

CP & SUSY breaking and E_6 GUT

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with M. Bando, T. Yamashita

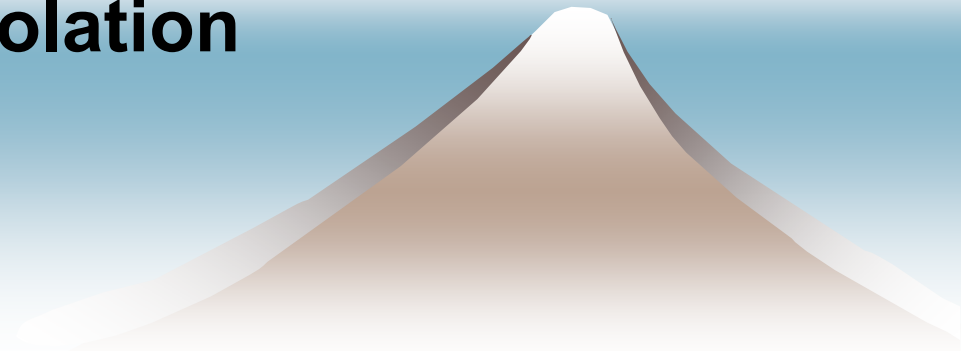
with S. Kim, A. Matsuzaki, K. Sakurai, T. Yoshikawa
M. Ishiduki

1. SUSY GUT and Yukawa structures
2. E_6 Unification for Matter sector

Why are larger neutrino mixings?

Horizontal symmetry to solve SUSY flavor problem

- New** 3. Prediction of E_6 GUT on FCNC
- New** 4. Spontaneous CP violation
5. Summary



Topics mentioned in this talk

- ◆ Quark&Lepton masses & mixings?
SU(5) GUT with an assumption \leftarrow E6 GUT
- ◆ SUSY flavor problem
E6+horizontal (family) symmetry
- New** ◆ Origin of CP violation(SUSY CP problem)
Spontaneous CP violation in E6+horizontal
KM phase &

CEDM constraints (Hisano-Shimizu) OK!!



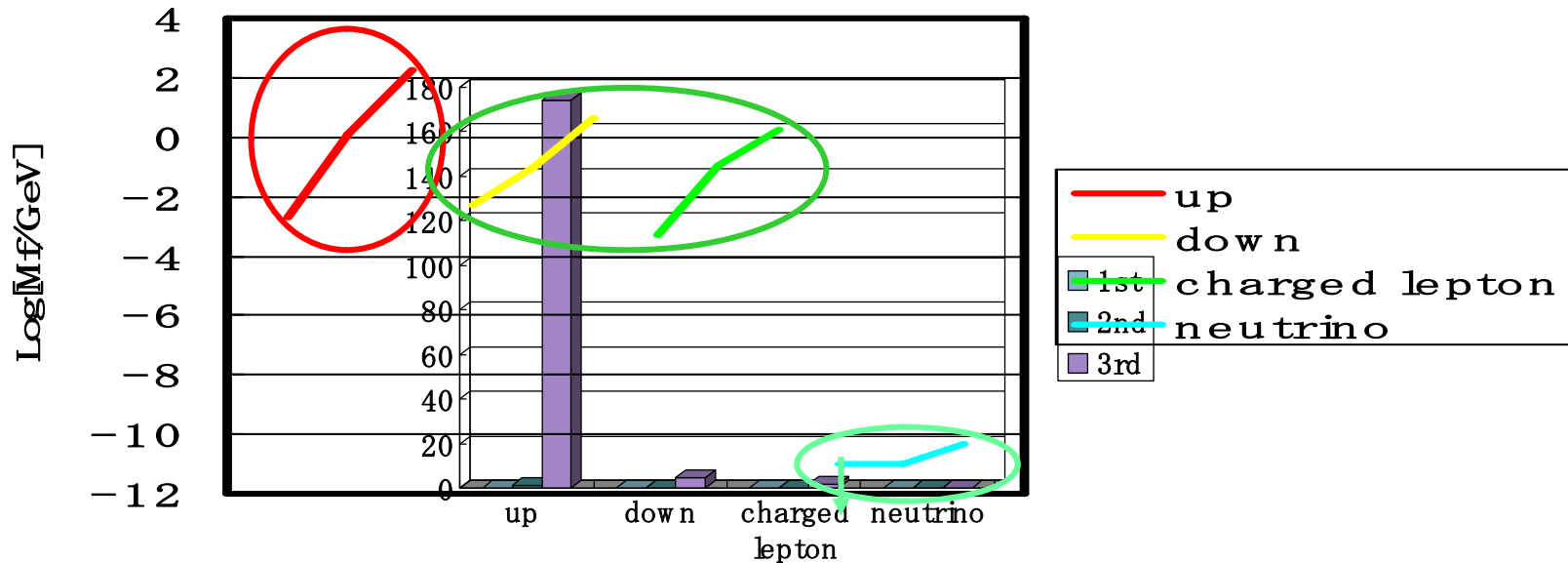
Quark&Lepton Masses & Mixings

SU(5) GUT with **an assumption** can explain these hierarchies



It may result in large FCNC in SUSY models

Masses & Mixings and GUT



u, c, t

Strongest

Neutrinos

Weakest

CKM

small mixings

e, μ , τ

Middle

MNS

large mixings

d, s, b

Middle

These can be naturally realized in SU(5) GUT!!

SU(5) SUSY GUT

Albright-Barr
Sato-Yanagida...

$$\mathbf{10} = (q, u_R^c, e_R^c) \quad \bar{\mathbf{5}} = (d_R^c, l) \quad \mathbf{1} = \nu_R^c$$

$$Y_u \mathbf{10}_i \mathbf{10}_j \mathbf{5}_H + Y_{(d,e)} \mathbf{10}_i \bar{\mathbf{5}}_j \bar{\mathbf{5}}_{\bar{H}} + \underbrace{Y_{\nu_D} \bar{\mathbf{5}}_i \mathbf{1}_j \mathbf{5}_H + M_{\nu_R} \mathbf{1}_i \mathbf{1}_j}_{\frac{Y_{\nu} \bar{\mathbf{5}}_i \bar{\mathbf{5}}_j \mathbf{5}_H \mathbf{5}_H}$$

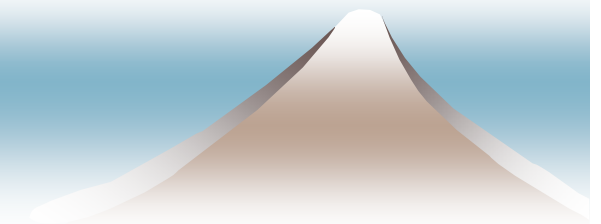
$$u \gg d, e \gg \nu$$

➔ $\mathbf{10}_i$ have stronger hierarchy than $\bar{\mathbf{5}}_i$

Stronger hierarchy leads to smaller mixings

➔ Quark mixings(CKM) \ll Lepton mixing(MNS)

$\mathbf{10}_i(q_i)$ $\bar{\mathbf{5}}_i(l_i)$



Mass hierarchy and mixings

- ◆ Stronger hierarchy leads to smaller mixings

$$\begin{pmatrix} (\epsilon^2) & \epsilon \\ \epsilon & 1 \end{pmatrix} \rightarrow 1, \epsilon^2$$
$$\begin{pmatrix} 1 & \epsilon \\ -\epsilon & 1 \end{pmatrix}$$

Stronger hierarchy \longleftrightarrow **Smaller** mixings

SU(5) SUSY GUT

$$\mathbf{10} = (q, u_R^c, e_R^c) \quad \bar{\mathbf{5}} = (d_R^c, l) \quad \mathbf{1} = \nu_R^c$$

$$Y_u \mathbf{10}_i \mathbf{10}_j \mathbf{5}_H + Y_{(d,e)} \mathbf{10}_i \bar{\mathbf{5}}_j \bar{\mathbf{5}}_{\bar{H}} + \underbrace{Y_{\nu_D} \mathbf{5}_i \mathbf{1}_j \mathbf{5}_H + M_{\nu_R} \mathbf{1}_i \mathbf{1}_j}_{\frac{Y_\nu}{M} \bar{\mathbf{5}}_i \bar{\mathbf{5}}_j \mathbf{5}_H \mathbf{5}_H}$$

$$u \gg d, e \gg \nu$$

→ $\mathbf{10}_i$ have stronger hierarchy than $\bar{\mathbf{5}}_i$

Stronger hierarchy leads to smaller mixings

→ Quark mixings(CKM) \ll Lepton mixing(MNS)

$$\mathbf{10}_i (q_i)$$

$$\bar{\mathbf{5}}_i (l_i)$$

Good agreement with masses & mixings

Large mixings and FCNC

- ◆ Even if universal sfermion masses at cutoff, radiative correction induces non-universality.

$$(\delta_{e_L})_{12} \rightarrow \text{Br}(l_i \rightarrow l_j \gamma) \propto (V_L^\dagger \hat{Y}_{\nu_D}^2 V_L \tan \beta)^2$$

Borzumati-Masiero 85

Hisano-Moroi-Tobe-Yamaguchi-Yanagida Barbieri-Hall-Strumia 95

$$(\delta_{d_R})_{23} \rightarrow S_{K\phi}, \Delta M_{B_s}$$

Moroi 00 (SU(5))

Chang-Masiero-Murayama 02 (SO(10))

- ◆ Large mixings for $\bar{\mathbf{5}} = (d_R^c, l)$ $V_L \sim V_{MNS}$

- ◆ SO(10) GUT relation $Y_{\nu_D} \sim Y_u \sim Y_d$

↔ Large $\tan \beta, Y_{\nu_D}$

→ Large FCNC are expected.

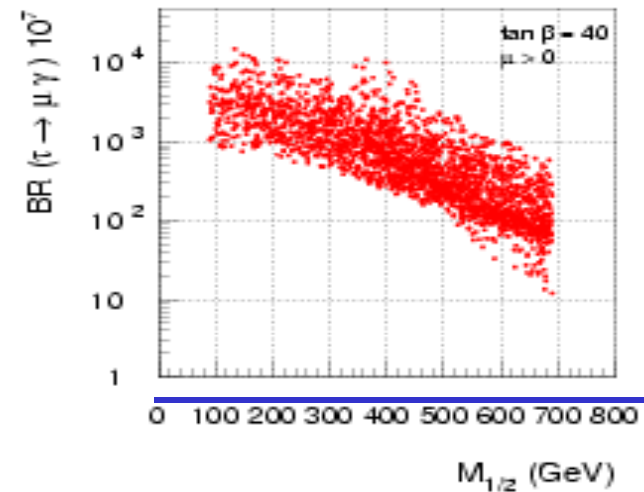
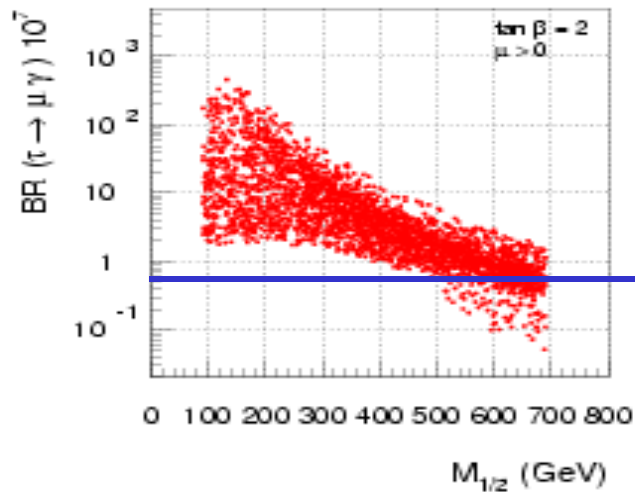
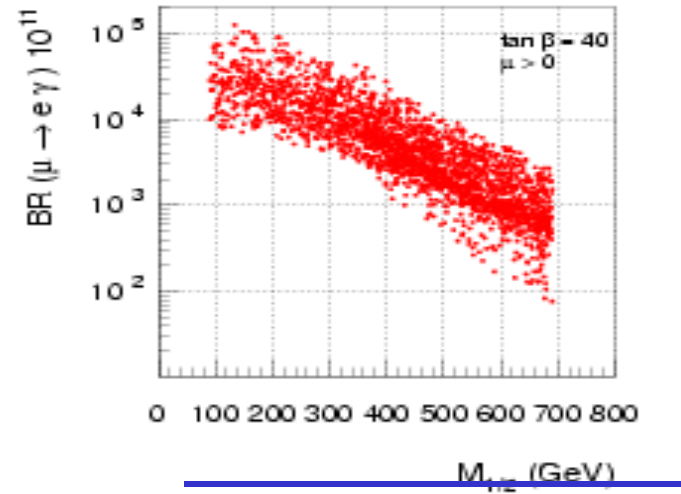
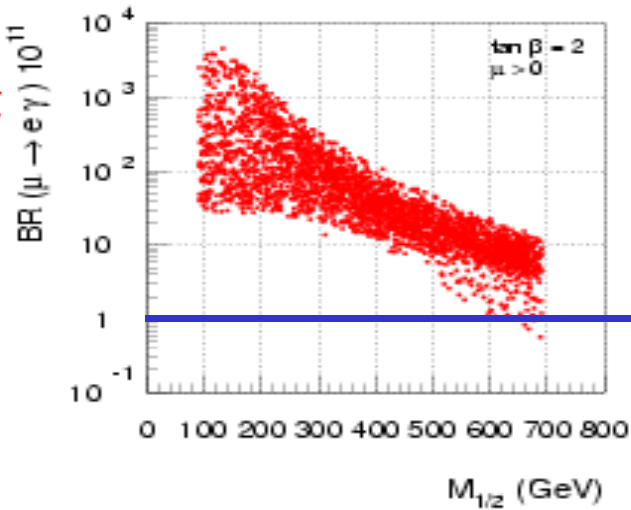
Lepton Flavor Vi

$\sim 10^{-6} + i \sim 10^{-6} \text{ SO}(10)$

$$V_L \sim V_{MNS}$$

Masiero, Vempati, Vives 02

$$Y_{\nu D} \sim Y_t$$



Constraints to $(\delta_{d_R})_{ij}$

- ◆ EDM of Hg (neutron) Hisano-Shimizu 03

$$\text{Im} (\delta_{d_L})_{13}(\delta_{d_R})_{31} < 3(0.2) \times 10^{-3}$$

$$(\delta_{d_L})_{13} \sim 0.01 \rightarrow (\delta_{d_R})_{13} < 0.02$$

- ◆ Bs mixing CDF 06

$$\sqrt{(\delta_{d_L})_{23}(\delta_{d_R})_{23}} < 0.02$$

$$(\delta_{d_L})_{23} \sim 0.04 \rightarrow (\delta_{d_R})_{23} < \mathbf{0.01}$$

$$\tilde{m}_q \sim 500 \text{ GeV}$$

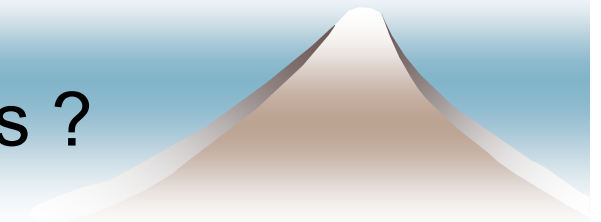
1st Summary & Questions

- ◆ SU(5) GUT is in good agreement with the hierarchies of quark & lepton masses and mixings.

$$\begin{array}{ccc}
 Y_u \sim \begin{array}{c} \text{red bars} \\ \text{decreasing} \end{array} & Y_{(d,e)} \sim \begin{array}{c} \text{purple bars} \\ \text{decreasing} \end{array} & Y_\nu \sim \begin{array}{c} \text{blue bars} \\ \text{decreasing} \end{array} \\
 V_{CKM} \sim \begin{array}{c} \text{red bars} \\ \text{decreasing} \end{array} & & V_{MNS} \sim \begin{array}{c} \text{blue bars} \\ \text{decreasing} \end{array}
 \end{array}$$

10_i has stronger hierarchy than $\bar{5}_i$

- ◆ Suppressed FCNC & EDM $\rightarrow \delta_{\bar{5}} \ll 1 \rightarrow Y_{\nu D} \ll 1$
SO(10) GUT relation $Y_{\nu D} \sim Y_u \sim Y_d$ looks bad.
- ◆ More unification of quark and lepton ?
- ◆ Origin of the assumption?
- ◆ Origin of various Yukawa hierarchies ?



**E6 GUT can answer
these questions !**



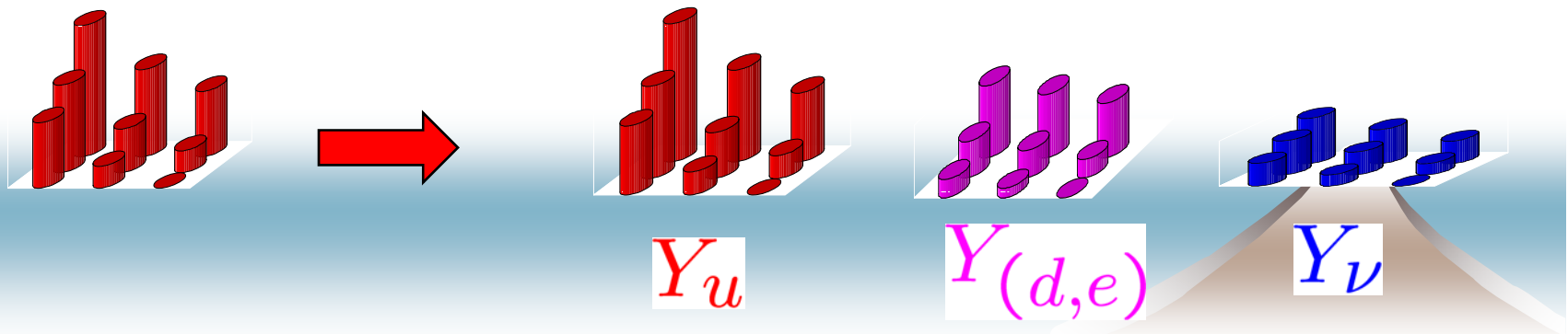
E6 GUT answers these questions

- ◆ More unification of quark and lepton ? Yes

$$E6 \supset SO(10) \quad 27 = 16 + 10 + 1$$

- ◆ Origin of the assumption ?
- ◆ Origin of various Yukawa hierarchies ?

Various Yukawa hierarchies can be induced from one Yukawa hierarchy in E6 GUT.



E_6 i f i c a t i o n

Guisey-Ramond-Sikivie,
Aichiman-Stech, Shafi,
Barbieri-Nanopoulos,
Bando-Kugo,...

$$27_i = 16_i [10_i + \bar{5}_i + 1_i] + 10_i [5_i + \bar{5}_i] + 1_i [1_i]$$

$(i = 1, 2, 3)$
 $\langle 16_C \rangle$
 $\langle 1_H \rangle$

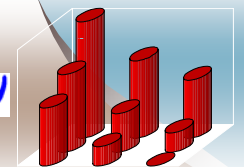
Three of six $\bar{5}$ become superheavy after the breaking

$$E_6 \xrightarrow{\langle 1_H \rangle} SO(10) \xrightarrow{\langle 16_C \rangle} SU(5) \quad \langle 1_H \rangle \geq \langle 16_C \rangle$$

$$W = Y^H 27_i 27_j \langle 27_H \rangle + Y^C 27_i 27_j \langle 27_C \rangle$$

Once we fix $Y^H, Y^C, \langle 27_H \rangle, \langle 27_C \rangle$,
three light modes of six $\bar{5}$ are determined.

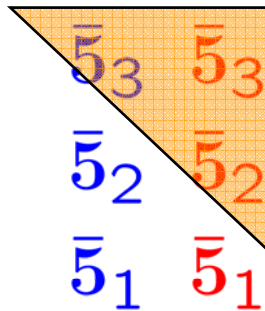
We assume all Yukawa matrices



Milder hierarchy for $\bar{5}_i(l)$

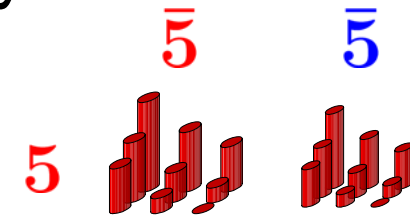
Bando-N.M. 01
N.M, T. Yamashita 02

- ◆ $\bar{5}$ fields from 27_3 become superheavy.



Superheavy

unless $\langle 27_H \rangle \gg \langle 27_C \rangle$



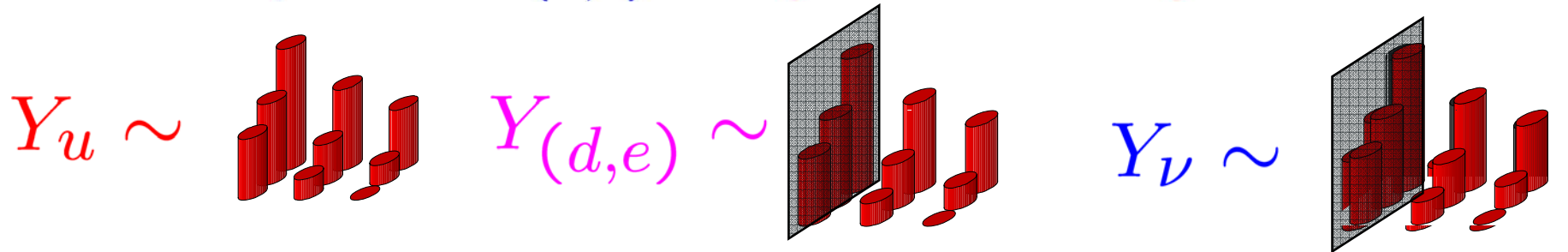
- ◆ Light modes ($\bar{5}_1, \bar{5}_1, \bar{5}_2$) have smaller Yukawa couplings and milder hierarchy than ($10_1, 10_2, 10_3$)

$$Y_{\nu D}, Y_d \ll Y_u$$

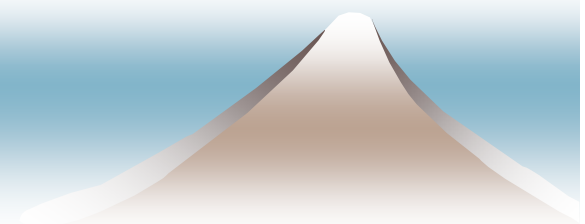
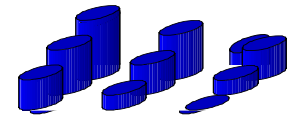
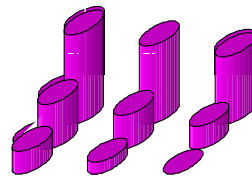
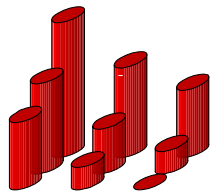
- Larger mixings in lepton sector than in quark sector.
 - Small $\tan \beta$
 - Small neutrino Dirac masses
- } Suppressed radiative LFV

How to obtain various Yukawas?

$$Y_u 10_i 10_j 5_H + Y_{(d,e)} 10_i \bar{5}_j \bar{5}_{\bar{H}} + Y_\nu \bar{5}_i \bar{5}_j 5_H 5_H$$



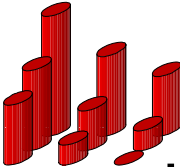
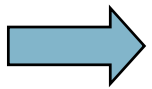
$$(\bar{5}_1, \bar{5}_2, \bar{5}_3) \rightarrow (\bar{5}_1, \bar{5}'_1, \bar{5}_2)$$

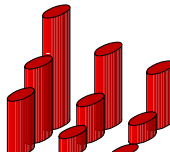


2nd Summary

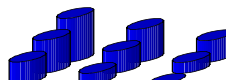
- ◆ E_6 unification explains why the lepton sector has larger mixings than the quark sector.
- ◆ Suppressed radiative LFV

Small $Y_{\nu D}$ Small $\tan \beta$

- ◆ A basic Yukawa hierarchy $Y \sim Y_u$ 
- 
 The other Yukawa hierarchies

$$Y_u \sim \text{$$

$$Y_{(d,e)} \sim \text{$$

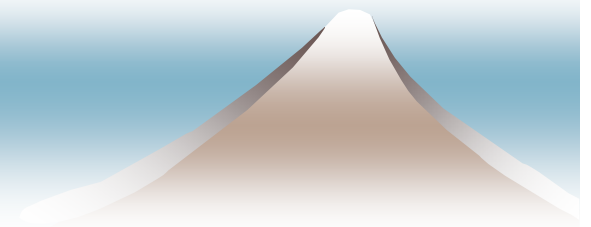
$$Y_\nu \sim \text{$$

Hierarchy of 10_i is stronger than that of $\bar{5}_i$

Three $\bar{5}_i$ come from the first 2 generation of $27_1, 27_2$

SUSY flavor problem

- ◆ E6+horizontal (family) symmetry
- ◆ More structures for suppressing FCNC.



o r i z o n t a l S y m m e t r y

Dine-Kagan-Leigh
Pomarol-Tommasini
Barbieri-Hall...

- Origin of Yukawa hierarchy
- Universal sfermion masses to suppress FCNC

$$\Phi_a, \Phi_3, H_u, H_d (\Phi = Q, U, D, L, E, N)$$

- ◆ The 1st 2 generation have universal sfermion masses.
- ◆ Large top Yukawa coupling

$$U(2)_H \longrightarrow U(1)_H \longrightarrow X$$

$$\langle \bar{F}^a \rangle / \Lambda \sim \epsilon \quad \langle A^{ab} \rangle / \Lambda \sim \epsilon'$$

$$Y \sim \begin{pmatrix} 0 & \epsilon' & 0 \\ \epsilon' & \epsilon^2 & \epsilon \\ 0 & \epsilon & 1 \end{pmatrix} \quad \tilde{m}_f^2 \sim \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 + \epsilon^2 & \epsilon \\ 0 & \epsilon & O(1) \end{pmatrix} \tilde{m}_{\frac{3}{2}}^2$$


$$Y_u \sim Y_d \sim Y_e \sim Y_\nu?$$

Not sufficient to suppress FCNC

Large neutrino mixings and FCNC

- ◆ The universal sfermion masses only for the 1st 2 generation $\bar{5}$ do not suppress FCNC sufficiently

$$\delta_{\bar{5}-1} \sim V_{\bar{5}}^\dagger \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & a \end{pmatrix} V_{\bar{5}} \sim a \begin{pmatrix} \lambda^2 & \lambda^{1.5} & \lambda \\ \lambda^{1.5} & \lambda & \lambda^{0.5} \\ \lambda & \lambda^{0.5} & 1 \end{pmatrix} \quad \lambda \sim 0.2$$

if $V_{\bar{5}} \sim V_{MNS}$.  $a \sim O(1)$

$$|\text{Im}(\delta_{d_R^c})_{12}| \leq 1.5 \times 10^{-3} \left(\frac{\tilde{m}_Q}{500\text{GeV}} \right)$$

$$|(\delta_l)_{12}| \leq 4 \times 10^{-3} \left(\frac{\tilde{m}_l}{100\text{GeV}} \right)^2$$

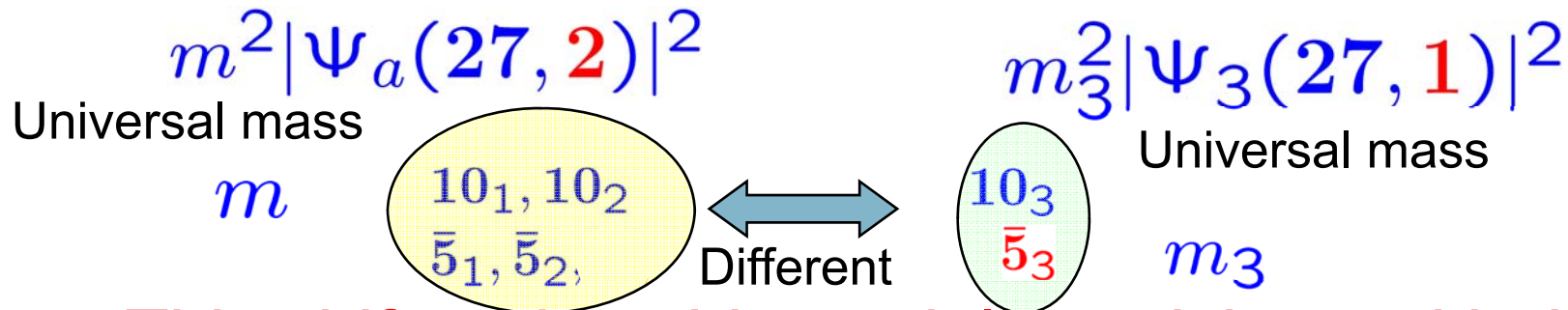
Universality for all three generations is required!

E_6 +horizontal symmetry N.M. 02,04 N.M, T. Yamashita 04

- ◆ In E_6 various Yukawa hierarchies are produced from one basic hierarchy from $U(2)$ breaking.
- ◆ Unification of generations by $U(2)_H$ (or $U(3)_H$) realizes the universality of sfermion masses.

$$\Psi_a(27, \mathbf{2}) + \Psi_3(27, \mathbf{1}) \quad (a = 1, 2)$$

- ◆ A prediction for sfermion mass spectrum



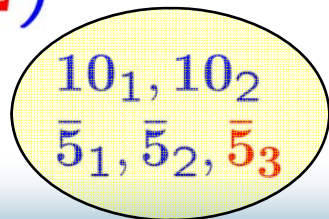
- ◆ This shift makes this model consistent with the present experiments. (Universal masses for 5̄_i)
- ◆ The 3rd generation FCNC can be large. (Testable)

Structures suppressing FCNC for $\bar{5}_i$

- ◆ Small Yukawa couplings
- ◆ Small $\tan \beta$
- ◆ Universal sfermion masses for $\bar{5}_i$
- ◆ m can increase without destabilizing the weak scale. (Effective SUSY)

$\Psi_a(27, 2)$

m



$\Psi_3(27, 1)$

10₃

m_3

$$m \gg m_3 \sim O(m_W)$$

How does FCNC processes take place in this model?

flavor violating

$$10: \begin{pmatrix} m^2 & & \\ & m^2 & \\ & & m_3^2 \end{pmatrix}$$

$10 : (Q, U_R^c, E_R^c)$

No source of flavor violation

$$\bar{5}: \begin{pmatrix} m^2 & & \\ & m^2 & \\ & & m^2 \end{pmatrix}$$

$\bar{5} : (L, D_R^c)$

For example, for the right-handed charged slepton sector,

$$\bar{e}_R^\dagger \begin{pmatrix} m^2 & & \\ & m^2 & \\ & & m_3^2 \end{pmatrix} \bar{e}_R \rightarrow \bar{e}_R^\dagger V^\dagger \begin{pmatrix} m^2 & & \\ & m^2 & \\ & & m_3^2 \end{pmatrix} V \bar{e}_R = \bar{e}_R^\dagger \tilde{m}_{\bar{e}_R}^2 \bar{e}_R$$

Since 10 contains Q, the form of unitary matrix V is CKM-like. We can parametrize it with Cabibbo angle λ .

$$V \sim \begin{pmatrix} 1 & \lambda & \lambda^3 \\ \lambda & 1 & \lambda^2 \\ \lambda^3 & \lambda^2 & 1 \end{pmatrix} \quad \lambda = 0.22$$

$$\Delta m^2 = (m_3^2 - m^2)$$

$$\tilde{m}_{\bar{e}_R}^2 = \tilde{m}_{\bar{q}_L}^2 = \tilde{m}_{\bar{u}_R}^2 = V^\dagger \begin{pmatrix} m^2 & & \\ & m^2 & \\ & & m_3^2 \end{pmatrix} V \sim \begin{pmatrix} m^2 & \Delta m^2 \lambda^5 & \Delta m^2 \lambda^3 \\ \Delta m^2 \lambda^5 & m^2 & \Delta m^2 \lambda^2 \\ \Delta m^3 \lambda^3 & \Delta m^2 \lambda^2 & m_3^2 \end{pmatrix}$$

Non universal SUSY breaking

- ◆ Universal sfermion masses for $\bar{5}$ fields

$$\delta_{\bar{5}} \sim \delta_{d_R^c} \sim \delta_l \sim 0$$

- ◆ Non universality for **10** fields

$$\delta_{10} \sim \delta_q \sim \delta_{u_R} \sim \delta_{e_R}$$
$$\sim V_{CKM}^\dagger \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix} V_{CKM} \sim \begin{pmatrix} \lambda^6 & \lambda^5 & \lambda^3 \\ \lambda^5 & \lambda^4 & \lambda^2 \\ \lambda^3 & \lambda^2 & 1 \end{pmatrix}$$

$$V_q \sim V_{u_R} \sim V_{e_R} \sim V_{CKM}$$

- ◆ Weak scale stability requires $m_{\frac{2}{3}} \sim O((100\text{GeV})^2)$
but almost no constraint for m_0

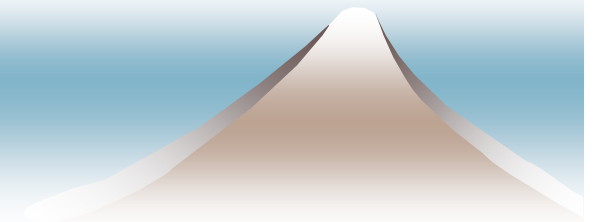
Predictions of E6 GUT +horizontal symmetry

Kim-N.M.-Matsuzaki-Sakurai-Yoshikawa

m_3 must be around the weak scale,
because of the **stability of the weak scale**,
while m can be taken larger.

$$10: \begin{pmatrix} m^2 & & \\ & m^2 & \\ & & m_3^2 \end{pmatrix}$$

$$5: \begin{pmatrix} m^2 & & \\ & m^2 & \\ & & m^2 \end{pmatrix}$$



Non decoupling feature of this model (in lepton flavor violation)

$$\tilde{m}_{\tilde{e}_R}^2 \sim \begin{pmatrix} m^2 & \Delta m^2 \lambda^5 & \Delta m^2 \lambda^3 \\ \Delta m^2 \lambda^5 & m^2 & \Delta m^2 \lambda^2 \\ \Delta m^3 \lambda^3 & \Delta m^2 \lambda^2 & m_3^2 \end{pmatrix} \quad \lambda = 0.22$$

$$\Delta m^2 = (m_3^2 - m^2)$$

- By picking up the 3-2 element, the size of $\tau \rightarrow \mu$ transition rate is order λ^2

$$\tau \rightarrow \mu \gamma \quad \tilde{\tau}_R \frac{(\tilde{m}_{\tilde{e}_R}^2)_{32}}{\times} \tilde{\mu}_R \approx \frac{1}{m_3^2} \Delta m^2 \lambda^2 \frac{1}{m^2} \longrightarrow \frac{\lambda^2}{m_3^2}$$

- For $\mu \rightarrow e \gamma$, there are two passes to change the flavor $\mu \rightarrow e$. Both they are order λ^5

$$\mu \rightarrow e \gamma \quad \tilde{\mu}_R \frac{(\tilde{m}_{\tilde{e}_R}^2)_{21}}{\times} \tilde{e}_R \approx \frac{1}{m^2} \Delta m^2 \lambda^5 \frac{1}{m^2} \longrightarrow 0$$

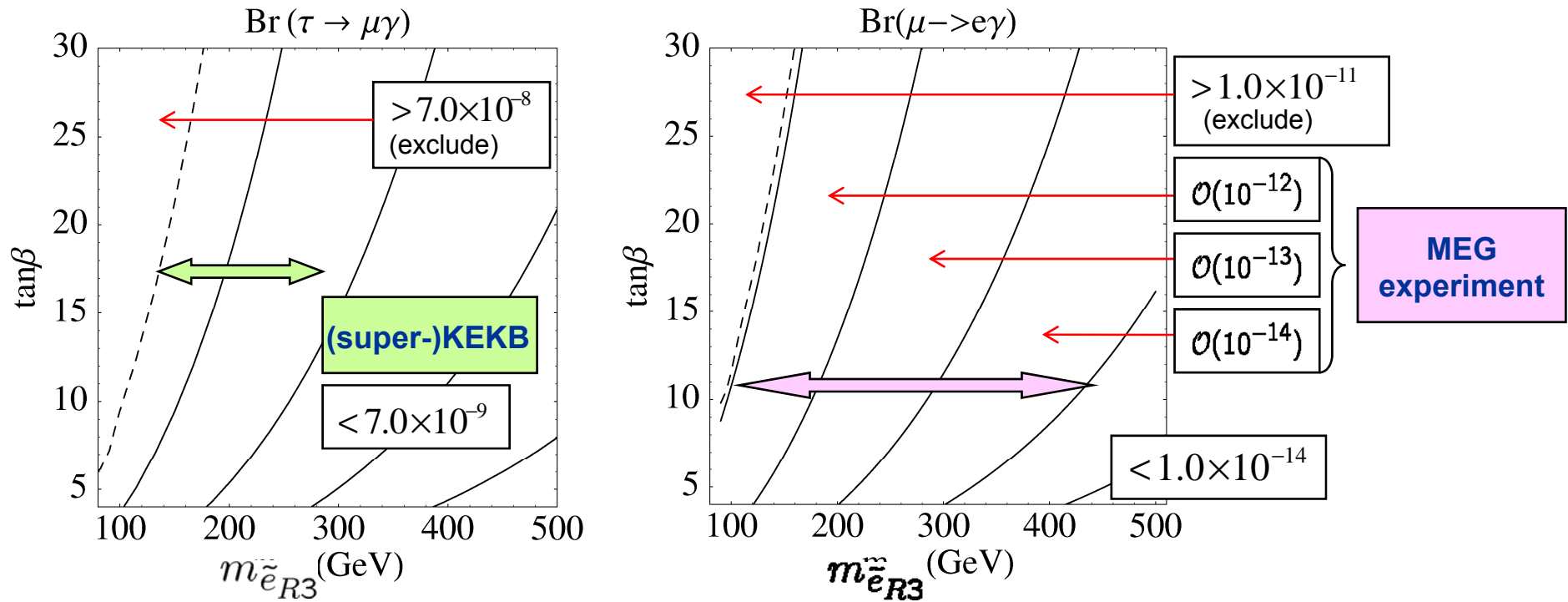
$$\tilde{\mu}_R \frac{(\tilde{m}_{\tilde{e}_R}^2)_{23}}{\times} \tilde{\tau}_R \frac{(\tilde{m}_{\tilde{e}_R}^2)_{31}}{\times} \tilde{e}_R \approx \frac{1}{m^2} \frac{(\Delta m^2)^2 \lambda^5}{m_3^2} \frac{1}{m^2} \longrightarrow \frac{\lambda^5}{m_3^2}$$

If we raise overall SUSY scale $m \dots$

$$m^2 \longrightarrow \infty$$

Propagator suppression from 1 or 2 generation becomes stronger, but mass difference Δm^2 increase. As a result, **both transition rate remain finite, and don't decouple!**

Can we discover the LFV at the future experiments?

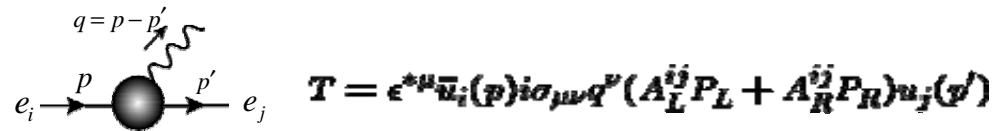


- $\tau \rightarrow \mu\gamma$ ➔ Detectable, when $\tan\beta$ is large and $m_{\tilde{e}_{R3}} < 250\text{GeV}$
- $\mu \rightarrow e\gamma$ ➔ Detectable unless $m_{\tilde{e}_{R3}} < 400\text{GeV}$

$M_2 = 120\text{GeV}$
 $\mu = 200\text{GeV}$
 $m_{\tilde{e}_{R1,2}} = 600\text{GeV}$

This model says that final state lepton tends to be right-handed.

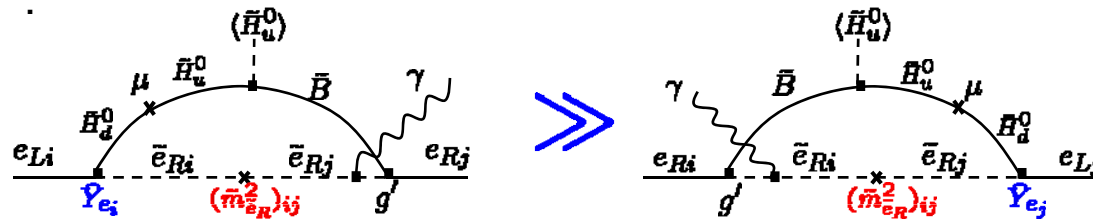
- Final state lepton has different chirality from initial one.



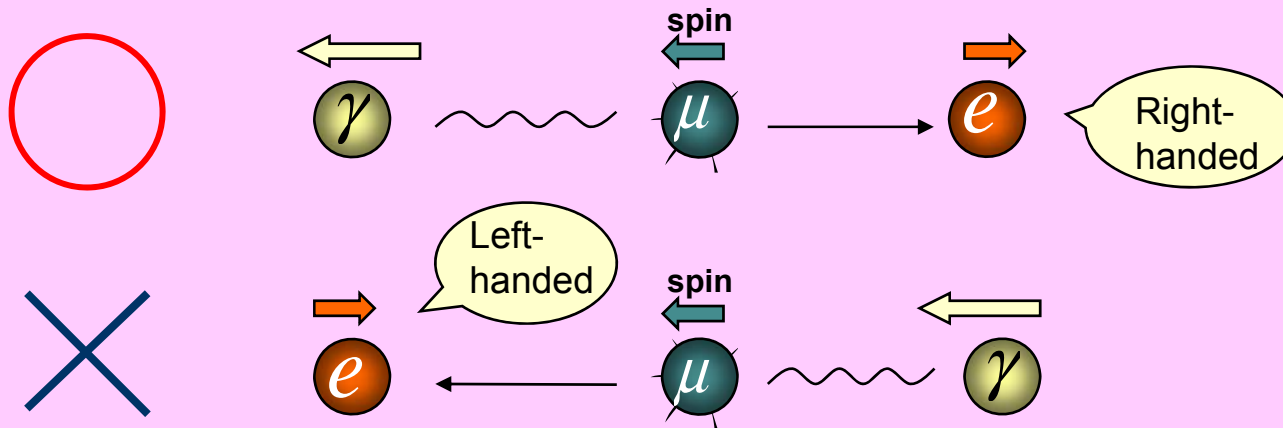
$$T = \epsilon^{\mu\nu\rho\sigma} \bar{u}_i(p) i\sigma_{\mu\nu} q^\rho (A_L^{ij} P_L + A_R^{ij} P_R) u_j(p')$$

Opposite from MSSM+ ν_R

- Intermediate state must be right-handed to pick up \tilde{m}_{eR}^2 the



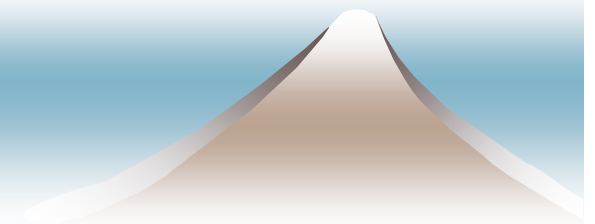
How can we see this feature experimentally?



It is possible to check this feature experimentally by measuring the angular distribution of final state lepton.

Predictions (Quark sector)

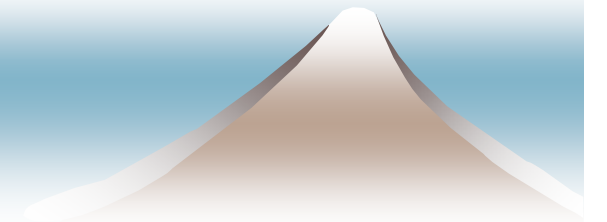
- ◆ The magnitudes are the same order as of the RGE effects in the universal mass case.
 $\tilde{m}_{\tilde{d}_L}^2 = \tilde{m}_{\tilde{u}_L}^2 = \tilde{m}_{\tilde{u}_R}^2 \sim \begin{pmatrix} m^2 & \Delta m^2 \lambda^5 & \Delta m^2 \lambda^3 \\ \Delta m^2 \lambda^5 & m^2 & \Delta m^2 \lambda^2 \\ \Delta m^3 \lambda^3 & \Delta m^2 \lambda^2 & m_3^2 \end{pmatrix}$
- ◆ **New CP phases!!**
The CP violation in B meson system may be detectable.



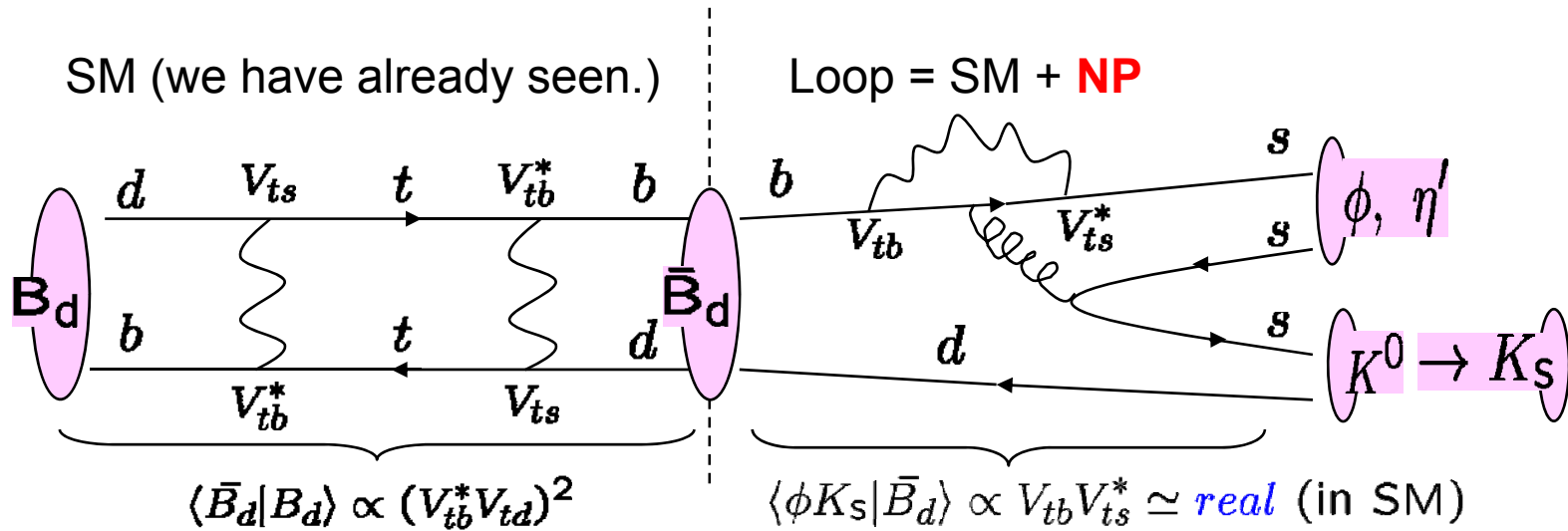
CP violation in B meson

	SM	E6	
$S_{B_s \rightarrow J/\psi \phi}$	O(0.04)	O(0.006)	$(\delta_{d_R})_{23} \sim 0$
$\Delta S_{K\phi}, \Delta S_{K\eta'}$		<0.15	
$A_{CP}(b \rightarrow s\gamma)$	0.006	<0.01	$(\delta_{d_L})_{23} \sim \lambda^2$
$A_{CP}(B \rightarrow V\gamma)$		very small	} $(\delta_{d_R})_{23} \sim 0$
$S_{B \rightarrow K_s \pi^0 \gamma}$		very small	

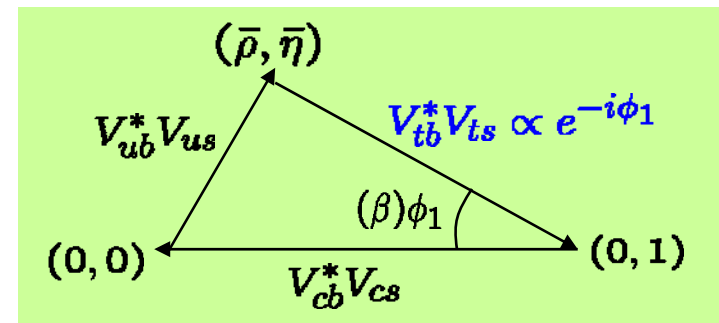
For $\tan \beta \sim 10$



$B_d \rightarrow \phi K_s, \eta' K_s$



$$S_{\phi K_s} = \frac{2\text{Im} \left[\sqrt{\frac{\langle \bar{B}_d | B_d \rangle^* \langle \phi K_s | \bar{B}_d \rangle}{\langle \bar{B}_d | B_d \rangle \langle \phi K_s | B_d \rangle}} \right]}{1 + \left| \frac{\langle \phi K_s | \bar{B}_d \rangle}{\langle \phi K_s | B_d \rangle} \right|^2} = \text{Im} \left[\frac{V_{tb} V_{td}^* \langle \phi K_s | \bar{B}_d \rangle}{V_{tb}^* V_{td} \langle \phi K_s | B_d \rangle} \right]$$



$$= \text{Im} \left[e^{2i\phi_1} \frac{\langle \phi K_s | \bar{B}_d \rangle}{\langle \phi K_s | B_d \rangle} \right] = \sin 2\phi_1 \text{Re} \left[\frac{\langle \phi K_s | \bar{B}_d \rangle}{\langle \phi K_s | B_d \rangle} \right] + \cos 2\phi_1 \text{Im} \left[\frac{\langle \phi K_s | \bar{B}_d \rangle}{\langle \phi K_s | B_d \rangle} \right]$$

$$\simeq \sin 2\phi_1 \simeq 0.68 \quad (\text{in SM})$$

Numerical Estimation

We estimate $S_{\phi K_s}$ by using parameters R, that is ratio of SUSY and SM amplitude.

$$\frac{\langle \phi K_s | \bar{B}_d \rangle}{\langle \phi K_s | B_d \rangle} \simeq \frac{A_{\phi K_s}^{SM} + e^{-i\theta} |A_{\phi K_s}^{SUSY}|}{A_{\phi K_s}^{SM} + e^{i\theta} |A_{\phi K_s}^{SUSY}|} = \frac{1 + e^{-i\theta} R_{\phi K_s}}{1 + e^{i\theta} R_{\phi K_s}} \quad \left(R_{\phi K_s} \equiv \frac{|A_{\phi K_s}^{SUSY}|}{|A_{\phi K_s}^{SM}|} \ll 1 \right)$$

$$\simeq 1 - i2(\sin \theta) R_{\phi K_s} \quad \left[\theta : \text{SUSY phase} \right]$$

$$S_{\phi K_s} = \sin 2\phi_1 \operatorname{Re} \left[\frac{\langle \phi K_s | \bar{B}_d \rangle}{\langle \phi K_s | B_d \rangle} \right] + \cos 2\phi_1 \operatorname{Im} \left[\frac{\langle \phi K_s | \bar{B}_d \rangle}{\langle \phi K_s | B_d \rangle} \right]$$

$$\simeq \sin 2\phi_1 - \cos 2\phi_1 \cdot 2 \sin \theta R_{\phi K_s} \simeq \underbrace{0.68}_{(\text{SM})} - \underbrace{1.46 R_{\phi K_s} \sin \theta}_{(\text{SUSY})}$$

$$S_{\eta' K_s} \simeq \underbrace{0.68}_{(\text{SM})} - \underbrace{1.46 R_{\eta' K_s} \sin \theta}_{(\text{SUSY})}$$

SUSY contribution to $S_{\phi K_s}$ is linearly depend on R.

$B_d \rightarrow \phi K_s, \eta' K_s$

$\Delta S_{\phi K_s}^{SUSY}, \Delta S_{\eta' K_s}^{SUSY} \sim O(0.1)$ is possible.

Glino contribution is decoupled.

Chargino contribution is not decoupled.

in the limit $m \gg m_3$

$O(0.1)$ deviation in B factory may be confirmed in SuperB factory.

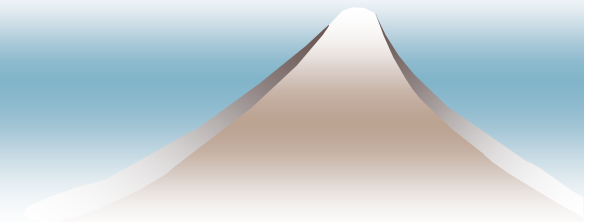
$$\sin(2\beta^{\text{eff}}) \equiv \sin(2\phi_1^{\text{eff}})$$

HFAG
ICHEP 2006
PRELIMINARY

$b \rightarrow ccs$	World Average		0.68 ± 0.03
ϕK^0	BaBar		$0.12 \pm 0.31 \pm 0.10$
	Belle		$0.50 \pm 0.21 \pm 0.06$
	Average		0.39 ± 0.18
$\eta' K^0$	BaBar		$0.55 \pm 0.11 \pm 0.02$
	Belle		$0.64 \pm 0.10 \pm 0.04$
	Average		0.59 ± 0.08

Summary table of E6 predictions

	SM	E6	
$\text{Br}(\mu \rightarrow e\gamma)$	~ 0	$10^{-11} - 10^{-14}$	⊙
$\text{Br}(\tau \rightarrow \mu\gamma)$	~ 0	$10^{-8} - 10^{-10}$	⊙
$S_{B_s \rightarrow J/\psi\phi}$	$O(0.04)$	$O(0.006)$	
$\Delta S_{K\phi}, \Delta S_{K\eta'}$		< 0.15	Δ
$A_{CP}(b \rightarrow s\gamma)$	0.006	< 0.01	
$A_{CP}(B \rightarrow V\gamma), S_{B \rightarrow K_s\pi^0\gamma}$		very small	



Discussions

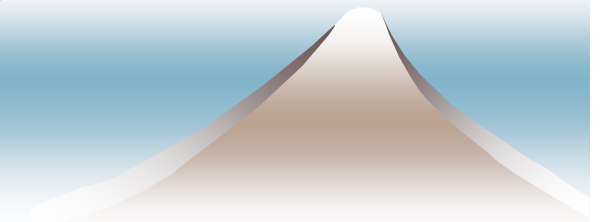
- ◆ Strictly speaking, $\delta_{\bar{5}} \neq 0$ $\delta_{LR} \neq 0$

when ~~$U(2)_H$~~ e.g. $\delta_{\bar{5}} \sim \begin{pmatrix} \lambda^5 & \lambda^4 & \lambda^{3.5} \\ \lambda^4 & \lambda^3 & \lambda^{2.5} \\ \lambda^{3.5} & \lambda^{2.5} & \lambda^2 \end{pmatrix}$

This can be consistent with the experiments,
but the predictions can be changed.

If we take $m_0 \gg m_3$, this model dependent
parts can be neglected.

No weak scale instability!!



SUSY CP problem

Spontaneous CP violation by the VEV of F .

The severe constraints from CEDM are also satisfied because of **real** up-Yukawa couplings.

SUSY CP problem

- ◆ EDM constraints from 1 loop

$$\mu = |\mu|e^{i\delta_\mu}, A = |A|e^{i\delta_A}$$
$$\delta_{\mu,A} < 10^{-2} \left(\frac{M_{SUSY}}{100\text{GeV}} \right)^2$$

- ◆ CEDM from Hg(neutron) even if $\delta_{\mu,A} = 0$

Hisano-Shimizu '04

$$\text{Im} (\delta_{d_L})_{13}(\delta_{d_R})_{31} < 3(0.2) \times 10^{-3}$$

$$\text{Im} (\delta_{u_L})_{23}(\delta_{u_R})_{32} < 3(0.2) \times 10^{-3} \quad \lambda^4 \sim 2 \times 10^{-3}$$

$$\text{Im} (\delta_{u_L})_{13}(\delta_{u_R})_{31} < 3(5) \times 10^{-5} \quad \lambda^6 \sim 10^{-4}$$

Contributions through stop loop are not decoupled.
Complex Yukawa couplings induce them generically.

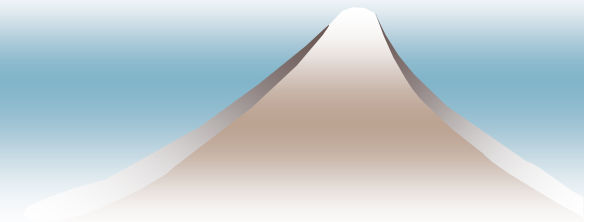
Difficulties in CEDM constraints

- ◆ We usually assume complex Yukawa couplings to induce KM phase.
- ◆ **Complex Yukawa couplings induce the complex $\delta_{u_L} \equiv U_{u_L} \tilde{m}_{u_L}^2 U_{u_L}^\dagger$ generically.**

Then CEDM constraints become severe.

$$\text{Im} (\delta_{u_L})_{23} (\delta_{u_R})_{32} < 3(0.2) \times 10^{-3}$$

$$\text{Im} (\delta_{u_L})_{13} (\delta_{u_R})_{31} < 3(5) \times 10^{-5}$$



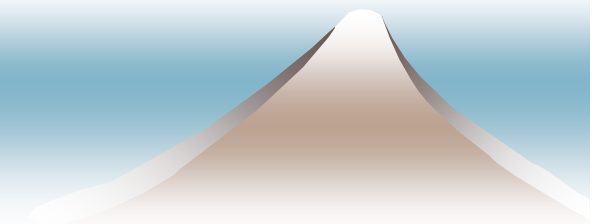
Spontaneous CP violation in $E6 + SU(2)_H(+U(1)_A)$

Ishiduki-Kim-N.M.-Sakurai

◆ Idea:

1, If the Higgs F which breaks horizontal symmetry has complex phase, we can obtain complex Yukawa couplings, and **KM phase**.

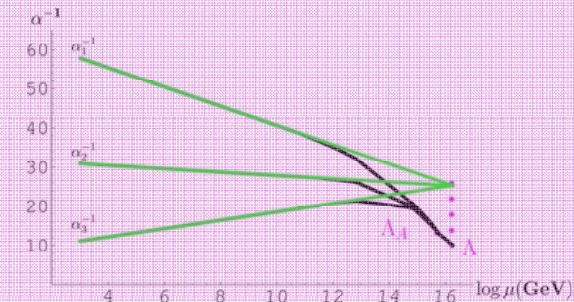
2, Naively the suppression $\langle \bar{F} F \rangle \sim \lambda^4$ can be expected for the phase of SUSY parameters.



Anomalous $U(1)_A$ GUT

N.M. 01, Bando-N.M. 01
N.M.-Yamashita 02

- ◆ Generic Interactions (Higher Dim. Op.) with $O(1)$ Coefficients
The symmetry define the theory
- ◆ **Doublet-triplet splitting**
- ◆ Suppressed Proton decay (dim. 5)
- ◆ Mass spectrum of Superheavy fields
- ◆ GUT scales and the other VEV scales
- ◆ **Natural gauge coupling unification (miraculous cancellation!)**
- ◆ Realistic quark&lepton masses and mixings (bi-large neutrino mixings)
- ◆ μ problem is solved



- ◆ To suppress the relative CP phase between μ and $B\mu$, non-trivial discrete charge must be imposed to F .

$$\langle \epsilon^{ab} F_a \rangle \Psi_b \sim \Psi_1$$

$$\langle \bar{F}^a \rangle \sim (0, v)$$

$$\langle \bar{F}^a \rangle \Psi_a \sim \Psi_2$$

$$\langle F_a \rangle \sim \begin{pmatrix} 0 \\ v e^{i\delta} \end{pmatrix}$$

A model with a discrete symmetry

(Z_{12})

Bonus 1

- ◆ Real up-type Yukawa couplings \Rightarrow real $\delta_{u_L}, \delta_{u_R}$
CEDM constraints can be satisfied.

- ◆ Complex down-type Yukawa couplings
KM phase can be induced.

$$\langle F_a \rangle \sim \begin{pmatrix} 0 \\ v e^{i\delta} \end{pmatrix}$$

- ◆ The point

$$Y^H (\cancel{F}, \bar{F}) : \text{real}, \quad H_u \sim 10_H \quad Y_u = Y^H$$

$$Y^C (F, \bar{F}) : \text{complex}$$

$$H_d \sim 10_H + 10_C + 16_C$$

$$W = Y^H 27_i 27_j \langle 27_H \rangle + Y^C 27_i 27_j \langle 27_C \rangle$$

A model with a discrete symmetry

- ◆ Bonus 2: **small up quark mass is realized.**

Usually, to obtain the CKM matrix $V_{CKM} \sim \begin{pmatrix} 1 & \lambda & \lambda^3 \\ \lambda & 1 & \lambda^2 \\ \lambda^3 & \lambda^2 & 1 \end{pmatrix}$

$$Y_u = \begin{pmatrix} \lambda^6 & \lambda^5 & \lambda^3 \\ \lambda^5 & \lambda^4 & \lambda^2 \\ \lambda^3 & \lambda^2 & 1 \end{pmatrix} \Rightarrow \begin{pmatrix} 0 & Q_{B-L}\lambda^5 & 0 \\ Q_{B-L}\lambda^5 & \lambda^4 & \lambda^2 \\ 0 & \lambda^2 & 1 \end{pmatrix}$$

$$\cancel{\epsilon^{ab}\psi_a\psi_b H}, \epsilon^{ab}\psi_a\langle A\rangle\psi_b H$$

$$\langle \epsilon^{ab} F_a \rangle \psi_b \sim \psi_1$$

$$\langle A \rangle \propto Q_{B-L}$$

$$y_u \sim \lambda^6$$

$$\Rightarrow \left(\frac{1}{3}\right)^2 \lambda^6$$

Too large \rightarrow good value!

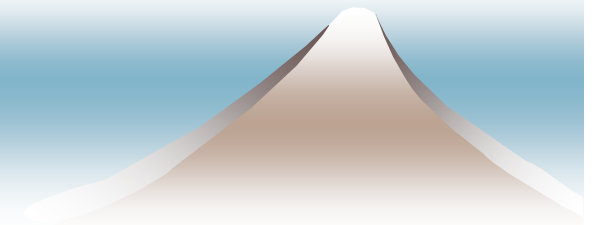
A model with a discrete symmetry

- ◆ Bonus 3?: # of $O(1)$ parameters = 10
13 physical parameters
 $\implies m_u, m_d, m_e, V_{CKM}$

- ◆ One of the relations

$$m_b = m_\tau(1 + O(\lambda))$$

$$\begin{pmatrix} m_s = O(1)m_\mu \\ m_d = O(1)m_e \end{pmatrix}$$



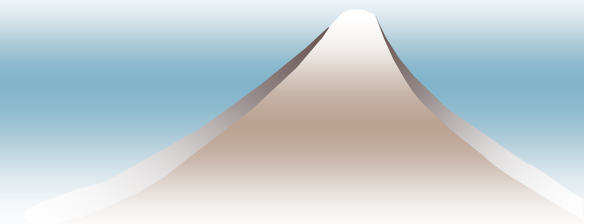
A model with a discrete symmetry

- ◆ The discrete symmetry is consistent with the E6 Higgs sector which realizes doublet-triplet splitting and

$$H_u \sim 10_H$$

$$H_d \sim 10_H + 10_C + 16_C$$

- ◆ 16_C mixing is required to avoid massless electron in this model.



Summary

- ◆ In E_6 GUT, one basic hierarchy for Yukawa couplings results in various hierarchical structures for quarks and leptons including larger neutrino mixings.
- ◆ Horizontal symmetry can easily reproduce the basic hierarchy, and suppress FCNC naturally in E_6 GUT. (not in SO(10) GUT)
- ◆ Spontaneous CP violation solves SUSY CP problem(CEDM)
- ◆ The simpler unification of quarks and leptons explains the more questions.

E_6 $3 \times 27 \implies$ larger neutrino mixings

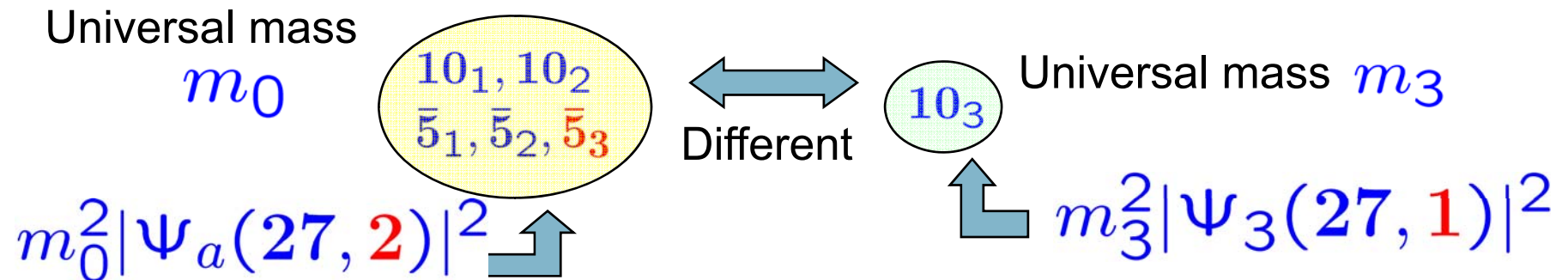
$$E_6 \times \begin{cases} U(2)_H & 2(27, 2 + 1) \\ U(3)_H & 1(27, 3) \end{cases}$$



SUSY Flavor Problem
SUSY CP Problem

Summary

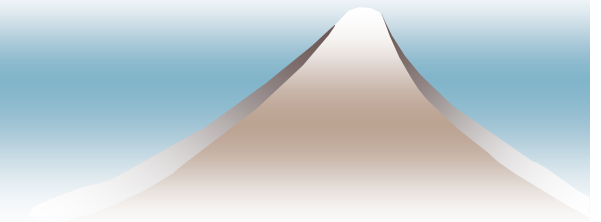
- ◆ Peculiar sfermion spectrum can be tested.



- ◆ $m_0 \gg m_3$ without unstability of weak scale. FCNC of 3rd generation becomes larger.
- ◆ LFV ($\mu \rightarrow e_R \gamma, \tau \rightarrow \mu_R \gamma$) and CP violation in B ($B \rightarrow \phi K_s, \eta' K_s$ etc) may be detectable in future.
- ◆ Polarization of final lepton can test the GUT scenario.

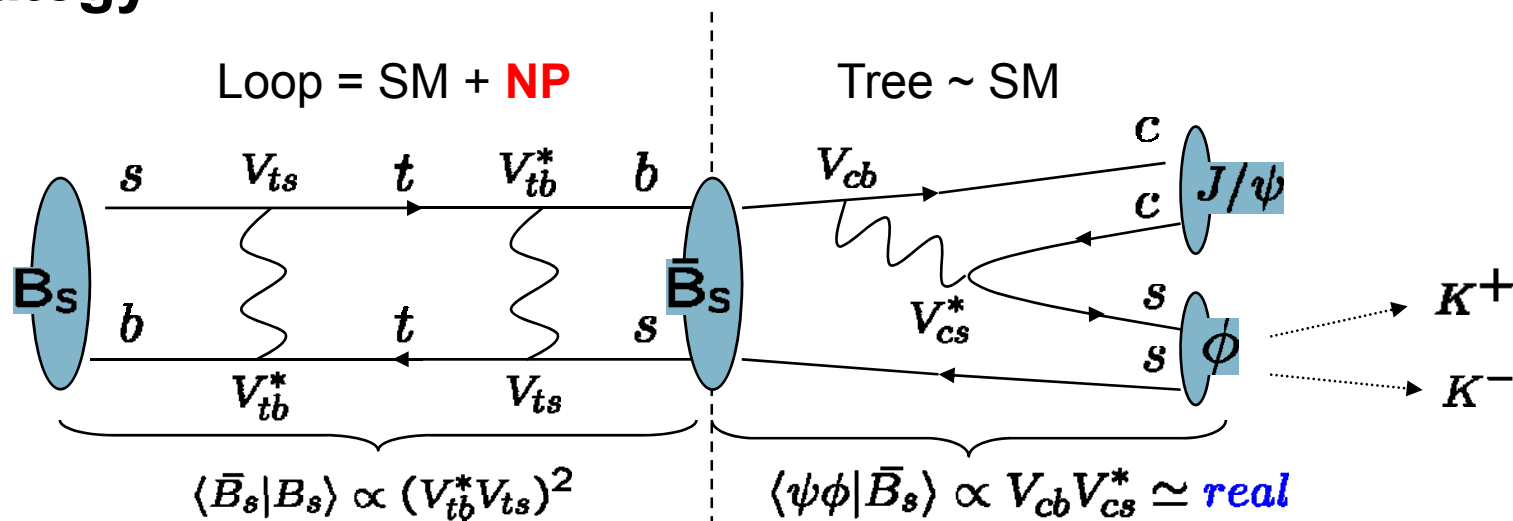
Future work

- ◆ What is the signature in LHC?
- ◆ CP violation in neutrino oscillation?
- ◆ The origin of SUSY breaking? The mediation mechanism of SUSY breaking?
- ◆ Can the D term contribution be sufficiently small?
 $D_{SU(2)_H}, D_{E_6}$
Decoupling feature mildens the constraints
Non-Abelian discrete symmetry?
- ◆ Cosmology?
- ◆



$B_s \rightarrow J/\psi\phi$

◆ Strategy



Time dependent CP asym.

$$\frac{\Gamma[B_S(t) \rightarrow \psi\phi] - \Gamma[\bar{B}_S(t) \rightarrow \psi\phi]}{\Gamma[B_S(t) \rightarrow \psi\phi] + \Gamma[\bar{B}_S(t) \rightarrow \psi\phi]}$$

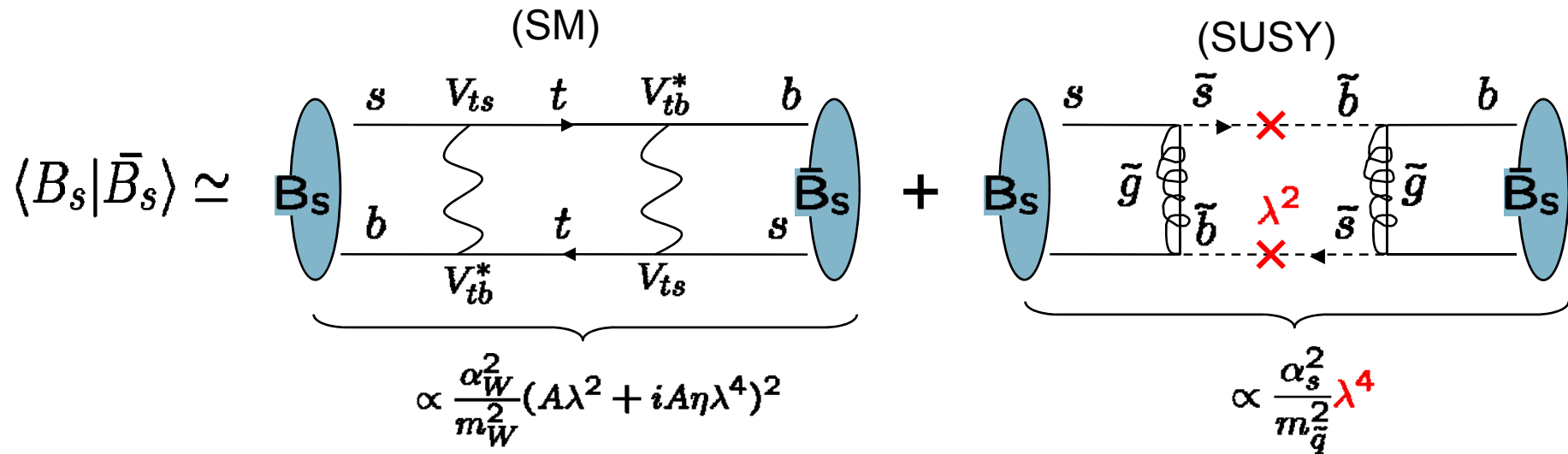
$$V_{tb} = 1 + \frac{1}{2}A^2\lambda^4 + \mathcal{O}(\lambda^6)$$

$$V_{ts} = -A\lambda^2 - i\eta A\lambda^4 + \dots$$

$$S_{\psi\phi} = \frac{2\text{Im}\left[\sqrt{\frac{\langle \bar{B}_s | B_s \rangle^* \langle \psi\phi | \bar{B}_s \rangle}{\langle \bar{B}_s | B_s \rangle \langle \psi\phi | B_s \rangle}}\right]}{1 + \left|\frac{\langle \psi\phi | \bar{B}_s \rangle}{\langle \psi\phi | B_s \rangle}\right|^2} \simeq \text{Im}\left[\frac{\langle B_s | \bar{B}_s \rangle^*}{\langle B_s | \bar{B}_s \rangle}\right] \simeq \text{Im}\left[\frac{V_{tb} V_{ts}^*}{V_{tb}^* V_{ts}}\right] \simeq -\eta\lambda^2 = \mathcal{O}(10^{-2}) \ll 1$$

In SM, $S_{\psi\phi}$ is very small ! NP contribution may be dominant !

Naïve Estimation



If λ^4 is pure imaginary...

$$S_{\psi\phi} \simeq \text{Im} \left[\frac{\langle \bar{B}_s | B_s \rangle^*}{\langle \bar{B}_s | B_s \rangle} \right] \sim \frac{\frac{\alpha_s^2}{m_{\tilde{q}}^2} \lambda^4}{\frac{\alpha_W^2}{m_W^2} A^2 \lambda^4} \simeq \frac{\alpha_s^2}{\alpha_W^2} \frac{m_W^2}{m_{\tilde{q}}^2} \sim \mathbf{0.4 !!}$$

$$\left(\begin{array}{l} \alpha_s \simeq 0.11 \\ \alpha_W \simeq 0.034 \\ m_{\tilde{q}} \sim 400 \text{ GeV} \end{array} \right)$$

Naively, it is expected that SUSY contribution make $S_{\psi\phi}$ 10 times larger than SM prediction !!

Numerical Estimation

$$S_{\psi\phi} \simeq \mathcal{O}(0.04) \quad \left. \vphantom{S_{\psi\phi}} \right\} \text{(SM)} \quad (m_{\tilde{q}} = 380\text{GeV}, m_{\tilde{g}} = 400\text{GeV}) \quad \text{Khalil et al. '03}$$

$$\left. \begin{aligned} &+ 2.65 [(\delta_{23}^d)_{LL}^2 + (\delta_{23}^d)_{RR}^2] + 57.3 [(\delta_{23}^d)_{LR} + (\delta_{23}^d)_{RL}] \\ &- 90.3 [(\delta_{23}^d)_{LR}(\delta_{23}^d)_{RL}] - 374 [(\delta_{23}^d)_{LL}(\delta_{23}^d)_{RR}] \end{aligned} \right\} \text{(SUSY)}$$

$$(\delta_{\alpha\beta}^q)_{AB} \equiv \frac{(m_{\tilde{q}}^2)_{\alpha\beta AB}}{m_{\tilde{q}}^2} \quad (\alpha, \beta = 1, 2, 3) \\ (A, B = L, R)$$

In E_6 + Horizontal model

$$(\delta_{23}^d)_{LL} \simeq \lambda^2 \quad (\delta_{23}^d)_{LR}, (\delta_{23}^d)_{RR} \simeq 0$$

$$(\bar{q}_\alpha)_A \xrightarrow{(\delta_{\alpha\beta}^q)_{AB}} (\tilde{q}_\beta)_B$$

(SM)

(SUSY)

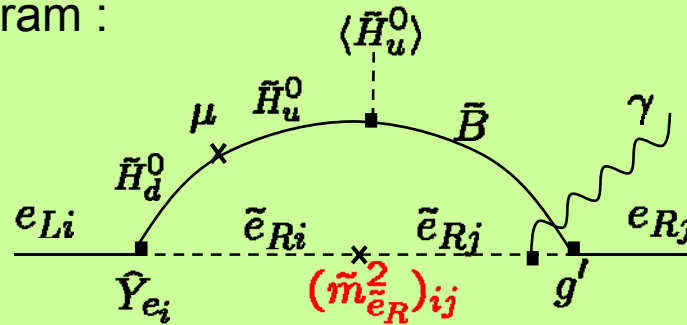
$$S_{\psi\phi} \simeq \mathcal{O}(0.04) + \mathcal{O}(0.006)$$

Actually, SUSY contribution is too tiny with compared to SM prediction.
It is difficult to observe it.

$$C_{SM} = \frac{\alpha_W^2}{m_W^2} (V_{tb}^* V_{ts})^2 \overbrace{S_0 \left(\frac{m_t^2}{m_W^2} \right)}^{\mathcal{O}(1)} \quad C_{SUSY} = -\frac{\alpha_s^2}{216 m_{\tilde{q}}^2} (\delta_{23}^d)_{LL}^2 \overbrace{\left[24 \left(\frac{m_{\tilde{g}}}{m_{\tilde{q}}} \right)^2 f_6 \left(\frac{m_{\tilde{g}}}{m_{\tilde{q}}} \right) + 66 \tilde{f}_6 \left(\frac{m_{\tilde{g}}}{m_{\tilde{q}}} \right) \right]}^{\mathcal{O}(1)}$$

Rough Estimation of the BRs

main diagram :



$$Br(l_i \rightarrow l_j \gamma) \sim \frac{1}{G_F^2} e(g')^2 \frac{\tan \beta (m_{\tilde{e}_R}^2)_{ij}}{m_{SUSY}^4} |^2 Br(l_i \rightarrow e \nu \bar{\nu})$$

$$\tilde{m}_{\tilde{e}_R}^2 \sim \Delta m^2 \begin{pmatrix} 1 & \lambda^5 & \lambda^3 \\ \lambda^5 & 1 & \lambda^2 \\ \lambda^3 & \lambda^2 & \mathcal{O}(1) \end{pmatrix}$$

$$(m_{SUSY} = 200 \text{ GeV}, \quad \tan \beta = 5)$$

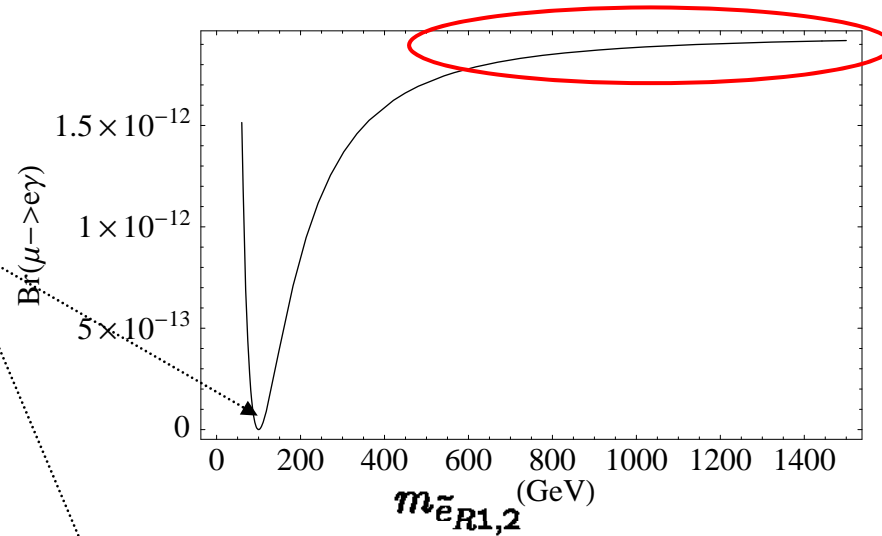
$$Br(\mu \rightarrow e \gamma) \sim 10^{-11} \quad Br(\tau \rightarrow \mu \gamma) \sim 10^{-8}$$

$$(Br(\mu \rightarrow e \gamma)_{\text{exp.}} < 1.2 \times 10^{-11}) \quad (Br(\tau \rightarrow \mu \gamma)_{\text{exp.}} < 4.8 \times 10^{-8})$$

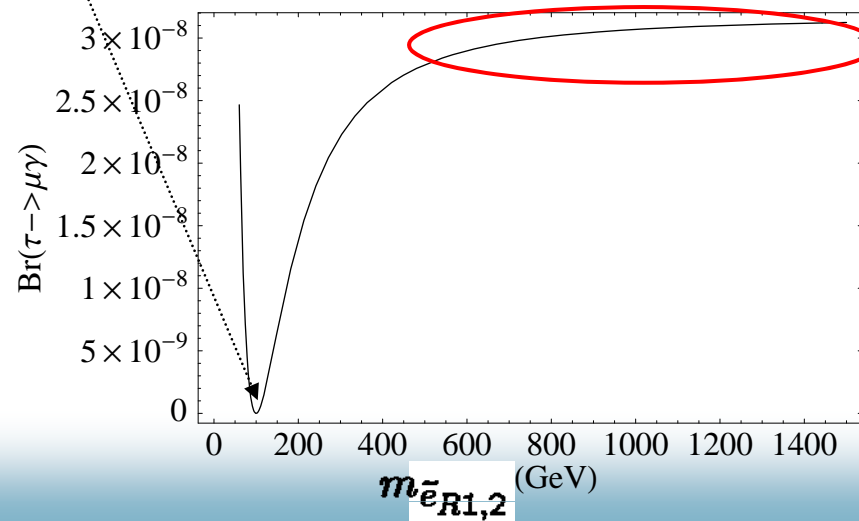
This model leads large LFV rate within reach of near future experiments.

$m_{\tilde{e}_{R1,2}}$ dependence

At this point, $m=m_3$
 $(\tilde{m}_{\tilde{e}_R}^2)_{ij} \propto \Delta m^2$
 $= (m_3^2 - m^2) = 0$.
 There are no source of LFV.



In this reason, the BRs are no longer changed for $m_{\tilde{e}_{R1,2}}$.
 BRs don't decouple as $m_{\tilde{e}_{R1,2}}$ becomes large.
 These behavior are explained qualitatively already.



$\tan \beta = 5$
 $M_2 = 120 \text{ GeV}$
 $\mu = 200 \text{ GeV}$
 $m_{\tilde{e}_{R3}} = 100 \text{ GeV}$

Branching Ratio is independent of $m_{\tilde{e}_{R1,2}}$ unless $m_{\tilde{e}_{R1,2}} < 600 \text{ GeV}$.

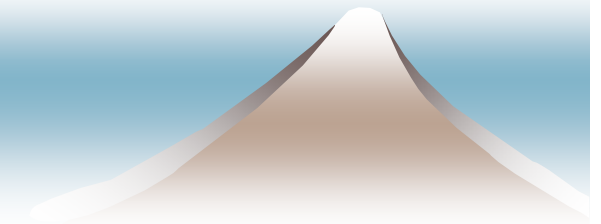
Quark sector

- ◆ Constraints from $B_s - \bar{B}_s$ mixing is OK.

$$\begin{aligned} (\delta_{d_L})_{23} &< 0.2 \\ \sqrt{(\delta_{d_L})_{23}(\delta_{d_R})_{23}} &< 0.02 \end{aligned} \quad \text{Khalil06}$$

- ◆ EDM of Hg(neutron) Hisano-Shimizu

$$\begin{aligned} \text{Im } (\delta_{d_L})_{13}(\delta_{d_R})_{31} &< 3(0.2) \times 10^{-3} && \text{OK} \\ \text{Im } (\delta_{u_L})_{23}(\delta_{u_R})_{32} &< 3(0.2) \times 10^{-3} && \lambda^4 \sim 2 \times 10^{-3} \\ \text{Im } (\delta_{u_L})_{13}(\delta_{u_R})_{31} &< 3(5) \times 10^{-5} && \lambda^6 \sim 10^{-4} \end{aligned}$$



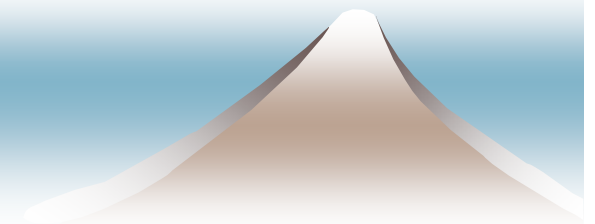
A structure in E_6 explains

- ◆ the suppression of the radiative FCNC with unifying Yukawa $E_6 \supset SO(10)$
- ◆ why the mixings in lepton sector is larger.
- ◆ the various mass hierarchies in quark & lepton sector.

$$\frac{m_{ui}}{m_{uj}} \gg \frac{m_{di}}{m_{dj}}, \frac{m_{ei}}{m_{ej}} \gg \frac{m_{\nu i}}{m_{\nu j}} \quad i < j$$

E_6 +horizontal symmetry

- ◆ A prediction for sfermion mass spectrum
Universality of sfermion masses are partly realized.
- ◆ FCNC can be suppressed without special assumption for SUSY breaking sector.
- ◆ Non universality makes the 3rd generation FCNC larger than in the universal case.



Radiatively induced FCNC

$$10 = (q, u_R^c, e_R^c)$$

$$\bar{5} = (d_R^c, l)$$

$$1 = \nu_R^c$$

$$W_Y = (Y_u)_{ij} 10_i 10_j \bar{5}_H + (Y_{(d,e)})_{ij} 10_i \bar{5}_j \bar{5}_{\bar{H}} \\ + (Y_{\nu_D})_{ij} \bar{5}_i 1_j 5_H + (M_R)_{ij} 1_i 1_j$$

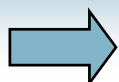
- ◆ Even if universal sfermion masses at the cutoff, radiative correction induces the non-universality. Borzumati-Masiero 85
Hisano-Moroi-Tobe-Yamaguchi-Yanagida Barbieri-Hall-Strumia 95

$$\text{Br}(l_i \rightarrow l_j \gamma) \propto (V_L^\dagger \hat{Y}_{\nu_D}^2 V_L \tan \beta)^2$$

- ◆ Large mixings $V_L \sim V_{MNS}$
- ◆ SO(10) GUT relations

$$Y_{\nu_D} \sim Y_u \sim Y_d \iff$$

$$\tan \beta, Y_{\nu_D}$$



too large LFV

Key Observation to understand the different mixings

- ◆ SU(5) GUT Albright, Barr, Sato, Yanagida, Ramond, ...,

$$10 = (q, u_R^c, e_R^c) \quad \bar{5} = (d_R^c, l)$$

Quark mixings(CKM) \leftrightarrow Mixings of $10_i(q_i)$

Lepton mixings(MNS) \leftrightarrow Mixings of $\bar{5}_i(l)$

\rightarrow $10_i(q_i)$ have stronger hierarchy than $\bar{5}_i(l)$

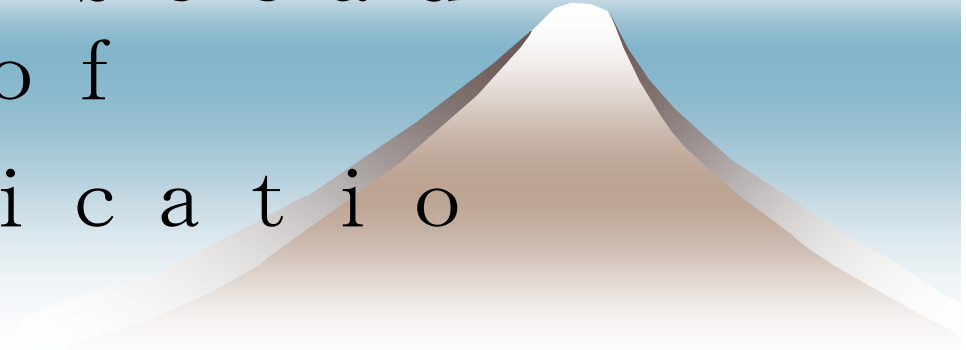
\rightarrow Y_u have stronger hierarchy than Y_d

$$Y_u 10_i 10_j 5_H + Y_d 10_i \bar{5}_j \bar{5}_{\bar{H}}$$

Next question

Why do $\bar{5}_i(l)$ have milder hierarchy than $10_i(q_i)$?

This is because
 E_6 sees of
Unification
n



Unification of Yukawa hierarchies

Bando-N.M.
N.M.

All Yukawa couplings have the same hierarchical flavor structure.

→ Various hierarchies of quarks and leptons

Only one basic hierarchical flavor structure
(to realize Y_u) Not fix the origin.

An example

$$Y^H \sim Y^C \sim Y \sim \begin{pmatrix} \lambda^6 & \lambda^5 & \lambda^3 \\ \lambda^5 & \lambda^4 & \lambda^2 \\ \lambda^3 & \lambda^2 & 1 \end{pmatrix} \quad \left[\begin{pmatrix} 0 & \lambda^5 & 0 \\ \lambda^5 & \lambda^4 & \lambda^2 \\ 0 & \lambda^2 & 1 \end{pmatrix} \right]$$

$\lambda \sim 0.22$

How to fix light modes of $\bar{5}$?

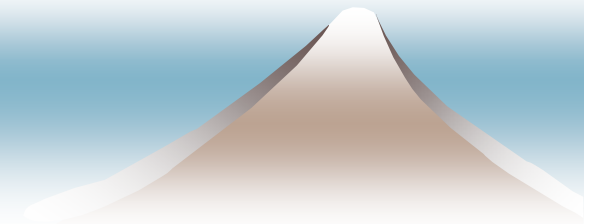
$$27_i = 16_i[10_i + \bar{5}_i + 1_i] + 10_i[5_i + \bar{5}_i] + 1_i[1_i]$$

$$M(\mathbf{5}, \bar{\mathbf{5}}) / \langle \mathbf{1}_H \rangle \sim$$

	$\bar{5}_1$	$\bar{5}_2$	$\bar{5}_3$	$\bar{5}_1$	$\bar{5}_2$	$\bar{5}_3$
5_1	λ^6	λ^5	λ^3	λ^{6+r}	λ^{5+r}	λ^{3+r}
5_2	λ^5	λ^4	λ^2	λ^{5+r}	λ^{4+r}	λ^{2+r}
5_3	λ^3	λ^2	1	λ^{3+r}	λ^{2+r}	λ^r

$$\lambda^r \equiv \frac{\langle 27_C \rangle}{\langle 27_H \rangle} \sim \lambda^{0.5}$$

Light modes $(\bar{5}_1, \bar{5}_1, \bar{5}_2)$



SO(10) GUT relations

$$Y_d = Y_e^T = Y_u = Y_{\nu D}$$

$$Y_u \sim \begin{matrix} 10_1 & 10_2 & 10_3 \\ 10_1 \left(\begin{matrix} \lambda^6 & \lambda^5 & \lambda^3 \\ \lambda^5 & \lambda^4 & \lambda^2 \\ \lambda^3 & \lambda^2 & 1 \end{matrix} \right) \\ 10_2 \\ 10_3 \end{matrix} \left. \vphantom{\begin{matrix} 10_1 \\ 10_2 \\ 10_3 \end{matrix}} \right\} \begin{matrix} \lambda \\ \lambda^2 \end{matrix} \quad Y_{\nu D} \sim \begin{matrix} 1_1 & 1_2 & 1_3 \\ \bar{5}_1 \left(\begin{matrix} \lambda^6 & \lambda^5 & \lambda^3 \\ \lambda^{5.5} & \lambda^{4.5} & \lambda^{2.5} \\ \lambda^5 & \lambda^4 & \lambda^2 \end{matrix} \right) \\ \bar{5}_1 \\ \bar{5}_2 \end{matrix} \left. \vphantom{\begin{matrix} \bar{5}_1 \\ \bar{5}_1 \\ \bar{5}_2 \end{matrix}} \right\} \begin{matrix} \lambda^{0.5} \\ \lambda^{0.5} \end{matrix}$$

$$(\bar{5}_1, \bar{5}_1, \bar{5}_2)$$

$$\bar{5}_1 + \lambda^\Delta \bar{5}_3 \quad (\Delta = 3 - r)$$

$$Y_d \sim Y_e^T \sim \begin{matrix} \bar{5}_1 & \bar{5}_1 & \bar{5}_3 \\ 10_1 \left(\begin{matrix} \lambda^6 & \lambda^{5.5} & \lambda^3 \\ \lambda^5 & \lambda^{4.5} & \lambda^2 \\ \lambda^3 & \lambda^{2.5} & 1 \end{matrix} \right) \\ 10_2 \\ 10_3 \end{matrix} \left. \vphantom{\begin{matrix} \bar{5}_1 \\ \bar{5}_1 \\ \bar{5}_3 \end{matrix}} \right\} \begin{matrix} \lambda \\ \lambda^2 \end{matrix}$$

$\underbrace{\lambda^6 \quad \lambda^{5.5}}_{\lambda^{0.5}} \quad \underbrace{\lambda^3 \quad \lambda^2}_{\lambda^{0.5}}$

$$\lambda^r \equiv \frac{\langle 27_C \rangle}{\langle 27_\Phi \rangle} \sim \lambda^{0.5}$$

Small $\tan \beta$

Small $Y_{\nu D}$

$$V_{CKM} \sim \begin{pmatrix} 1 & \lambda & \lambda^3 \\ \lambda & 1 & \lambda^2 \\ \lambda^3 & \lambda^2 & 1 \end{pmatrix}$$

$$V_{MNS} \sim \begin{pmatrix} 1 & \lambda^{0.5} & \lambda \\ \lambda^{0.5} & 1 & \lambda^{0.5} \\ \lambda & \lambda^{0.5} & 1 \end{pmatrix}$$

E_6 GUT is
an interesting target of B factory



Weak scale stability requirement

◆ $m_3, M_{\frac{1}{2}}, m_H, \mu \sim O(100\text{GeV})$

◆ Not strong constraint

for m_0 because of small Yukawa couplings

◆ If we take $m_0 \gg m_3$, constraints from FCNC, EDM, g-2 etc. become much weaker.

◆ The little hierarchy problem?

Natural parameters \Rightarrow ~~$m_h > 114.4 \text{ GeV}$~~

\Rightarrow Kim's talk

