

# Exclusive diffractive photon bremsstrahlung at high energies

Antoni Szczurek

Institute of Nuclear Physics (PAN), Cracow, Poland  
Rzeszów University, Rzeszów, Poland

DIS2013

Marseille, France, April 22-27, 2013



# Contents

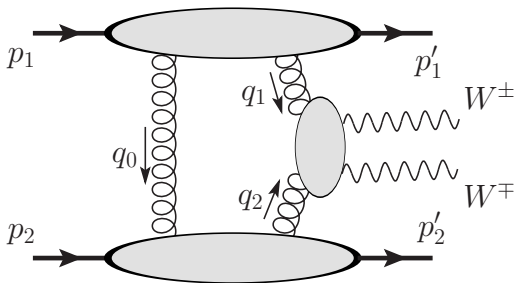
- Introduction
- Theoretical framework of *exclusive*  $pp \rightarrow ppy$ 
  - Classical bremsstrahlung
  - Vector meson rescattering
  - Pion cloud
  - Photon rescattering
- Results
  - Distributions for **photon** observables
  - Distributions for **proton** observables
- Conclusions

# Introduction

- Inclusive photon production at high energies in several Monte Carlo
- Exclusive **single** photon bremsstrahlung at high energy was practically not studied.
- Only recently **classical bremsstrahlung**, approximate formulas, point-like particles.
- Monte Carlo generators do not include it.
- What are relevant mechanisms ?
- Vector mesons bremsstrahlung (**Cisek, Lebedowicz, Schäfer, A.S.**)
- Differential distributions ?
- New **test of dynamics** of diffractive processes.
- Good test for and of **forward detectors**.



$pp \rightarrow pp\gamma\gamma$

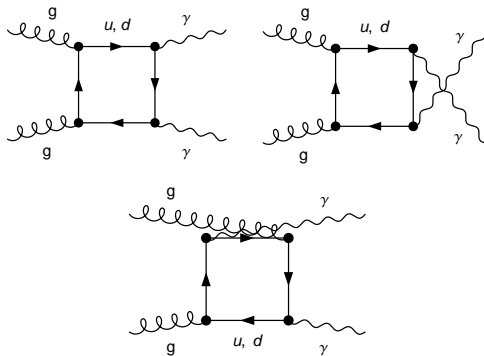


Lebiedowicz, Pasechnik, Szczurek

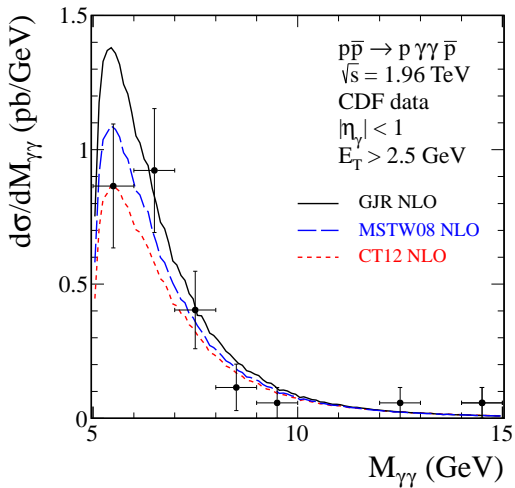
Nucl. Phys. **B867** (2013) 61.



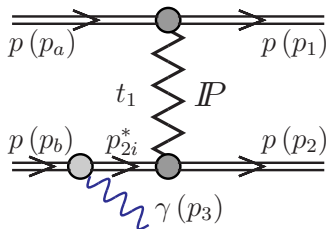
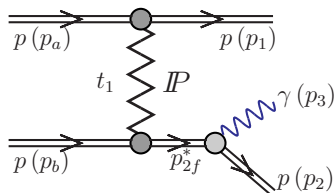
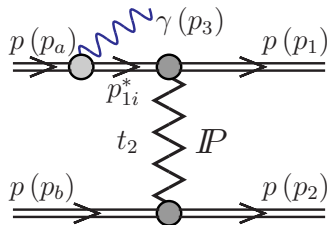
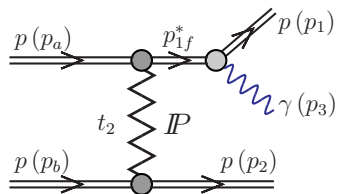
$pp \rightarrow pp\gamma\gamma$



$pp \rightarrow pp\gamma\gamma$



# Classical bremsstrahlung



# Classical bremsstrahlung

$$\begin{aligned}
 M_{\vec{n}_a \vec{n}_b \rightarrow \vec{n}_1 \vec{n}_2 \vec{n}_3}^{(a)} &= e \bar{u}(\rho_1, \vec{n}_1) \not{\epsilon}^*(\rho_3, \vec{n}_3) S_N(\rho_{1f}^2) \gamma^\mu u(\rho_a, \vec{n}_a) F_{\gamma N^* N}(\rho_{1f}^2) F_{PNN^*}(\rho_{1f}^2) \\
 &\times i s C_P^{NN} \left( \frac{s}{s_0} \right)^{\alpha_P(t_2)-1} \exp\left( \frac{B_P^{NN} t_2}{2} \right) \frac{1}{2s} \bar{u}(\rho_2, \vec{n}_2) \gamma_\mu u(\rho_b, \vec{n}_b), \quad (1)
 \end{aligned}$$

$$\begin{aligned}
 M_{\vec{n}_a \vec{n}_b \rightarrow \vec{n}_1 \vec{n}_2 \vec{n}_3}^{(b)} &= e \bar{u}(\rho_2, \vec{n}_2) \not{\epsilon}^*(\rho_3, \vec{n}_3) S_N(\rho_{2f}^2) \gamma^\mu u(\rho_b, \vec{n}_b) F_{\gamma N^* N}(\rho_{2f}^2) F_{PNN^*}(\rho_{2f}^2) \\
 &\times i s C_P^{NN} \left( \frac{s}{s_0} \right)^{\alpha_P(t_1)-1} \exp\left( \frac{B_P^{NN} t_1}{2} \right) \frac{1}{2s} \bar{u}(\rho_1, \vec{n}_1) \gamma_\mu u(\rho_a, \vec{n}_a), \quad (2)
 \end{aligned}$$

$$\begin{aligned}
 M_{\vec{n}_a \vec{n}_b \rightarrow \vec{n}_1 \vec{n}_2 \vec{n}_3}^{(c)} &= e \bar{u}(\rho_1, \vec{n}_1) \gamma^\mu S_N(\rho_{1i}^2) \not{\epsilon}^*(\rho_3, \vec{n}_3) u(\rho_a, \vec{n}_a) F_{\gamma NN^*}(\rho_{1i}^2) F_{PN^* N}(\rho_{1i}^2) \\
 &\times i s_{12} C_P^{NN} \left( \frac{s_{12}}{s_0} \right)^{\alpha_P(t_2)-1} \exp\left( \frac{B_P^{NN} t_2}{2} \right) \frac{1}{2s_{12}} \bar{u}(\rho_2, \vec{n}_2) \gamma_\mu u(\rho_b, \vec{n}_b), \quad (3)
 \end{aligned}$$

$$\begin{aligned}
 M_{\vec{n}_a \vec{n}_b \rightarrow \vec{n}_1 \vec{n}_2 \vec{n}_3}^{(d)} &= e \bar{u}(\rho_2, \vec{n}_2) \gamma^\mu S_N(\rho_{2i}^2) \not{\epsilon}^*(\rho_3, \vec{n}_3) u(\rho_b, \vec{n}_b) F_{\gamma NN^*}(\rho_{2i}^2) F_{PN^* N}(\rho_{2i}^2) \\
 &\times i s_{12} C_P^{NN} \left( \frac{s_{12}}{s_0} \right)^{\alpha_P(t_1)-1} \exp\left( \frac{B_P^{NN} t_1}{2} \right) \frac{1}{2s_{12}} \bar{u}(\rho_1, \vec{n}_1) \gamma_\mu u(\rho_a, \vec{n}_a), \quad (4)
 \end{aligned}$$





# Classical bremsstrahlung

$$t_{1,2} = (p_{a,b} - p_{1,2})^2 = a_{1,2}^2,$$

$$p_{1i,2i}^{*2} = (p_{a,b} - p_\gamma)^2,$$

$$p_{1f,2f}^{*2} = (p_{1,2} + p_\gamma)^2$$

$$s_{ij} = (p_i + p_j)^2$$

$$a_p(t) = 1.0808 + 0.25 t.$$

$$B(s) = B_0 + 2a'_p \ln\left(\frac{s}{s_0}\right)$$

where we use the value  $s_0 = 1 \text{ GeV}^2$  and  $B_0 = 9 \text{ GeV}^{-2}$ .

off-shell nucleon

$$F(p^*) = \frac{\Lambda_N^4}{(p^{*,2} - m_N^2)^2 + \Lambda_N^4} \quad (5)$$

this form used at low-energy bremsstrahlung



# Classical bremsstrahlung

When the off-shell effects of the participating protons are neglected, hence the  $\gamma NN^*$  and  $\gamma N^* N$  vertices are parametrized by the on-shell proton e.m. form factors in terms of the Dirac and Pauli form factors  $F_1(p_\gamma^2)$  and  $F_2(p_\gamma^2)$  given by

$$e\gamma^\mu F_{\gamma NN} F_{PNN} \rightarrow e \left[ F_1(p_\gamma^2) \gamma^\mu - \frac{i}{2m_N} F_2(p_\gamma^2) \sigma^{\mu\nu} p_{\gamma\nu} \right] \quad (6)$$

where  $\sigma^{\mu\nu} = \frac{i}{2} [\gamma^\mu, \gamma^\nu] = \frac{i}{2} (\gamma^\mu \gamma^\nu - \gamma^\nu \gamma^\mu)$ .

$F_1(p_\gamma^2 = 0) = 1$  and  $F_2(p_\gamma^2 = 0) = \kappa_p = 1.79$ .

The propagators of the intermediate nucleons

$$S_N(p_{1f,2f}^{*2}) = \frac{i(p_{1f,2f}^* \gamma^\nu + m_N)}{p_{1f,2f}^{*2} - m_N^2}, \quad S_N(p_{1i,2i}^{*2}) = \frac{i(p_{1i,2i}^* \gamma^\nu + m_N)}{p_{1i,2i}^{*2} - m_N^2}. \quad (7)$$



# Classical bremsstrahlung

The polarization vectors of real photon are defined in the proton-proton center-of-mass frame

$$\varepsilon(p_\gamma, \pm 1) = \frac{1}{\sqrt{2}}(0, i \sin \varphi \mp \cos \vartheta \cos \varphi, -i \cos \varphi \mp \cos \vartheta \sin \varphi, \pm \sin \vartheta), \quad (8)$$

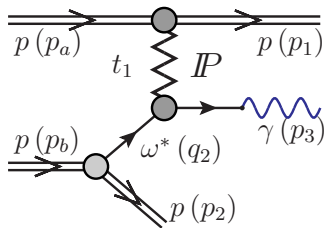
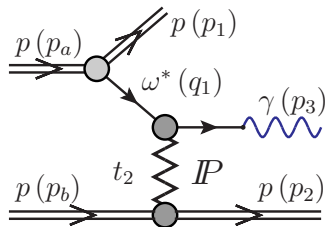
It is easy to check that:

$$\varepsilon^\mu(p, \hat{n}) \varepsilon_\mu^*(p, \hat{n}) = -1, \quad (9)$$

$$p^\mu \varepsilon_\mu(p, \hat{n}) = 0. \quad (10)$$



# Vector meson rescattering



Vector meson is off-mass shell.

Similar diagrams for  $\rho^0$  meson.

The diagrams for  $\omega$  and  $\rho^0$  interfere.



# Vector meson rescattering

$$\begin{aligned}
 \mathcal{M}_{\vec{\rho}_a \vec{\rho}_b \rightarrow \vec{\rho}_1 \vec{\rho}_2 \vec{\rho}_\gamma}^{(e)} &= \bar{u}(\rho_1, \vec{\rho}_1) \gamma^\mu u(\rho_a, \vec{\rho}_a) S_{\mu\nu}(t_1) \varepsilon^{\nu*}(\rho_\gamma, \vec{\rho}_\gamma) g_{\omega NN} F_{\omega^* NN}(t_1) F_{P\omega^* \omega}(t_1) C_{\omega \rightarrow \gamma} \\
 &\times i s_{23} C_P^{\omega N} \left( \frac{s_{23}}{s_0} \right)^{\alpha_P(t_2)-1} \left( \frac{s_{13}}{s_{th}} \right)^{\alpha_\omega(t_1)-1} \exp\left( \frac{B_P^{\omega N} t_2}{2} \right) \delta_{\vec{\rho}_2 \vec{\rho}_b}, \quad (11)
 \end{aligned}$$

$$\begin{aligned}
 \mathcal{M}_{\vec{\rho}_a \vec{\rho}_b \rightarrow \vec{\rho}_1 \vec{\rho}_2 \vec{\rho}_\gamma}^{(f)} &= \bar{u}(\rho_2, \vec{\rho}_2) \gamma^\mu u(\rho_b, \vec{\rho}_b) S_{\mu\nu}(t_2) \varepsilon^{\nu*}(\rho_\gamma, \vec{\rho}_\gamma) g_{\omega NN} F_{\omega^* NN}(t_2) F_{P\omega^* \omega}(t_2) C_{\omega \rightarrow \gamma} \\
 &\times i s_{13} C_P^{\omega N} \left( \frac{s_{13}}{s_0} \right)^{\alpha_P(t_1)-1} \left( \frac{s_{23}}{s_{th}} \right)^{\alpha_\omega(t_2)-1} \exp\left( \frac{B_P^{\omega N} t_1}{2} \right) \delta_{\vec{\rho}_1 \vec{\rho}_a}, \quad (12)
 \end{aligned}$$



# Vector meson rescattering

$$C_P^{\omega N} = C_P^{\pi N} = 13.63 \text{ mb (Donnachie-Landshoff)}$$

$$B_P^{\omega N} = B_P^{\pi N} = 5.5 \text{ GeV}^{-2}.$$

The amplitudes are corrected to reproduce the high-energy Regge dependence.

$$a_\omega(t) = 0.5 + 0.9 t \text{ and } s_{th} = (m_N + m_\omega)^2.$$

$g_{\omega NN}^2/4\pi = 10$  (different values have been used in the literature)

$$C_{\omega \rightarrow \gamma} = \sqrt{a_{em}/20.5}$$

within vector dominance model with finite width corrections

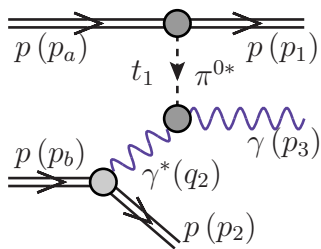
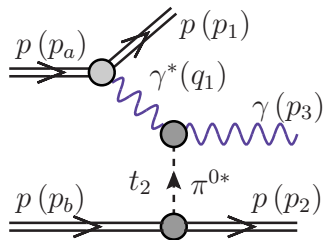
(Uleshchenko-Szczurek)

$$S_{\mu\nu}(t) = \frac{-g_{\mu\nu} + \frac{q_\mu q_\nu}{m_\omega^2}}{t - m_\omega^2}. \quad (13)$$

$$F_{\omega^* NN}(t_{1,2}) = \exp\left(\frac{t_{1,2} - m_\omega^2}{\Lambda_{\omega NN}^2}\right), F_{P\omega^* \omega}(t_{1,2}) = \exp\left(\frac{t_{1,2} - m_\omega^2}{\Lambda_{P\omega\omega}^2}\right), \quad (14)$$



## Pion cloud contribution



Anomalous coupling, but off-shell pion



# Pion cloud contribution

$$\begin{aligned}
 M_{\hat{n}_a \hat{n}_b \rightarrow \hat{n}_1 \hat{n}_2 \hat{n}_\gamma}^{\gamma \pi^0 \text{-exch.}} &= \bar{u}(\rho_1, \hat{n}_1) \Gamma_{\gamma^* NN}^a(q_1^2) u(\rho_a, \hat{n}_a) \\
 &\times \frac{-g_{\alpha\beta}}{t_1} F_{\gamma\pi \rightarrow \gamma}(t_1, t_2) \varepsilon^{\beta\mu\nu\hat{n}} q_{1\mu} p_{3\nu} \varepsilon_{\hat{n}}^*(\rho_\gamma, \hat{n}_\gamma) \\
 &\times g_{\pi^0 NN} F_{\pi NN}(t_2) \frac{1}{t_2 - m_\pi^2} \bar{u}(\rho_2, \hat{n}_2) i\gamma_5 u(\rho_b, \hat{n}_b), \tag{15}
 \end{aligned}$$

$$\begin{aligned}
 M_{\hat{n}_a \hat{n}_b \rightarrow \hat{n}_1 \hat{n}_2 \hat{n}_\gamma}^{\pi^0 \gamma \text{-exch.}} &= g_{\pi^0 NN} F_{\pi NN}(t_1) \frac{1}{t_1 - m_\pi^2} \bar{u}(\rho_1, \hat{n}_1) i\gamma_5 u(\rho_a, \hat{n}_a) \\
 &\times \frac{-g_{\alpha\beta}}{t_2} F_{\gamma\pi \rightarrow \gamma}(t_2, t_1) \varepsilon^{\beta\mu\nu\hat{n}} q_{2\mu} p_{3\nu} \varepsilon_{\hat{n}}^*(\rho_\gamma, \hat{n}_\gamma) \\
 &\times \bar{u}(\rho_2, \hat{n}_2) \Gamma_{\gamma^* NN}^a(q_2^2) u(\rho_b, \hat{n}_b), \tag{16}
 \end{aligned}$$





# Pion cloud

## Pion-to-nucleon form factor

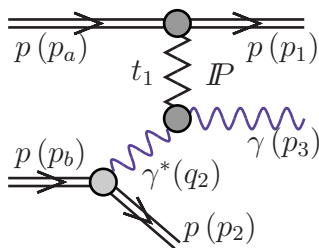
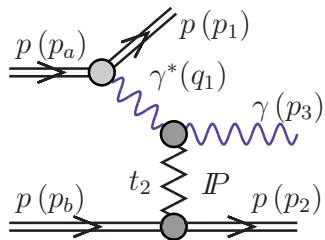
$$F_{\pi NN}(t_{1,2}) = \exp\left(\frac{t_{1,2} - m_\pi^2}{\Lambda_{\pi NN}^2}\right). \quad (17)$$

## On-shell normalization

$$F_{\gamma\pi\rightarrow\gamma}(t_1, t_2) = \frac{N_c}{12\pi^2 f_\pi} \frac{m_\rho^2}{m_\rho^2 - t_1} \exp\left(\frac{t_2 - m_\pi^2}{\Lambda_{\gamma\pi\rightarrow\gamma}^2}\right) \quad (18)$$



# Photon rescattering



photon-proton quasi-elastic scattering



# Photon rescattering

## Equivalent Photon Approximation

$$\begin{aligned} \frac{d\sigma}{dydp_{\perp}^2} &= z_1 f(z_1) \frac{d\sigma_{\gamma p \rightarrow \gamma p}}{dt_2} (W_{23}, t_2 \approx -p_{\perp}^2) \\ &+ z_2 f(z_2) \frac{d\sigma_{\gamma p \rightarrow \gamma p}}{dt_1} (W_{13}, t_1 \approx -p_{\perp}^2) . \end{aligned} \quad (19)$$

$$E = p_{\perp} \cosh y \quad , \quad p_z = \pm \sqrt{E^2 - p_{\perp}^2} . \quad (20)$$

$$z_1 s = s_{13} \quad , \quad z_2 s = s_{23} \quad (21)$$

$$\begin{aligned} \frac{d\sigma_{\gamma p \rightarrow \gamma p}}{dt_1} (W_{13}, t_1) &= \frac{[\sigma_{tot}^{\gamma p}(W_{13})]^2}{16\pi} \exp(B_{\gamma p}(W_{13})t_1) \\ \frac{d\sigma_{\gamma p \rightarrow \gamma p}}{dt_2} (W_{23}, t_2) &= \frac{[\sigma_{tot}^{\gamma p}(W_{23})]^2}{16\pi} \exp(B_{\gamma p}(W_{23})t_2) . \end{aligned} \quad (22)$$



# From the amplitude to cross section

Four-dimensional integration in:

$p_{1t}, p_{2t}, Y, \phi_{12}$

This is rather difficult !

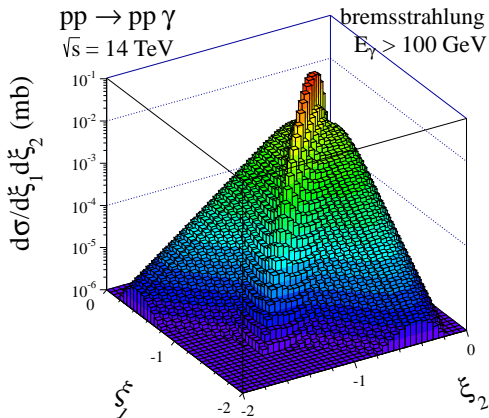
**Better:**

$$p_{1t} \rightarrow \xi_1 = \log_{10}(p_{1t}),$$

$$p_{2t} \rightarrow \xi_2 = \log_{10}(p_{2t})$$



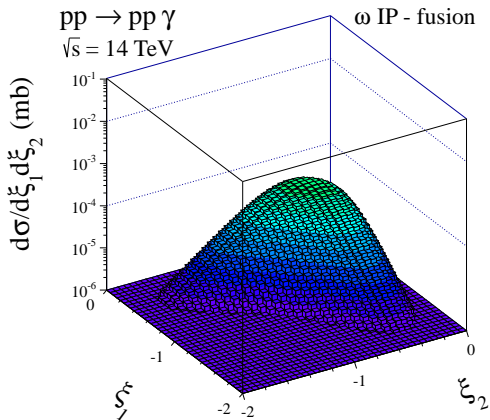
## Two-dimensional integration, bremsstrahlung



Enhancement on the diagonal (elastic scattering)



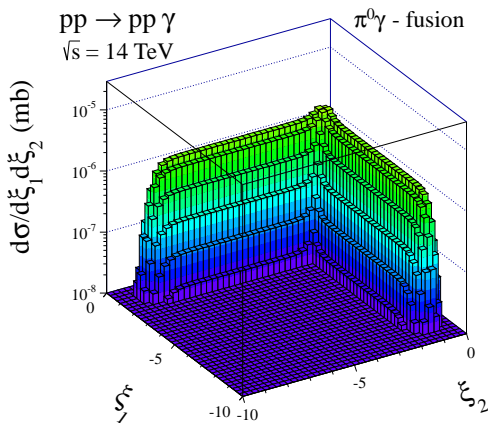
# Two-dimensional integration, omega rescattering



NN absorption should be included.



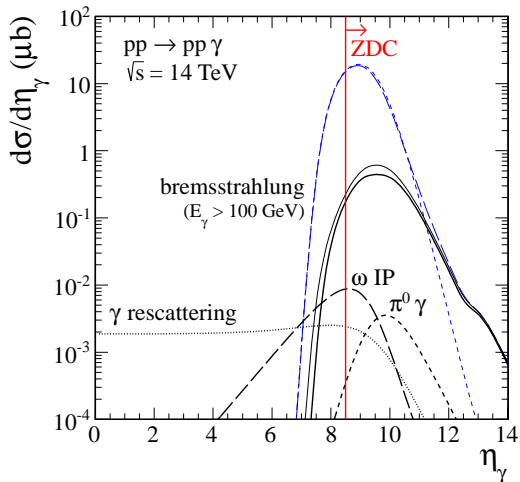
# Two-dimensional integration, pion cloud



Absorption probably small

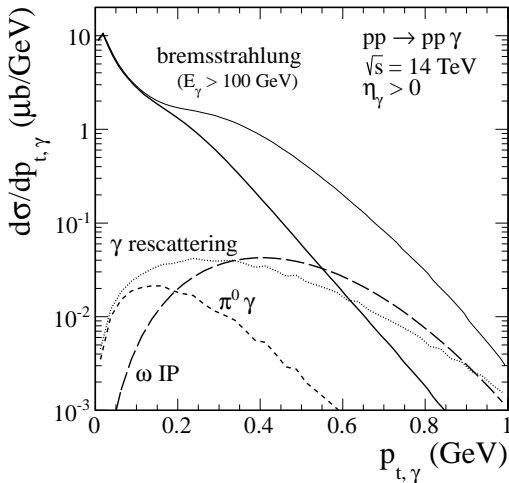


# Pseudorapidity distribution of photons

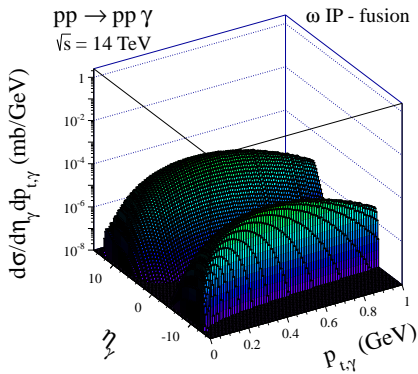
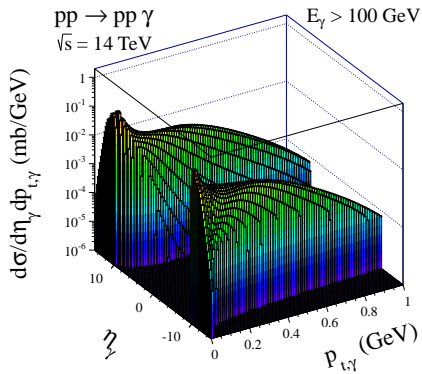




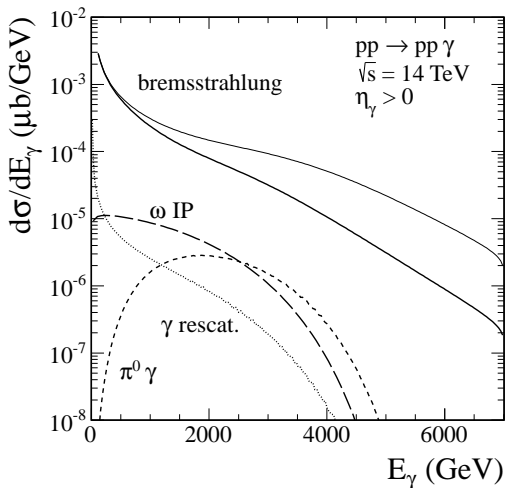
# Transverse momentum of photons



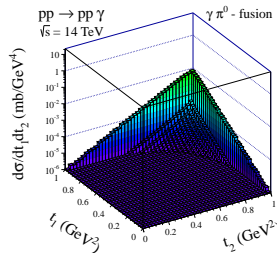
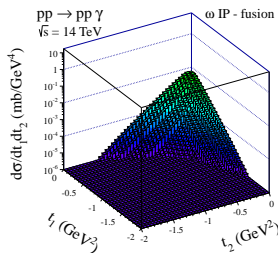
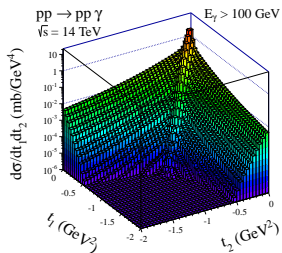
# Pseudorapidity-transverse momentum distributions



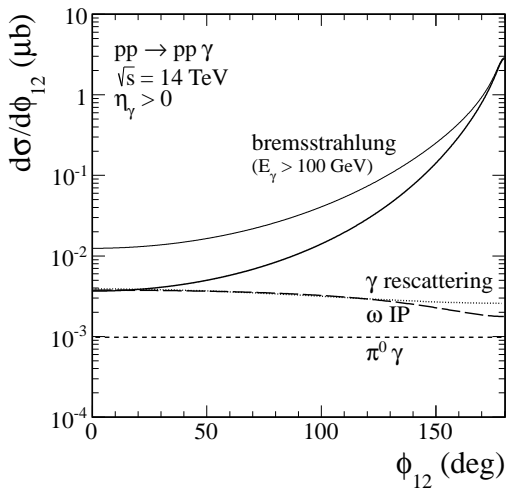
# Energy of photons



# Proton-proton correlations



# Proton-proton azimuthal correlations



reminiscence of elastic scattering



## Comment on absorption effects

- **classical bremsstrahlung**  
Includes effective nucleon-nucleon interaction.  
Including absorption would be (to large extend) a double-counting.
- **omega rescattering**  
effective  $\omega N$  interaction -- effectively includes higher orders in  $\omega N$  rescattering.  
**missing  $NN$  interaction**
- **pion cloud**  
probably small  
never both  $t_1$  and  $t_2$  are large.
- **photon rescattering**  
probably small  
Either  $p_{1t}(t_1)$  or  $p_{2t}(t_2)$  are small.



# Conclusions

- Exclusive photon production
- New mechanisms calculated for the first time.
- Classical bremsstrahlung gives the largest contribution.
- Photons are emitted very forward.  
Could be measured by ZDC detectors.
- Protons are emitted at small angles.  
Could be measured by ALPHA or TOTEM.
- Combination of both devices simultaneously would allow real measurement.
- Very small contribution at midrapidities.
- Next step → include absorption.

