

Muon Colliders R. B. Palmer (BNL) June 16, 2012 Crete

For the Muon Accelerator Program (MAP)

- A National Program
- Administered through Fermi Lab
- MAP Director Mark Palmer (no relation of mine)
- \bullet Funded now at \approx 10 M\$ per year
 - Why
 - Concept
 - Comparison with CLIC
 - Conclusion

Why a Muon Collider ?

- Electron Linear Colliders
 - synchrotron radiation ($\propto \gamma^4$) forces Linear Colliders to be linear
 - electrons intersect once and are thrown away
 - beamstrahlung causes huge energy variation (70% of Luminosity has dE > 1% at 3 TeV)
- Muon Collider
 - Acceleration can be in rings, using much less rf
 - Collisions can be in rings ≈ 1000 collisions before decay allowing larger emittances and spot sizes and requiring less beam power
 - $-\operatorname{Beamstrahlung}$ now negligible dE/E \approx 0.1 %



- Muon Colliders certainly smaller, and use less power ?
- They may be cheaper
- Main challenge is emittance reduction (cooling)

Schematic





- 2 10¹⁴ proton with σ_t =2 nsec at 15 Hz
- 8 GeV Linac, Accumulator & Buncher: \geq 4 bunches \leq 5 10 13
- Kicker and Trombone (Ankenbrandt)
- Intersecting liquid metal target, in time, from multiple directions

20 T Capture Solenoid



- Copper coil gives 6 T, (uses 15 MW of wall power)
- 14 T Super-conducting solenoid, tapering to 3 T
- Tungsten Carbide in water shielding

Liquid metal (eg mercury) jet target MERIT Experiment at CERN



• 15 T pulsed magnet

- 1 cm rad mercury jet
- Up to 30 Tp
- Splash velocities were moderate
- Density persists for 100 micro sec

Images of Jet Flow at Viewport 3, B=10T, N=10Tp, L=17cm, 2ms/frame



t = 6 ms

t = 8 ms



t = 10 ms

t = 14 ms



- \approx 70% efficiency into 12 bunches
- \bullet rms dE/E from 100 % to $\approx 15\%$

Conventional cooling methods

Synchrotron radiation cooling of electrons Proton/ion cooling by co-moving electron beam Stochastic cooling of protons/ions Laser cooling of ions

negligible radiation for μ s too slow cf muon decay too slow cf muon decay too slow cf muon decay

Only known way to cool muons is by ionization energy loss



Emittances in Cooling Sequence ICOOL Simulations of 6D cooling are for Guggenheim lattices



Candidates for 6D cooling lattices

Initial I0 m B=3 T f=201 MHz Later 2.5 m B=18 T f=805 MHz





- Guggenheim and HCC have similar simulated performance
- A third system that cools both signs: 'The Snake' (not shown) does not cool to low emittances & would only be used at start

Muon Ionization Cooling Experiment (MICE) International collaboration at RAL, US, UK, Japan (Blondel)

- Early Experiment to demonstrate Emittance Exchange
 - Cooling in all dimensions
 But no re-acceleration
- Will then demonstrate transverse cooling in liquid hydrogen, including rf re-acceleration



• Experiment should run in two years time

FNAL Exp's on High Pressure Gas rfwith Muons IncAs required for HCC 6D cooling

• rf works well with/without magnetic field





- rf tested in proton beam
 - No breakdown with magnetic field and/or beam
 - Beam loading, with 0.2% O_2 acceptable
- Problems remain in fitting rf inside HCC coils

FNAL R&D on vacuum rf with mag fields As required for Guggenheim 6D cooling

- Observed damage & reduced gradients with fields
- But recent tests with Be buttons show
 - Evidence that Be better resists damage in magnetic fields
 - Be walled cavity now under construction

Copper button after 28 MV/m & 3 T

Beryllium button after 33 MV/m & 3 T





Final Transverse Cooling



- Cooling in hydrogen simulated for all 13 stages
- Matching and re-acceleration still only simulated last stages
- Consequences of a limitation to 30 T probably acceptable but we believe that 40 T is attainable and leave as baseline

BNL R&D on HTS magnets

with PBL As required for final cooling





- \bullet When tested together we expect 25 T
- If tested in NHMFL 20 T, should demonstrate 40 T

Acceleration

Must be fast:

Linacs, recirculating linacs (RLA) and pulsed synchrotrons (RCS)



- Appears straight foreword
- Impedance questions not yet studied



$$\begin{split} R_B &= 4.4 \, 10^{-24} \, \frac{N_\mu \ f \ E^3 \ t \ < B >}{D \ B} \quad \text{Sv} \quad \text{from regions of uniform B} \\ R_L &= 6.7 \, 10^{-24} \, \frac{N_\mu \ f \ E^3 \ t \ < B > \ L}{D} \quad \text{Sv} \quad \text{from straight sections} \end{split}$$

For $R_B = R_L = 10\%$ Fed limit = 0.1 mSv (10 mRad)

E	B(min)	L(max)
TeV	Т	m
1.5	0.25	2.4
3.0	1.5	0.28

These appear hard, but not impossible

MC Rings

C of m Energy	1.5	3	6	TeV
Luminosity	1	4	12	$10^{34} \text{ cm}^2 \text{sec}^{-1}$
Muons/bunch	2	2	2	10^{12}
Total muon Power	7.2	11.5	11.5	MW
Ring <bending field=""></bending>	6.04	8.4	11.6^{1}	Т
Ring circumference	2.6	4.5	6	km
eta^* at $IP=\sigma_z$	10	5	2.5 ¹	mm
rms momentum spread	0.1	0.1	0.1	%
Depth	135	135	540 ²	m
Repetition Rate	15	12	6	Hz
Proton Driver power	4	3.2	1.6	MW
Muon Trans Emittance	25	25	25	μ m
Muon Long Emittance	72	72	72	mm

Note 1: This is a blind extrapolation from 1.5 and 3 TeV designs Note 2: For the same neutrino radiation Muon source the same for all energies \rightarrow natural upgrades

Detector Shielding



Wall Power Requirement for 1.5 TeV

From summer 2011 PRELIMINARY and approximate

	Len	Static	Dynamic				Tot
		4 ⁰	rf	PS	4 ⁰	20 ⁰	
	m	MW	MW	MW	MW	MW	MW
p Driver (SC linac)							(20)
Target and taper	16			15.0	0.4		15.4
Decay and phase rot	95	0.1	0.8		4.5		5.4
Charge separation	14						
6D cooling before merge	222	0.6	7.2		6.8	6.1	20.7
Merge	115	0.2	1.4				1.6
6D cooling after merge	428	0.7	2.8			2.6	6.1
Final 4D cooling	78	0.1	1.5			0.1	1.7
NC RF acceleration	104	0.1	4.1				4.2
SC RF linac	140	0.1	3.4				3.5
SC RF RLAs	10400	9.1	19.5				28.6
SC RF RCSs	12566	11.3	11.8				23.1
Collider ring	2600	2.3		3.0	10		15.3
Totals	26777	24.6	52.5	18.0	21.7	8.8	145.6

 \approx 160 MW for 3 TeV $~~\approx$ 200 MW for 6 TeV

Compare 3 TeV $\mu^+\mu^-$ with e^+e^- CLIC

		$\mu^+\mu^-$	e^+e^-
Luminosity	$10^{34} \text{ cm}^2 \text{sec}^{-1}$	4	2
Detectors		2	1
eta^* at IP = σ_z	mm	5	0.09
rms bunch height σ_y	μ m	4	0.001
Total lepton Power	MW	11.5	28
Comparable Wall power	MW	pprox 160	450 (570 tot)

- $\mu^+\mu^-$ luminosity twice CLIC's (for dE/E < 1%) & 2 detectors
- Spot sizes and tolerances much easier than CLIC's
- \bullet Lepton and Wall power \approx 1/3 CLIC's
- \bullet Because muons interact \approx 1000 times, but electrons only once
- But Muon Collider less developed

CONCLUSION I

- Much simulation progress this year
 - new capture magnet design, chikane, new merge designs, Nonflip cooling lattices, lower final emittances, detector background studies, a start on space charge in cooling
- Progress in needed technologies
 - $-\ln$ HP Gas cavity in a beam
 - In rf-in-magnetic fields using Beryllium
 - In High Temp Super-Conductor YBCO coils
- Favorable comparisons with CLIC:
 - Luminosity greater than CLIC's
 - $-\operatorname{Estimated}$ wall power $\approx 1/3$ of CLIC
- Extrapolation to higher energies thinkable

PERSONAL CONCLUSION II

We have long argued that a detailed study of 'New Physics' such as Super Symmetry requires a lepton collider with appropriate energy

- \bullet If 'New Physics' < 1 TeV Go for ILC
- \bullet If 'New Physics' < 2 TeV \qquad Go for CLIC
- \bullet But if 'New Physics' $> 2 \mbox{ TeV}$ then Muon Collider the only way
- Note: Plasma acceleration does not solve the energy problem

