

Remarks

A. Yu. Smirnov

A. Yu. Smirnov

without conclusions

Content:

1. From special to normal

2. Back to the Sun

3. Physics with HAND's

exceptional

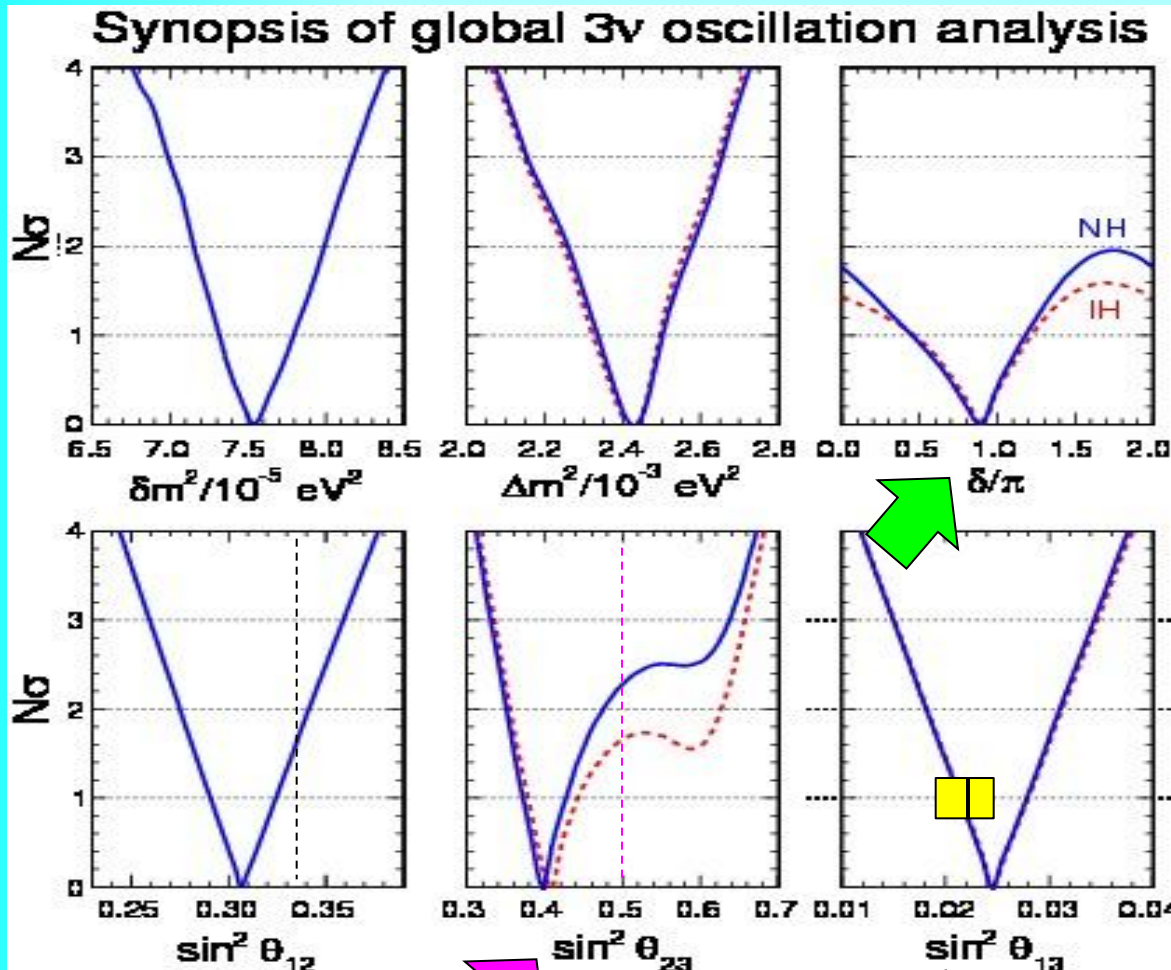
From special
to normal



Synopsis

before
nu2012

G. L. Fogli



Serious implications
for theory

Non-zero, relatively
Large 1-3 mixing

Substantial deviation
of the 2-3 mixing
from maximal

$$\delta_{CP} \sim \pi$$

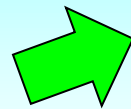
DB new

Robust ?

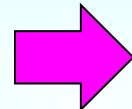
Deviation of 2-3 mixing from maximal

$$d_{23} = \frac{1}{2} - \sin^2 \theta_{23}$$

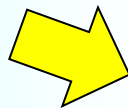
the key to (probe)
understand the
underlying physics



$\nu_\mu - \nu_\tau$ symmetry
violation

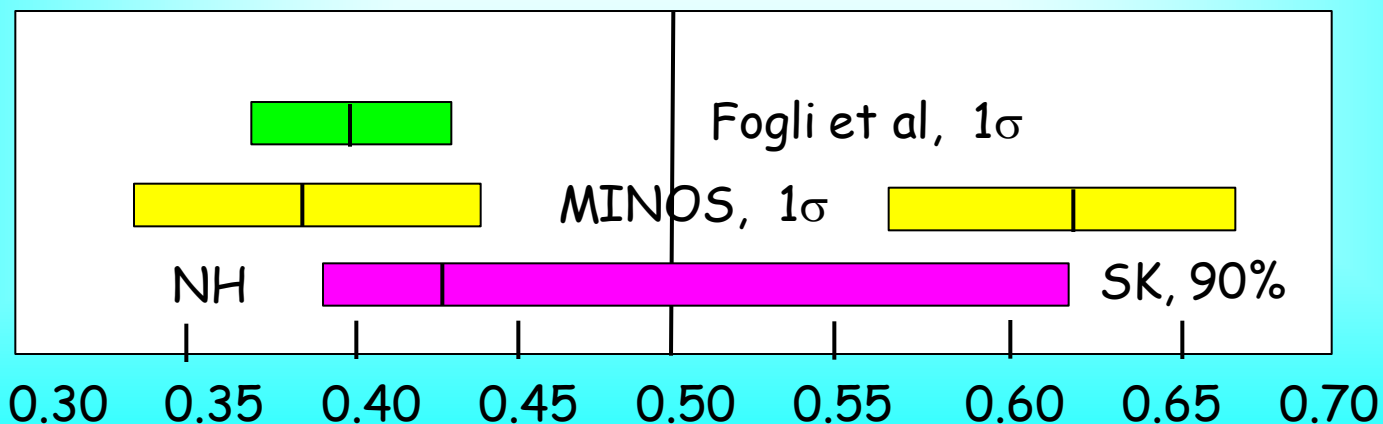


Connection to
1-3 mixing



Quark -Lepton
Complementarity

$$\theta_{23} \sim \pi/2 - V_{cb}$$



$\sin^2 \theta_{23}$

Measuring the deviation

Atmospheric neutrinos

ν_e - oscillation effects

$$\begin{aligned} \frac{F_e}{F_e^0} - 1 &= P_{e2} (r c_{23}^2 - 1) \\ &= P_{e3} (r s_{23}^2 - r) \end{aligned}$$

sub-GeV range

multi-GeV range

$r = F_\mu^0 / F_e^0 \sim 2$ ``screening factor''

The e -like event excess - at low energies and deficit at higher energies - signature of deviation of the 2-3 mixing from maximal (first quadrant)

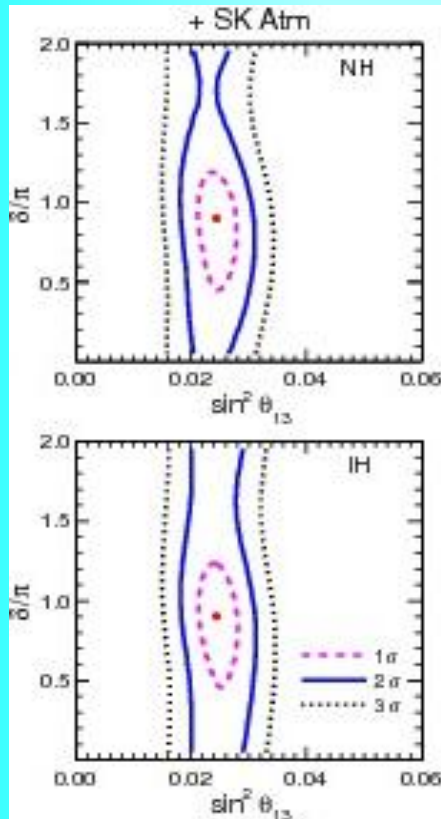
T2K
MINOS

$P_{\mu e} \sim \sin^2 \theta_{13} \sin^2 \theta_{23}$ - appearance

$P_{\mu\mu} \sim \sin^2 2\theta_{23}$ - disappearance

PINGU

CP-phase: measurements and predictions



G. L. Fogli

First glimpses?

T. Yanagida $\delta_{CP} \sim \pi/2 \pm 0.02$

Neutrino-antineutrino asymmetry

Dependence of probabilities on energy in wide range

Reconstruction of unitarity triangle

Key measurement: amplitudes of the $\nu_\mu - \bar{\nu}_\mu$ oscillations due to solar and atmospheric mass splittings

Third way

Do we have predictions for the phase in quark sector?
 Why do we think that we can predict leptonic mixing?
 Again because of neutrinos are special? Symmetries?

Implications for theory

$\sin^2\theta_{13} \sim 0.025$ The same 1-3 mixing with completely different implications

$\sin^2\theta_{13} =$

$$O(1) \frac{\Delta m_{21}^2}{\Delta m_{32}^2}$$

“Naturalness” of mass matrix

$$\sim \frac{1}{2} \sin^2\theta_c$$

Quark Lepton Complementarity

$$\sim \frac{1}{2} \cos^2 2\theta_{23}$$

$\nu_\mu - \nu_\tau$ - symmetry violation

$$> 0.025$$

Mixing anarchy

*A. De Gouvea,
H. Murayama*

$$\theta_{13} + \theta_{12} = \theta_{23} \sim \pi/4$$

Self-complementarity

Mixing pattern



**Tri-bimaximal
mixing**



**Quark-Lepton
complementarity**



**Quark-lepton
universality**

Assuming that it is not accidental and there is certain fundamental physics behind

Based on observation:
lepton mixing =
maximal mixing -
quark mixing

The same principle
as in quark sector

Large mixing is related
to smallness of neutrino
mass and weak mass
hierarchy of neutrinos

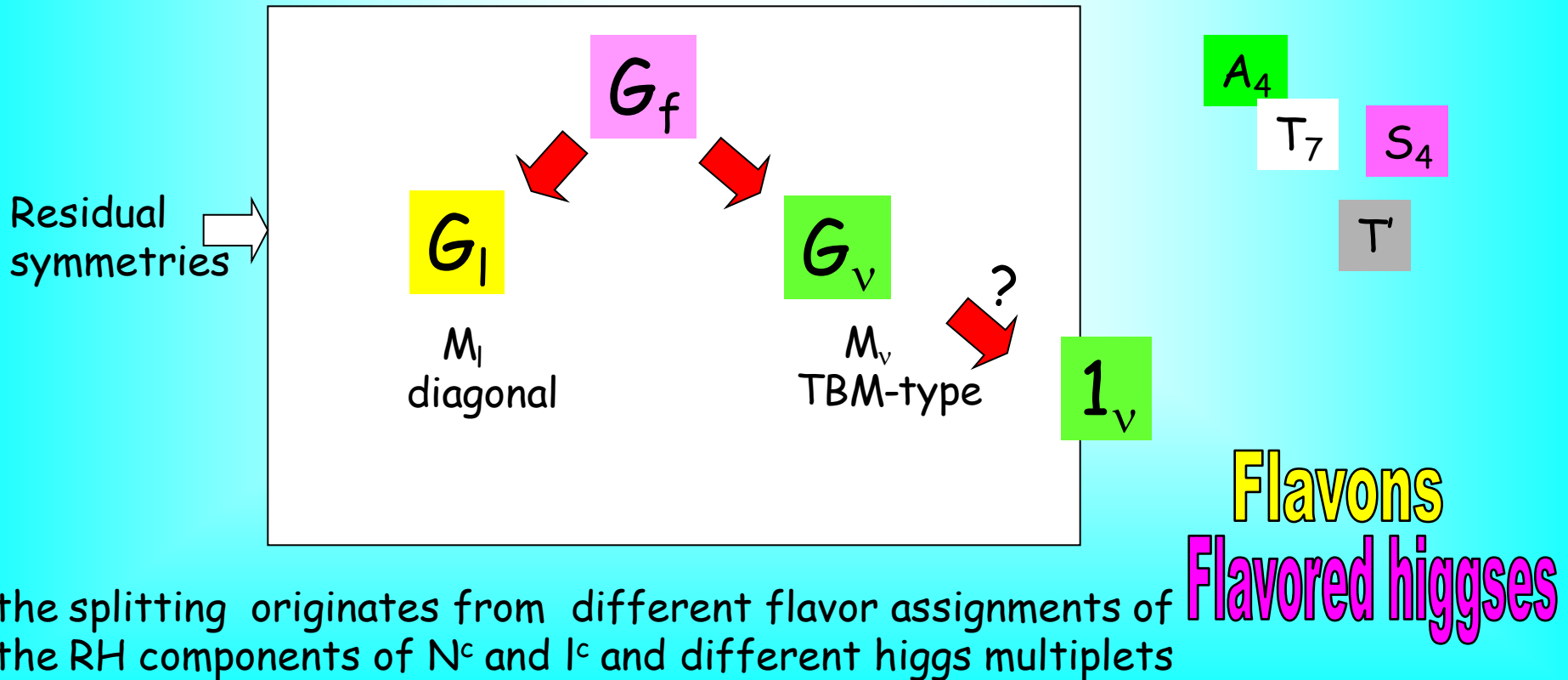
With different
implications

Still something special?

Symmetry in sector responsible for
smallness of neutrino mass?
RH neutrinos?

Symmetry without TBM

Mixing appears as a result of different ways of the flavor symmetry breaking in neutrino and charged lepton sectors



Relations between mixing parameters

If G is von Dyck group $D(2, m, p)$

For column of the mixing matrix:

$$|U_{\beta i}|^2 = |U_{\gamma i}|^2$$

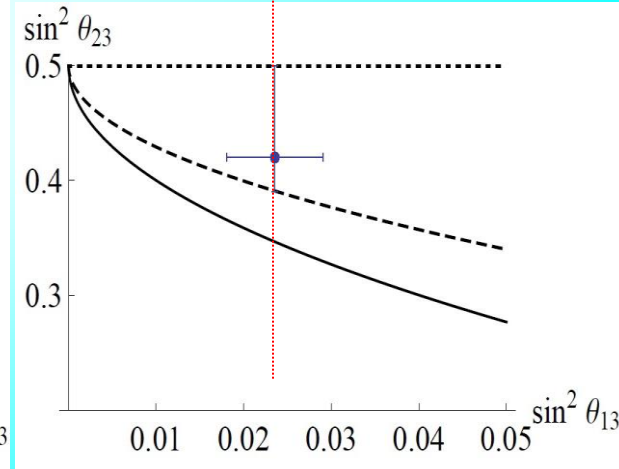
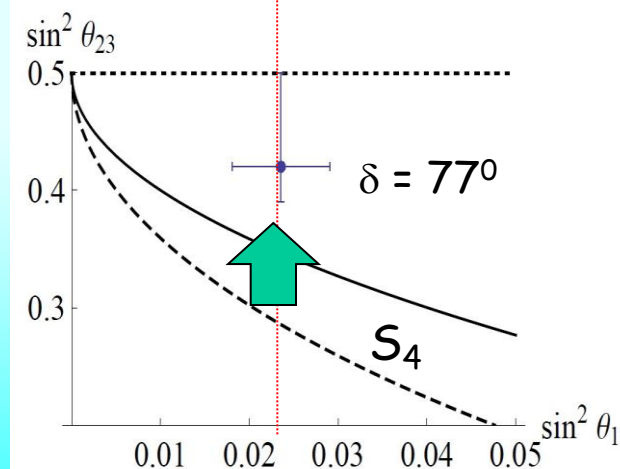
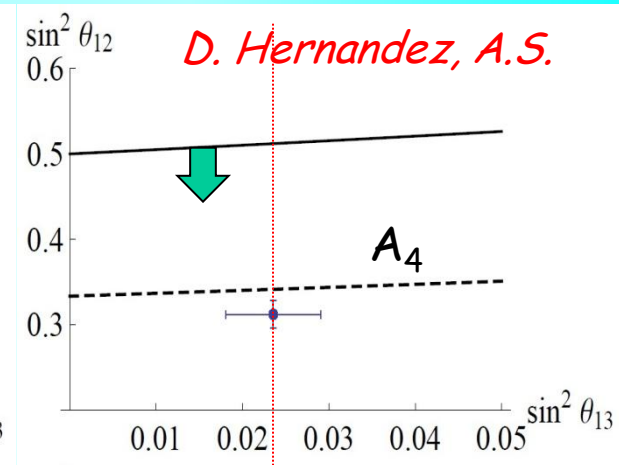
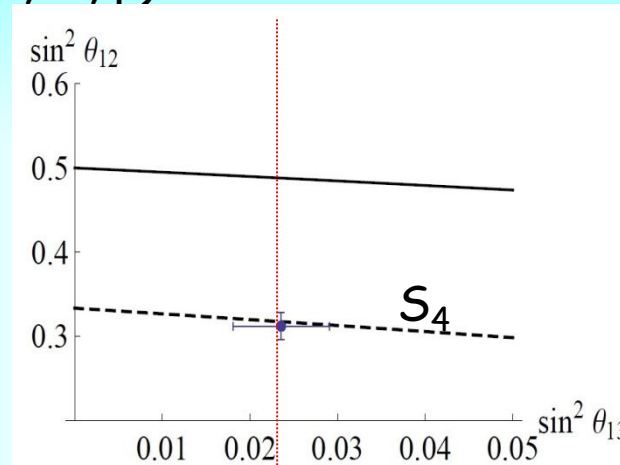
$$|U_{\alpha i}|^2 = \frac{1 - a}{4 \sin^2(\pi k/m)}$$

A is determined from condition

$$\lambda^3 + a \lambda^2 - a^* \lambda - 1 = 0$$

$$\lambda_i^p = 1$$

k, m, p integers which determine symmetry group



*Also S. F. Ge, D. A. Dicus,
W. W. Repko, PRL 108 (2012) 041801*

$$\theta_{13} \sim \frac{1}{\sqrt{2}} \theta_c$$

$$\sin^2 \theta_{13} \sim \frac{1}{2} \sin^2 \theta_c$$

First obtained in the context of
Quark-Lepton Complementarity

H. Minakata, A Y S

Follows from
permutation of
matrices

$$U_{12}(\theta_c) U_{23}(\pi/2)$$

From charged
leptons

Maximal from
neutrinos

Permutation - to reduce the lepton
mixing matrix to the standard form

Again neutrinos
are special!

Related to smallness
of mass

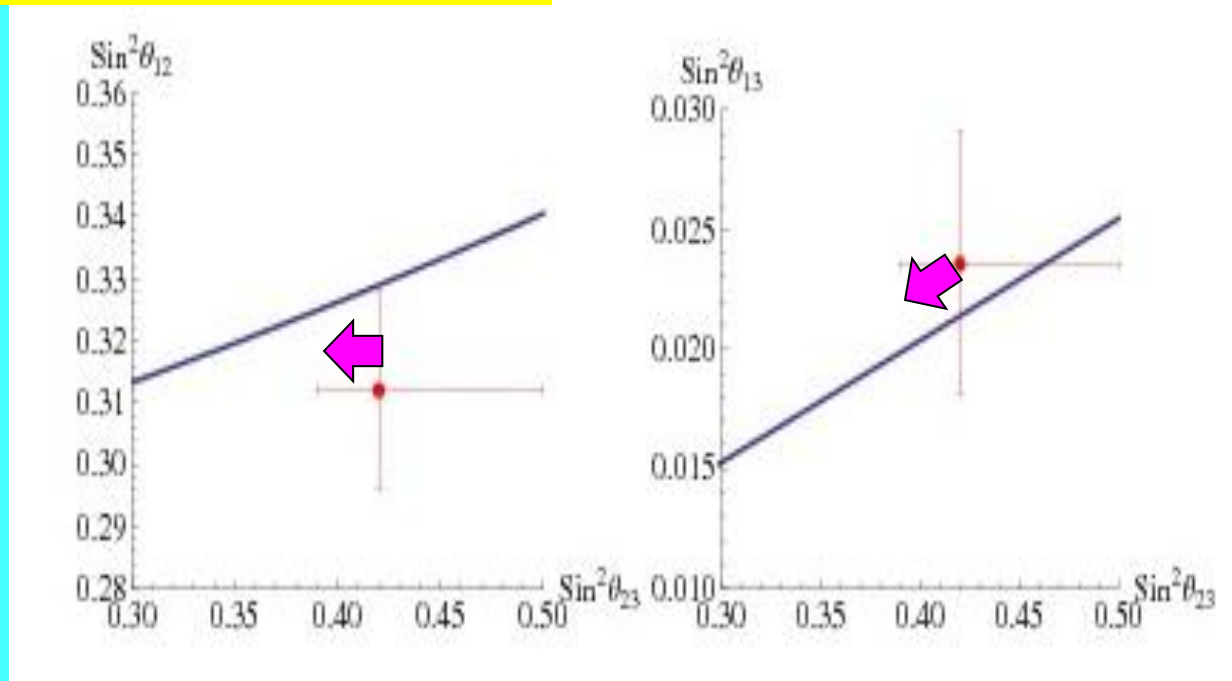
See-saw,
symmetry of RH
neutrino sector

Realizations: QLC

QLC: generalization

$$\sin^2\theta_{13} \sim \sin^2\theta_{23} \sin^2\theta_C$$

*D. Hernandez,
A.S.*



Improves also
predictions
for 1-2 mixing

Bi-maximal mixing?

RGE effect

$$\sin^2\theta_{13} \sim \sin^2\theta_{23} \sin^2\theta_C$$

QLC or Cabibbo "haze"

P. Ramond

Deviations from BM due to high order corrections

Complementarity:
implies quark-lepton
symmetry or GUT,
or horizontal symmetry

Weak complementarity or
Cabibbo haze

Altarelli et al

Corrections from high order
flavon interactions generate
Cabibbo mixing and deviation from BM,
GUT is not necessary

$$\sin\theta_C = \sqrt{\frac{m_\mu}{m_\tau}}$$



$$\sin\theta_C = 0.22$$

as ``quantum'' of
flavor physics

Self-complementarity relations

Xinyi Zhang Bo-Qian Ma, arXiv:1202.4258

Quark-lepton universality?

Similar Ansatz for
structure of mass matrices

Relations between
masses and mixing

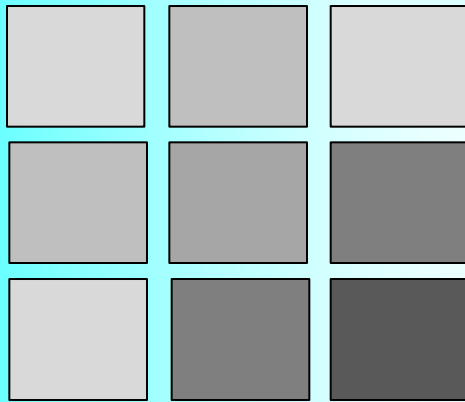
M Fukugita T. Yanagida

Fritsch Ansatz
similar to quark sector
3 RH neutrinos with equal
masses →
Normal mass hierarchy,
Right value of 13 mixing

Flavor ordering

Flavor alignment

Values of elements gradually decrease from $m_{\tau\tau}$ to m_{ee}



corrections wash out sharp difference of elements of the dominant $\mu\tau$ -block and the subdominant e -line

This can originate from power dependence of elements on large expansion parameter $\lambda \sim 0.7 - 0.8$.

Another complementarity: $\lambda = 1 - \theta_c$

Froggatt-Nielsen?

Masses and mixing

$$\sin^2\theta_{13} \sim \frac{\Delta m_{21}^2}{\Delta m_{32}^2}$$

1. Two mass scales in the mass matrix

$$\sqrt{\Delta m_{21}^2}$$

$$\sqrt{\Delta m_{31}^2}$$

2. Two large mixing angles

3. Normal mass hierarchy

4. No fine tuning - no equalities of matrix elements

Assumptions



$$\sin\theta_{13} \sim \sqrt{\Delta m_{21}^2 / \Delta m_{31}^2} = 0.17 - 0.20$$

Implications: - no particular (for leptons) flavor symmetries,
- normal mass hierarchy

Old does not mean wrong

After many speculations back to good old picture?

High scale seesaw

GUT

???

Something is still missed

Flavor physics at very high scales above GUT?

The same mechanism which explains smallness of neutrino mass is responsible for large lepton mixing

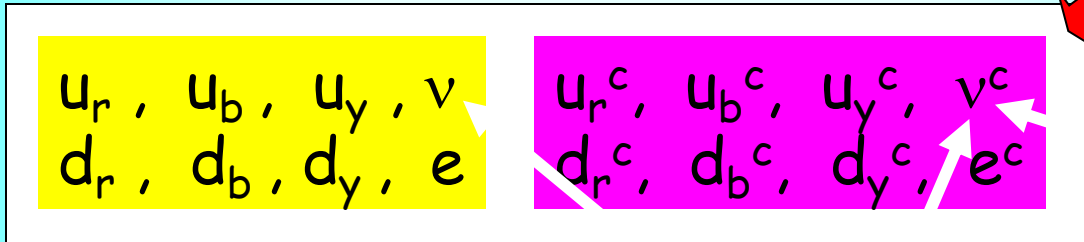
Seesaw enhancement of mixing
No particular symmetry?

Difference of quark and lepton mixings is related to smallness of neutrino mass

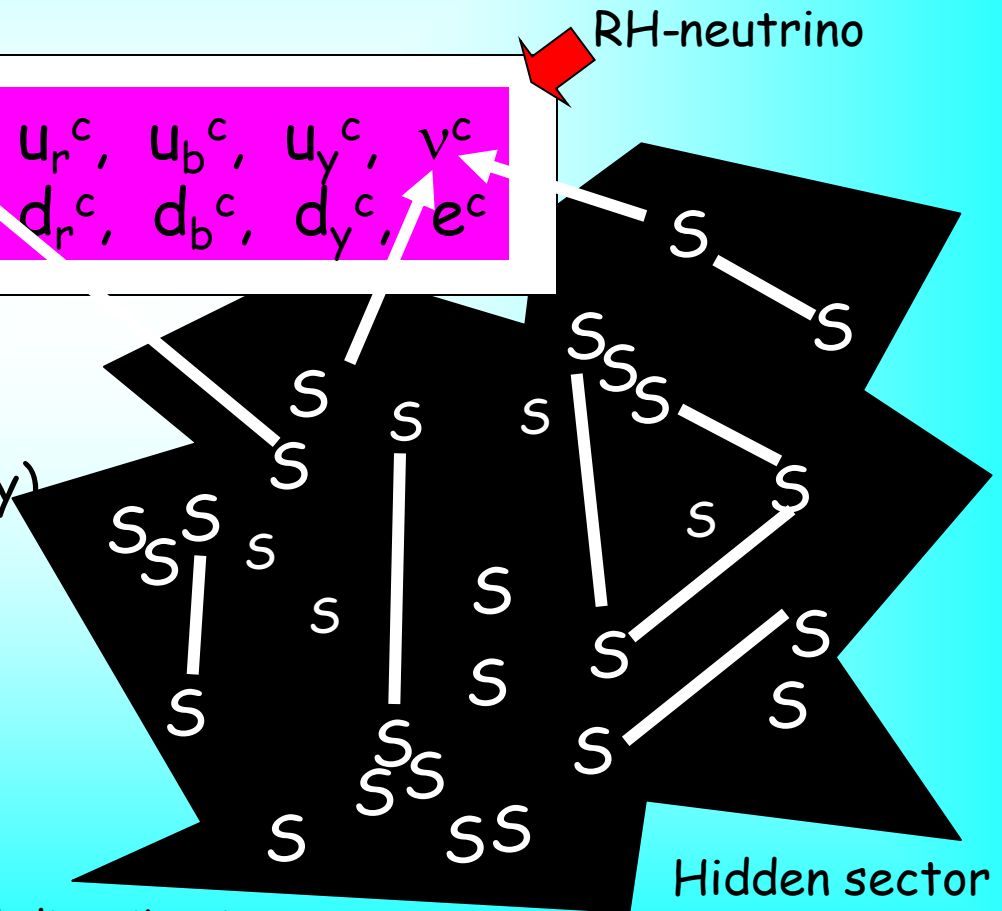
Hidden sector at GUT, Planck scales

SO(10) GUT + ...

16



- Enhance mixing
- Produce randomness (anarchy)
- Seesaw symmetries
- Increase seesaw scale
- produce bi-maximal mixing



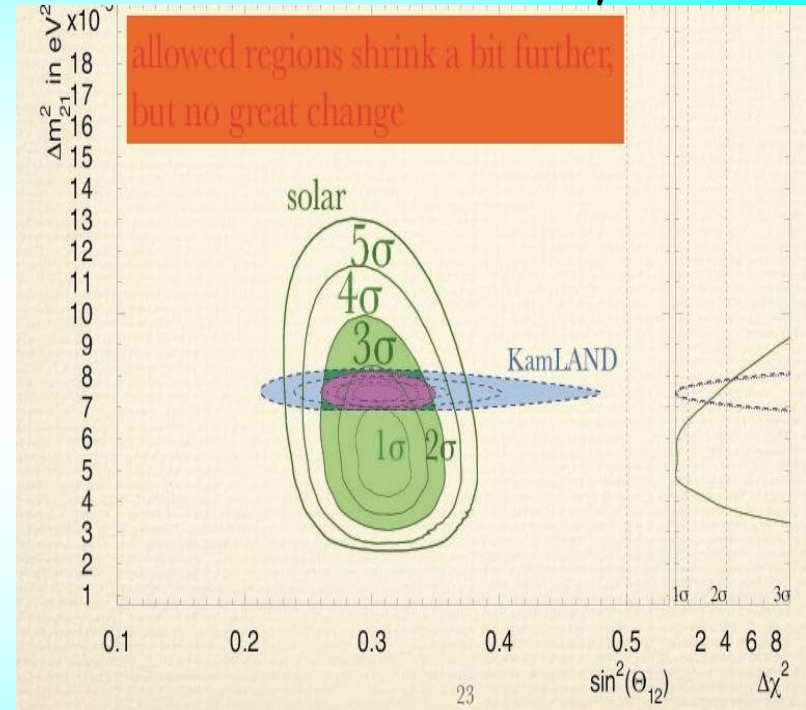
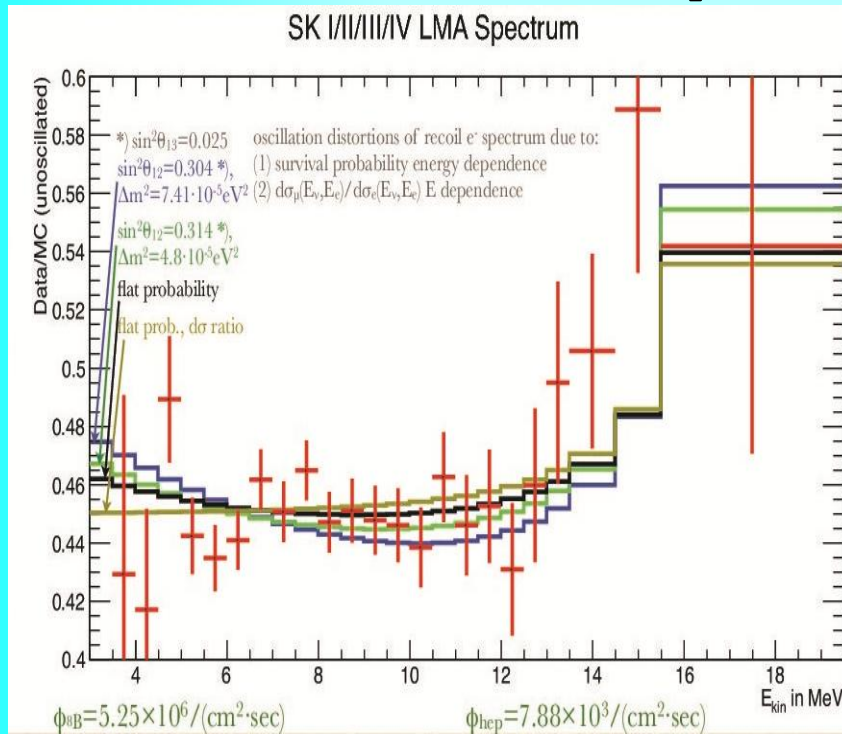
Back to the Sun



Solar neutrinos: two features

SuperKamiokande

M. Smy

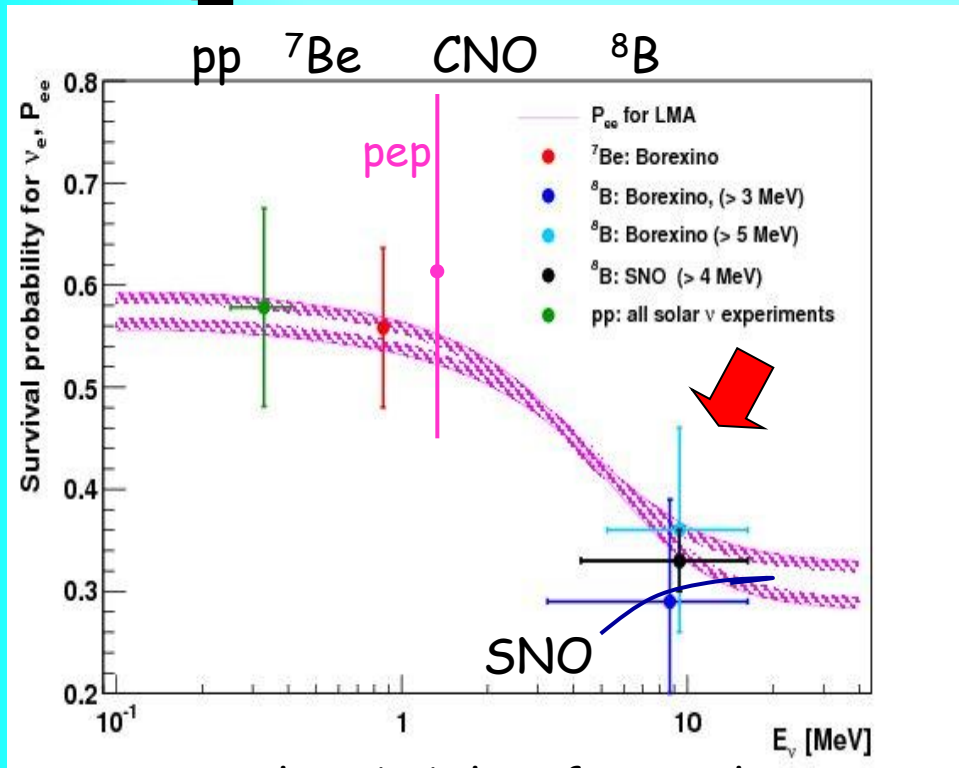


No distortion of the energy spectrum at low energies : the upturn is disfavored at $(1.1 - 1.9) \sigma$ level

Increasing tension between Δm^2_{21} measured by KamLAND and in solar neutrinos 1.3σ level

This is how new physics may show up

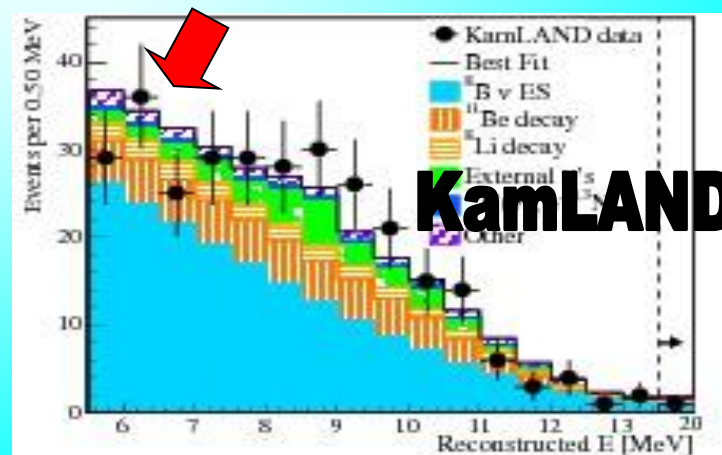
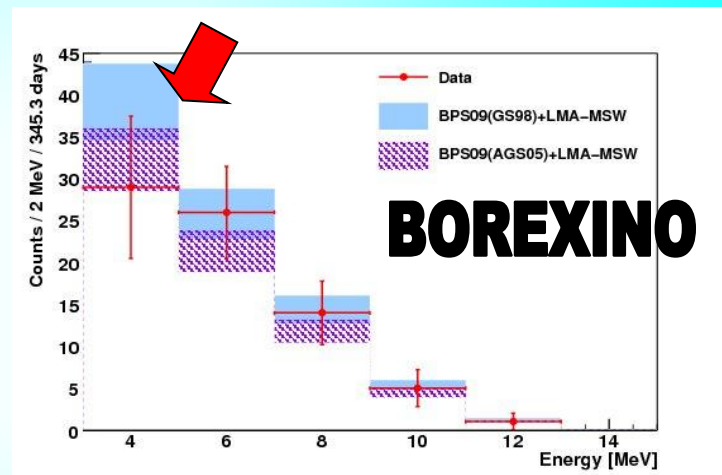
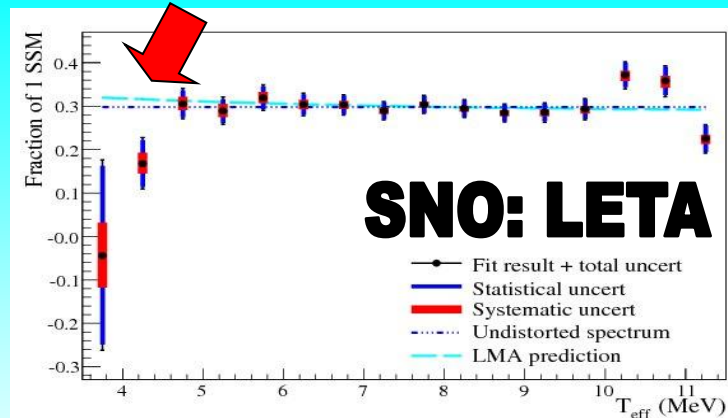
Up-turn?



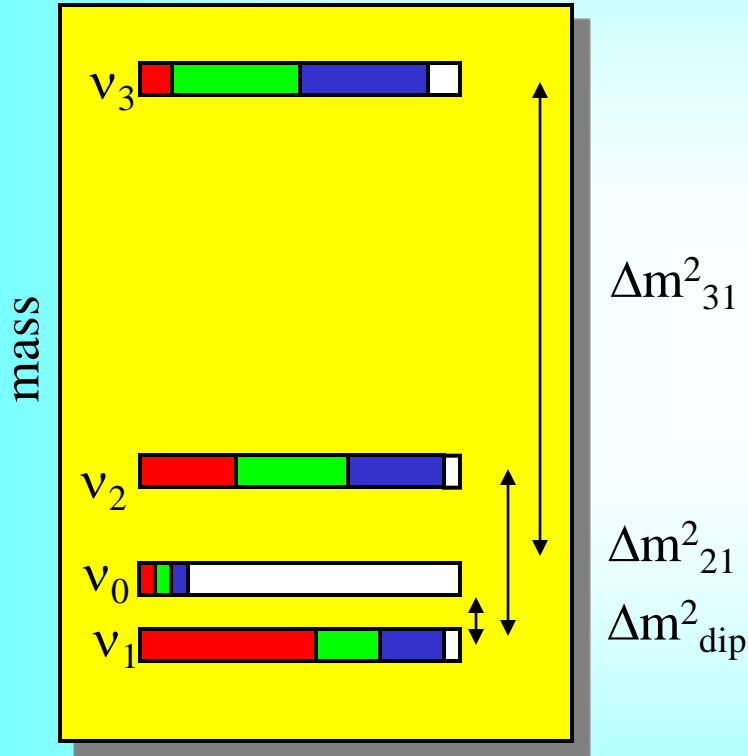
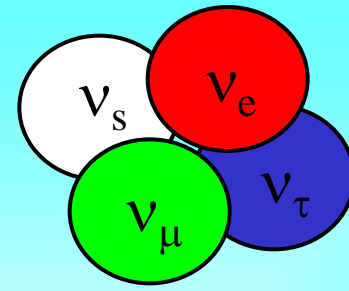
ν_e - survival probability from solar neutrino data vs LMA-MSW solution

HOMESTAKE
low rate

SNO+



Very light sterile



Very light sterile neutrino

$$m_0 \sim 0.003 \text{ eV}$$

DE scale?

$$\frac{M^2}{M_{\text{Planck}}}$$

$$M \sim 2 - 3 \text{ TeV}$$

- solar neutrino data

- additional radiation in the Universe if mixed in ν_3

no problem with LSS (bound on neutrino mass)

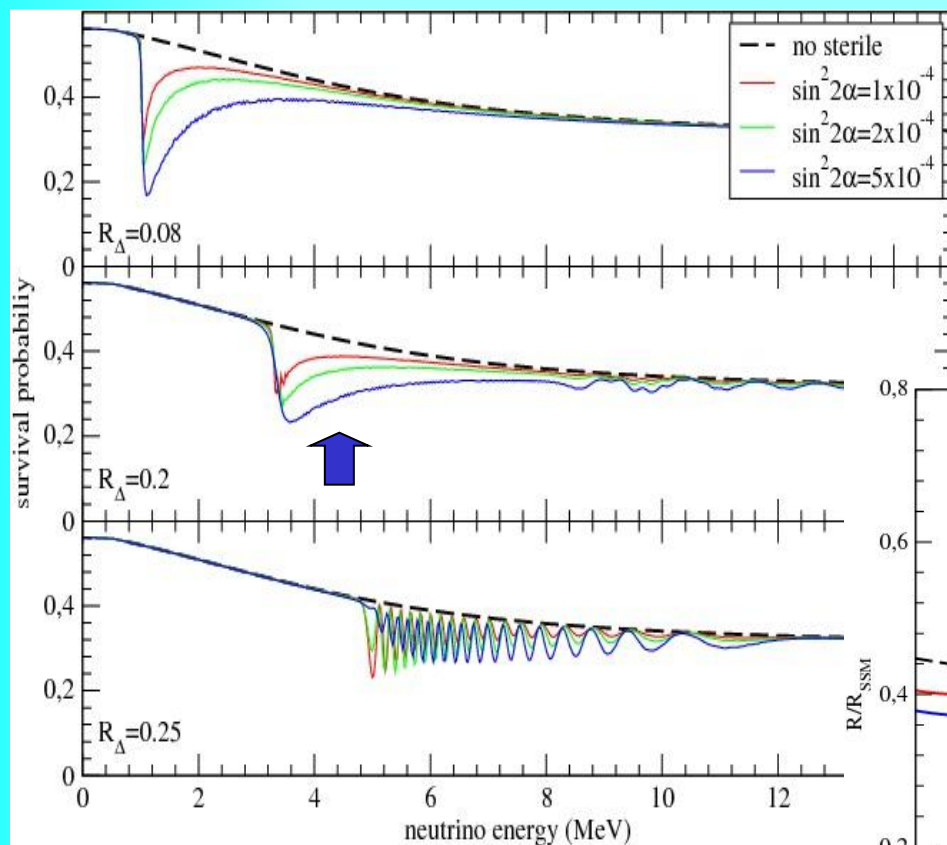
can be tested in atmospheric neutrinos with DC IceCube

$$\sin^2 2\alpha \sim 10^{-3}$$

$$\sin^2 2\beta \sim 10^{-1}$$

Survival probability

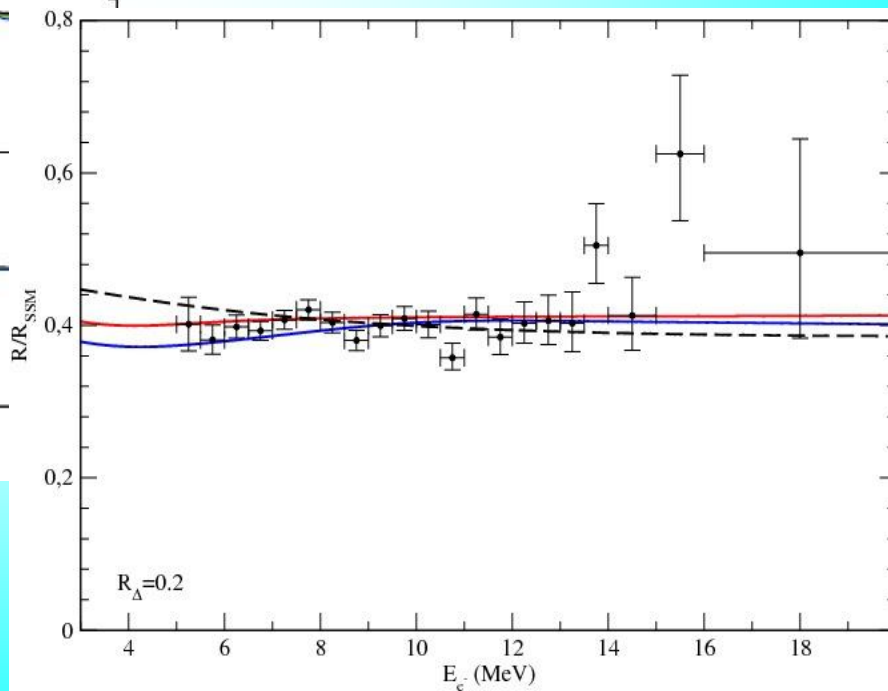
P. de Holanda,
AYS



$$m_0 \sim 0.003 \text{ eV}$$

$$m_0 = \frac{M^2}{M_{\text{Planck}}}$$

$$M \sim 2 - 3 \text{ TeV}$$



Day and Night

Accumulating data at SK SK I - IV

Day-Night effect: at 2.3σ - level in agreement with the LMA MSW solution

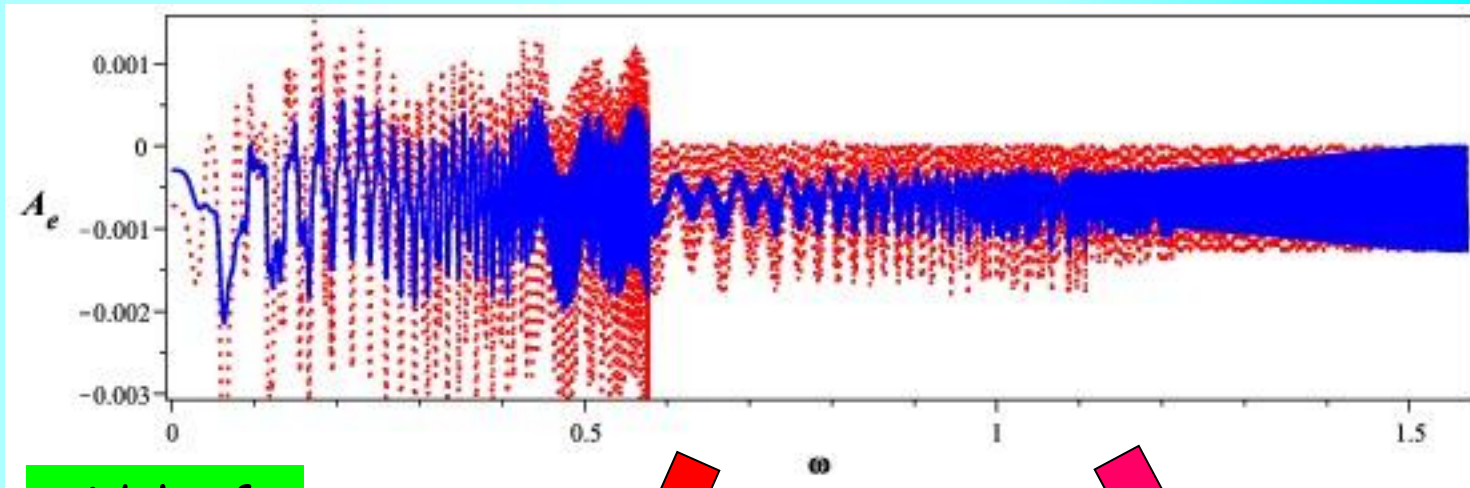
Provides overall test of the solution

New precision level - new possibilities: HyperKamiokande, LENA, MICA

LENA and Earth matter effect

Be neutrino line

A Ioanissian, AYS



Period of oscillations in energy scale

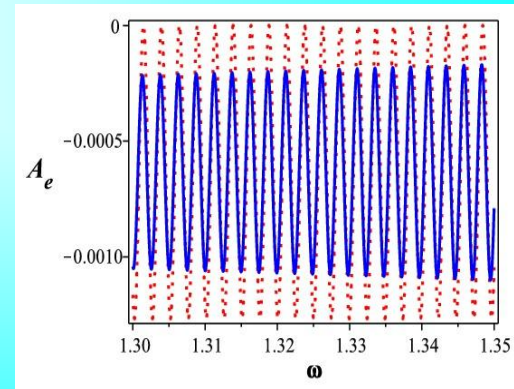
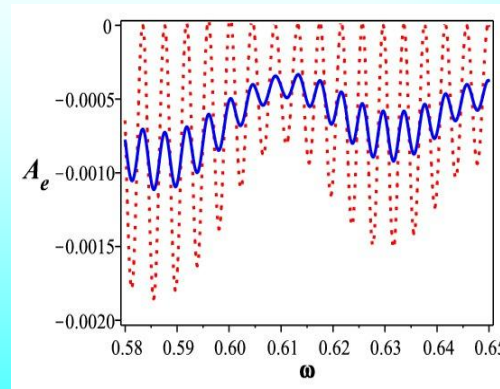
\sim

width of Beryllium nu line

Width of the Be nu line \rightarrow central temperature of the Sun

Precise measurements of Δm_{21}^2

Tomography of the Earth with resolution 20 km



Physics with HAND'S

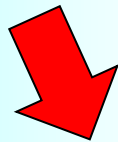
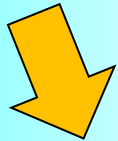
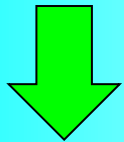
Huge
Atmospheric
Neutrinos
Detectors



or race for hierarchy

Race for hierarchy

Matter effect
on 1-3 mixing



Precise
measurements
of Δm^2
at reactors

Cosmology
 Σm

Atmospheric
neutrinos

LBL
experiments

Supernova
neutrinos

Double beta
decay m_{ee}

PINGU

NH \leftrightarrow IH
nu \leftrightarrow antinu

INO

NOvA
Neutrino beam
Fermilab-PINGU(W. Winter)

Earth matter
effect
Energy spectrs

Sterile neutrinos
may help?

Atmospheric neutrinos

Oscillation physics with Huge atmospheric neutrino detectors

P. Coyle

ANTARES

Oscillations 2.7σ

DeepCore

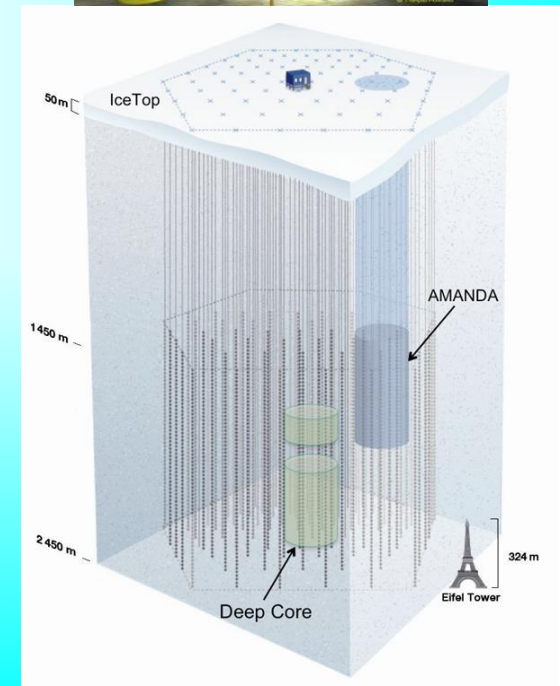
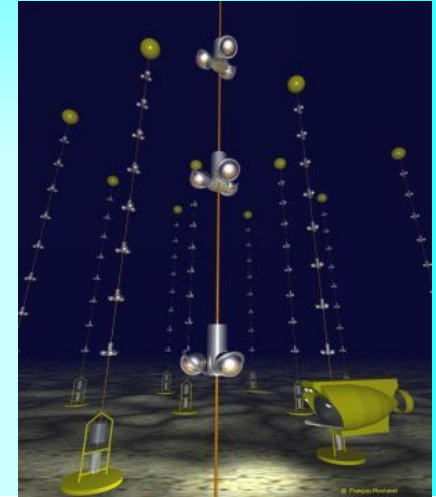
Oscillations at high energies 10 - 100 GeV in agreement with low energy data

Ice Cube

no oscillation effect at $E > 100$ GeV



Bounds on non-standard interaction, Lorentz violation etc



PINGU Geometry

Precision IceCube Next Generation Upgrade

Denser array

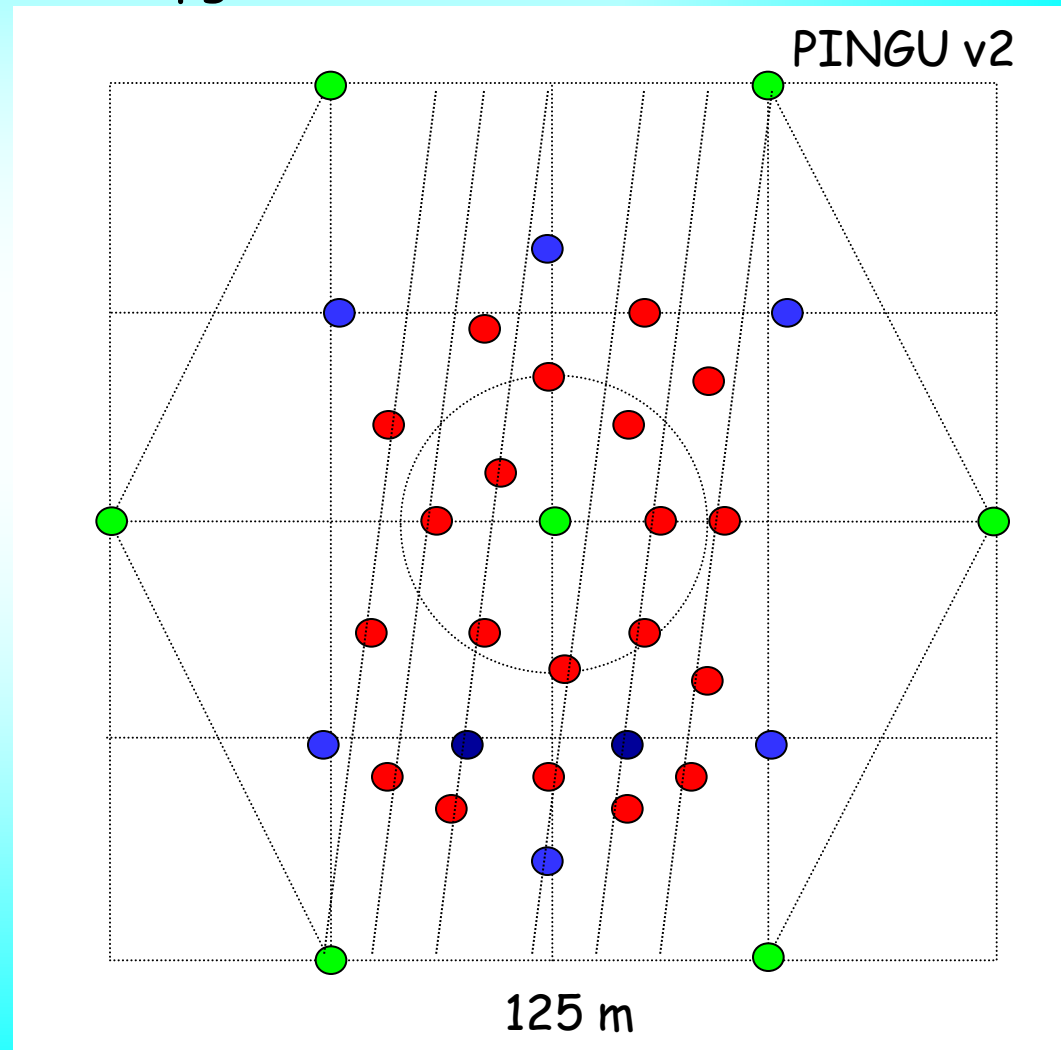
20 new strings (~60 DOMs each)
in 30 Mton DeepCore volume



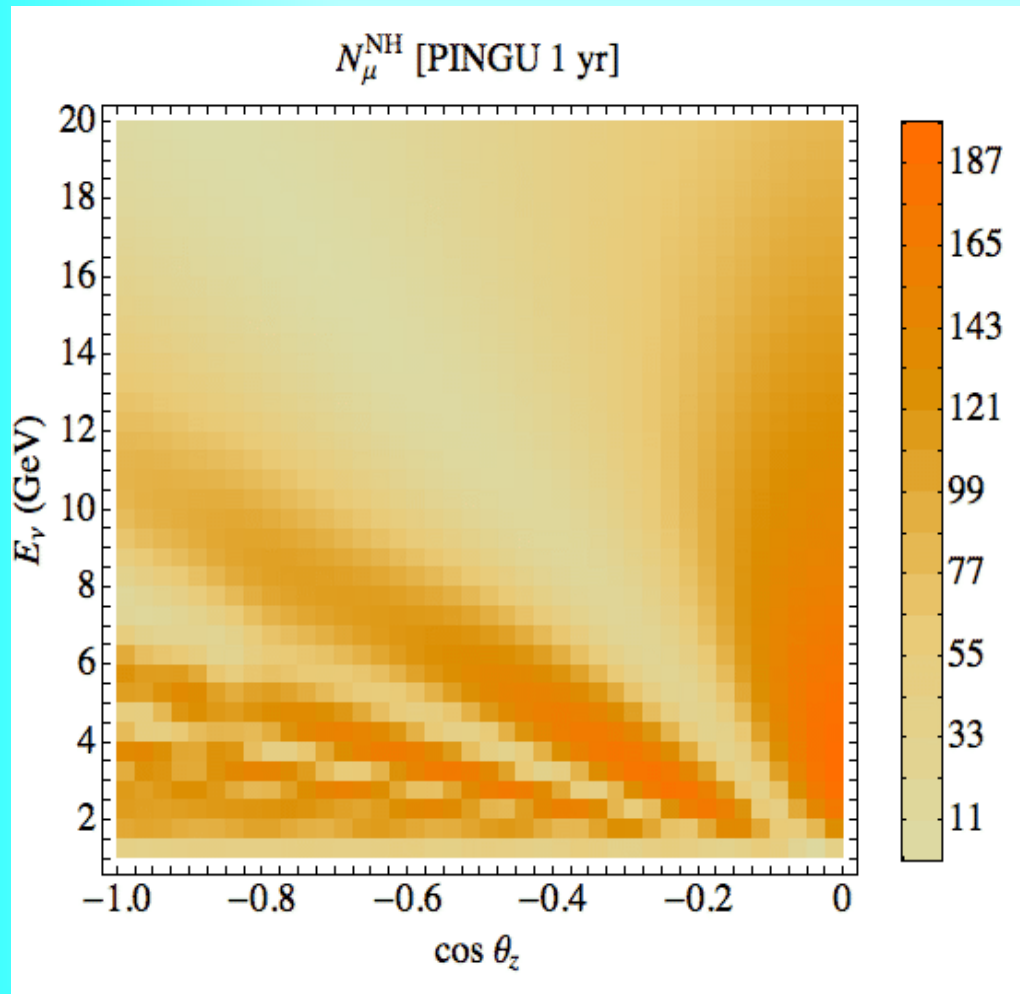
Few GeV threshold in inner
10 Mton volume

Energy resolution ~ 3 GeV

- Existing IceCube strings
- Existing DeepCore strings
- New PINGU-I strings

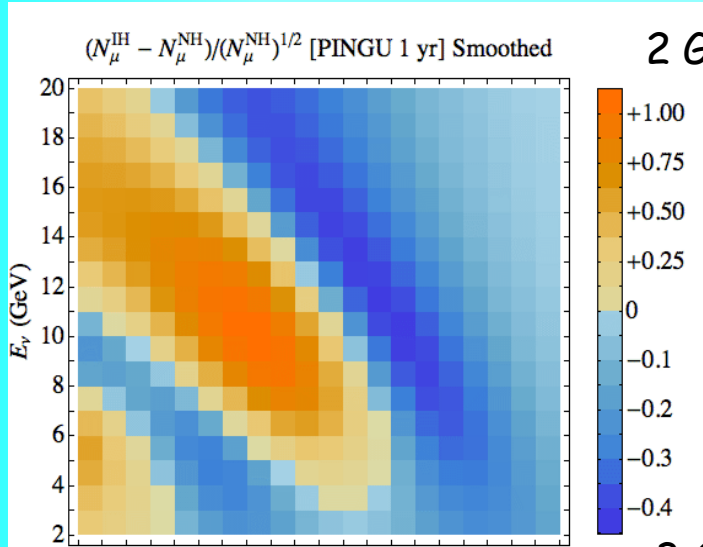


Muon neutrino events



High statistics can
cure other problems

PINGU and mass hierarchy

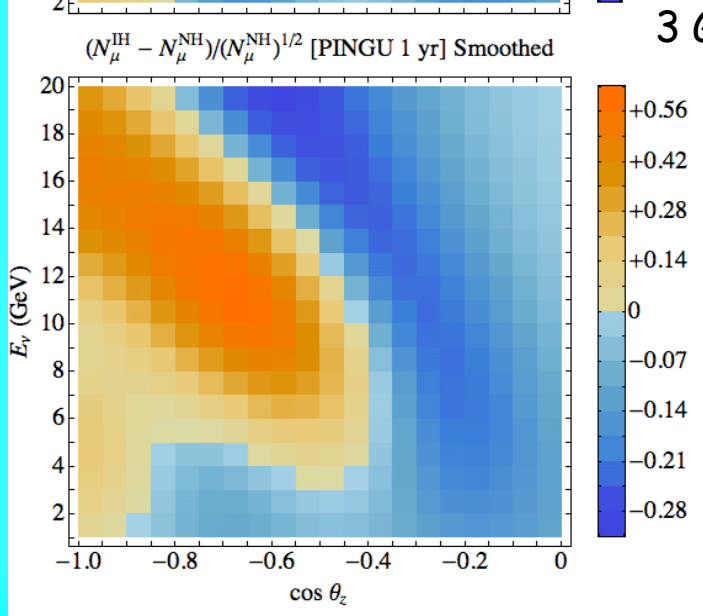


2 GeV, 11.25°

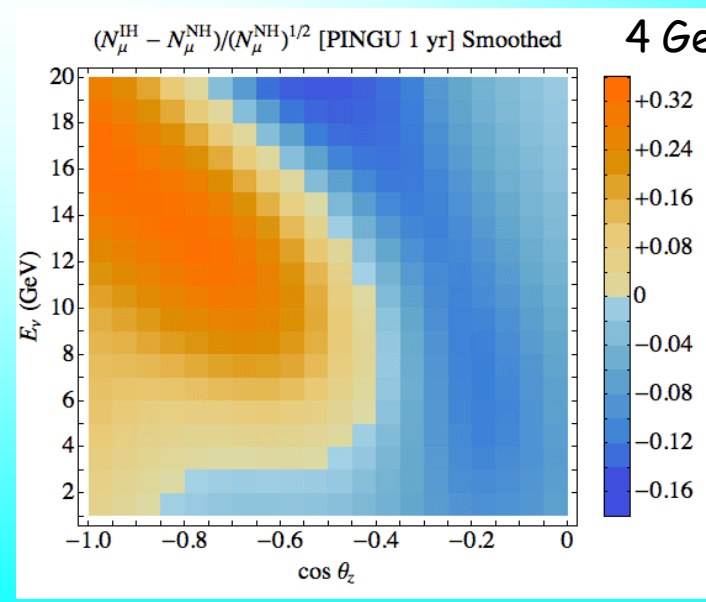
E. Akhmedov, S. Razzaque, A. Y. Smirnov
arXiv: 1205.7071

Smearing with Gaussian reconstruction functions characterized by (half) widths

$(\sigma_E, \sigma_{\theta})$

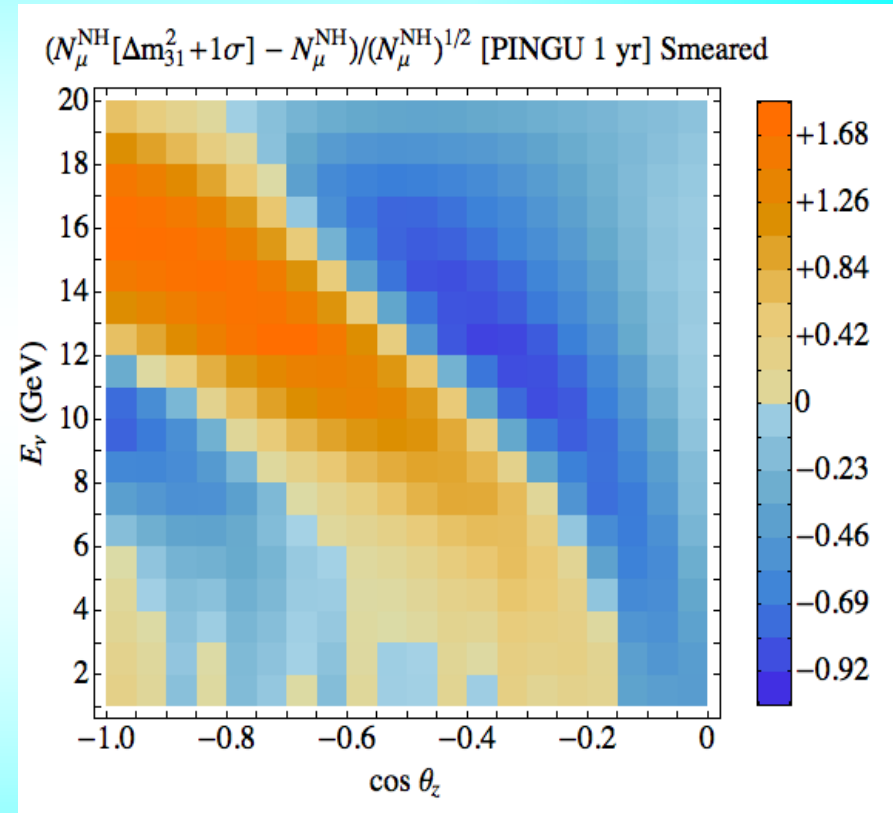
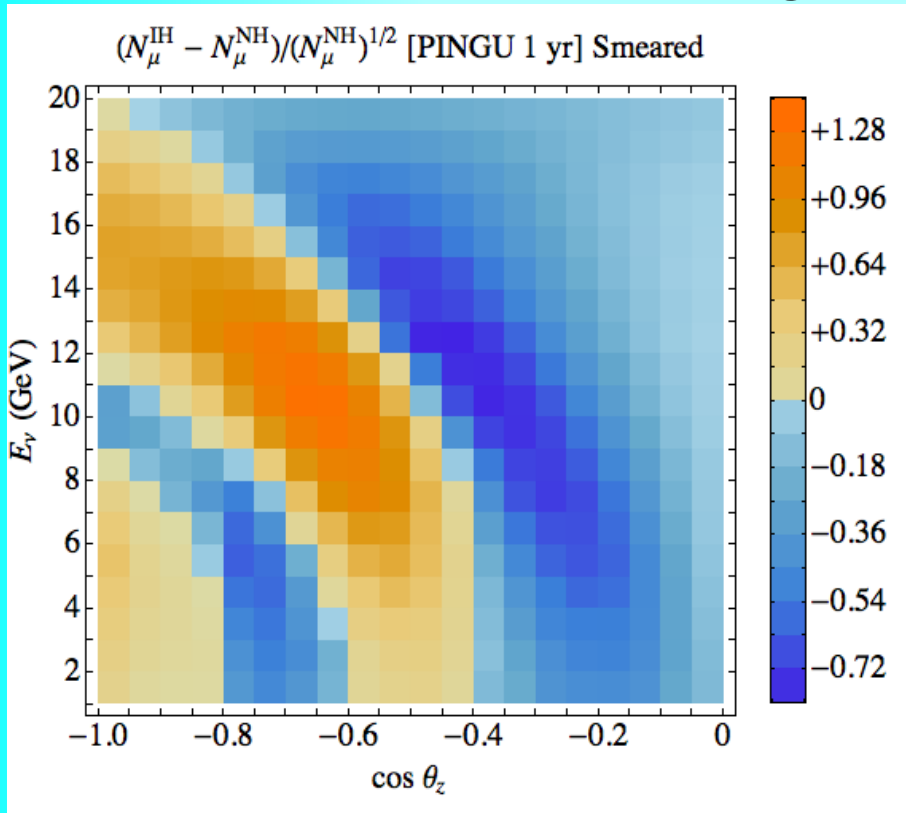


3 GeV, 15°



4 GeV, 22.5°

Hierarchy with PINGU

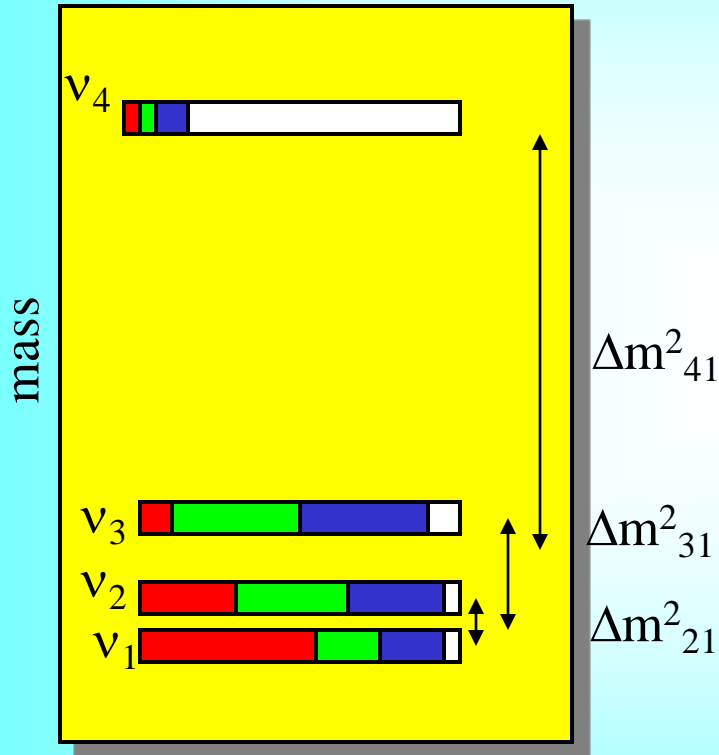
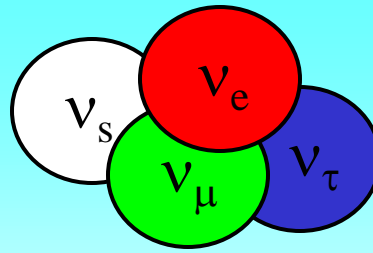


$$\sigma_E = 0.2E$$

$$\sigma_{\theta} \sim 1/E^{0.5}$$

Degeneracy

(3 + 1) scheme



LSND/MiniBooNE: vacuum oscillations

$$P \sim 4 |U_{e4}|^2 |U_{\mu 4}|^2$$

restricted by short baseline exp.
BUGEY, CHOOZ, CDHS, NOMAD

For reactor and source experiments

$$P \sim 4 |U_{e4}|^2 (1 - |U_{e4}|^2)$$

With new reactor data:

$$\Delta m_{41}^2 = 1.78 \text{ eV}^2 \quad (0.89 \text{ eV}^2)$$

$$U_{e4} = 0.15 \quad U_{\mu 4} = 0.23$$

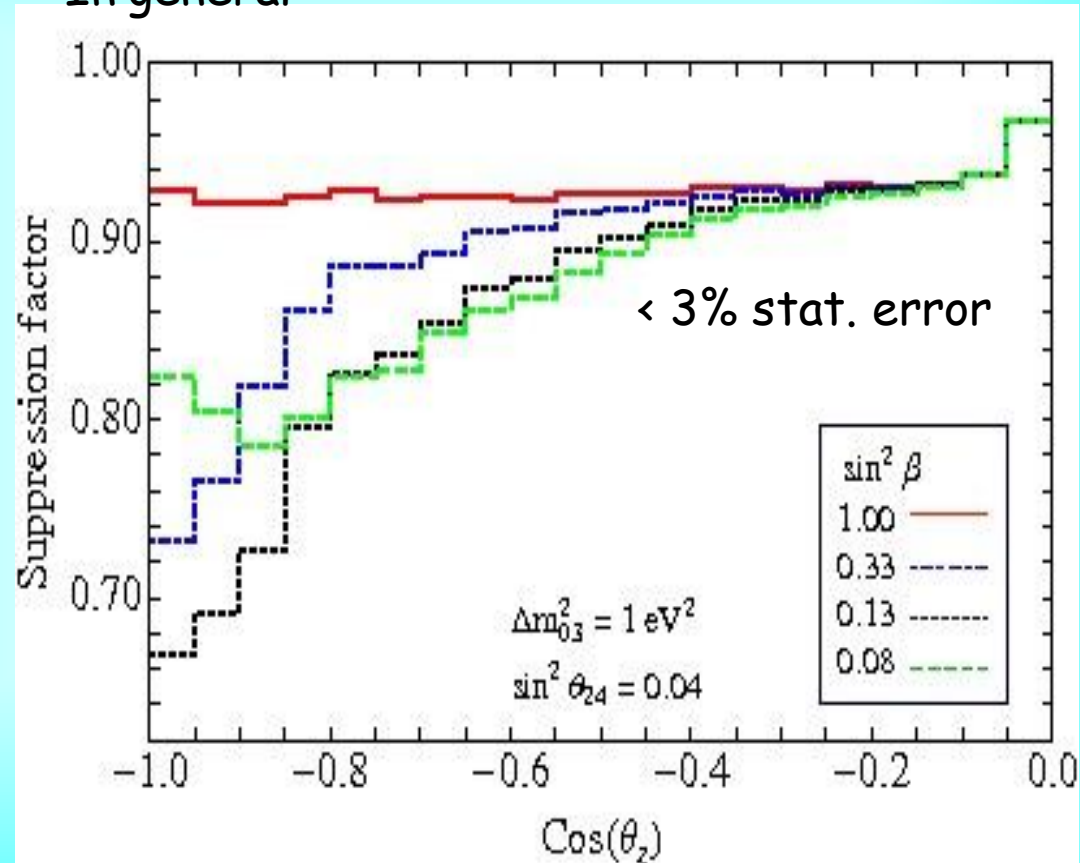
- additional radiation in the universe
- bound from LSS?

Zenith angle distributions

In general

For different
mixing schemes

Varying $|U_{\tau 0}|^2$



Zenith angle distribution depends on admixture of ν_τ in 4th mass state

without
conclusions

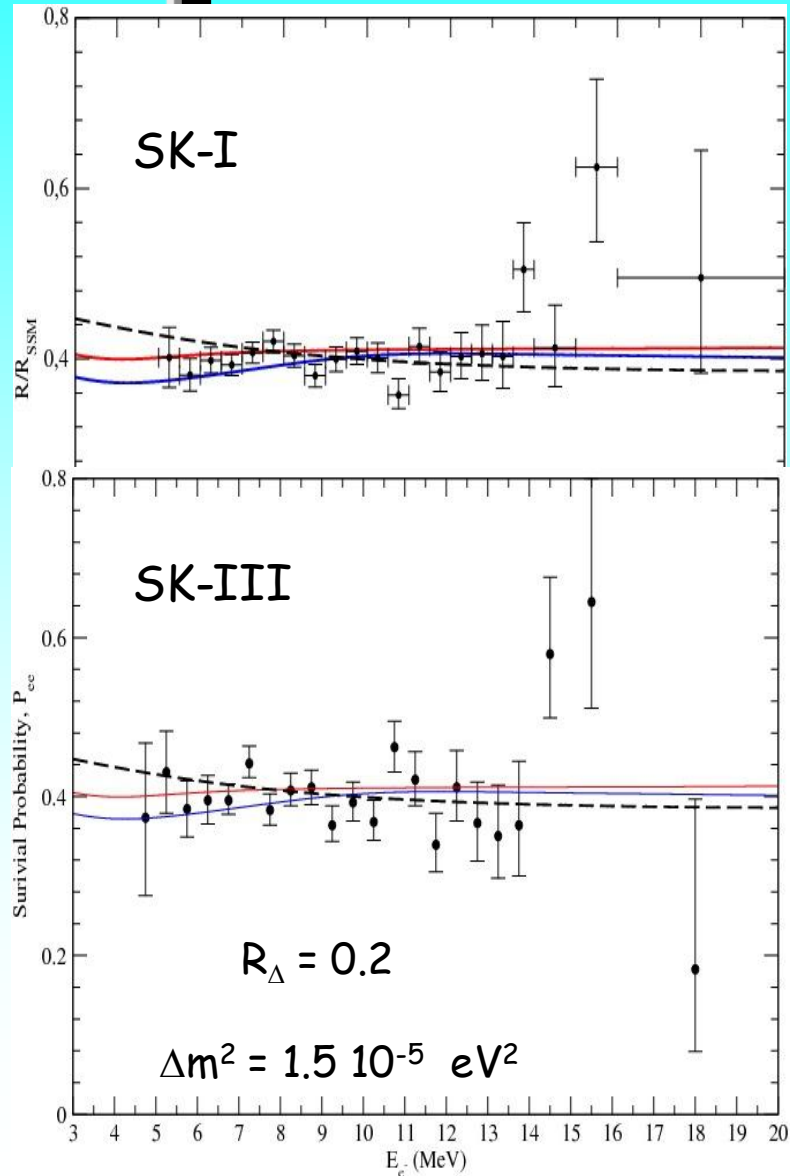
but



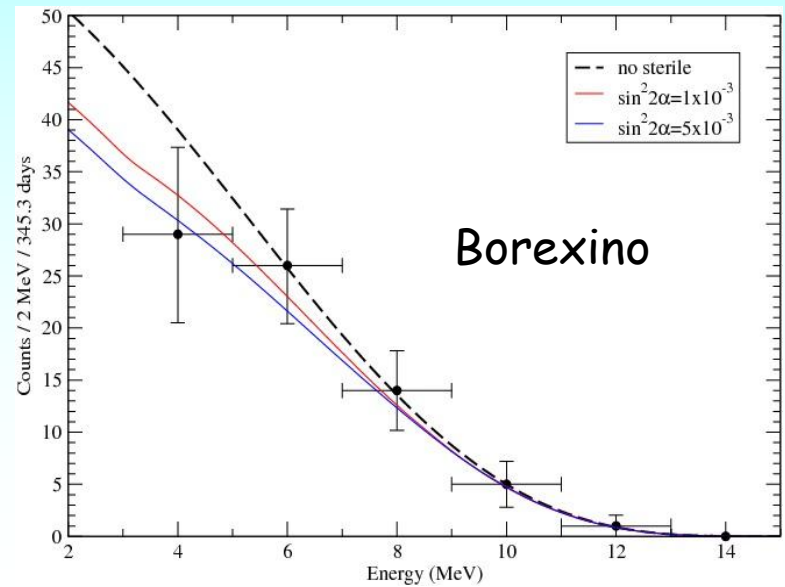
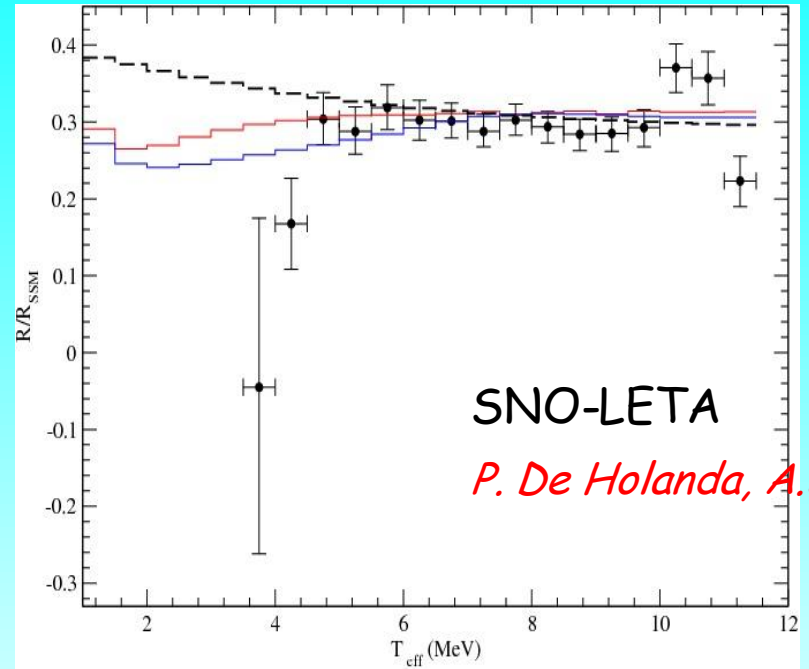
A scenic view of a river at sunset. The sun is low on the horizon, casting a warm glow over the scene. A bridge with multiple arches spans the river. Buildings line the banks, and a row of parked cars is visible on the left. The text "with many thanks to you!" is overlaid in a large, yellow, 3D-style font.

with many thanks to you!

Up-turns

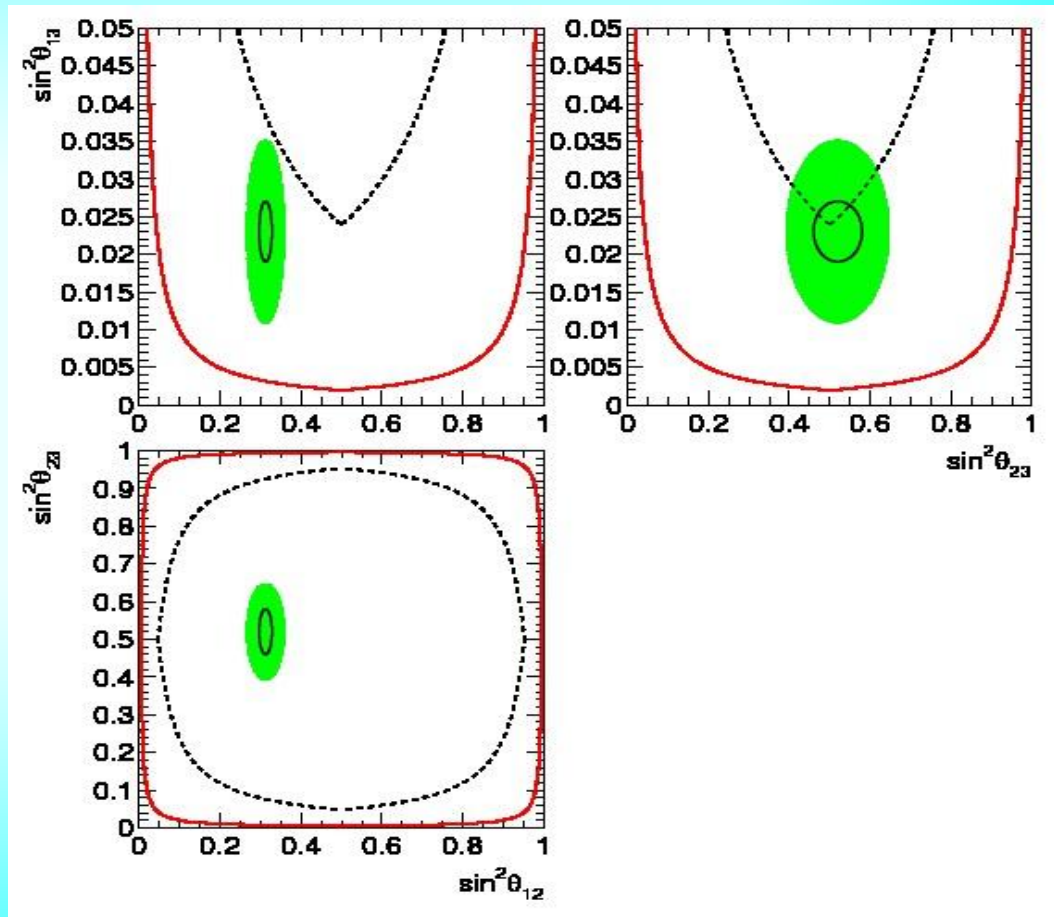


$\sin^2 2\alpha = 10^{-3}$ (red), $5 \cdot 10^{-3}$ (blue)



Anarchy

De Gouvea,
Murayama



Global view

from global fits

with salient probably features

smallness of mass

related

Peculiar (?) pattern of mixing

strongly differs from quark mixing

Standard
3 neutrino
framework

- Mass hierarchy (ordering)
- Deviation of 2-3 mixing from maximal
- CP violation
- Majorana nature
- Absolute scale

Usual ``hard'' masses

Generated at the electroweak and higher mass scales

challenge:

Sterile neutrinos

Not a small perturbation of the standard framework

Tri-bimaximal mixing

L. Wolfenstein

P. F. Harrison

D. H. Perkins

W. G. Scott

$$U_{\text{tbm}} = \begin{pmatrix} \sqrt{2/3} & \sqrt{1/3} & 0.15 \\ -\sqrt{1/6} & \sqrt{1/3} & -\sqrt{1/2} \\ -\sqrt{1/6} & \sqrt{1/3} & \sqrt{1/2} \end{pmatrix}$$

- maximal 2-3 mixing
- zero 1-3 mixing, no CP-violation

$$- \sin^2\theta_{12} = 1/3$$

v_3 is bi-maximally mixed
 v_2 is tri-maximally mixed

Mass matrix in flavor basis:

$$m_{\text{TBM}} = \begin{pmatrix} a & b & b \\ \dots & c & d \\ \dots & \dots & c \end{pmatrix}$$

Mass relations

$$m_{e\mu} = m_{e\tau}$$

$$m_{\mu\mu} = m_{\tau\tau}$$

$$m_{ee} + m_{e\mu} = m_{\mu\mu} + m_{\mu\tau}$$

Symmetry

Should be broken

SN neutrinos and large 1-3 mixing

Level crossing in the H-resonance is highly adiabatic

Strong suppression of the neutronization peak:

NH

$\nu_e \rightarrow \nu_3$

Permutations of flavor spectra which depend on mass hierarchy

Earth matter effects

Normal mass hierarchy:
in the antineutrino channel only

Inverted mass hierarchy:
in the neutrino channel only

If the earth matter effect is observed for antineutrinos NH is established!

Shock wave effect

Adiabaticity is broken in shock front if the relative width of the front:

$$\Delta R/R < 10^{-4} \rightarrow 10 \text{ km}$$

if larger - no shock wave effect:
probe of the width of front

Complication?

Collective effects

neutrinosphere
R = 20 - 50 km

usual matter potential:

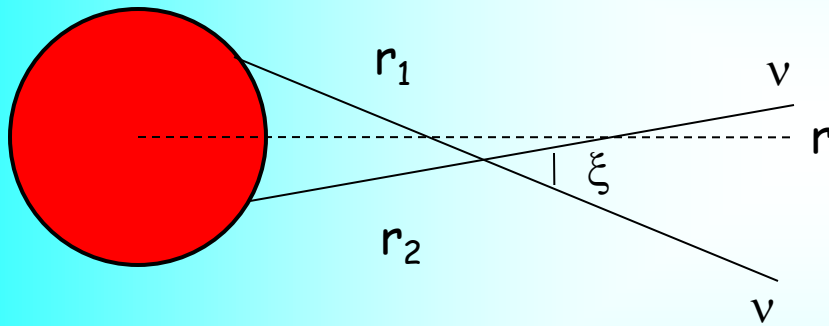
$$\lambda = V = \sqrt{2} G_F n_e$$

neutrino potential:

$$\mu = \sqrt{2} G_F (1 - \cos \xi) n_\nu$$

$$n_\nu \sim 1/r^2$$

$$\xi \sim 1/r \quad \text{for large } r$$



Multiple spectral splits -swaps

Multi-angle effect:

$$r_2 < r_1$$

$$\phi_2 < \phi_1$$

Different phases from different directions due to usual matter potential

$$n_\nu \sim 10^{33} \text{ cm}^{-3}$$

$$n_e \sim 10^{35} \text{ cm}^{-3}$$



$$\lambda \gg \mu$$

decoherence

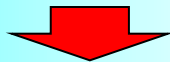
QLC predictions

H. Minakata, A.S.

Leptons

$$U_\nu = U_{\text{bm}} \text{ seesaw}$$

$$U_l = U_{\text{CKM}} \quad \text{q-l symmetry}$$

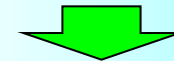


$$U_{\text{PMNS}} = U_l^\dagger U_\nu = U_{\text{CKM}}^\dagger U_{\text{bm}}$$

Quarks

$$V_u = I$$

$$V_d = V_{\text{CKM}}$$



$$V_{\text{quarks}} = V_u^\dagger V_d = V_{\text{CKM}}$$

1-3 mixing is generated by permutation of U_{12} and U_{23}

$$\sin\theta_{13} = \sin\theta_{23} \sin\theta_C \sim 0.16$$

$$\sin\theta_{12} = \sin(\pi/4 - \theta_C) + 0.5 \sin\theta_C (\sqrt{2} - 1 - V_{cb} \cos \delta) \rightarrow$$

$$D_{23} = 0.5 \sin^2\theta_C + \cos^2\theta_C V_{cb} \cos \delta = 0.02 \pm 0.04$$

$$\sin^2\theta_{12} = 0.3345$$

RGE \rightarrow can reduce
M. Schmidt, A.S.

Symmetry and mixing

If G is von Dyck group $D(2, m, p)$

$$D(2,3,3) = A_4$$

D. Hernandez, A.S.

$$D(2,3,4) = S_4$$

$$D(2,3,5) = A_5$$

the mixing matrix should satisfy condition

$$(S_i U_{PMNS}^\dagger T U_{PMNS})^p = I$$

$$i = 1, 2, 3$$

S_i is the symmetry transformation of the neutrino mass matrix in mass basis

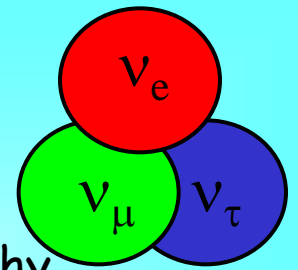
$$S_1 = \text{diag}(1, -1, -1) \quad S_2 = \text{diag}(-1, 1, -1) \quad S_i^2 = I$$

T is the symmetry transformation of the charged lepton mass matrix in mass basis

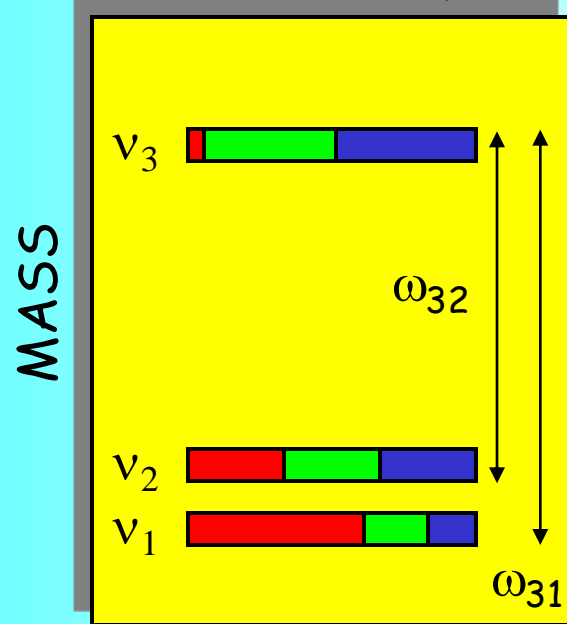
$$T = \text{diag}(e^{i\phi_1}, e^{i\phi_2}, e^{i\phi_3}) \quad \phi_i = 2\pi k_i / m$$

$$T^m = I$$

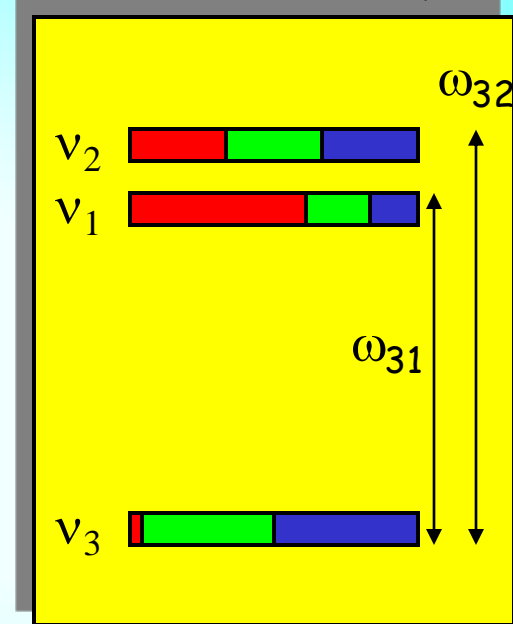
Mass hierarchy (ordering)



Normal hierarchy



Inverted hierarchy



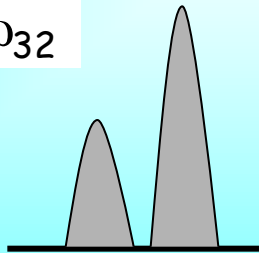
Cosmology

$\beta\beta$ -decay

$$\omega_{ij} = \Delta m^2_{ij} / 2E$$

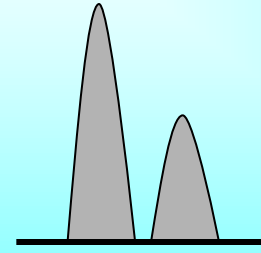
Mass states can be marked by ν_e - admixtures

$$\omega_{31} > \omega_{32}$$



Fourier analysis

$$\omega_{31} < \omega_{32}$$



ω

Oscillations

$$D_{31} \sim 2D_{32}$$

Matter effect

makes the e-flavor heavier \rightarrow
changes two spectra differently

S. Petcov
M. Piai