



Exploring SUSY parameter space with MultiNest

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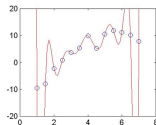
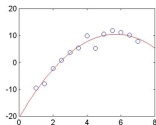
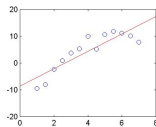
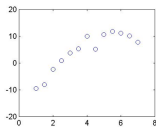


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- 4 An aside on Bayesian statistics
- 5 The MultiNest algorithm
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The problem of vast parameter spaces

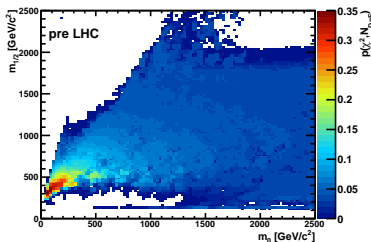


- Theories contain parameters with a priori unknown values
 - determined from experimental data
- If many free parameters
 - can accommodate ‘any’ observations
 - low predictability
- Unbroken SUSY: 1 new parameter
- Broken SUSY: 4 – 105 new parameters
- Need methods for exploring huge parameter spaces
 - to exclude models
 - to identify interesting parameter regions





- Given a parameter space
 - e.g. CMSSM: $m_0, m_{1/2}, A_0, \tan \beta, \text{sgn}(\mu)$
- Given a set of observations
 - $m_h, \Omega h^2, \text{decay rates}, \dots$
- Pick parameter points according to some algorithm
- For each point:
 - calculate theory predictions
 - compare with observed values
- Quantify the level of agreement in some way
 - e.g. calculate a likelihood value

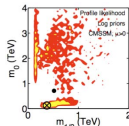
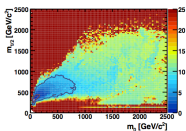
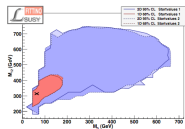
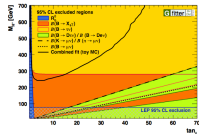


[MasterCode]

Oh joy, another scanning tool...



- **Present goal:**
Construct a scanning tool to explore various SUSY models
- There exist several scanning tools already (*GFitler*, *Fittino*, *MasterCode*, *SuperBayes*, ...) – why make another one?
- For the existing codes:
 - some codes are not public
 - interesting scans already performed by code authors
- Important to have ‘complete’ control of the code
- Easier to tailor the code for future needs
 - mass degenerate models
 - only stau coannihilation regions
 - other special parameter relations





- Algorithm for choosing points: *MultiNest*
 - based on Bayesian statistics
- Bayes' theorem:

$$P(\theta|D, H) = \frac{P(D|\theta, H)P(\theta, H)}{P(D|H)} = \frac{L(\theta)\pi(\theta)}{Z}$$

H : model

θ : model parameters

D : data

$P(\theta|D, H)$: posterior probability (dist.) of parameters

$P(D|\theta, H) = L(\theta)$: likelihood of parameters

$P(\theta|H) = \pi(\theta)$: prior probability (dist.) of parameters

$P(D|H) = \int L(\theta)\pi(\theta) d\theta = Z$: the 'Bayesian evidence'

- Consistent approach for updating one's *degree of belief* in θ, H in light of the data D (what theorists are interested in. . .)



■ Bayesian inference:

- parameter estimation
- model comparison

■ Parameter estimation:

- use information contained in posterior $P(\theta|D, H)$
(mean, maximum, credible intervals, ...)
- Z can often be ignored (only normalization)
- coincide with frequentist results when data are strong enough

■ Model comparison:

- rank models H_0, H_1 based on posterior probabilities $P(H_0|D), P(H_1|D)$

$$\frac{P(H_1|D)}{P(H_0|D)} = \frac{P(D|H_1)P(H_1)}{P(D|H_0)P(H_0)} = \frac{Z_1}{Z_0} \frac{P(H_1)}{P(H_0)}$$

- numerically challenging to calculate $Z = \int L(\theta)\pi(\theta) d\theta$
(many parameters, complicated $L(\theta)$)
- MultiNest is designed to calculate Z efficiently



- MultiNest (Feroz, Hobson) is a Fortran code based on the *nested sampling* algorithm (Skilling)
- Calculates Z
- Produces a set of samples of $P(\theta|D, H)$ as a by-product
- Main advantage: good coverage of multi-modal posteriors (important for SUSY)
- Main steps:
 1. draw a set of N points from prior distribution $\pi(\theta)$
 2. calculate likelihoods $L(\theta_1), \dots, L(\theta_N)$
 3. discard the point with lowest likelihood
 4. pick a new point from $\pi(\theta)$, given that $L(\theta_{\text{new}}) > L(\theta_{\text{discarded}})$
 5. repeat from 3
- Keep the set of *discarded* points θ_i



■ What happens?

- new variable $dX = \pi(\theta) d\theta$ (prior mass)
- integral transformed to 1D integral in X :

$$Z = \int_0^1 L(X) dX$$

- at step i : X_i is the prior mass within the parameter region defined by

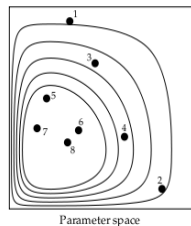
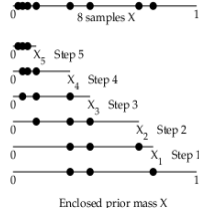
$$L(\theta) > L(\theta_i)$$

- integral approximated by the trapezoidal rule:

$$Z \approx \sum w_i L(\theta_i) = \sum \frac{1}{2} (X_{i-1} - X_{i+1}) L(\theta_i)$$

- set of θ_i is a sample of $P(\theta|D, H)$ if weighted by

$$p_i = \frac{w_i L(\theta_i)}{Z}$$



[Skillig]



■ Current status:

- main steering code is python based
- code written for parallel computing
- complete scan setup (incl. $L(\theta)$ contributions) given in a SLHA file
- python wrapper for MultiNest (Vallisneri)
- predictions from calls to SoftSusy, micrOMEGAs, FeynHiggs, HiggsBounds
- interface for adding user-defined calculations
- SLHA file stored for each accepted (discarded) point
- scripts for summarizing and plotting results

■ To-do list (ever expanding):

- add more observables/predictions
- separate MultiNest from code for easy replacement
- move from 4 core laptop to a more powerful system
- explore some interesting SUSY models!

```
BLOCK ParSetup # settings for the parameters in
# Note: At the moment derived parameters can only depend on fixed
1 m_squarks MINPAR 1 1 1000.0 4000.0 # -- m_squ
2 m_eL EXTPAR 31 1 0.0 1000.0 # -- mL
3 m_muL EXTPAR 32 2 1.0*p[2] #
4 m_tauL EXTPAR 33 2 1.0*p[2] #
5 m_eR EXTPAR 34 1 0.0 1000.0 # -- mR
6 m_muR EXTPAR 35 2 1.0*p[5] #
7 m_tauR EXTPAR 36 2 1.0*p[5] #
8 M1 EXTPAR 1 1 0.0 1000.0 # -- M1
9 M2 EXTPAR 2 1 0.0 1000.0 # -- M2
10 M3 EXTPAR 3 1 1000.0 4000.0 # -- M3
11 tanb MINPAR 3 1 10.0 50.0 # -- tanb
12 mu EXTPAR 23 1 0.0 3000.0 # -- mu
13 mA0 EXTPAR 26 1 1000.0 4000.0 # -- mA0
14 A_t/b MINPAR 5 1 -5000.0 5000.0 # -- At, A
15 A_tau EXTPAR 13 1 -5000.0 5000.0 # -- Atau
#
#
BLOCK SSSetup # settings for SoftSusy (1-7 are
1 1.000000000e-03 # numerical precision: suggested
2 0 # quark mixing parameter: 0.2
3 0 # printout level: 0.3
4 1.0 # multiplicative factor to change
5 1 # include 2-loop scalar mass squ
6 8 # precision -- number of signifi
7 2 # number of loops in REWSB/mh ca
8 _SS_OUTPUT # suffix for temporary files
BLOCK FHSetup # settings for FeynHiggs (for de
1 4 # msspart = 0.4, specifies
2 0 # fieldren = 0.4, determine
3 0 # tanren = 0.2, determine
4 2 # higgsmix = 1.3, determine
5 0 # p2approx = 0.4, determine
6 2 # looplevel = 0.2, determine
7 1 # runningMT = 0.1, determine
8 1 # botResum = 0.1, determine
9 0 # t1plxApprox = 0.6, determine
```

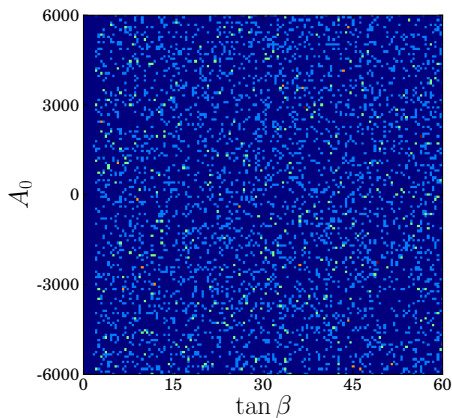
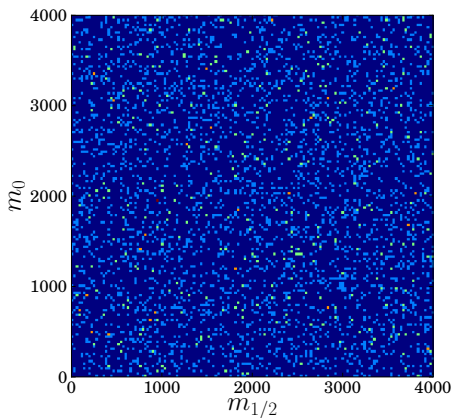
Plots with shiny colours



■ A few preliminary plots: CMSSM/mSUGRA scan

■ $m_0 \in (0, 4000)$ $m_{1/2} \in (0, 4000)$ $A_0 \in (-6000, 6000)$ $\tan \beta \in (2, 60)$ $\text{sgn}(\mu) = +1$

■ Flat priors



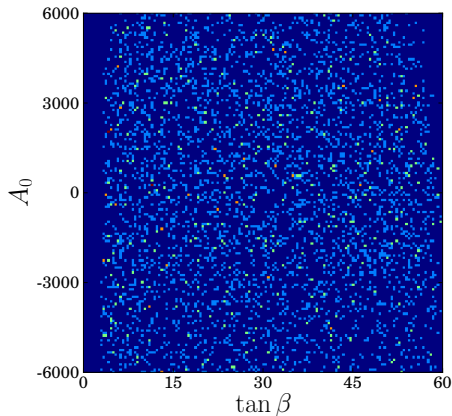
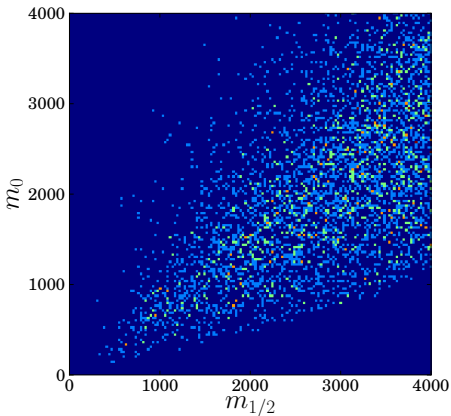
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■ + discrete constraints: neutral LSP, $m_{\tilde{\tau}_1} < 2 m_{\tilde{\chi}_1^0}$, LEP bounds on $m_{\tilde{\tau}_1}, \dots$



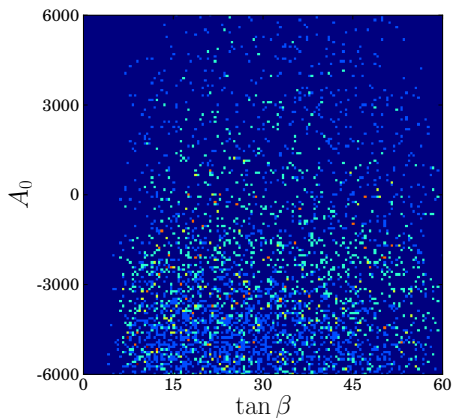
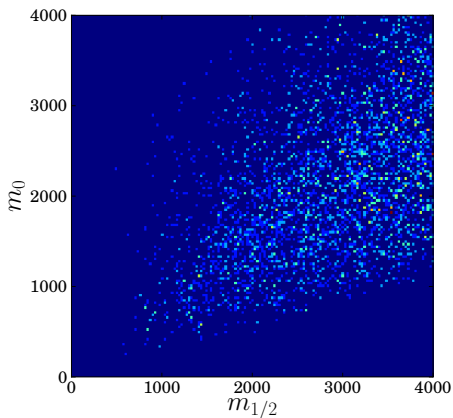
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■ + gaussian likelihood contribution for m_{h^0}



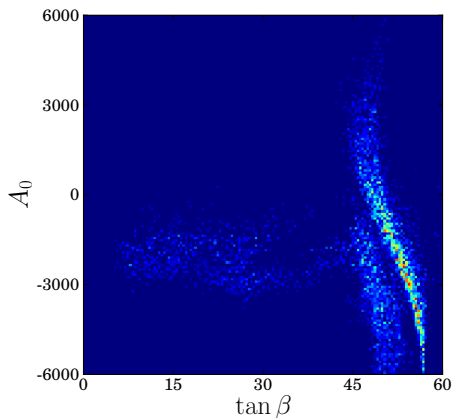
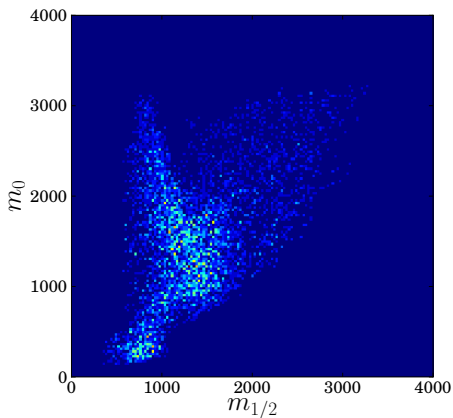
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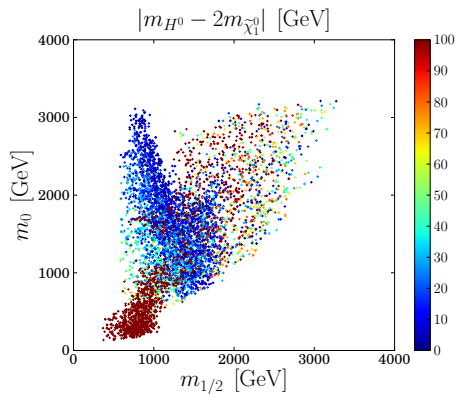
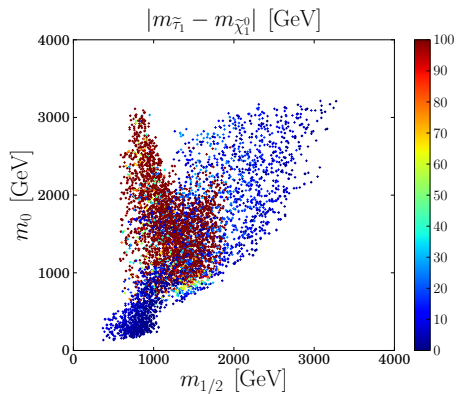
■ + gaussian likelihood contribution for Ωh^2



Plots with shiny colours (cont.)



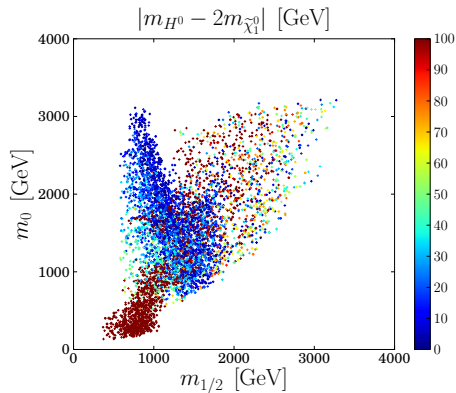
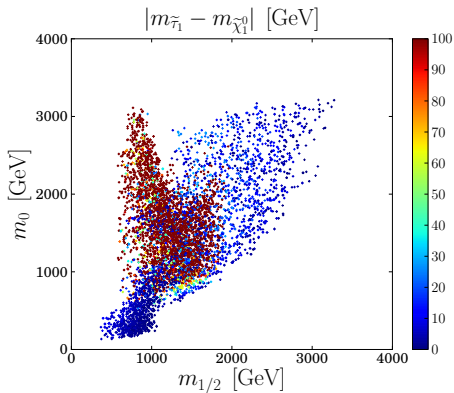
- Found 'stau coannihilation' region and 'Higgs funnel' region



Plots with shiny colours (cont.)



- Found 'stau coannihilation' region and 'Higgs funnel' region
- This was only an illustration — more interesting results to come (hopefully...)





Thank you

Backup slide 1: Bayesian vs frequentist



- Bayesian: all forms of uncertainty are expressed in terms of probabilities
- Bayesian: 'how probable is a parameter value, given the data?'
Frequentist: 'how probable is the data, given a parameter value?'
- Very informative data \rightarrow identical results
- F: priors are subjective!
B: yes, but thinking that high $L \leftrightarrow$ more probable parameter value already assumes some prior
- B: prior dependence give us and estimate of the how robust the inference is
- Model comparison using the Bayesian evidence (Z) incorporates 'Occam's razor'
- Choice of method depends on the question you are asking...