

TH/LPCC Institute on SM at the LHC CERN, 4 October 2012

#### LUND UNIVERSITY

Faculty of Science

# Inclusion of W/Z emission in showers

## 

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#### based on ongoing work together with

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 $\Rightarrow$  no definite conclusions

#### Motivation — Status — Prospects

#### Motivation 1

Why still try to improve showers? Did not matching solve it all?

- curiosity: what are limits of shower approach?
- Sudakovs for CKKW-L merging
- showers below merging scale
- rapid implementation and studies of new processes
- ditto for showers in new sectors (gauge groups)

Past experience: showers often do better than expected

- FSR in  $e^+e^-\to \gamma^*/Z^0\to q\overline{q}$  (1986), and other resonances
- ISR in  ${\rm q}\overline{\rm q}\to\gamma^*/{\rm Z}^0/{\rm W}^\pm$  (1998), and other colour singlets
- ISR+FSR in QCD jets, with  $p_{\perp \mathrm{shower}} < p_{\perp \mathrm{ME}}$
- ISR(+FSR) in other colour exchange processes (e.g.  $t\bar{t}$ ), with  $dp_{\perp}^2/p_{\perp}^2 \rightarrow dp_{\perp}^2/p_{\perp}^4$  at characteristic scale of process

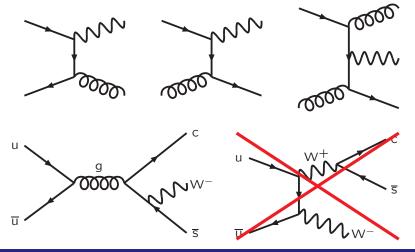
Third let v While showers work for W/Z + 1 jet do/dp⊥ [pb/GeV] AS date vthia8 default they fail for  $W/Z + \ge 2$  jets: Pythia8 ME<sub>2</sub>PS vthia8 ME<sub>3</sub>PS  $\sigma_{\perp}^{\text{jet}} > 20 \text{ GeV}$ 10 Inclusive Jet Multiplicity  $\sigma(W + \ge N_{jet} \text{ jets}) \text{ [pb]}$ 10 ATLAS data Pvthia8 default 103 Pythia8 ME2PS Pythia8 ME<sub>3</sub>PS  $p_{\perp}^{\text{jet}} > 20 \text{ GeV}$ MC/data 1.4 1.2  $10^{2}$ 0.6 0.4 20 80 100 120  $p_{\perp}$  [GeV] 10<sup>1</sup> Azimuthal Distance of Leading Jets ldq] φΔb/bb] 250 Pythia8 default Pythia8 ME2PS 12 Pythia8 ME2P > 20 GeV 150 1.5 MC/data 100 50 0.5 1.4 0 1.3 5 MC/data 0 1 2 3 4 N<sub>jet</sub> 0.8 0.6 (CKKW-L merging by Stefan Prestel) 0.5 1.5

Torbjörn Sjöstrand

Δφ(First Jet, Second Jet)

#### Histories for W/Z production

Q: So what is unique about W/Z + 2 jets? A: First order in which core "hard process" cannot be chosen as W/Z production!



## $\mathrm{W}/\mathrm{Z}$ production in showers

Need to start from QCD 2  $\rightarrow$  2 and add shower emission of W/Z: • FSR: final-state radiation  $q \rightarrow q' W^{\pm}$ ,  $q \rightarrow q Z^{0}$ .

 $\bullet$  ISR: largely already covered by W/Z production processes.

Project at a primitive stage; for now only  $e^+e^-$  annihilation. Viewed as interleaved evolution:

$$\begin{aligned} \frac{\mathrm{d}\mathcal{P}}{\mathrm{d}p_{\perp}} &= \left( \frac{\mathrm{d}\mathcal{P}_{\mathrm{QCD}}}{\mathrm{d}p_{\perp}} + \frac{\mathrm{d}\mathcal{P}_{\mathrm{QED}}}{\mathrm{d}p_{\perp}} + \frac{\mathrm{d}\mathcal{P}_{\mathrm{W/Z}}}{\mathrm{d}p_{\perp}} \right) \\ &\times & \exp\left( - \int_{p_{\perp}}^{p_{\perp i-1}} \left( \frac{\mathrm{d}\mathcal{P}_{\mathrm{QCD}}}{\mathrm{d}p_{\perp}'} + \frac{\mathrm{d}\mathcal{P}_{\mathrm{QED}}}{\mathrm{d}p_{\perp}'} + \frac{\mathrm{d}\mathcal{P}_{\mathrm{W/Z}}}{\mathrm{d}p_{\perp}'} \right) \mathrm{d}p_{\perp}' \right) \end{aligned}$$

Dipole formulation of shower, with each end radiating, e.g.

$$\mathrm{d}\mathcal{P}_{\mathrm{q}\to\mathrm{qg}} = \frac{\mathrm{d}p_{\perp}^2}{p_{\perp}^2} \frac{\alpha_{\mathrm{s}}}{2\pi} C_{\mathsf{F}} \frac{1+z^2}{1-z} \,\mathrm{d}z$$

+ ME corrections where available/convenient ( $\sim$ POWHEG)

## Matrix Elements merging

Consider 
$$e^+e^- \rightarrow \gamma^*/Z^0 \rightarrow q\overline{q} \rightarrow q\overline{q}Z^0$$
:

 $\frac{1}{\sigma_0} \frac{\mathrm{d}\sigma}{\mathrm{d}x_1 \,\mathrm{d}x_2} = \frac{\alpha_{\mathrm{em}}}{2\pi} \frac{v_{\mathrm{q}}^2 + a_{\mathrm{q}}^2}{16 \sin^2 \theta_W \cos^2 \theta_W} \\ \times \left( \frac{x_1^2 + x_2^2 + 2r(x_1 + x_2) + 2r^2}{(1 - x_1)(1 - x_2)} - \frac{r}{(1 - x_1)^2} - \frac{r}{(1 - x_2)^2} \right)$ 

with 
$$x_1 = 2E_q/\sqrt{s}$$
,  $x_2 = 2E_{\overline{q}}/\sqrt{s}$ , and  $r = M_Z^2/s$ .

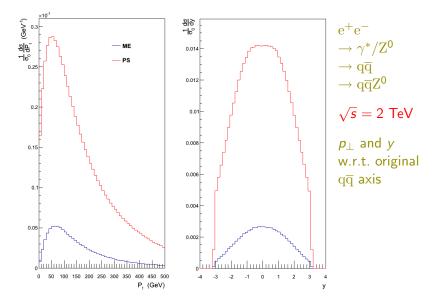
Split ME in "propagator" proportions  $(1 - x_1) : (1 - x_2)$  for comparison with PS  $q \rightarrow qZ : \overline{q} \rightarrow \overline{q}Z$ .

$$\mathrm{d}\mathcal{P}_{\mathrm{q}\to\mathrm{qZ}} = N \frac{\mathrm{d}\mathrm{p}_{\perp\mathrm{evol}}^2}{\mathrm{p}_{\perp\mathrm{evol}}^2(+M_Z^2)} \frac{1}{1-z'} = f_{\mathrm{Z}}(x_1,x_2) \,\mathrm{d}x_1 \,\mathrm{d}x_2$$

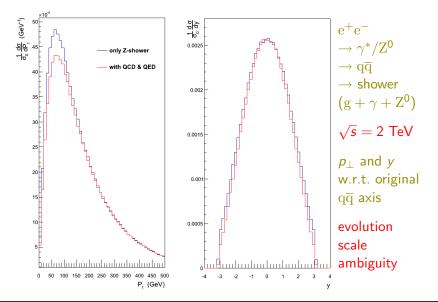
Gives weight  $ME(x_1, x_2)/PS(x_1, x_2) < 1 + veto algorithm.$ 

(Detailed kinematics: L. Carloni, J. Rathsman, T.S., JHEP 04(2011)091)

#### Matrix Elements correction step

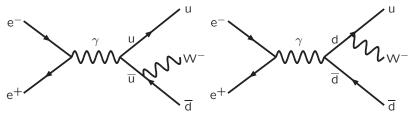


#### Competition from interleaving



#### Flavour interference in W emission

For W emission interference between two dipole ends is replaced by interference between two flavour topologies:

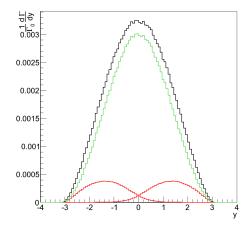


 $|M|^{2}(\gamma^{*} \to u\bar{d}W^{-}) = e_{u}^{2}A(x_{1}, x_{2}) + e_{d}^{2}A(x_{1}, x_{2}) + 2e_{u}e_{d}B(x_{1}, x_{2}) +$ diagrams with WW intermediate states

•  $Z \to u \overline{d} W^-$ : different couplings, same structure (& problem).

- $\bullet \ \mathrm{g} \to \mathrm{u} \overline{\mathrm{d}} \mathrm{W}^- \text{:} \ \textbf{\textit{e}}_\mathrm{u} \to 1, \textbf{\textit{e}}_\mathrm{d} \to 1.$
- $\mathrm{H} \to \mathrm{u}\overline{\mathrm{d}}\mathrm{W}^-$ :  $e_\mathrm{u} \to m_\mathrm{u}, e_\mathrm{d} \to m_\mathrm{d}$ , and also  $A \to A', B \to B'$ .
- $Z^0$  emission: common couplings, same A, B, no  $Z^0 \rightarrow Z^0 Z^0$ .

#### Dipole radiation is it



$$\mathrm{g} 
ightarrow \mathrm{q} \overline{\mathrm{q}} \mathrm{Z}^{0}$$
  
 $\sqrt{s} = 2 \ \mathrm{TeV}$   
 $y = \frac{1}{2} \ln \frac{1 - x_{1}}{1 - x_{2}}$ 

black	=	total emission
ыаск	=	total emission

- green = interference (B)
  - red = individual (A)

## Non-unitary Sudakov factors (1)

Leading electroweak corrections of type  $\alpha_{\rm w} \ln^2(Q^2/M_{\rm W}^2)$ :



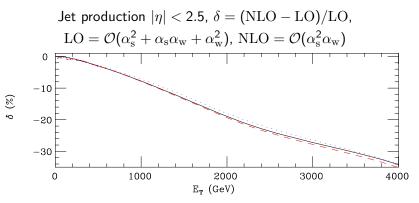
Bloch-Nordsieck violation: real/virtual non-cancellation

- W/Z in final state is another class of events
  - $\Rightarrow$  large negative correction to no-W/Z cross sections!

even if not separated, negative virtual win out (by how much?)

- P. Ciafaloni and D. Comelli, Phys. Lett. B446 (1999) 278;
- M. Beccaria, P. Ciafaloni, D. Comelli, F.M. Renard, C. Verzegnassi, Phys. Rev. D61 (2000) 073005;
- M. Melles, Phys. Rep. 375 (2003) 219;
- U. Baur, Phys. Rev. D75 (2007) 013005;
- A. Banfi, G. Salam, G. Zanderighi, JHEP 0707 (2007) 026;
- G. Bell, J.H. Kühn, J. Rittinger, Eur. Phys. J. C70 (2010) 659;
- S. Dittmaier, A. Huss, C. Speckner, arXiv:1210.0438;
- + many more (apologies to anybody present who is forgotten)

S. Moretti, M.R. Nolten and D.A. Ross, Nucl. Phys. B759 (2006) 50



How much compensated by real emission? To come! What to do with uncompensated part? Introduce by hand?

## Outlook

Many issues remain to address before complete framework:

- step from  $e^+e^-$  to pp
- ullet mass spectrum and complete interference for  $\gamma^*/Z^0$
- $\bullet\,$  flavour separation for W production
- polarized  $\gamma^*/{\rm Z/W}$  decays
- multiple emissions
- evolution scale ambiguity for competition
- on non-unitary Sudakovs
- . . .