

# Searches for dark matter and dark forces at *BABAR*

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*Caltech*

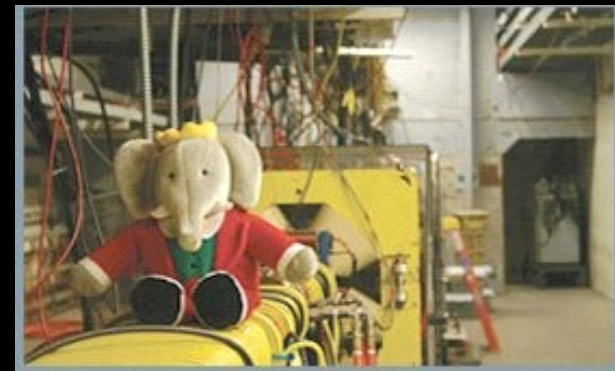
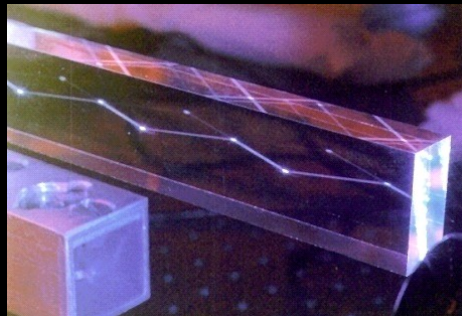
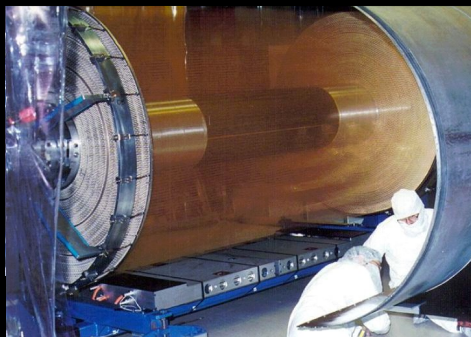
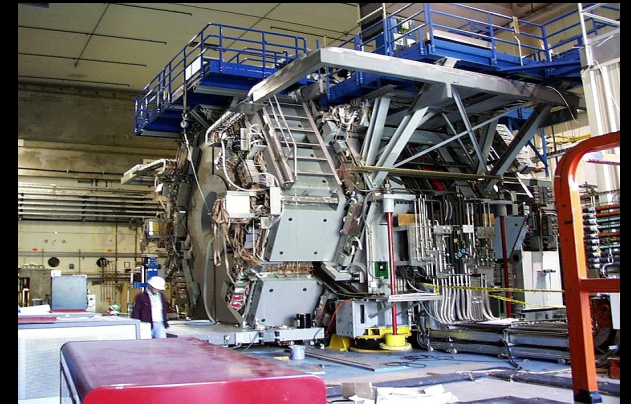
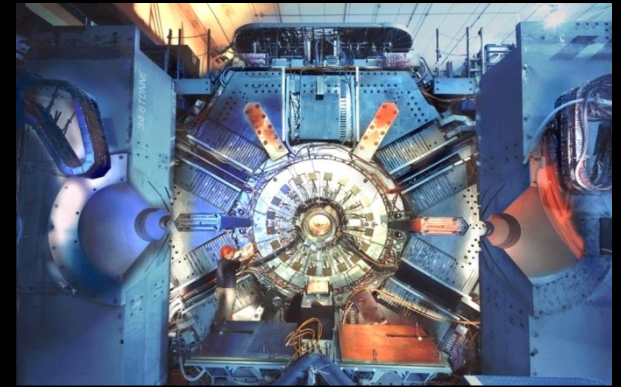
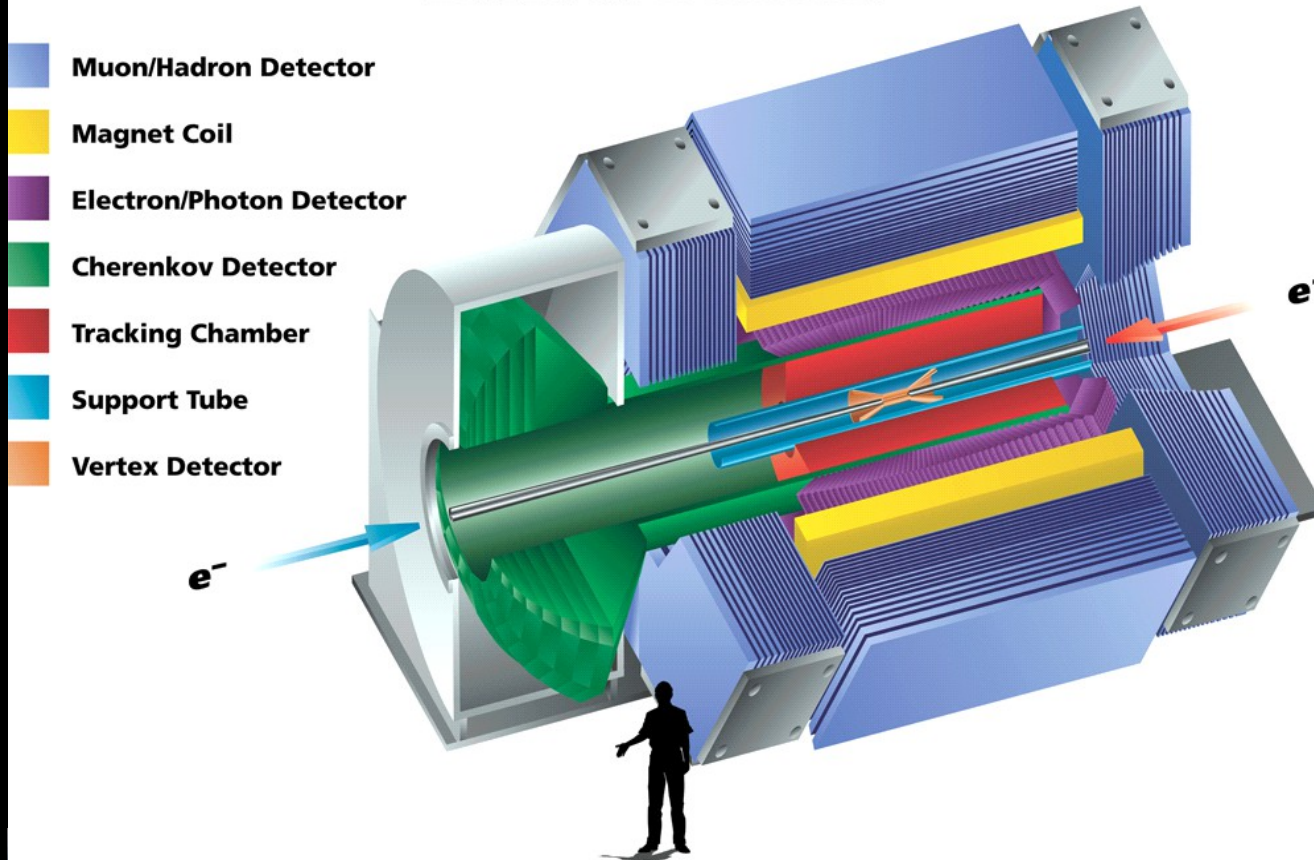
*on behalf of the BABAR collaboration*

*Closing in on dark matter*

*Aspen, January 2013*

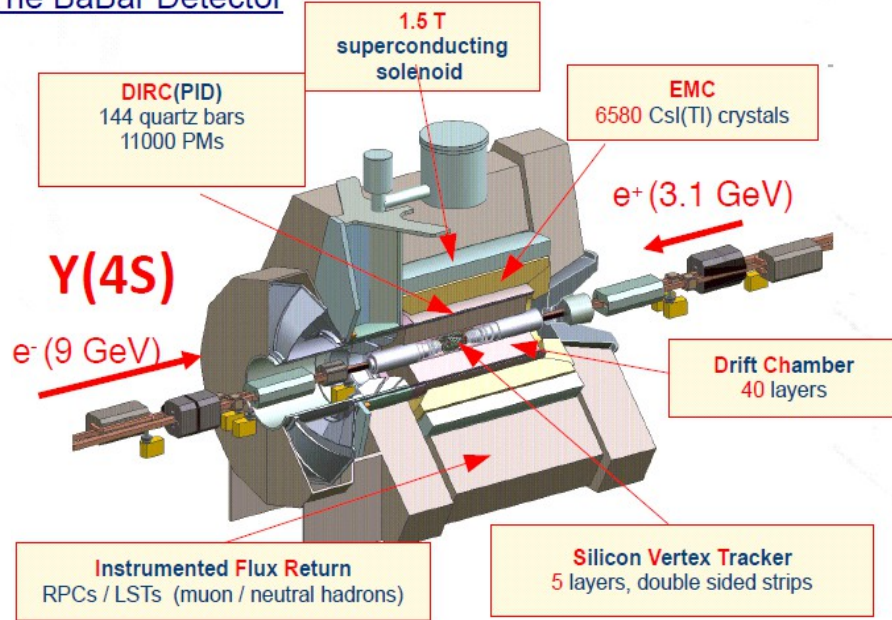
# The BABAR experiment at SLAC

## BABAR Detector

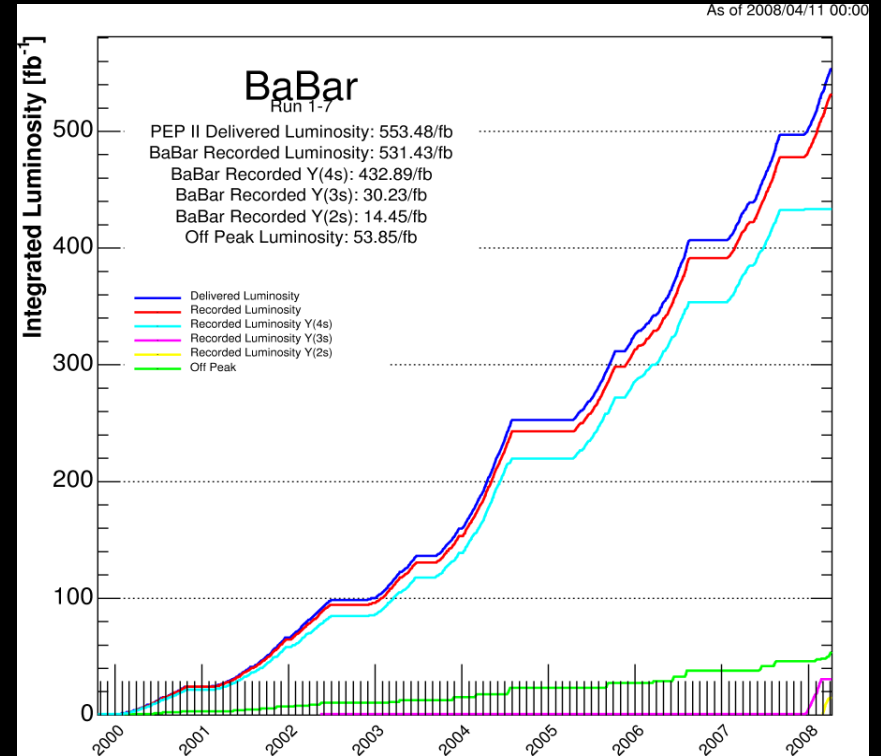
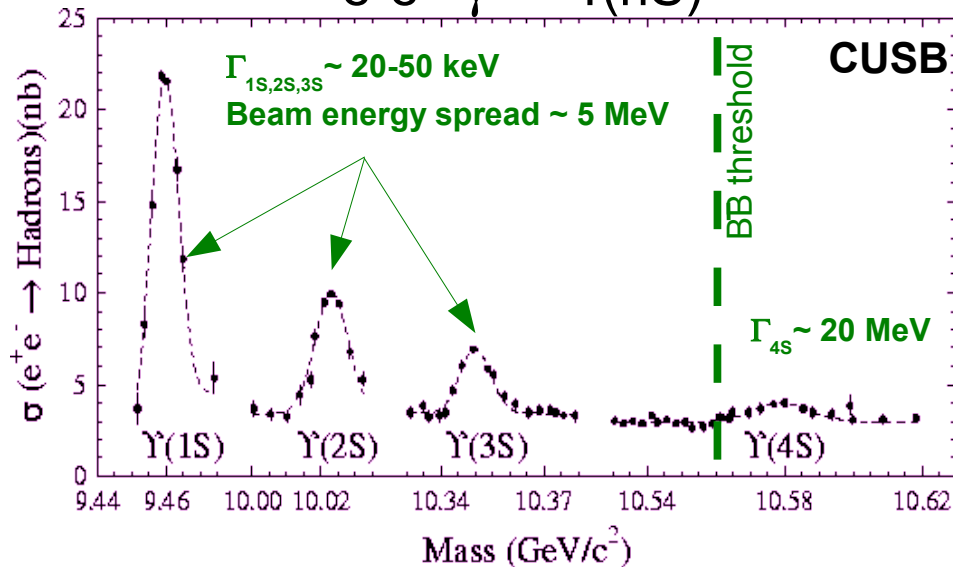


# The BABAR experiment at SLAC

## The BaBar Detector



$$e^+e^- \rightarrow \gamma^* \rightarrow \Upsilon(nS)$$



## BABAR data sample contains

- $\sim 470 \times 10^6$   $\Upsilon(4S)$
- $\sim 120 \times 10^6$   $\Upsilon(3S)$  (10x Belle)
- $\sim 100 \times 10^6$   $\Upsilon(2S)$  (10x CLEO)
- $\sim 18 \times 10^6$   $\Upsilon(1S)$  from  $\Upsilon(2S) \rightarrow \pi^+\pi^-\Upsilon(1S)$



Search for dark matter  
in Upsilon decays

## Generic Dark Matter model

- ⇒ Minimal model introducing a light dark matter particle  $\chi$  ( $m_\chi < m_\gamma/2$ ) and a new scalar or gauge boson  $A'$  to serve as an s-channel annihilation mediator
- ⇒ Scalar/boson  $A'$  couples to SM via Higgs mixing or kinetic mixing with SM hypercharge
- ⇒ Could explain 511 keV gamma rays excess from the galactic center reported by INTEGRAL and electron/positron excess by PAMELA/FERMI
- ⇒ Could increase the invisible decay width of the  $\Upsilon(1S)$  predicted by SM<sup>1)</sup> by orders of magnitude. Rate estimates are fairly model independent, based on cosmological observations and assuming time-reversal symmetry

## Rate predictions

$$\text{BF}(\Upsilon(1S) \rightarrow \chi\chi) \sim 4.2 \times 10^{-4} \text{ (s-wave)} \quad (\text{PRD 72, 103508 (2005)})$$

$$\text{BF}(\Upsilon(1S) \rightarrow \chi\chi) \sim 1.8 \times 10^{-3} \text{ (p-wave)} \quad (\text{PRD 72, 103508 (2005)})$$

$$\text{BF}(\Upsilon(1S) \rightarrow \nu\nu) \sim 9.9 \times 10^{-6} \quad (\text{PLB 441, 419 (1998)})$$

**Large increase from SM predictions**

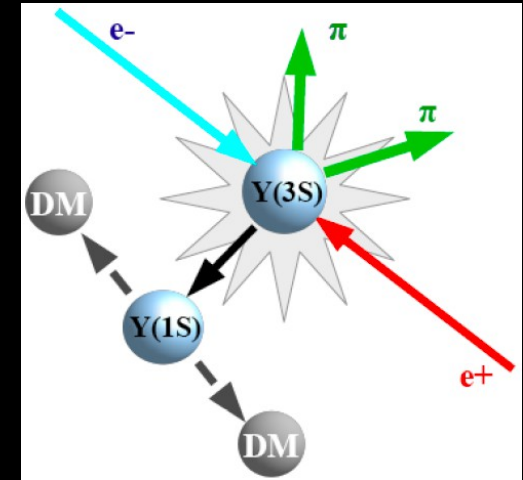
1) PLB 441, 419 (1998).

## Analysis strategy

- ⇒ Tag  $\Upsilon(1S)$  mesons in  $\Upsilon(3S) \rightarrow \pi^+\pi^-\Upsilon(1S)$  transition
- ⇒ Select event containing two oppositely-charged tracks only (no extra activity)

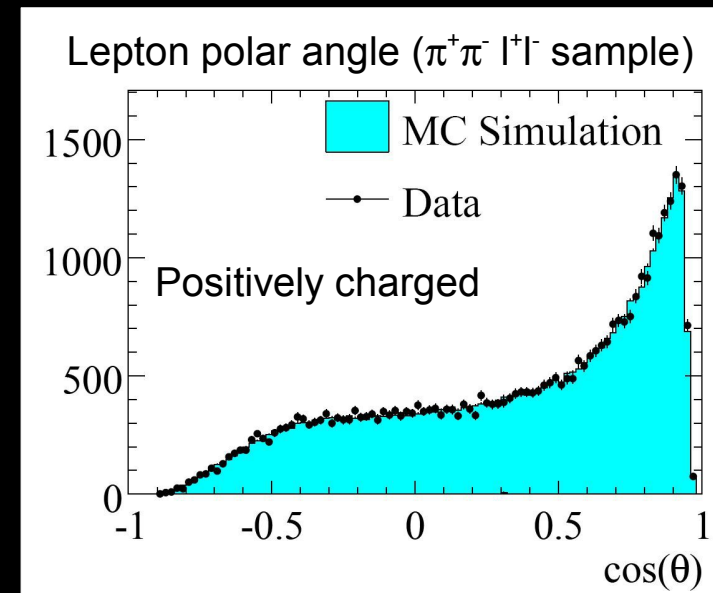
## Data sample contains

- ⇒ Non-peaking background from random  $\pi^+\pi^-$  combinations
- ⇒ Peaking background (indistinguishable from signal)  
 $\Upsilon(3S) \rightarrow \pi^+\pi^-\Upsilon(1S)$ ,  $\Upsilon(1S) \rightarrow X$  ( $X$  undetected, mainly  $l^+l^-$  pairs)
- ⇒ Signal:  $\Upsilon(3S) \rightarrow \pi^+\pi^-\Upsilon(1S)$ ,  $\Upsilon(1S) \rightarrow \text{invisible}$



## Signal extraction

- ⇒ Fit recoil mass  $M_{\text{rec}} = (s + M_{\pi\pi}^2 - 2sE_{\pi\pi}^*)^{1/2}$ , should peak at  $\Upsilon(1S)$
- ⇒ Subtract peaking background estimated from MC. Use  $\Upsilon(1S, 2S) \rightarrow l^+l^-$  with one or two reconstructed leptons to check and correct simulations



**Fit**

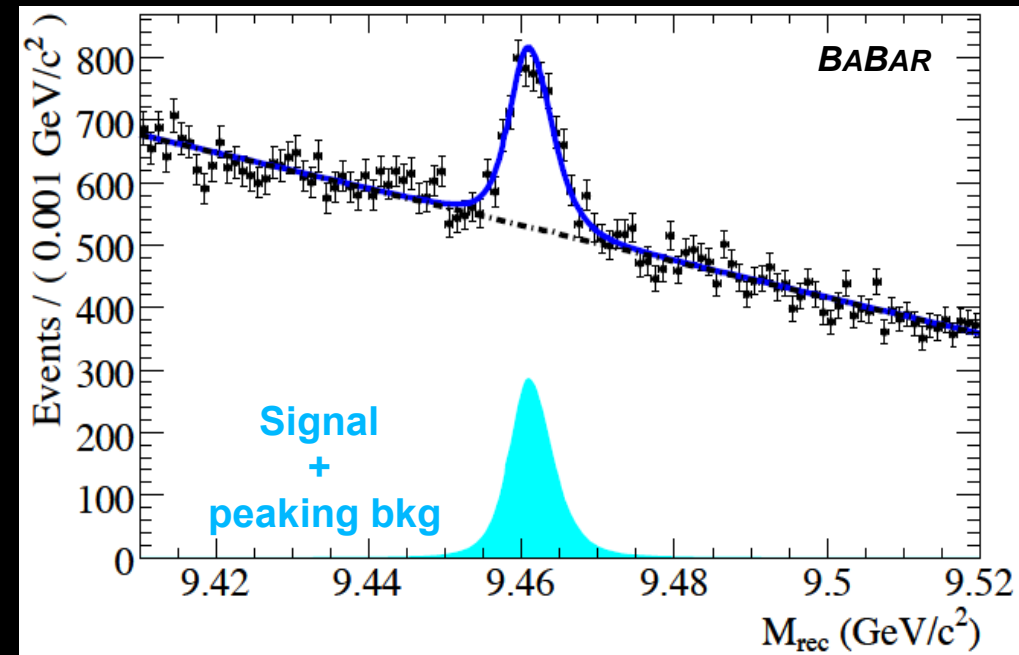
- ⇒ Extended unbinned maximum likelihood fit of recoil mass  $M_{\text{rec}}$
- ⇒ Sum of signal + peaking background
  - Crystal-Ball (Gaussian with power-law tail)
- ⇒ Non-peaking background
  - 1<sup>st</sup> order polynomial

**Results**

Fit yield	$2326 \pm 105$
Peaking bkg (MC)	$2444 \pm 124$
<b>Signal only</b>	<b><math>-118 \pm 105 \pm 124</math></b>

**Upper limit (90% CL)**

$$\text{BF}(\Upsilon(1S) \rightarrow \text{invisible}) < 3.0 \times 10^{-4}$$

**Previous measurements  $\text{BF}(\Upsilon(1S) \rightarrow \text{invisible})$** 

CLEO:	$\text{BF} < 3.9 \times 10^{-3}$ @ 90% CL	PRD 75 (2007) 031104
Belle:	$\text{BF} < 2.5 \times 10^{-3}$ @ 90% CL	PRL 98 (2007) 132001

**No evidence of dark matter contribution to invisible  $\Upsilon(1S)$  decays**

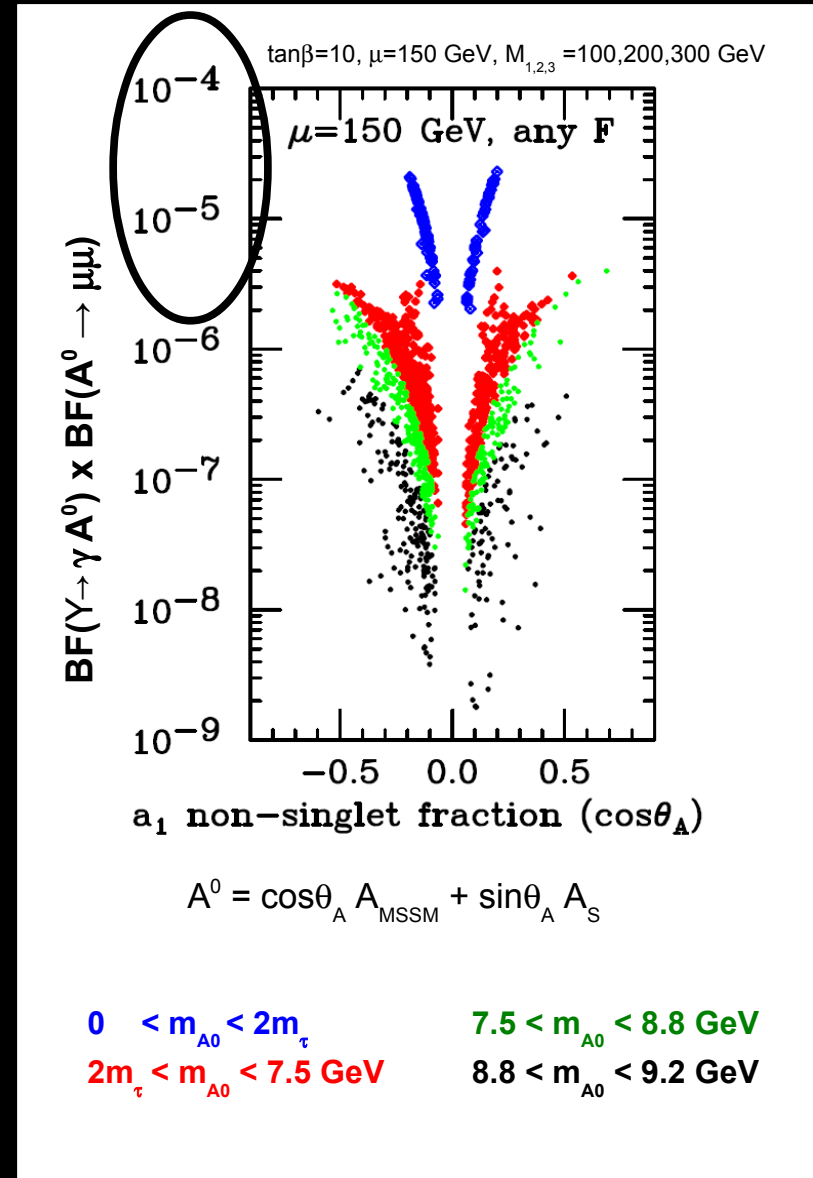


Search for dark matter  
in radiative Upsilon decays



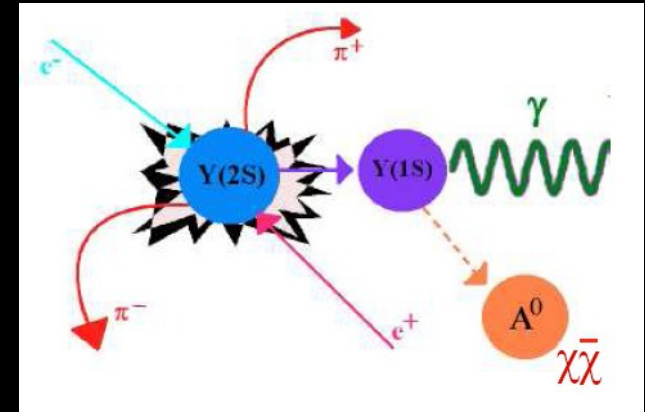
## Dark matter and light Higgs boson

- ⇒ Many SM extensions, such as the NMSSM, include the possibility of a light Higgs boson
- ⇒ NMSSM proposed to solve the “ $\mu$  problem”, adding one CP-odd Higgs, one CP-even Higgs and one neutralino to MSSM content
- ⇒ A light CP-odd Higgs  $A^0$  with mass lower than  $2m_b$  is not excluded by LEP constraints.
- ⇒ Radiative decays  $\Upsilon(nS) \rightarrow \gamma A^0$  ( $n=1,2,3$ ) can have a large branching fraction.
- ⇒ The lightest neutralino  $\chi_1$  is a natural dark matter candidate and the  $A^0$  decays dominantly in  $A^0 \rightarrow \chi_1 \chi_1$  if  $m_{A^0} > 2m_{\chi_1}$
- ⇒ More generally one can look for dark matter production in radiative  $\Upsilon(nS) \rightarrow \gamma \chi \chi$  decays (e.g G.K. Yeghiyan, PRD 80 (2009) 115019)



### Event selection

- ⇒ Tag the dipion transition  $\Upsilon(2S) \rightarrow \pi^+\pi^-\Upsilon(1S)$
- ⇒ Require two charged tracks and one energetic photon, no additional activity
- ⇒ Missing energy and momentum



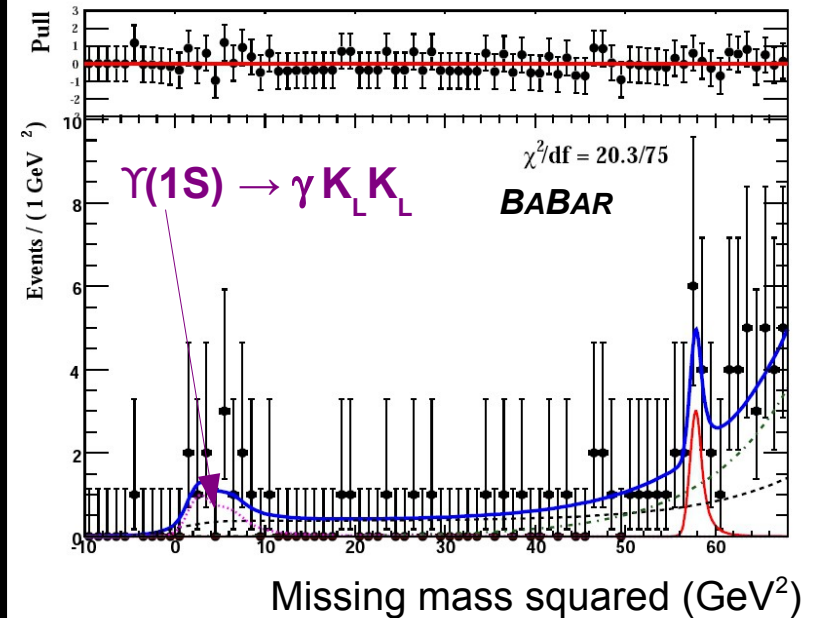
### Signal extraction

- ⇒ Consider both  $\Upsilon(1S) \rightarrow \gamma A^0$  (two-body) and  $\Upsilon(1S) \rightarrow \gamma \chi \chi$  (multi-body) decays
- ⇒ 2D fit the the recoil mass squared ( $M_{\text{recoil}}^2$ ) and missing mass ( $M_X^2$ ) squared

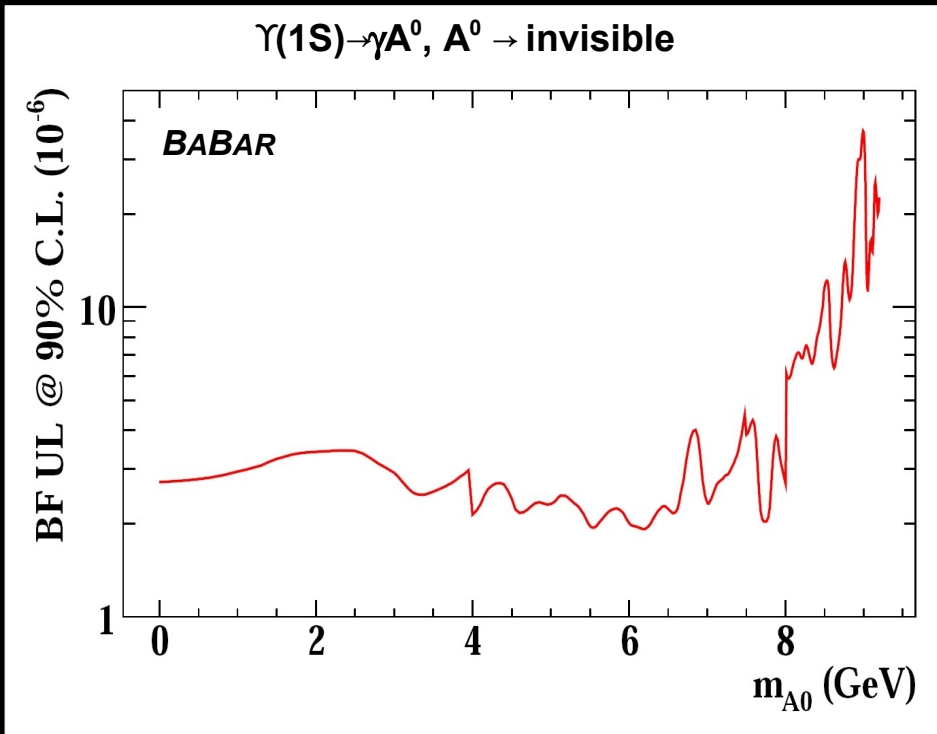
$$M_{\text{recoil}}^2 = M_{\Upsilon(2S)}^2 + m_{\pi\pi}^2 - 2M_{\Upsilon(2S)}E_{\pi\pi}^*$$

$$M_X^2 = (\mathcal{P}_{e^+e^-} - \mathcal{P}_{\pi\pi} - \mathcal{P}_{\gamma})^2$$

Most significant peak ( $m_A = 7.58 \text{ GeV}$ ) with local significance  $2.0\sigma$ .

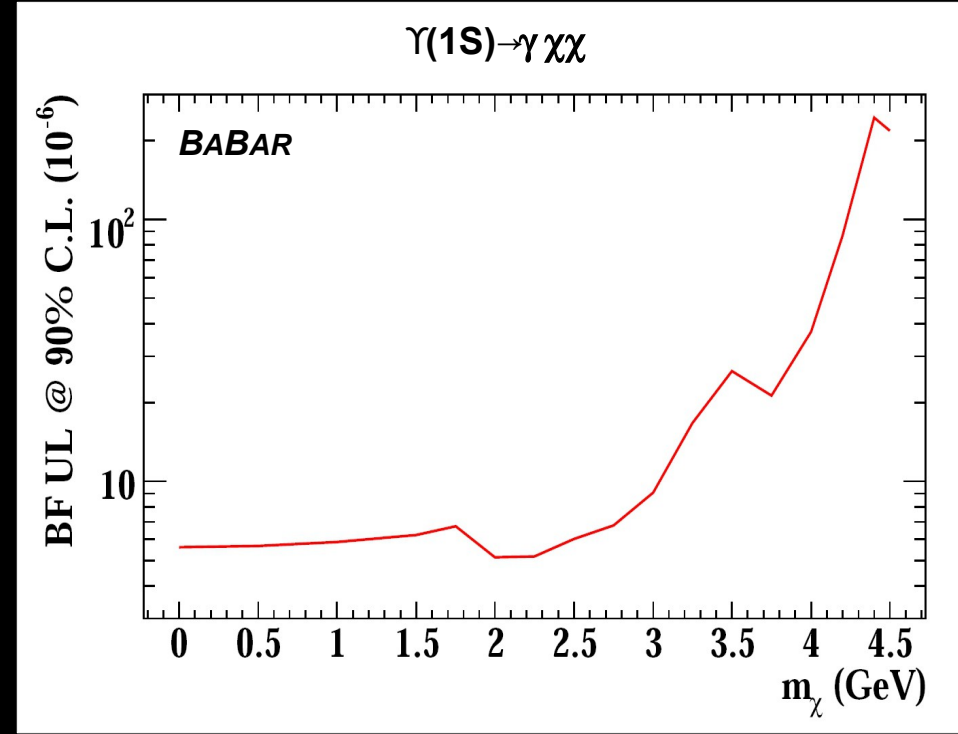


**No significant signal observed**

**Upper limits (90% CL)**

$$\text{BF}(\Upsilon(1S) \rightarrow \gamma A^0, A^0 \rightarrow \text{invisible}) < (1.9 - 37) \times 10^{-6}$$

$$\text{BF}(\Upsilon(1S) \rightarrow \gamma \chi \chi) < (0.5 - 25) \times 10^{-5}$$

**Predictions<sup>1,2)</sup>**

$$\text{BF}(\Upsilon(1S) \rightarrow \gamma A^0, A^0 \rightarrow \text{invisible}) \sim 10^{-4}$$

$$\text{BF}(\Upsilon(1S) \rightarrow \gamma \chi \chi) \sim 10^{-5} - 10^{-4}$$

**Significant constraints on these models**

# Search for dark forces

The image features a dark, almost black background. In the center, there is a cluster of bright, glowing stars in shades of orange, red, and yellow. Surrounding this central cluster is a faint, diffuse nebula with a color gradient from blue on the left to purple and red on the right. The overall appearance is that of a deep space or astronomical scene.

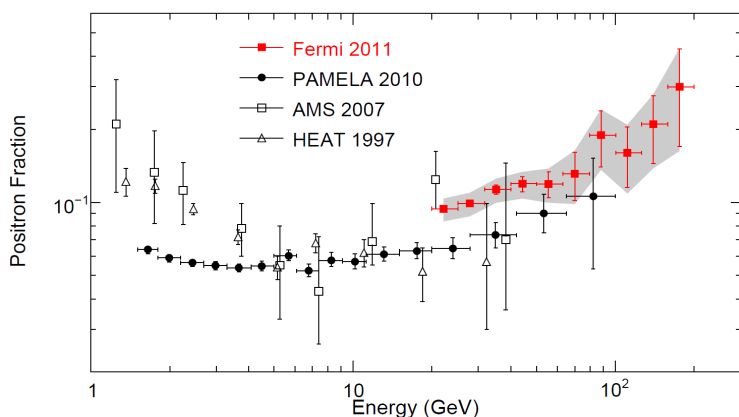
# Dark sector in a nutshell

Models introducing a **new 'dark' force** mediated by a **new gauge boson with a mass around a GeV** have been proposed to explain the observations of PAMELA, FERMI, DAMA/LIBRA, CREST,...

**Wimp-like TeV-scale dark matter** particles can **annihilate into pairs of dark photons**, which subsequently decay to lepton pairs (protons are kinematically suppressed)

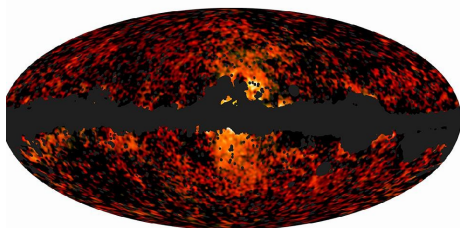
Other explanations of these anomalies have been proposed, but the possibility of a light hidden sector is **poorly constrained** and I recently learned that this idea only.....

## FERMI



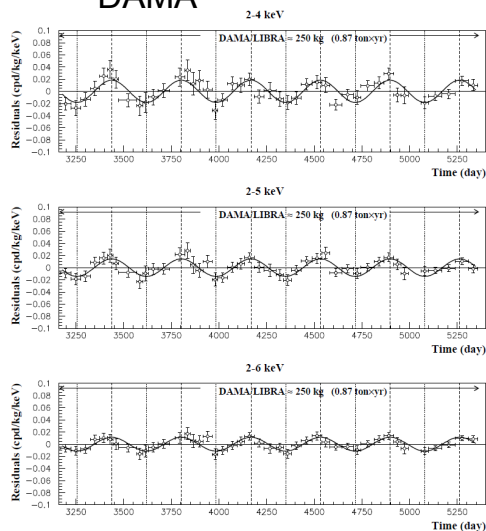
FERMI Collab., PRL 108, 011103 (2012), arXiv:1109.0521

## FERMI HAZE



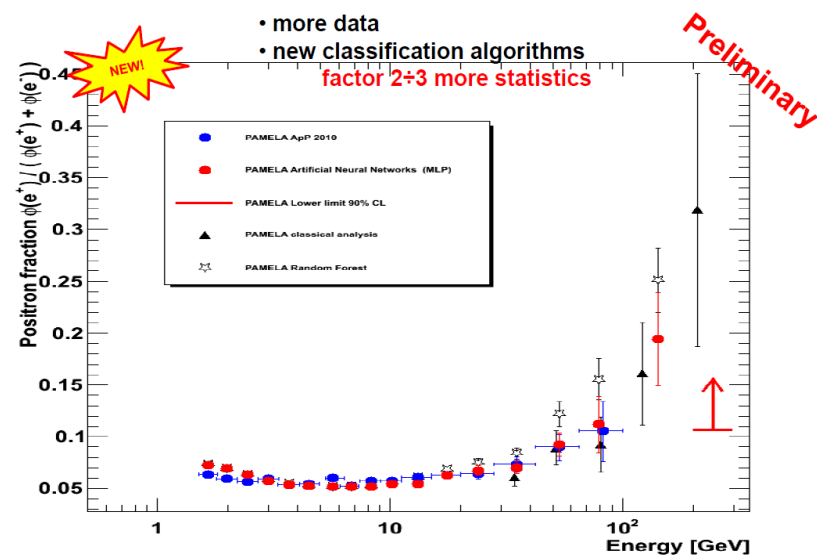
[http://www.nasa.gov/mission\\_pages/planck/multimedia/pia15228.html](http://www.nasa.gov/mission_pages/planck/multimedia/pia15228.html)

## DAMA



DAMA Collab., Eur. Phys. J. C (2010) 67: 39-49

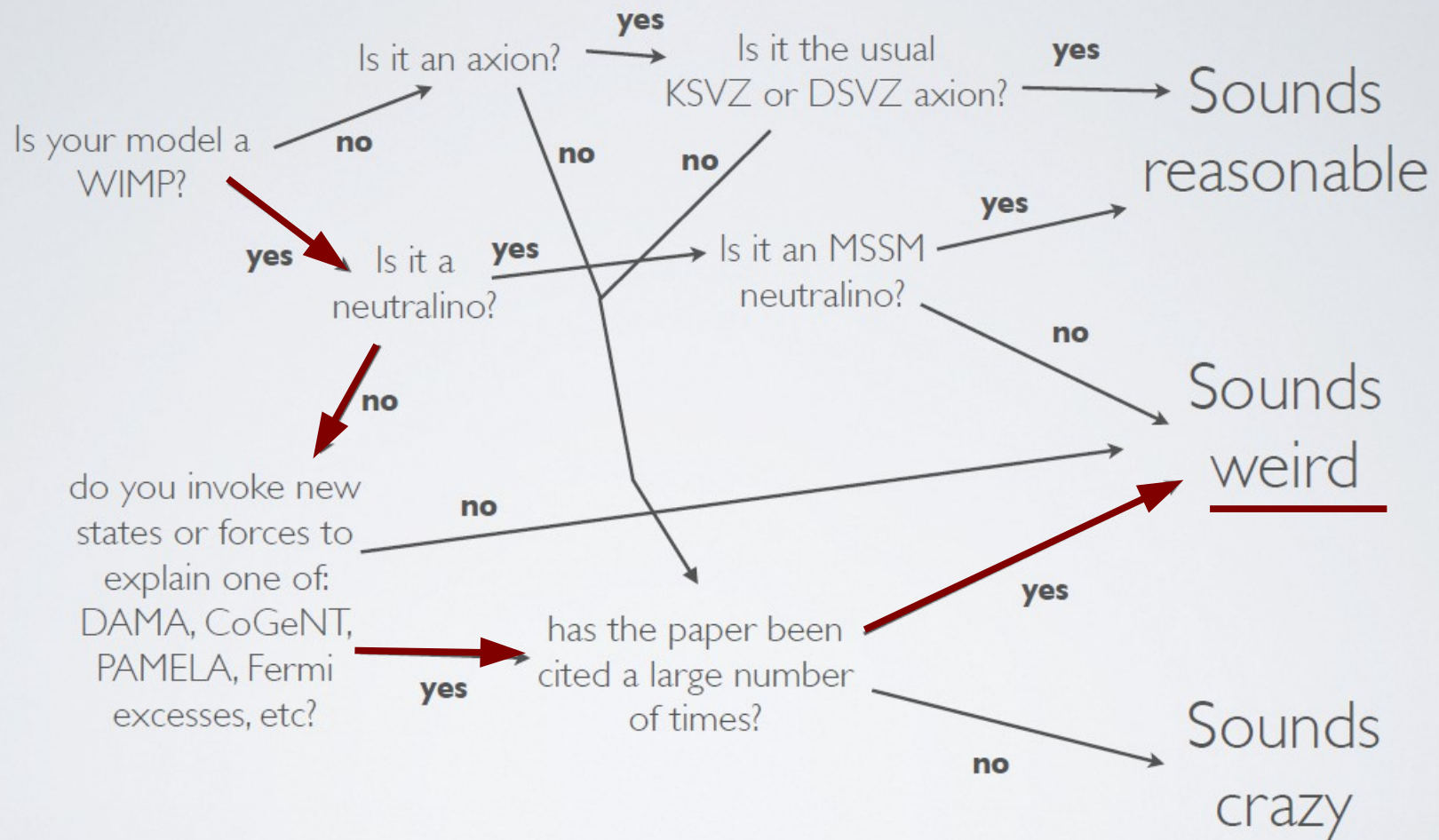
## PAMELA



E. Mocchiutti, International workshop on positrons in Astrophysics 2012

# Neal Weiner classification scheme

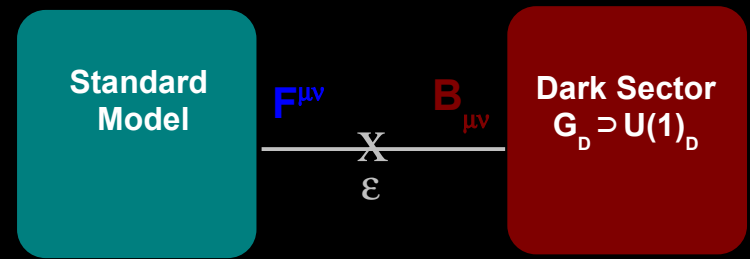
N. Weiner



# Dark sector and dark forces

⇒ New dark sector with a  $U(1)_D$  gauge group

⇒ New gauge boson: **dark photon  $A'$** , which could have  **$O(\text{GeV})$  mass**



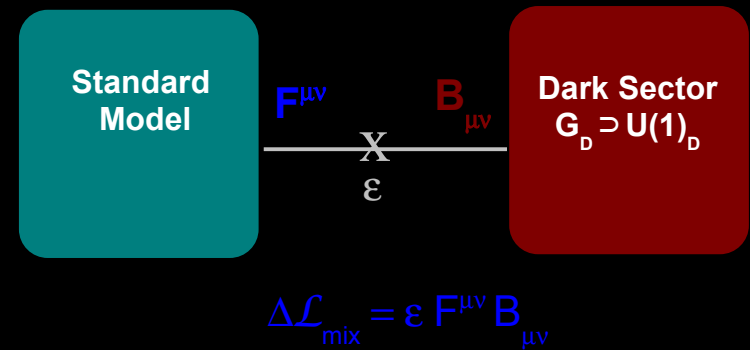
$$\Delta\mathcal{L}_{\text{mix}} = \epsilon F^{\mu\nu} B_{\mu\nu}$$

# Dark sector and dark forces

- ⇒ New dark sector with a  $U(1)_D$  gauge group
- ⇒ New gauge boson: **dark photon  $A'$** , which could have  **$O(\text{GeV})$  mass**
- ⇒ Interaction with the SM is via **kinetic mixing**

$$\varepsilon F^{\mu\nu} B_{\mu\nu}$$

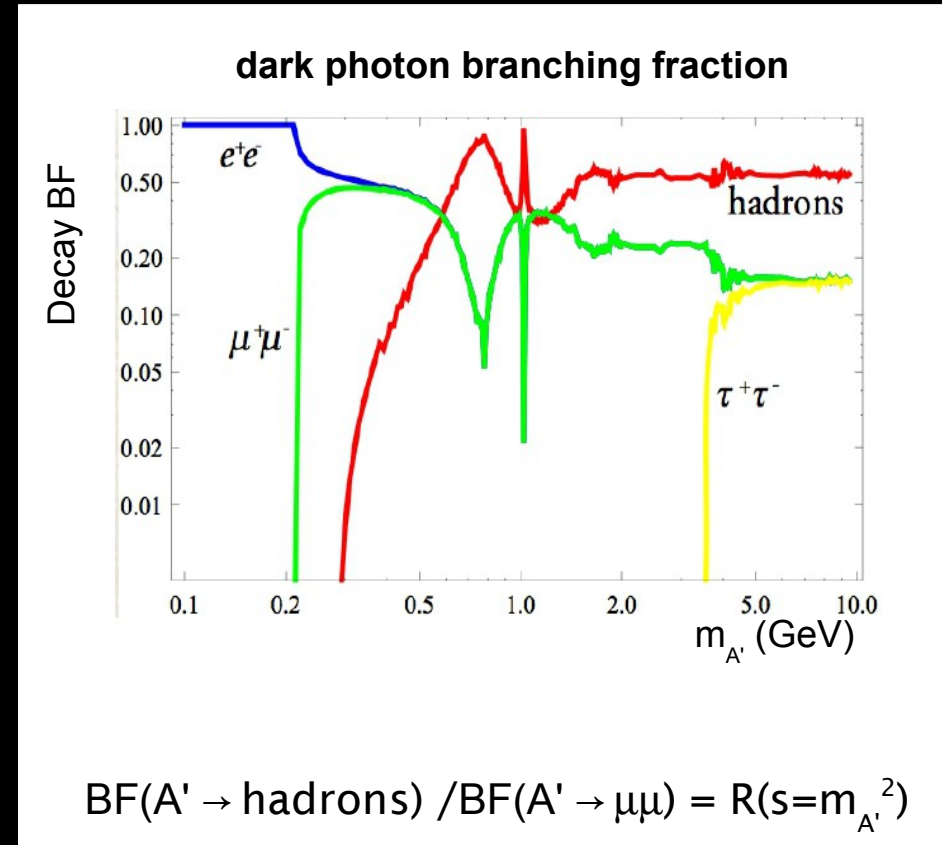
with a **mixing strength  $\varepsilon$** .





# Dark sector and dark forces

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- ⇒ Interaction with the SM is via **kinetic mixing**  
$$\varepsilon F^{\mu\nu} B_{\mu\nu}$$
with a **mixing strength  $\varepsilon$** .
- ⇒ The dark photon acquires a charge  $\varepsilon e$ , and the **coupling of the dark photon to SM fermions** is characterized by  **$\alpha' = \alpha \varepsilon^2$**



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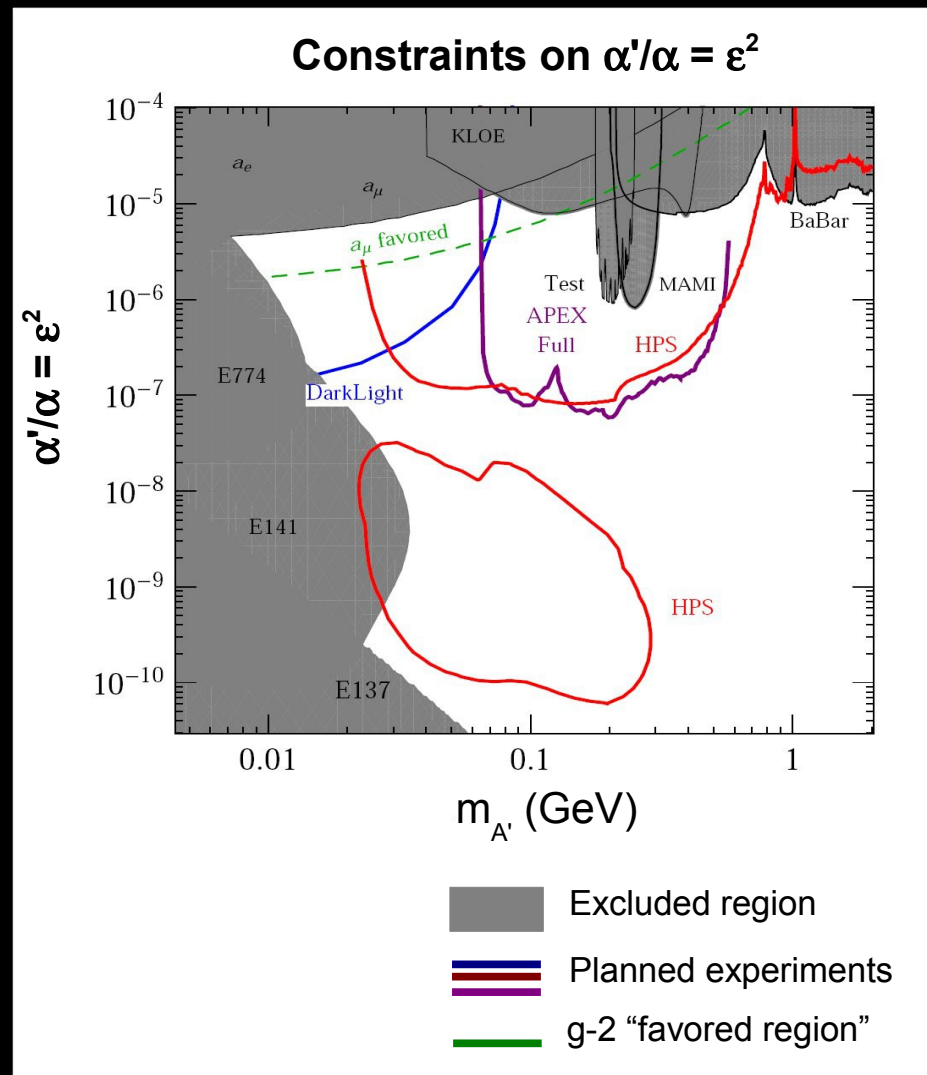
- ⇒ Interaction with the SM is via **kinetic mixing**

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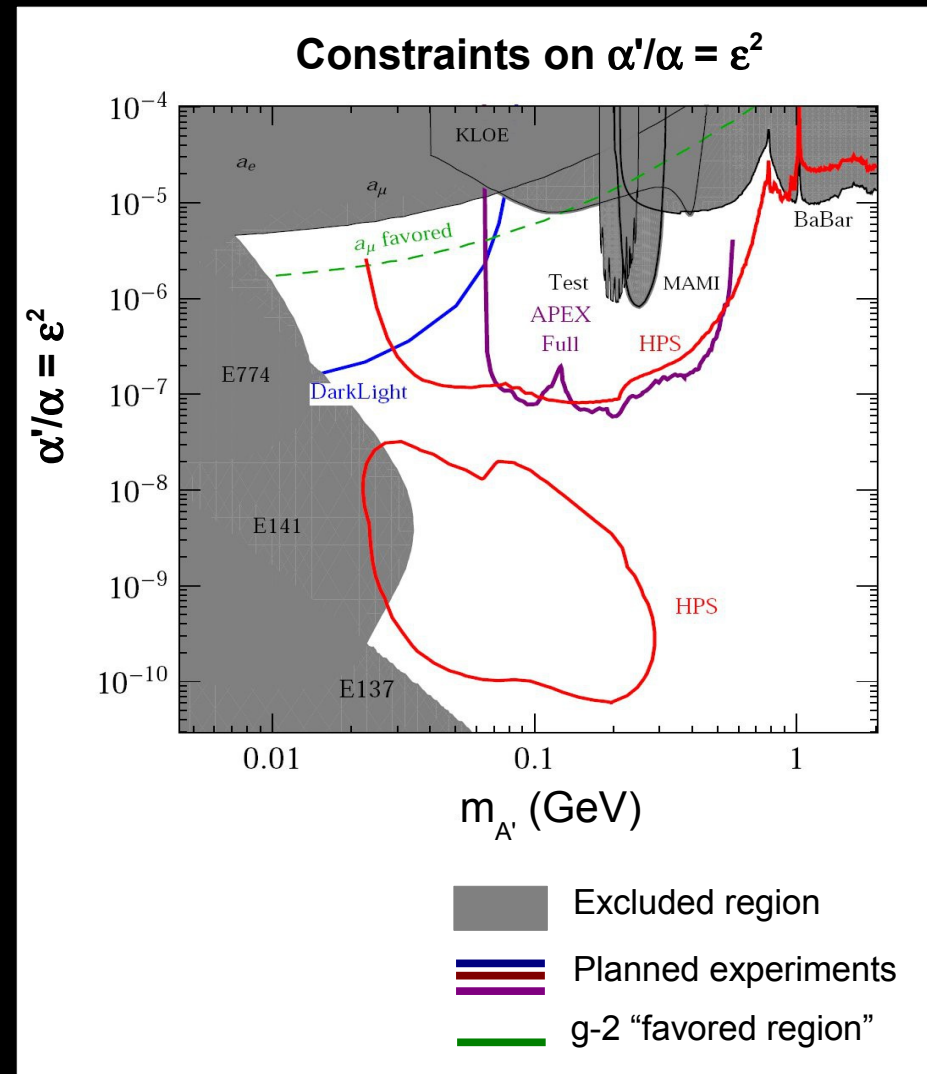
- ⇒ Current limits on the mixing strength  $\varepsilon^2$  are shown as a function of the dark photon mass for existing measurements and a few planned experiments.



Low-energy high-luminosity  $e^+e^-$  colliders offer a **low-background environment** to search for **MeV/GeV-scale hidden sector signatures** and **probe their structure**

- ⇒ New dark sector with a  $U(1)_D$  gauge group
- ⇒ New gauge boson: **dark photon  $A'$** , which could have  **$O(\text{GeV})$  mass**
- ⇒ Interaction with the SM is via **kinetic mixing**

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 with a **mixing strength  $\varepsilon$** .
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Other portals can connect the dark sector to the SM (Higgs portal, axion portal,...)  
I'll only discuss the vector portal in this talk

## Search for dark photon

$$e^+e^- \rightarrow \gamma A', \quad A' \rightarrow e^+e^-, \mu^+\mu^-, \pi^+\pi^-$$

## Search for dark photon in meson decay

$$\pi^0 \rightarrow \gamma l^+l^-, \eta \rightarrow \gamma l^+l^-, \phi \rightarrow \eta l^+l^-, \dots$$

## Search for dark boson(s)

$$e^+e^- \rightarrow A'^* \rightarrow W'W',$$
$$e^+e^- \rightarrow \gamma A' \rightarrow W'W''$$

## Search for dark scalar (s) / pseudoscalar (a)

$$B \rightarrow K^{(*)}s \rightarrow K^{(*)}l^+l^- \text{ and } B \rightarrow K^{(*)}a \rightarrow K^{(*)}l^+l^-$$
$$B \rightarrow ss \rightarrow 2(l^+l^-)$$
$$B \rightarrow K 2(l^+l^-)$$
$$B \rightarrow 4(l^+l^-)$$

## Search for dark Higgs boson

$$e^+e^- \rightarrow h' A', \quad h' \rightarrow A' A'$$

+ related searches (hidden warped extra dimensions,...)

## Search for dark hadrons

$$e^+e^- \rightarrow \pi_D + X, \quad \pi_D \rightarrow e^+e^-, \mu^+\mu^-$$

Can probe the hidden sector structure

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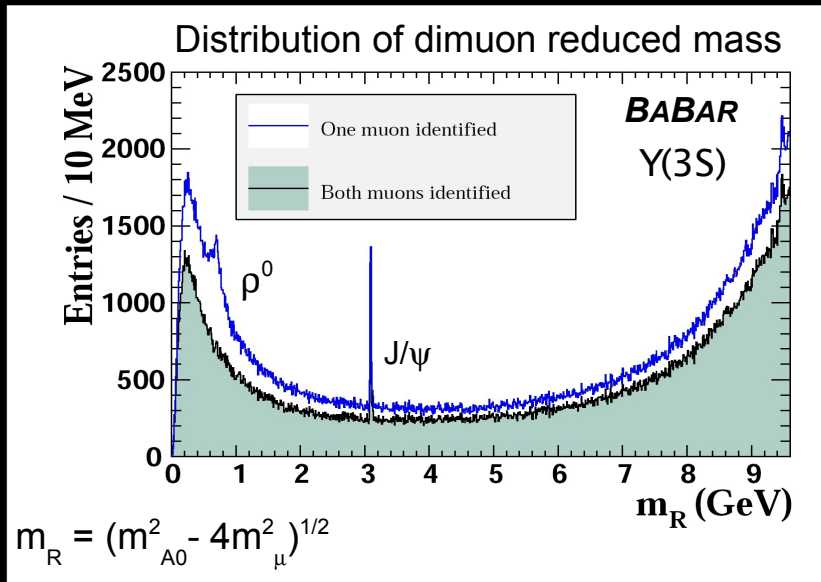
Can probe the hidden sector structure

A dark photon can be readily produced in

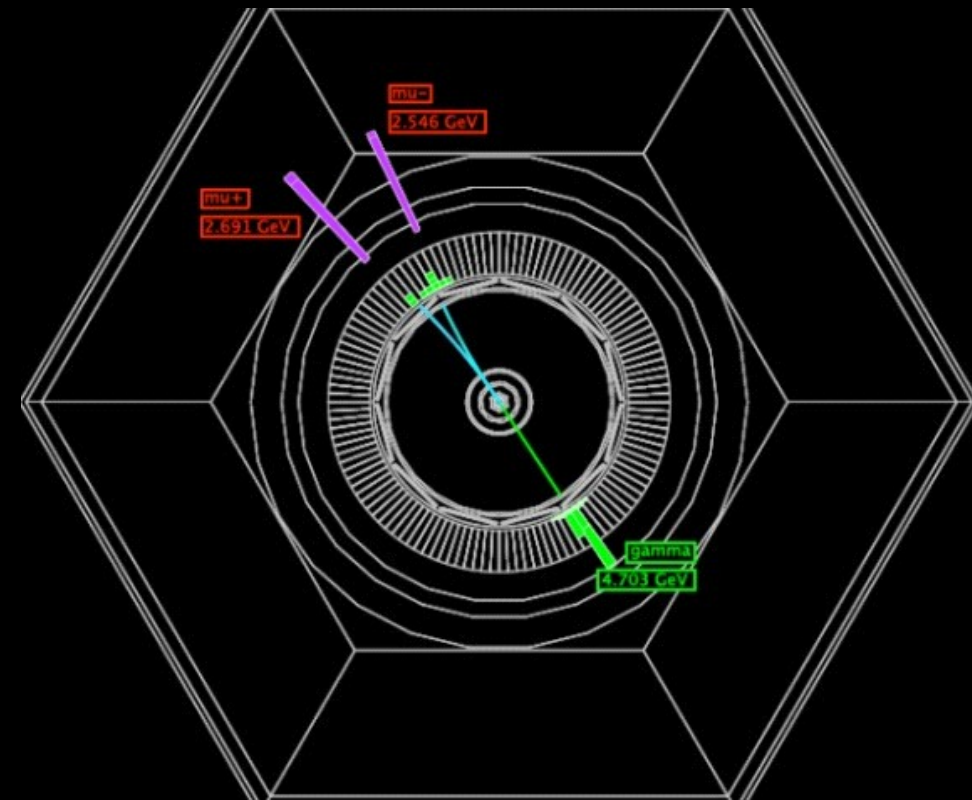
$$e^+e^- \rightarrow \gamma A' \rightarrow \gamma l^+l^-, \gamma q\bar{q}$$

So far, only one measurement of this final state at *BABAR* from light CP-odd Higgs search in  $\Upsilon(2S,3S)$  decays based on  $\sim 45 \text{ fb}^{-1}$  of data:

$$e^+e^- \rightarrow \gamma A^0, A^0 \rightarrow \mu^+\mu^-$$



## Candidate event



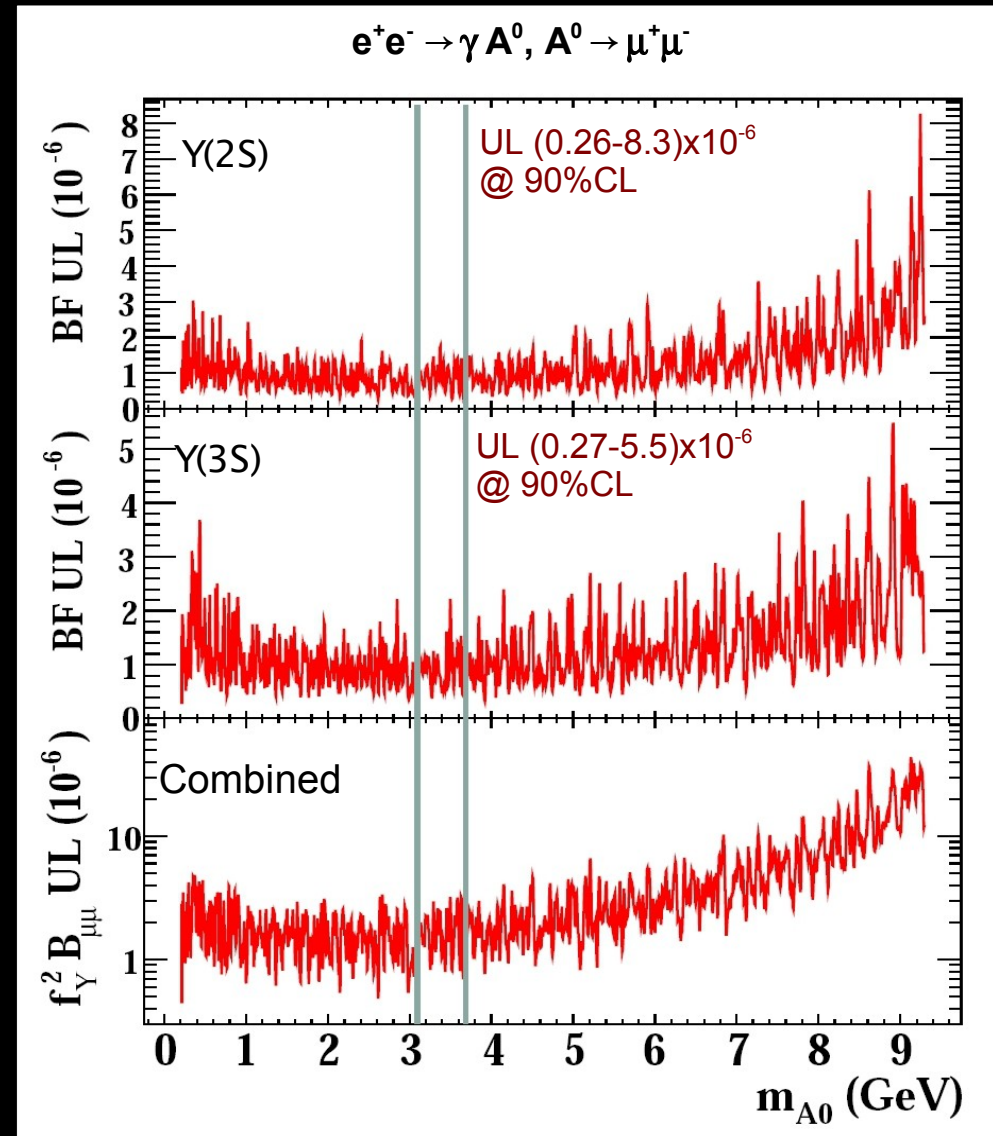
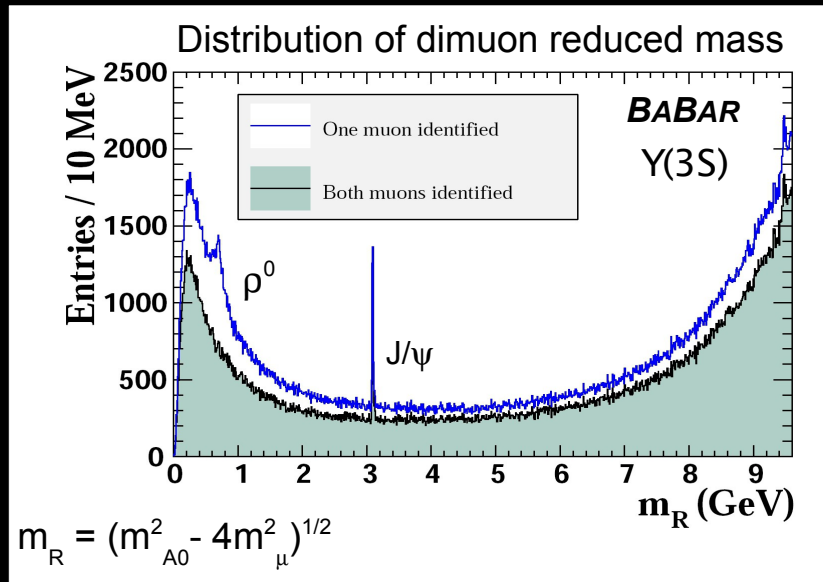
- Tracks
- Photon
- Signal in muon/hadron detector

A dark photon can be readily produced in

$$e^+e^- \rightarrow \gamma A' \rightarrow \gamma l^+l^-, \gamma q\bar{q}$$

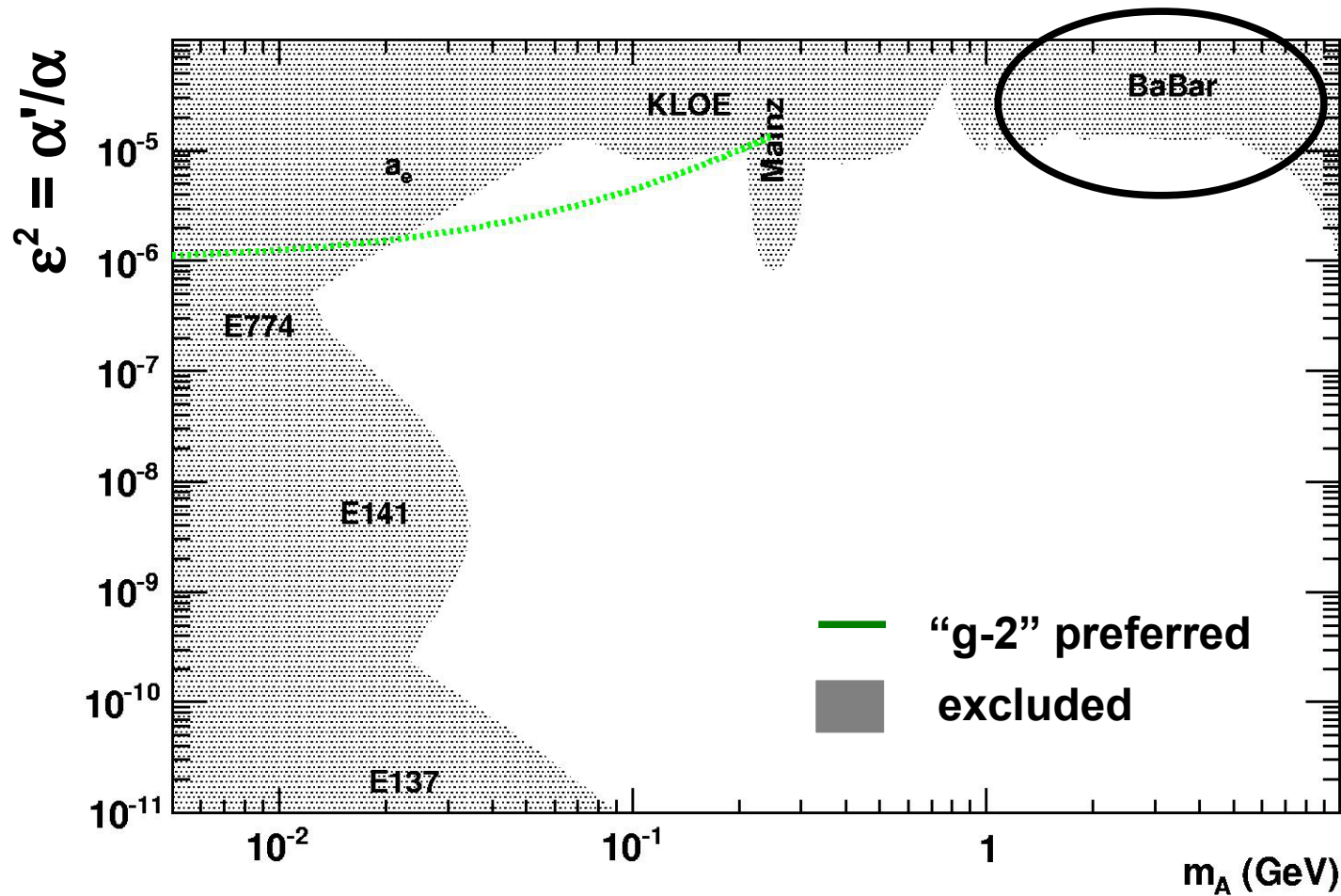
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**No sign of narrow resonance**

Limit obtained by reinterpreting the  $Y(2S,3S) \rightarrow \gamma A^0, A^0 \rightarrow \mu^+\mu^-$  measurements<sup>1)</sup>

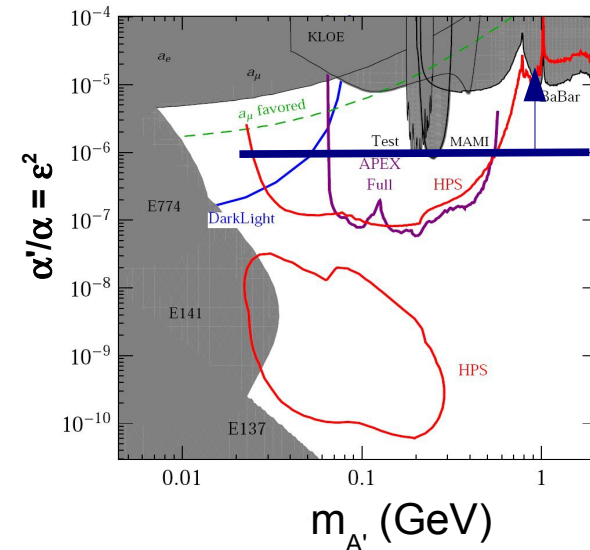
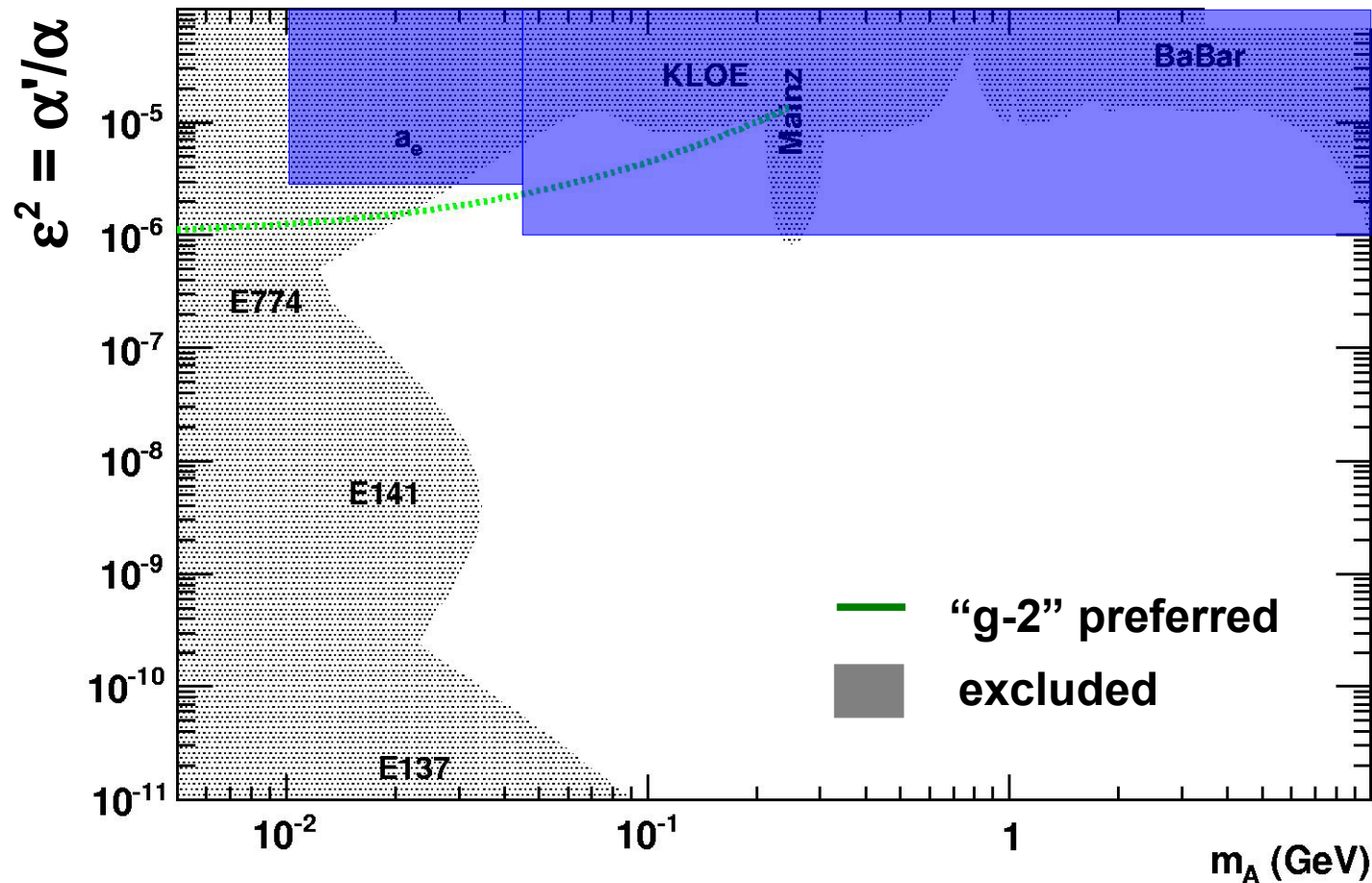


1) J.D. Bjorken et al., PRD 80 (2009) 075018

Note:  $\sigma(e^+e^- \rightarrow \gamma A') \sim 1/s$ , so need low-energy collider to probe direct production

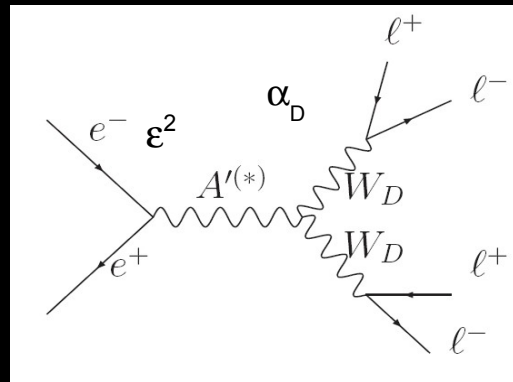


Crude estimate of expected limits using the full *BABAR* dataset for  $e^+e^- \rightarrow \gamma e^+e^-, \gamma \mu^+\mu^-, \gamma \pi^+\pi^-$



Expect significant improvement, exclude almost all the “g-2” preferred region. Stay tuned!

$$e^+e^- \rightarrow A'^* \rightarrow WW', W^{(\prime)} \rightarrow e^+e^-, \mu^+\mu^-$$



$$\alpha_D = g_D^2 / 4\pi$$

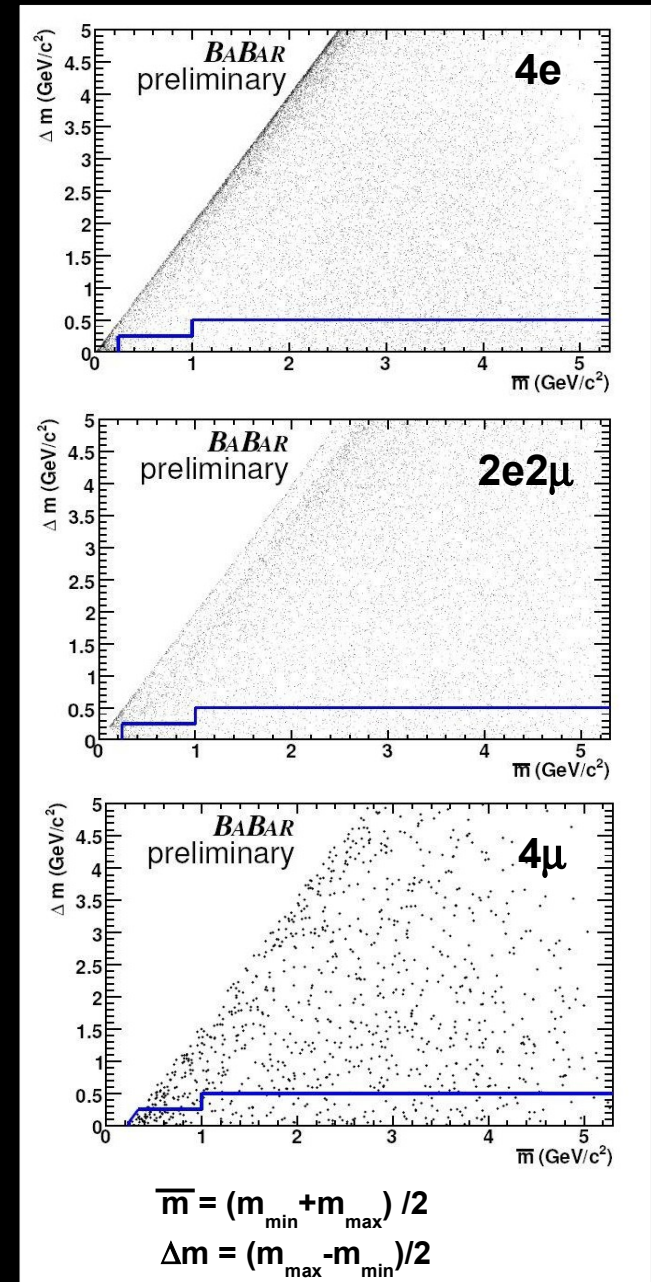
$g_D$  dark sector gauge coupling

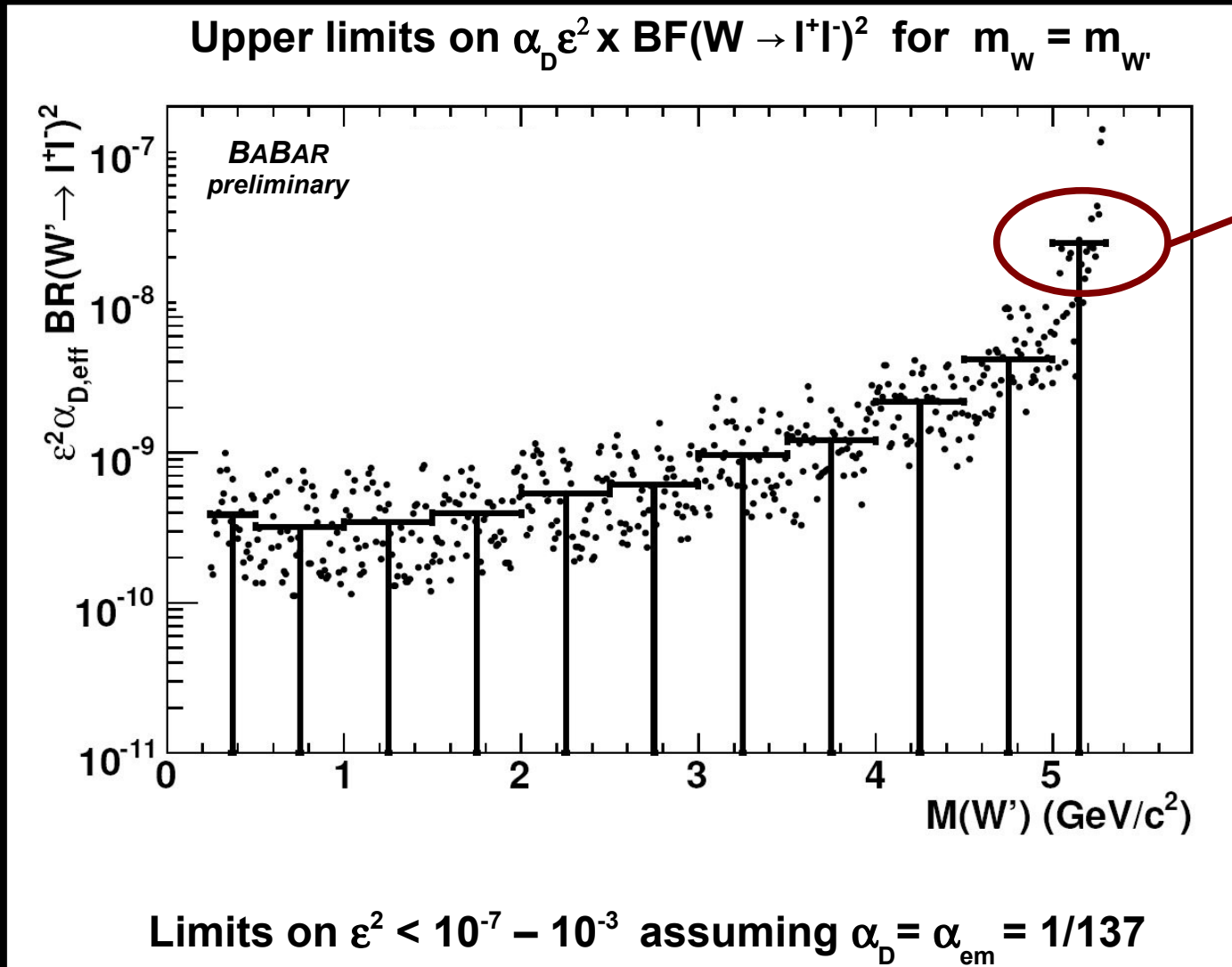
⇒ The simplest extension to a non-Abelian case is SU(2)xU(1), which has **4 bosons: A', W, W' and W''**

⇒ Sensitive to the dark sector gauge coupling  $g_D$

⇒ Can produce a pair of dark bosons through an off-shell A'.  
Process suppressed only by  $\alpha_D \epsilon^2$  where  $\alpha_D = g_D^2 / 4\pi$

⇒ Search for **two dileptonic resonances with similar mass**





Average limit over many bins

Expect similar limits for  $m_W - m_{W'} \gg 0$

# Search for dark Higgs boson

B. Batell et al., PRD 79 (2009) 115008  
 R. Essig et al., PRD 80 (2009) 015003

⇒ Dark photon mass is generated via the Higgs mechanism, **adding a dark Higgs boson (h')**

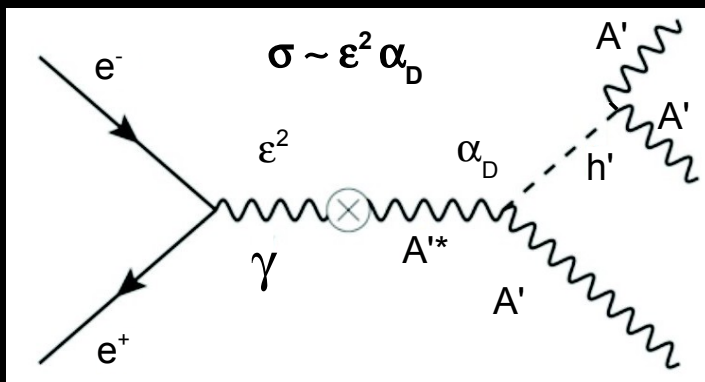
⇒ A minimal scenario has a **single dark photon** and **a single dark Higgs boson**.

⇒ **Dark Higgs mass** could be at **the GeV scale**

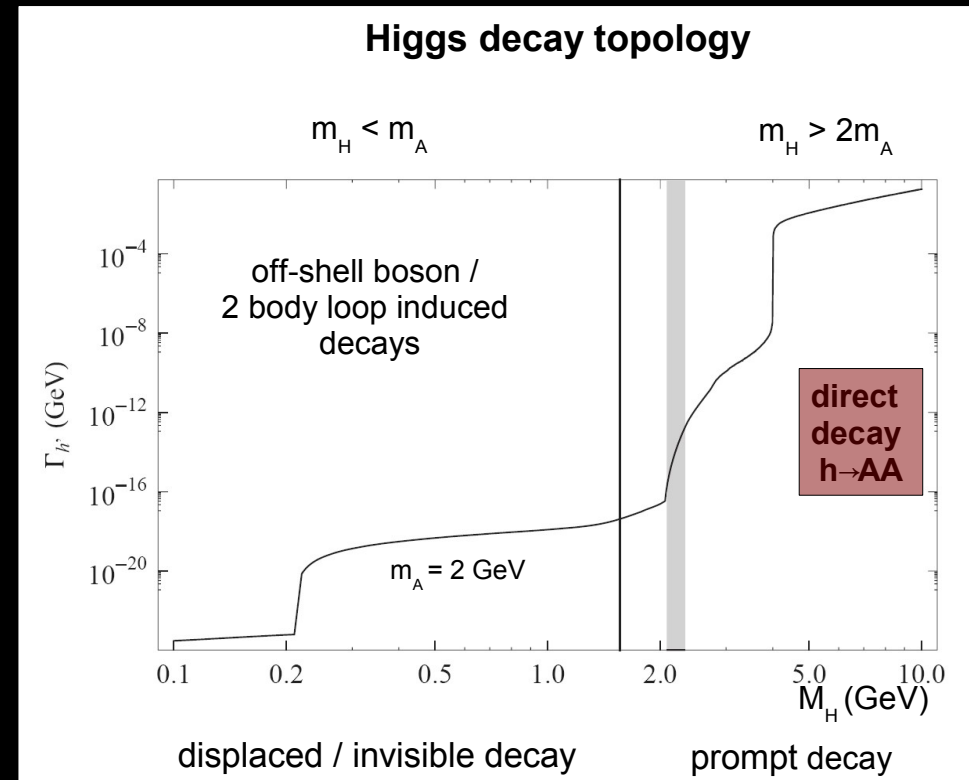
⇒ The **Higgs-strahlung process**

$$e^+e^- \rightarrow A'^* \rightarrow h' A', h' \rightarrow A' A'$$

is very interesting, as it is **only suppressed by  $\epsilon^2$**  and is expected to have a **very small background**.



$\alpha_D = g_D^2 / 4\pi$   
 $g_D$  is the dark sector gauge coupling

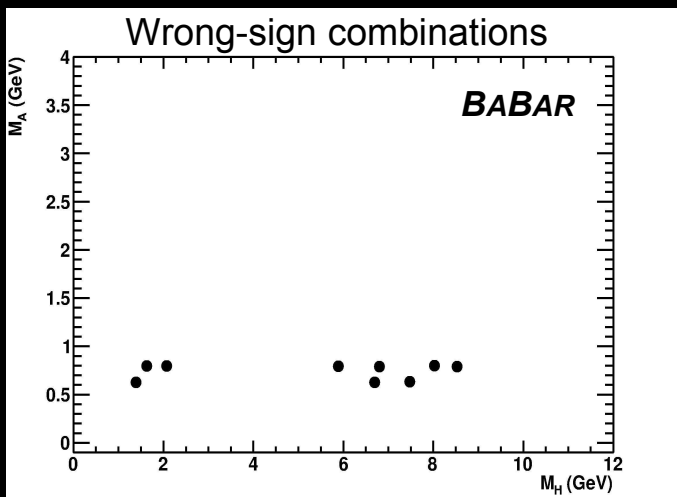
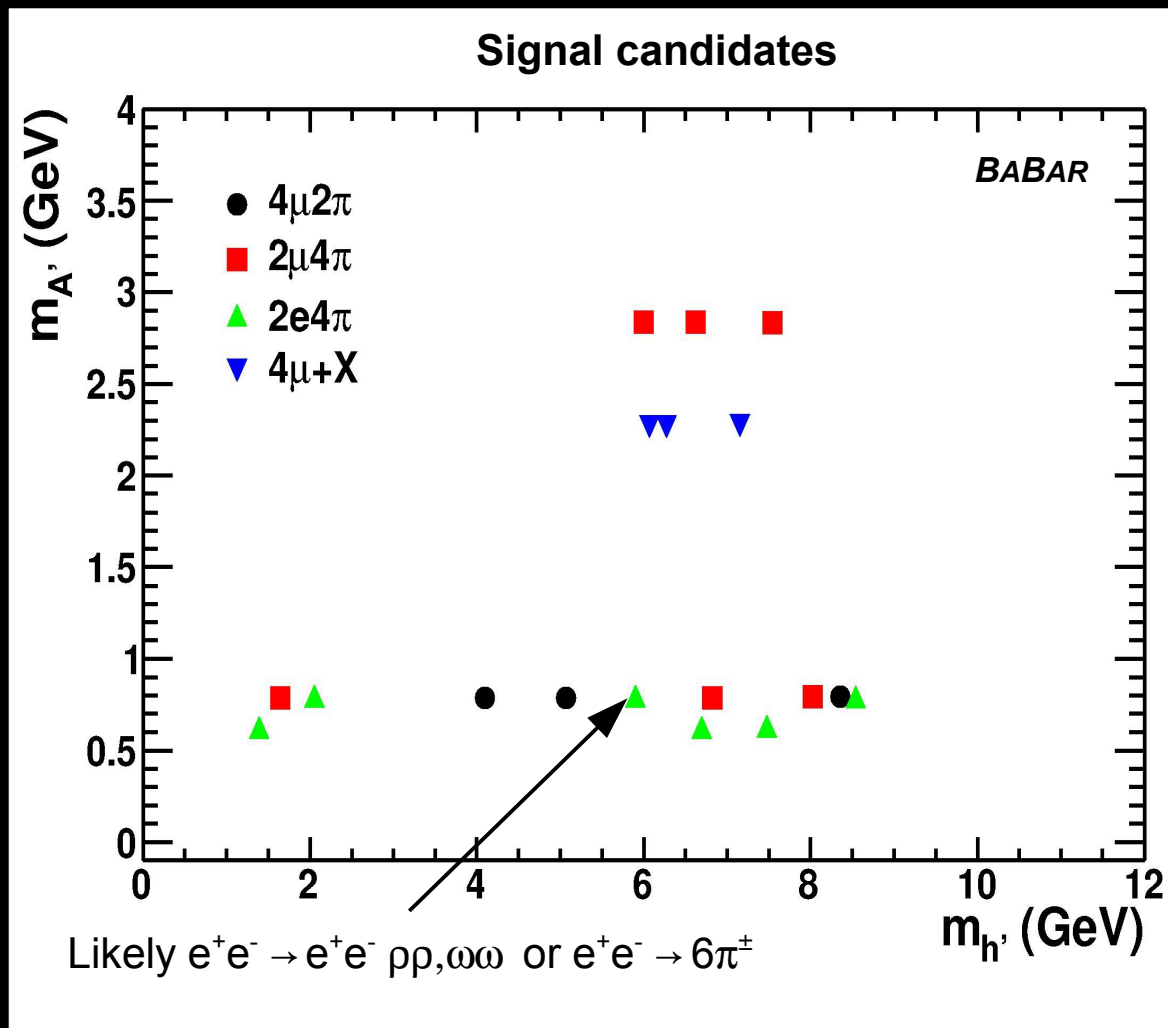


**Focus on prompt decays**

$$m_h > 2m_A$$

**Search for three resonances with similar masses**

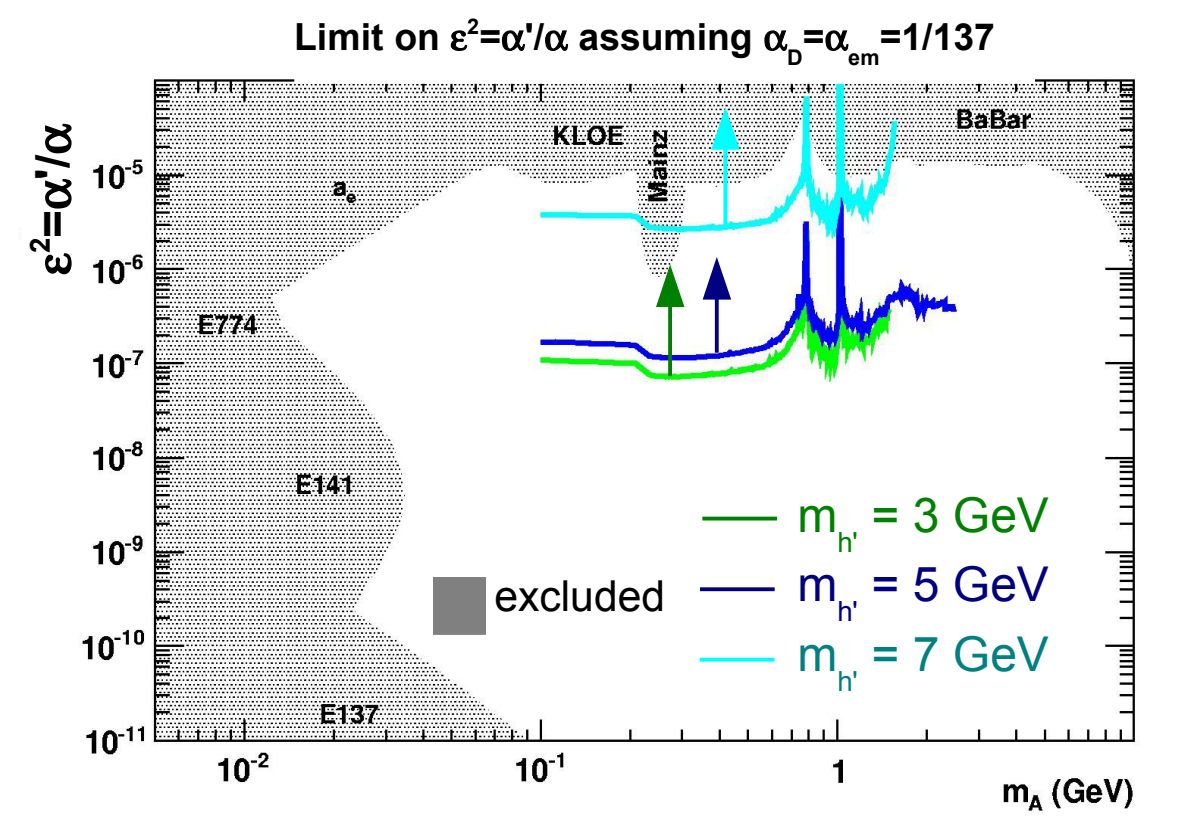
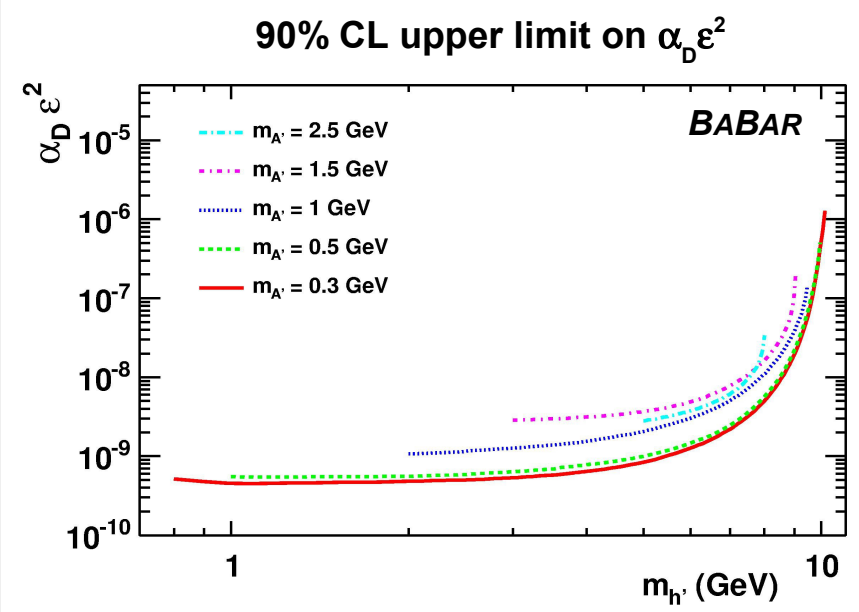
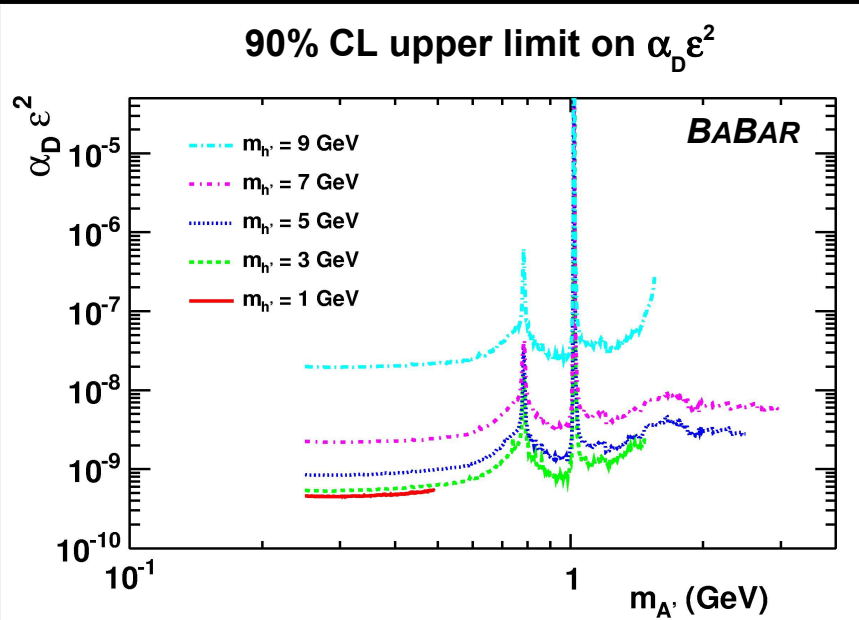
- ⇒ Six candidates are selected from the full *BABAR* dataset ( $\sim 500 \text{ fb}^{-1}$ )
- ⇒ Three entries for each event, corresponding to the three possible assignments of the  $h' \rightarrow A'A'$  decay
- ⇒ Estimate background from
  - wrong-sign combinations, e.g.  $e^+e^- \rightarrow (e^+e^+)(e^-e^-)(\mu^+\mu^-)$
  - sidebands from final sample
  - rate for 6 leptons  $\sim 100\times$  rate for  $4\pi+2l$  above 1.5 GeV



**No events with 6 leptons, consistent with the pure background hypothesis**

# Search for dark Higgs boson

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**Substantial improvement over existing limits for  $m_{h'} < 5 - 7$  GeV if light dark Higgs boson exists**

Low-energy  $e^+e^-$  colliders provide a clean environment to explore MeV-GeV scale dark matter and dark sector. Colliders can not only investigate their existence, but also their structure.

Several results in  $\Upsilon$  decays have already set stringent limits on theoretical models of generic dark matter or BSM physics.

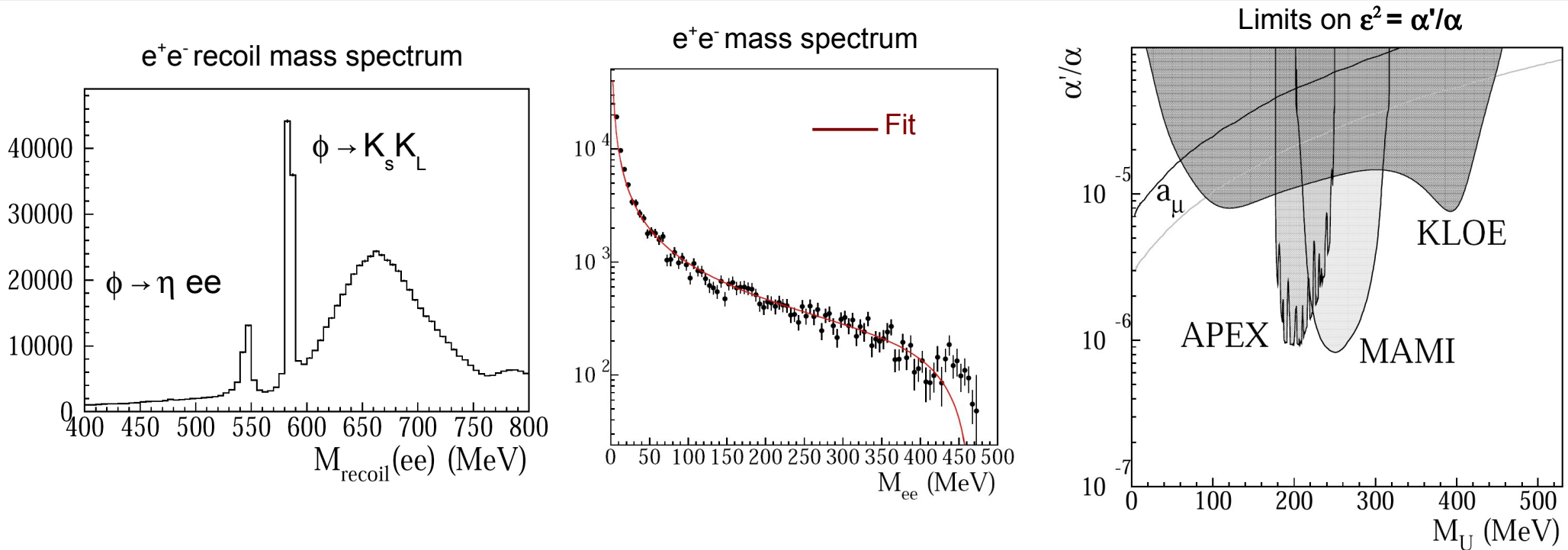
Searches for dark sector particles have already set the best limits on dark sector particles, and many searches are currently being performed at *BABAR*.

Extra material

The image features a dark, almost black background. In the center, there is a cluster of approximately 15-20 bright, orange-red stars of varying sizes. Surrounding this central cluster is a diffuse, glowing nebula-like structure with a color gradient from blue on the left to purple and red on the right. The overall appearance is that of a star-forming region or a young stellar cluster.



Search for dark photon in  $e^+e^- \rightarrow \phi \rightarrow \eta A', \eta \rightarrow \pi^+\pi^-\pi^0, A' \rightarrow e^+e^-$  at KLOE with  $\sim 7 \times 10^9$   $\phi$  mesons



Exclude values of  $\epsilon^2$  down to  $\sim 10^{-5}$  for  $80 < m_{A'} < 400$  MeV  
 KLOE II is expected to improve these limits by a factor 3-5.

## Search for invisible dark photon decay

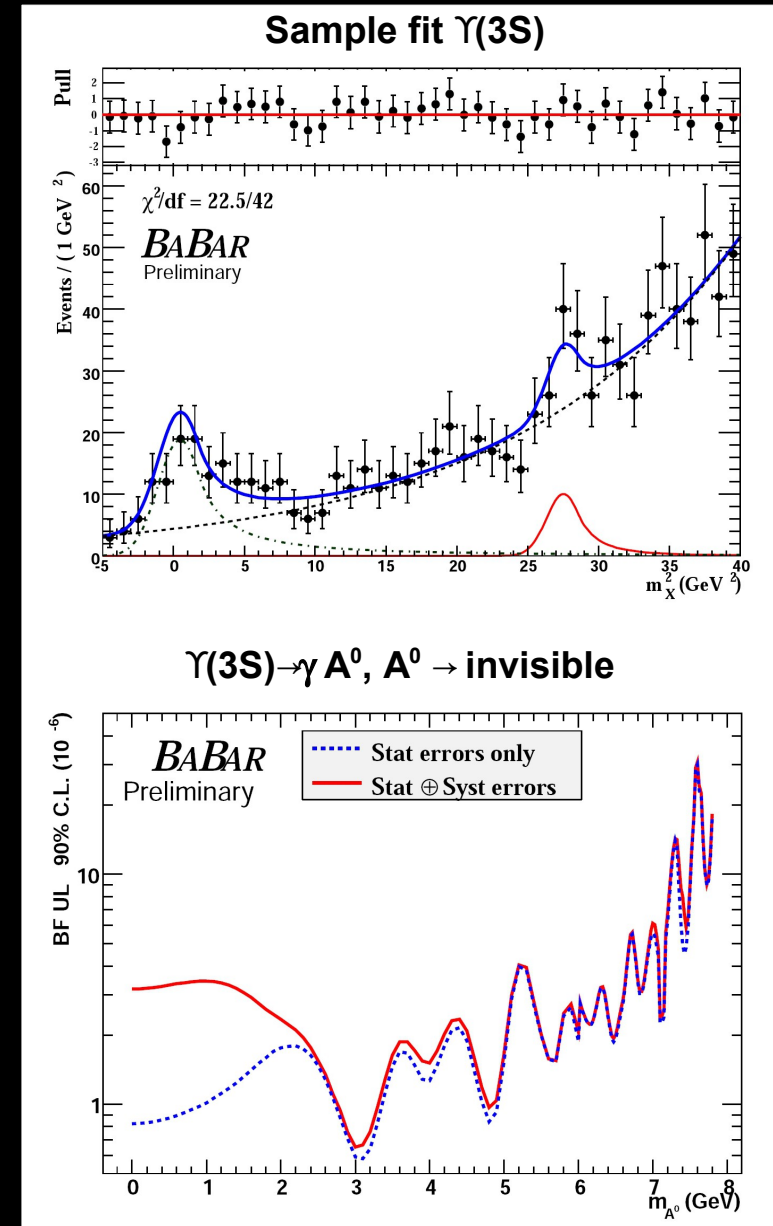
$$e^+e^- \rightarrow \gamma A', A' \rightarrow \text{invisible}$$

⇒ Several scenarios with **dark photon decay to invisible particles** (light dark matter for example, see *Cline et al., arXiv 1108.1391*).

⇒ Dark photon or similar particle may be long lived and escape detector.

⇒ Reinterpret  $\Upsilon(2S,3S) \rightarrow \gamma + \text{invisible}$  measurement from *BABAR*, limits on  $\epsilon^2$  should be at the level of  $10^{-5} - 10^{-4}$ .

⇒ Current analysis will improve the preliminary results by a factor  $\sim 2$ .



## Fully reconstructed

$$e^+e^- \rightarrow h' A', h' \rightarrow A' A'$$

with  $A' \rightarrow e^+e^-, \mu^+\mu^-, \pi^+\pi^-$

### Fully reconstructed signal

⇒ Three dark photons fully reconstructed

### Modes included

⇒  $e^+e^- \rightarrow (l^+l^-) (l^+l^-) (l^+l^-)$   $l=e,\mu$

⇒  $e^+e^- \rightarrow (l^+l^-) (l^+l^-) (\pi^+\pi^-)$

⇒  $e^+e^- \rightarrow (l^+l^-) (\pi^+\pi^-) (\pi^+\pi^-)$

### Selection

⇒ 6 tracks with an invariant mass  $m_{\text{tot}} > 0.95 \sqrt{s}$

⇒ apply particle identification

⇒ cosine helicity angle of  $A' \rightarrow e^+e^-$  candidates  $< 0.9$

⇒ three dark photon candidates have similar mass

## Partially reconstructed

$$e^+e^- \rightarrow h' A'_1, h' \rightarrow A'_2 A'_3$$

$A'_{1,2} \rightarrow e^+e^-, \mu^+\mu^-, A'_3 \rightarrow X + \text{perm.}$

### Partially reconstructed signal

⇒ In the high mass region ( $m_A > 1.2 \text{ GeV}$ ), the decay of the dark photon is dominated by  $A' \rightarrow q\bar{q}$

⇒ Measure 2  $A'$  decaying to leptons and 1  $A' \rightarrow q\bar{q}$

⇒ Assign recoiling system to  $A'_3$ ,  $P_3 = P_{ee} - P_1 - P_2$

### Modes included

⇒  $e^+e^- \rightarrow (l^+l^-) (\mu^+\mu^-) + X$  where  $X$  is not  $l^+l^- / \pi^+\pi^-$

### Selection

⇒ 4 or more tracks

⇒ apply particle identification for  $A' \rightarrow l^+l^-$  decays

⇒ cosine helicity angle of  $A' \rightarrow e^+e^-$  candidates  $< 0.9$

⇒ three dark photon candidates have similar mass