

Electromagnetic and muonic cosmic ray shower profiles

Francisco Diogo
on behalf of the Auger Group at LIP



23rd April 2012

Extensive air showers

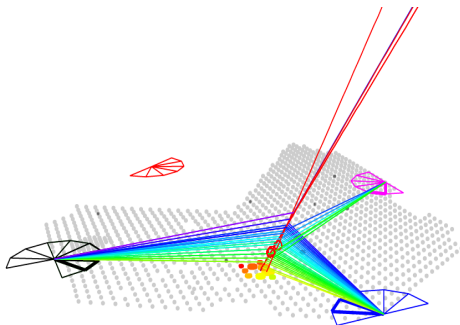
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and muonic
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Electromagnetic
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Profile
 X_{max}
distribution
interpretation

Muonic
Longitudinal
profiles

Measuring
the muonic
profile -
Muon
Production
Depth
(MPD)
method



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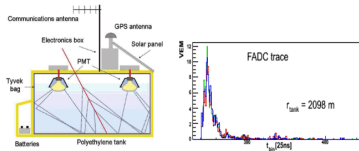
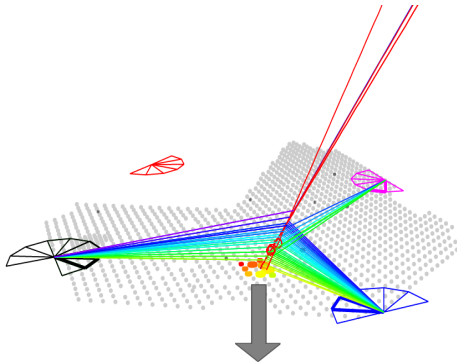
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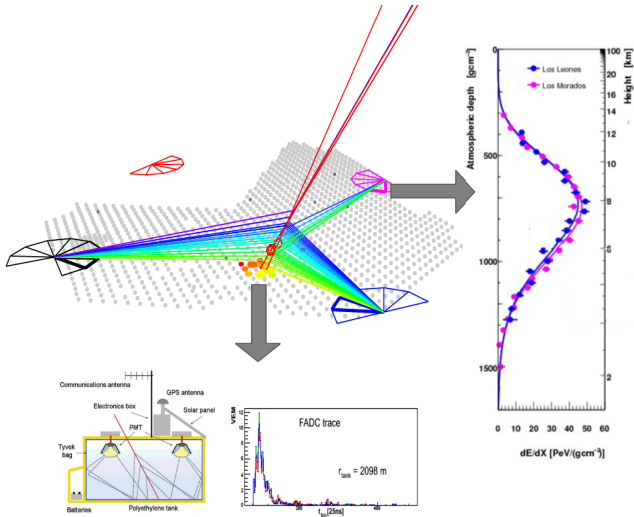
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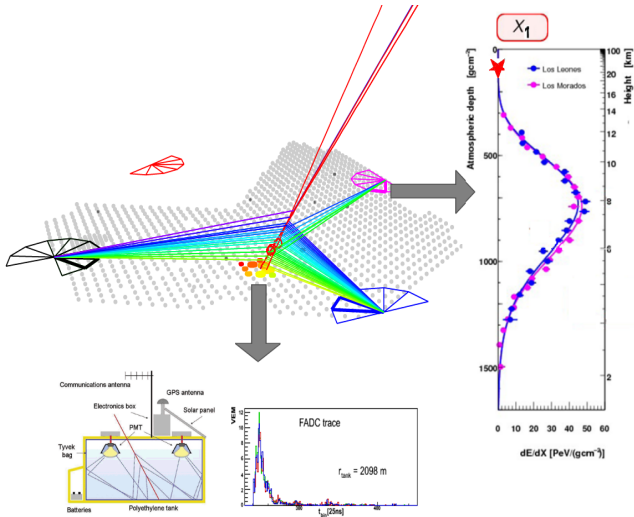
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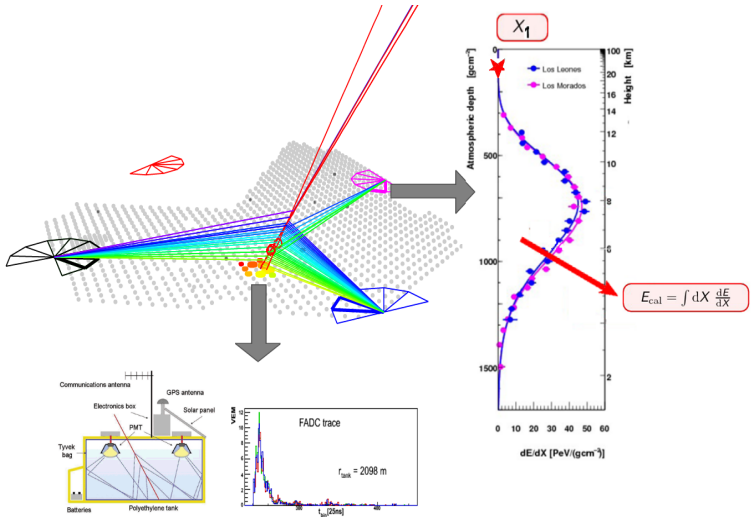
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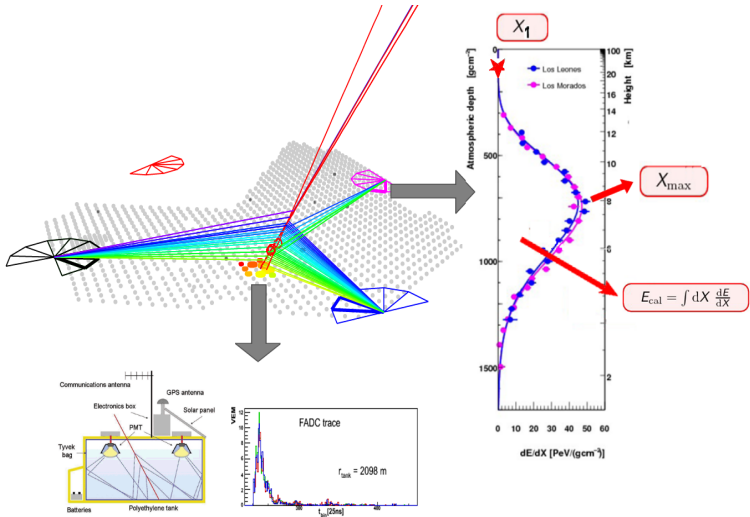
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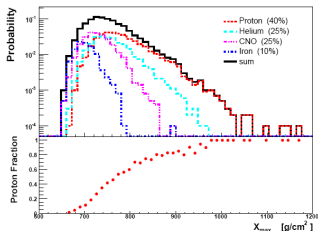
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The electromagnetic longitudinal profile is a very powerful analysis tool, and its main variable is X_{max} :

- sensitive to the mass of the primary particle

Electromagnetic Longitudinal Profile - X_{max}

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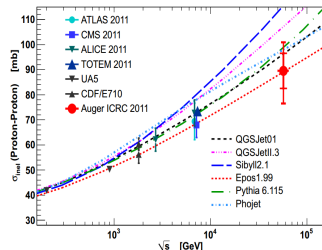
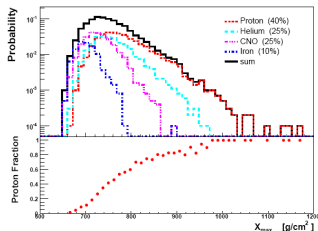
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The electromagnetic longitudinal profile is a very powerful analysis tool, and its main variable is X_{max} :

- sensitive to the mass of the primary particle
- the $X_{max} = X_1 + \Delta$ distribution tail is dominated by showers with large X_1 - proton-Air inelastic cross-section measurement
- Particle physics at energies greater than LHC

X_{max} and $RMS(X_{max})$ joint interpretation

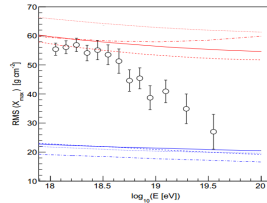
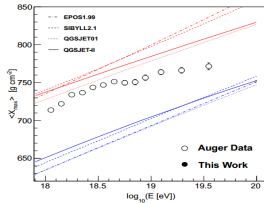
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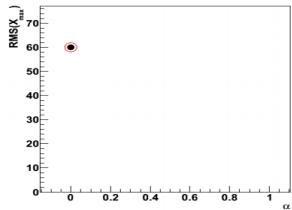
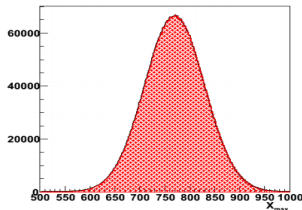
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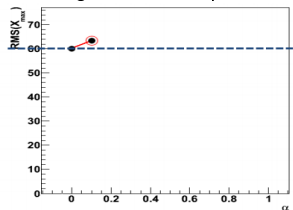
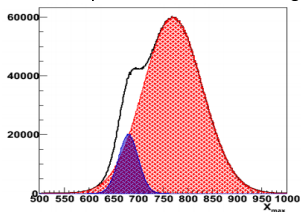
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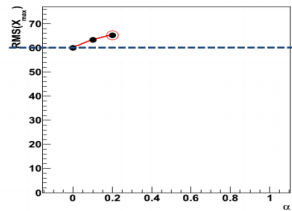
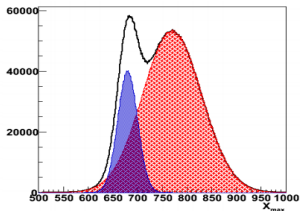
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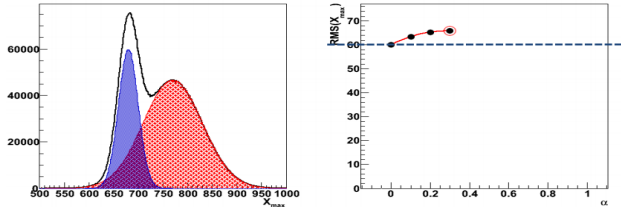
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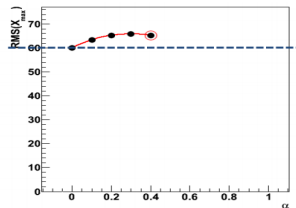
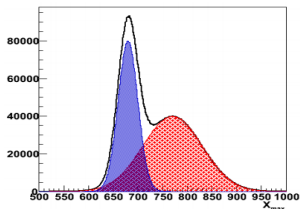
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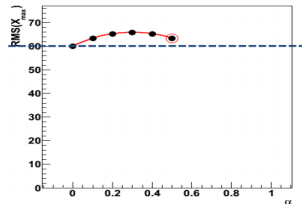
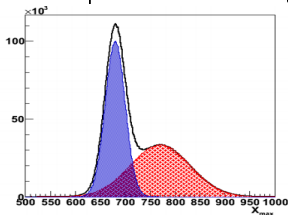
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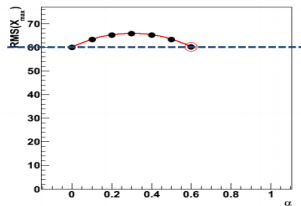
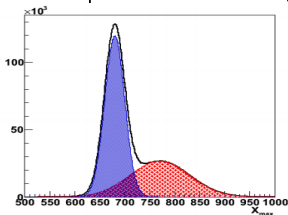
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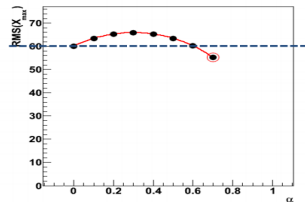
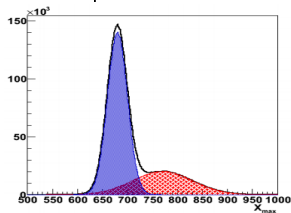
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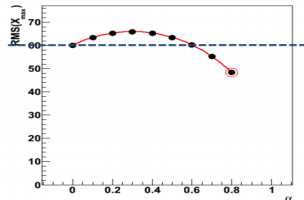
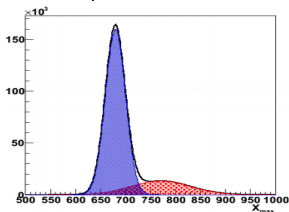
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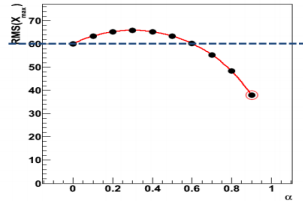
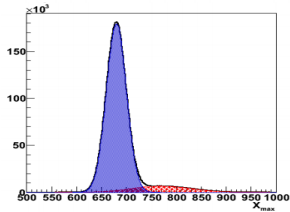
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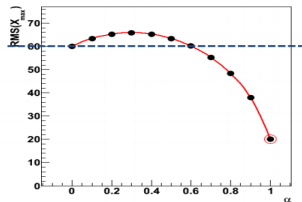
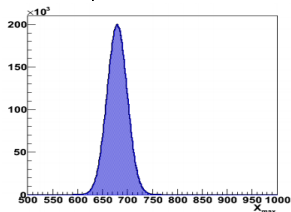
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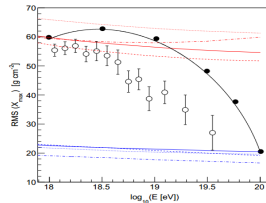
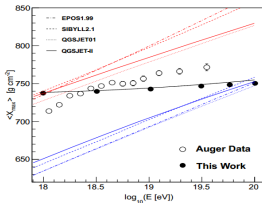
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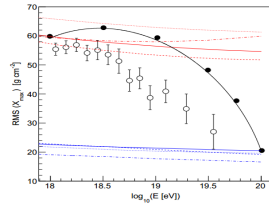
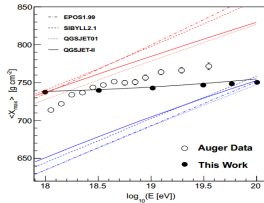
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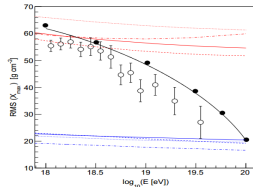
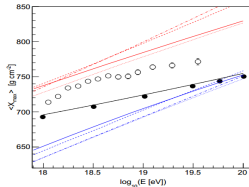
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Transition from 50% proton/50% iron



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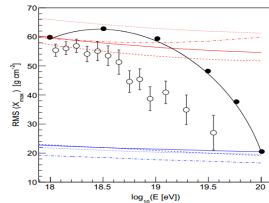
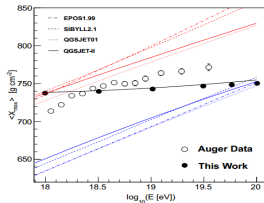
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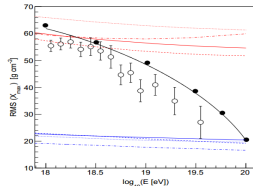
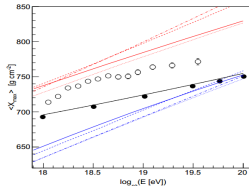
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Transition from 50% proton/50% iron



No simple scenario of transition from proton to iron explains data!

Grey Disk Model

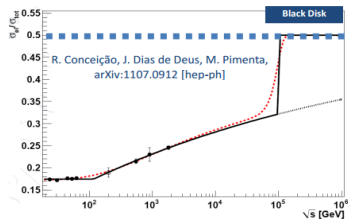
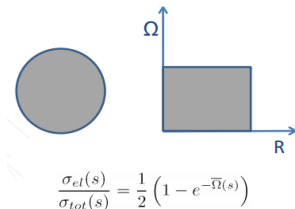
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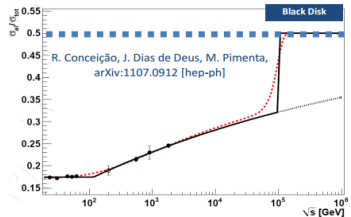
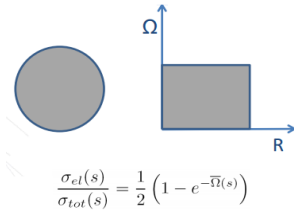
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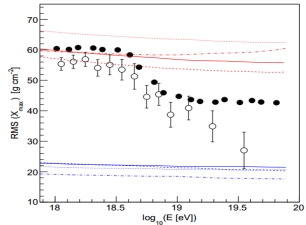
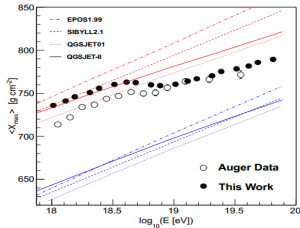
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Proton turning into a black disk at 10^{19} eV



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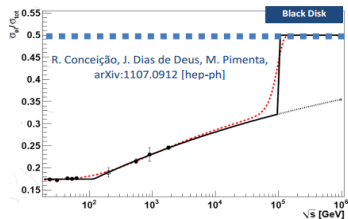
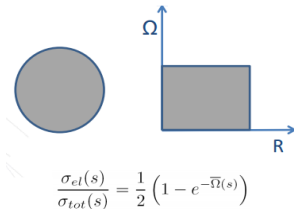
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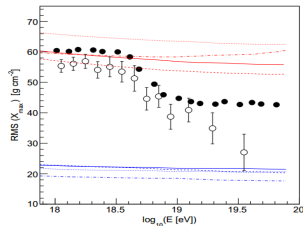
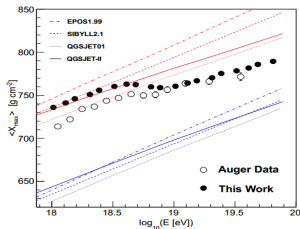
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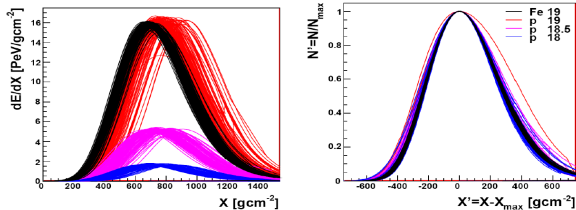


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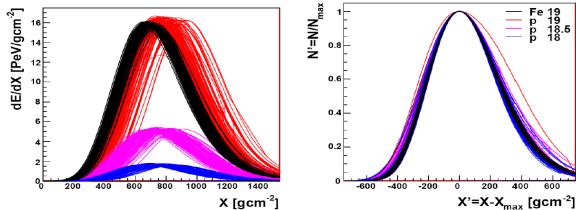
To interpret data we need a better understanding of the first interaction

Universal Shower Profile



Deconvoluting information from N_{max} and X_{max} the shape of the electromagnetic profile is (almost) universal

Universal Shower Profile



Deconvoluting information from N_{max} and X_{max} the shape of the electromagnetic profile is (almost) universal
In LIP, a new parameterization was introduced, in which the profile assymetry is described by a parameter R'

- is a measure of the shower development speed $\Delta = X_{\text{max}} - X_1$
- is independent of X_1
- is sensitive to composition of the primary particle

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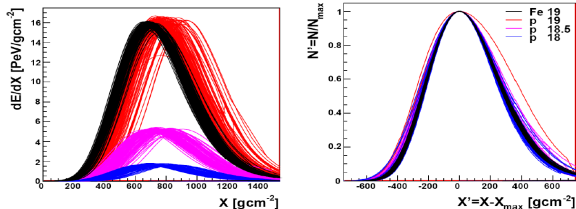
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Possible to measure $X_1 = X_{max} - \Delta(R')$

$$X_1(R')$$

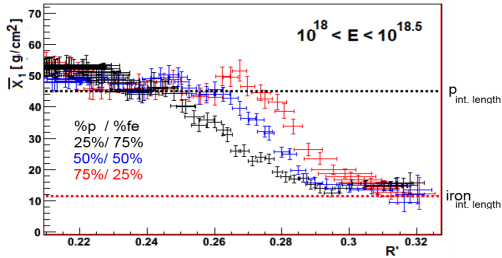
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Master Thesis

- $\overline{X_1}$ for mixed iron/proton samples
- Extreme bins in R' allow for a measure of both cross-sections
- Only done in simulations, still need to analyze data

EAS components

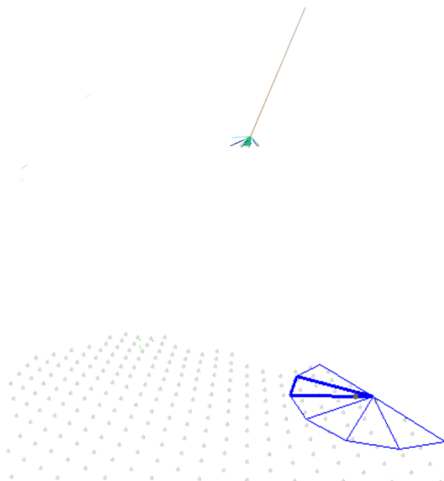
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EAS components

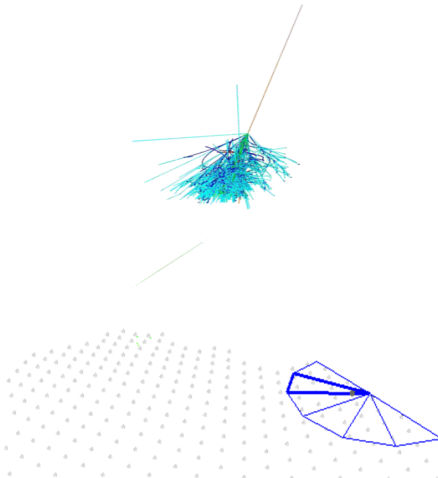
Electromagnetic
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Francisco
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on behalf of
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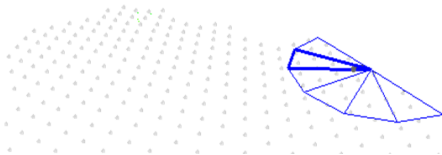
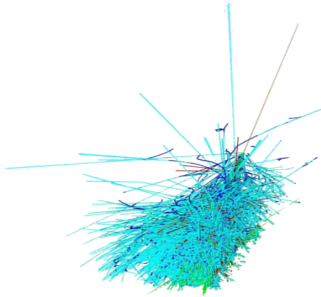
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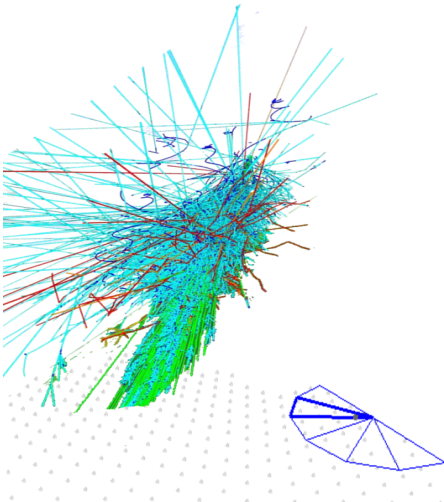
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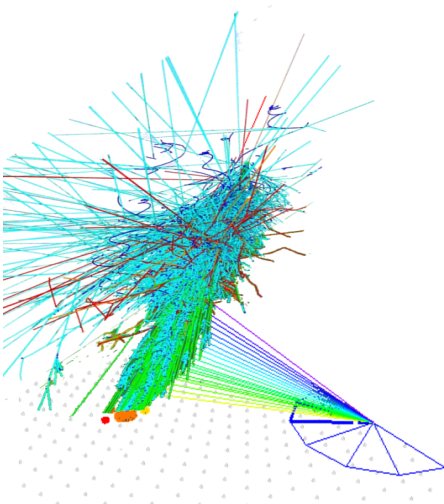
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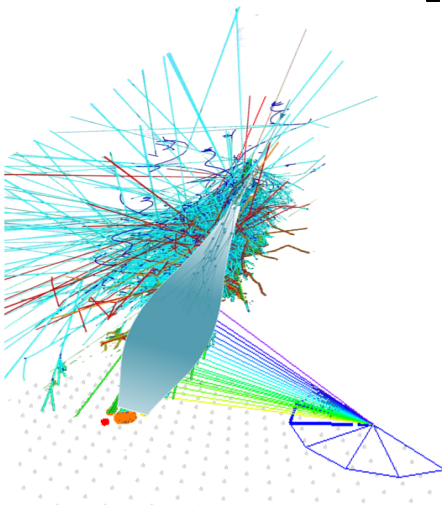
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Electromagnetic

- carries most energy (90%) of developed showers
- visible in both SD and FD
- easy to measure near maximum
- only visible in dark nights with clear sky
- very difficult to measure near the initial region
- first interaction information is 'dilluted'

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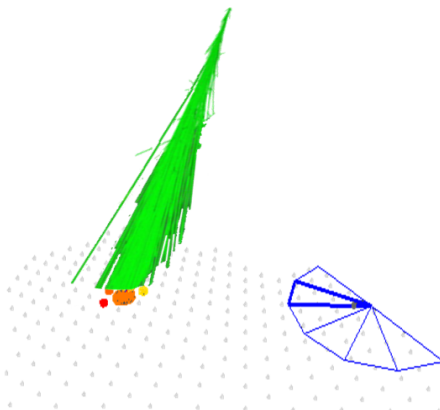
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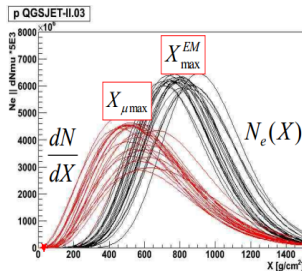
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Muonic

- hard to measure - no signal in FD, difficult to disentangle from e.m signal in SD
- muons interact scarcely with the atmosphere
- high energy muons preserve information from first interaction



The muonic longitudinal shower profiles at production



The number of muons produced in the shower as a function of depth can be described by the same function as the e.m. profile

The fast rise allows a more direct probing of the initial interactions

The muonic longitudinal shower profiles at production

Electromagnetic and muonic cosmic ray shower profiles

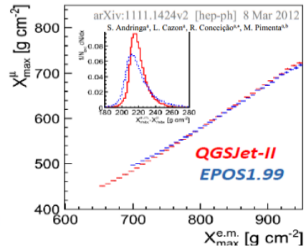
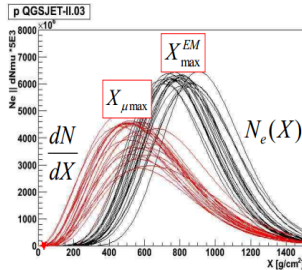
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X_{max} distribution interpretation

Muonic Longitudinal profiles

Measuring the muonic profile - Muon Production Depth (MPD) method



The number of muons produced in the shower as a function of depth can be described by the same function as the e.m. profile
 The fast rise allows a more direct probing of the initial interactions
 The profiles are heavily correlated: both X_{max} and the shape parameters of one profile can be estimated by knowing the other

- Inconsistencies give important information on hadronic models

Muon Production Depth (MPD)

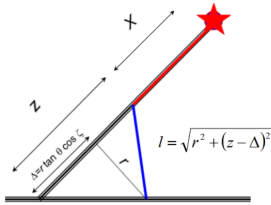
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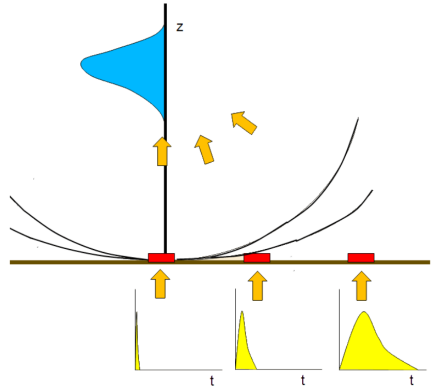
Electromagnetic Shower Profile
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Muonic Longitudinal profiles

Measuring the muonic profile - Muon Production Depth (MPD)



Muons travel in straight lines. Knowing their arrival time, we can get the depth of their production, z .



Measuring the muonic profile - Muon Production Depth (MPD) method

Electromagnetic and muonic cosmic ray shower profiles

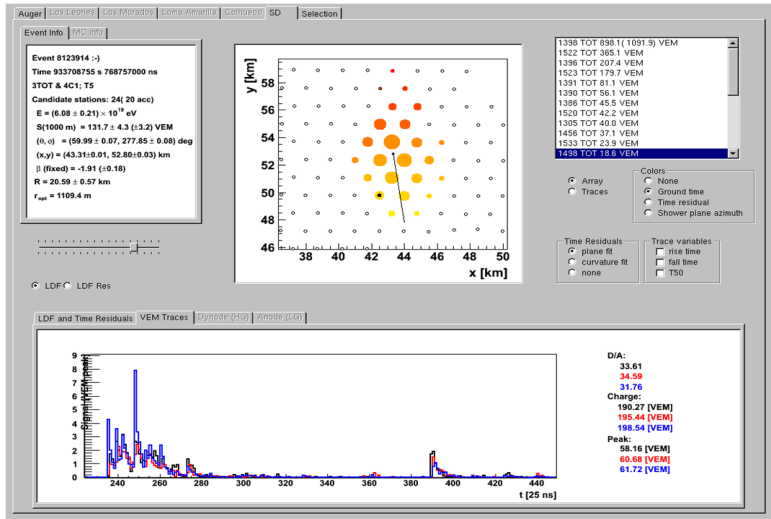
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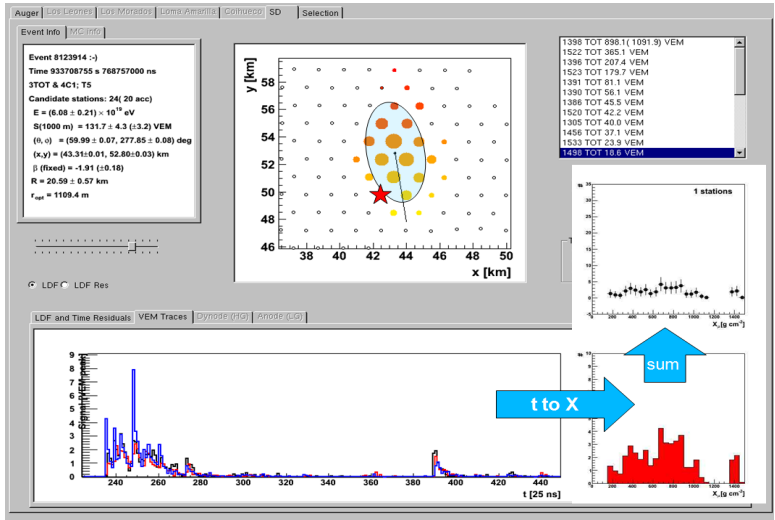
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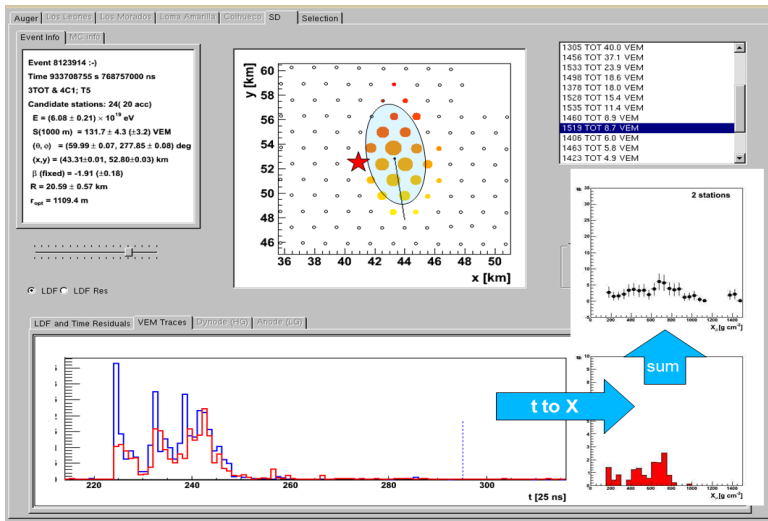
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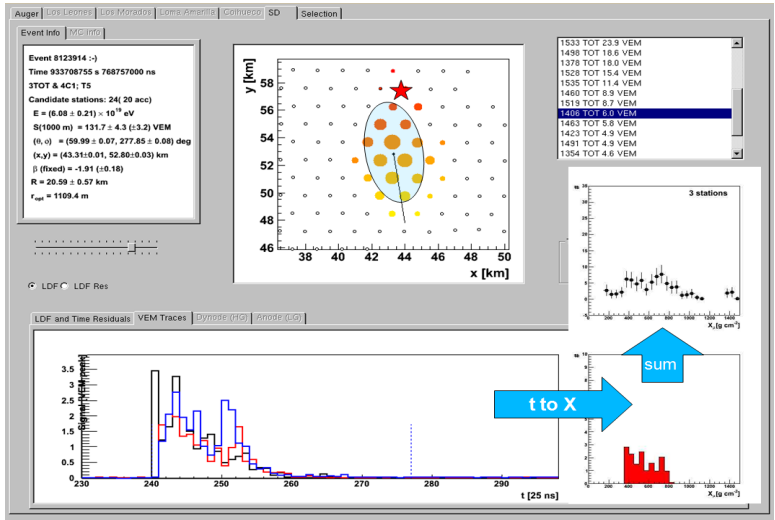
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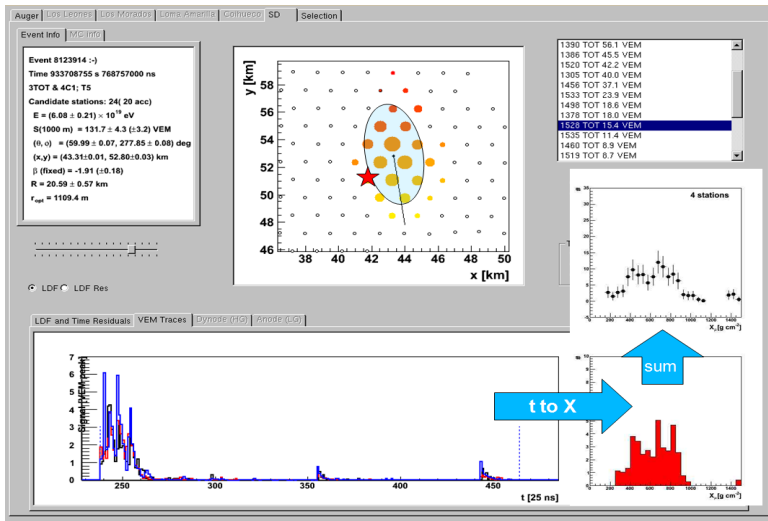
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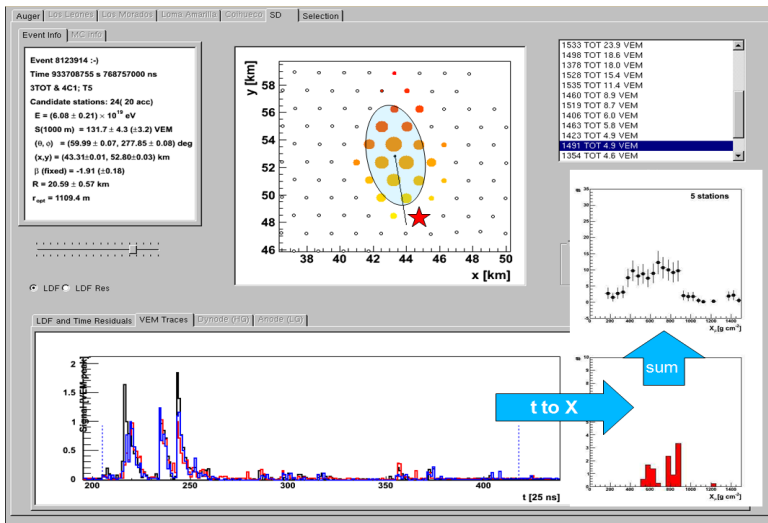
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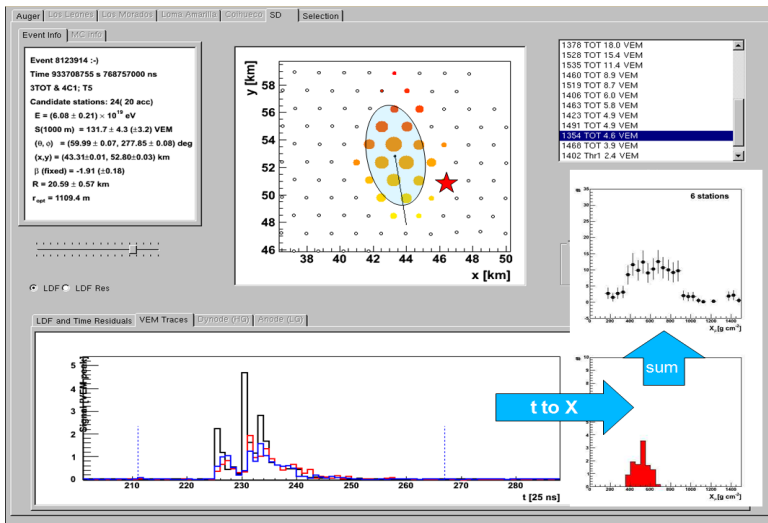
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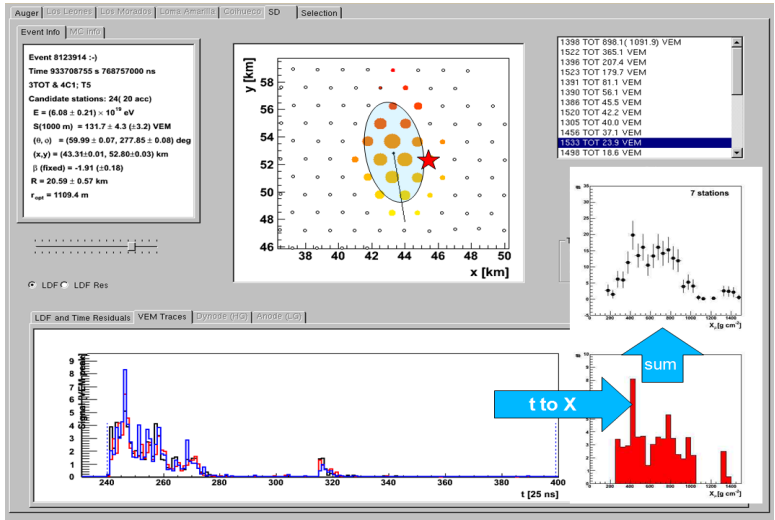
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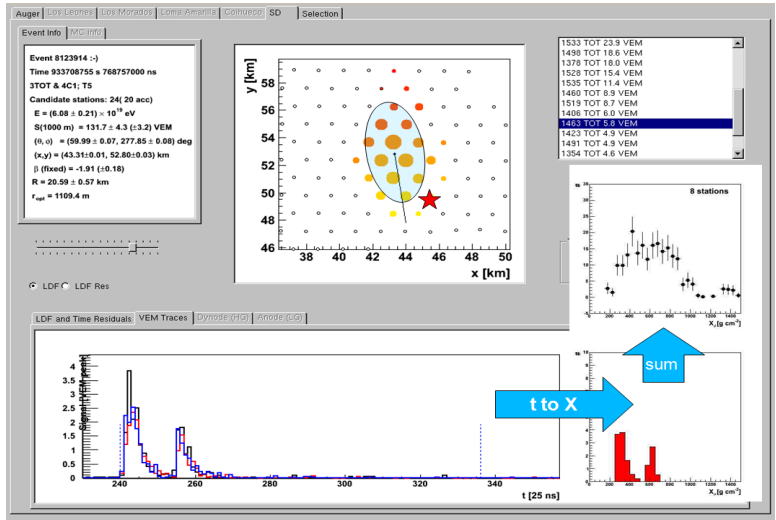
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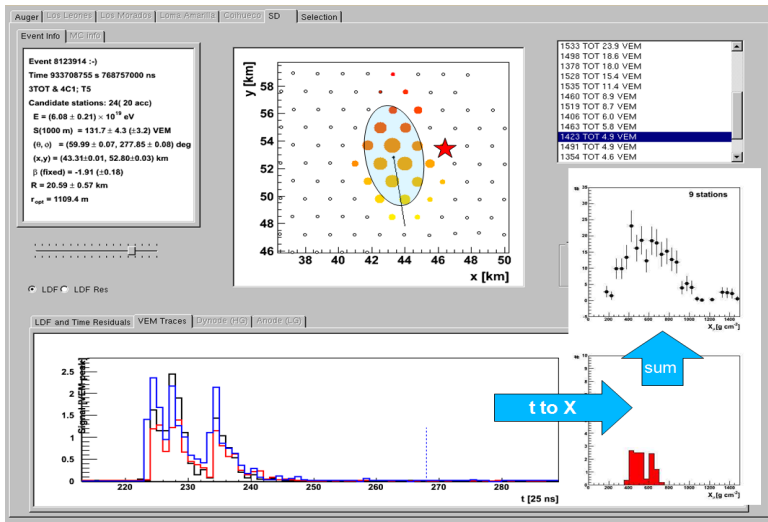
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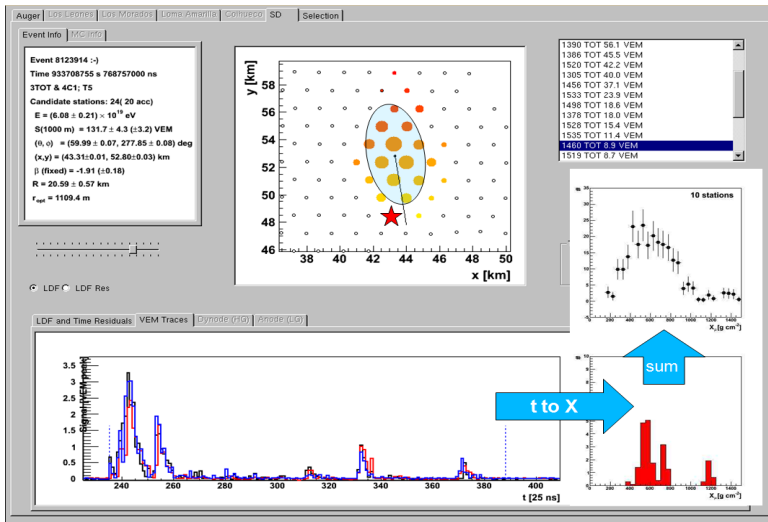
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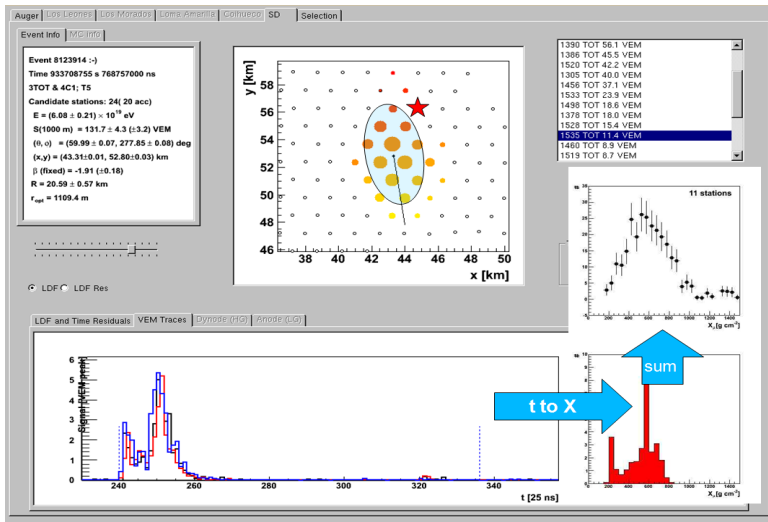
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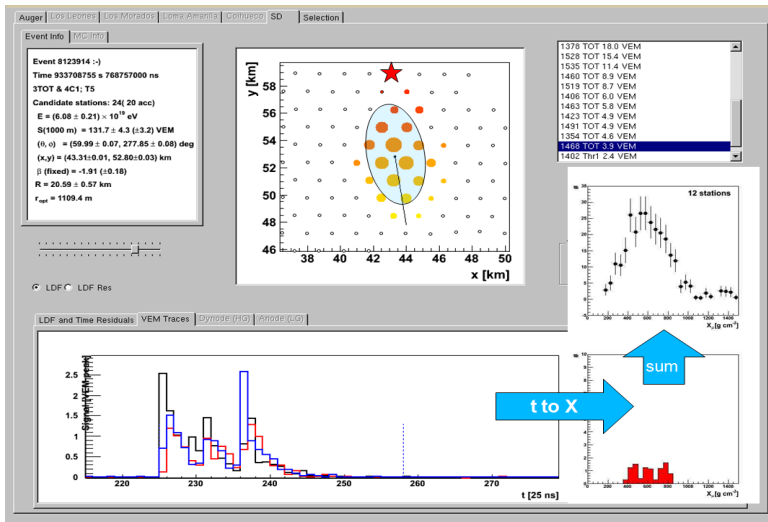
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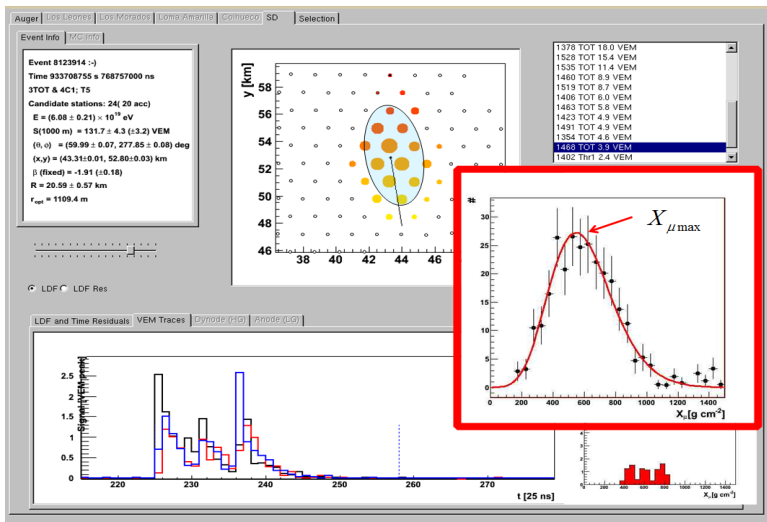
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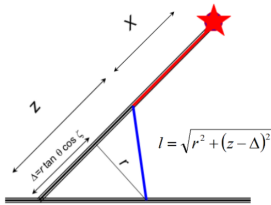
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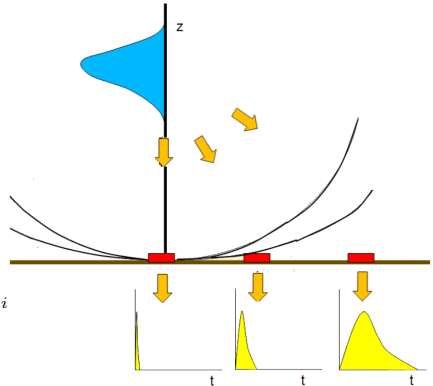
X_{max} distribution interpretation

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Measuring the muonic profile - Muon Production Depth (MPD)



Muons travel in straight lines. Modelling their p_t distribution, and correcting for decays and atmospheric effects we can calculate the likelihood that a muon reaches a tank at time t_i



MPD 2 global fit

Electromagnetic and muonic cosmic ray shower profiles

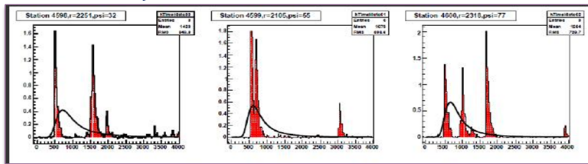
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 X_{max} distribution interpretation

Muonic Longitudinal profiles

Measuring the muonic profile - Muon Production Depth (MPD)

- Fits directly over the stations FADC traces



Eva dos Santos
PhD thesis

$$\log L = \sum \sum \log[(t_i) * pdf(t_i | X_{max}^{\mu}, L, R)] \rightarrow \text{USP}$$

[Detector \otimes (Geometry \otimes Kinematics)] && Muon decay

MPD 2 global fit

Electromagnetic and muonic cosmic ray shower profiles

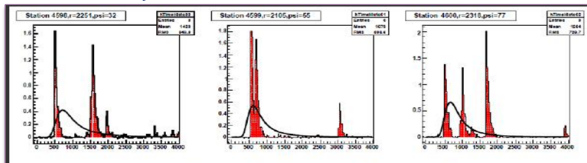
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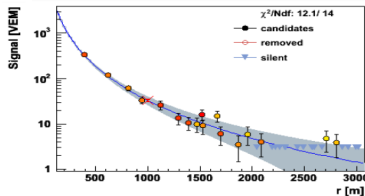
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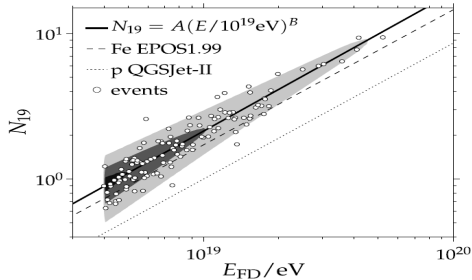
[Detector \otimes (Geometry \otimes Kinematics)] && Muon decay



Micael Oliveira
Master Thesis

Muon excess

Understand the 'muon excess' problem at Auger



$$N_{\mu} = 2.13 \pm 0.04(\text{stat}) \pm 0.11(\text{syst})$$

with respect to QGSJET-II protons at 10^{19}eV

Conclusion

- Auger data presents particle physics puzzles
- the longitudinal profiles of both the electromagnetic and muonic component are the key for building powerful observables
- the LIP Auger team is deeply involved in this task

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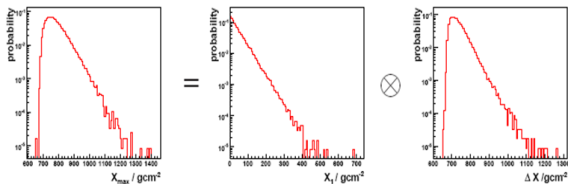
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BACKUP SLIDES

Cross-section measurement



X_1 is an interaction length, and therefore follows an exponential distribution

Cross-section measurement

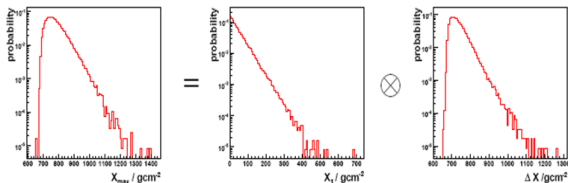
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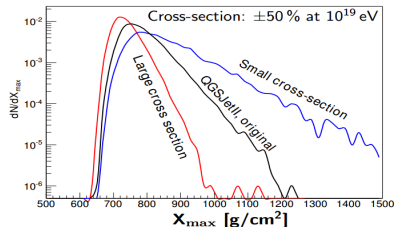
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X_1 is an interaction length, and therefore follows an exponential distribution

The end of the X_{max} distribution tail is dominated by this exponential \rightarrow cross section measurement

$$\sigma_{in} = 1/\overline{X_1}$$



Cross-section measurement

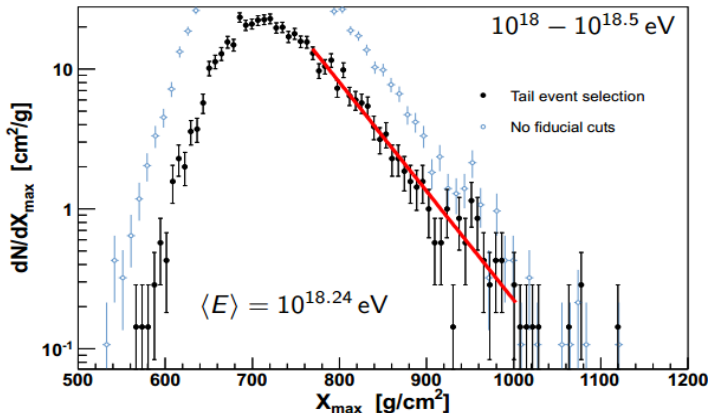
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$$\Lambda_{\eta} = [55.8 \pm 2.3_{\text{stat}} \pm 1.6_{\text{sys}}] \text{ g}/\text{cm}^2$$

Learning from the LHC

Electromagnetic and muonic cosmic ray shower profiles

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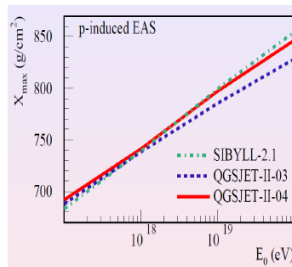
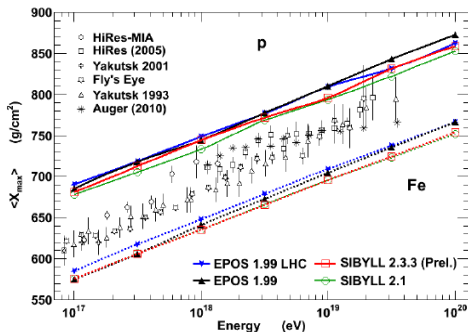
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● Cross section and multiplicity fixed at 7 TeV

- ➔ smaller $\langle X_{max} \rangle$ for EPOS and larger for QGSJETII
- ➔ re-tuned model converge to old Sibyll 2.1 predictions
- ◆ reduced uncertainty from $\sim 50 \text{ g/cm}^2$ to $\sim 20 \text{ g/cm}^2$ (difference proton/iron is about 100 g/cm^2)



from S. Ostapchenko