

## Outline

- A landmark study: Explore the phase diagram of strongly interacting matter phase it's possible to go  $\mu$ nasc pmatter 37  $\delta$  and the transition
- RHIC capabilities for the study:  $\mathbf{v}$  .
	- Accelerator status
	- Experimental status, planned upgrades  $P_{\rm s}$
- What do we know? What to look for? Physics observables in the energy scan  $\alpha$ k t $\alpha$ r $\prime$  $\sum_{i=1}^{n}$ droplets in the cloud of plasma, he cloud of plasma, he cloud of plans in the cloud of plans in the cloud of p  $\mathcal{T}_{\mathbf{v}}$



w.sciencemag.org SCIENCE VOL 312 14 APRIL 2006 1914

the search in 2012.

when the transition from bound  $\alpha$ quarks becomes violent like boiling, says

3 years, he says. Meanwhile, researchers at Germany's GSI laboratory in Darmstadt are developing an accelerator that could also take up

Still, Gazdzicki and others agree that with its wide energy range and two large particle detectors, RHIC isideally suited to hunt for the critical point. The best opportunity should come after researchers upgrade the STAR detector in 2009. In the meantime, interest in the critical point is building like pressure in a boiling tea kettle.

## Motivation

Explore the phase diagram of strongly interacting matter

- Theoretical constraints:
	- $T = 0$  First order phase transition
	- $\mu_B = 0$ , T = T<sub>C</sub> Crossover
	- Critical Point at finite μ<sub>B</sub>
- Experimental constraints
	- Signatures for deconfinement at RHIC and top SPS
	- Disappearance of these signatures at SPS energies?



## Introduction RHIC @ BNL

- Has already explored an important part of the phase diagram:
	- Signatures for deconfinement
	- Studying properties of deconfined matter
- Energy/System size scan up to now





Critical point and 1st order phase transition line are *the* landmarks are they accessible at RHIC?



#### Introduction First Ideas for RHIC-Low Energy <u>RHIC Beam Region</u>

Tim Schuster RHIC Energy Scan VI-SIM Bad Liebenzell 12.-14. 09. 2007

- Fixed target RHIC program  $10 < E_{\text{Beam}} < 100 \text{ AGeV}$
- Cross-check the structures seen in hadron production excitation functions
- Use BRAHMS, NA49 detector?

100 cm

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Forward Spectrometer (FS)

**BRAHMS Experimental Setup** 

**Mid Rapidity Spectrometer** 

 $C<sub>4</sub>$ 

TOFW-

TPM<sub>2</sub>

 $DS$ TPM1

SiMA

& TMA



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## Accelerator Fixed Target vs Collider Mode

Advantages of collider mode over fixed target:

- Acceptance stays constant with energy
- Spatial track density rises slower



# Accelerator Collider Mode

Potential collider mode drawbacks:

## Trigger

Zero Degree Calorimeters not usable at low energy

 $\rightarrow$  Beam Beam Counters receive sufficient hits to be used  $\checkmark$ 

## Rate

Injection energy from AGS: 9.8 GeV/u per beam  $(\sqrt{s_{NN}} = 19.6 \text{ GeV})$ 

 $\gamma^2$  scaling of luminosity at higher energies.

Scaling for energies below normal injection energy unknown



## Accelerator Status

First test runs below standard injection energy:

- June 2006:  $\sqrt{s} = 22.5$  GeV p+p
- July 2007:  $\sqrt{s_{NN}} = 9.2$  GeV Au+Au

Au+Au collisions  $\overline{\omega}$   $\sqrt{s_{NN}}$  = 9.2 GeV seen in the STAR detector on June 7, 2007:



## Accelerator Status

First test run below design injection energy exceeded optimistic estimate for low energy luminosity



Planned energy scan: Au+Au at 7 energies (NA49 + 2)

 $=$  3 months of run X

• Test run at  $\sqrt{s_{NN}}$  = 5 GeV (at the end of run VIII) will show the scaling for lower energies

• Electron cooling in AGS (RHIC) would increase luminosity by another factor of 10 (100)

# Accelerator Conclusion



- RHIC is capable to extend existing program into the low energy region in collider mode
- Theoretical predictions see critical point in energy range  $5 < \sqrt{s_{NN}} < 20 \text{ GeV}$
- RHIC gives access to the whole range with sufficient statistics

## Experiments STAR and PHENIX



- Two commissioned, proven detectors: STAR and PHENIX ...with forming low energy working groups
- Large acceptance:  $2\pi$  (STAR) and wide  $p_T$  range for PID

Tim Schuster **RHIC Energy Scan** VI-SIM Bad Liebenzell 12.-14. 09. 2007

# Physics Observables • Spectra and yields • Fluctuations  $K/\pi$  $\langle p_T \rangle$ • Flow  $v_2$  scaling behavior  $Ω$  and φ  $v_2$ Disappearance of proton  $v_2$  at 40 GeV? • HBT • Heavy flavor mesons, di-leptons

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# Physics Observables

- Spectra and yields
- **Fluctuations** 
	- $K/\pi$
	- $\langle p_T \rangle$
- Flow
	- $v_2$  scaling behaviour
	- $Ω$  and φ  $v_2$
	- Disappearance of proton  $v_2$  at 40 GeV?
- HBT
- Heavy Flavor Mesons, Di-Leptons

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## Observables Spectra and Yields



Non-monotonous structures in SPS energy region



Interpretation under discussion - Systematic re-measurement of all strange hadrons for  $5 < \sqrt{s_{NN}} < 200$  GeV will hopefully solve the issue

Current SPS data seen as "*suggestive but inconclusive*" (Paul Sorensen)

Hyperons at AGS energies: FAIR? NICA?



## $Observables$   $K/\pi$  Fluctuations



- Positive dynamical fluctuations, rising fast towards low SPS energies
- Only excitation function of a fluctuation observable not being described by hadronic model **n**ction of a fluc

**STAR Preliminary**

## Observables  $K/\pi$  Separation: STAR TOF

### Measurement is tricky at RHIC:

- Kaon efficiency in collider lower (due to decay,  $c\tau = 3.7$ m)
- $K/\pi$  separation by  $dE/dx$  in TPC ambiguous above  $p = 0.5$  GeV/ $c$
- Misidentification has large impact on flucuations
- STAR TOF will enhance unambiguous kaon sample - full barrel TOF completed in 2009





## Observables Flow



Probe the early stage of the collision:

Test for initial pressure and degrees of freedom



## Observables Flow



## Observables Flow **0**  $\sim$  and from cumulants for four-particle correlations (v2{4}). The smooth solid lines are from Blast Wave model fits.



Proton *v*<sub>2</sub> collapse as signal for deconfinement

- Difference between methods: Depends on  $v_2$  fluctuations and non-flow contributions **-0.05**
- Azimuthally symmetric detector STAR can measure event-by-event flow vector **-0.1**
- STAR event plane resolution makes measurement with smaller error possible collisions, while we would have expected a significantly smaller value following the observations at 158A GeV. **0.1** - Azimuthally symmetric detector STAR<br>- STAR event plane resolution makes m<br>- Event plane detector as upgrade under<br>Tim Schuster RHIC Energy Scan
	- Event plane detector as upgrade under discussion **0.05**

# Conclusion Comparison of Programs

Large worldwide efforts to scan the phase diagram:

- The RHIC energy scan will provide a systematic study over a wide energy range (total covered range:  $5 < \sqrt{s_{NN}} < 200$  GeV) with large acceptance independent of energy
- CBM at FAIR will add the measurement of rare probes at lower energies
- The program at SPS adds the complementary system size scan and a larger rapidity coverage