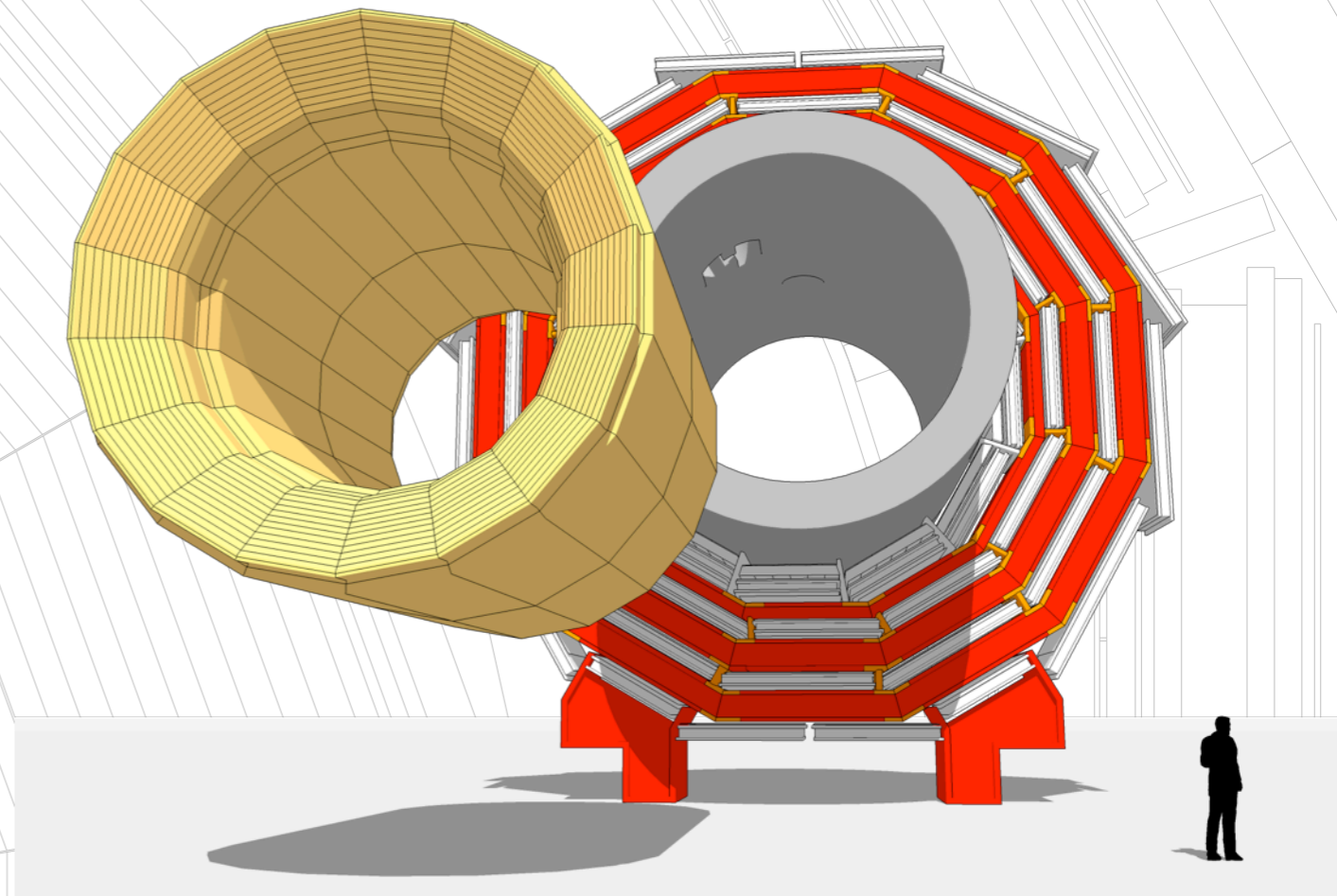




# Dark Matter Searches at CMS

**Tai Sakuma (Texas A&M)  
for the CMS collaboration**



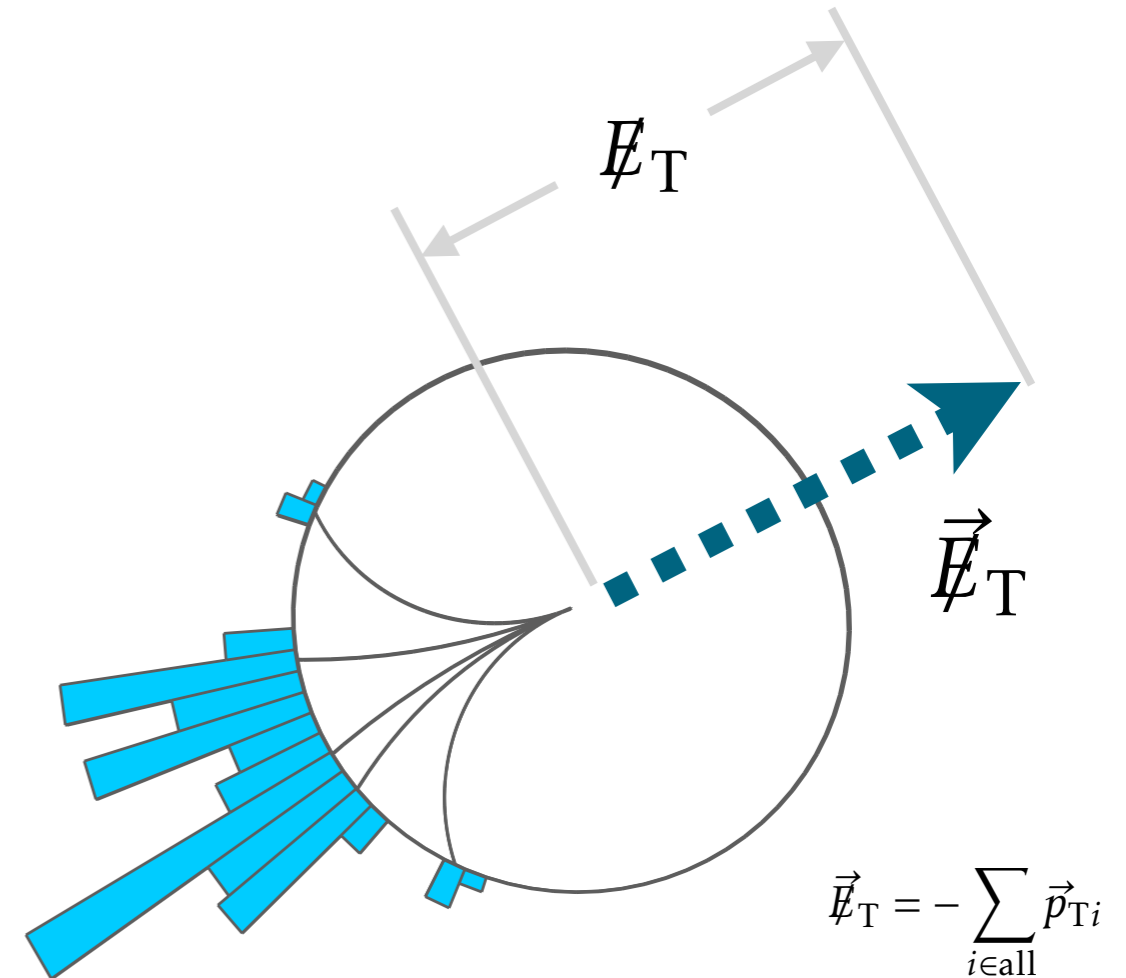
**3 February 2013**

**Aspen 2013 Closing in on Dark Matter**

<https://indico.cern.ch/contributionDisplay.py?contribId=65&confId=197862>

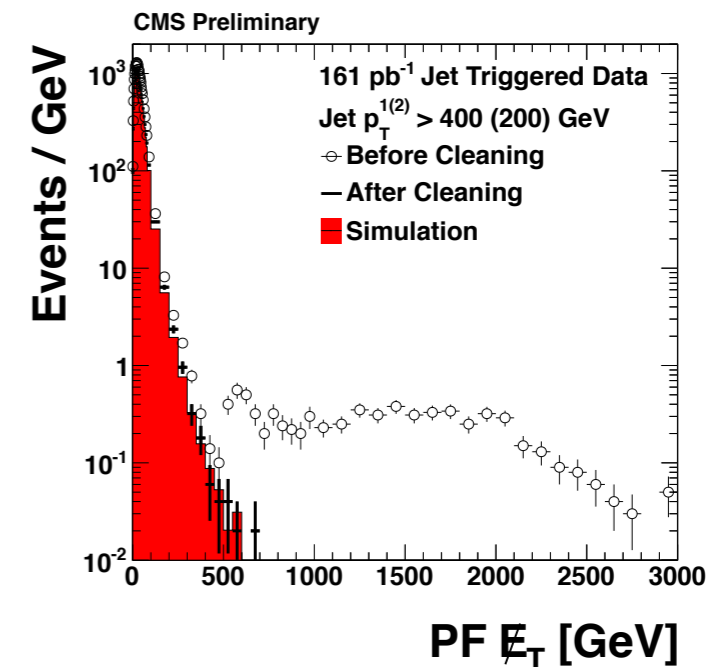
## • MET

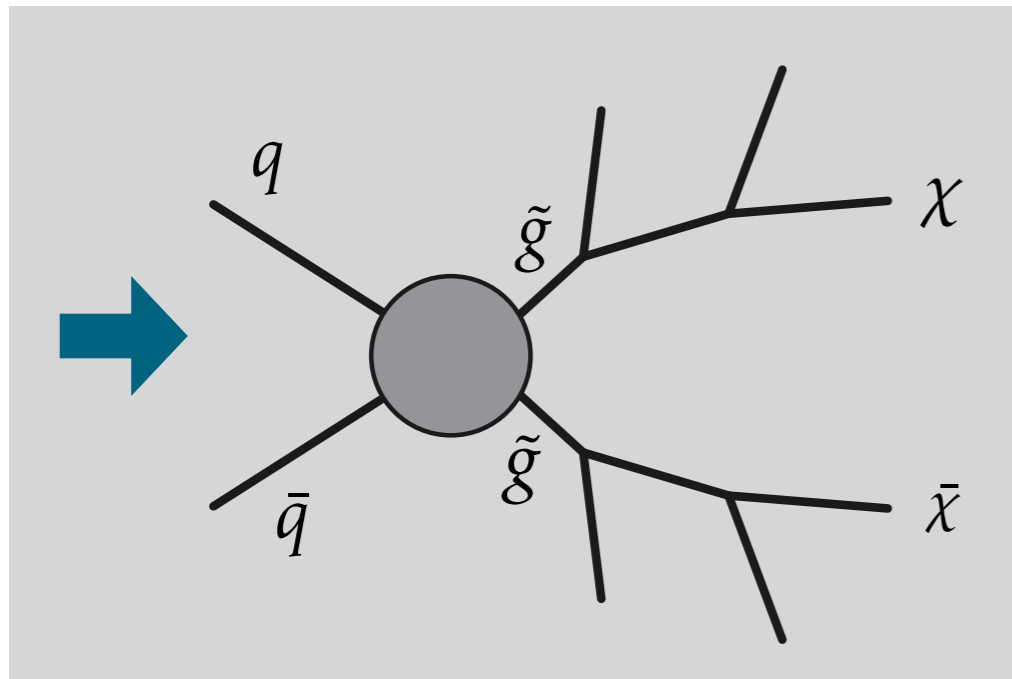
- the form in which the dark matter particles would appear in CMS
- the imbalance in the transverse momentum of all visible particles
  - ➔ the transverse momentum that must have been carried by invisible particles



## • MET reconstruction

- ➔ requires hermetic detector
- ➔ entails reconstruction of all particles with electromagnetic and strong interactions with precision
- ➔ susceptible to many types of imperfections, e.g., hot calorimeter cells, detector noise, beam-halo particles

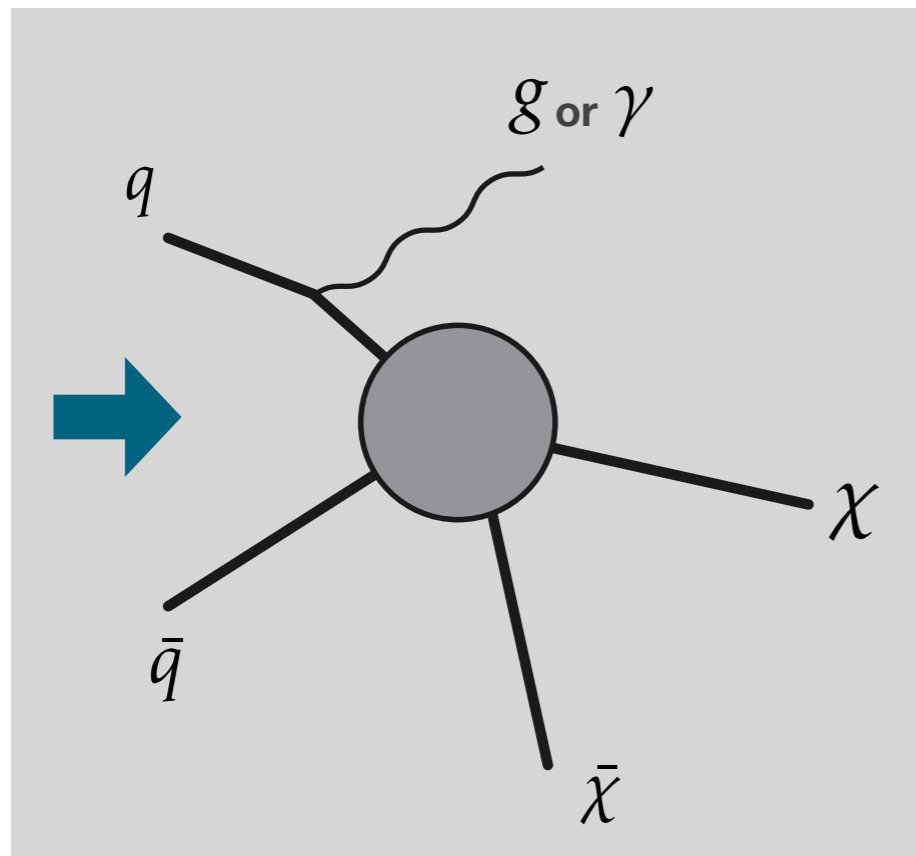




Heavier unstable new particles are produced  
Their decay chains end with dark matter particles

e.g. SUSY models

Signature: large MET + jets (+ leptons) (+ photons)



Dark Matter particles are directly produced in pairs  
after initial-state radiation (ISR)

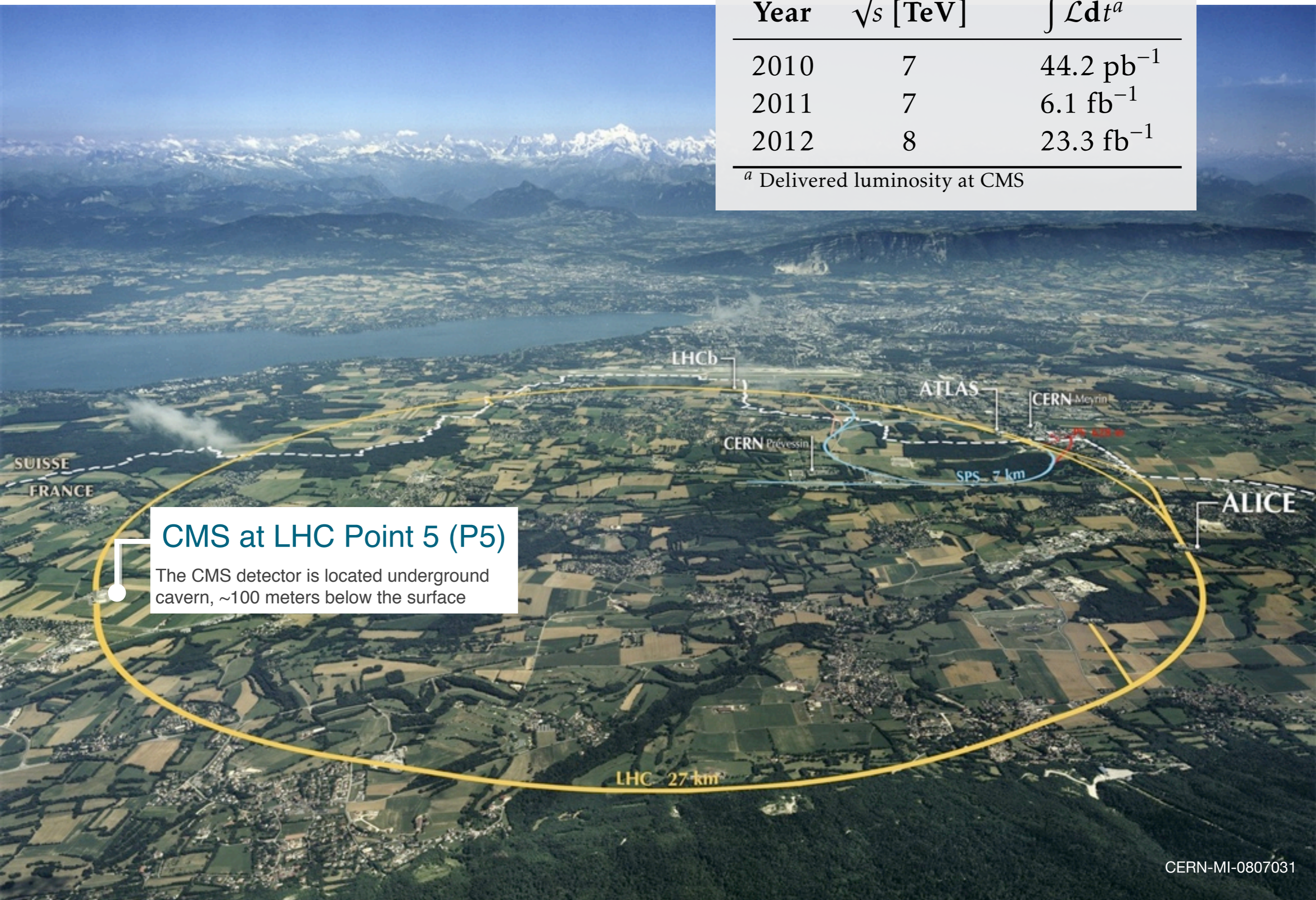
Signature: large MET + mono-jet (or mono-photon)

# Large Hadron Collider (LHC)

Proton-proton runs in the last 3 years

Year	$\sqrt{s}$ [TeV]	$\int \mathcal{L} dt^a$
2010	7	44.2 pb <sup>-1</sup>
2011	7	6.1 fb <sup>-1</sup>
2012	8	23.3 fb <sup>-1</sup>

<sup>a</sup> Delivered luminosity at CMS



## CMS at LHC Point 5 (P5)

The CMS detector is located underground cavern, ~100 meters below the surface

# CMS Collaboration

- 3275 physicists
  - 790 engineers and technicians
  - 179 institutes
  - 41 countries
- <http://cms.web.cern.ch/content/people-statistics>



summer 2012, in a surface building at P5  
(CMS-PHO-COLLAB-2012-004)



# CMS DETECTOR

Total weight : 14,000 tonnes  
Overall diameter : 15.0 m  
Overall length : 28.7 m  
Magnetic field : 3.8 T

STEEL RETURN YOKE  
12,500 tonnes

SILICON TRACKERS  
Pixel ( $100 \times 150 \mu\text{m}$ )  $\sim 16\text{m}^2 \sim 66\text{M}$  channels  
Microstrips ( $80 \times 180 \mu\text{m}$ )  $\sim 200\text{m}^2 \sim 9.6\text{M}$  channels

SUPERCONDUCTING SOLENOID  
Niobium titanium coil carrying  $\sim 18,000\text{A}$

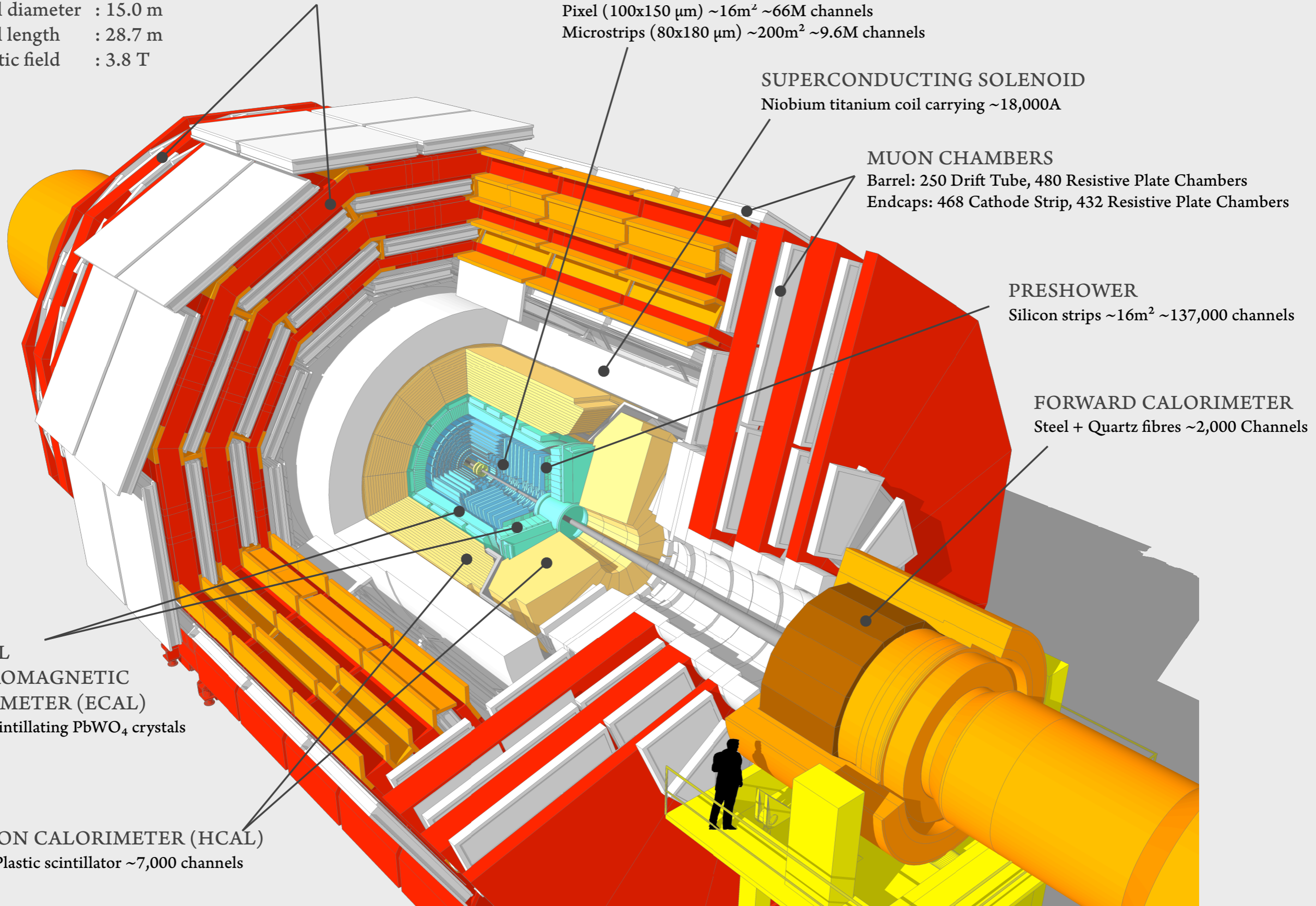
MUON CHAMBERS  
Barrel: 250 Drift Tube, 480 Resistive Plate Chambers  
Endcaps: 468 Cathode Strip, 432 Resistive Plate Chambers

PRESHOWER  
Silicon strips  $\sim 16\text{m}^2 \sim 137,000$  channels

FORWARD CALORIMETER  
Steel + Quartz fibres  $\sim 2,000$  Channels

CRYSTAL  
ELECTROMAGNETIC  
CALORIMETER (ECAL)  
 $\sim 76,000$  scintillating  $\text{PbWO}_4$  crystals

HADRON CALORIMETER (HCAL)  
Brass + Plastic scintillator  $\sim 7,000$  channels

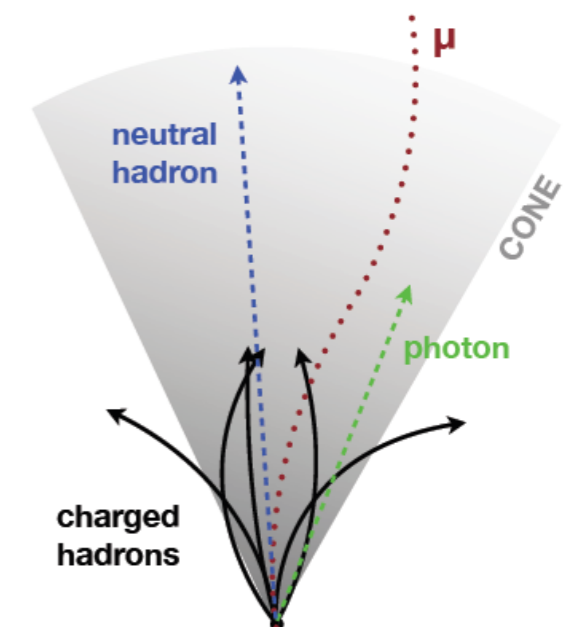
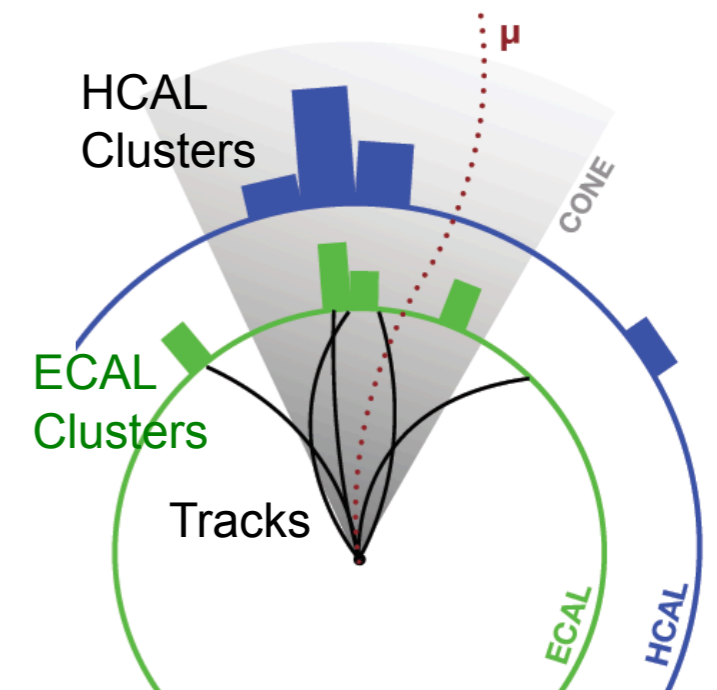


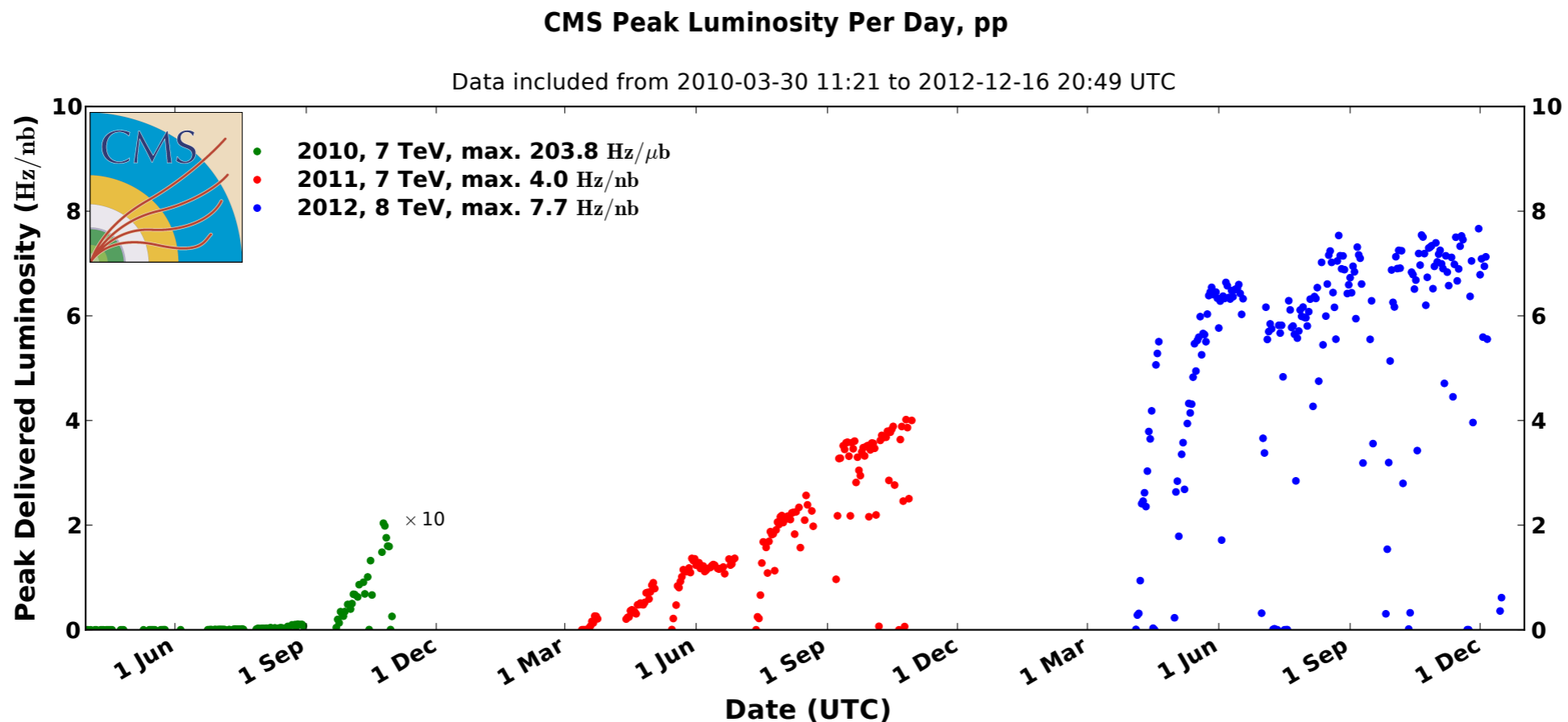
## Online Event Selection

- LHC makes bunches of protons ( $10^{11}$  protons in each bunch) cross each other at CMS at 20~40 MHz
- **L1 Trigger (custom hardware processors, underground P5)**
  - selects interesting events based on signals from muon systems and calorimeters, reducing the event rate to 100 kHz
- **HLT, High-Level Trigger (computing farm, surface building at P5)**
  - reconstructs full events and selects interesting events, reducing the event rate further down to 300~500 Hz, the recording rate of the storage system

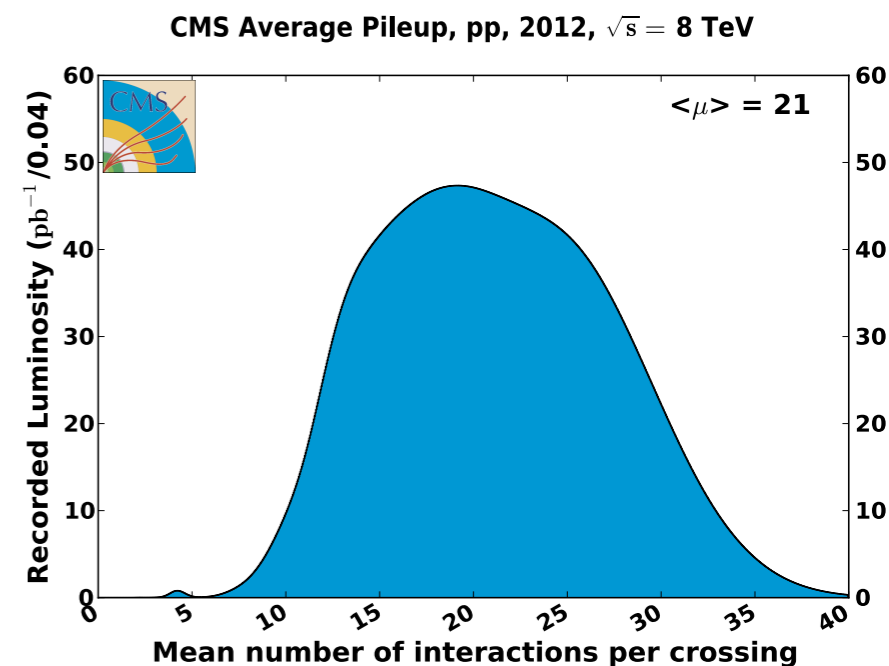
- **Offline Event Reconstruction (on grid, ~60 computing sites worldwide)**

- **Global Event Description, GED**
- Visible particles (muons, electrons, charged hadrons, photons, neutral hadrons)
  - reconstructed and identified by **particle-flow (PF) algorithm**, which uses all CMS detector subsystems, i.e., trackers, calorimeters, muon systems
- **Jets**
  - defined as sets of particles clustered by jet-clustering algorithms, e.g., anti- $k_T$
  - corrected for detector effects (jet energy corrections, JEC)
  - can be tagged to indicate possible origins, e.g., b-quarks, tau leptons, boosted-W bosons, boosted-top quarks.
- **MET (missing transverse momentum)**
  - reconstructed from all jets to which JEC is applied and all remaining visible unclustered particles reconstructed by the PF algorithm
  - cleaned for detector noise, cosmic rays, beam halos and corrected for pile-up events, detector mis-alignment



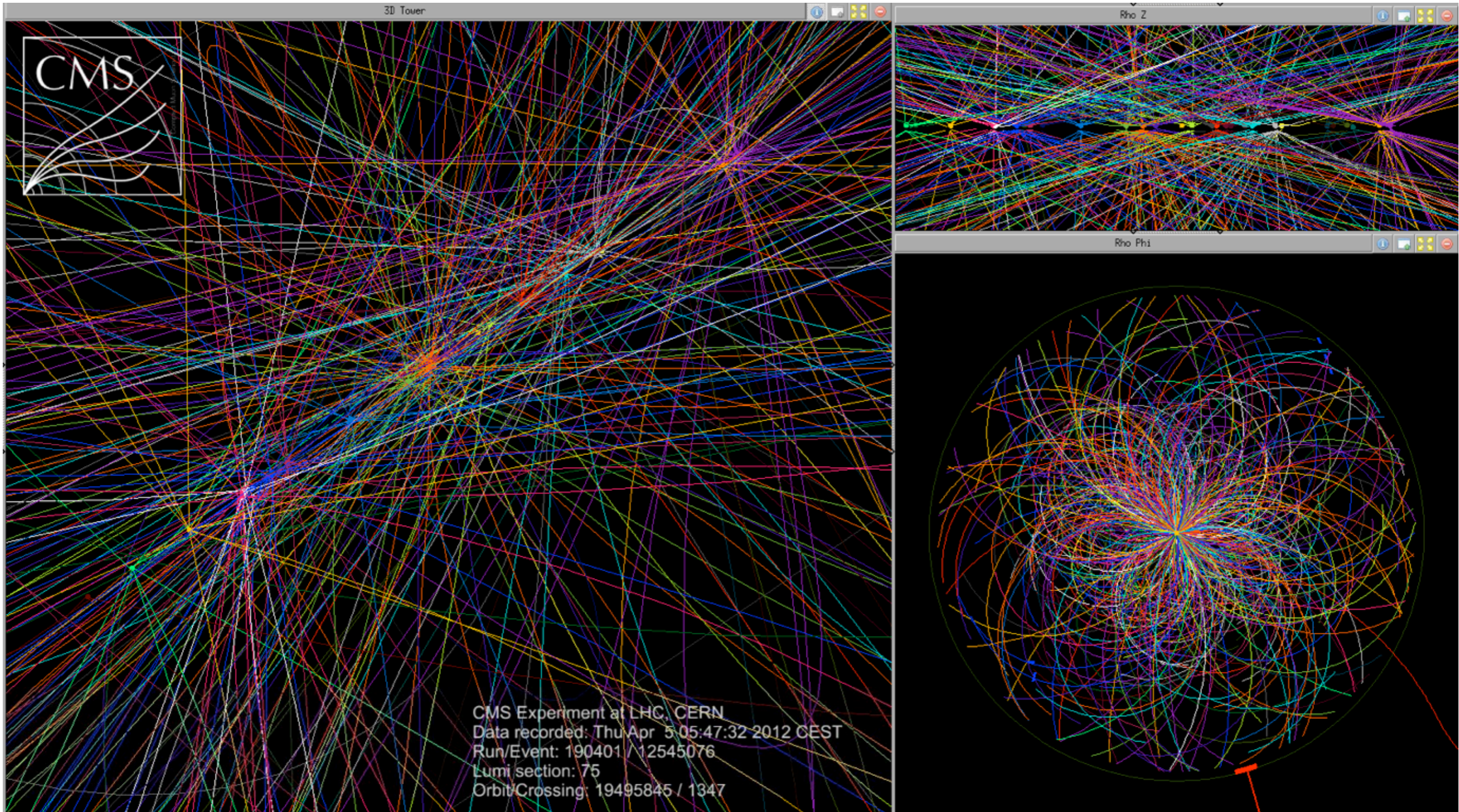


Inelastic proton-proton cross section at 8 TeV:  $\sim 70$  mb  
 for example, if the luminosity is  $7\text{nb}^{-1}/\text{s}$ ,  
 $490\text{M}(= 7\text{nb}^{-1} \times 70 \text{ mb})$  interactions per second  
 $20\text{M}$  times proton bunches cross each other per second (when bunch spacing is 50ns)  
 The average numbers of interaction per crossing (**pile-up events**) would be  $24.5(=490/20)$



note:  $\text{nb}^{-1}/\text{s} = 10^{33}\text{cm}^{-2}/\text{s}$





<http://cms.web.cern.ch/news/new-world-record-first-pp-collisions-8-tev>

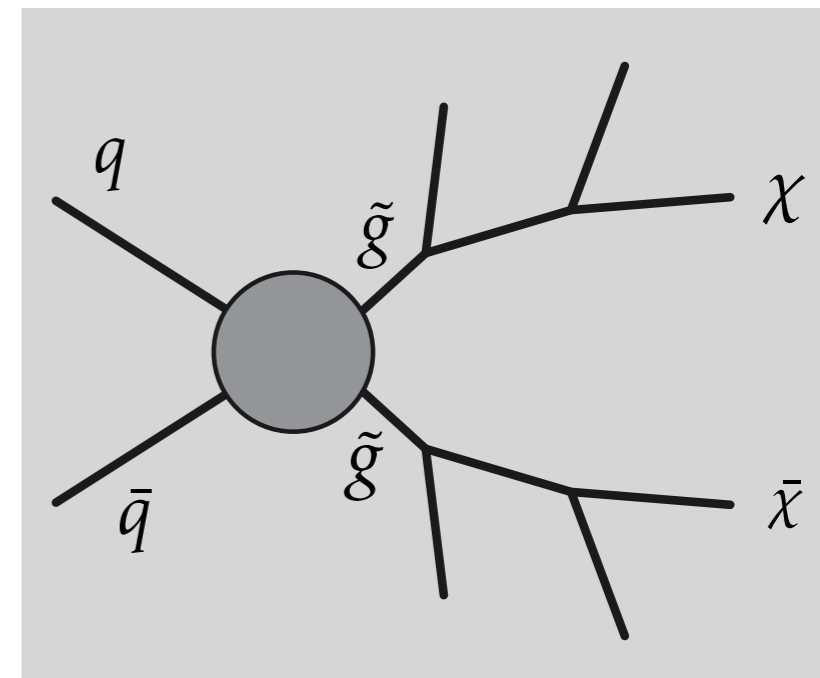
The image above is the CMS event display showing 29 vertices reconstructed, which correspond to 29 distinct proton-proton collisions in the same bunch crossing.

Charged particles produced in the interesting collision are easy to identify from the vertices

CMS develops various techniques to mitigate the effect of neutral particles produced in the pile-up events

- SUSY

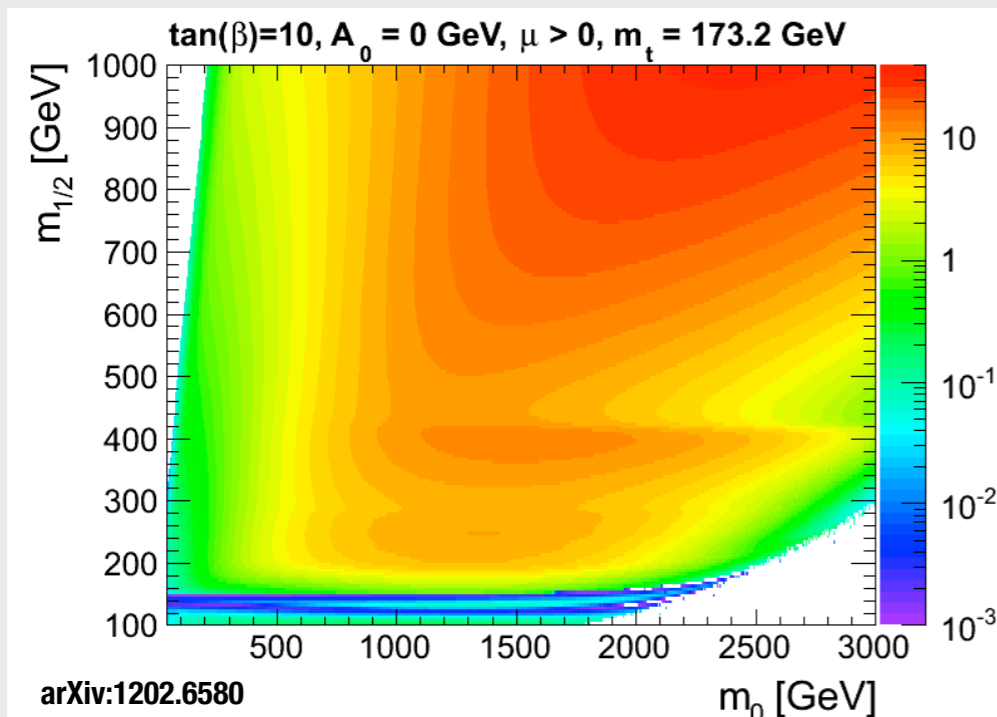
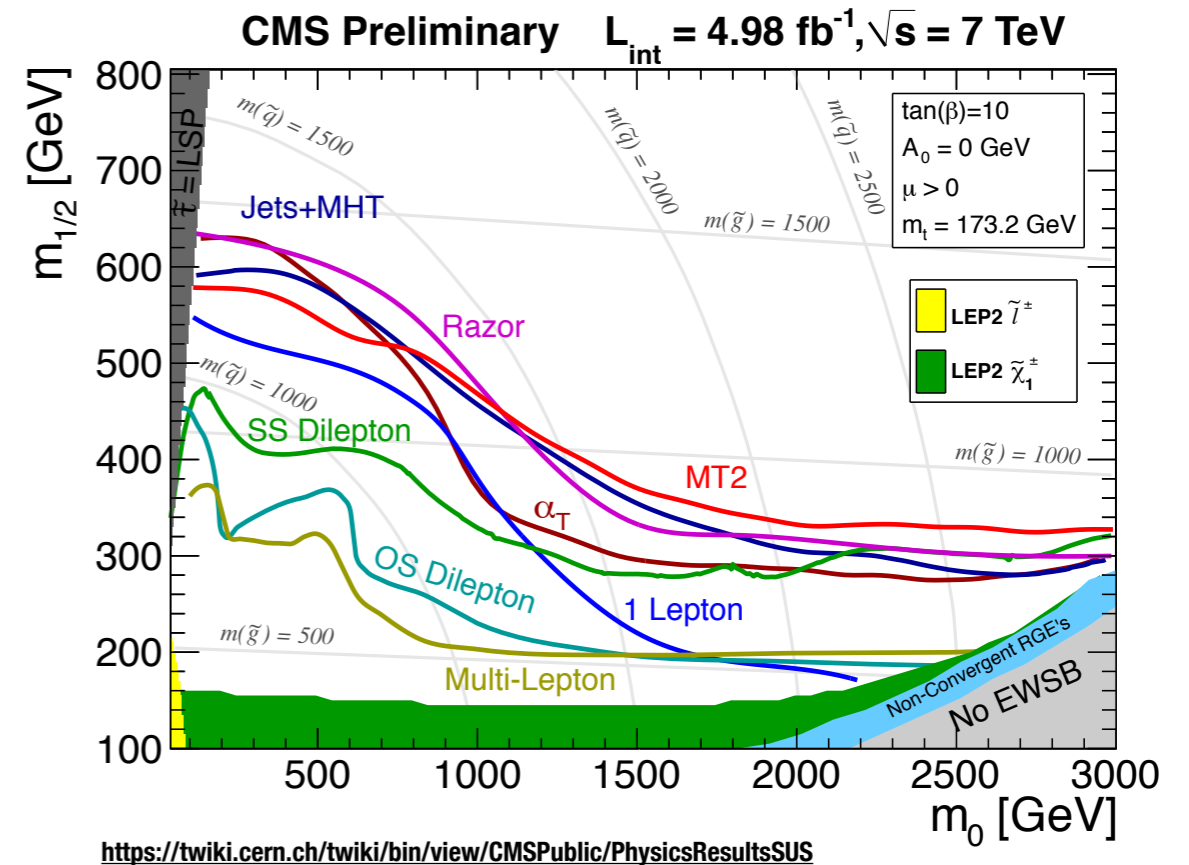
- a hypothetical extension of space-time symmetry
- solves several outstanding problems in particle physics
  - ➔ stabilizes radiative corrections to Higgs boson mass
  - ➔ unifies the three gauge couplings
  - ➔ provides a dark matter candidate
- a broken symmetry as the exact symmetry contradicts the real world
- has many parameters ( $>100$ ), most of which are for its breaking mechanism
  - impractical to constrain by experimental data
- further hypotheses can reduce the number of parameters, which could be explored at LHC
  - e.g.,  $m_0 - m_{1/2}$  plane in CMSSM



- SUSY searches at CMS

- SUSY models predict that SUSY particles will be pair-produced at LHC and each of their decay chains will end with a LSP
  - the final states will contain large MET, jets, and possibly leptons and photons
- CMS selects events with such final states. Some analyses use discriminating variables such as  $\alpha_T$ ,  $M_{T2}$ , *razor* variables
- CMS evaluates if the rate of such events exceed the SM predictions
  - the SM predictions are estimated by data-driven methods, MC simulation
- No such excess has been observed by now
- The results have been interpreted in terms of CMSSM, other SUSY models, and Simplified Models (SMS)
  - exclusion limits have been placed on parameters of the models

- CMSSM is one of the popular SUSY models in interpreting the results of CMS SUSY analyses
- CMSSM adds five new parameters to SM
  - $m_0, m_{1/2}, \tan\beta, A_0, \text{sign}(\mu)$
- In most of the its parameter spaces, LSP is a neutralino, a dark matter candidate
- The right figure summarizes the exclusion limits of several CMS SUSY analyses on the  $m_0 - m_{1/2}$  plane for  $\tan\beta=10, A_0=10\text{GeV}$ , and positive  $\mu$

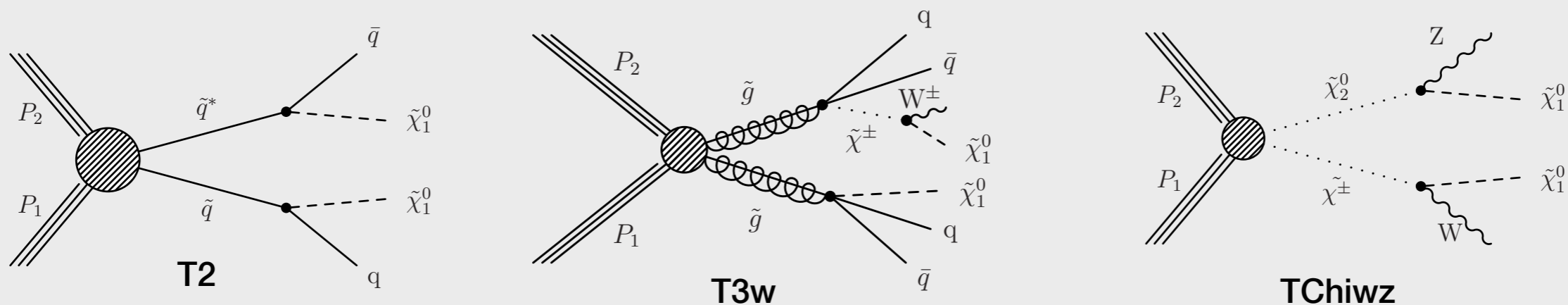


## Dark matter relic density

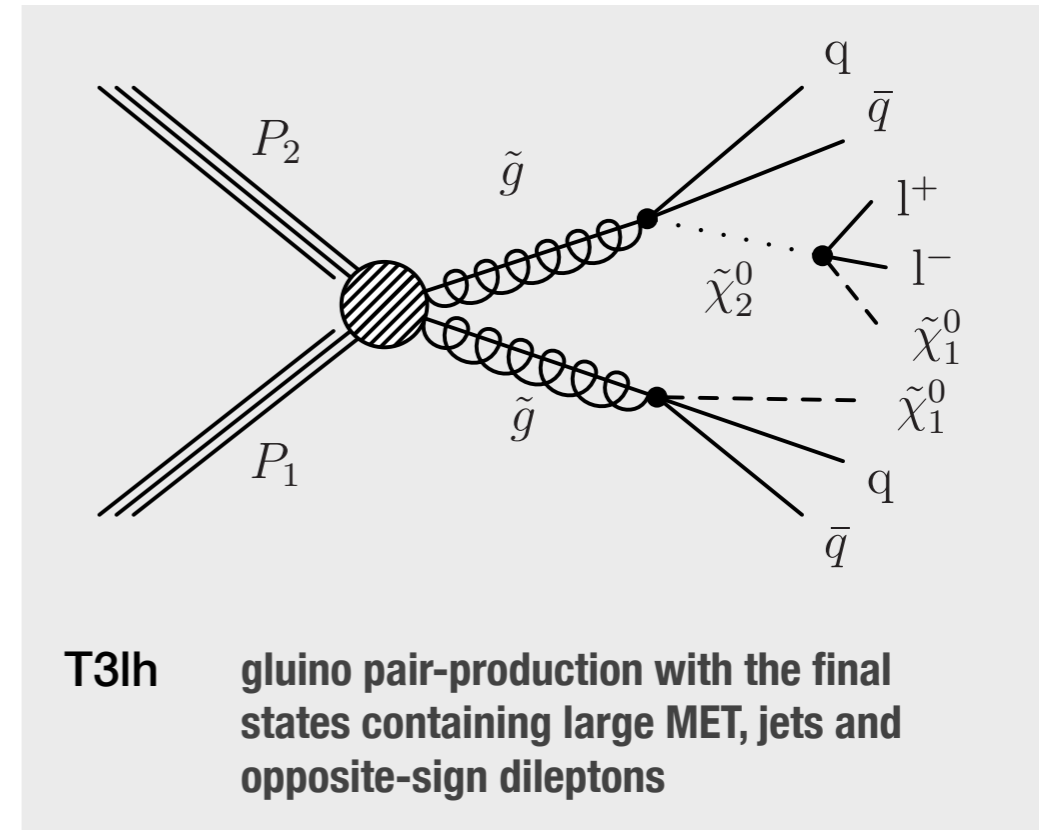
- the corresponding neutralino relic density  $\Omega_{\tilde{\chi}_1^0} h^2$  on the  $m_0 - m_{1/2}$  plane can be calculated under certain assumptions (arxiv:1202.6580)
- which can be compared with the cold dark matter relic density derived from cosmological observations
  - $\Omega_{\text{cdm}} h^2 = 0.112 \pm 0.006$  [Phys. Rev. Derived 86, 010001 (2012)]

- A **simplified model** is defined by a few new particles and their production and decay
  - New particles have the same names as the SUSY particles
    - e.g., *gluinos*, *squarks*, *neutralinos*, *charginos*, *LSP* (neutral)
  - New particles are produced in pairs. Each decay chain ends with LSP
- SUSY analysis results are used to set upper limits on the product of the production cross section and branching fraction ( $\sigma \times \text{BR}$ ) as a function of new particle masses
- SMS allows us to explore wider kinematic phase space of final states than particular models predict
- The upper limits can be compared with predictions of particular SUSY (or non-SUSY) models
  - This comparison can be used to set limits on new particle masses

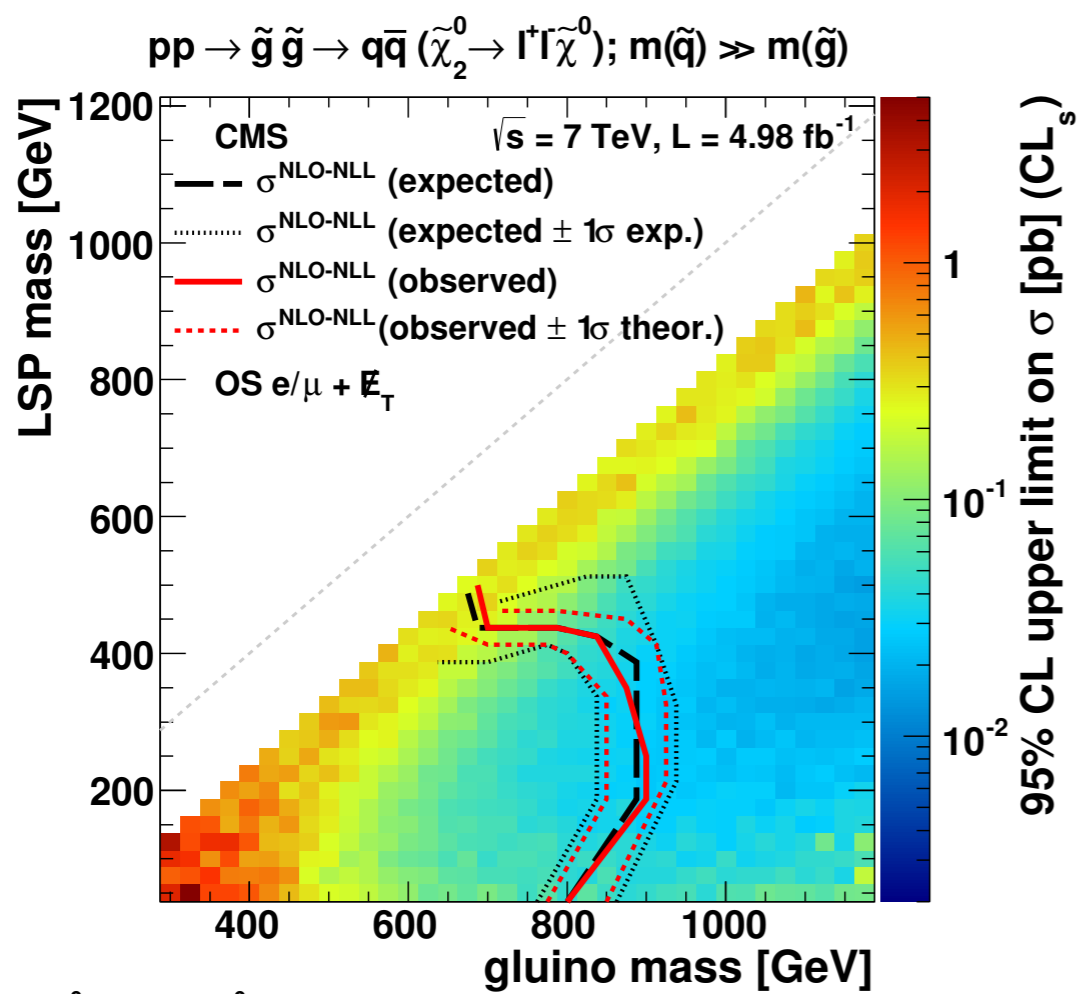
## Example simplified models



- An interpretation of a CMS SUSY result [*Phys.Lett. B718 (2013) 815*] in the SMS T3lh



**T3lh** gluino pair-production with the final states containing large MET, jets and opposite-sign dileptons



The color scale shows the upper limit on  $\sigma \times BR$  (the product of the cross section and the branching fraction)

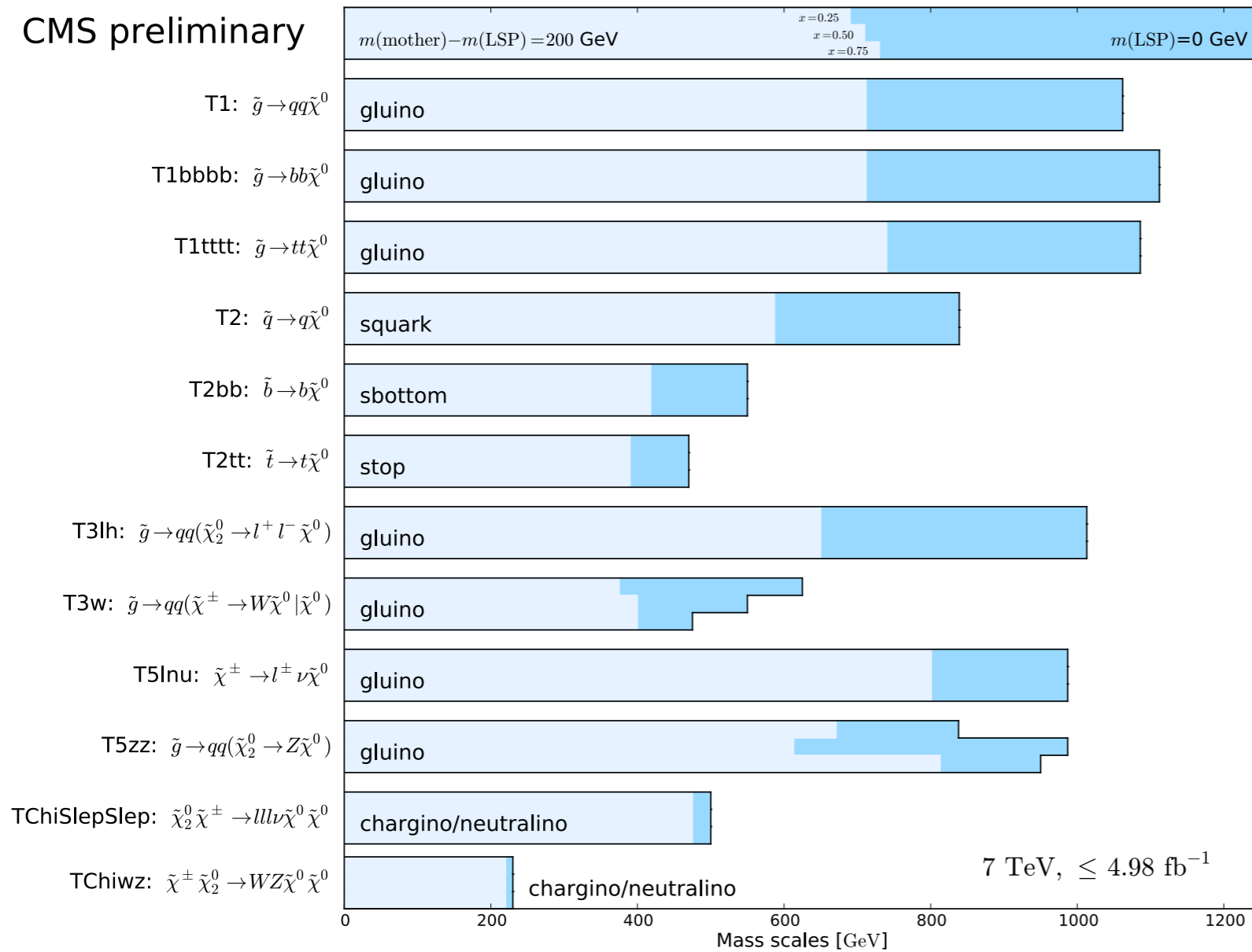
The lines show the contour at which the reference cross sections calculated in SUSY NLO+NLL intersect with the upper limit, indicating the lower mass limits in the reference theory

$m(\tilde{\chi}_2^0) = 0.5 m(\tilde{\chi}_1^0) + 0.5 m(\tilde{g})$

arXiv:1301.2175

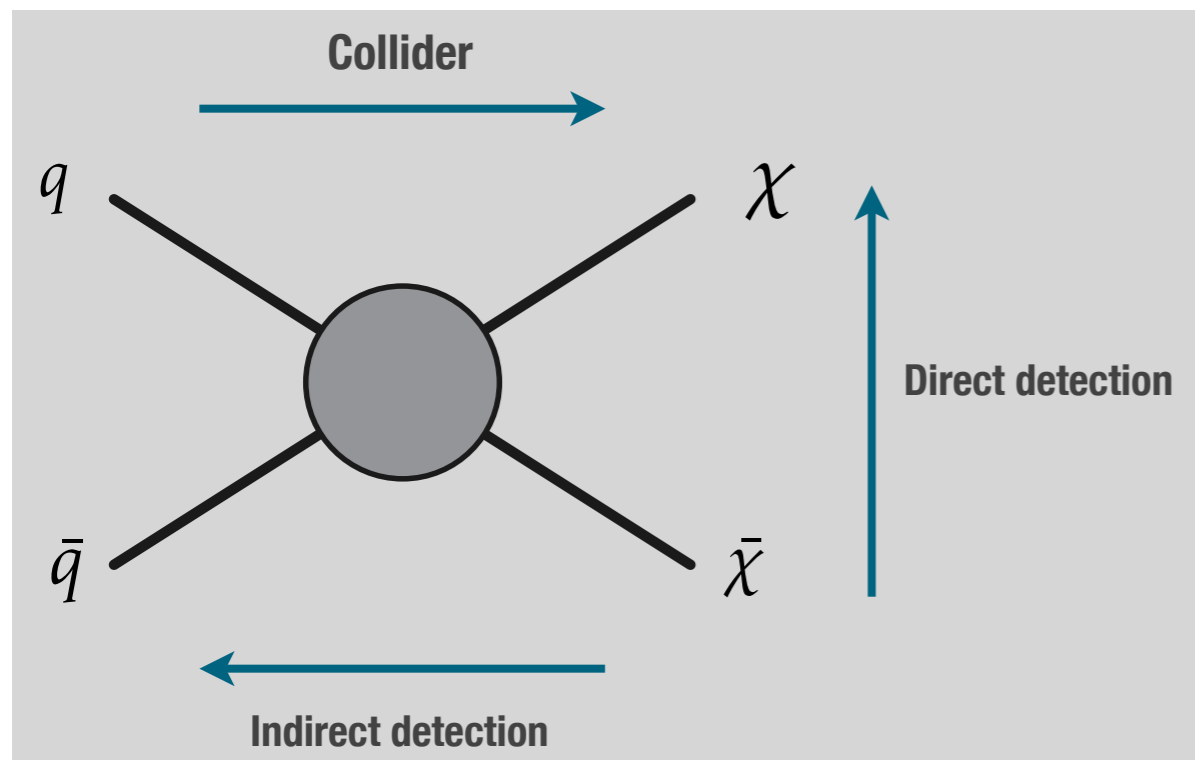
<https://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResultsSUS11016>

CMS preliminary

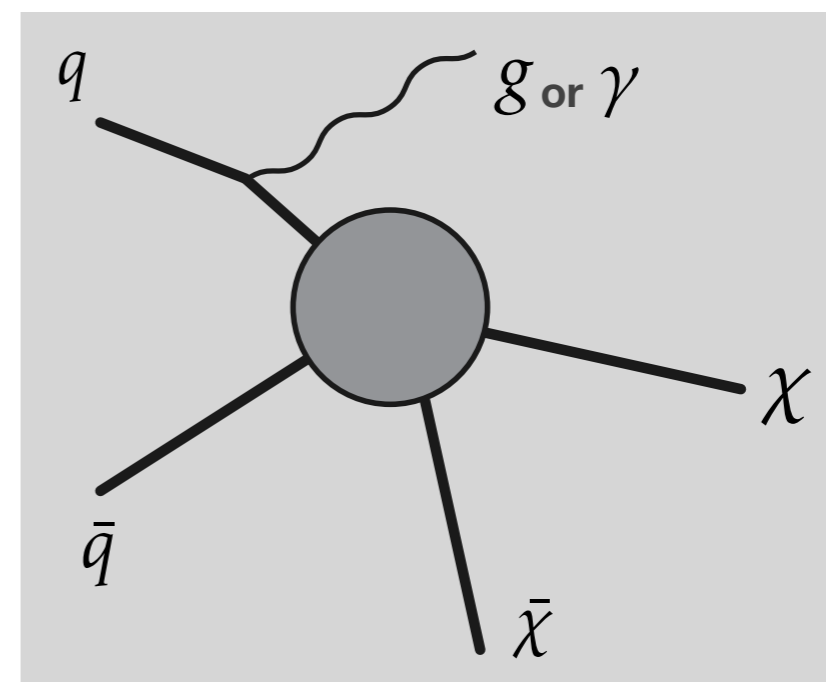


This bar chart summarizes exclusion limits of masses of the produced new particle in each SMS with the reference theory in the cases in which LSP is massless and in which LSP is 200 GeV lighter than the produced particle. For the SMSs with intermediate new particles, limits with up to three different assumptions of the intermediate particles are shown

Instead of being at the end of decay chains, dark matter particles could be directly produced in pairs



The DM pair-production event can be tagged if it occurs after ISR (initial-state radiation)



It will appear as monojet or mono-photon event

As the coupling between SM and DM can be evaluated, results can be compared with direct detection results

- The event rates did not exceed the standard model and background expectation in the analysis of  $5.0 \text{ fb}^{-1}$  of data at  $\sqrt{s}=7 \text{ TeV}$

• **The dark matter-nucleon cross section limit:**

- The following assumptions were made in the limit setting:
  - The mediator is heavy (an effective contact operator)
  - The dark matter particles are Dirac fermions
  - The interaction is vector or axial-vector interaction
- The lower limits were set on the cutoff  $\Lambda$
- The upper limits were set on the dark matter-nucleon cross section
  - The vector interactions correspond the spin-independent interactions
  - The axial-vector interactions correspond the spin-dependent interactions
- The limits were compared with the limits from the direct detection results

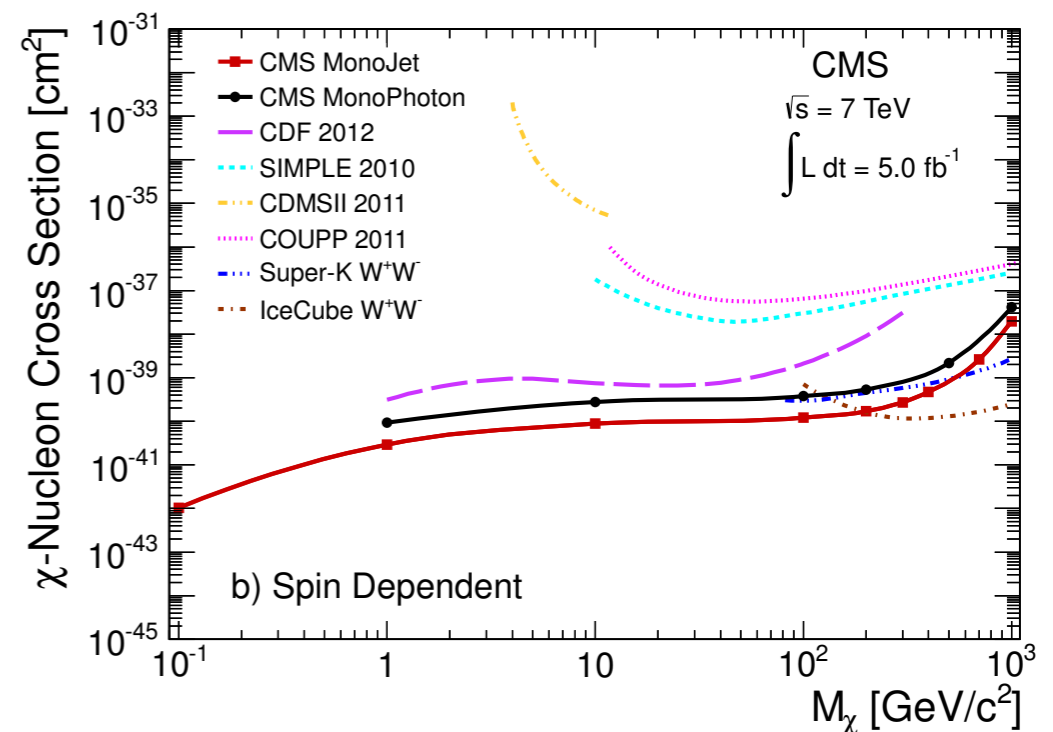
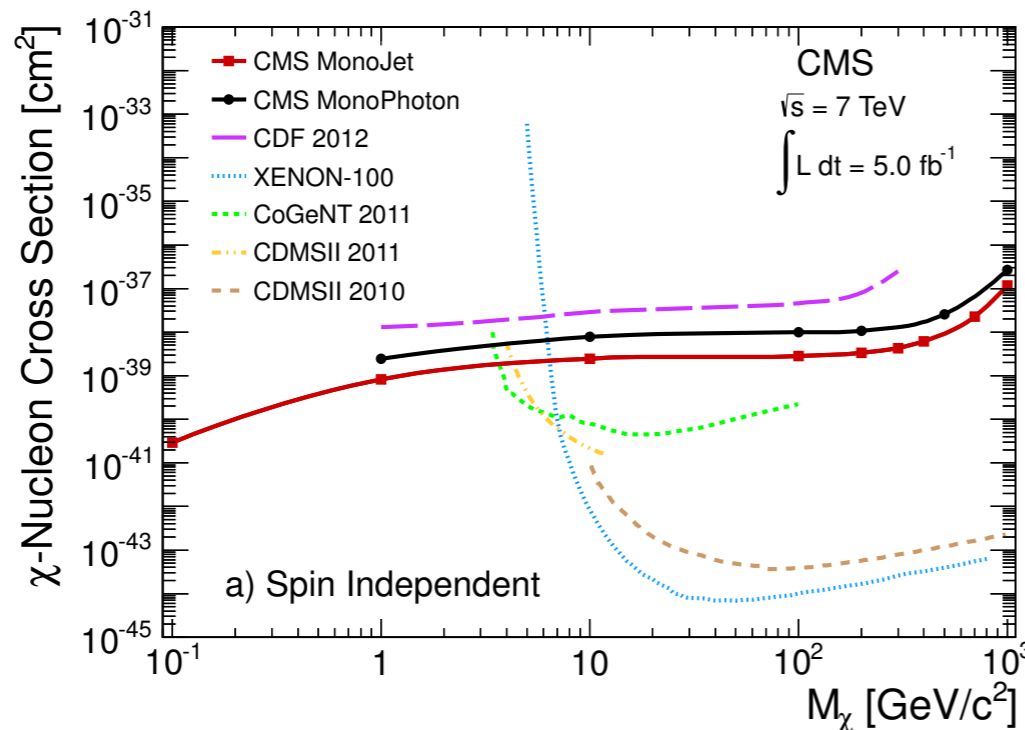
**vector interaction**

$$\mathcal{O}_V = \frac{(\bar{\chi}\gamma_\mu\chi)(\bar{q}\gamma^\mu q)}{\Lambda^2}$$

**axial-vector interaction**

$$\mathcal{O}_A = \frac{(\bar{\chi}\gamma_\mu\gamma_5\chi)(\bar{q}\gamma^\mu\gamma_5q)}{\Lambda^2}$$

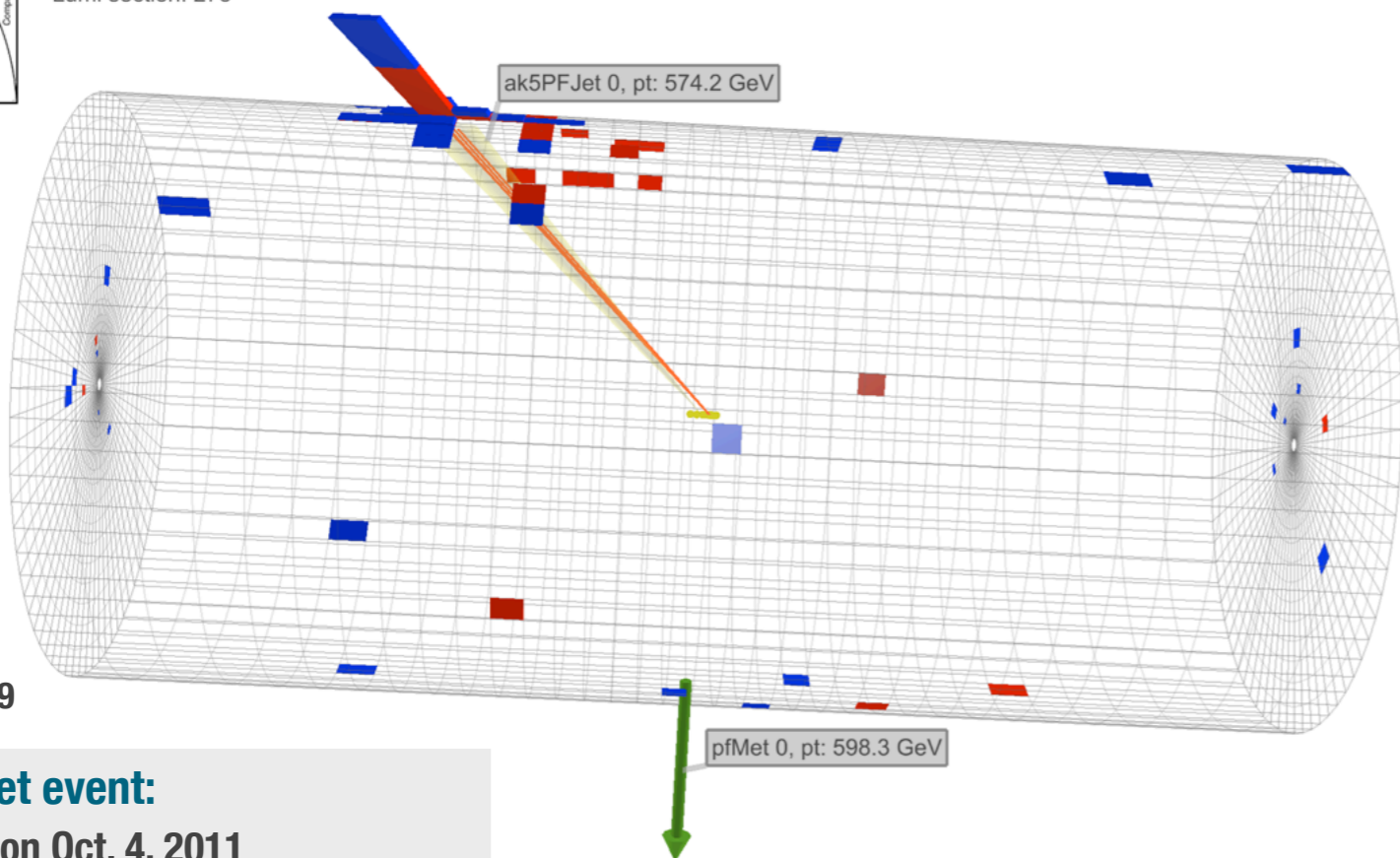
**cutoff**  $\Lambda \equiv M/\sqrt{g_\chi g_q}$







CMS Experiment at LHC, CERN  
Data recorded: Tue Oct 4 02:50:32 2011 CEST  
Run/Event: 177783 / 442962676  
Lumi section: 273

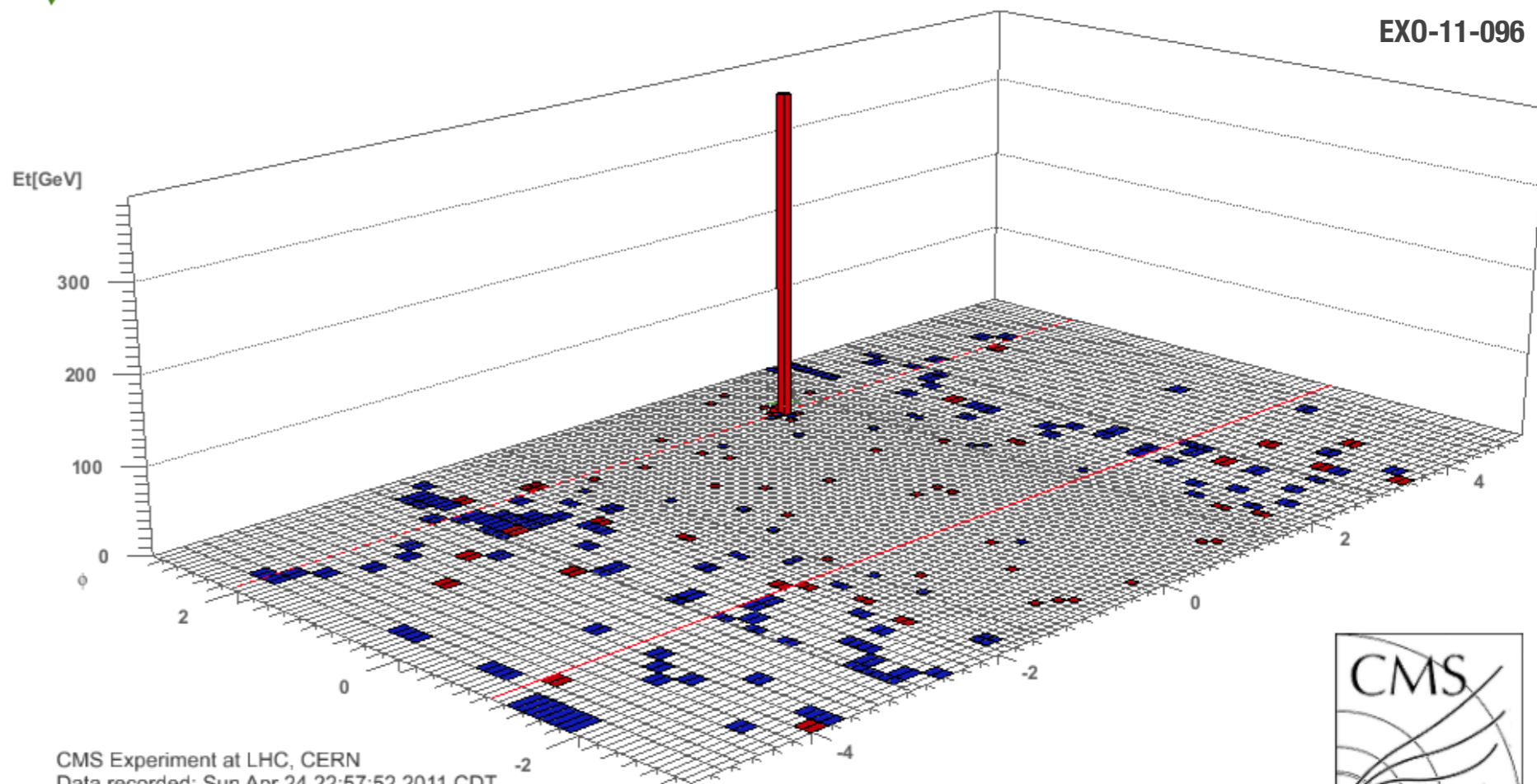


EX0-11-059

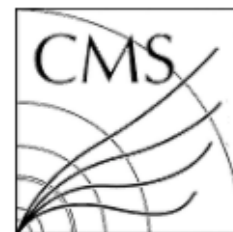
**A monojet event:**  
recorded on Oct. 4, 2011  
 $p_T(\text{jet})=574.2 \text{ GeV}$ ,  $\text{MET} = 598.3 \text{ GeV}$

**A mono-photon event:**  
recorded on Apr. 24, 2011  
 $p_T(\text{photon})=384 \text{ GeV}$ ,  $\text{MET} = 407 \text{ GeV}$

EX0-11-096



CMS Experiment at LHC, CERN  
Data recorded: Sun Apr 24 22:57:52 2011 CDT  
Run/Event: 163374 / 314736281  
Lumi section: 604



## Summary

- This presentation showed the results of dark matter searches in up to  $5\text{fb}^{-1}$  of proton-proton collision data at  $\sqrt{s}=7$  TeV collected with the CMS detector at LHC in 2011
- If dark matter particles are produced in the collision, they will be invisible to the CMS detector; their existence can be inferred by large MET
- The results of CMS SUSY searches were summarized as exclusion limits on a  $m_0$ - $m_{1/2}$  plane in CMSSM. The limits can be compared with calculated corresponding neutralino relic density on the plane and the cold dark matter density derived from cosmological observations
- CMS SUSY search results were interpreted in Simplified Models (SMSes), in which LSPs are dark matter candidates. In each SMS, upper limits were placed on the product of the production cross section and branching fraction. With the benchmark theory, lower limits were set on the mass of the produced new particles in the SMS.
- Direct productions of the pairs of dark matter particles can be tagged by the initial-state radiation (ISR). CMS monojet and mono-photon analyses searched for events with only the final state of ISR and large MET and placed upper limits on the dark matter-nucleon cross sections, which can be compared with direct detection results.
- We are currently analyzing  $20\text{fb}^{-1}$  of proton-proton collision data at  $\sqrt{s}=8\text{TeV}$  collected in 2012 and continuing the search



*End*