

# Forward Neutral Pion Cross Section and Spin Asymmetry Measurements at STAR

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STAR Collaboration

Argonne National Laboratory

XI International Workshop on  
Deep-Inelastic Scattering and Related Subjects

Marseille, France

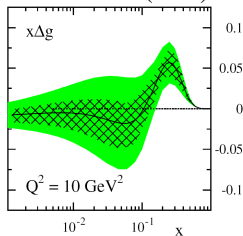
April 23<sup>rd</sup>, 2013



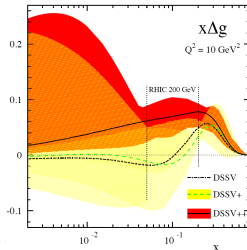
# Motivation: $\Delta g(x)$ and pQCD

- ▶ The gluon helicity distribution is one contributor to the total nucleon spin.
- ▶ While initially measured via (SI)DIS, measuring  $A_{LL}$  in polarized  $pp \rightarrow \pi^0 + X$ 
  - ▶ Provides complimentary access, both kinematically and in relation to partonic sub-processes
  - ▶ Has significant effects on global  $\Delta g(x)$  fits
- ▶ The global fits of  $\Delta g(x)$  are poorly constrained at  $x < 0.05$ .
- ▶ How to reach  $\Delta g(x)$  at lower  $x$ ?
  - ▶ Measure  $A_{LL}$  farther forward ( $\eta$  in 1-2), i.e. the STAR endcap electromagnetic calorimeter (EMC)
  - ▶ Main subprocess is now  $gq$  scattering, with small  $x$  of the gluon.
- ▶ First step: measure  $pp \rightarrow \pi^0 + X$  cross section and compare with pQCD.
- ▶  $\pi^0$  mesons are also a background to the prompt photon + jet, another channel to access  $\Delta g(x)$

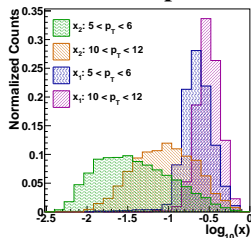
DSSV, PRD 80 (2009),  
PRL 101 (2008)



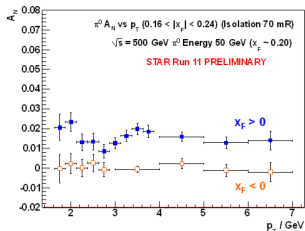
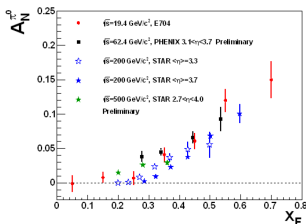
DSSV+/DSSV++,  
arXiv:1304.0079



$\pi^0$ s in STAR  
EMC Acceptance



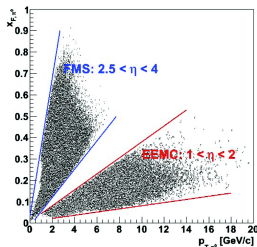
# Transverse Spin Asymmetry $A_N$



- ▶ Non-zero  $A_N$  for  $x_F > 0$  observed over wide energy range.
- ▶ Includes contributions from leading twist Sivers and Collins effects or higher twist effects.

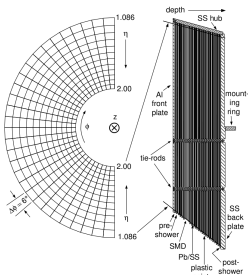
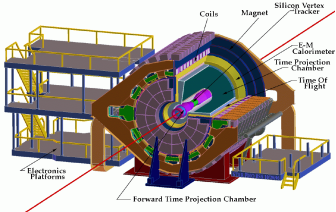
- ▶  $A_N$  expected to
  - ▶ Increase with  $x_F$
  - ▶ Go to zero in the limit of  $p_T \rightarrow 0$
  - ▶ Scale as  $1/p_T$  at high  $p_T$ .
- ▶ STAR EEMC has larger dynamic range in  $p_T$  and covers an unmeasured ( $p_T, x_F$ ) region
  - ▶  $5 < p_T < 12$  GeV,  $0.06 < x_F < 0.27$ , and  $0.8 < \eta < 2.0$
- ▶  $A_N$  will be small due to small  $x_F$ , but may show  $p_T$  dependence.

## STAR $\sqrt{s} = 200$ GeV

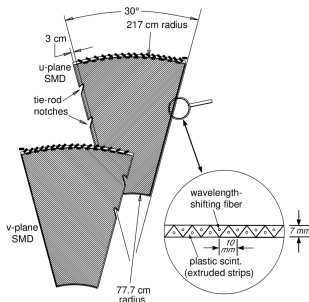


# STAR's Endcap Electromagnetic Calorimeter

STAR Detector



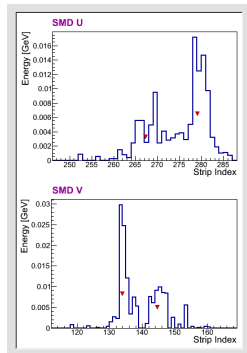
- ▶ Scintillating strip SMD
  - ▶  $\phi$  segmented into 12 sectors
  - ▶ Two active planes
  - ▶ 288 strips per plane
- ▶ Full  $\phi$  coverage—no gaps
- ▶ Resolution of a few mm



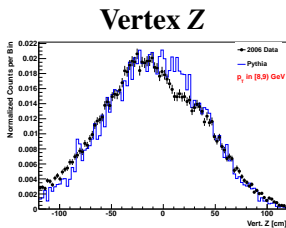
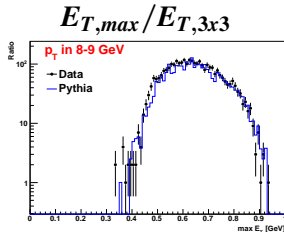
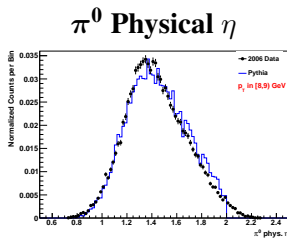
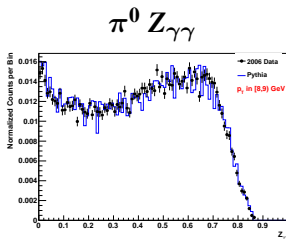
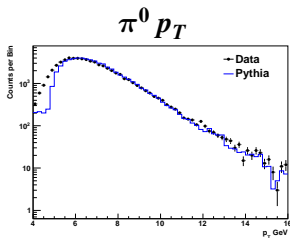
- ▶ Nucl. Instrum. Meth. A 499 (2003) 740.
- ▶ Lead/scintillator sampling EM calorimeter
- ▶ Covers  $1.09 < \eta < 2$  over full azimuth
- ▶ 720 optically isolated projective towers ( $\approx 22X_0$ )
- ▶ 2 pre-shower, 1 post-shower layers, and an additional shower max. detector (SMD)
- ▶ Trigger involves thresholds on the maximum tower energy and the  $3 \times 3$  patch of surrounding towers.

# Particle Reconstruction

- ▶ EM Particle ( $\gamma$ ,  $e^\pm$ , etc.) Reconstruction Procedure
  1. Identify clusters in the  $u$  and  $v$  strips
  2. Determine which  $u$  and  $v$  clusters to associate with incident particles
  3. Compute energy of incident particle using the towers.
- ▶ SMD clusters are found by
  - ▶ Smoothing the histogram using the method of J. Tukey (TH1::Smooth).
  - ▶ Identify clusters as a strip above an energy threshold, with  $\pm 3$  strips having monotonically decreasing energy.
  - ▶ Cluster position is set to energy-weighted mean position
- ▶ We expect cluster to be larger than  $1 \pm 3 = 7$  strips, but
  - ▶ Expect central strip position & energy to be sufficiently correlated to cluster position & energy.
  - ▶ Correlation increased by smoothing
- ▶ SMD response in fairly clean  $\pi^0$  candidate (data) event is plotted on the right.
  - ▶ Blue histograms show energy response per strip.
  - ▶ Inverted red triangles represent clusters, drawn at  $x=\text{mean}$ ,  $y=10\%$  cluster energy.
- ▶ General reconstruction difficulties include
  - ▶ Upstream material:  $\pi^0$  opening angle on the same order as opening angle for  $\gamma \rightarrow e^+e^-$
  - ▶ Single particle sometimes looks like two particles, and vice versa



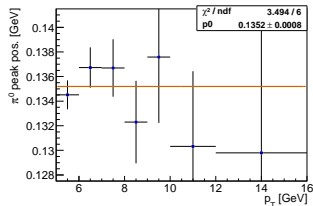
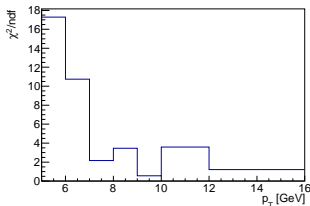
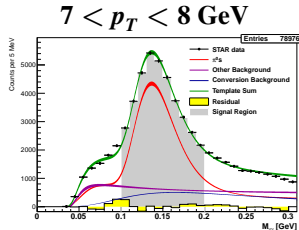
# Data/Monte Carlo Comparison



- ▶ Plots shown for  $\pi^0 p_T$  in 8-9 GeV
- ▶ Pythia tune 329, “Pro-pT0”
- ▶ Agreement generally good for  $\pi^0 p_T > 6$  GeV
- ▶  $5 < p_T < 6$  GeV usable, but has higher systematic uncertainty
- ▶ Sampled lumi. of  $8 \text{ pb}^{-1}$



# Background Subtraction



- ▶ There exist a variety of backgrounds, both due to physics and reconstruction; for example,
  - ▶  $\gamma \rightarrow e^+e^-$  conversions, and  $\pi^0$  candidate could be  $\gamma e^+$ ,  $\gamma e^-$ ,  $e^+e^-$ , etc.
  - ▶ Reconstructing the wrong number of photons in an event
- ▶ Sufficient to use three template functions to model signal + background
  - ▶  $\pi^0$  signal, direct conversion background, all other backgrounds
- ▶ Template function parameters fixed by fitting functions to reconstructed Pythia Monte Carlo.
- ▶ Normalizations of the templates and an energy scale factor determined by fitting template functions to the data

$$f_T(M_{\gamma\gamma}) = \sum_{i=1}^3 w_i f_i(M_{\gamma\gamma}/\alpha)$$



# Computing the Cross Section

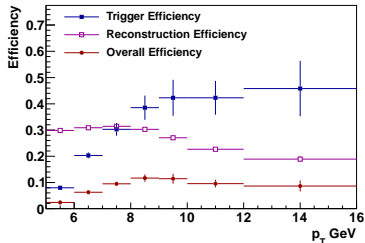
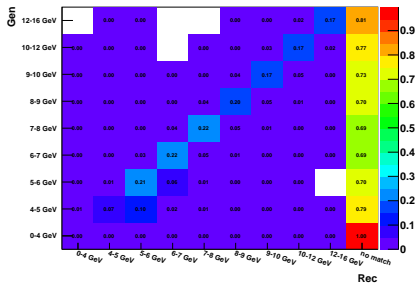
- ▶ The unfolded number of  $\pi^0$ s per  $p_T$  bin is computed as

$$N_i^{(\pi^0)} = \sum_j S_{i,j}^{-1} f_j s_j N_j^{(\text{raw})}$$

- ▶  $S$  is the smearing matrix
- ▶  $f$  accounts for  $\pi^0$ s smeared into the  $p_T$  range from outside
- ▶  $s$  is the signal fraction
- ▶  $N^{(\text{raw})}$  is the raw number of counts in the  $\pi^0$  peak window.
- ▶ The cross section is computed as

$$E \frac{d^3\sigma}{dp^3} = \frac{1}{2\pi} \frac{1}{\Delta\eta} \frac{1}{\Delta p_T} \frac{1}{\langle p_T \rangle} \frac{1}{\epsilon} \text{B.R.} \frac{N^{(\pi^0)}}{\mathcal{L}}$$

- ▶ Physical  $\eta$  is in (0.8, 2.0), thus  $\Delta\eta = 1.2$ .
- ▶ The  $p_T$  bin width,  $\Delta p_T$ , varies between 1 and 4 GeV.
- ▶ The total efficiency  $\epsilon$  is the product of the trigger and reconstruction efficiencies.
- ▶ The branching ratio for  $\pi^0 \rightarrow \gamma\gamma$  is 0.98823 (PDG)



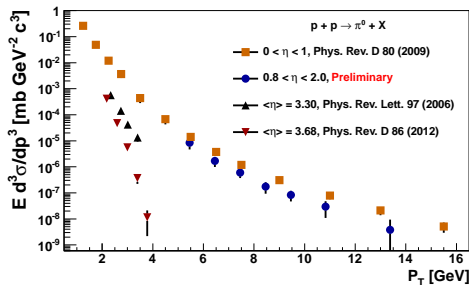
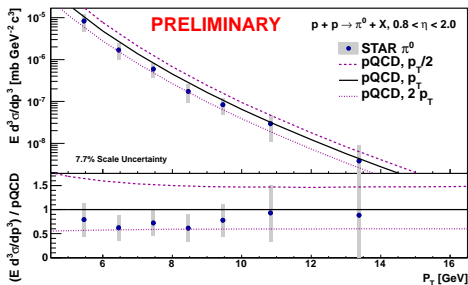


# Cross Section Uncertainties

- ▶ The statistical uncertainty is the Poisson uncertainty on the raw number of counts
- ▶ The following  $p_T$  dependent systematic uncertainties are included in the analysis
  - ▶ On the signal fraction
    - ▶ Uncertainty on template function parameters, energy scale and signal weight
    - ▶ Uncertainty on integrals to determine the fraction within the signal (peak) region
    - ▶ Uncertainty based on the fit residual, related to accuracy of template shapes
  - ▶ On the unfolded number of  $\pi^0$ s
    - ▶ Uncertainty on the smearing matrix  $S$  and factor  $f$  (related to Monte Carlo statistical uncertainty)
    - ▶ Uncertainty related to effectiveness of accounting for “smeared in” background
  - ▶ On the final cross section
    - ▶ Uncertainty on  $\langle p_T \rangle$ , assuming EEMC resolution is  $\delta E/E = 0.16/\sqrt{E}$
    - ▶ Uncertainties on reconstruction and trigger efficiencies (related to Monte Carlo statistical uncertainty)
    - ▶ Overall energy scale uncertainty of 3%—dominant systematic uncertainty
- ▶ All uncertainties are propagated analytically



# Cross Section Preliminary Results



- ▶ Theory curve from private communication with Marco Stratmann
  - ▶ Uses CTEQ65M distribution functions and DSS fragmentation function
  - ▶ Does not include propagated uncertainty on distribution and fragmentation functions
- ▶ Points plotted at Lafferty-Wyatt points, by fitting entire result to an exponential.
  - ▶ Error bars indicate statistical uncertainty (barely visible)
  - ▶ Error boxes indicate systematic uncertainty.
- ▶ Experimental uncertainties are on the order of the theoretical scale uncertainty.
- ▶ Result covers an unmeasured region of phase space
  - ▶ It is observed that the cross section for  $0.8 < \eta < 2.0$  is not much different than that for  $0 < \eta < 1$ .

# Analysis for Longitudinal Spin Asymmetries

- ▶ The raw asymmetries are computed from number of  $\pi^0$  candidates within  $0.1 < M_{\gamma\gamma} < 0.2$  GeV.

$$A_{L,B} = \frac{1}{\langle P_B \rangle} \left( \frac{N^{++} + N^{+-} - N^{-+} - N^{--}}{N^{++} + N^{+-} + N^{-+} + N^{--}} - \frac{L^{++} + L^{+-} - L^{-+} - L^{--}}{L^{++} + L^{+-} + L^{-+} + L^{--}} \right),$$

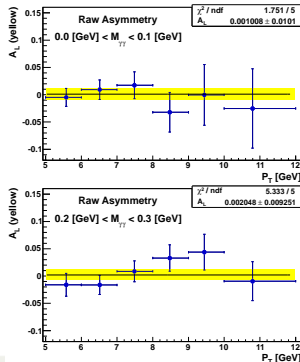
$$A_{L,Y} = \frac{1}{\langle P_Y \rangle} \left( \frac{N^{++} - N^{+-} + N^{-+} - N^{--}}{N^{++} + N^{+-} + N^{-+} + N^{--}} - \frac{L^{++} - L^{+-} + L^{-+} - L^{--}}{L^{++} + L^{+-} + L^{-+} + L^{--}} \right),$$

$$A_{LL} = \frac{1}{\langle P_B P_Y \rangle} \left( \frac{N^{++} - N^{+-} - N^{-+} + N^{--}}{N^{++} + N^{+-} + N^{-+} + N^{--}} - \frac{L^{++} - L^{+-} - L^{-+} + L^{--}}{L^{++} + L^{+-} + L^{-+} + L^{--}} \right),$$

- ▶ Luminosity weighted polarizations are  $\langle P_B \rangle = 0.56$ ,  $\langle P_Y \rangle = 0.59$ , and  $\langle P_B P_Y \rangle = 0.33$ .
- ▶ Asymmetries corrected for background using

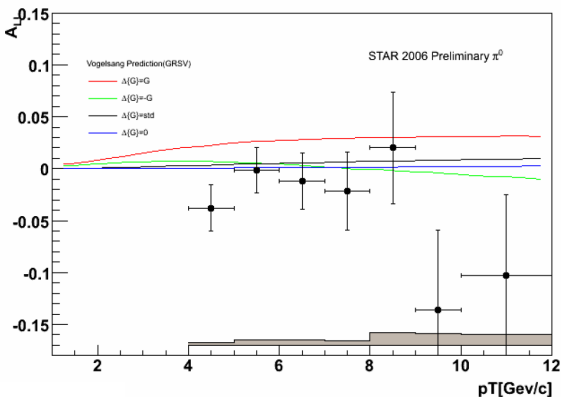
$$A^{sig} = \frac{1}{s} \left( A^{raw} - (1 - s)A^{bkg} \right),$$

- ▶ Background asymmetries are estimated from mass sideband regions and are observed to be within  $1\sigma$  of zero with  $\sigma \approx 0.01$ .



# Preliminary Results for Longitudinal Spin Asymmetries

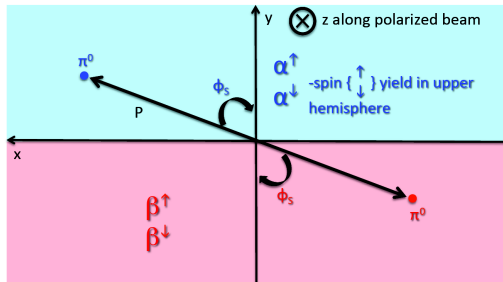
$\vec{p}+\vec{p}\rightarrow\pi^0X, \sqrt{s}=200\text{Gev}, 1.0\leq\eta\leq 2.0$



- ▶ Preliminary results shown at SPIN'08 use older procedures
  - ▶ Significant time invested in improving simulations, reconstruction code, and background subtraction procedure.
- ▶ Main message unchanged
- ▶ Systematics well under control
- ▶ First measurement in this  $\eta$  range
- ▶ Consistent with predictions
- ▶ May have some impact due to lower Bjorken- $x$  coverage
- ▶ Additional statistics already recorded at STAR during other running years.
  - ▶ Need effort verifying and improving simulations for the other years.



# Transverse Spin Asymmetry $A_N$ Analysis



- ▶ Raw asymmetry

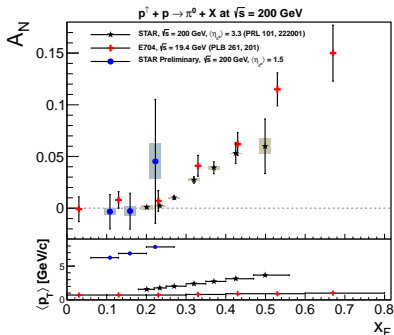
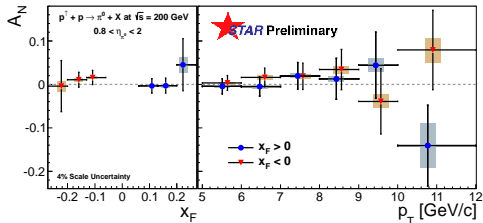
$$\mathcal{E}(\phi) = \frac{\sqrt{\alpha^\uparrow \beta^\downarrow} - \sqrt{\alpha^\downarrow \beta^\uparrow}}{\sqrt{\alpha^\uparrow \beta^\downarrow} + \sqrt{\alpha^\downarrow \beta^\uparrow}}$$

- ▶ Fit  $\mathcal{E}(\phi)$  to  $p_0 + \varepsilon \sin \phi$

- ▶ The background is subtracted from  $\varepsilon$  using same procedure as for longitudinal asymmetries.
  - ▶ Background asymmetries again within  $1\sigma$  of zero with  $\sigma \approx 0.01$ .
- ▶  $A_N$  obtained by scaling background subtracted  $\varepsilon$  by one over the luminosity weighted polarization.
- ▶ Systematics include
  - ▶ Uncertainty on the signal fraction
  - ▶ Uncertainty on the background asymmetry estimate
  - ▶  $\phi$ -dependent single beam backgrounds.



# Transverse Spin Asymmetry $A_N$ Preliminary Results



- ▶ Results consistent with zero for both  $x_F < 0$  and  $x_F > 0$
- ▶ Sensitivity not great enough to discern shape of  $p_T$  dependence
- ▶ Results lie in unmeasured region of  $(x_F, p_T)$  phase space
- ▶ Would like to include theory curves for comparison before publication



# Conclusions and Outlook

- ▶ Results represent measurements in previously unmeasured kinematic regions
  - ▶ First  $pp \rightarrow \pi^0 + X$  cross section within this  $\eta$  range
  - ▶  $A_{LL}$  reaches lower Bjorken- $x_2$  than published measurements
  - ▶  $A_N$  covers unmeasured  $(x_F, p_T)$  region
- ▶ Publication of these results is currently under review within the collaboration.
- ▶ These results demonstrate reconstruction with the EEMC is in a mature state.
- ▶ Thus far only 200 GeV data from one year analyzed
  - ▶ Several more years of data to analyze
  - ▶ More recent years have higher integrated luminosity and less upstream material
  - ▶ Data available for both  $\sqrt{s}$  at 200 GeV and 500 GeV.
  - ▶ Just need to finalize some details regarding the simulations.
- ▶ The STAR EEMC is also sensitive to other final states, such as prompt photons and  $\eta$ 's
- ▶ These results open the door for many STAR EEMC results to come.



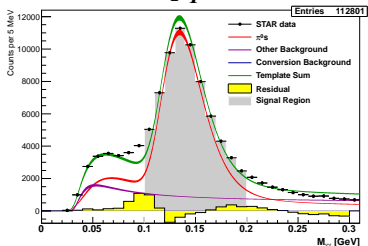
# Backup Slides



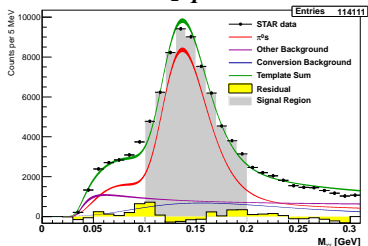


# Additional Template Fit Results

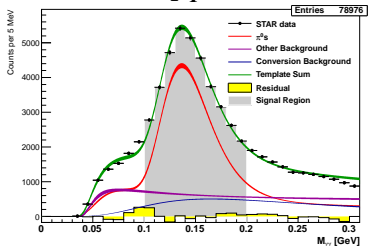
5 [GeV] <  $p_T$  < 6 [GeV]



6 [GeV] <  $p_T$  < 7 [GeV]



7 [GeV] <  $p_T$  < 8 [GeV]



8 [GeV] <  $p_T$  < 9 [GeV]

