✓	factorization and universality ✓	

convolution multiplicative factor from scratch (nDS) (EPS) (nCTEQ) first neutrino DIS first NLO analysis first dAu DIS & DY DIS & DY DIS & DY RHIC dAu pion production neutrino DIS huge anti-shadowing/ tension electron DIS/ small effect on gluons EMC effect for gluon neutrino DIS factorization and universality ✓ factorization and universality ✓

convolution (nDS)	multiplicative factor (EPS)	<u>from scratch</u> (nCTEQ)
first NLO analysis	first dAu	first neutrino DIS
DIS & DY	DIS & DY	DIS & DY
	RHIC dAu pion production	neutrino DIS
	To equilibrium. To experience the second sec	
small effect on gluons	huge anti-shadowing/ EMC effect for gluon	tension electron DIS/ neutrino DIS
factorization and universality ✓	factorization and universality \checkmark	factorization and universality?

D. de Florian, R. Sassot, M. Stratmann, P. Z. [PRD85(2012)074028]

DSSZ

electron DIS $x\sim[0.01-1]$ (valence at higher x, sea at $x\sim0.01$)

Drell-Yan $x\sim[0.01-0.2]$ (valence/sea but large errors)

hadroproduction dAu

neutrino DIS (valence/sea)

global

D. de Florian, R. Sassot, M. Stratmann, P. Z. [PRD85(2012)074028]

DSSZ

electron DIS $x\sim[0.01-1]$ (valence at higher x, sea at $x\sim0.01$)

Drell-Yan $x\sim[0.01-0.2]$ (valence/sea but large errors)

hadroproduction dAu

neutrino DIS (valence/sea)

reference PDFs: NLO MSTW (GM-VFNS) A. Martin et al. 2009

same conventions for evolution, alphas, F2 coeff.

massive coeff. for neutrino F2 and F3

J. Bluemlein et al. 2011

massless coeff. for Drell Yan and dAu

global

up-to-date

D. de Florian, R. Sassot, M. Stratmann, P. Z. [PRD85(2012)074028]

DSSZ

electron DIS $x\sim[0.01-1]$ (valence at higher x, sea at $x\sim0.01$)

Drell-Yan $x\sim[0.01-0.2]$ (valence/sea but large errors)

hadroproduction dAu

neutrino DIS (valence/sea)

reference PDFs: NLO MSTW (GM-VFNS) A. Martin et al. 2009

same conventions for evolution, alphas, F2 coeff.

massive coeff. for neutrino F2 and F3

J. Bluemlein et al. 2011

massless coeff. for Drell Yan and dAu

up-to-date

improved hessian approach: eigenvector representation

J. Pumplin et al. 2001

"error sets"

uncertainties

D. de Florian, R. Sassot, M. Stratmann, P. Z. [PRD85(2012)074028]

DSSZ

parameterization

$$f_i^A(x_N, Q_0) = R_i^A(x_N, Q_0) f_i^p(x_N, Q_0)$$

$$R_{\nu}^{A}(x,Q_{0}^{2}) = \varepsilon_{1} x^{\alpha_{\nu}} (1-x)^{\beta_{1}} \times (1+\varepsilon_{2}(1-x)^{\beta_{2}}) (1+a_{\nu}(1-x)^{\beta_{3}})$$

$$R_{s}^{A}(x,Q_{0}^{2}) = R_{\nu}^{A}(x,Q_{0}^{2}) \frac{\varepsilon_{s}}{\varepsilon_{1}} \frac{1+a_{s}x^{\alpha_{s}}}{a_{s}+1}$$

$$R_g^A(x, Q_0^2) = R_v^A(x, Q_0^2) \frac{\varepsilon_g}{\varepsilon_1} \frac{1 + a_g x^{\alpha_g}}{a_g + 1}$$

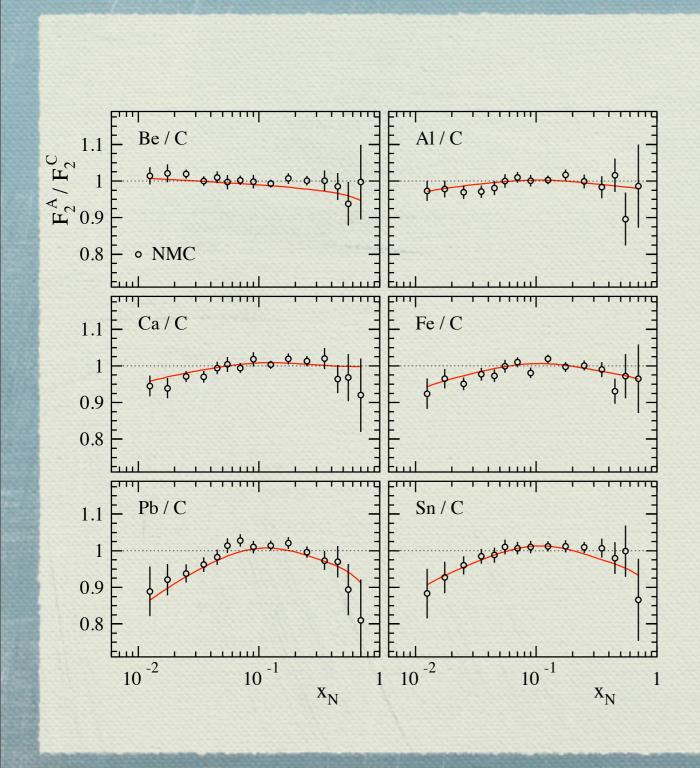
$$\alpha_{\nu}, \alpha_{s}, \alpha_{g}, \beta_{1}, \beta_{2}, \beta_{3}, a_{\nu}, a_{s}, a_{g}$$

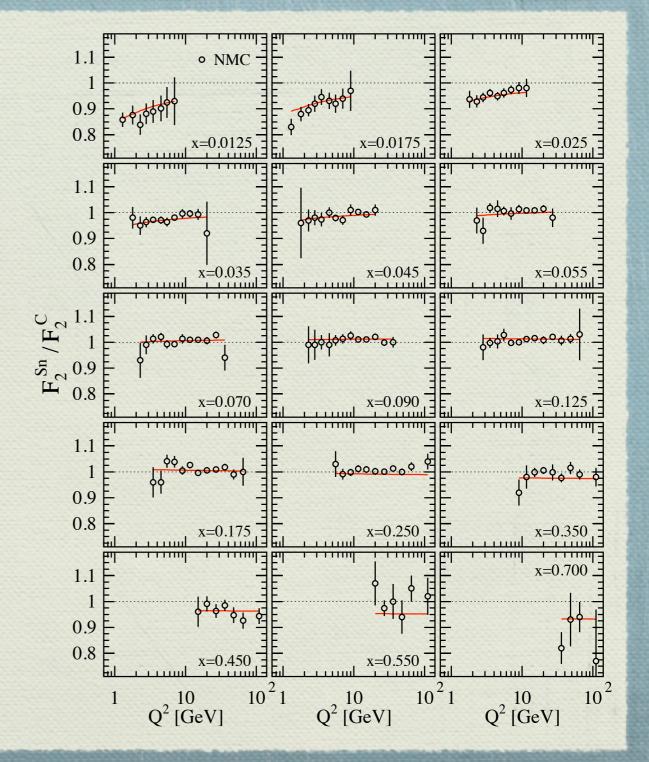
$$\xi = \gamma_{\xi} + \lambda_{\xi} A^{\delta_{\xi}}$$

25 free parameters

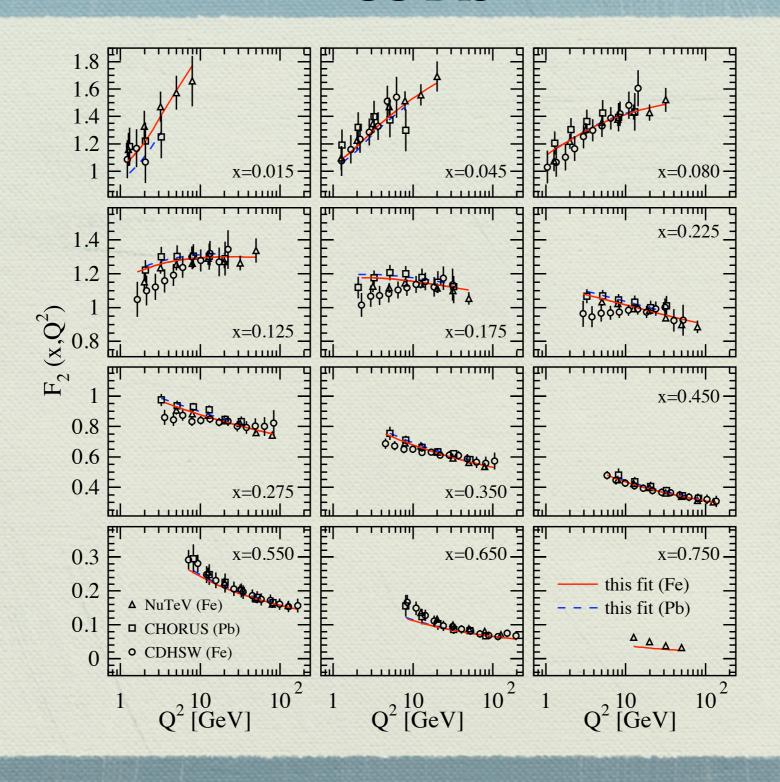
What does DSSZ look like?

NC-DIS



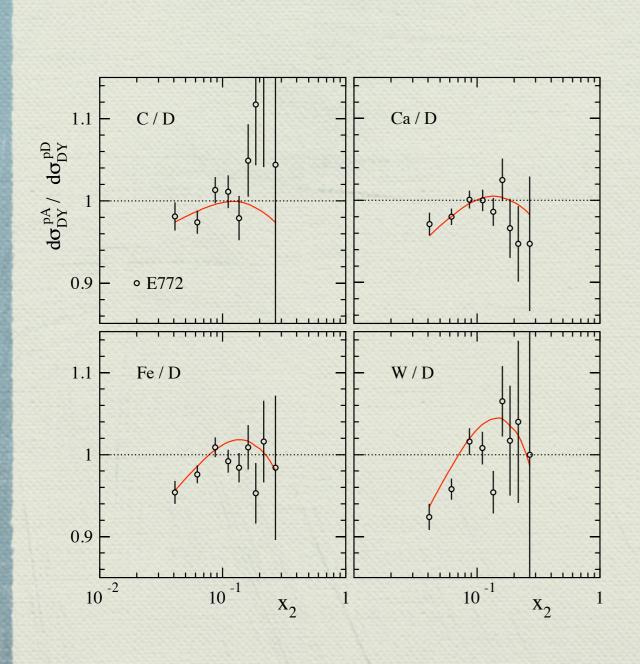


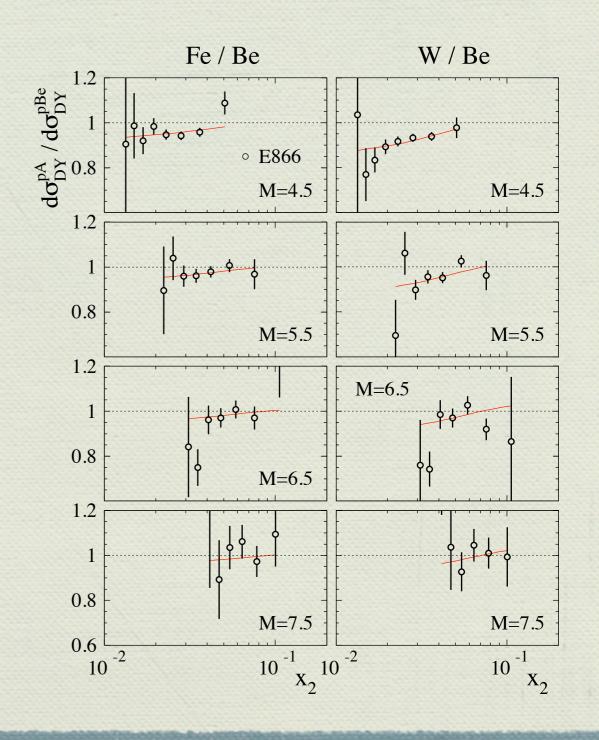
CC-DIS



What does DSSZ look like?

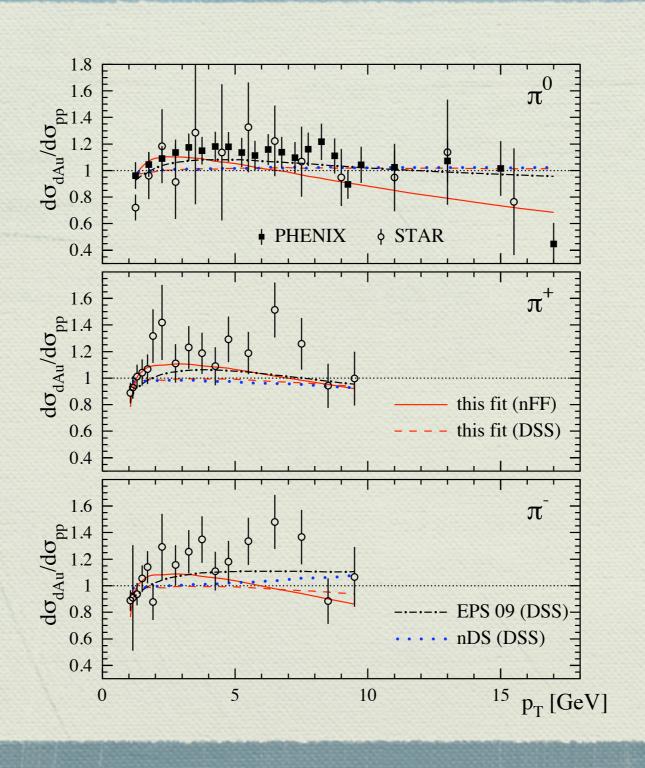
Drell-Yan



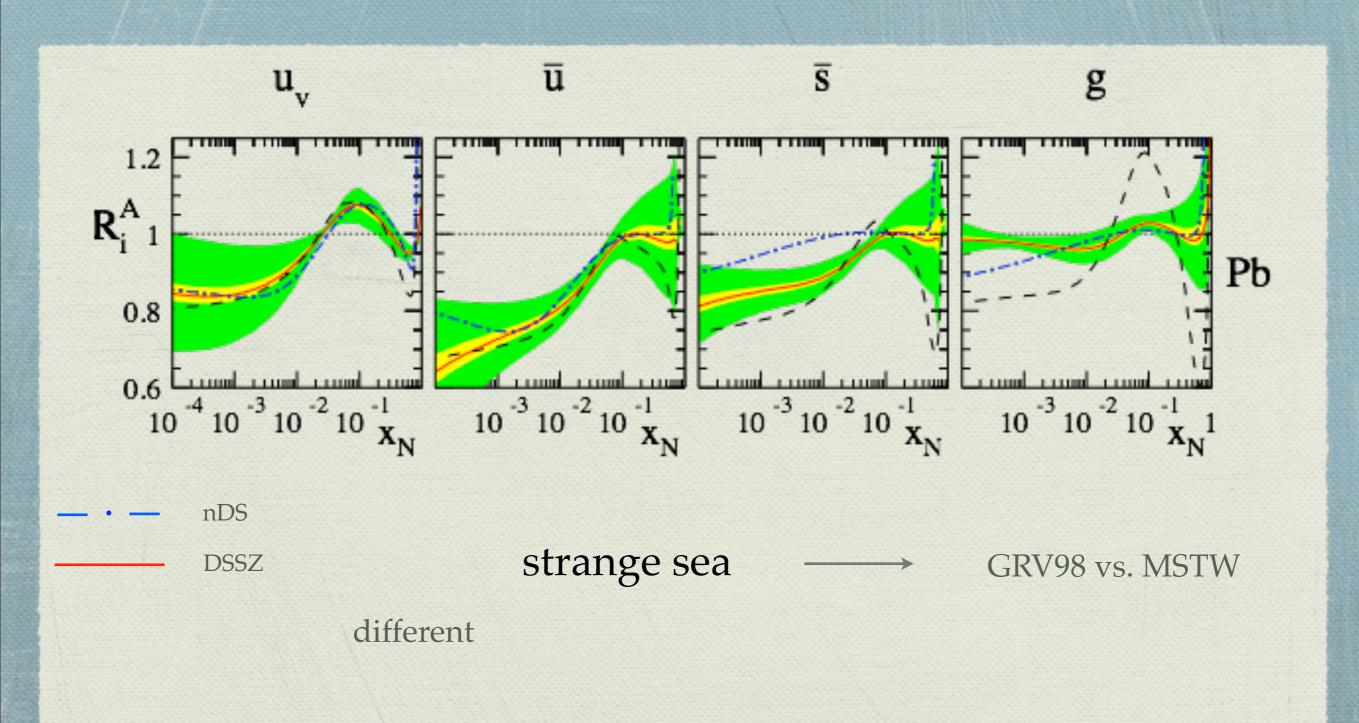


What does DSSZ look like?

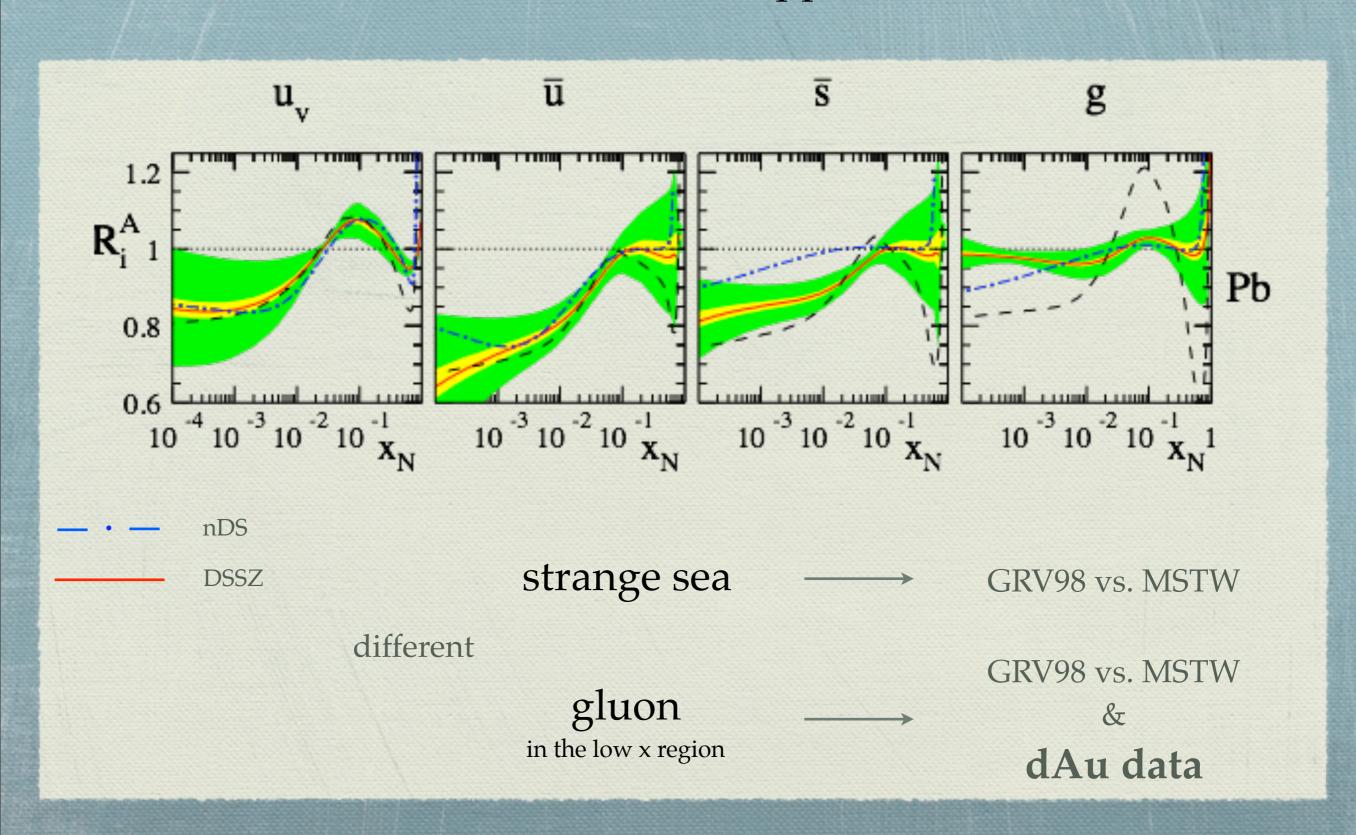
single-inclusive hadroproduction



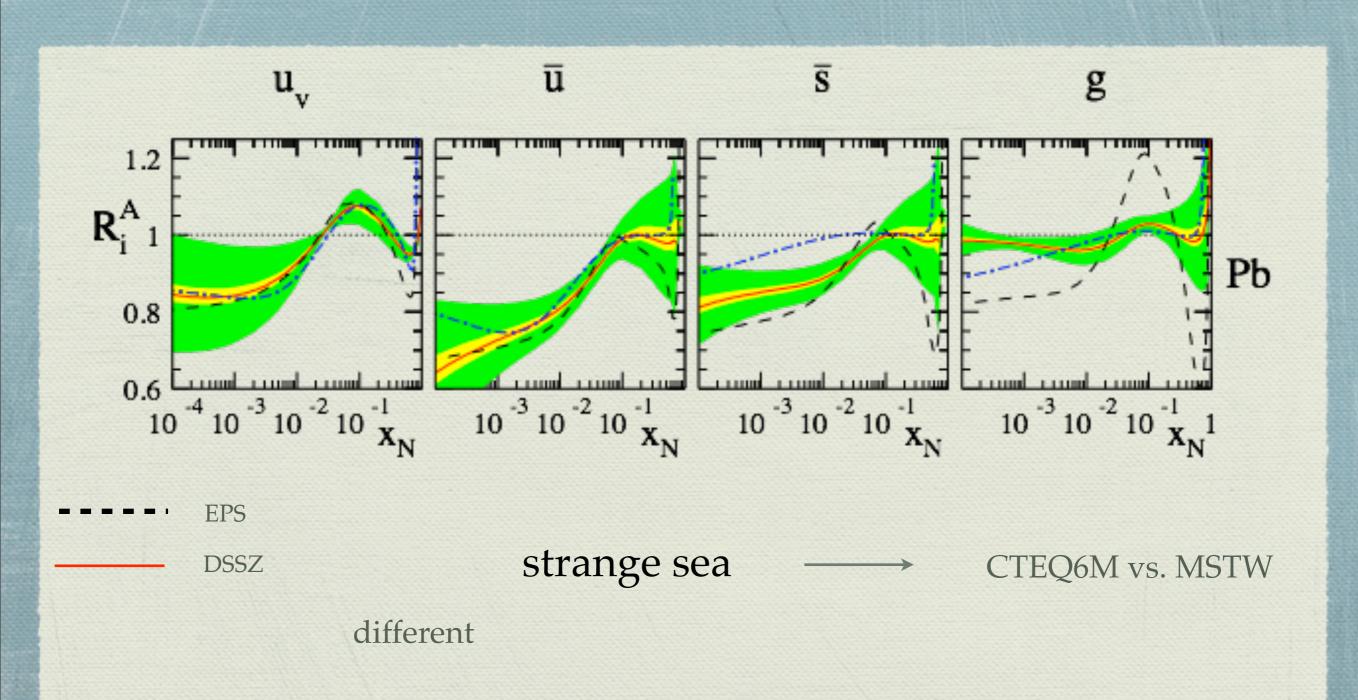
with the convolutional approach (nDS)



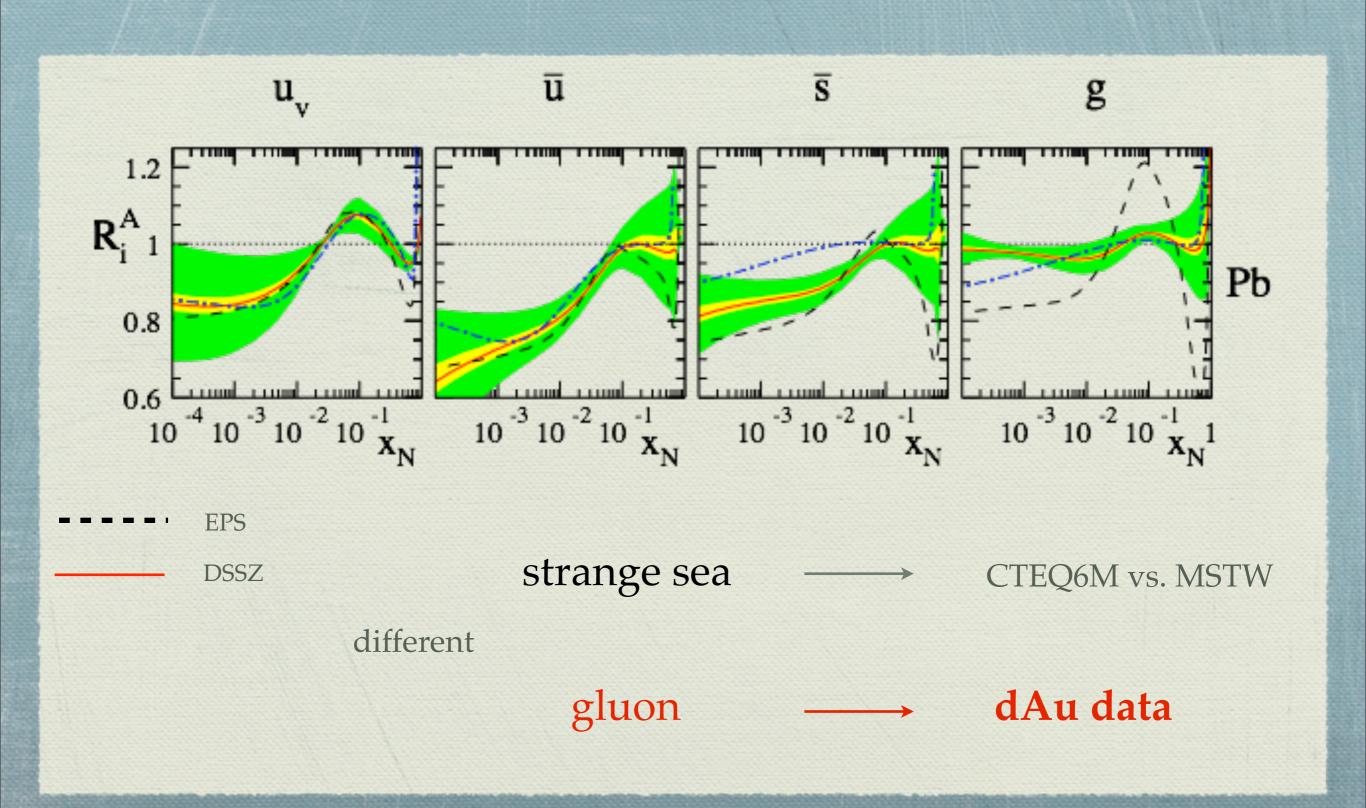
with the convolutional approach (nDS)



with other multiplicative ansatz (EPS)



with other multiplicative ansatz (EPS)



why, if both analyses use similar strategies?

EPS

ZM-VFNS

CTEQ

no neutrino DIS data

parameterization in sectors

compatible
valence
and
u-sea
distributions

DSSZ

GM-VFNS

MSTW

neutrino DIS data

continuous parameterization

why, if both analyses use similar strategies?

EPS

ZM-VFNS

CTEQ

no neutrino DIS data

parameterization in sectors

compatible valence and

u-sea

distributions

DSSZ

GM-VFNS

MSTW

neutrino DIS data

continuous parameterization

no effect for fragmentation in Au

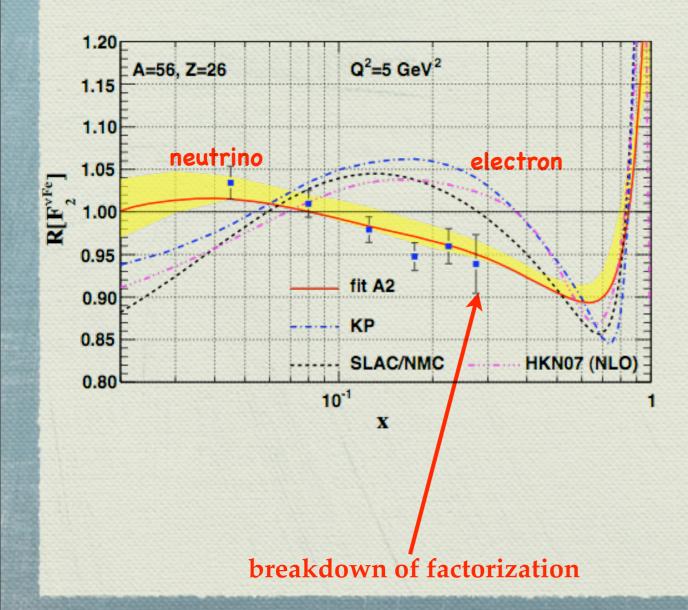
medium-modified fragmentation functions

R.Sassot, M.Stratmann, P. Z. 2010

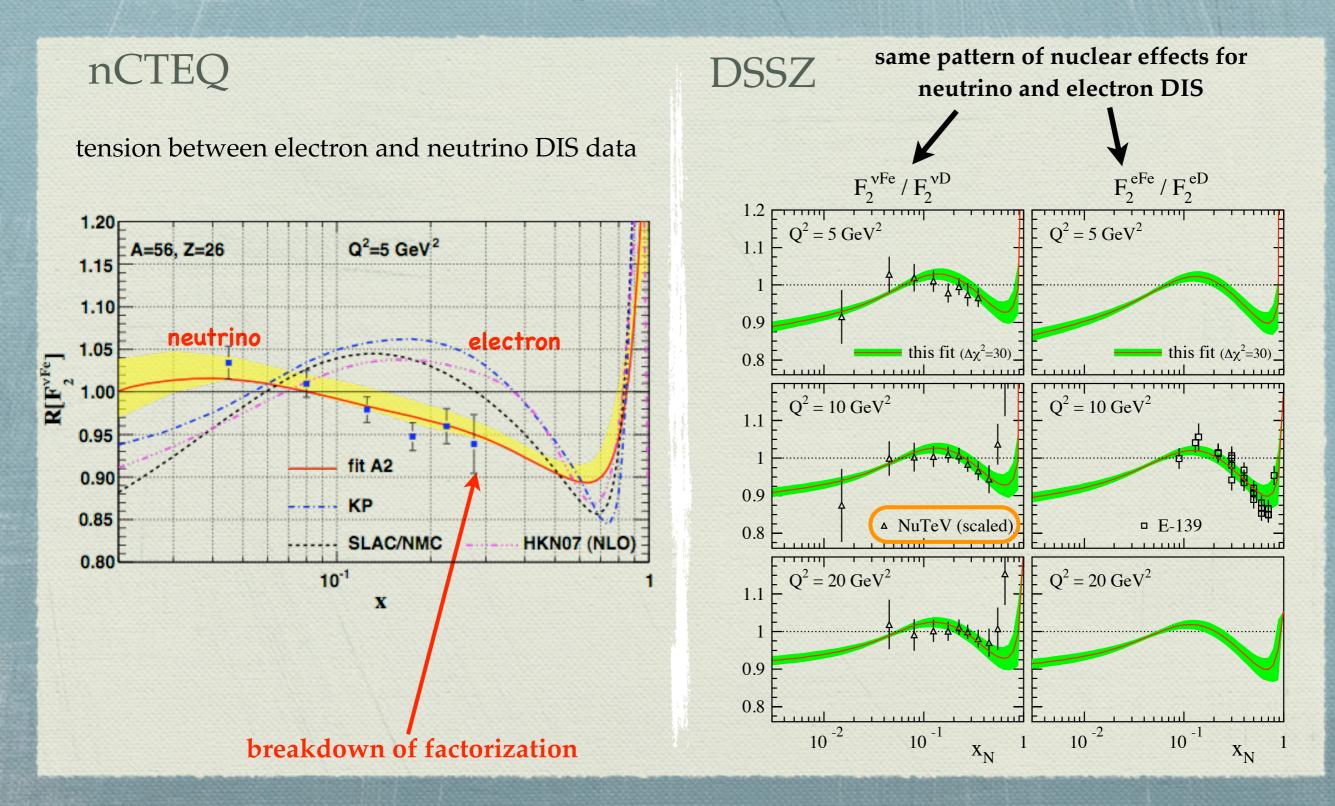
with the direct parameterization (nCTEQ)



tension between electron and neutrino DIS data



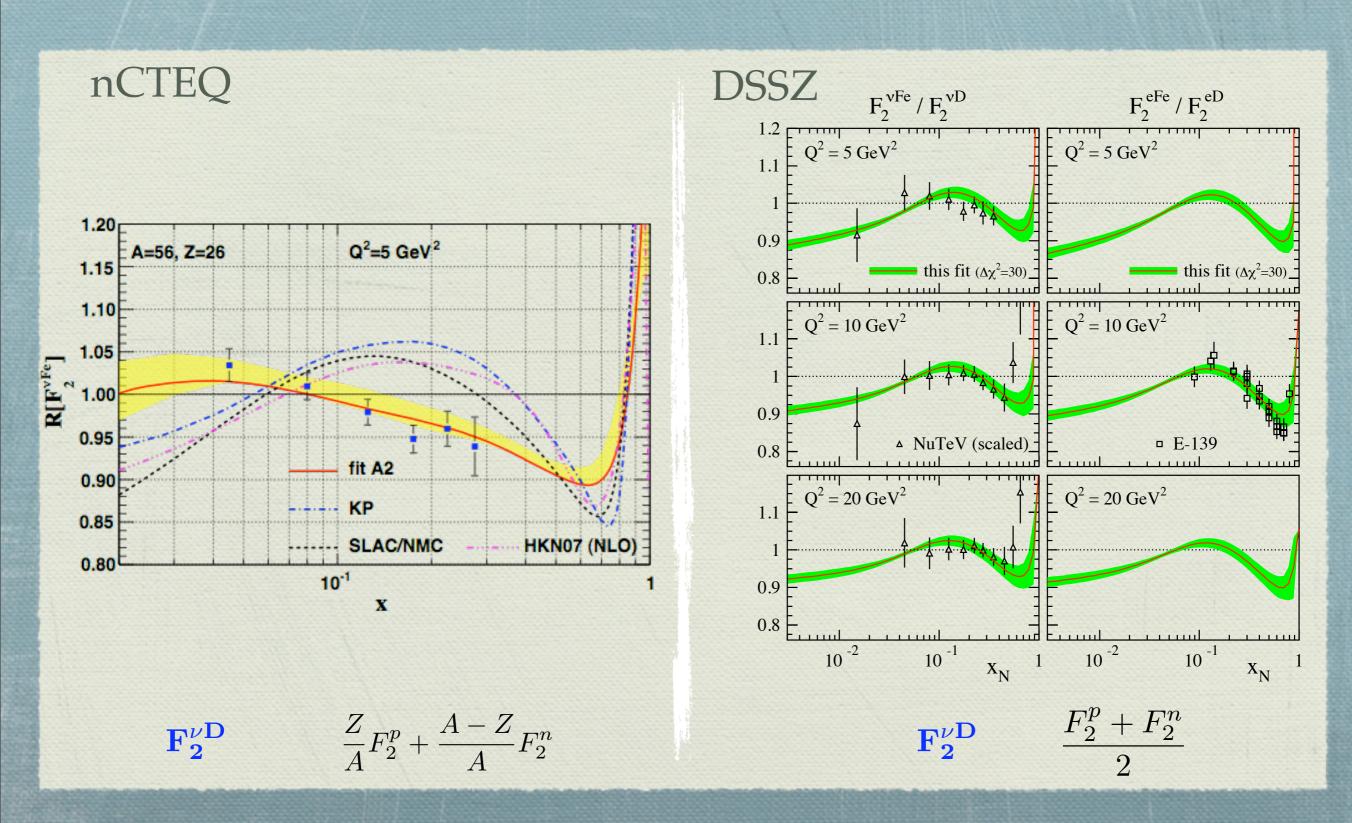
with the direct parameterization (nCTEQ)



what is different?

"theoretical data": $\mathbf{F_2^{\nu D}}$ not measured

must be computed



Summing up...

support of factorization & universality

different nPDFs sets lead to fairly similar distributions

gluons not well determined: very different shapes for Rg

more precise gluon-dependent observables needed

Summing up...

support of factorization & universality

different nPDFs sets lead to fairly similar distributions

gluons not well determined: very different shapes for Rg

more precise gluon-dependent observables needed

What's next?

p-Pb@LHC

what can we look?

for initial state effects



Drell-Yan & prompt photons

p-Pb@LHC

what can we look?

for initial state effects



Drell-Yan & prompt photons

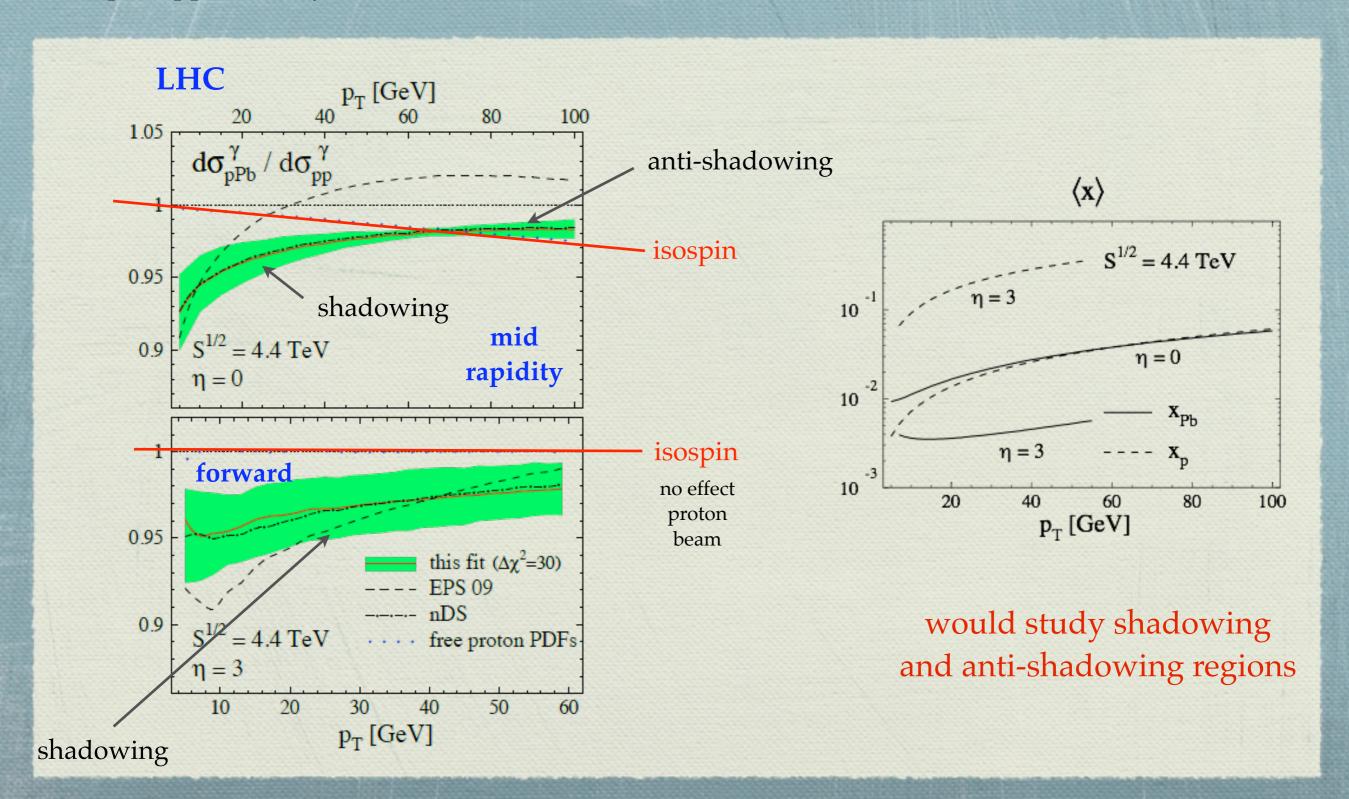
to check for final state effects



Prompt-photons

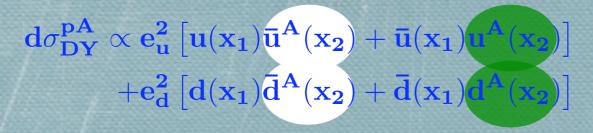
"isospin effects" = dilution of u-quark density from neutrons $\mathbf{u}^{\mathbf{A}}(\mathbf{x}) < \mathbf{u}^{\mathbf{p}}(\mathbf{x})$

ratio pPb/pp not unity even without nuclear modifications



Drell-Yan

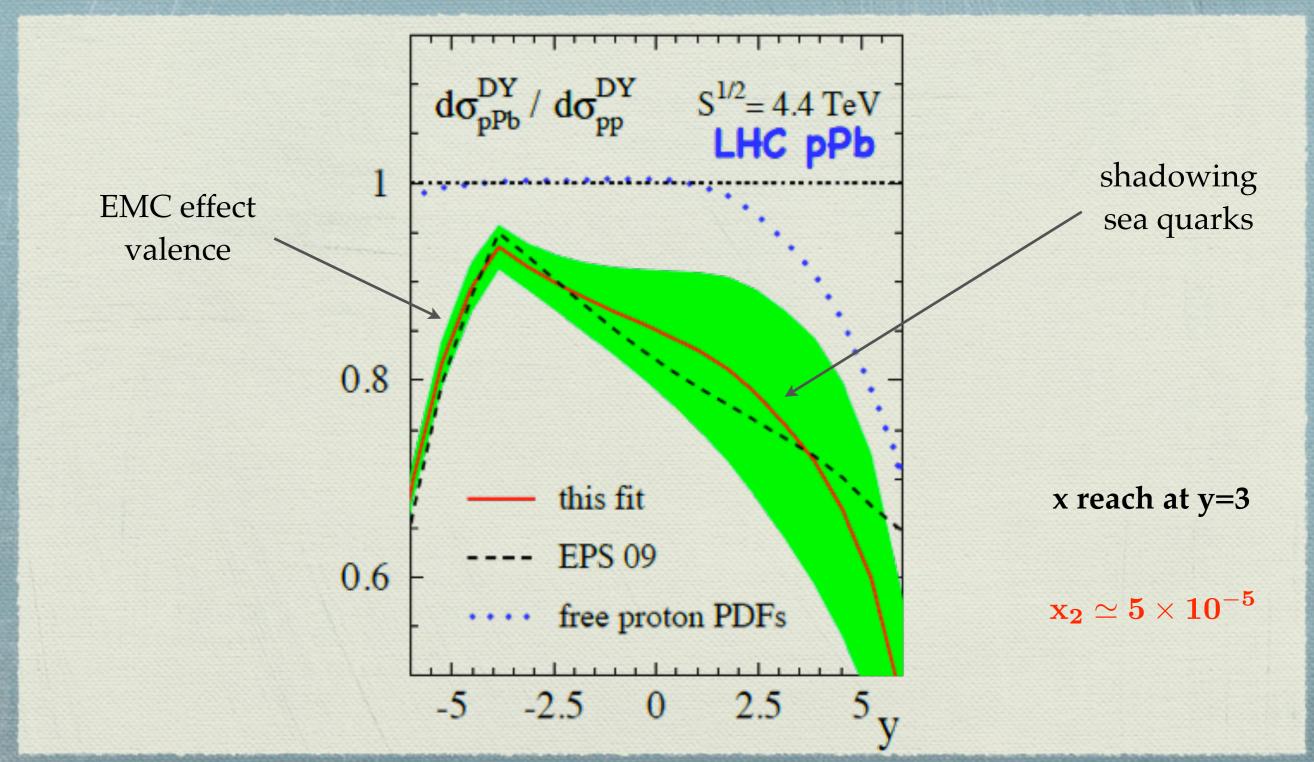
LO

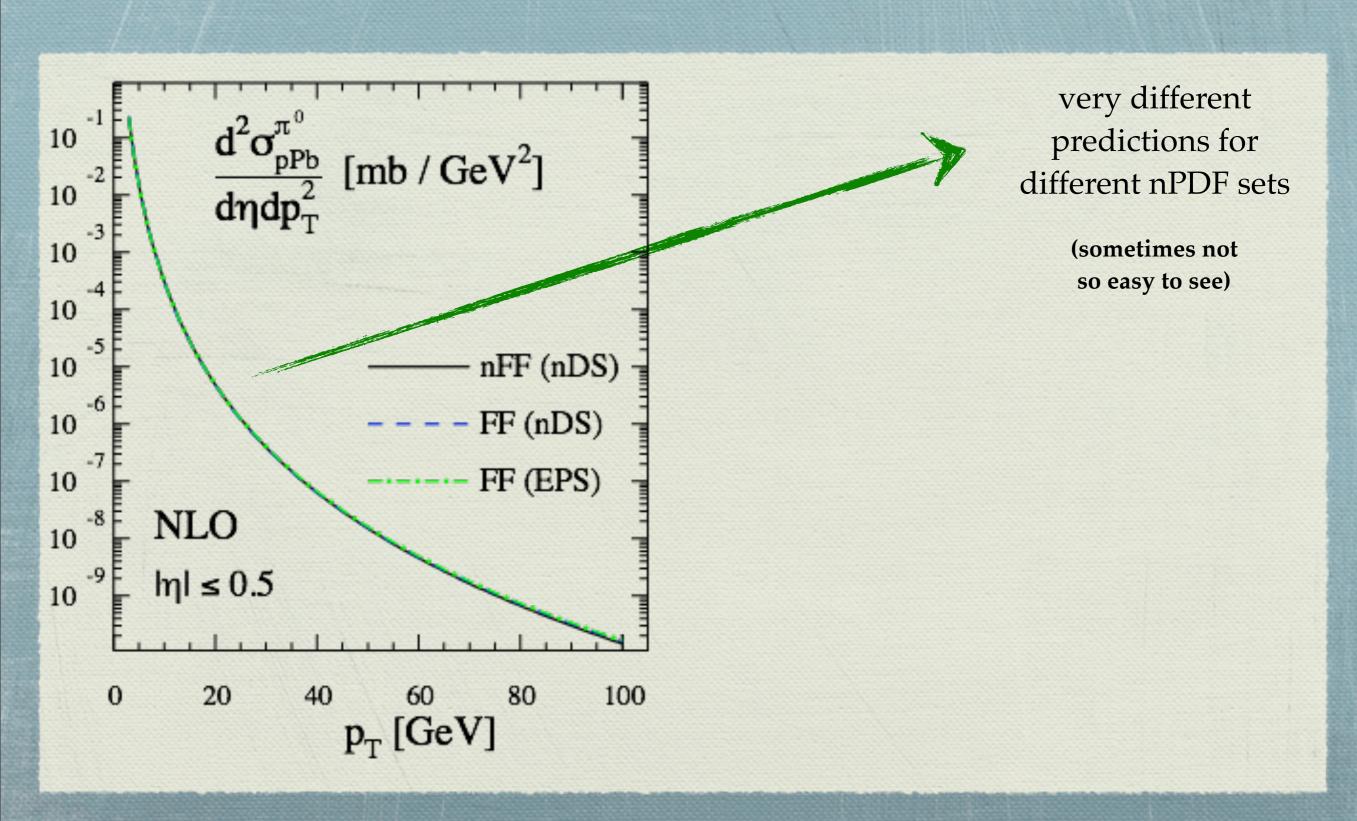


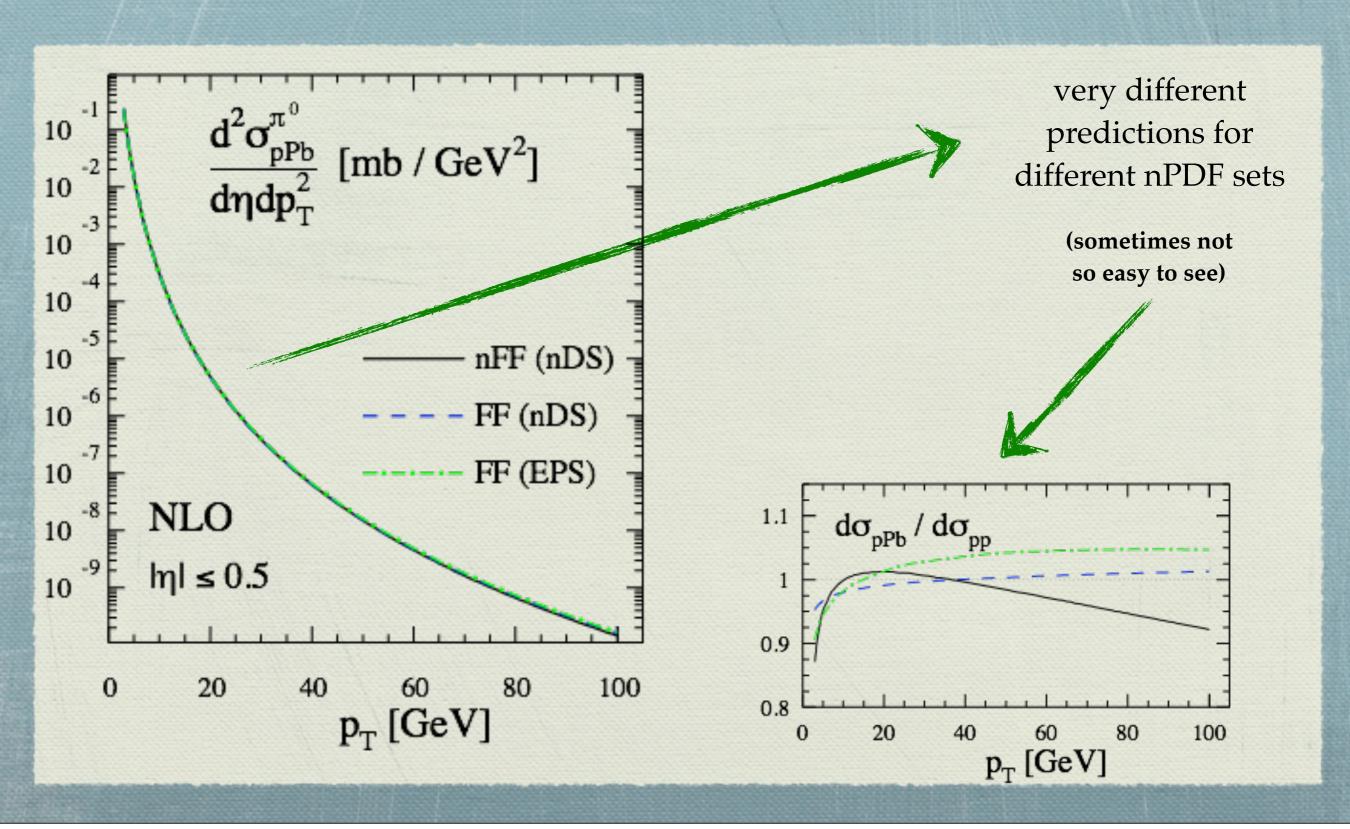
large positive y

large negative y

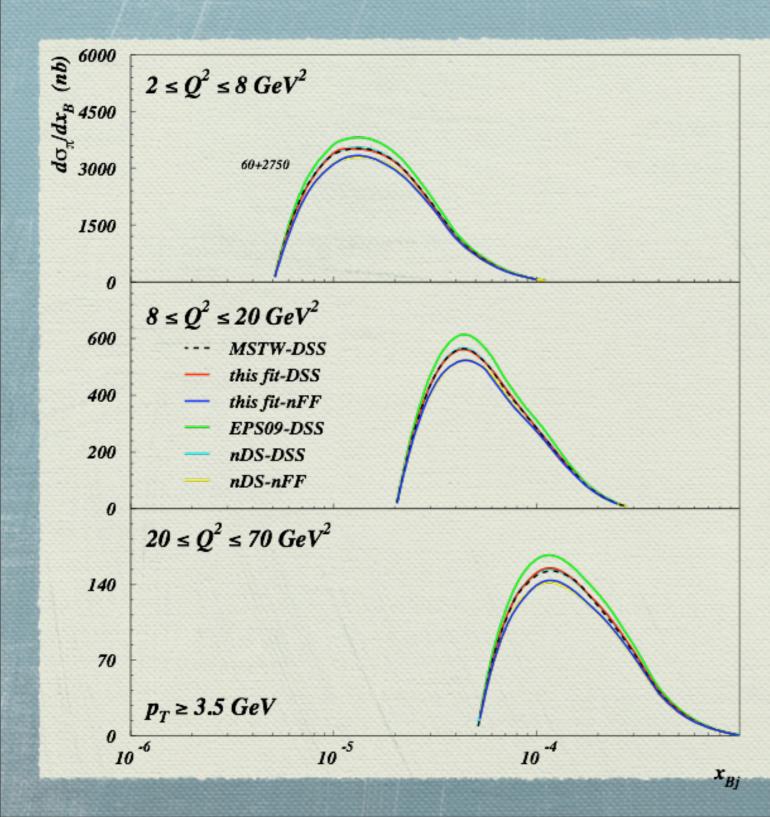




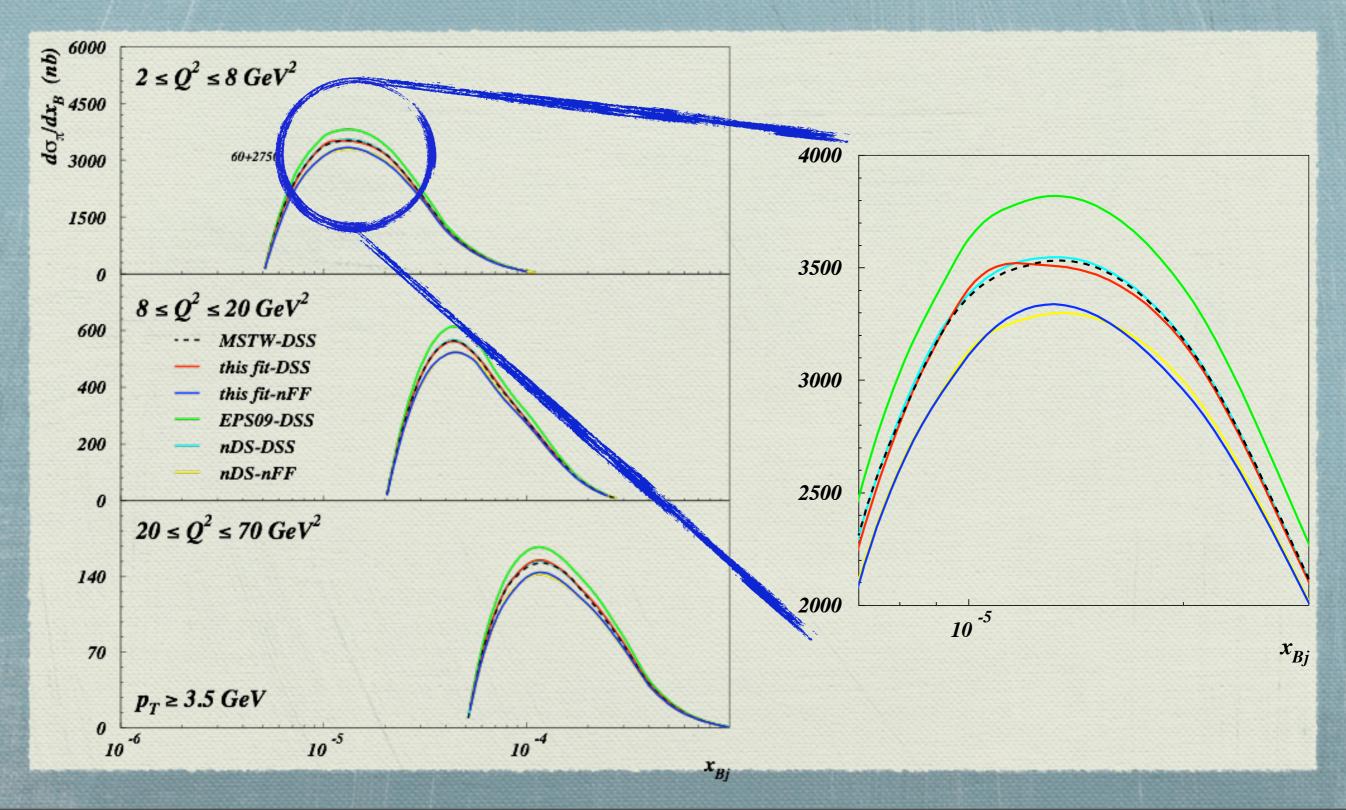




What about LHeC?



What about LHeC?



Summary

several sets of nPDFs available, all (but one) agree on factorization

strange and gluon distributions not yet fully constrained by data

p-Pb @LHC will be crucial to study the low x region

LHeC to cross check results



Why do we need nuclear parton distributions?

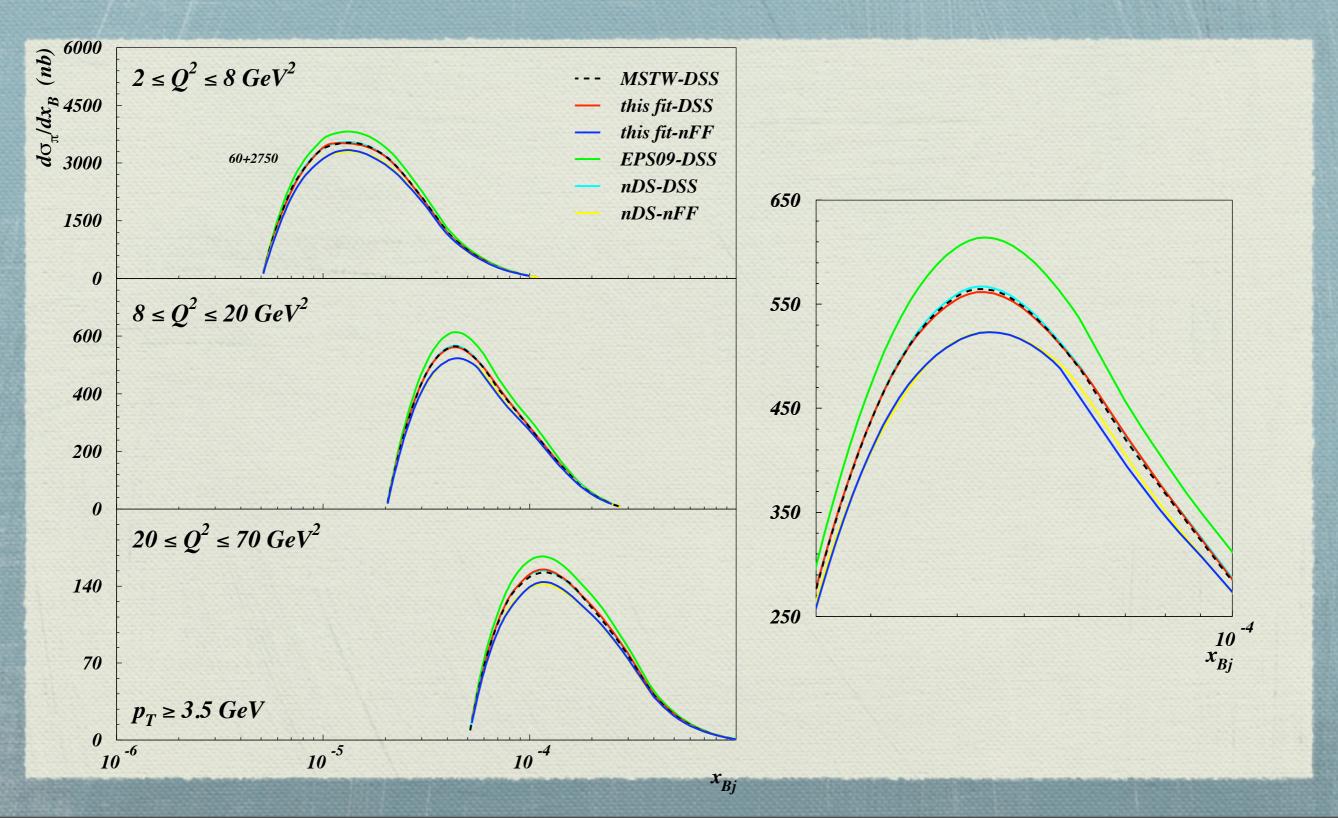
crucial ingredient for perturbative description of

o neutrino DIS with heavy targets (NuTeV, CHORUS, CDHSW)

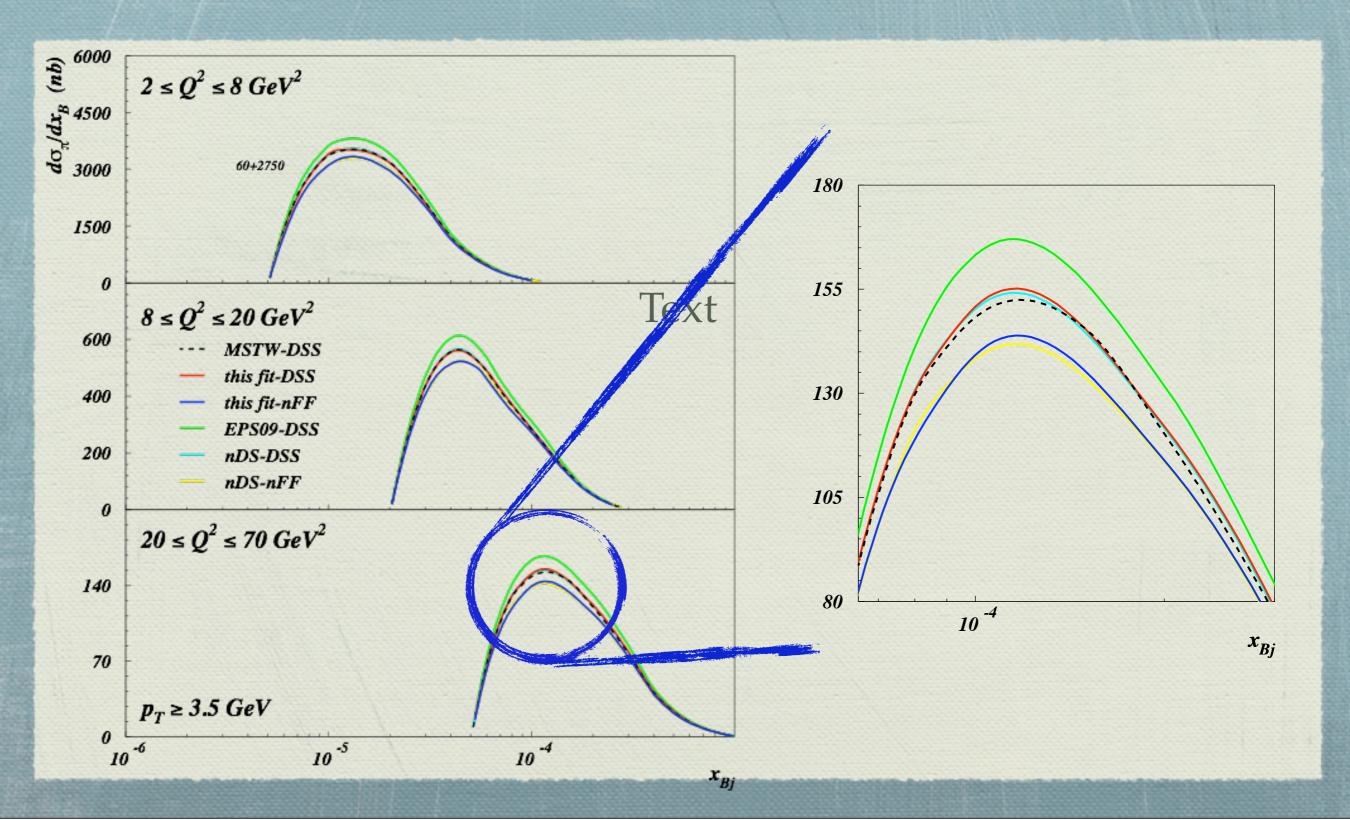
O proton(deuteron)-heavy ion collision (dAu @ RHIC, pPb @LHC)

O electron DIS with nuclei (LHeC, EIC)

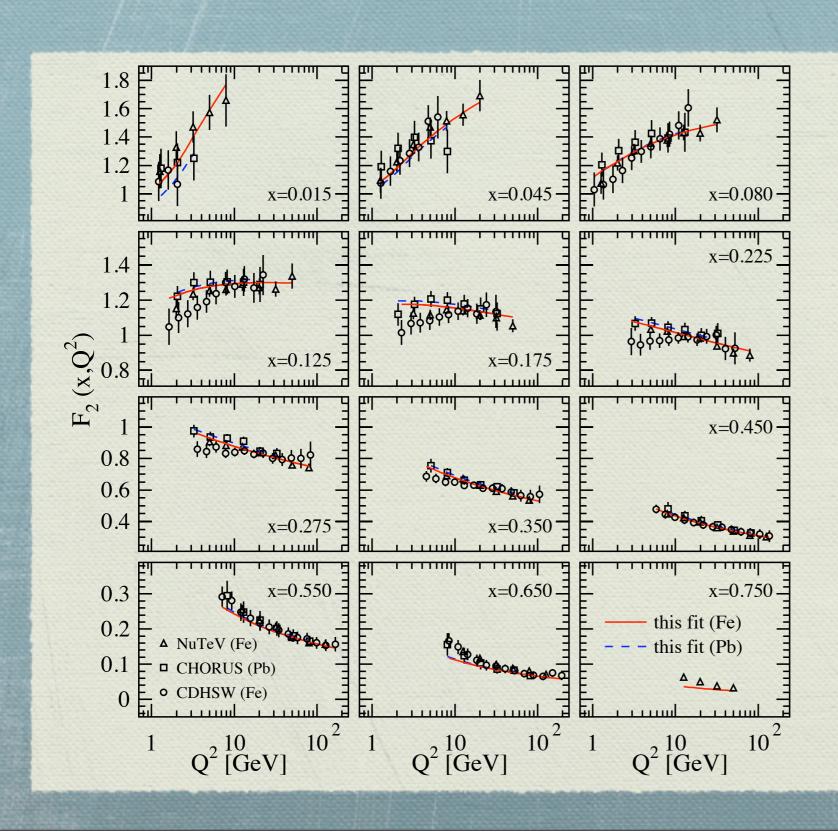
What about LHeC?



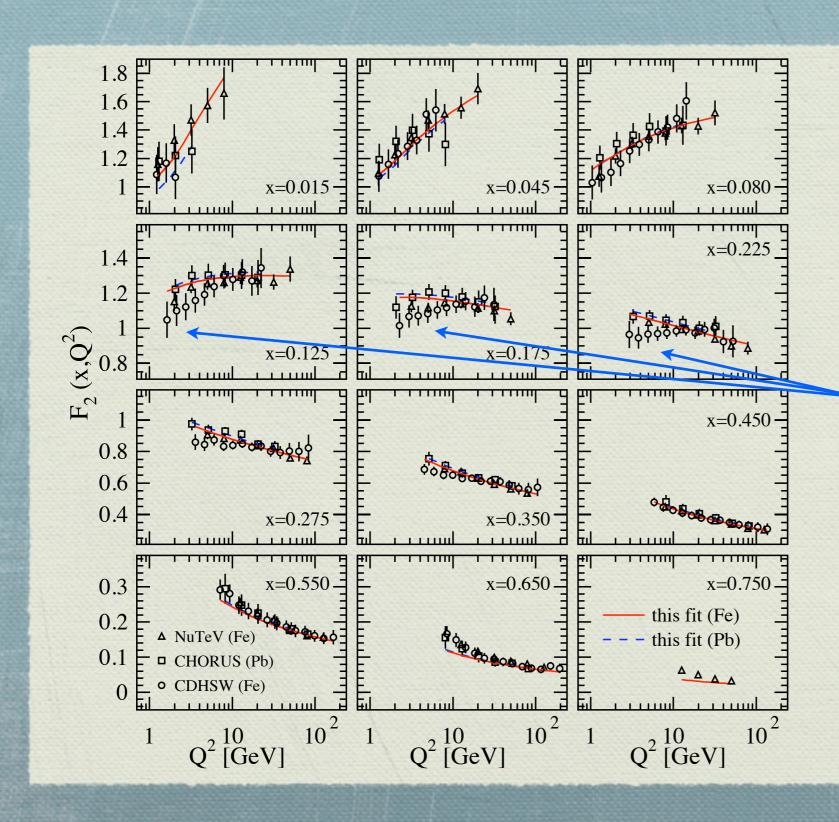
What about LHeC?



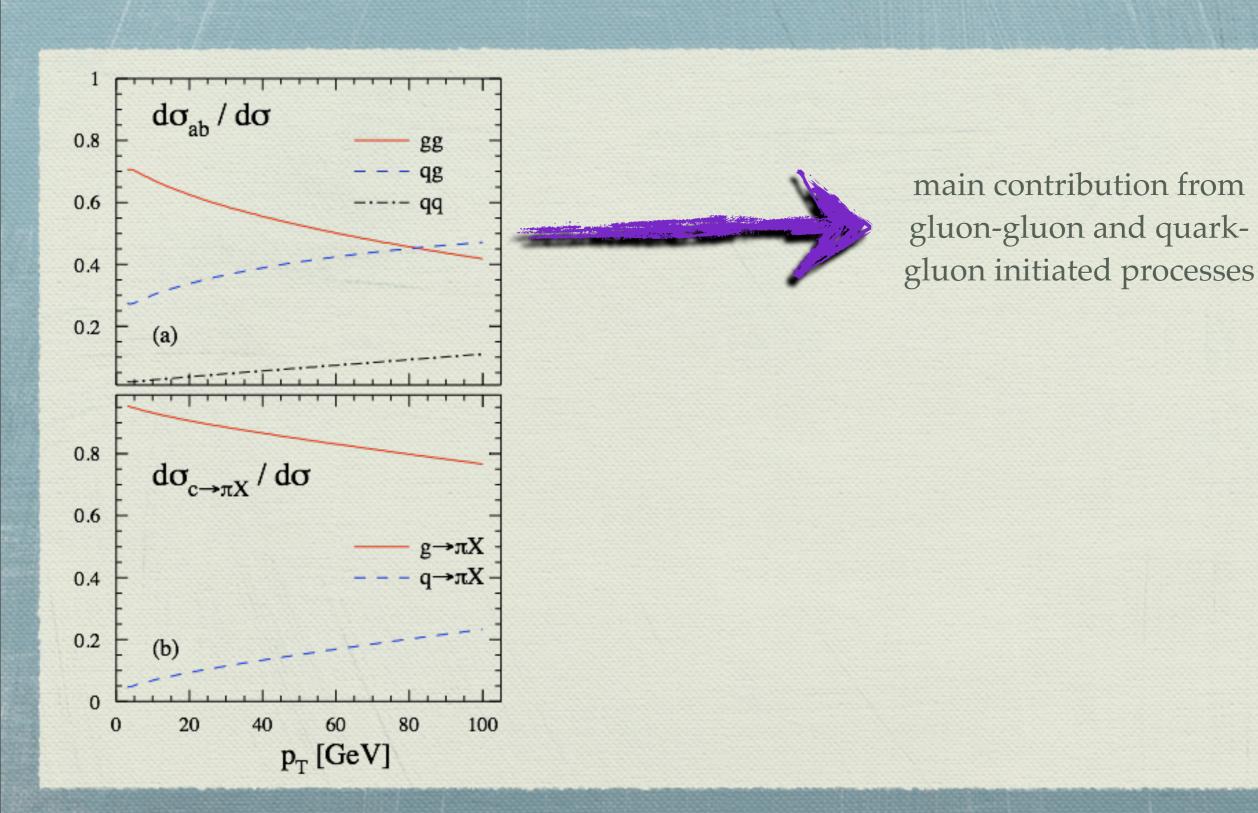
DSSZ gives a good description of CC-DIS

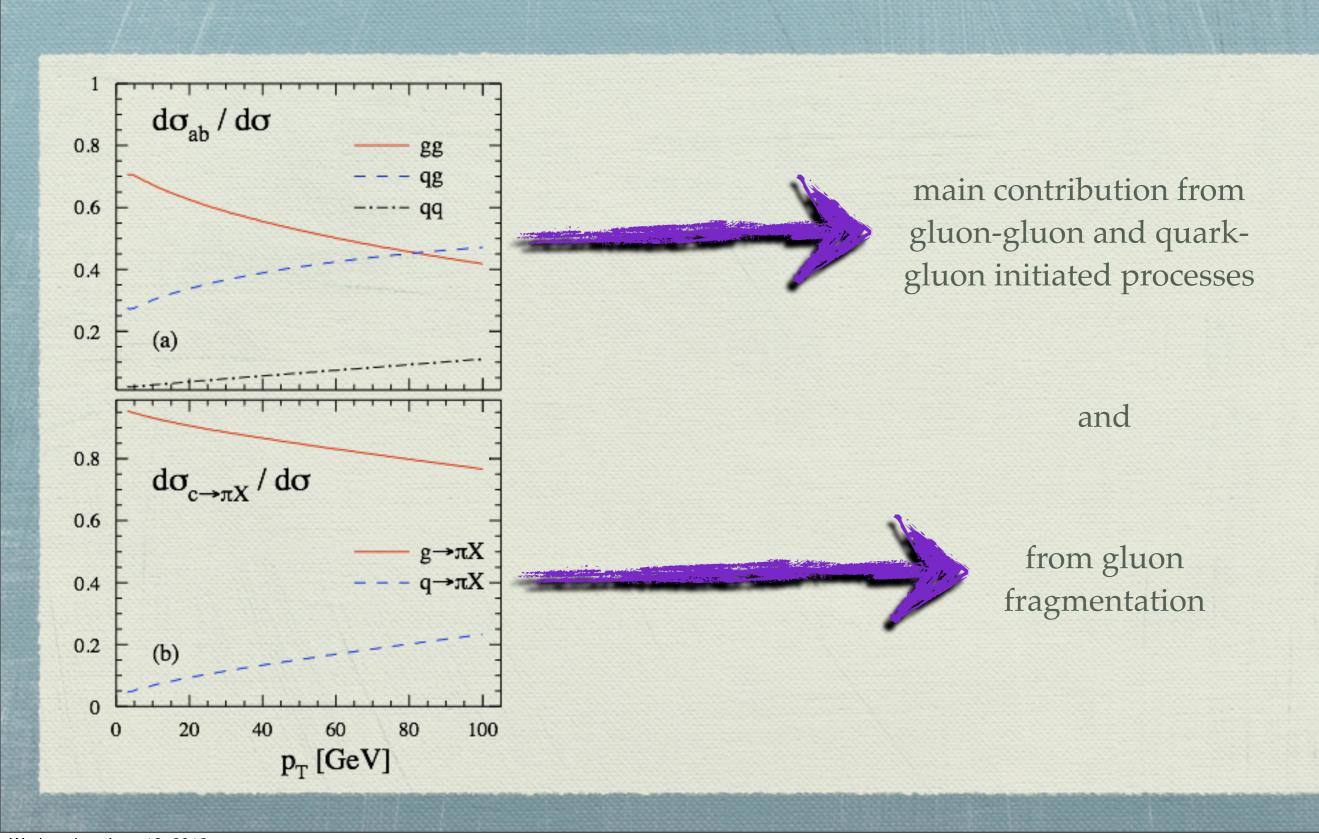


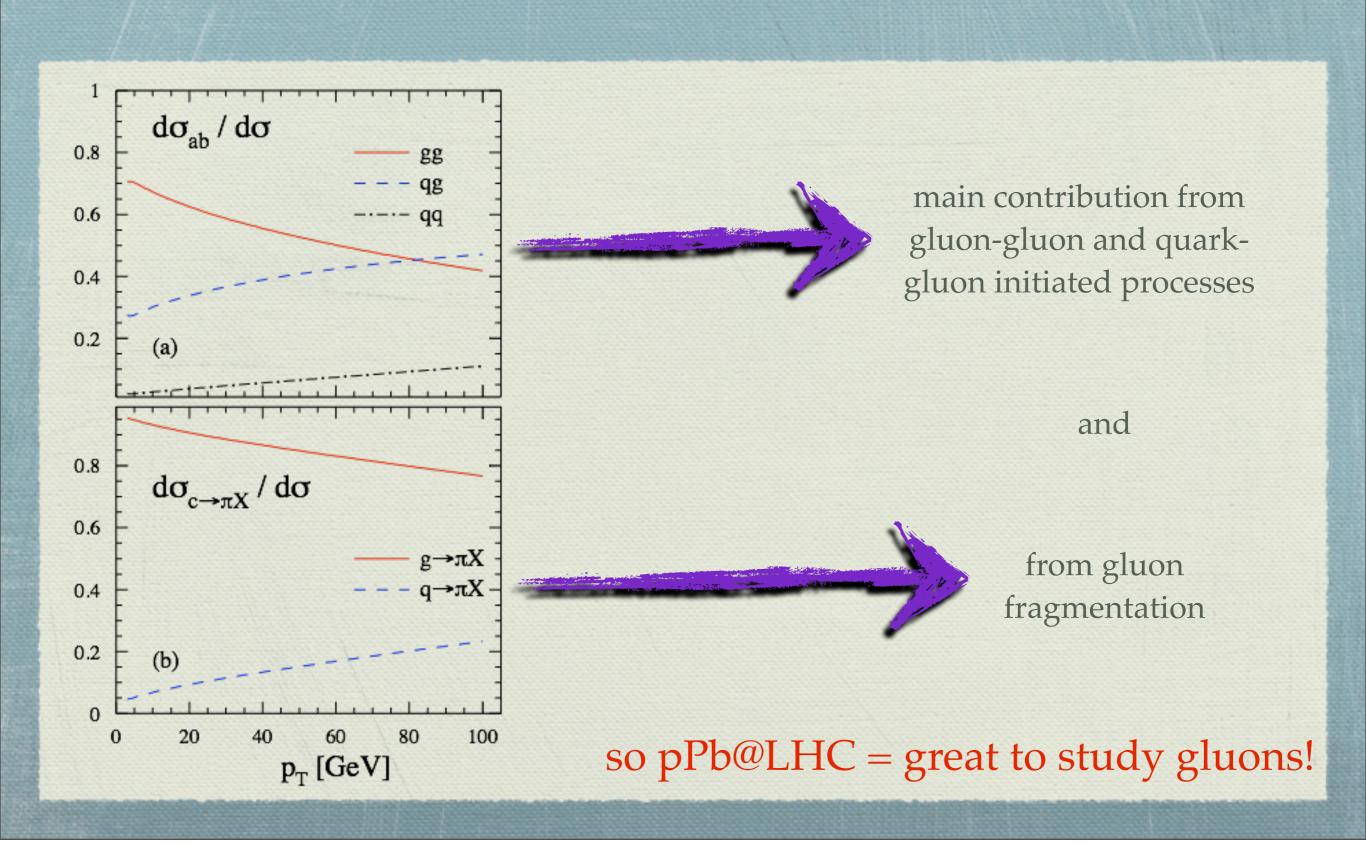
DSSZ gives a good description of CC-DIS



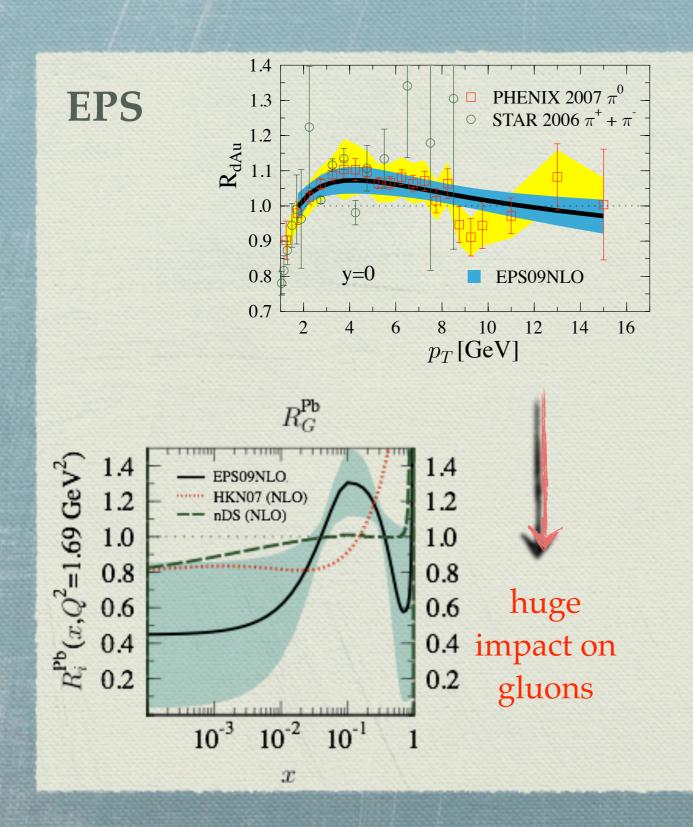
some tension between NuTeV and CDHSW (Fe) for specific x bins



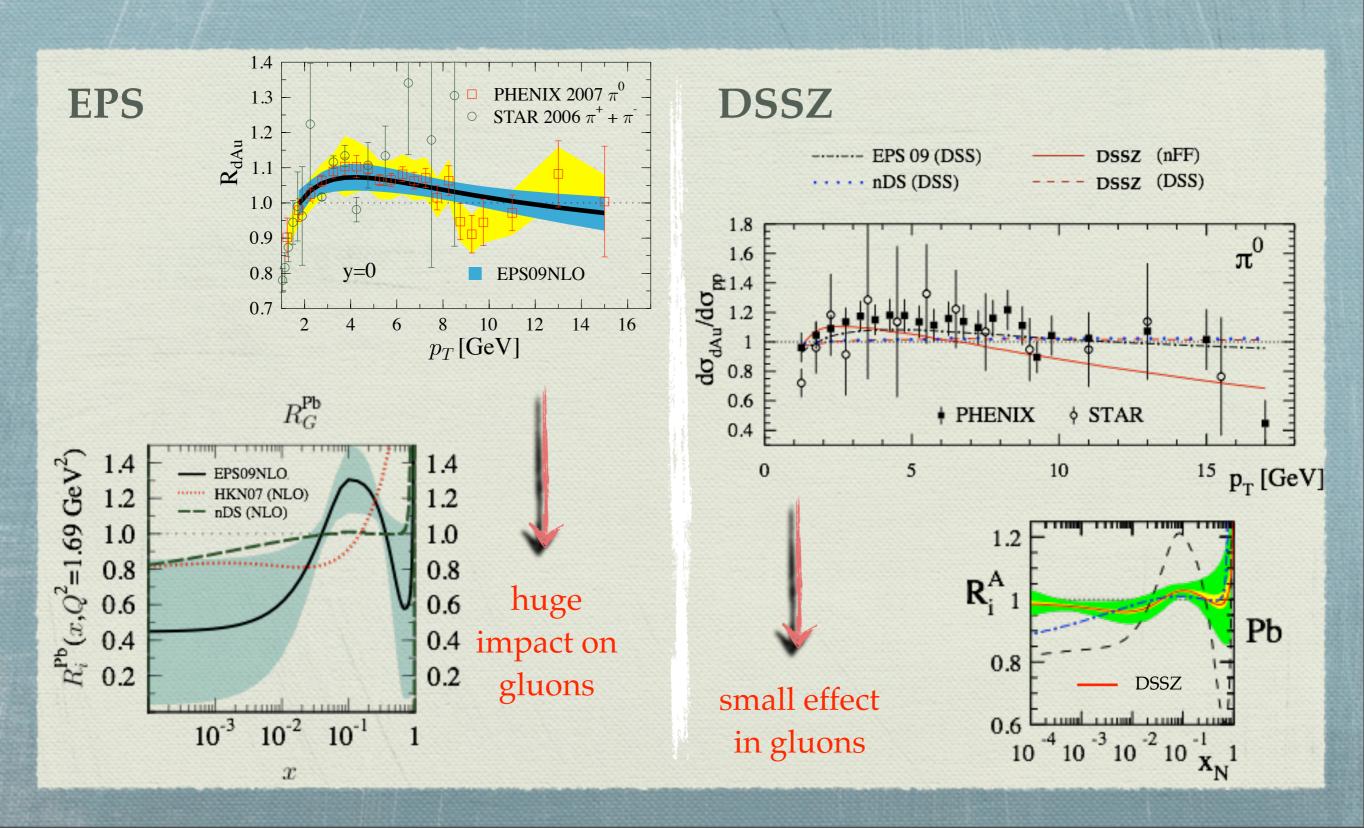




Both analysis include data from hadroproduction in dAu collisions (highly sensitive to gluon distribution)



Both analysis include data from hadroproduction in dAu collisions (highly sensitive to gluon distribution)



How do these different choices compare?

convolution (nDS)	multiplicative factor (EPS)	<u>from scratch</u> (nCTEQ)
first NLO analysis	first dAu	first neutrino DIS
DIS & DY	DIS & DY	DIS & DY
	RHIC dAu pion production	neutrino DIS
convolutional ansatz in Mellin space	piecewise multiplicative ansatz	direct ansatz (recovering proton for A=1)
small effect on gluons	huge anti-shadowing/ EMC effect for gluon	tension electron DIS/ neutrino DIS
factorization and universality ✓	factorization and universality ✓	factorization and universality?