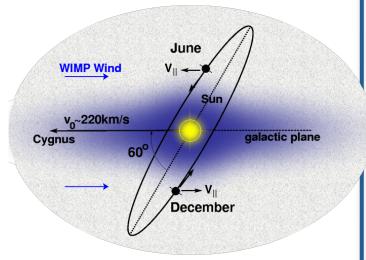
On the annual modulations in direct-detection experiments and the DAMA results Itay Yavin McMaster University Perimeter Institute

This research was supported in part by 🐨

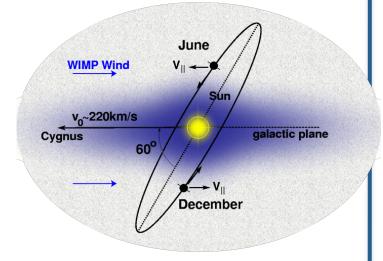
Aspen - Closing in on Dark Matter

29 January, 2013

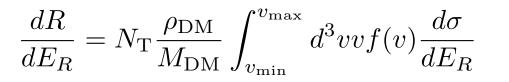


We measure the differential recoil rate,

$$\frac{dR}{dE_R} = N_{\rm T} \frac{\rho_{\rm DM}}{M_{\rm DM}} \int_{v_{\rm min}}^{v_{\rm max}} d^3 v v f(v) \frac{d\sigma}{dE_R}$$

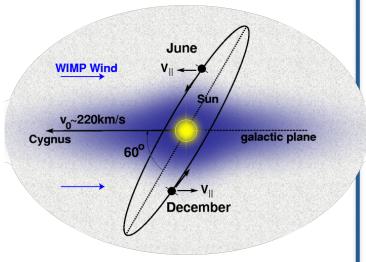


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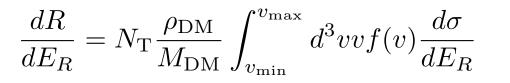


But our velocity w.r.t to halo is modulating

$$\mathbf{v}_{\rm obs} = \mathbf{v}_{\odot} + V_{\oplus} \left(\varepsilon_1 \cos \omega \left(t - t_1 \right) + \varepsilon_2 \sin \omega \left(t - t_1 \right) \right)$$



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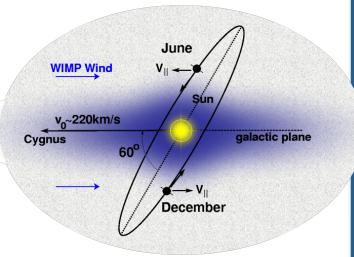
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This leads to the well known annual modulations (Drukier, Freese, Spergel, 1986) and more (Chang, Pradler, IY, 2011):

$$egin{aligned} rac{dR}{dE_R} \propto & \int_{v_{min}}^\infty rac{f(v)}{v} dv \ &pprox & \sum_{n=0,1,...} ilde{c}_n(v_{min}) \left[\epsilon_v \cos \omega (t-t_0)
ight]^n \ &= & \sum_{n=0,1,...} ilde{c}_n(v_{min}) \cos \left[n \omega (t-t_0)
ight] \end{aligned}$$

Thursday, January 31, 13



expanding in powers of

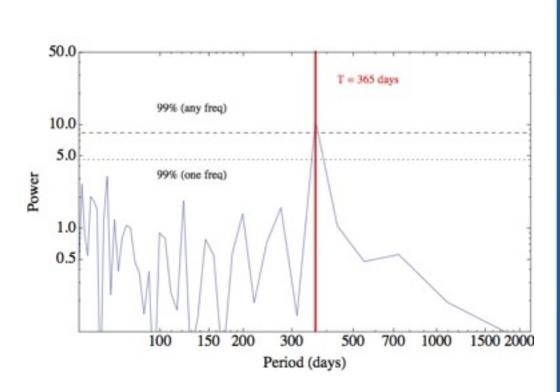
 $\epsilon_v = V_\oplus/2v_\odot \approx 0.06$

Use the Lomb-Scargle periodogram to test for dominant frequencies,

$$\begin{split} \mathrm{LS}(\omega) &= \frac{1}{\sum_{i} \cos^{2}\left(\omega \tilde{t}_{i}\right)} \left[\sum_{i} d_{i} \cos\left(\omega \tilde{t}_{i}\right)\right]^{2} \\ &+ \frac{1}{\sum_{i} \sin^{2}\left(\omega \tilde{t}_{i}\right)} \left[\sum_{i} d_{i} \sin\left(\omega \tilde{t}_{i}\right)\right]^{2} \end{split}$$

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DAMA/LIBRA

DAMA/LIBRA Use the Lomb-Scargle periodogram to test for dominant frequencies, 50.0 T = 365 days 99% (any freq) $\mathrm{LS}(\omega) = \frac{1}{\sum_{i} \cos^{2}\left(\omega \tilde{t}_{i}\right)} \left[\sum_{i} d_{i} \cos\left(\omega \tilde{t}_{i}\right)\right]^{2}$ 10.0 5.0 Power 99% (one freq) $+ \frac{1}{\sum_{i} \sin^2(\omega \tilde{t}_i)} \left[\sum_{i} d_i \sin(\omega \tilde{t}_i) \right]^2$ 1.0 0.5 100 150 200 300 500 700 1000 1500 2000 Period (days) There also seems to be power in 1/3 year (saw-tooth signal?)

Testing the Null Hypothesis CoGeNT 0.8 combined efficiency 0.2 3.2 0 0.8 1.21.6 $\mathbf{2}$ 2.42.8 0.4energy (keVee)

160

140

120

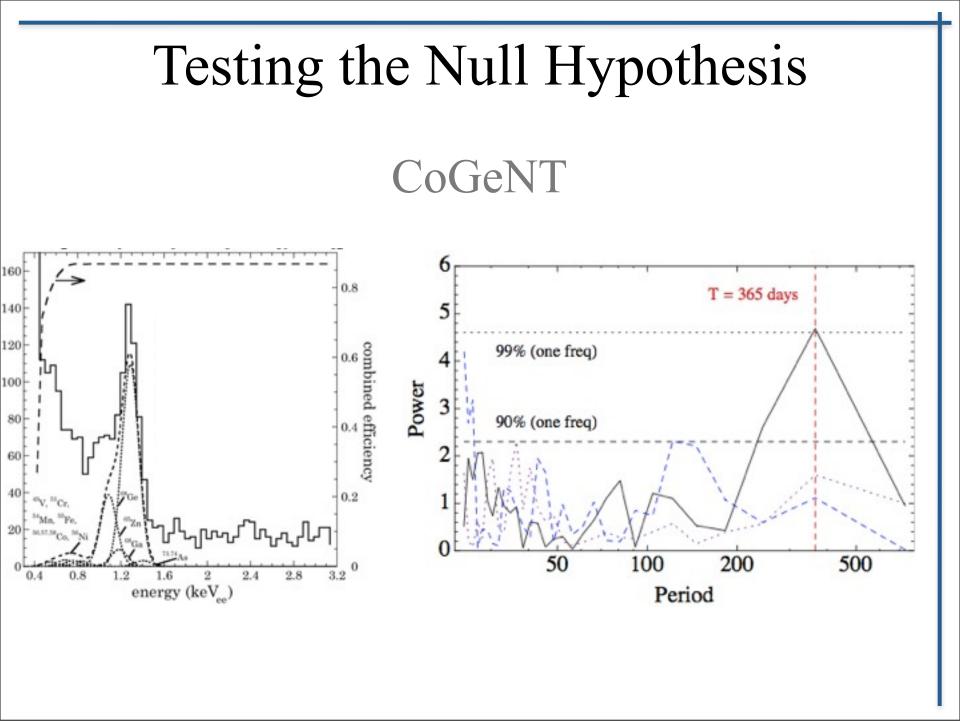
100

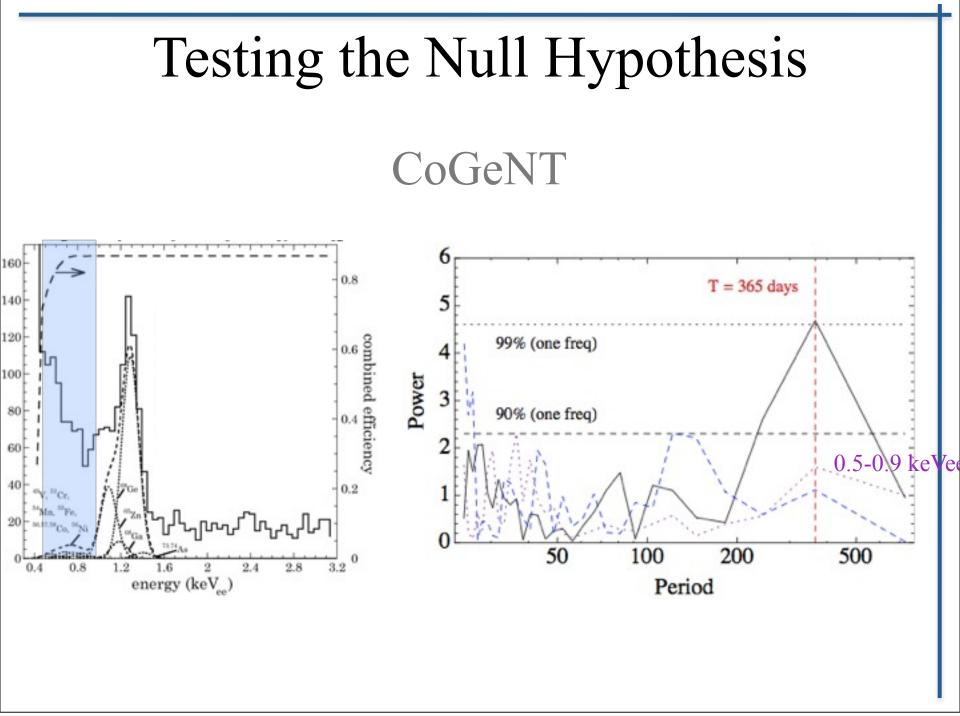
80

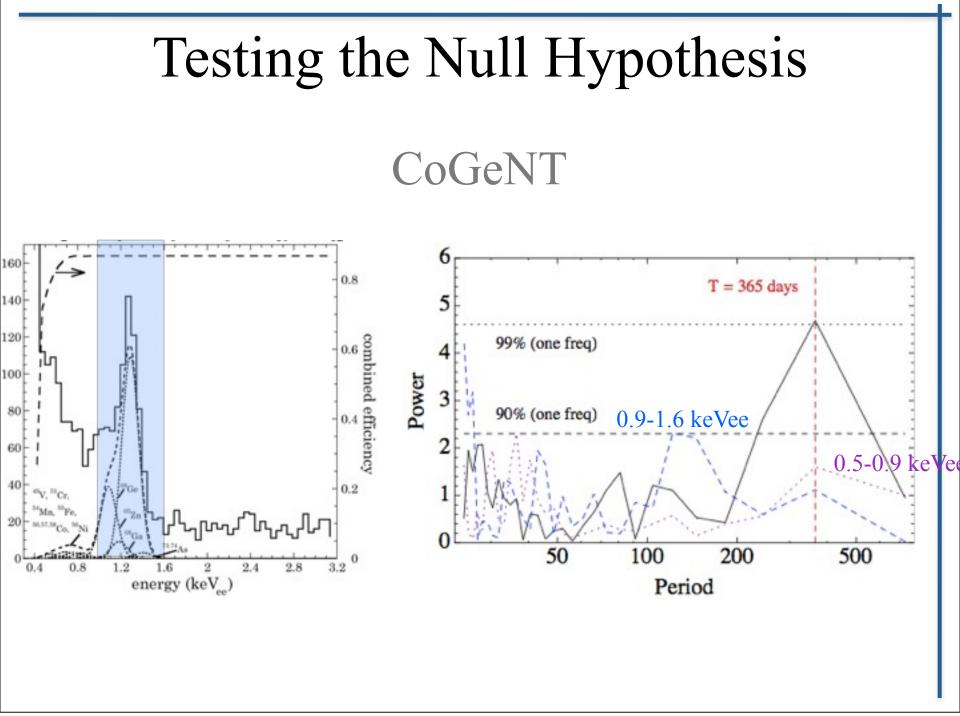
60

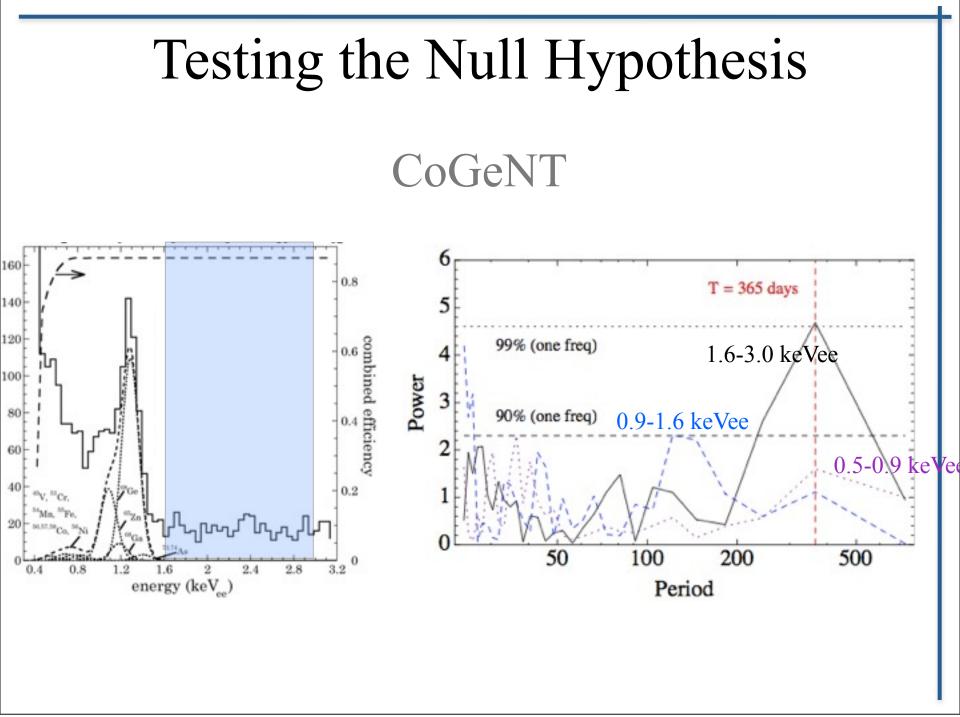
40

20



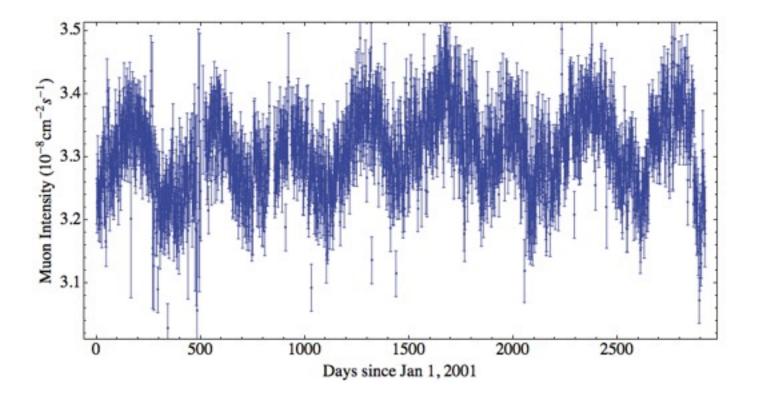






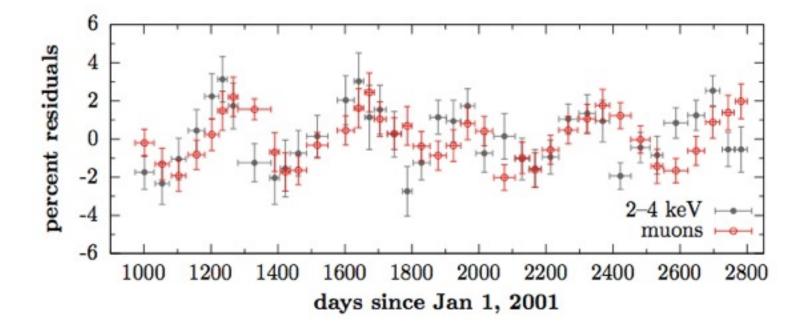
The Muon Hypothesis

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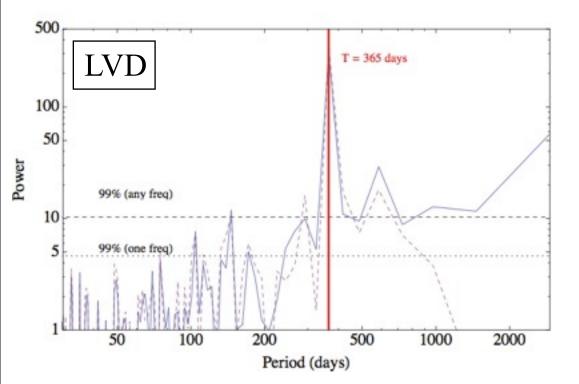


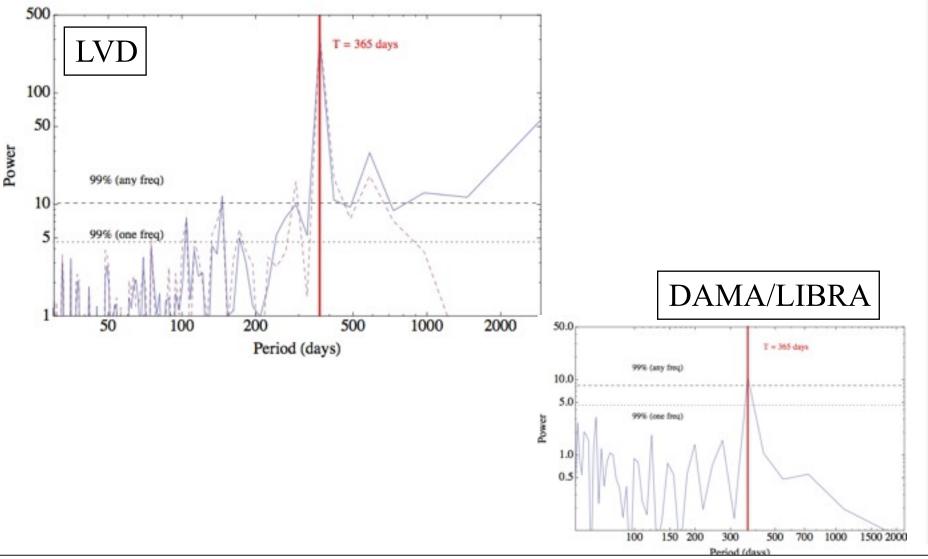
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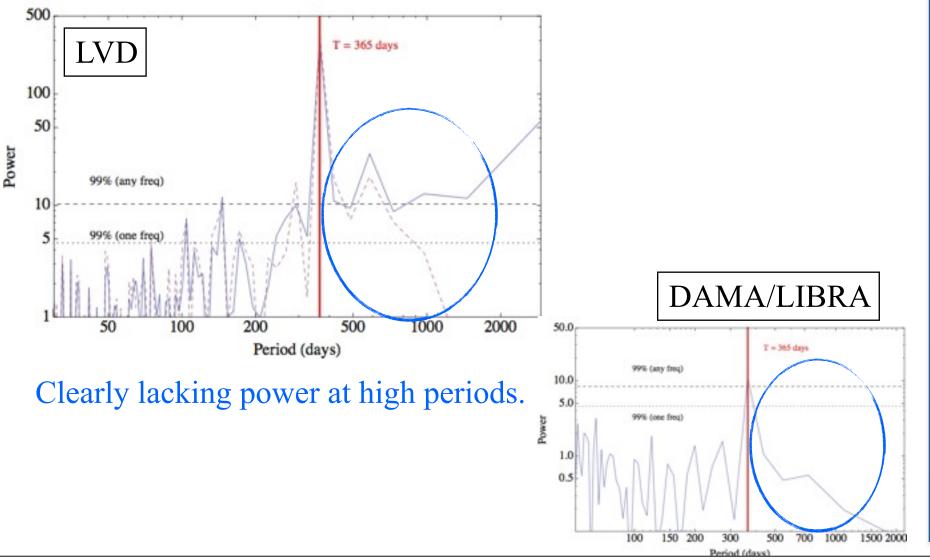
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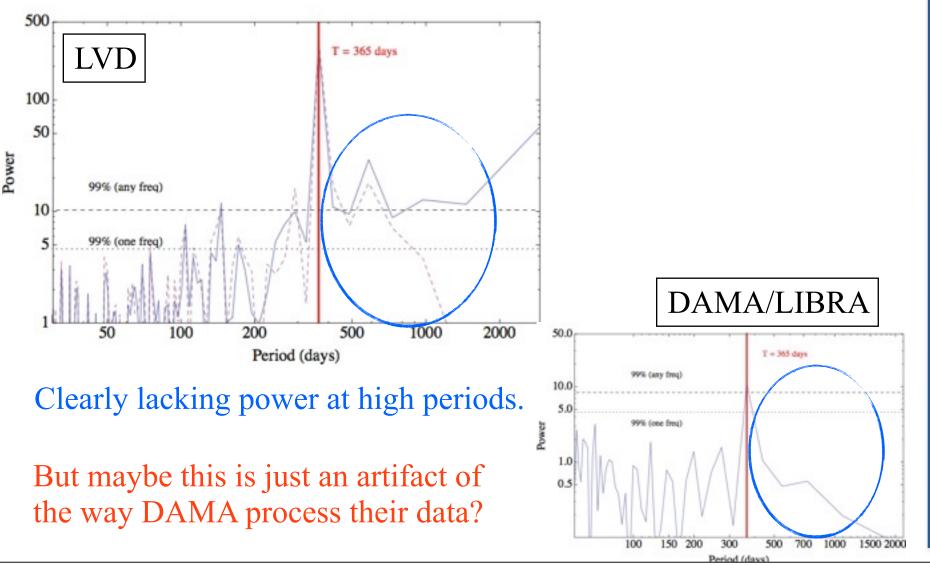


This has been discussed by several independent authors (Ralston 2010, Nygren 2011, Blum 2011).









We can compare the two datasets in phase-period space using a generalization of the Lomb-Scargle periodogram

Chang, Pradler, IY, 1111.4222

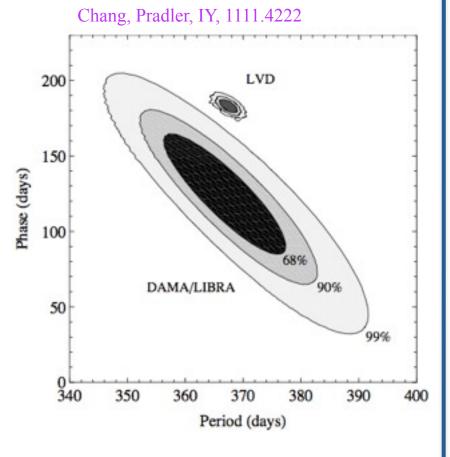
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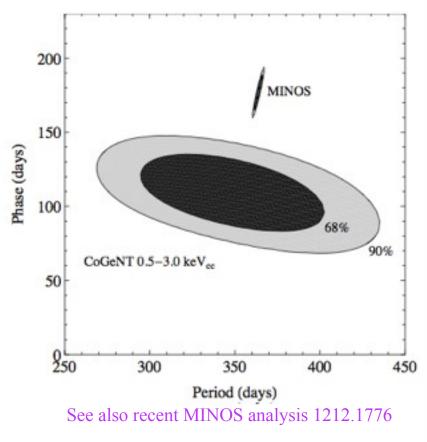
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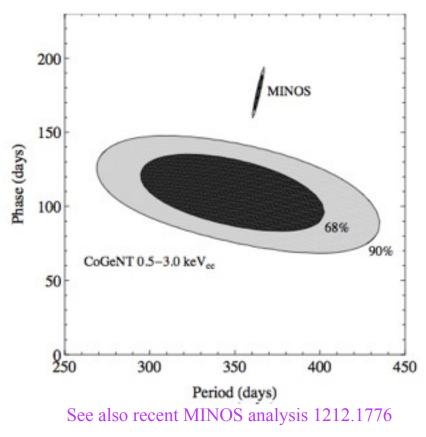


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But, maybe this is just an artifact of forcing the data into a sinusoidal modulation?

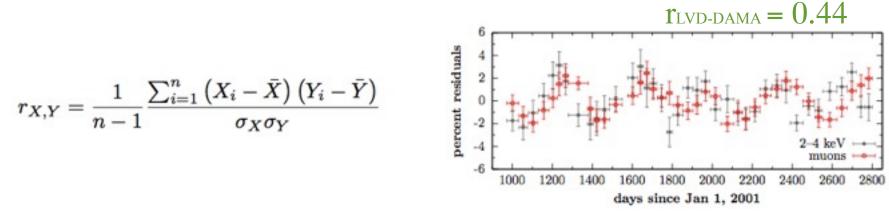
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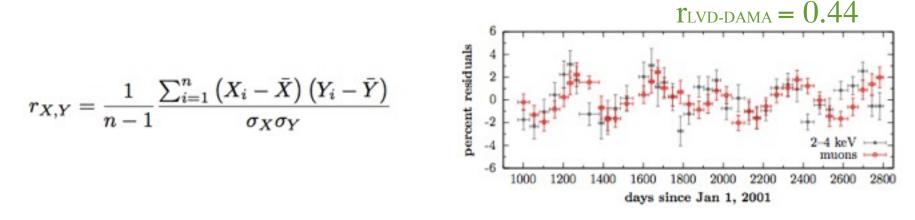
Without assuming any model, we can compute Pearson's coefficient of correlation between the two datasets.

$$r_{X,Y} = \frac{1}{n-1} \frac{\sum_{i=1}^{n} \left(X_i - \bar{X}\right) \left(Y_i - \bar{Y}\right)}{\sigma_X \sigma_Y}$$

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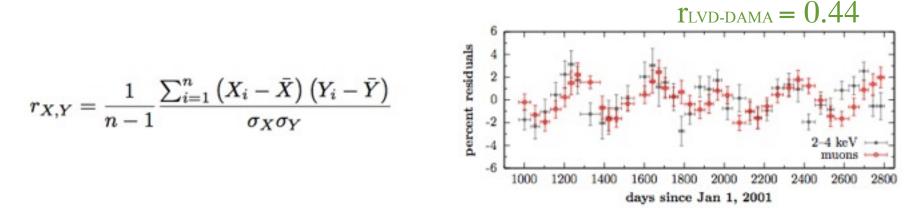


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$$s_i = rac{y N_{\mu,i}}{M \Delta E \epsilon_i \Delta t_i} egin{array}{c} ext{where } N_{\mu,i} ext{ is Poisson distributed with mean} \ & \langle N_{\mu,i}
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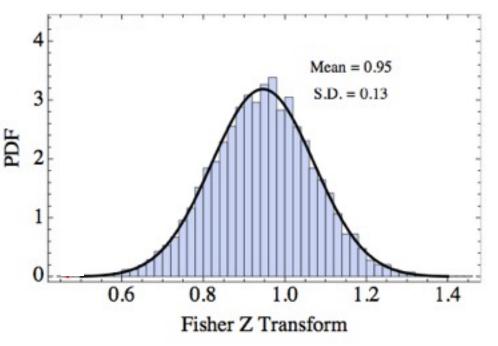
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- 3. Correlation is too weak
- 4. Phaseogram seems robust even under biased binning.

Models

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1) MiDM (Tucker-Smith and Weiner 2001, Chang, Weiner, IY, 2010)

2) LDM (Feldstein, Graham, & Rajendran 2010)

3) Light DM (Bottino, Donato, Fornengo, and Scopel, 2008, and others)

4) Channeling (Bozorgnia, Gelmini, and Gondolo 2010 and others)

5) RDM (Bai, Fox, 2009)

6) Leptonic DM (Bernabei 2008 and Kopp, Niro, Schwetz & Zupan 2009)

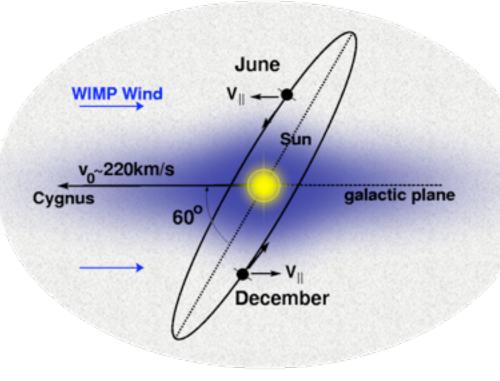
7) Neutrinos with new baryonic currents (Pospelov 2012)

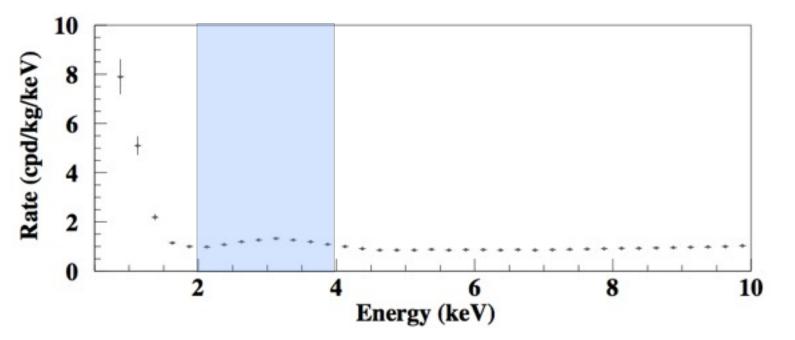
8) Isospin violating (Chang, Liu, Pierce, Weiner, IY, 2010 and Feng, Kumar, Marfatia & Sanford and others)

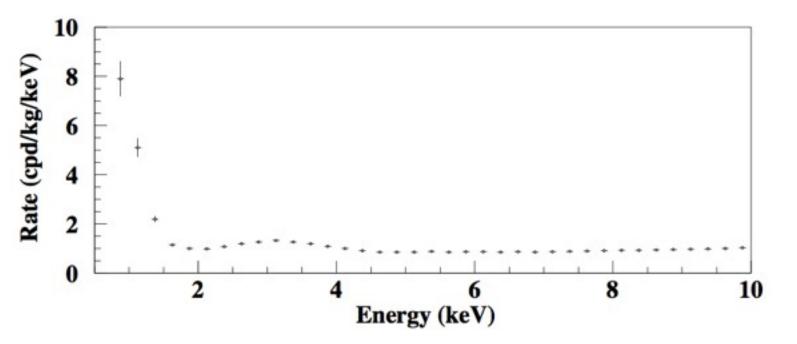
9) . . .

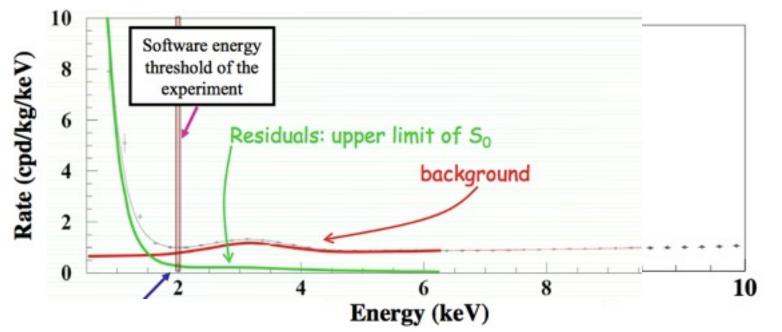
Modulation Fraction

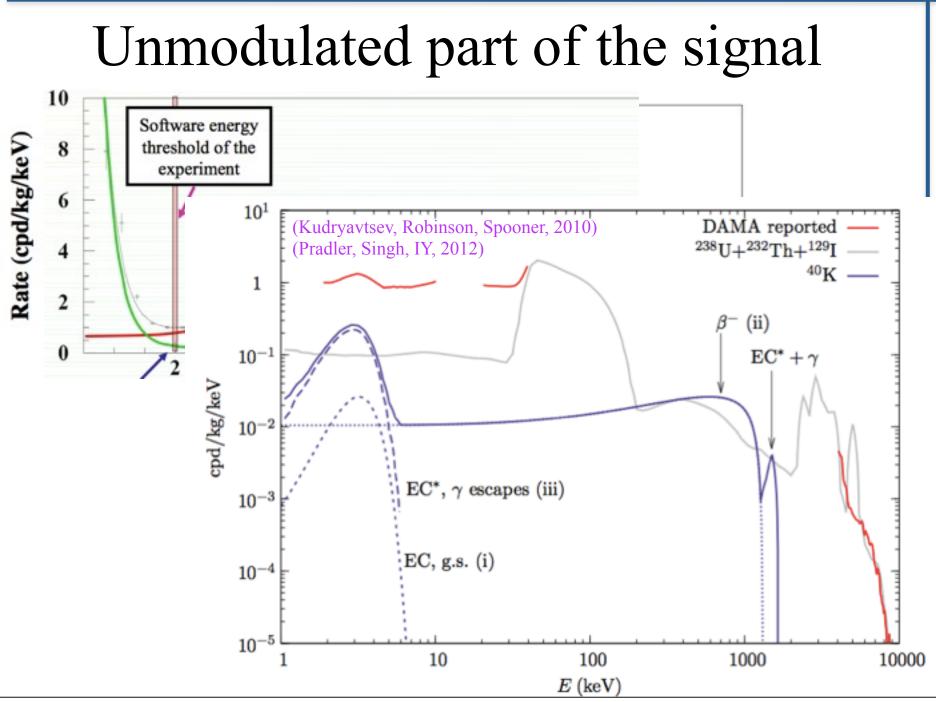
The modulation fraction expected in most models is no more than 10%. Is that compatible with the total number of events seen?



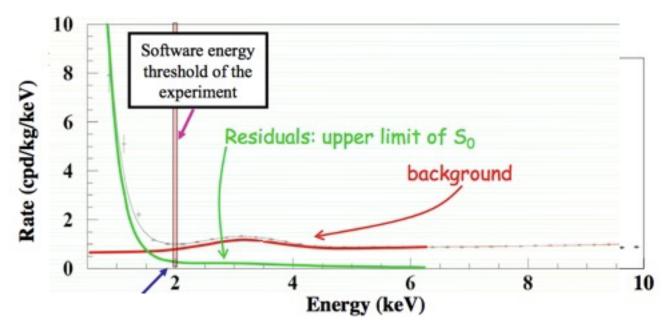




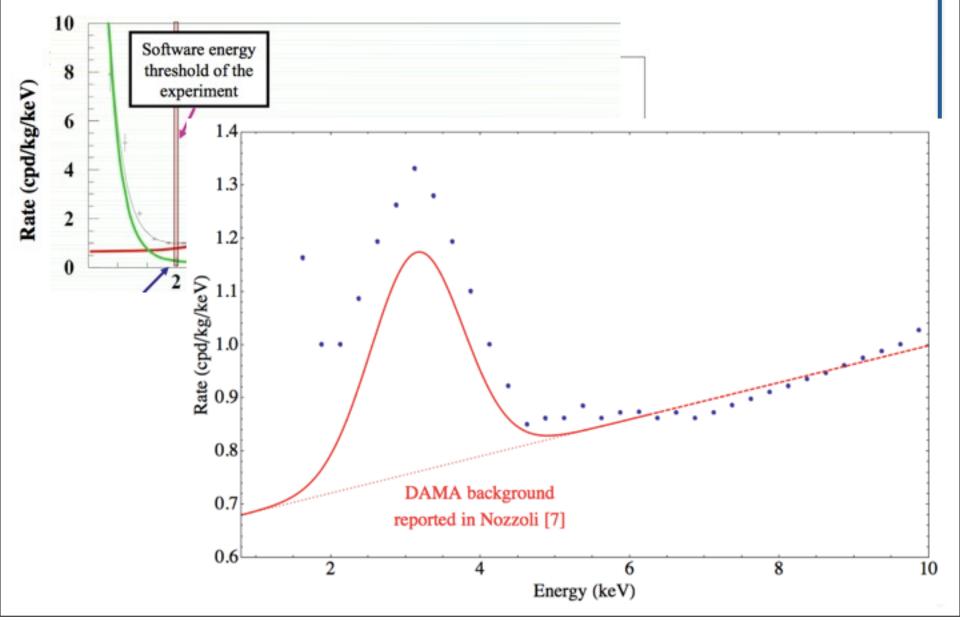




A Simple(r) Fit to the Background

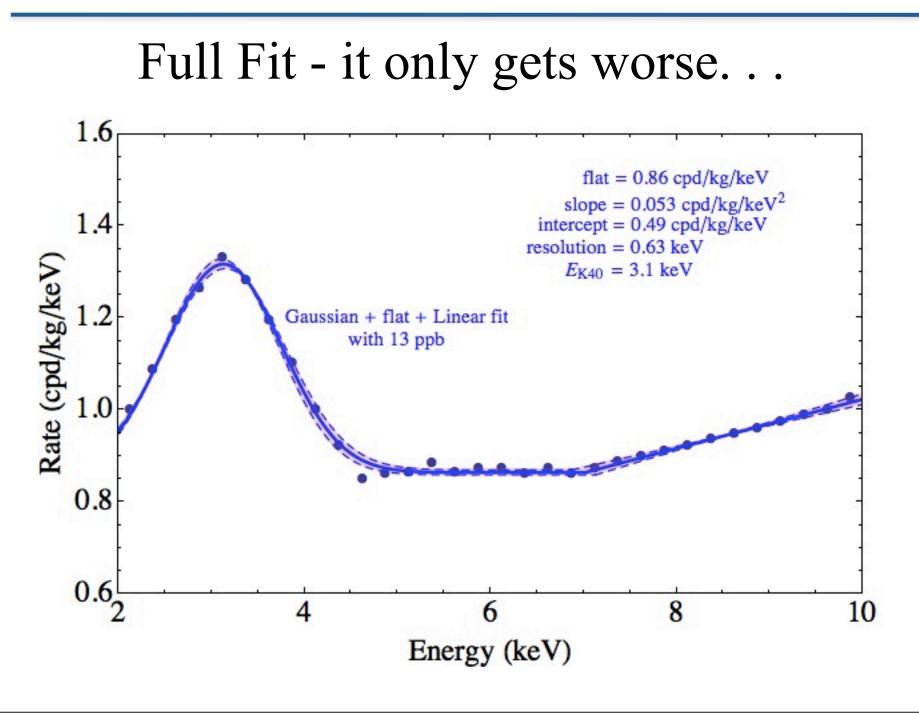


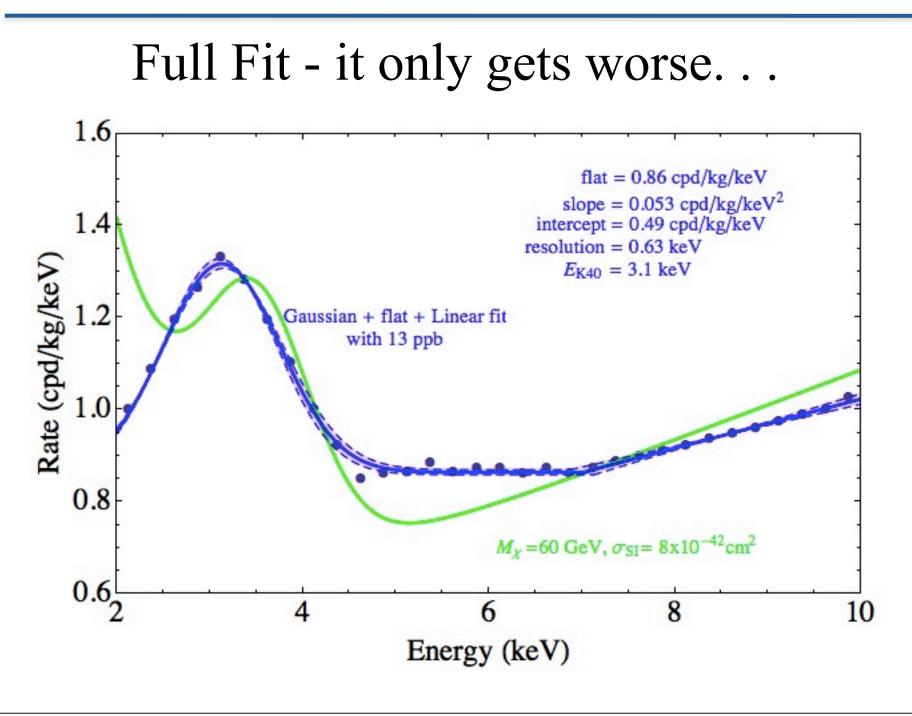
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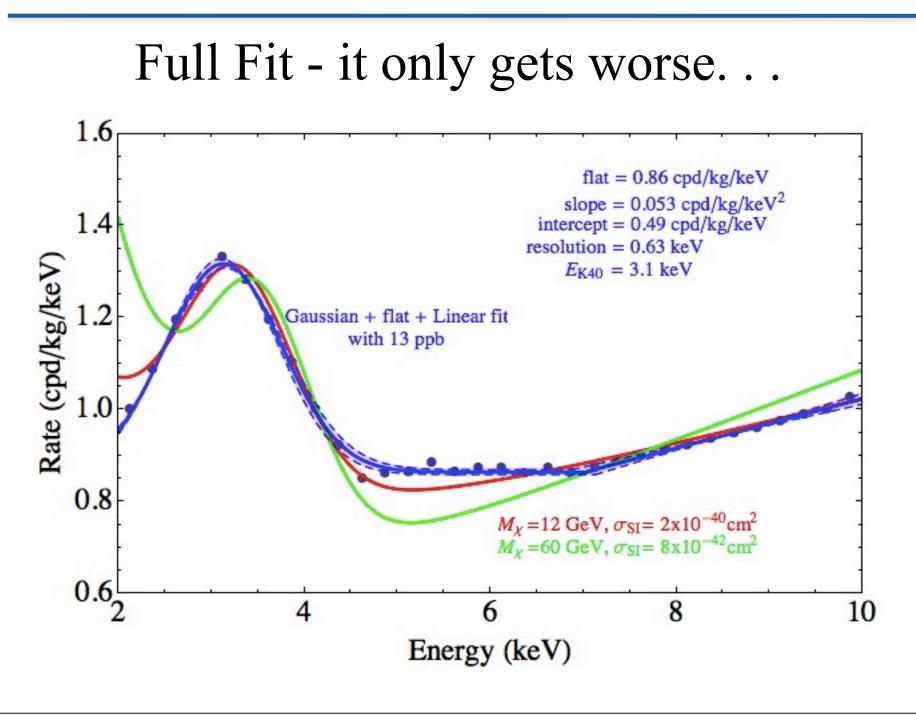


A Simple(r) Fit to the Background 10 Software energy threshold of the 8 Rate (cpd/kg/keV) experiment 6 1.4(Pradler, Singh, IY, 2012) (Pradler, IY, 2012) 1.3 2 1.2 0 Gaussian + Flat + Linear Fit 1.1 Rate (cpd/kg/keV) with 13 ppb 1.0 0.9 Flat background 0.85 cpd/keV/kg 0.8 used in arXiv:1210.5501 DAMA background 0.7 reported in Nozzoli [7] 0.6 2 6 8 10 Energy (keV)

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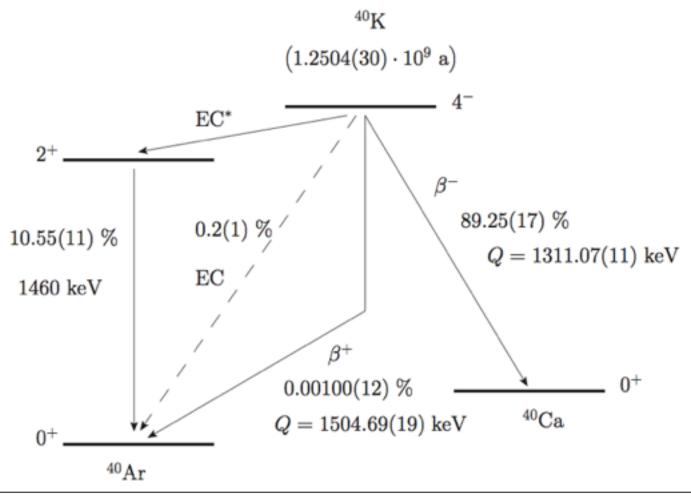






Unmeasured Nuclear Decay

One piece of nuclear physics that emerged out of it is the identification of an hitherto unmeasured special nuclear transition in potassium.



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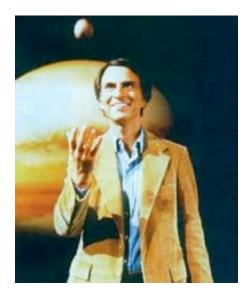
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3) But, something is modulating in DAMA, what is it? Maybe DM, but higher modulation fraction, maybe something more exotic? Maybe something more mundane . . .

4) Our aim is not to claim that we now understand the background. This might be complicated and messy. But considering the gravity of the claims (DM), it seems only appropriate to devote more attention to this issue.

Extraordinary Claims

"Extraordinary claims require extraordinary evidence." - Carl Sagan

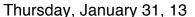


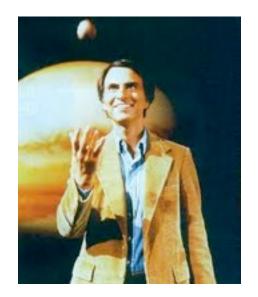
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Really, an experiment claiming something so extraordinary should release all its data. . .

