

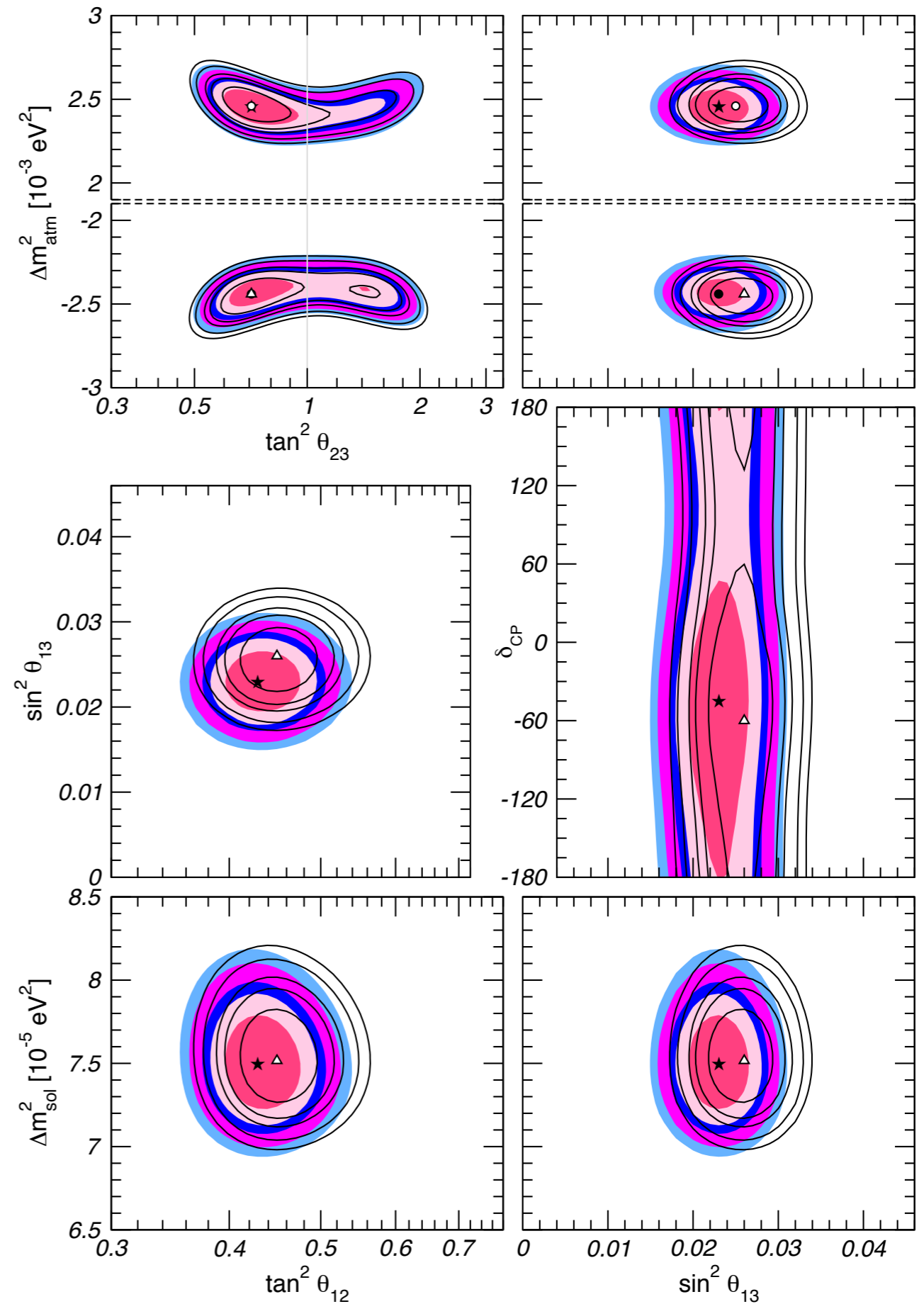
# *Review on global fits*

*What's  $\nu$  - Invisibles12, GGI, Florence, 25 June 2012*

*Thomas Schwetz*



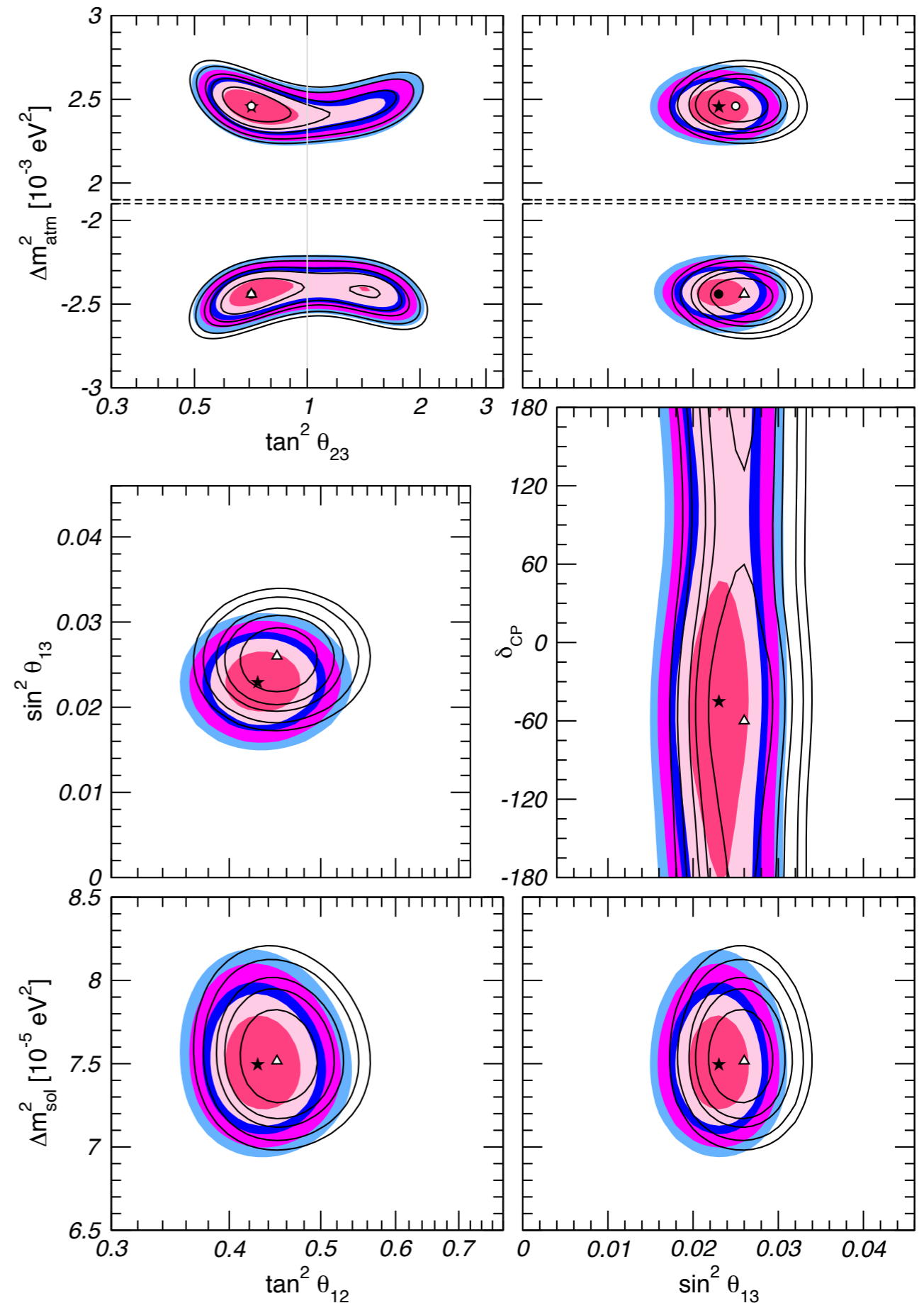
preliminary results from  
work in collab. with  
**C. Gonzalez-Garcia,**  
**M. Maltoni, J. Salvado**  
including latest data  
from Neutrino2012



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 work in collab. with  
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 including latest data  
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*Outline:*

- *the  $\theta_{13}$  revolution*
- *non-maximality of  $\theta_{23}$*
- *CP phase and hierarchy*



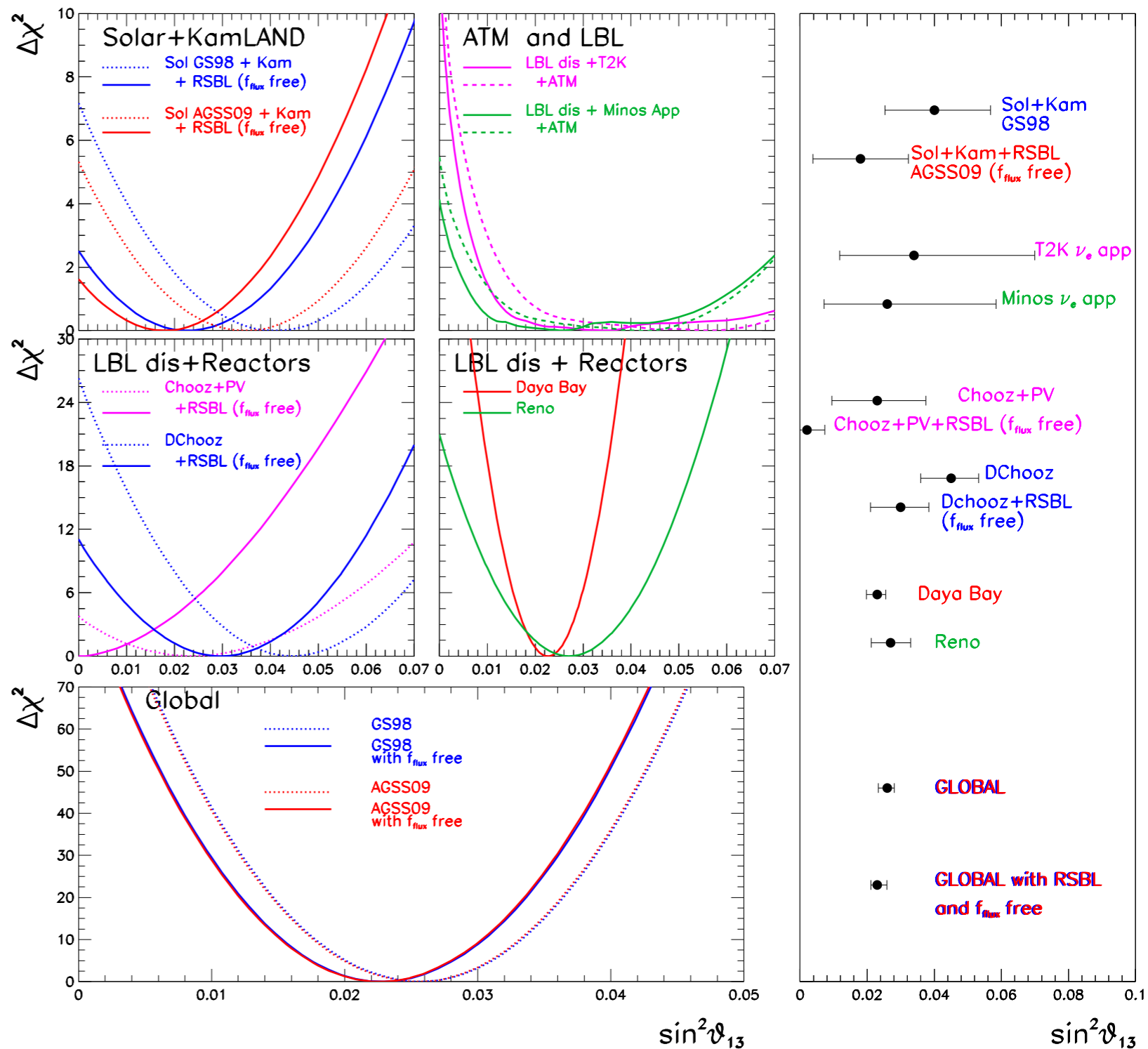
# The $\theta_{13}$ revolution

- About 1 year ago:  
6 events in T2K:  $2.5\sigma$
- global fits gave  
 $>3\sigma$  for the first  
time

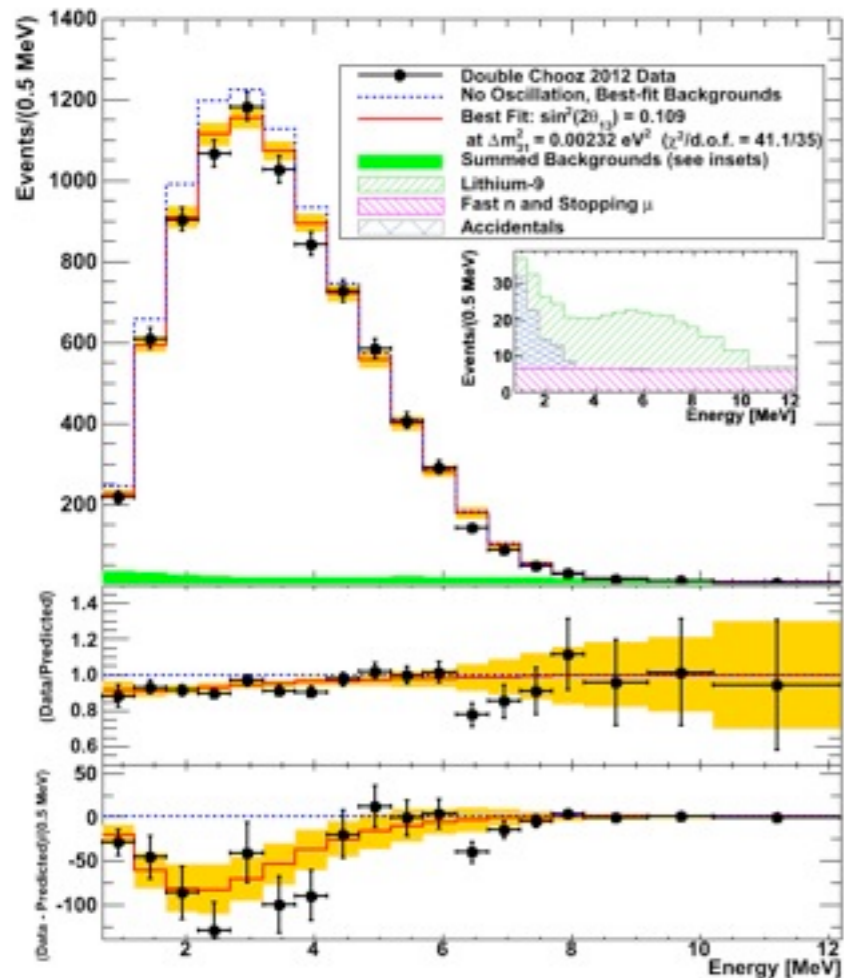
Fogli et al, *I 106.6028*  
TS, Tortola, Valle *I 108.1376*

- DoubleChooz,  
DayaBay, RENO

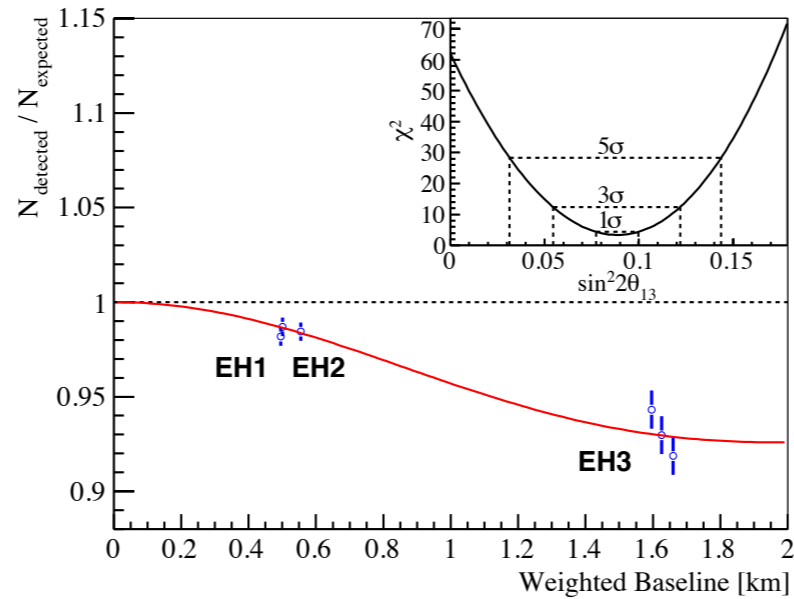
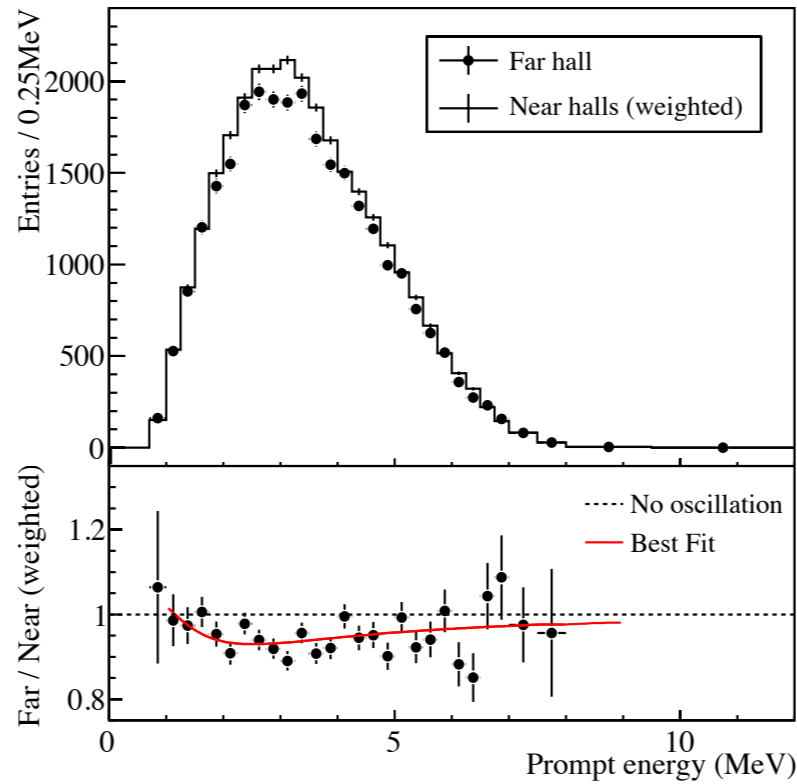
- post-Neutrino 12:  
 $\Delta\chi^2 \approx 100$  in the  
global fit



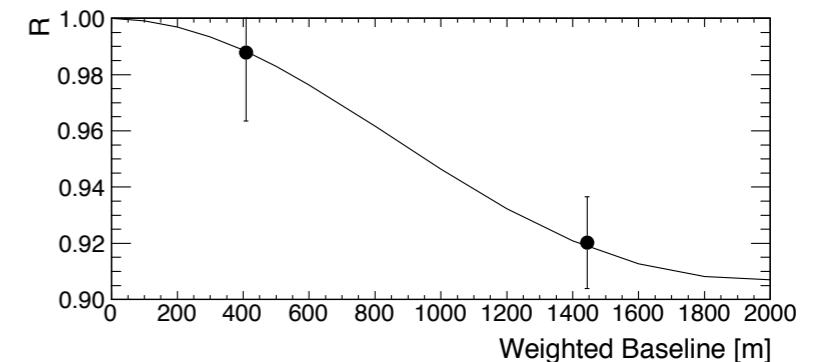
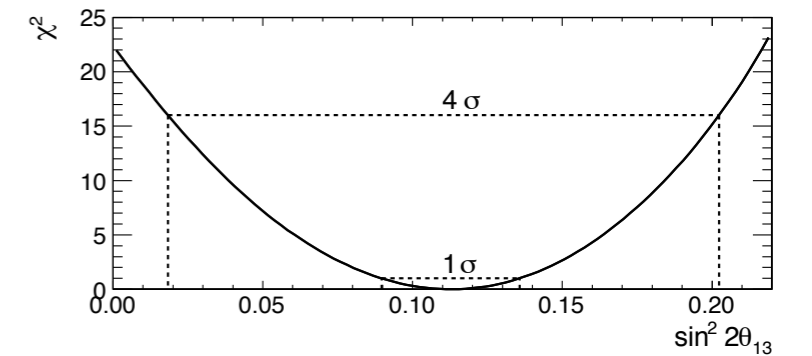
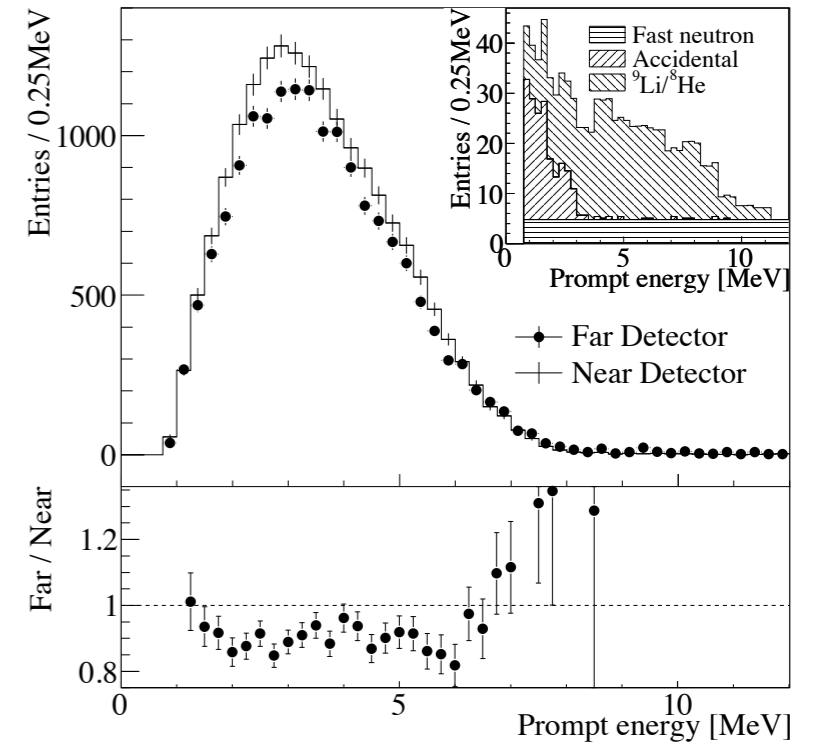
# The $\theta_{13}$ revolution - the reactors



DoubleChooz @ Neutrino2012



DayaBay @ Neutrino2012



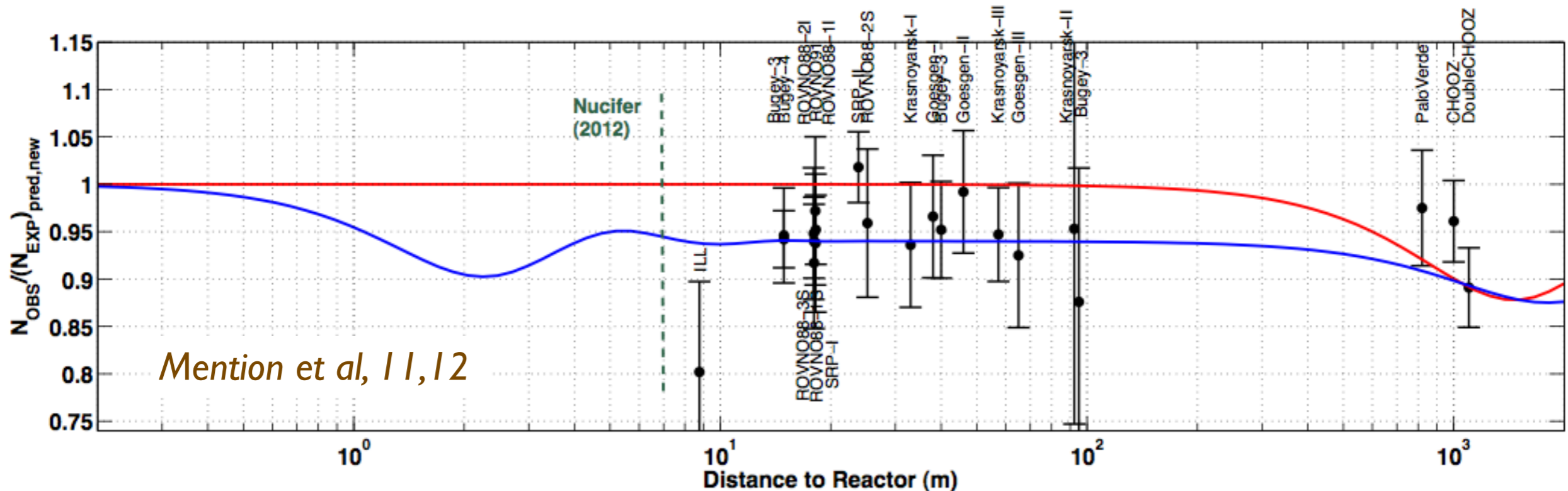
RENO, 1204.0626

# *The reactor anomaly*

- ▶ to predict the  $\bar{\nu}_e$  flux from nuclear reactors one has to convert the measured  $e^-$  spectra from  $^{235}\text{U}$ ,  $^{239}\text{Pu}$ ,  $^{241}\text{Pu}$  into neutrino spectra  
*Schreckenbach et al., 82, 85, 89*
- ▶ recent improved calculation *Mueller et al., 1101.2663*  $\sim 3\%$  higher fluxes (ab initio calculations + virtual branches for missing part)
- ▶ confirmed by independent calculation *P. Huber, 1106.0687* (virtual branches)
- ▶ increase of predicted number of neutrino-induced events compared to old flux calculations:

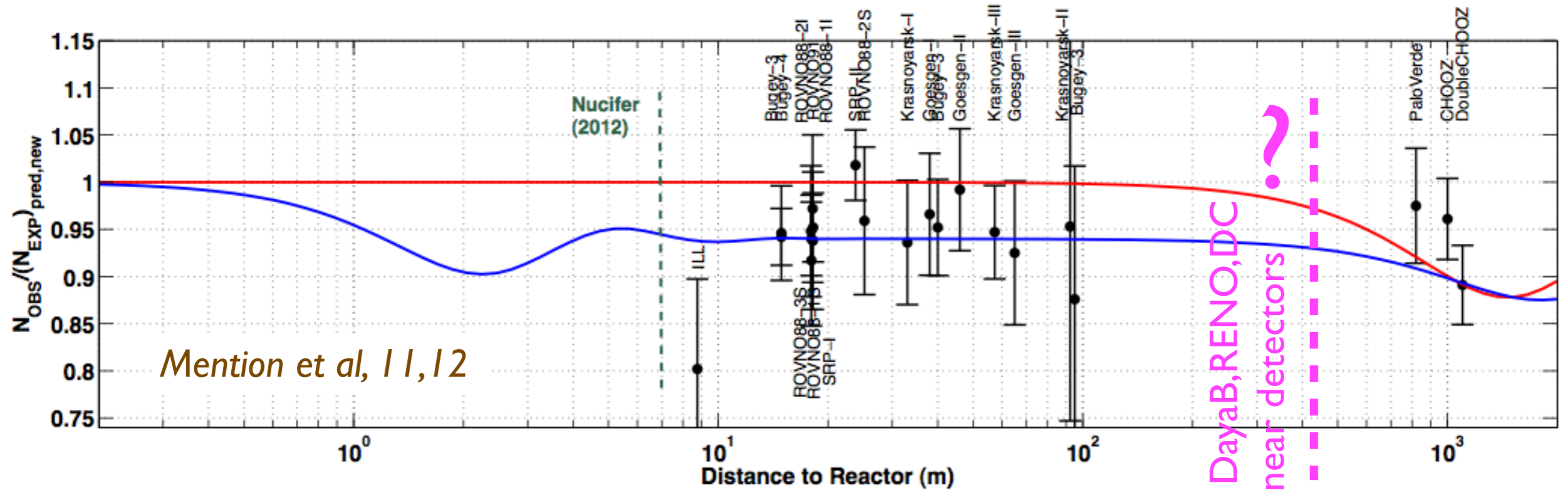
$^{235}\text{U}$	$^{239}\text{Pu}$	$^{241}\text{Pu}$	$^{238}\text{U}$
3.7%	4.2%	4.7%	9.8%

# The reactor anomaly



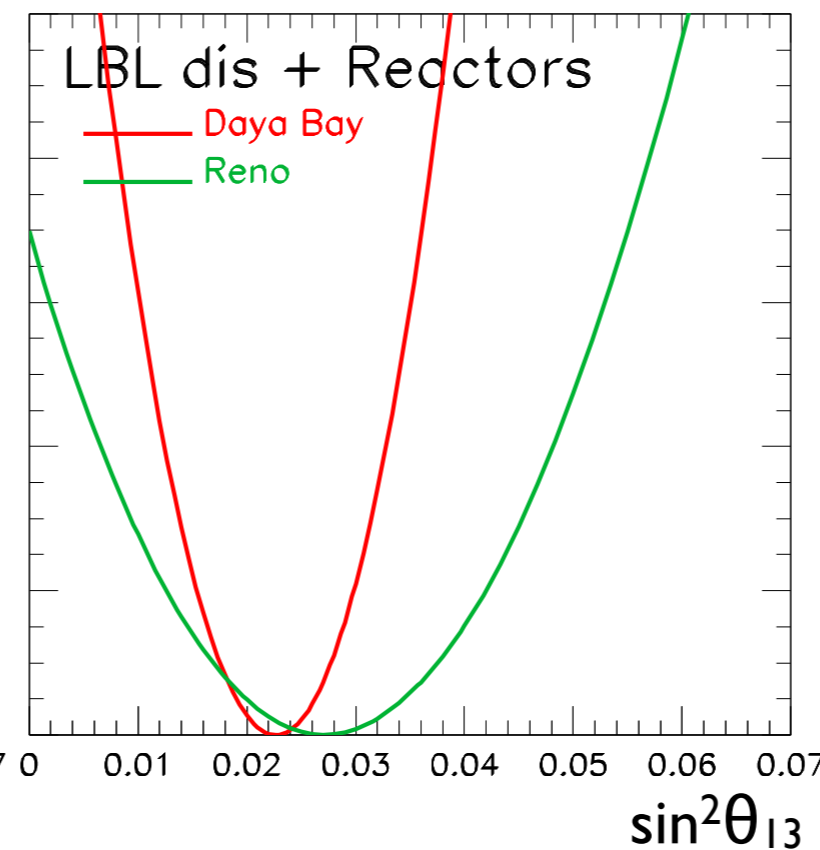
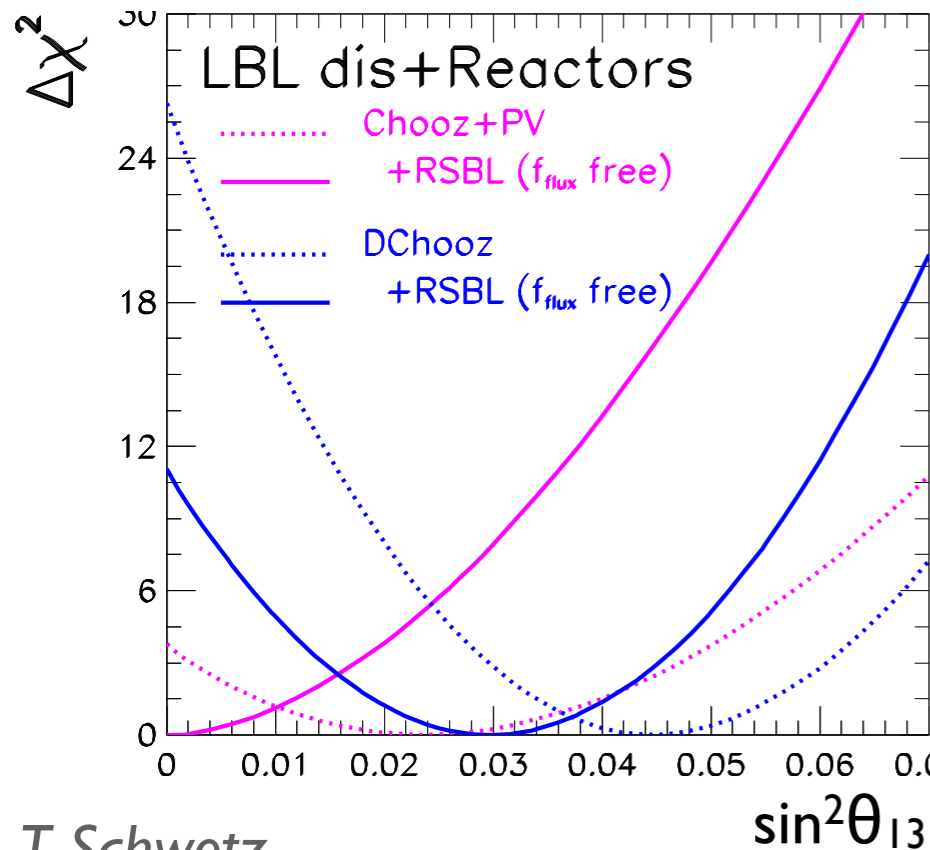
- SBL reactor data ( $L < 100\text{m}$ ) in tension with predicted flux  $f = 0.935 \pm 0.024$  (different from 1 @  $2.7\sigma$ )
- systematics?
  - ▶ normalization of ILL electron spectra
  - ▶ neutron lifetime (use 2012 PDG value)
- sterile neutrinos at the eV scale? *talk by Maltoni*

# The reactor anomaly and the $\theta_{13}$ determination



Mention et al, 11,12

DayaB, RENO, DC  
near detectors

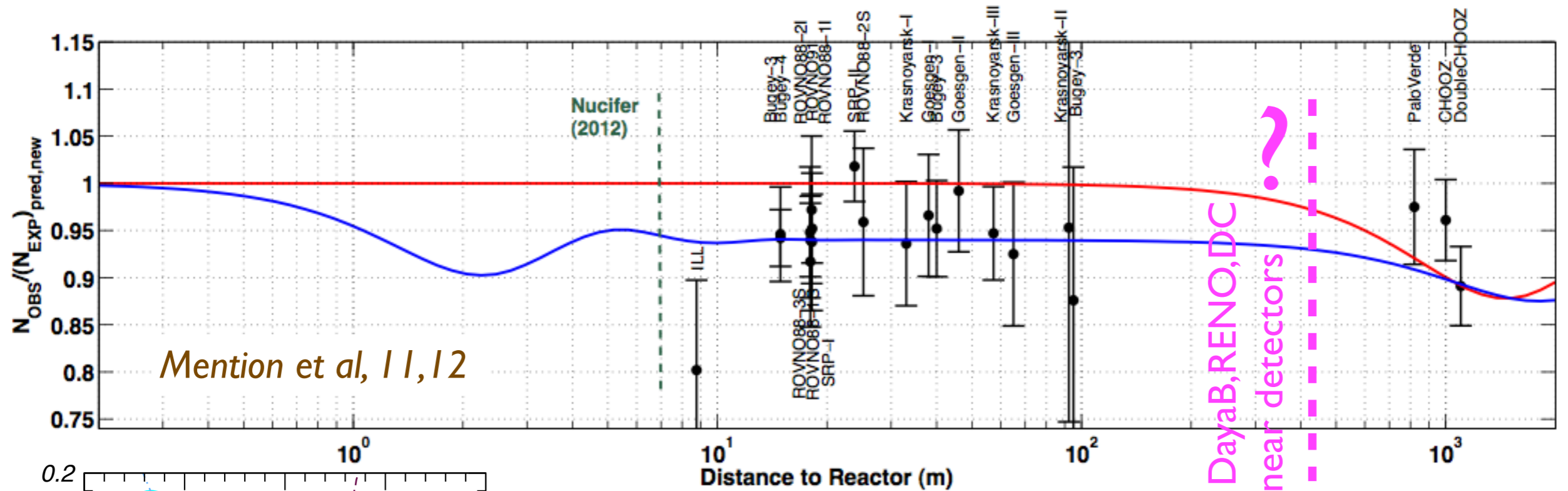


two extreme assumptions:

- use fluxes from *Huber, 1106.0687* without SBL reactor data
- leave reactor flux free and include SBL data in fit

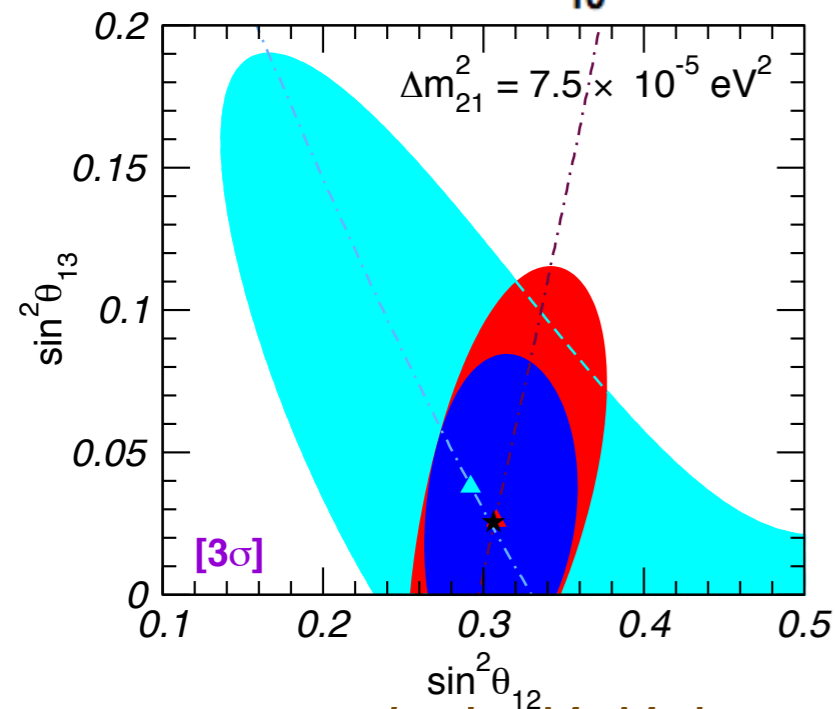


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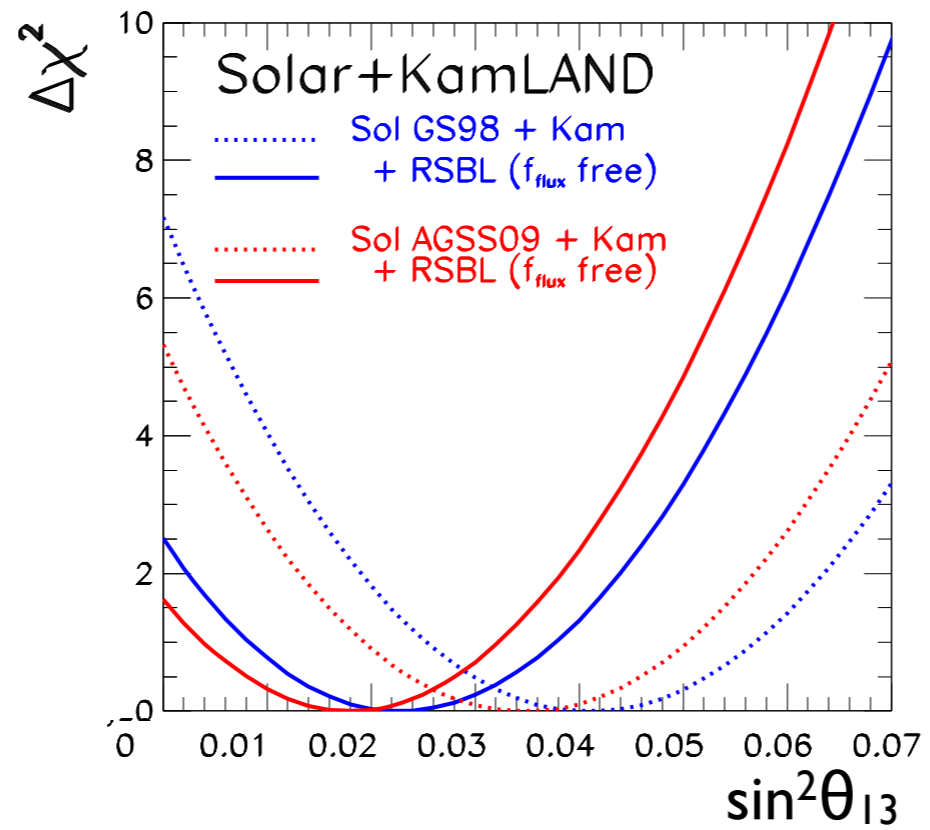


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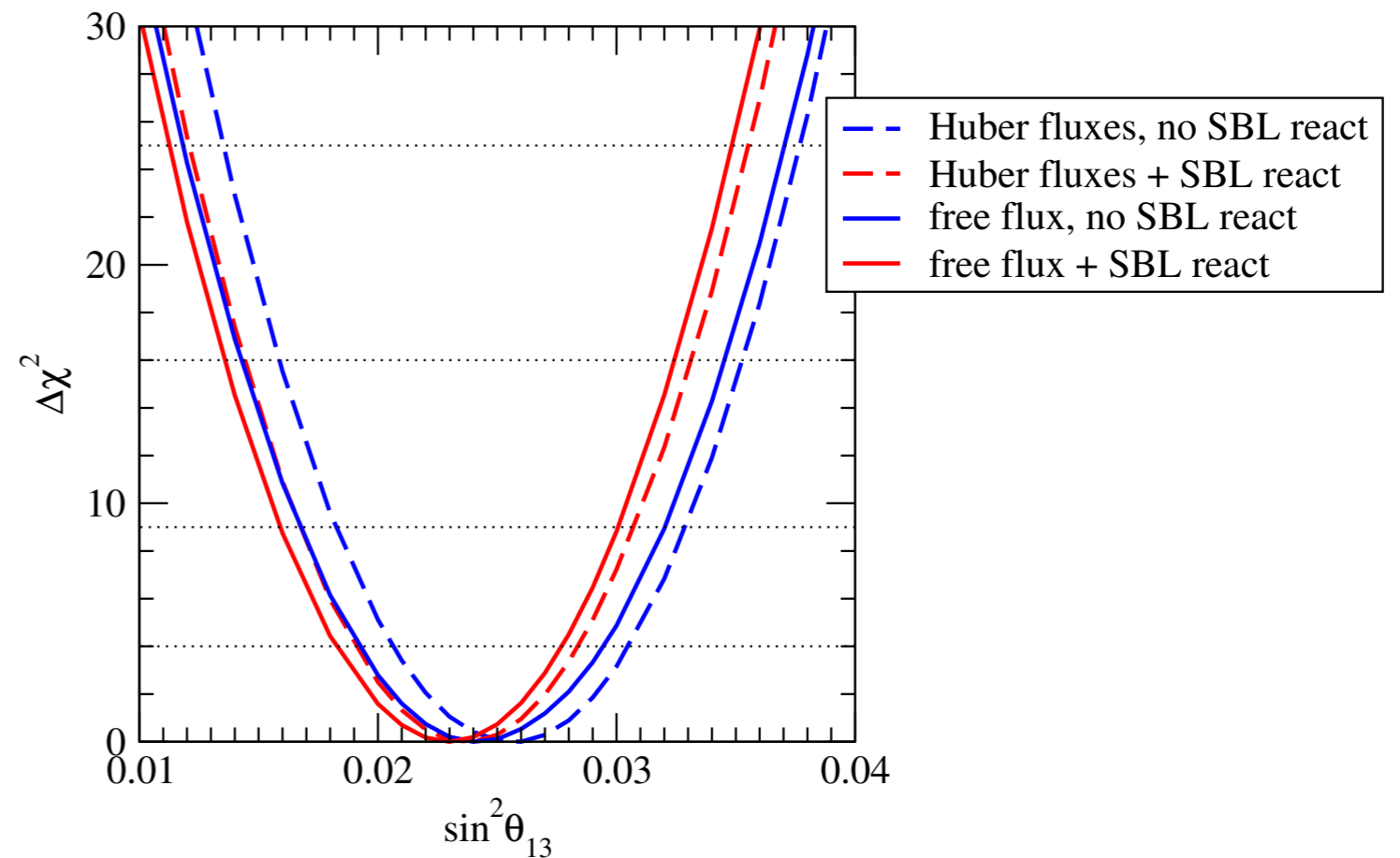
plot by M. Maltoni;  
Fogli et al, 08;  
TS, Tortola, Valle, 08



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# *The reactor anomaly and the $\theta_{13}$ determination*



*two extreme assumptions:*

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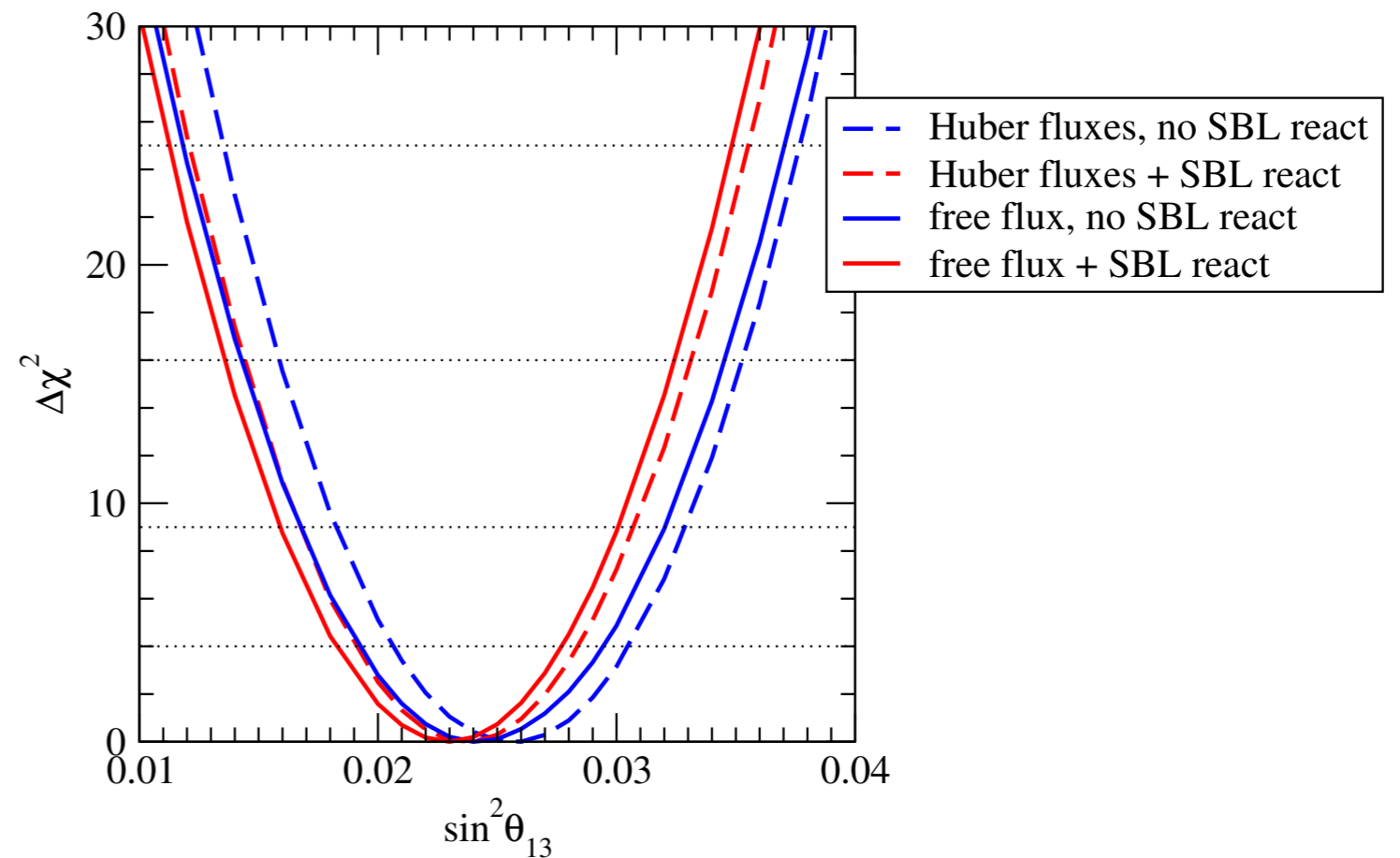
$$\sin^2 \theta_{13} = 0.0257 \pm 0.0025, \quad \theta_{13} = (9.2 \pm 0.46)^\circ, \quad \sin^2 2\theta_{13} = 0.100 \pm 0.0095$$

- *leave reactor flux free and include SBL data in fit*

$$\sin^2 \theta_{13} = 0.0230 \pm 0.0023, \quad \theta_{13} = (8.7 \pm 0.44)^\circ, \quad \sin^2 2\theta_{13} = 0.090 \pm 0.0090$$

# The reactor anomaly and the $\theta_{13}$ determination

- result depends on data which in principle is not sensitive to  $\theta_{13}$
- shift at the level of  $1\sigma$



two extreme assumptions:

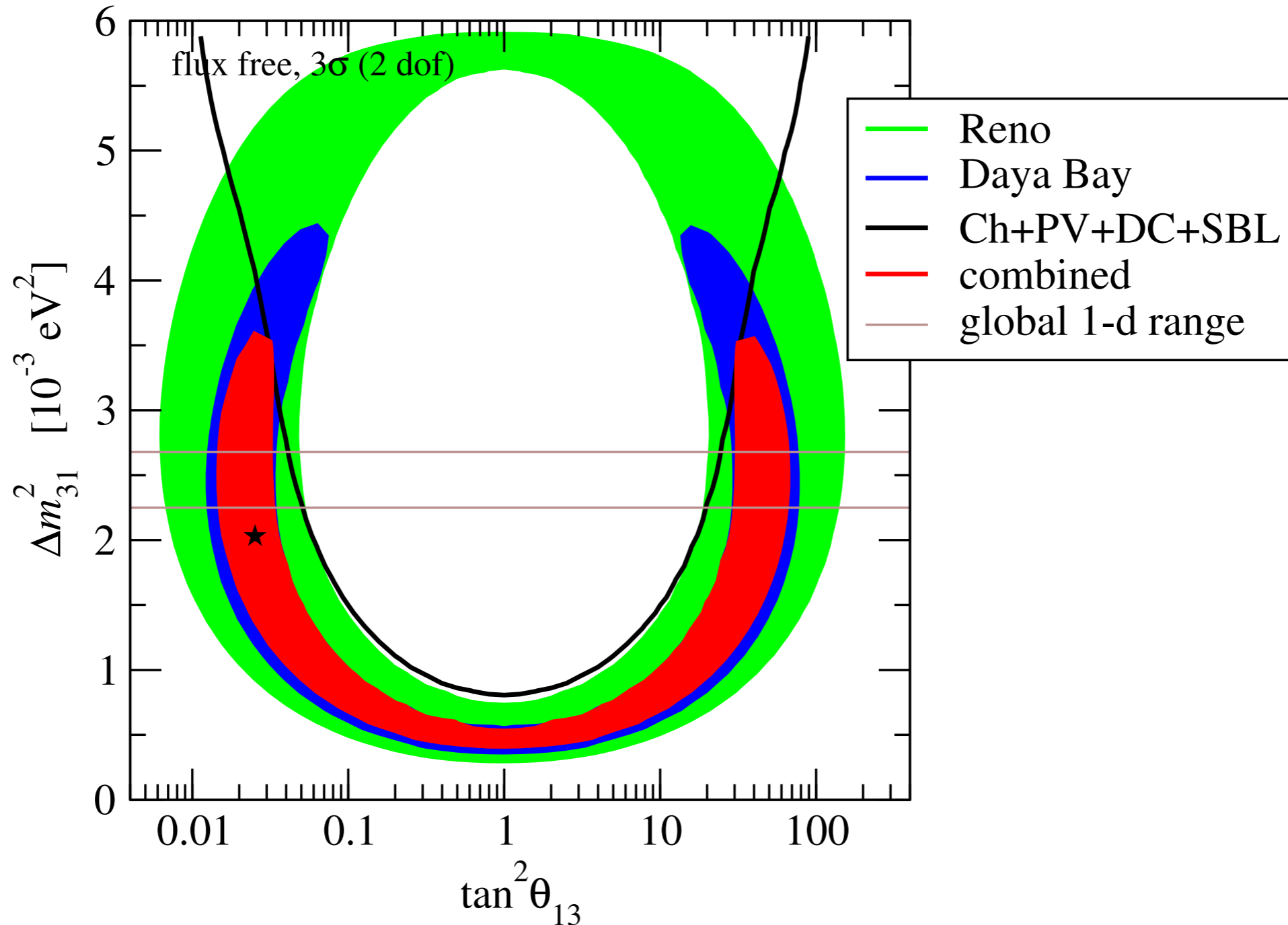
- use fluxes from *Huber, 1106.0687* without SBL reactor data

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- leave reactor flux free and include SBL data in fit

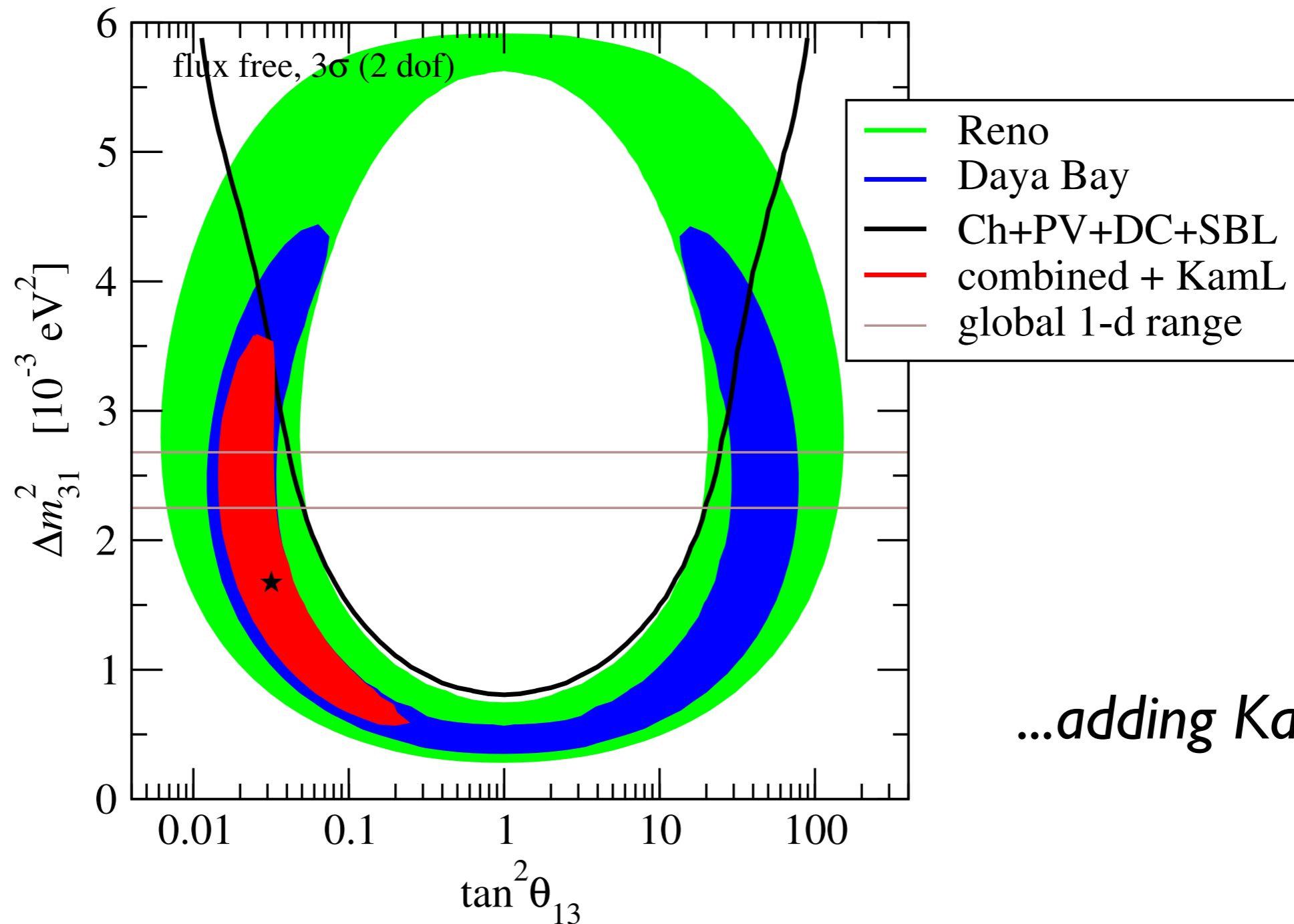
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# *Measuring $\Delta m^2_{31}$ with reactors*



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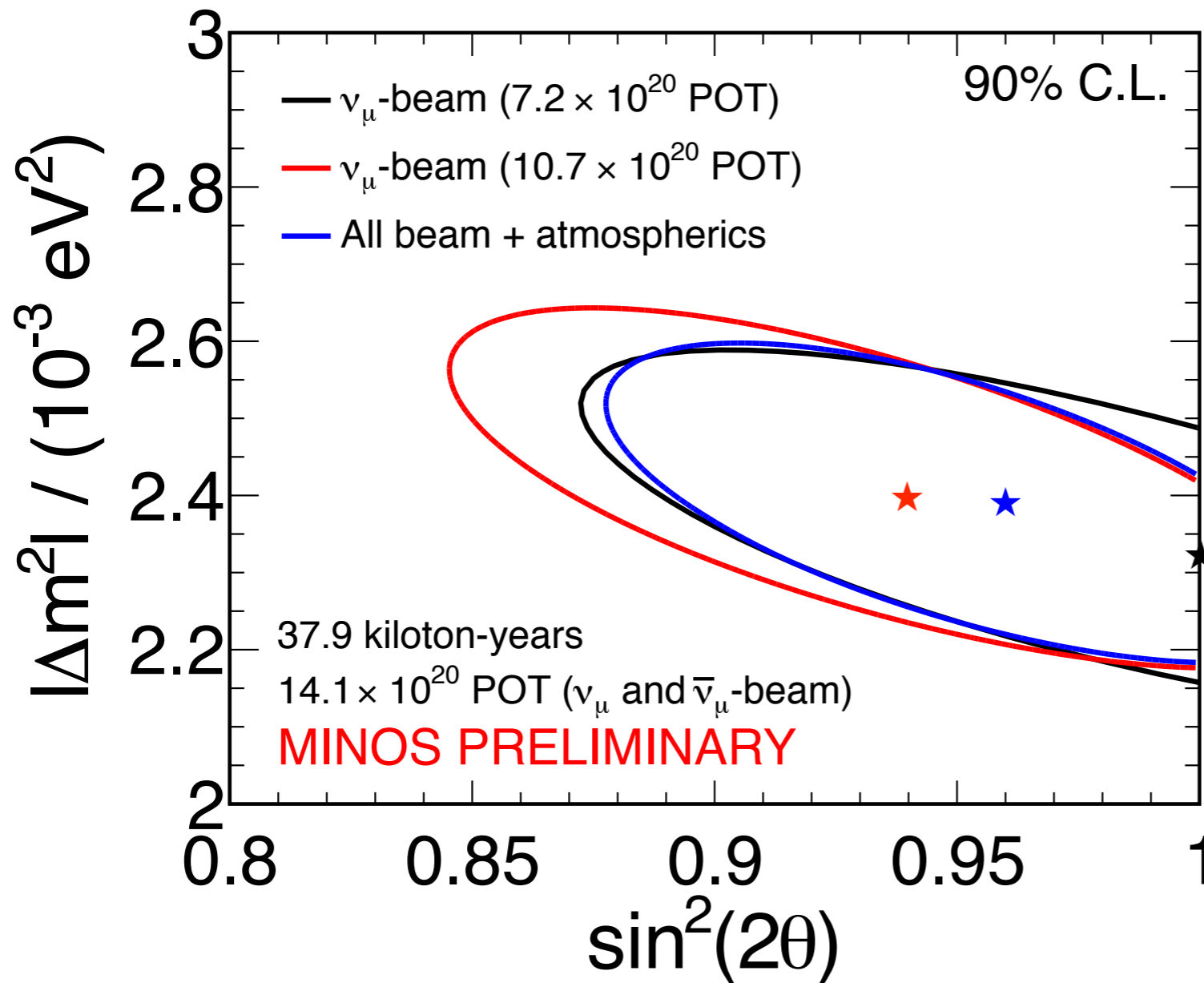


*...adding KamLAND*

*will improve with spectral data from DayaBay / RENO*

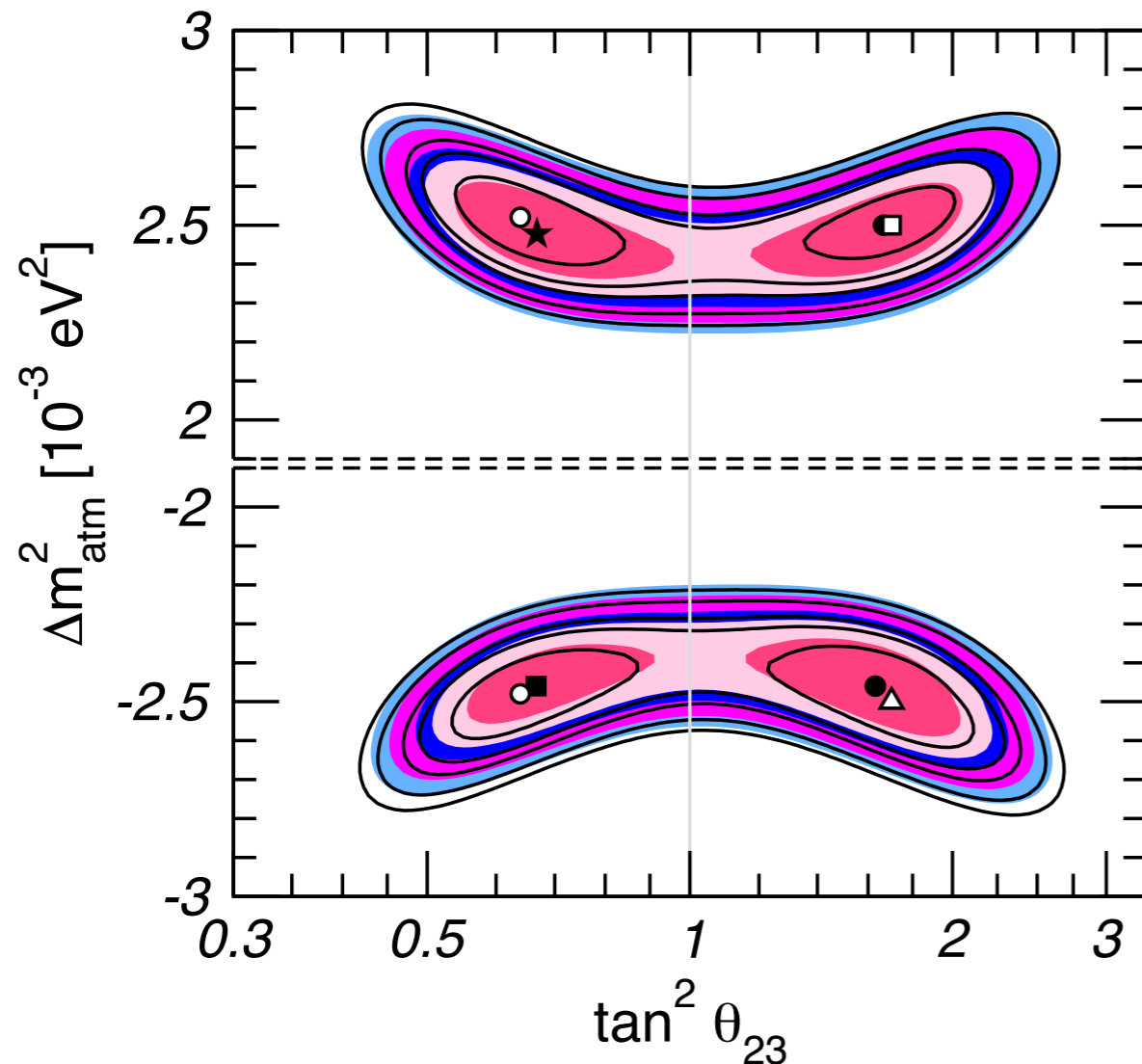
# *On non-maximal 23 mixing*

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Nichol (MINOS), talk  
at Neutrino2012

# *On non-maximal 23 mixing*



global data without  
atmospheric  
(MINOS and T2K  
disappearance most  
important)

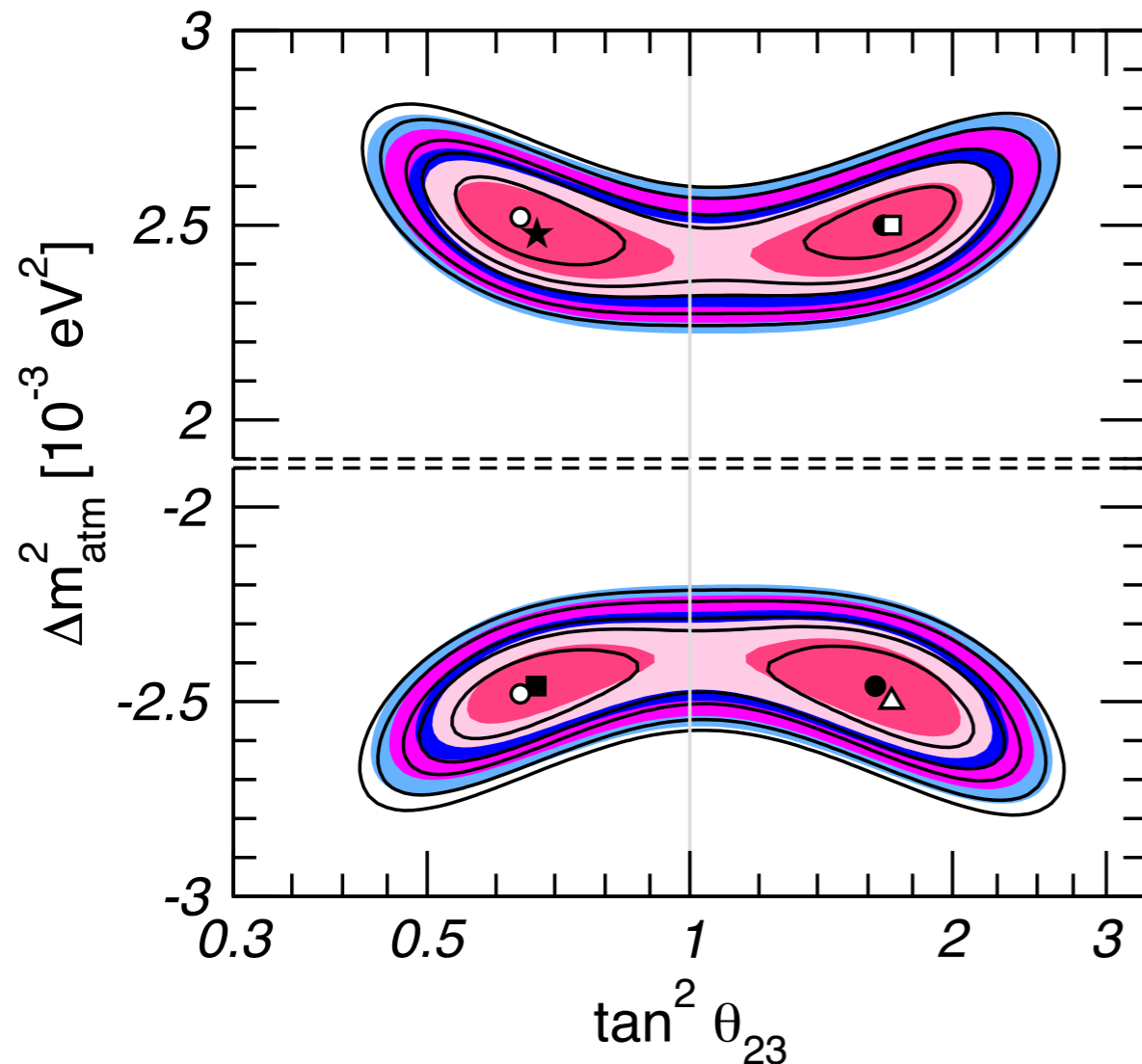
$$\sin^2 \theta_{23} \approx 0.40$$

$$\sin^2 \theta_{23} \approx 0.62$$

*degeneracy between the two  $\theta_{23}$  octants*



# On non-maximal 23 mixing



global data without  
atmospheric  
(MINOS and T2K  
disappearance most  
important)

$$\sin^2 \theta_{23} \approx 0.40$$

$$\sin^2 \theta_{23} \approx 0.62$$

degeneracy between the two  $\theta_{23}$  octants

neglecting  $\Delta m^2_{21}$ :  $P_{\mu\mu} \approx 1 - 4|U_{\mu 3}|^2(1 - |U_{\mu 3}|^2) \sin^2 \frac{\Delta m^2_{\text{atm}} L}{4E} \Rightarrow \sin^2 \theta_{23} = \frac{|U_{\mu 3}|^2}{\cos^2 \theta_{13}}$

slight shift to larger values of  $\sin^2 \theta_{23}$

# Octant degeneracy and LBL appearance

Fogli, Lisi, hep-ph/9604415

$$P_{\mu e} \simeq \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2(1-A)\Delta}{(1-A)^2} + \sin 2\theta_{13} \hat{\alpha} \sin 2\theta_{23} \frac{\sin(1-A)\Delta}{1-A} \frac{\sin A\Delta}{A} \cos(\Delta + \delta_{\text{CP}}) + \hat{\alpha}^2 \cos^2 \theta_{23} \frac{\sin^2 A\Delta}{A^2}$$

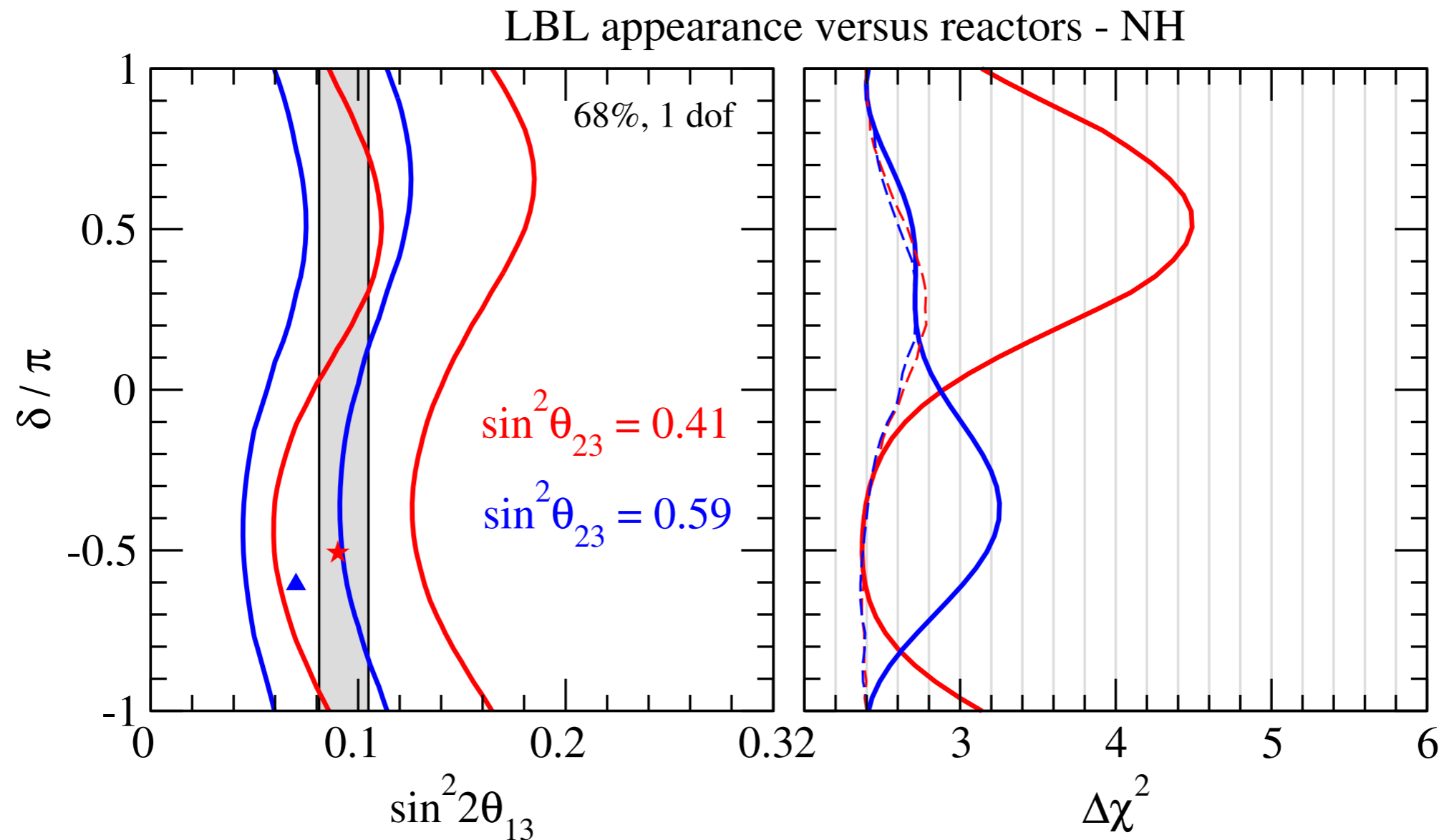
with

$$\Delta \equiv \frac{\Delta m_{31}^2 L}{4E_\nu}, \quad \hat{\alpha} \equiv \frac{\Delta m_{21}^2}{\Delta m_{31}^2} \sin 2\theta_{12}, \quad A \equiv \frac{2E_\nu V}{\Delta m_{31}^2}$$

- for large  $\theta_{13}$  the leading term depends on octant
- beam+reactor combination may be sensitive to octant

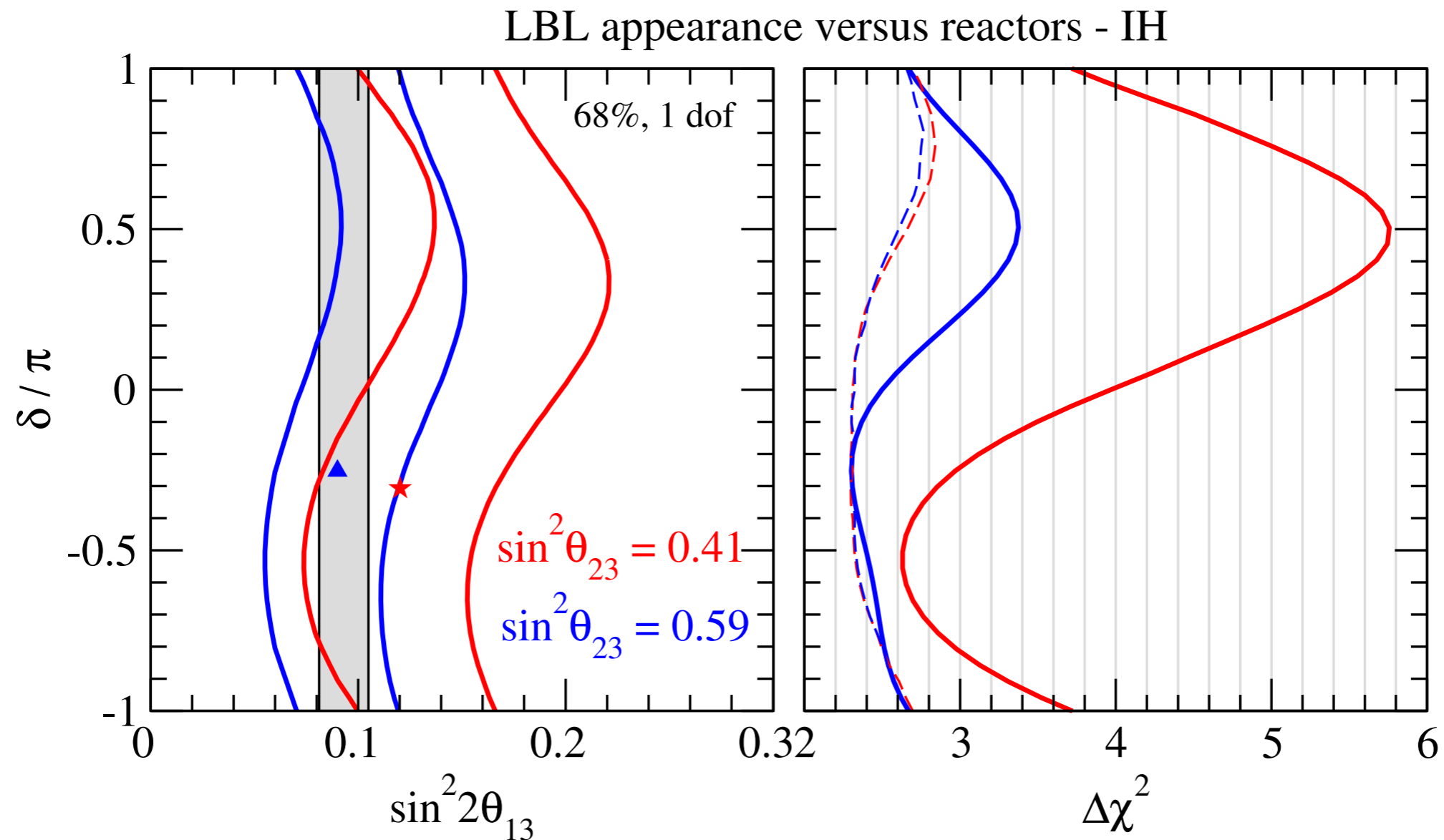
Minakata et al. hep-ph/0211111; McConnel, Shaevitz, hep-ex/0409028

# Octant degeneracy and LBL appearance



*present data from LBL appearance versus reactor cannot discriminate between the octants*

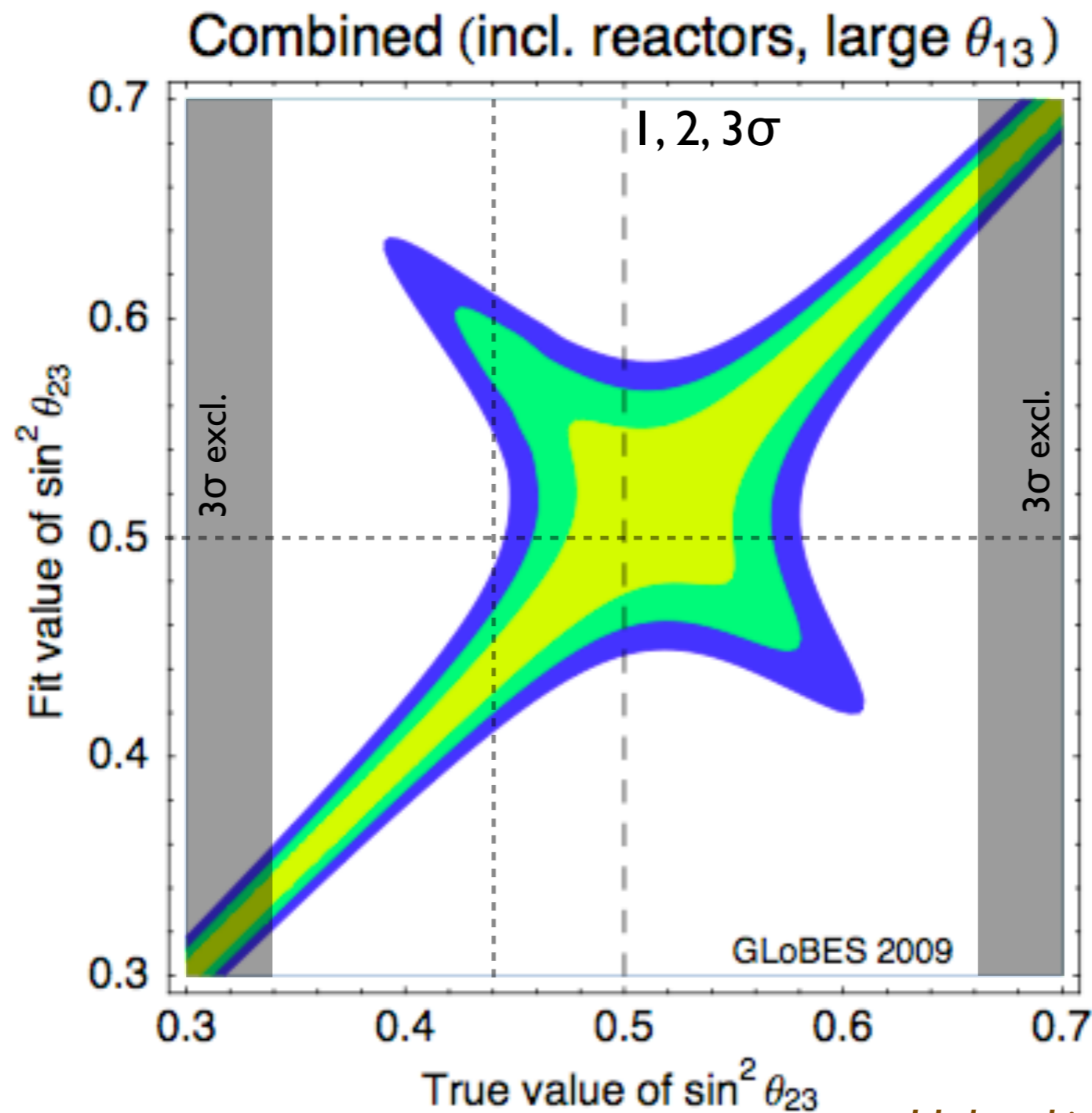
# Octant degeneracy and LBL appearance



*present data from LBL appearance versus reactor cannot discriminate between the octants*

# Global fit $\sim 2020$ - $\theta_{23}$ octant

final exposure of T2K, NOvA, DayaBay combined



$$\sin^2 2\theta_{13} = 0.1$$
$$\delta = 0$$

Huber, Lindner, TS, Winter, 0907.1896

# 3-flavor effects in atmospheric neutrinos

excess in electron-like events:

$$\begin{aligned} \frac{N_e}{N_e^0} - 1 &\simeq (r s_{23}^2 - 1) P_{2\nu}(\Delta m_{31}^2, \theta_{13}) && \theta_{13}\text{-effects} \\ &+ (r c_{23}^2 - 1) P_{2\nu}(\Delta m_{21}^2, \theta_{12}) && \Delta m_{21}^2\text{-effects} \\ &- 2s_{13}s_{23}c_{23} r \operatorname{Re}(A_{ee}^* A_{\mu e}) && \text{interference: } \delta_{\text{CP}} \end{aligned}$$

$$r = r(E_\nu) \equiv \frac{F_\mu^0(E_\nu)}{F_e^0(E_\nu)} \quad \begin{array}{l} r \approx 2 \quad (\text{sub-GeV}) \\ r \approx 2.6 - 4.5 \quad (\text{multi-GeV}) \end{array}$$

# 3-flavor effects in atmospheric neutrinos

Peres, Smirnov, 99;

Gonzalez-Garcia, Maltoni, Smirnov, 04

excess in electron-like events:

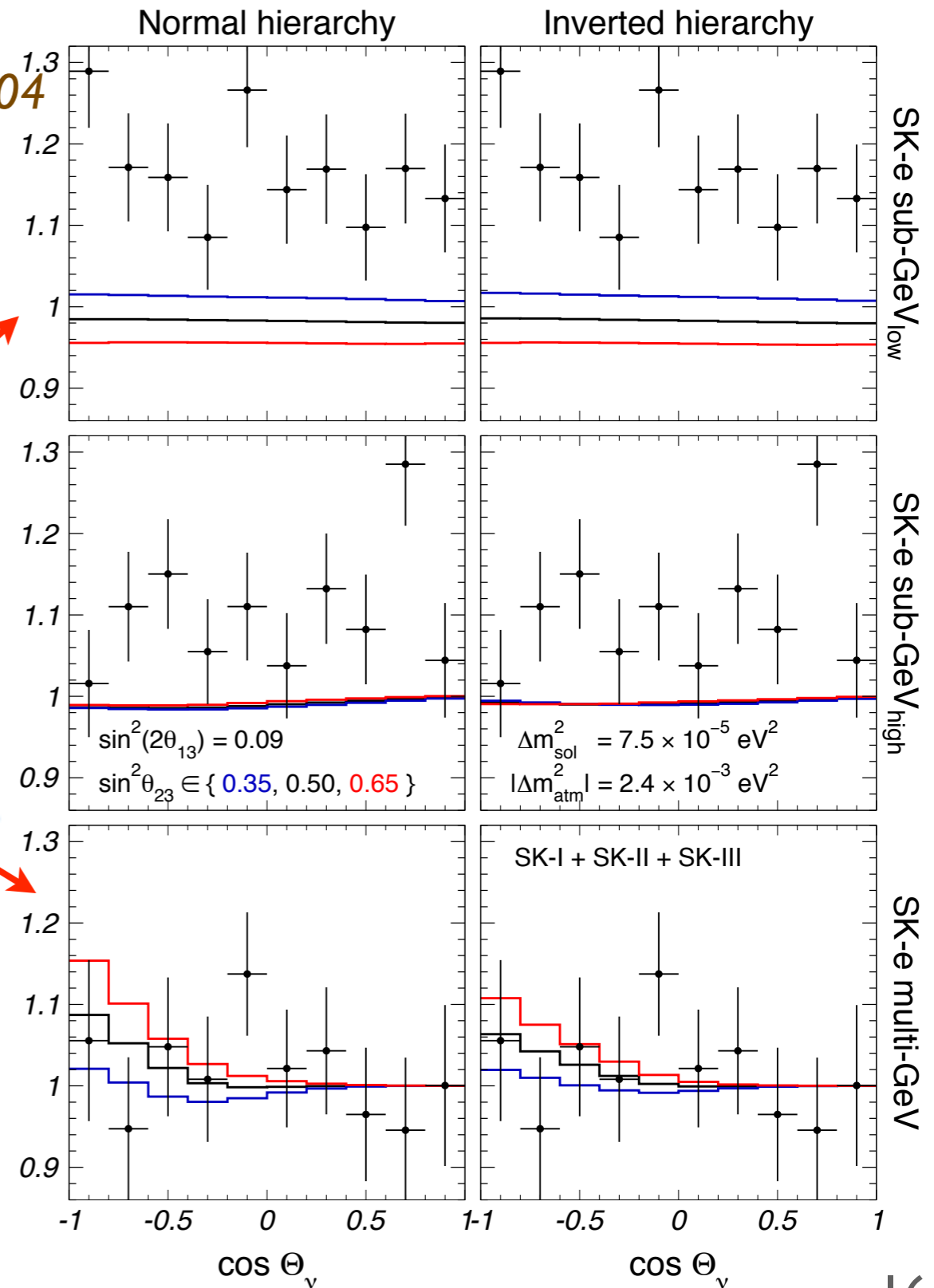
$$\frac{N_e}{N_e^0} - 1 \simeq (r s_{23}^2 - 1) P_{2\nu}(\Delta m_{31}^2, \theta_{13}) \quad \theta_{13}\text{-effects}$$

$$+ (r c_{23}^2 - 1) P_{2\nu}(\Delta m_{21}^2, \theta_{12}) \quad \Delta m_{21}^2\text{-effects}$$

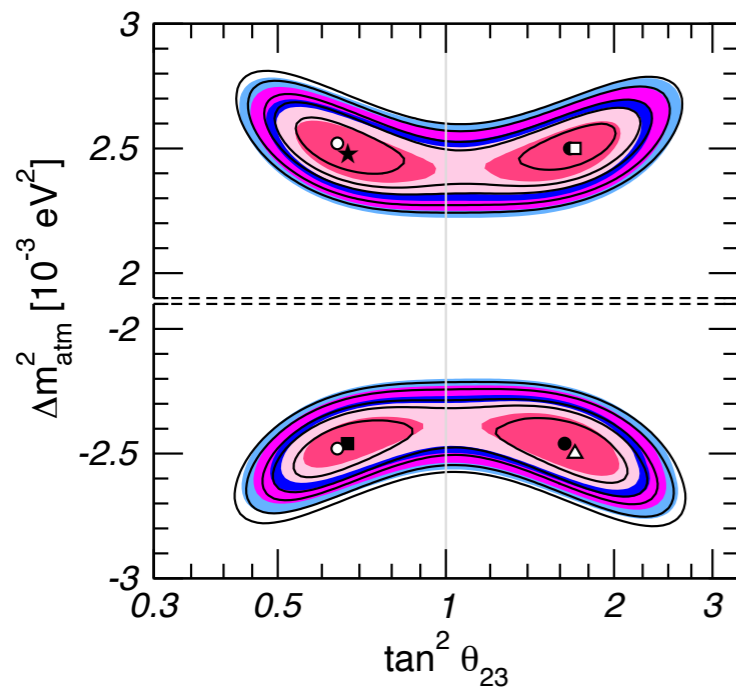
$$- 2s_{13}s_{23}c_{23} r \operatorname{Re}(A_{ee}^* A_{\mu e}) \quad \text{interference: } \delta_{CP}$$

$$r = r(E_\nu) \equiv \frac{F_\mu^0(E_\nu)}{F_e^0(E_\nu)} \quad r \approx 2 \quad (\text{sub-GeV})$$

$$r \approx 2.6 - 4.5 \quad (\text{multi-GeV})$$



# The octant and atmospheric neutrino data



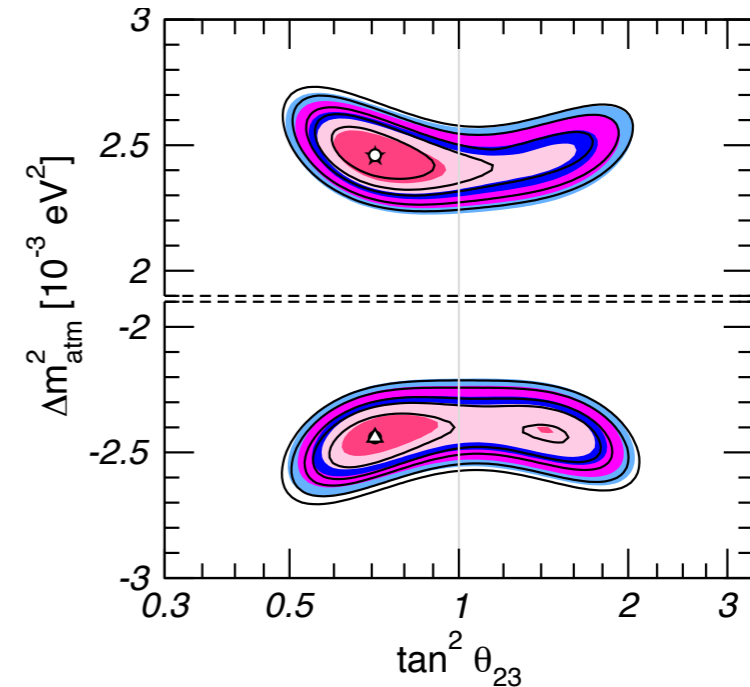
adding atmospheric



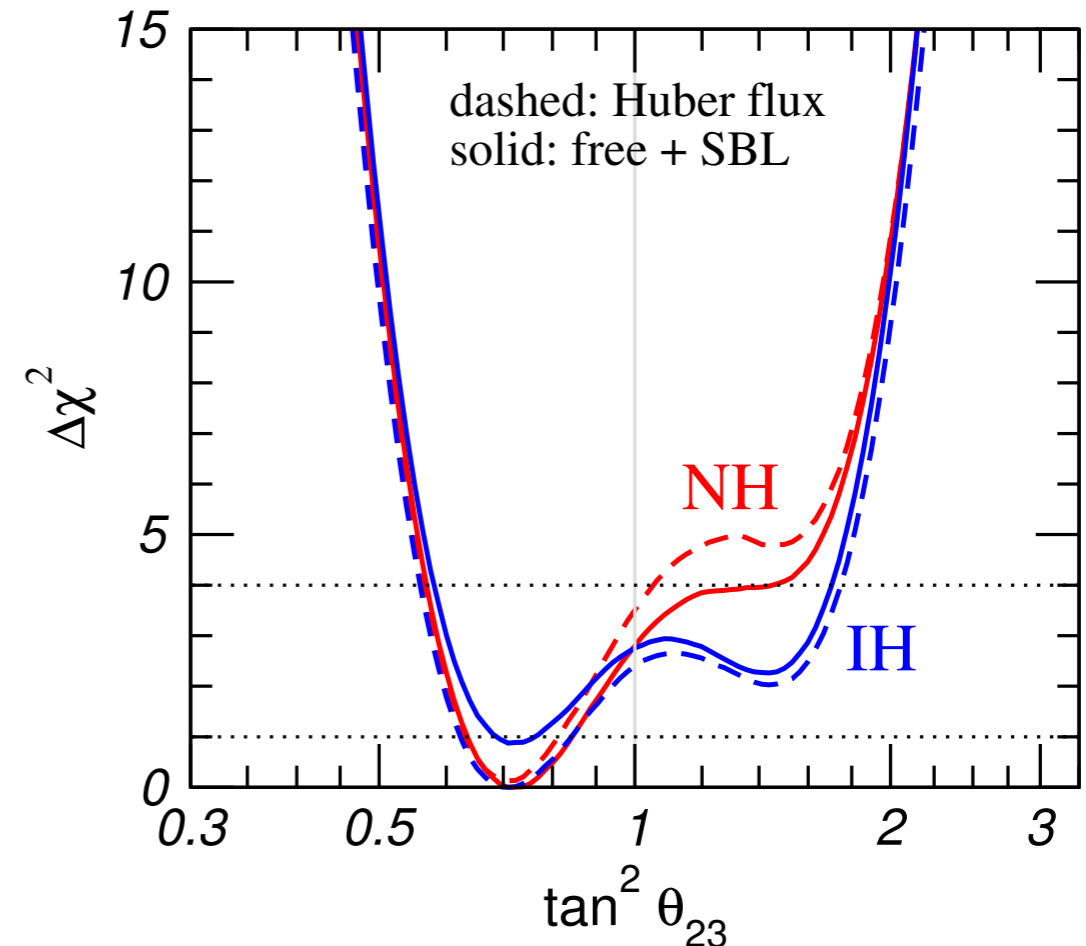
$$\sin^2 \theta_{23} = 0.42^{+0.037}_{-0.031}$$

$$0.39 - 0.66 @ 3\sigma$$

$$(\theta_{23} \approx 40^\circ)$$



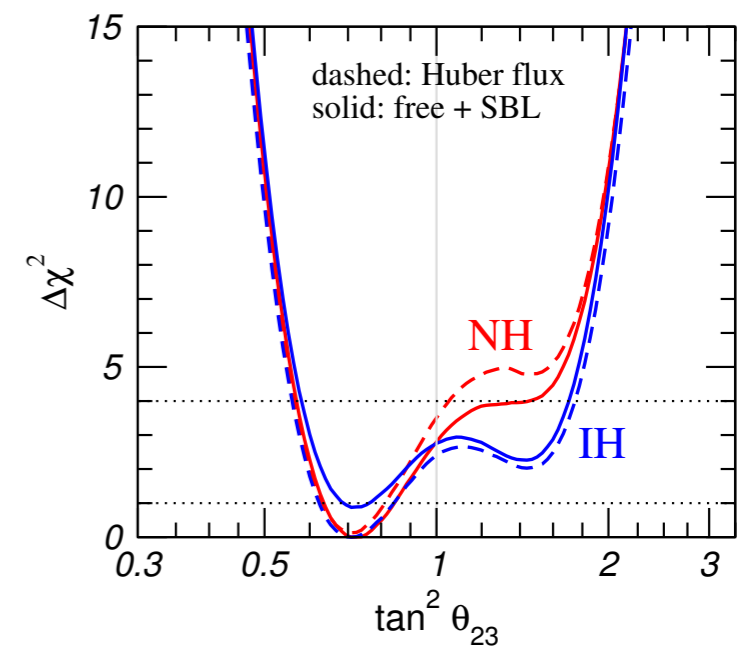
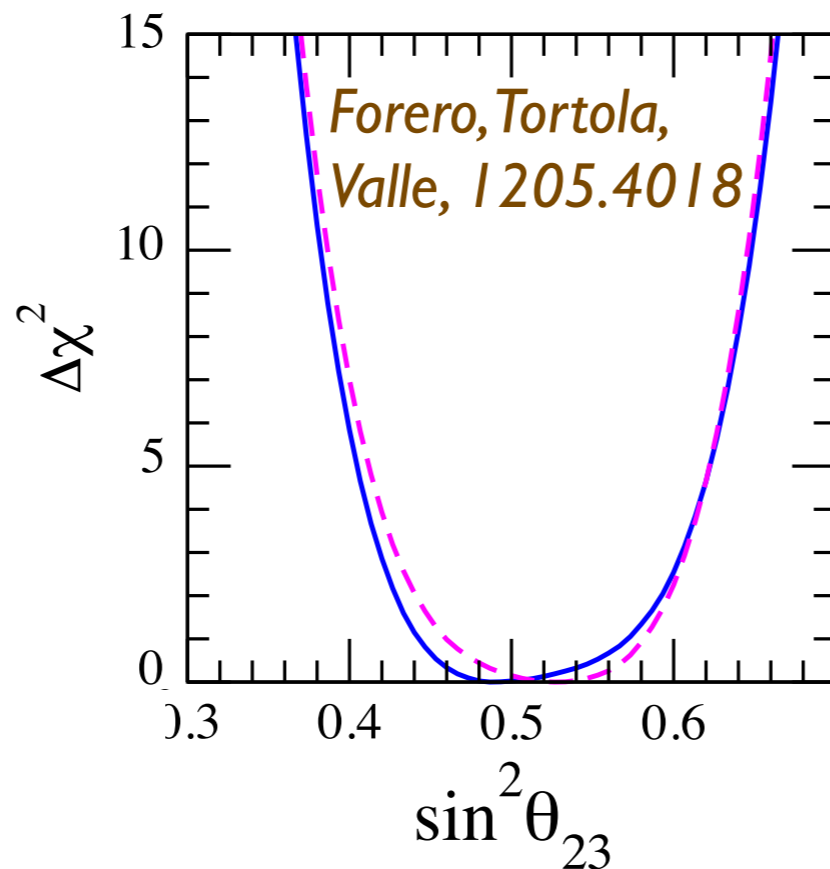
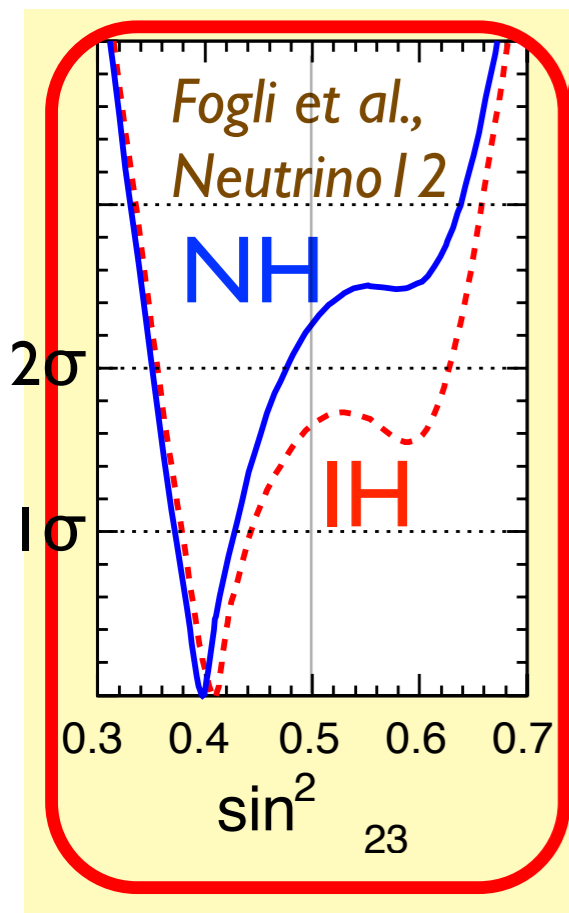
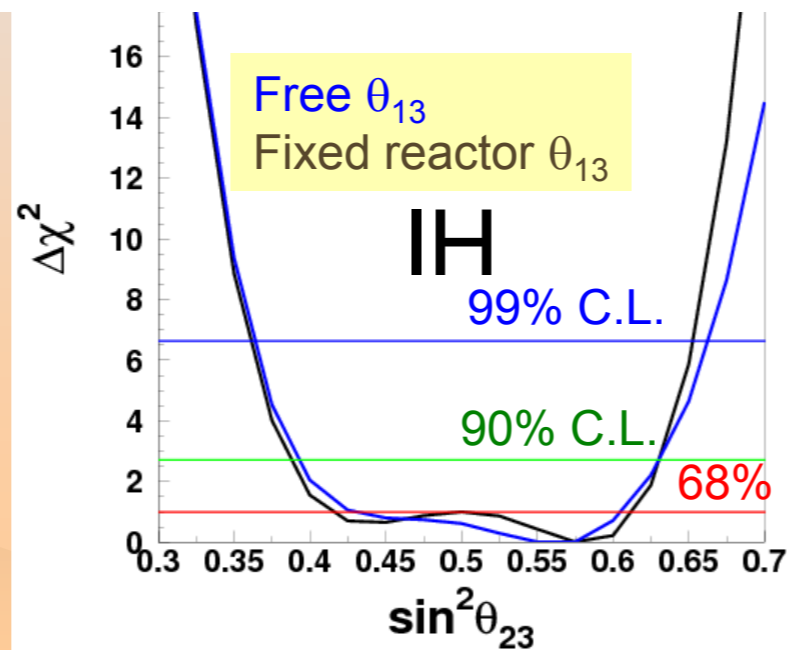
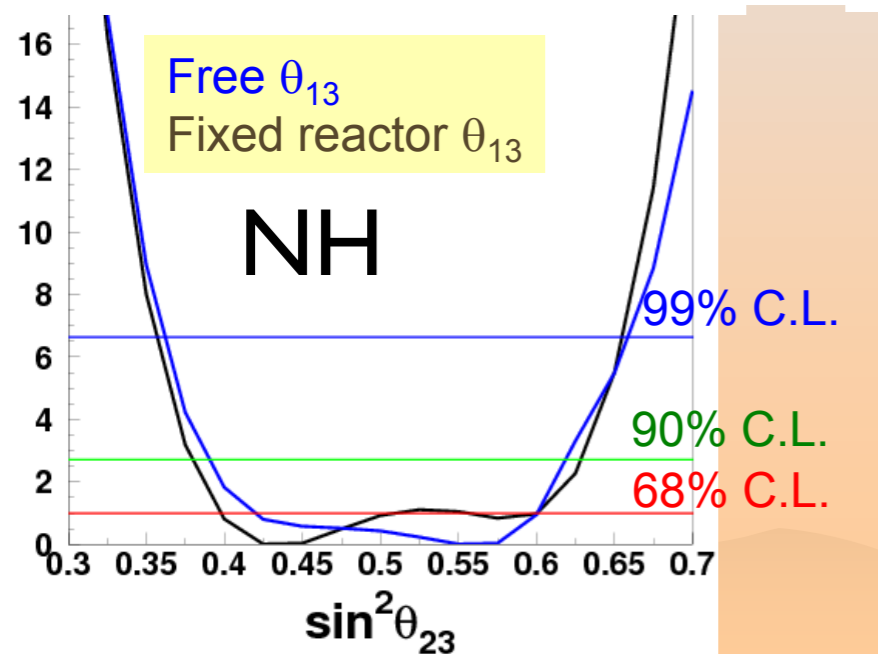
- slight preference for 1st octant by atmospheric data
- more pronounced for NH ( $\sim 2\sigma$ )
- depends on  $\theta_{13}$





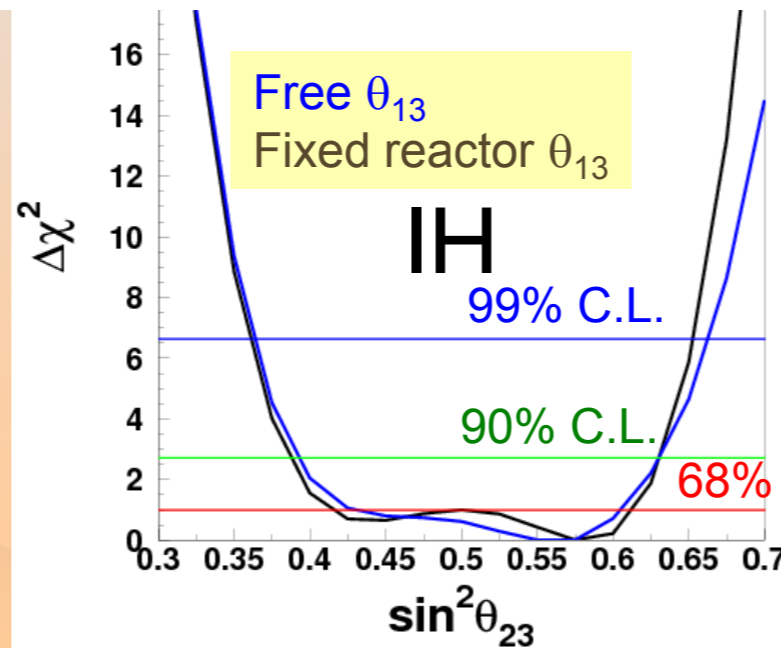
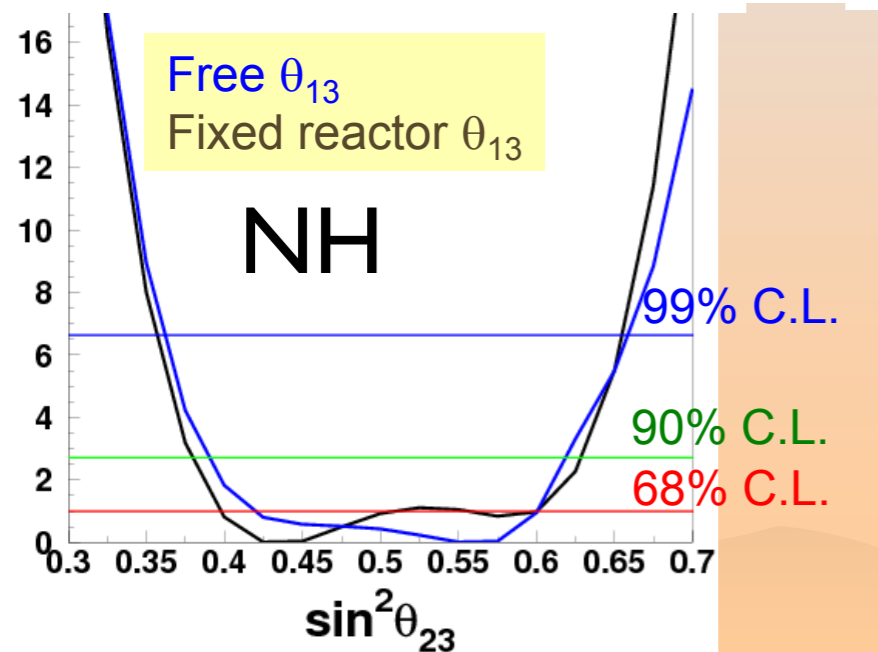
# A word of warning

Itow (SuperK), talk at Neutrino2012

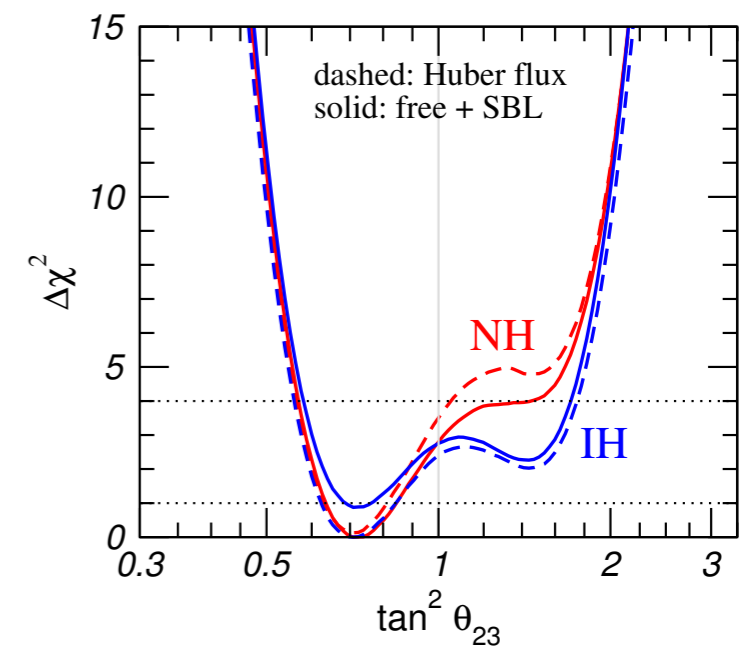
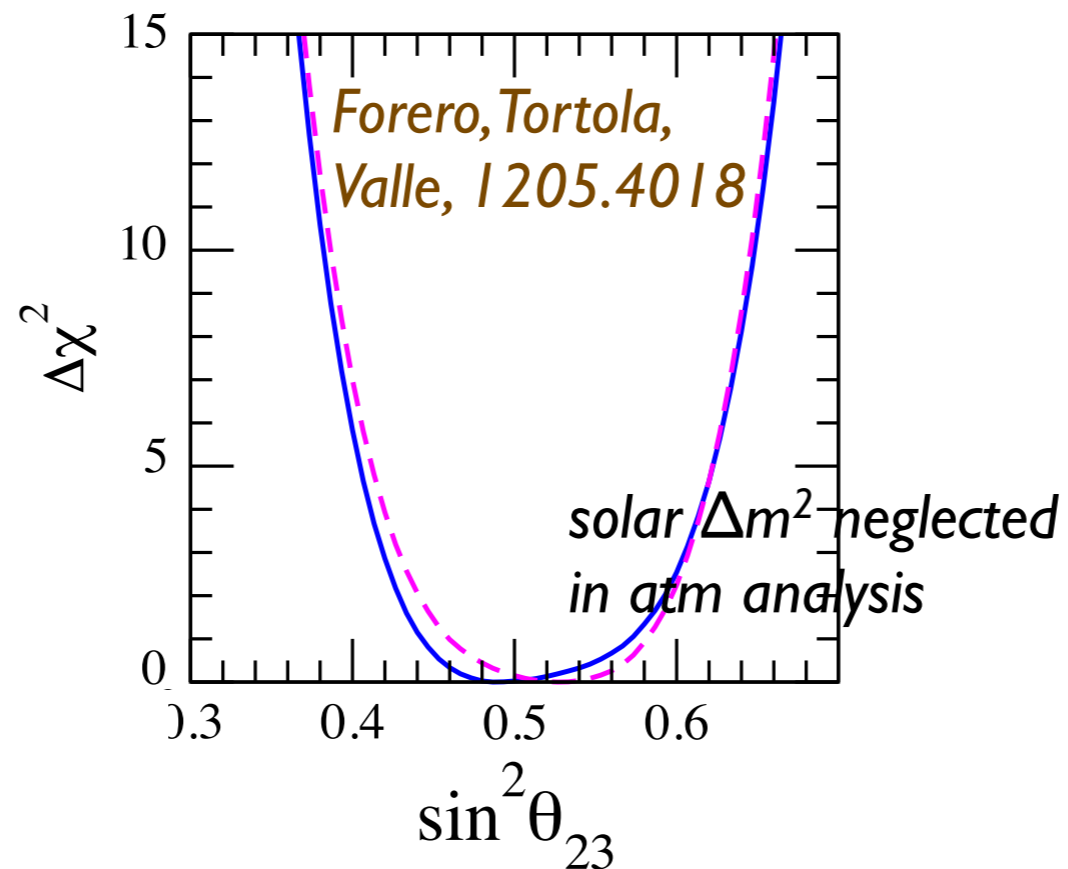
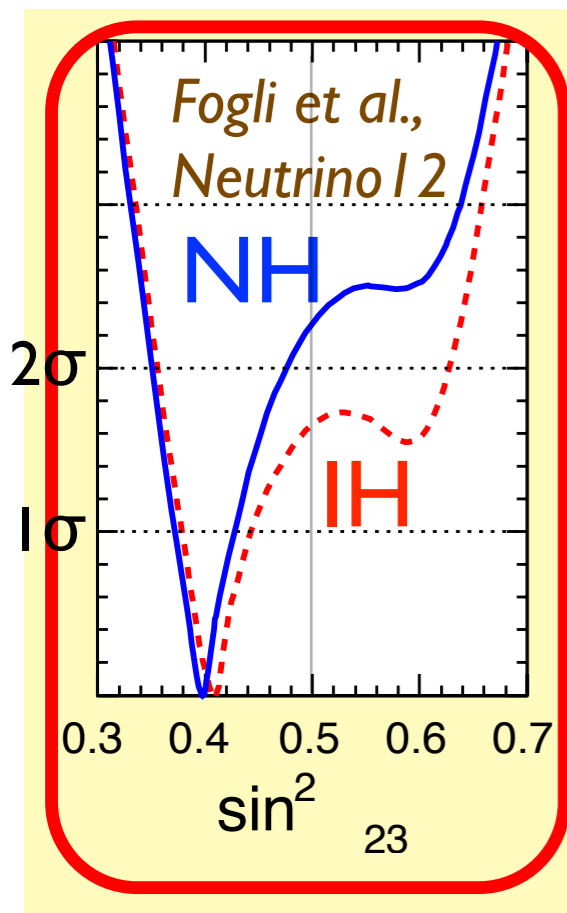


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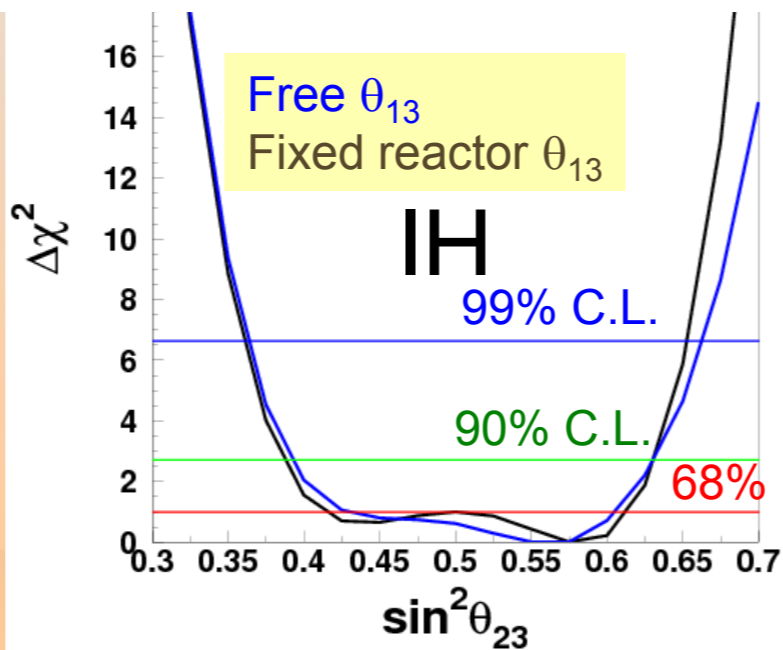
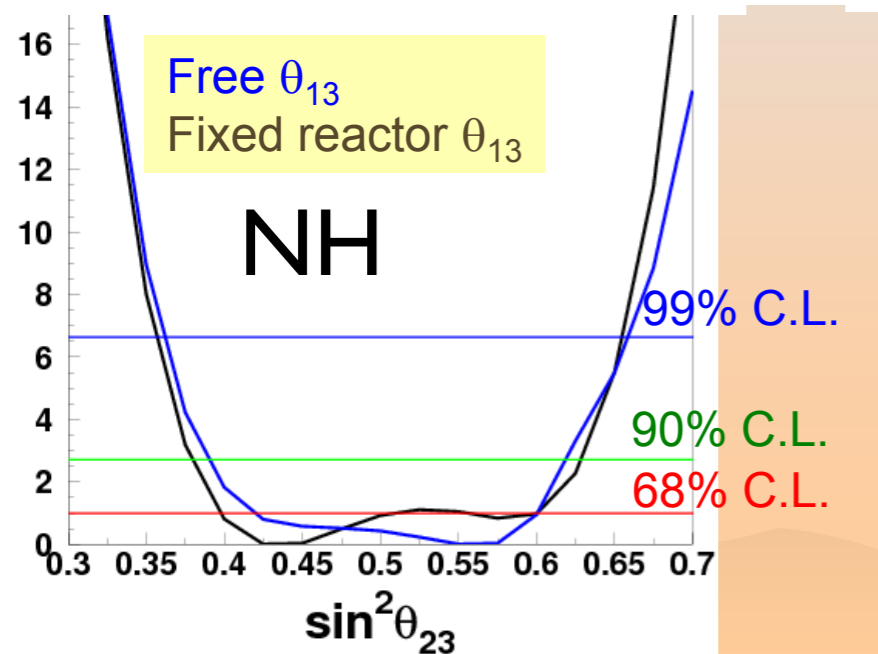
Itow (SuperK), talk at Neutrino2012



MINOS data from Neutrino2012 not included yet



# *A word of warning*



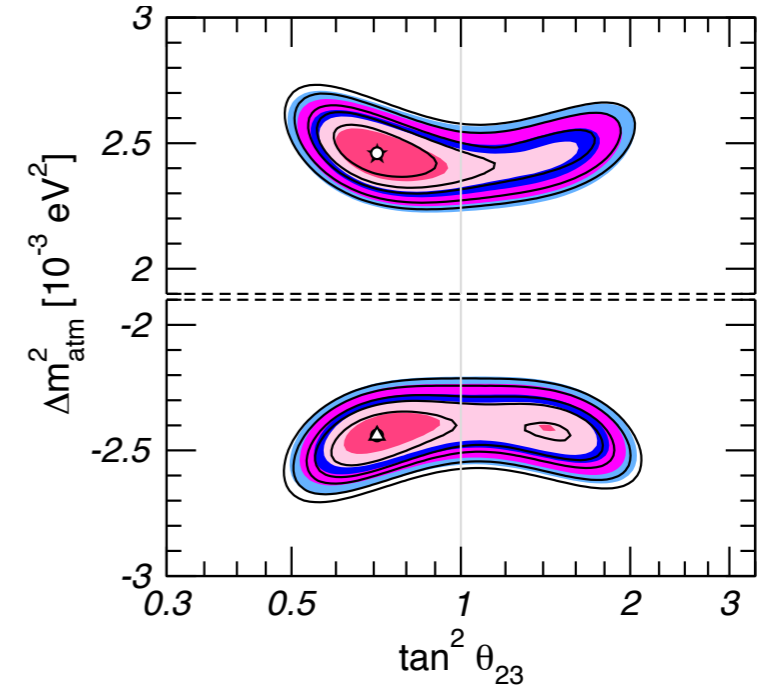
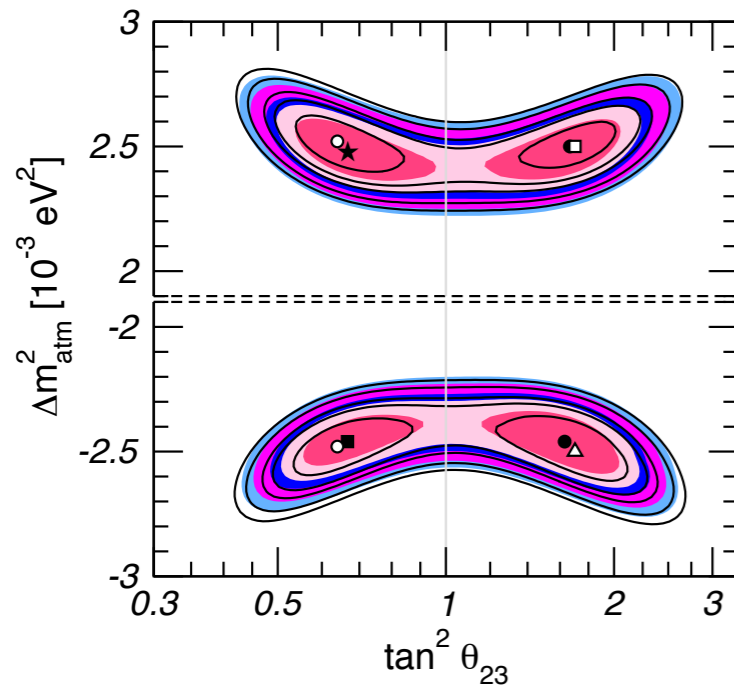
Itow (SuperK), talk  
at Neutrino2012

*personal interpretation:*

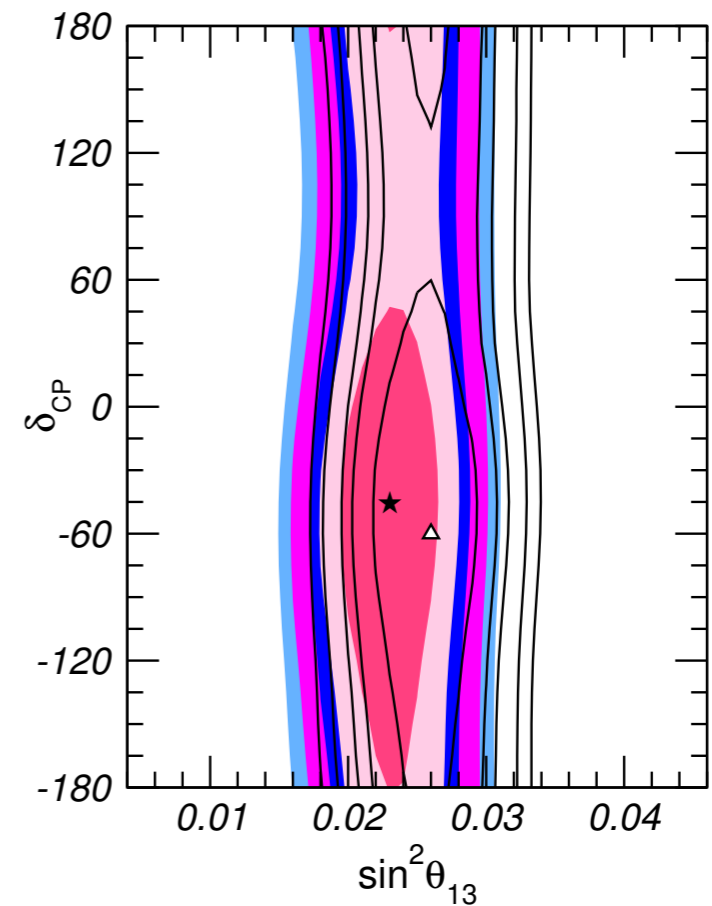
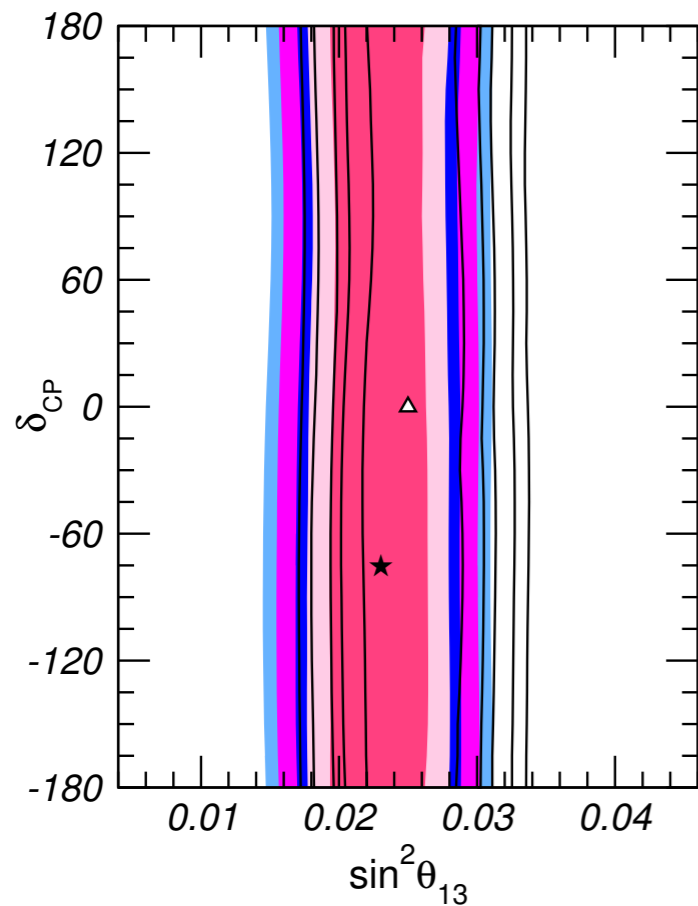
- *there seem to be subtle effects in SK atm data pulling into different directions*
- *sub-GeV data vs multi-GeV (?) data ( $\theta_{13}$  / hierarchy dependent)*
- *which effect wins depends on fine details of the analysis / system. treatment*

*before drawing definite conclusions it is mandatory to identify the physics and understand the origin of the different results from SK and phenomenologists*

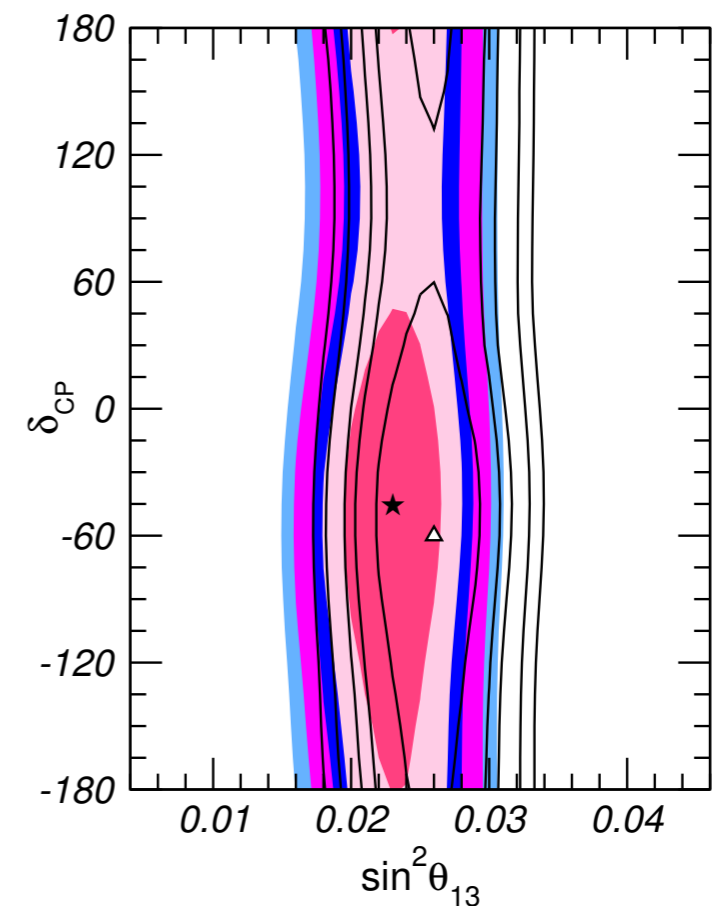
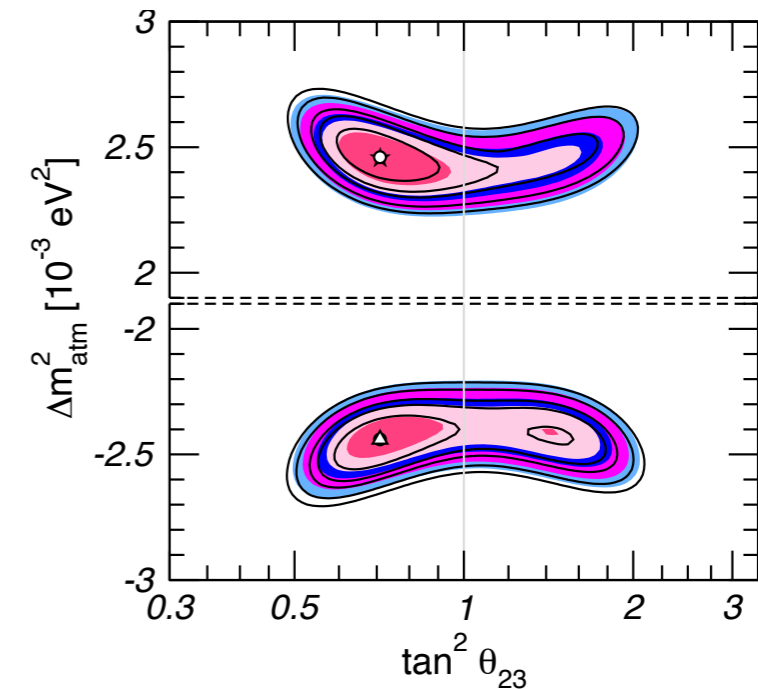
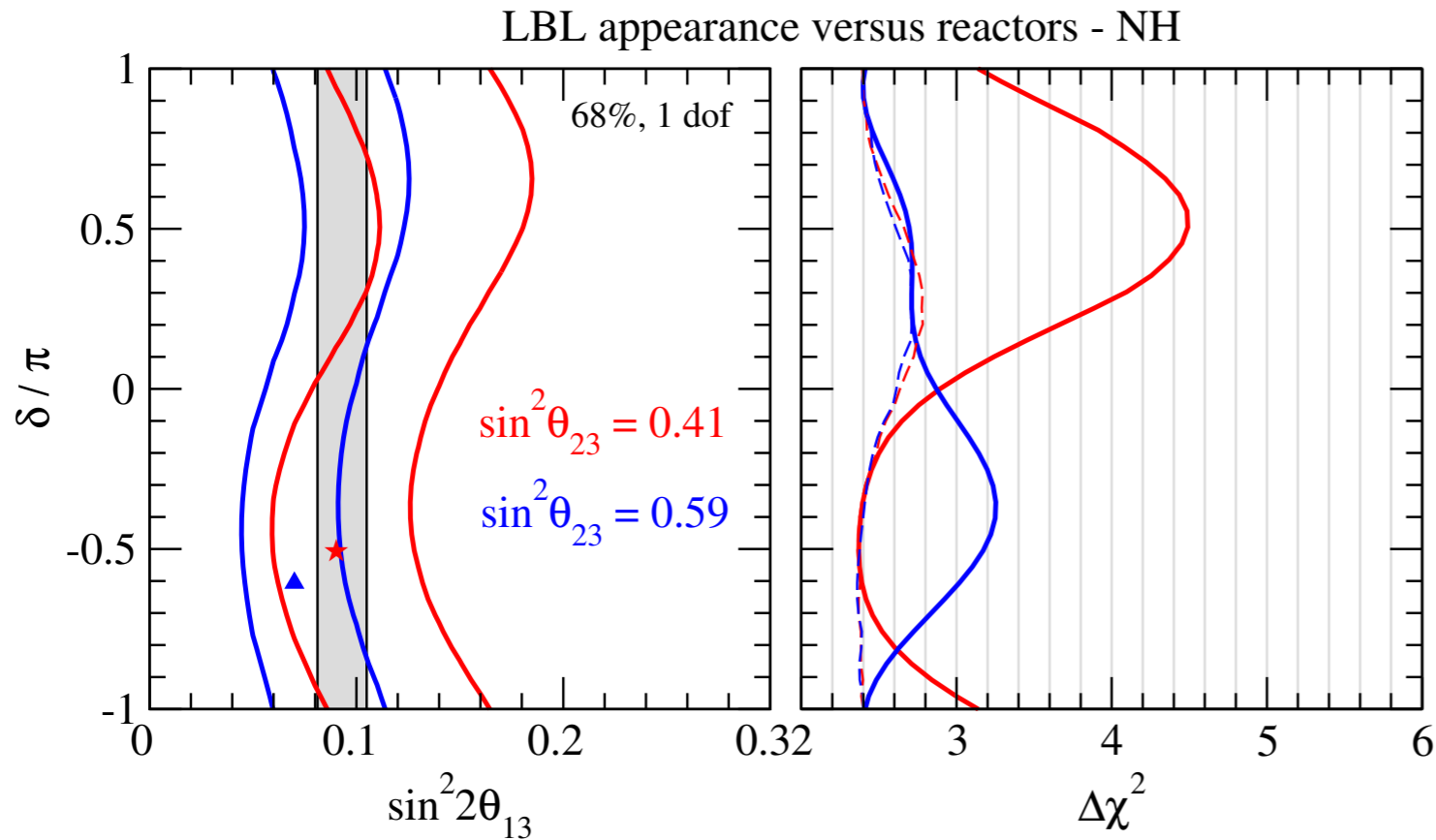
# The CP phase and atmospheric neutrino data



adding atmospheric

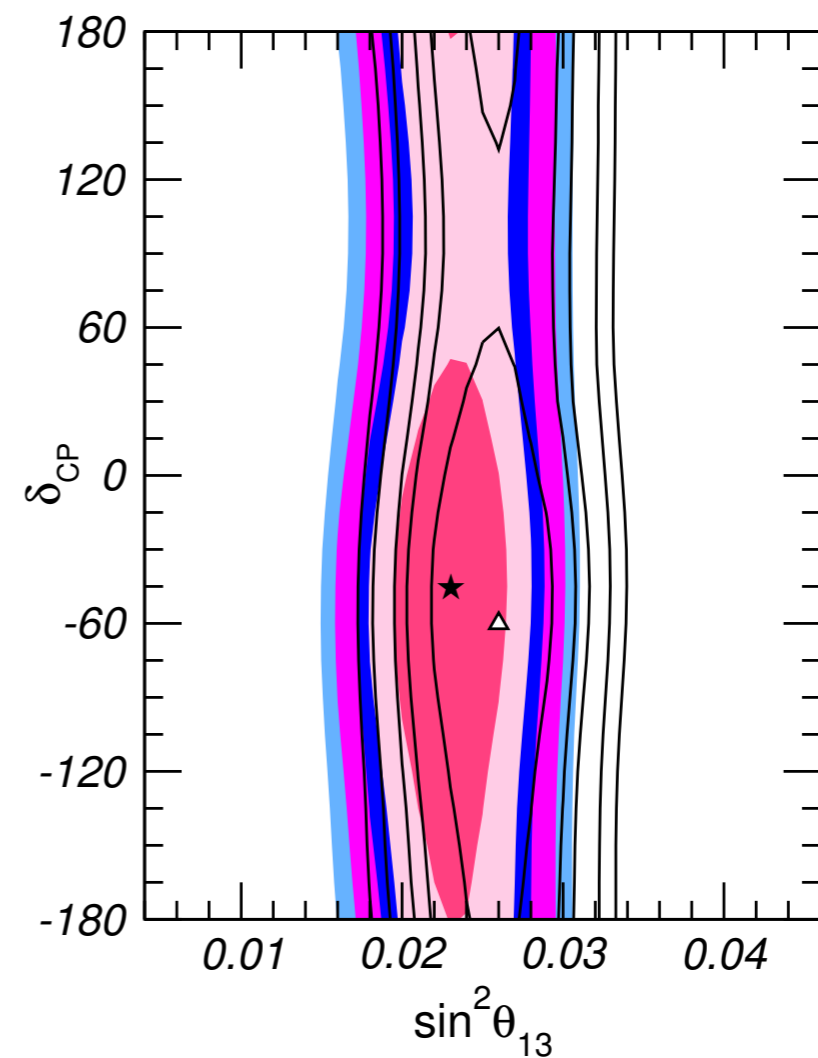
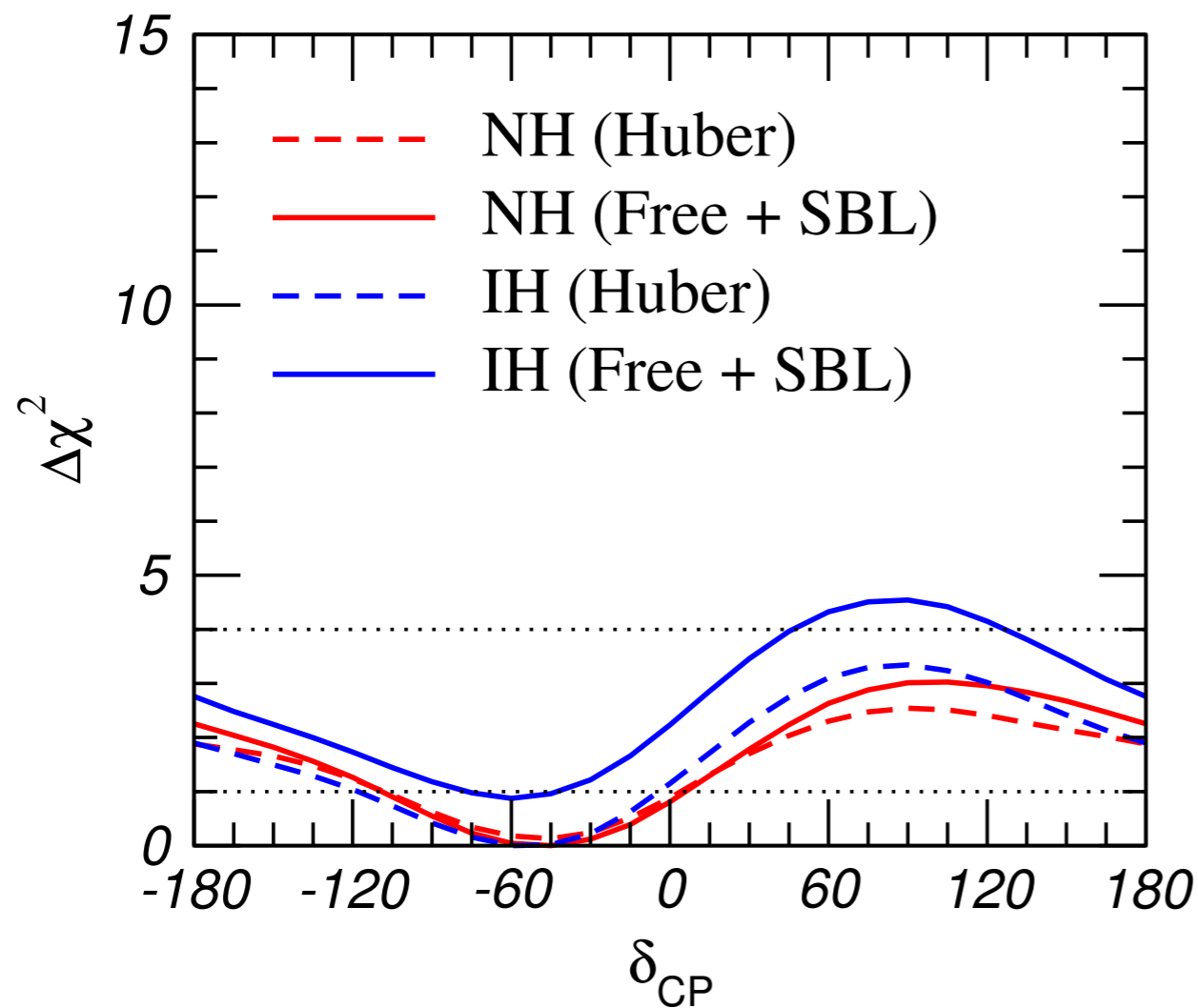


# The CP phase and atmospheric neutrino data



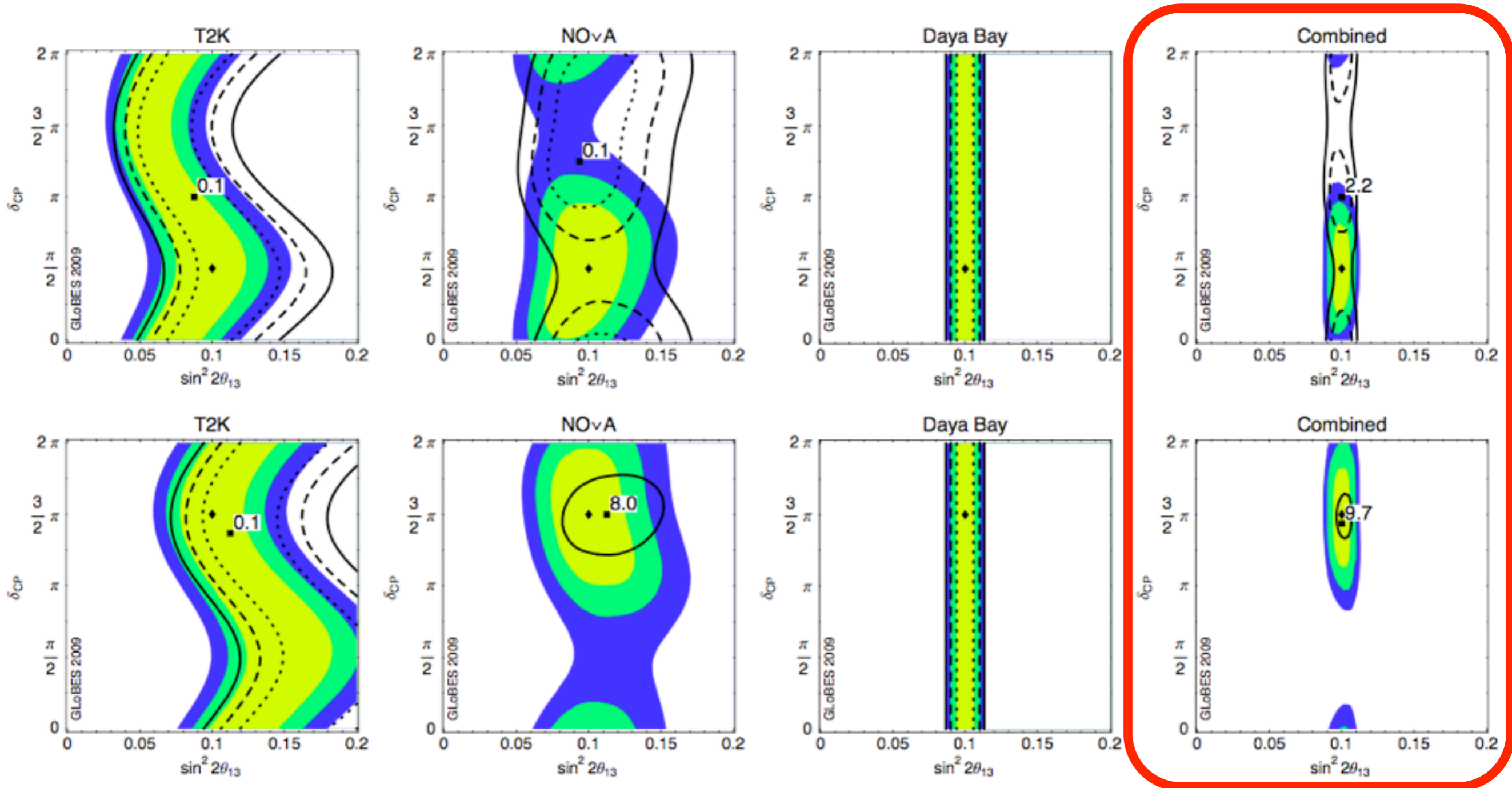
- ➔ non-maximality from LBL disapp
- ➔ preference for 1st octant from atm
- ➔ sensitivity to  $\delta$  from LBL appearance + reactors

# The CP phase



- “preferred” regions for  $\delta \sim -60^\circ$  at  $1\sigma$  (everything allowed at  $2\sigma$ )
- difference between NH and IH of  $\Delta\chi^2 \approx 1$

# Global fit $\sim 2020$ Huber, Lindner, TS, Winter, 0907.1896

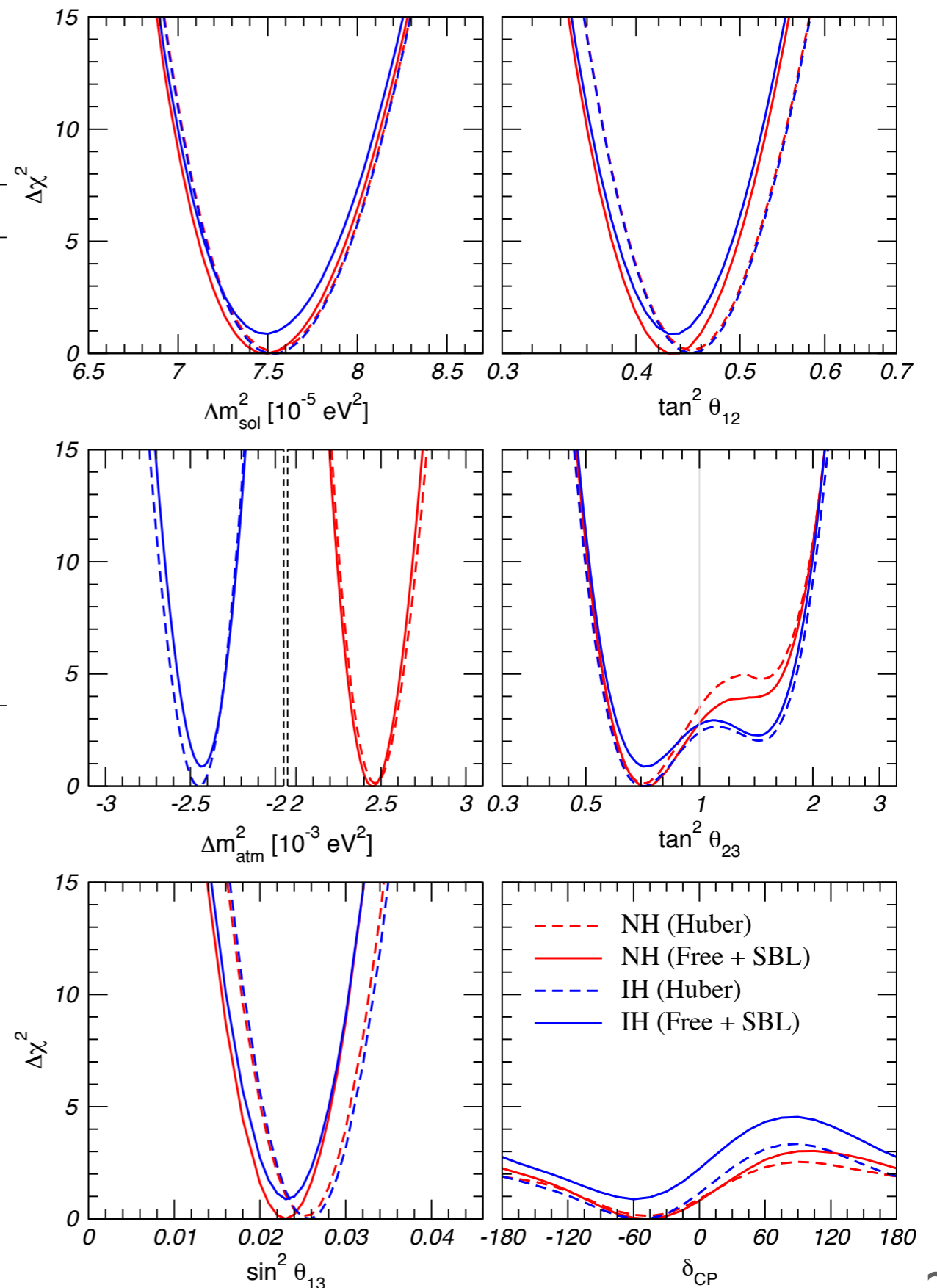


*some regions of  $\delta$  will appear but CP violation and hierarchy will be very hard with this generation of experiments*

# Summary

	bfp $\pm 1\sigma$	$3\sigma$ range
$\sin^2 \theta_{12}$	$0.30^{+0.13}_{-0.12}$	0.27 – 0.34
$\sin^2 \theta_{23}$	$0.42^{+0.037}_{-0.031}$	0.39 – 0.66
$\sin^2 \theta_{13}$	$0.023 \pm 0.0023$	0.016 – 0.030
$\delta/^\circ$ (NH)	$-48^{+53}_{-59}$	–
$\delta/^\circ$ (IH)	$-59^{+49}_{-60}$	–
$\Delta m_{21}^2/10^{-5}\text{eV}^2$	$7.50 \pm 0.185$	7.00 – 8.09
$\Delta m_{31}^2/10^{-3}\text{eV}^2$ (NH)	$2.45^{+0.067}_{-0.071}$	2.25 – 2.67
$ \Delta m_{32}^2 /10^{-3}\text{eV}^2$ (IH)	$2.43 \pm 0.068$	2.23 – 2.65

(reactor flux-free analysis + SBL react.)

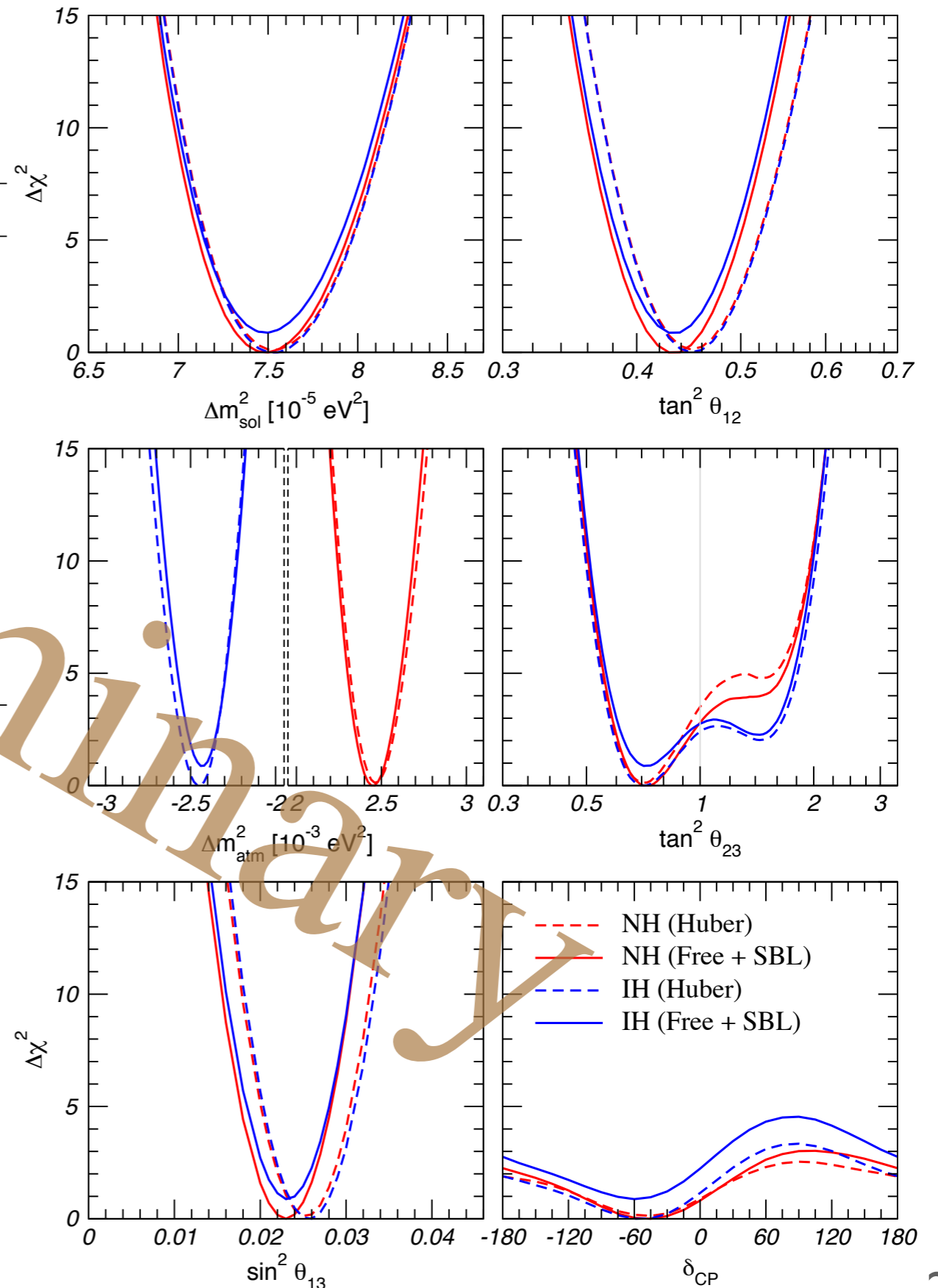




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$\sin^2 \theta_{12}$	$0.30^{+0.13}_{-0.12}$	0.27 – 0.34
$\sin^2 \theta_{23}$	$0.42^{+0.037}_{-0.031}$	0.39 – 0.66
$\sin^2 \theta_{13}$	$0.023 \pm 0.0023$	0.016 – 0.030
$\delta/^\circ$ (NH)	$-48^{+53}_{-59}$	–
$\delta/^\circ$ (IH)	$-59^{+49}_{-60}$	–
$\Delta m_{21}^2/10^{-5} \text{eV}^2$	$7.50 \pm 0.185$	7.00 – 8.09
$\Delta m_{31}^2/10^{-3} \text{eV}^2$ (NH)	$2.45^{+0.067}_{-0.071}$	2.25 – 2.67
$ \Delta m_{32}^2 /10^{-3} \text{eV}^2$ (IH)	$2.43 \pm 0.068$	2.23 – 2.65

(reactor flux-free analysis + SBL react.)



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***Thank you for your attention!***