

Present and future strategies for Neutrinoless Double Beta Decay

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Open questions in ν physics

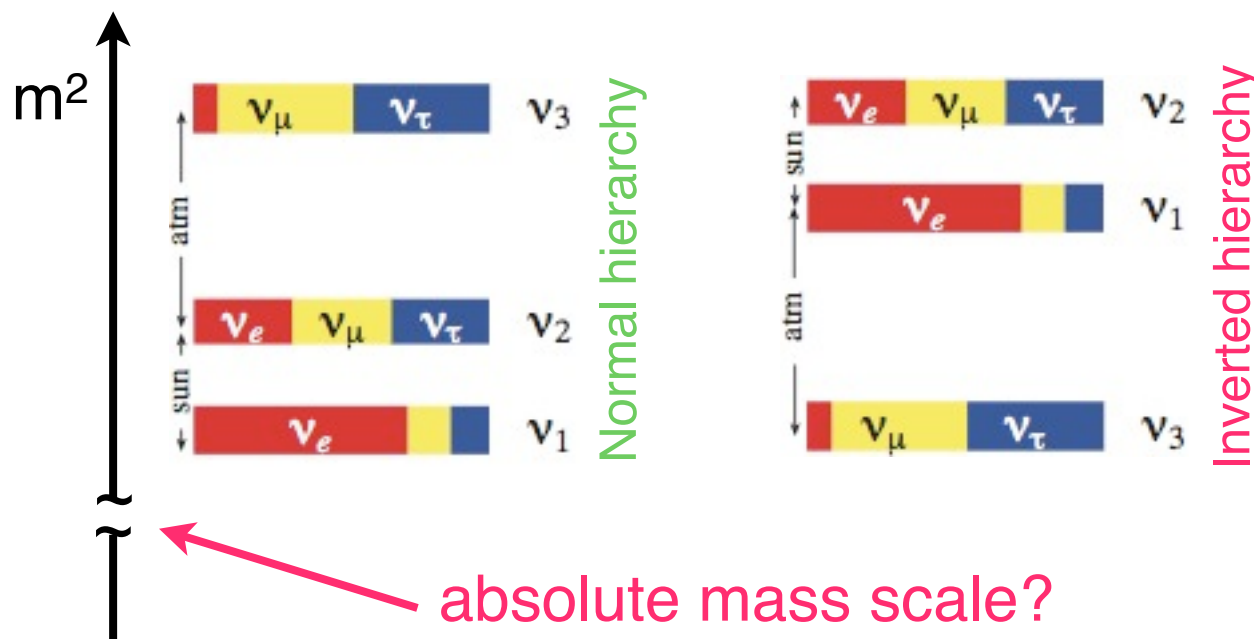
- ~~What is the size of the θ_{13} mixing angle?~~
 - ▶ Is there CP violation?
- Are there sterile neutrinos?
- Is the total Lepton Number conserved?
 - ▶ Is the neutrino a Majorana particle?
- What is value of the neutrino mass?
- What is the hierarchy of neutrino masses?

Oscillation Results

- Almost all oscillation parameters have been measured ([arXiv:1205.5254](https://arxiv.org/abs/1205.5254))

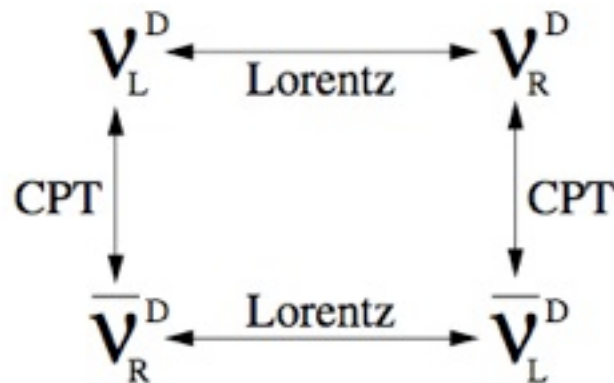
Oscillation Parameter		Value	Precision
solar mass splitting	Δm_{21}^2	$= 7.54^{+0.26}_{-0.22} \times 10^{-5} \text{ eV}^2$	3%
atmospheric mass splitting	$ \Delta m_{23}^2 $	$= 2.43^{+0.07}_{-0.09} \times 10^{-3} \text{ eV}^2$	3%
solar mixing angle	$\sin^2 \theta_{12}$	$= 0.307^{+0.018}_{-0.016}$	6%
atmospheric mixing angle	$\sin^2 \theta_{23}$	$= 0.398^{+0.030}_{-0.026}$	7%
'CHOOZ' mixing angle	$\sin^2 \theta_{13}$	$= 0.025 \pm 0.003$	12%

- The sign of Δm_{23}^2 is still unknown, thus allowing two configurations:

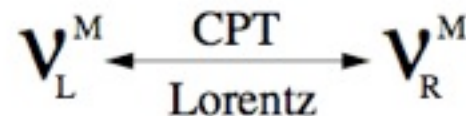


Neutrino nature

- Except for the total leptonic number the neutrino is a neutral fermion.
- So if the total leptonic number is not conserved neutrinos can be Majorana particles:
 - ▶ particle and antiparticle are the same. Chirality determines the charge of the lepton produced in interactions:



Dirac

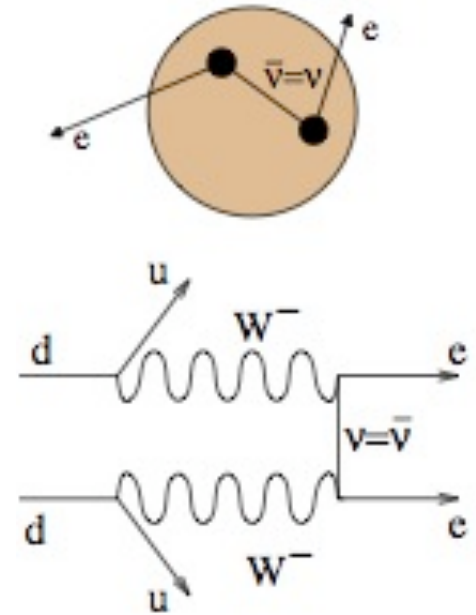


Majorana

- It is still not clear today whether neutrinos are **Dirac or Majorana particles**. The distinction make sense only because they are massive.

Neutrinoless double beta decay

- Nuclear process: $(A, Z) \rightarrow (A, Z+2) + 2 e^-$
- Can only happen if lepton number is not conserved, unlike the Standard Model allowed 2ν mode: $(A, Z) \rightarrow (A, Z+2) + 2 e^- + 2 \bar{\nu}$
- The decay probability depends on the effective Majorana mass $m_{\beta\beta}$ of the neutrino exchanged between the two electron vertexes.
- The measurable quantity is the half-life $(\tau_{1/2}^{0\nu})$ of the decay:



Phase space factor: $\sim Q^5$

Effective neutrino mass

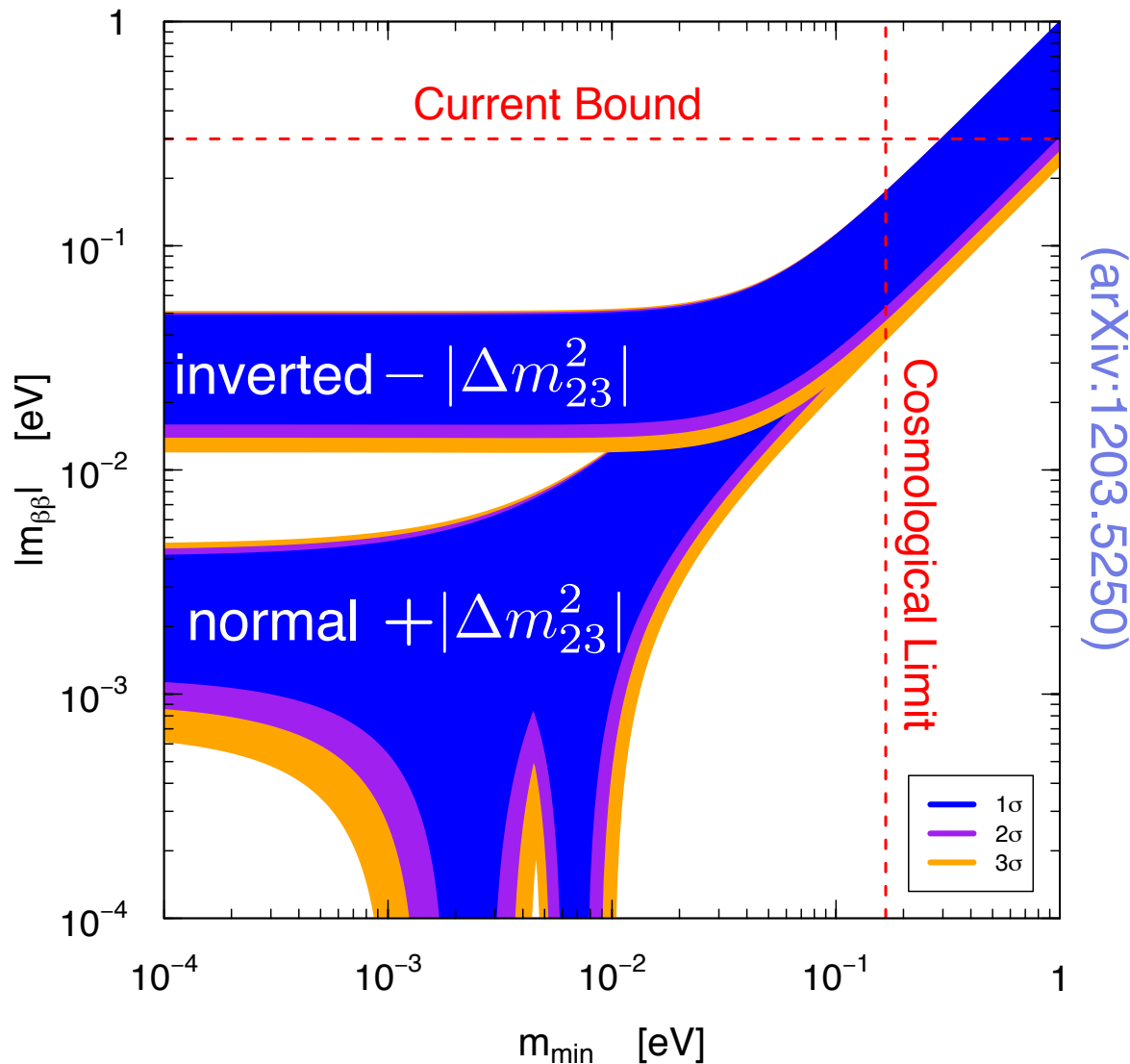
$$\frac{1}{(\tau_{1/2}^{0\nu})} = G(Q, Z) |M_{nucl}|^2 |m_{\beta\beta}|^2 = F_N \frac{|m_{\beta\beta}|^2}{m_e^2}$$

Nuclear Matrix element

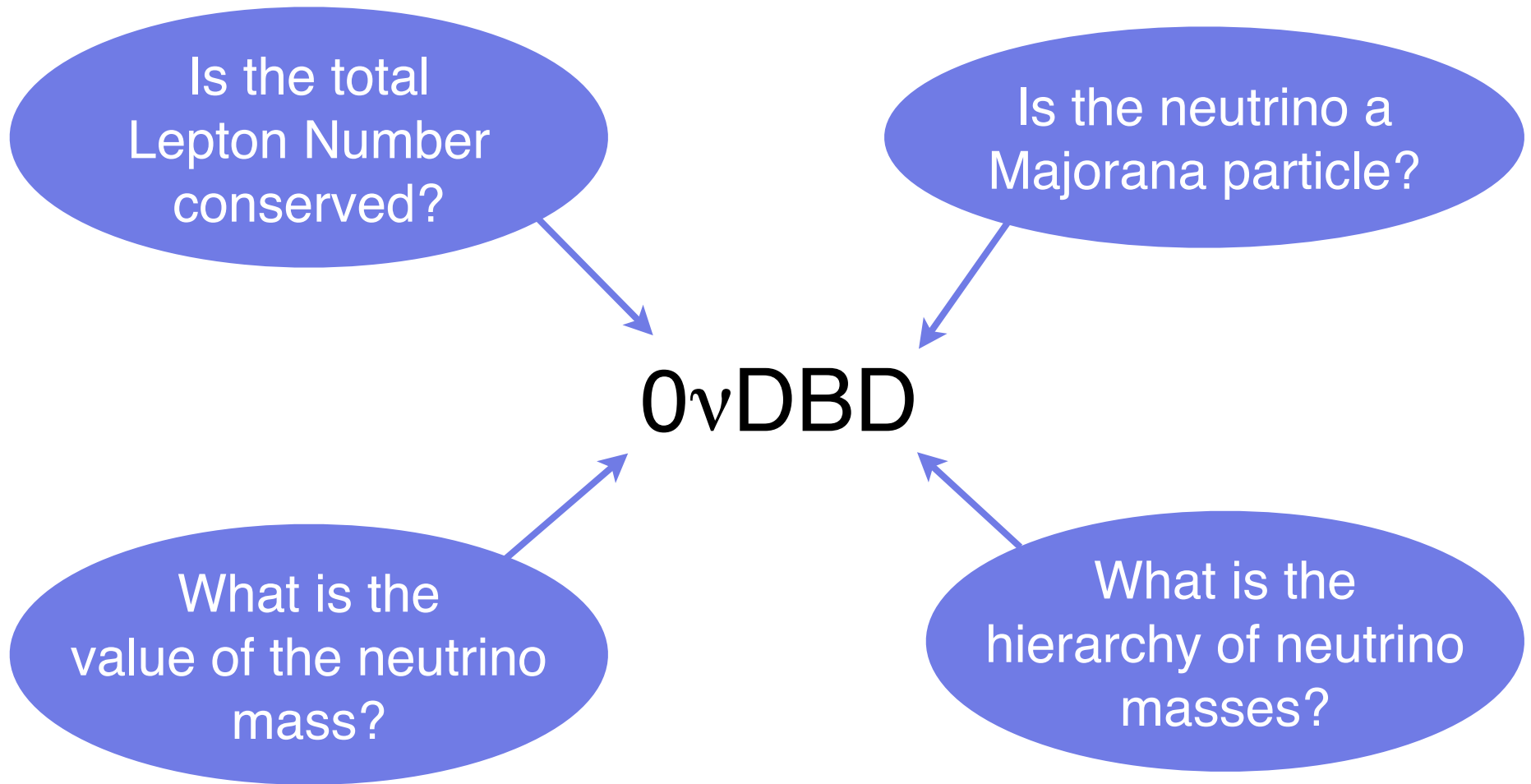
Effective Majorana mass

- Effective neutrino Majorana mass in terms of the measured oscillation parameters and the unknown lightest neutrino mass:

$$m_{\beta\beta} = f(\theta_{12}, \theta_{13}, \theta_{23}, \Delta m_{12}^2, \pm |\Delta m_{23}^2|, m_{\min})$$



Summary so far

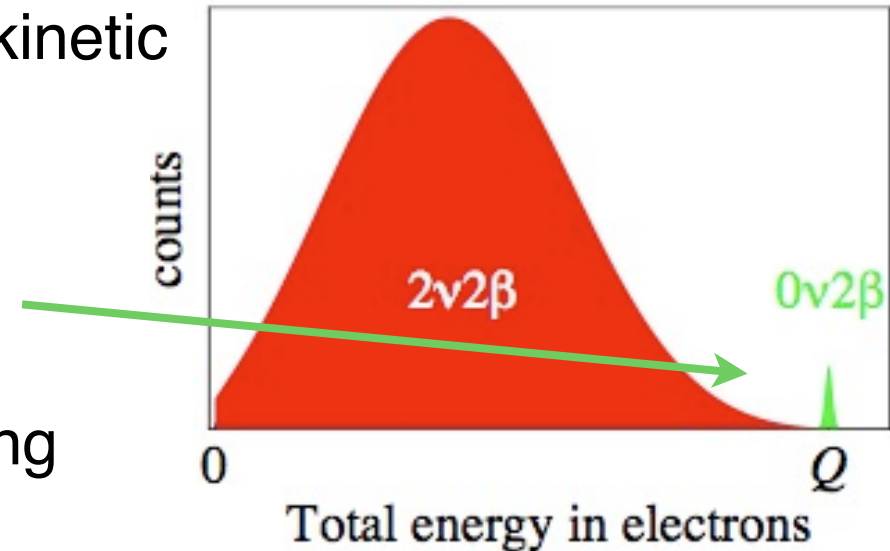


1 measure → many answers...

...extremely challenging though.

0ν DBD in Experiments

- Experiments measure the sum of the kinetic energies of the two emitted electrons.
- Signature:** monochromatic line at the Q-value of the decay.
- Sensitivity ($S^{0\nu}$):** lifetime corresponding to the minimum number of detectable events above background at a given C.L.:



$$S^{0\nu} = \ln 2N_A \cdot \frac{a}{A} \left(\frac{Mt}{B\Delta E} \right)^{1/2} \cdot \epsilon$$

Isotopic abundance (%) → a
 Detector mass (kg) → M
 Measurement time (y) → t
 Atomic mass → A
 Background (counts/keV/kg/y) → B
 Energy Resolution (keV) → ΔE
 detection efficiency → ϵ

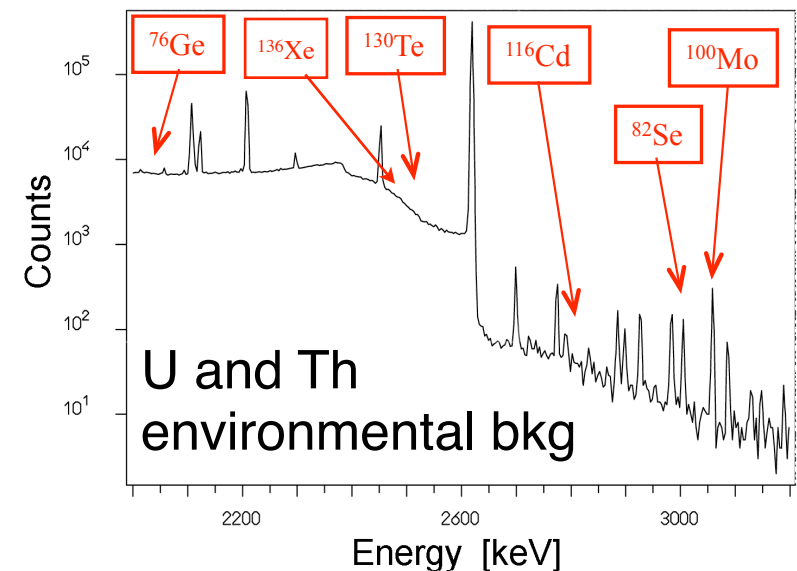
Isotope choice

$$|m_{\beta\beta}|^2 = 1 / \left[G(Q, Z) |M^{0\nu}|^2 \tau_{1/2}^{0\nu} \right]$$

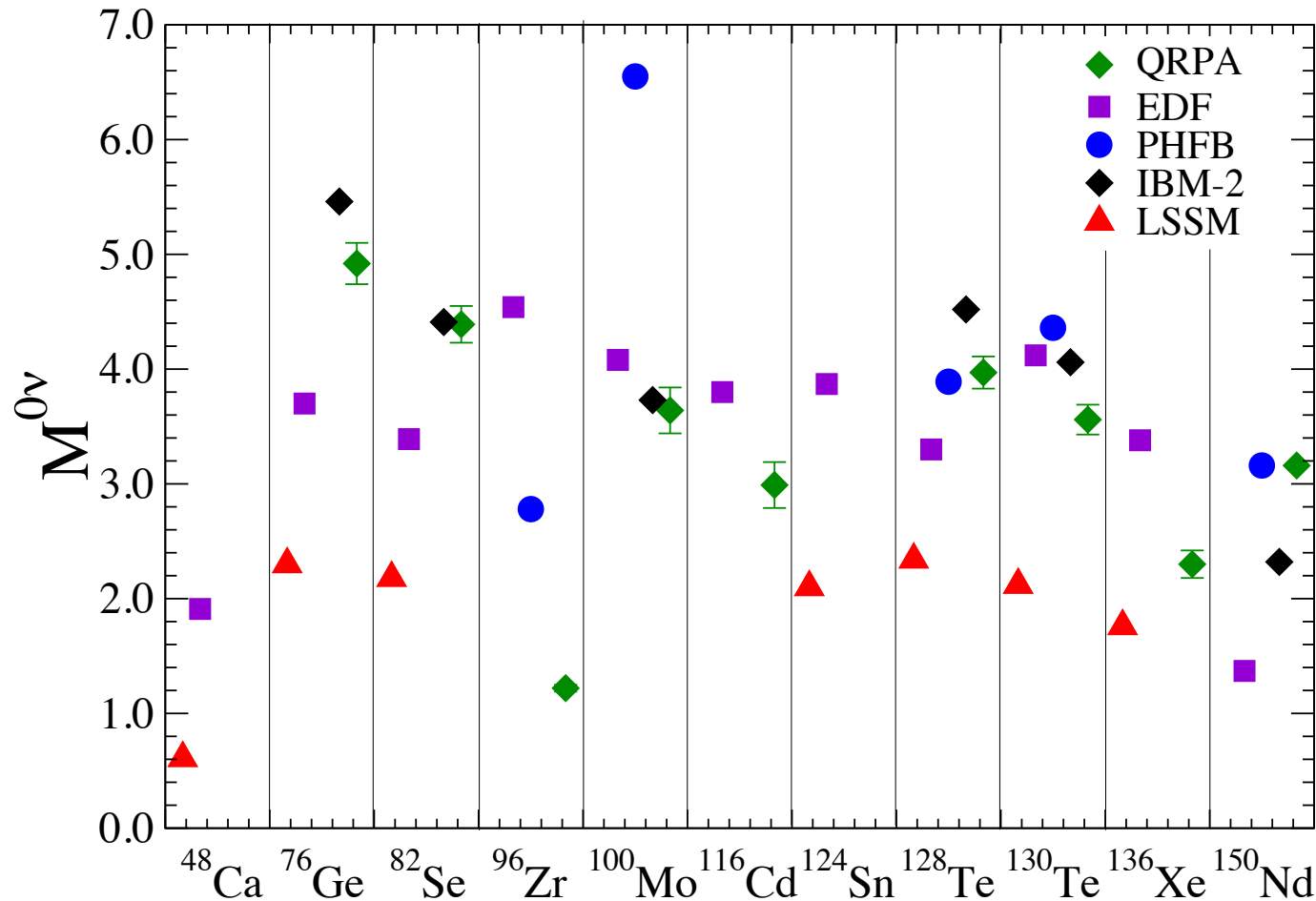
- 0ν DBD candidates of experimental interest: ([arXiv:1201.4916](https://arxiv.org/abs/1201.4916))

isotope	$G^{0\nu}$ [$10^{-14}y^{-1}$]	$Q_{\beta\beta}$ [keV]	nat. abund. [%]	$T_{1/2}^{2\nu}$ [$10^{20}y$]
^{48}Ca	6.3	4273.7	0.187	0.44
^{76}Ge	0.63	2039.1	7.8	15
^{82}Se	2.7	2995.5	9.2	0.92
^{100}Mo	4.4	3035.0	9.6	0.07
^{116}Cd	4.6	2809	7.6	0.29
^{130}Te	4.1	2528	34.2	9.1
^{136}Xe	4.3	2461.9	8.9	21
^{150}Nd	19.2	3367.3	5.6	0.08

- In general, $Q > 2615$ keV isotopes are preferred because they lie above the natural radioactivity edge.
- However the choice has been dominated so far by technology compromises.



Nuclear amplitude $|m_{\beta\beta}|^2 = 1 / \left[G(Q, Z) |M^{0\nu}|^2 \tau_{1/2}^{0\nu} \right]$

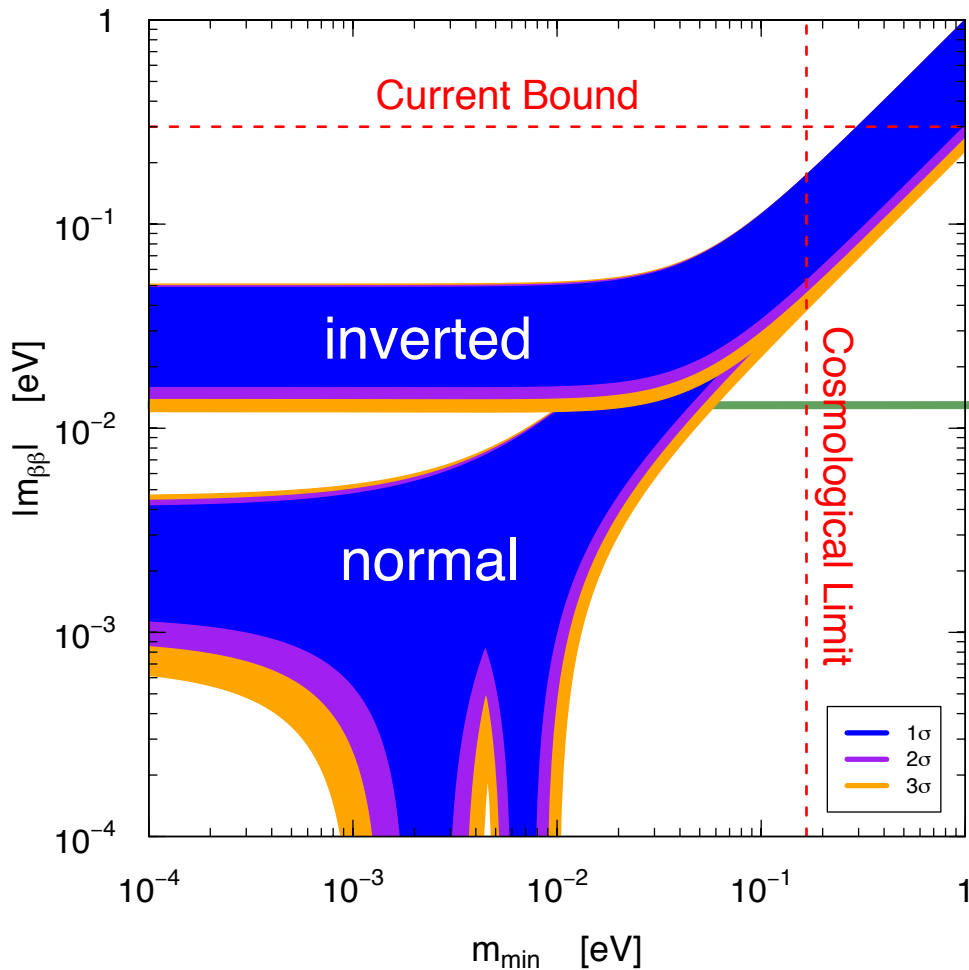


arXiv:1201.4916

- The spread in the values does not influence the isotope choice.
- Generate problems when comparing exclusions and evidences from experiments running different isotopes.

Experimental race

$$S^{0\nu} = \ln 2N_A \cdot \frac{a}{A} \left(\frac{Mt}{B\Delta E} \right)^{1/2} \cdot \epsilon$$



Entire inverted hierarchy:

$M \sim 1$ ton

$a \sim 90\%$

$B \Delta E \sim 1$ count/ton/year

Normal hierarchy:

Dream

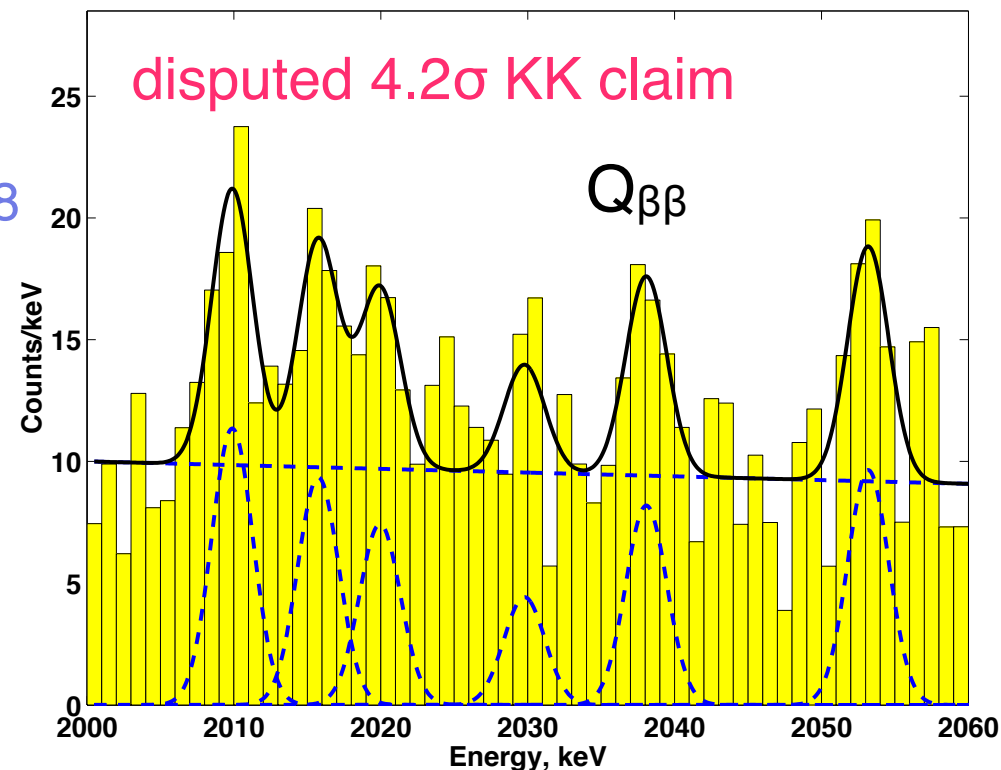
Germanium diodes

The Klapdor-Kleingrothaus claim

- Past Germanium-diodes-based experiment: **Heidelberg-Moscow**
 - ▶ background 0.12 cpy/keV/kg
 - ▶ **no evidence** in 10 kg(^{76}Ge) \times 3.5y
- A small group departed from the collaboration and published their own (improved) analysis of the data.

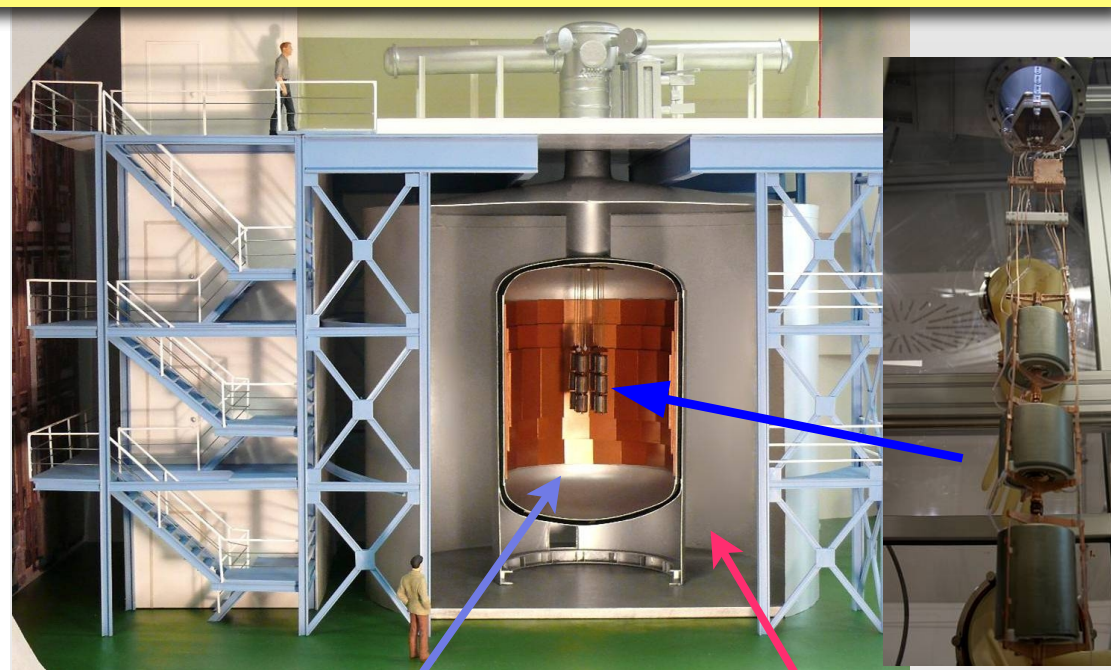
- The group claimed for the peak observation: [Phys.Lett.B 586\(2004\)198](#)

- ▶ $T^{0\nu}_{1/2} = (2.23 \pm 0.44) \times 10^{25} \text{ yr}$
- ▶ $\langle m_{\beta\beta} \rangle = 0.32 \pm 0.03 \text{ eV}$



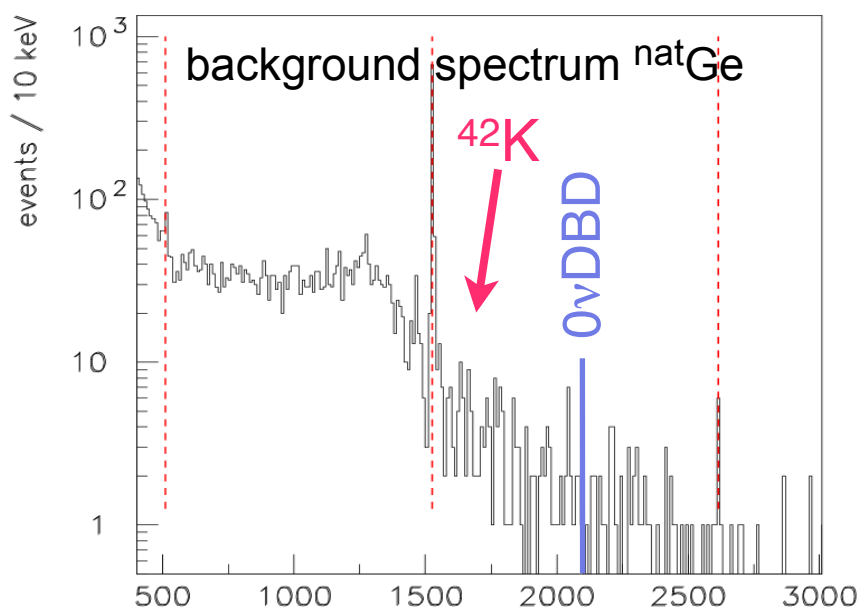
Gerda - @LNGS

- Present: **Gerda phase I**
3 ^{nat}Ge diodes: 7.6 kg
6 ^{enr}Ge diodes 14.6 kg
- $\Delta E = 3 \text{ keV FWHM}$
- Data taking delayed by an unexpected ^{42}Ar contamination in the LAr



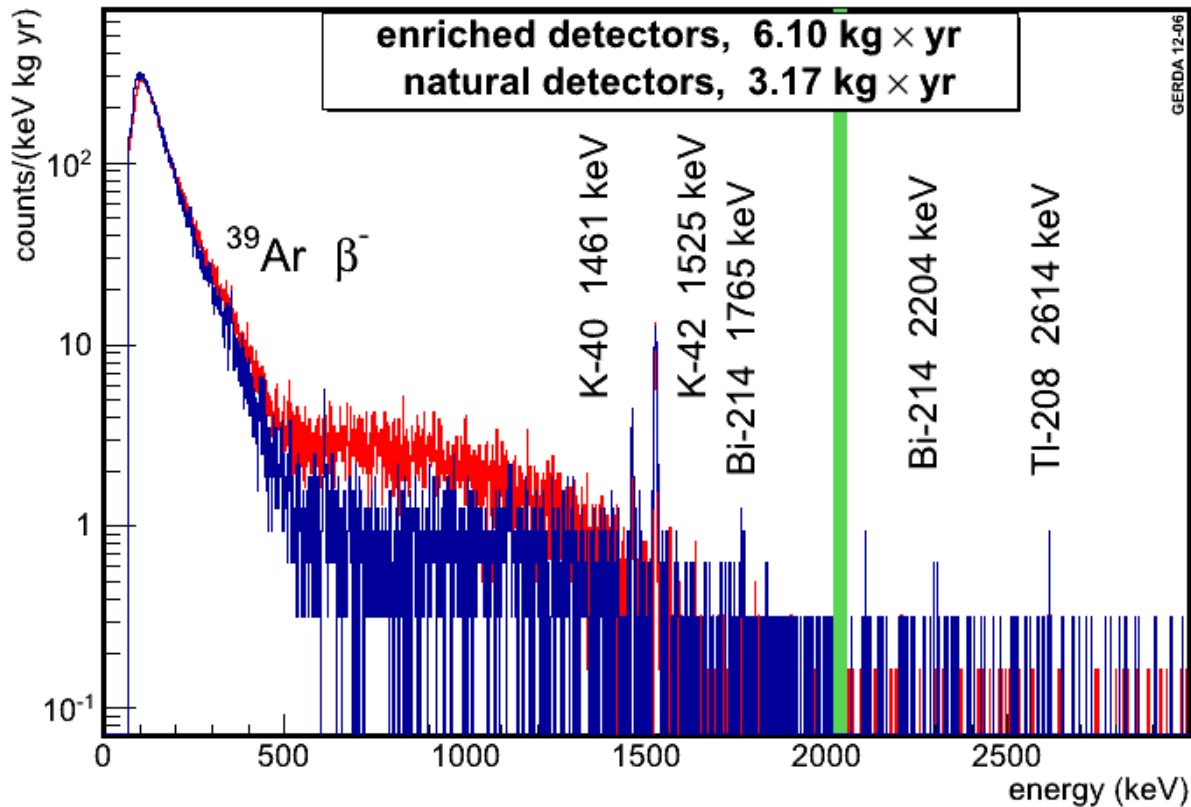
Liquid Argon
(cooler)

Water Cherenkov
(muon veto)

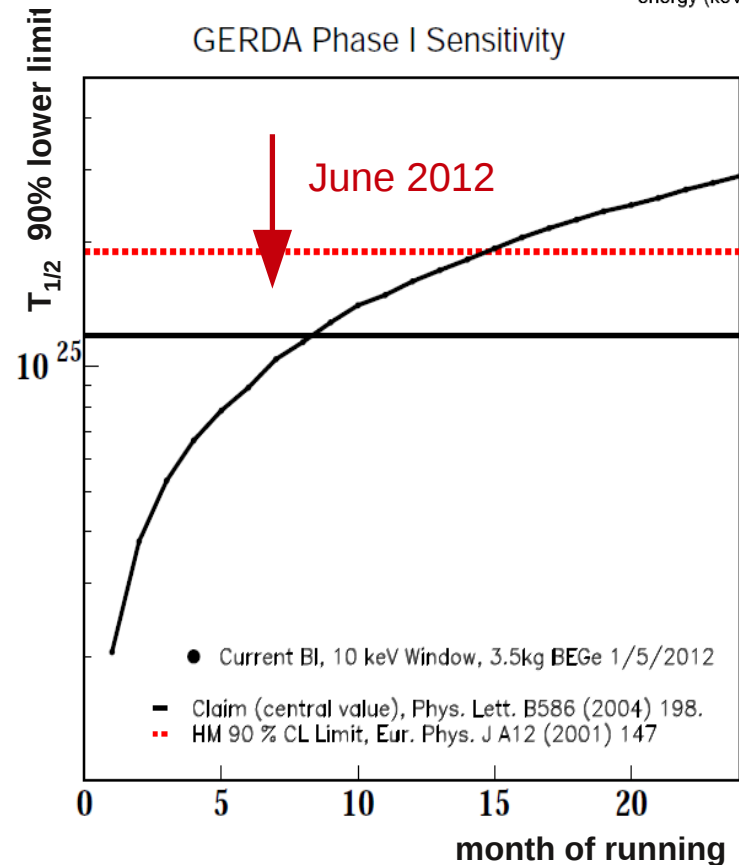
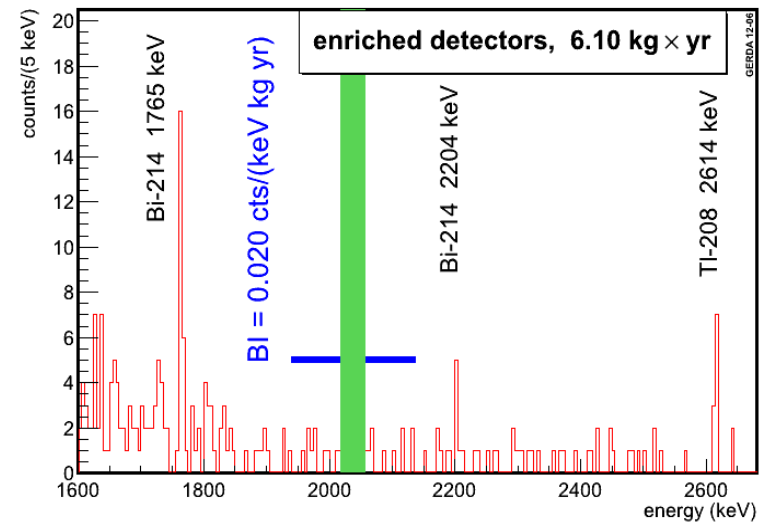


- Diodes p-contact attracts the ^{42}K ions which then decay in the detector: bkg 6x higher than the design.
- Bkg successfully reduced by catching the ^{42}K ions with electric fields inside the LAr.

Gerda: good data from 11/11 *P. Grabmayr at Neutrino 2012*



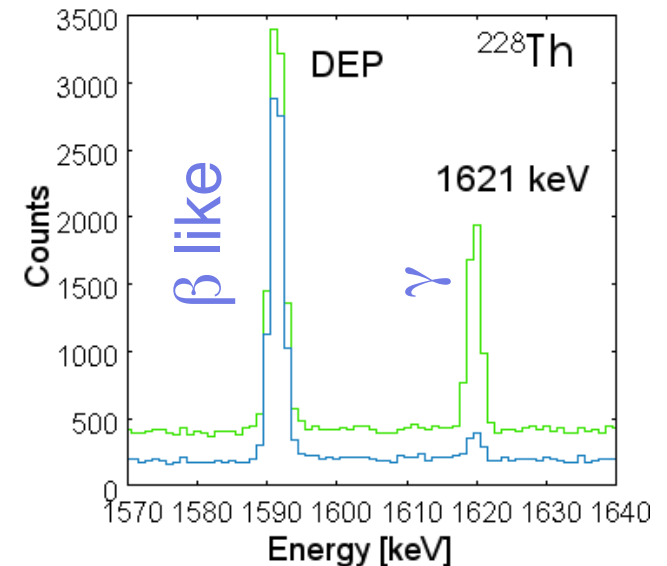
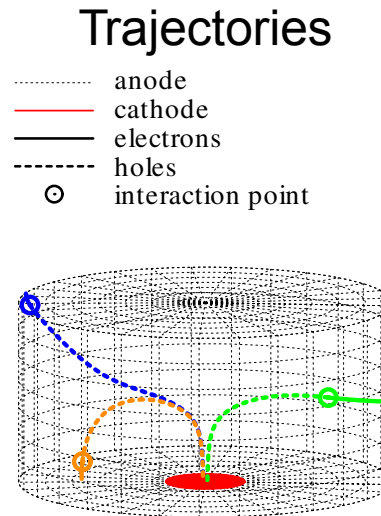
- $0\nu\text{DBD}$ region blinded.
- Should check KK claim in ~ 1.5 y.



Germanium diodes: future

- Gerda phase II (2013)

- ▶ Increase mass to ~ 25 kg
- ▶ New detectors: BEGe, Allow γ/β discrimination via Multi/Single site identification



- ▶ Veto external gammas with readout of scintillation light of LAr
- ▶ New bkg. lower by 1 order of mag.: 10^{-3} counts/keV/kg/y

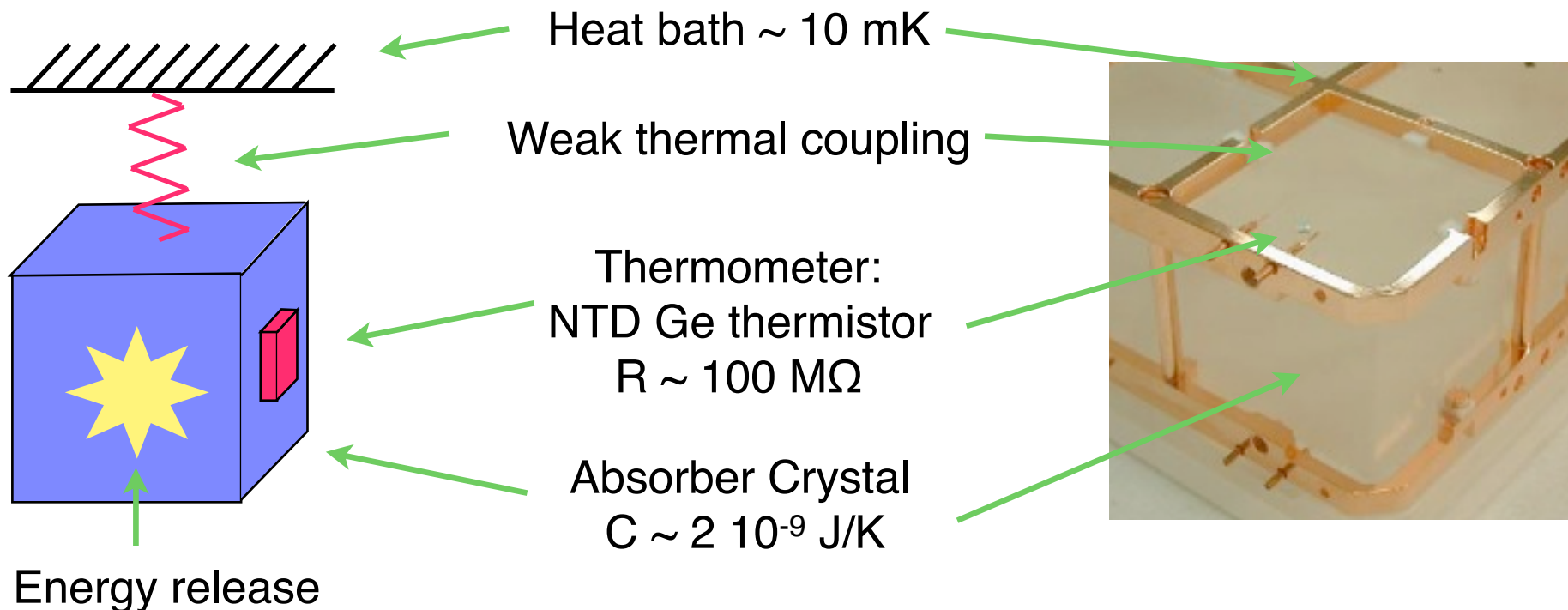
- Majorana Demonstrator (2013)

- ▶ Deploy 40 kg of $^{\text{enr}}\text{Ge}$
- ▶ Determine the best technology for a 1 ton Ge experiment to be pursued together with the Gerda collaboration.

Bolometers

Bolometric technique

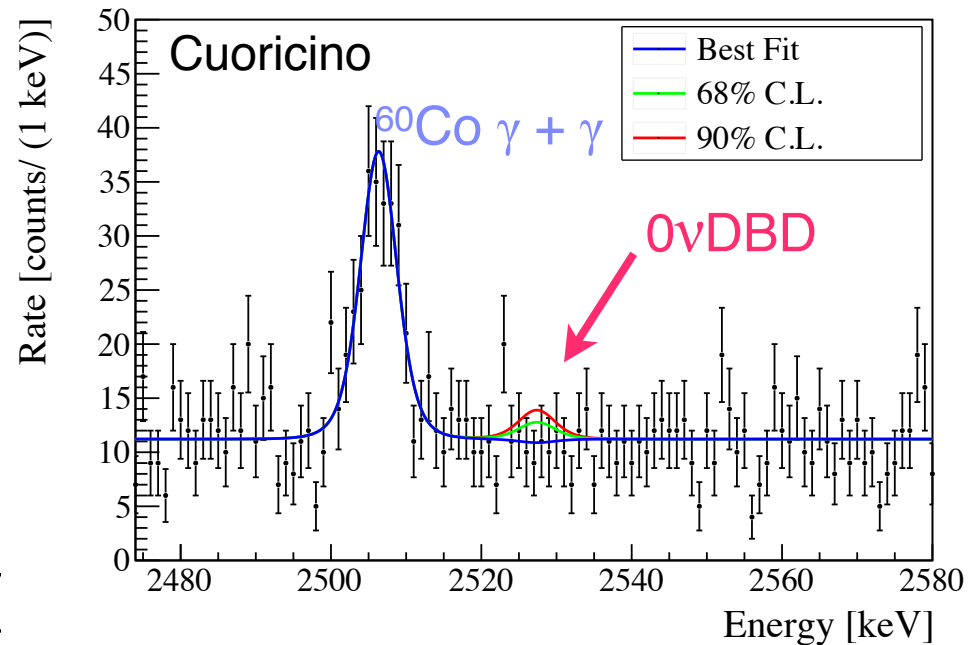
- Particle energy converted into phonons → temperature variation.
- Crystals embedding $0\nu\text{DBD}$ source.
- Low crystal heat capacitance and low base temperature to see small temperature variations → $\Delta T \sim E/C$



- Detector response in this configuration: ~ 0.1 mK / MeV
- Resolution @ $0\nu\text{DBD}$ \sim few keV

CUORE - @ LNGS

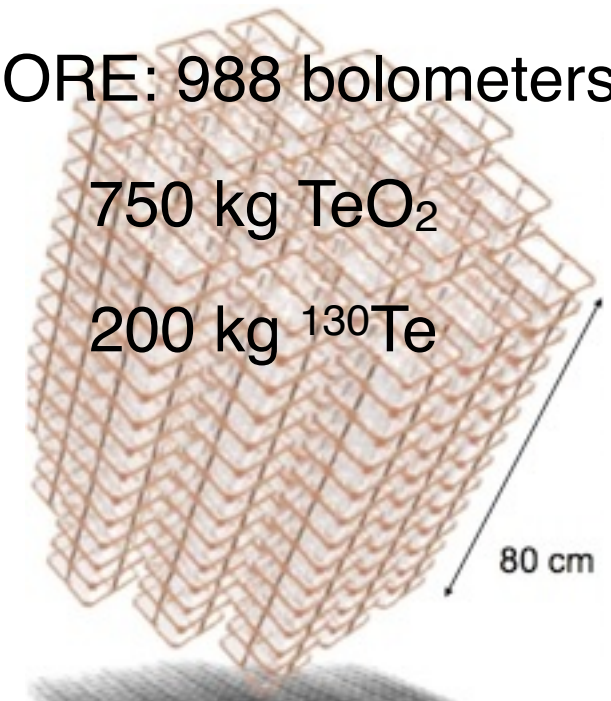
- natTeO_2 bolometers (34% ^{130}Te), 750g each ($\Delta E = 5$ keV FWHM)
- Past: **Cuoricino**
 - ▶ 62 bolometers
11 kg (^{130}Te) $\times 2$ y,
Bkg: 0.16 cpy/keV/kg
 - ▶ $T^{0\nu}_{1/2} > 2.8 \times 10^{24}$ years (90% CL)
 $\langle m_{\beta\beta} \rangle < 300 \sim 700$ meV
- Future: **Cuore** (data taking in 2015)
 - ▶ Expected bkg: 0.01~0.04 cpy/keV/kg
 - ▶ Exp. $T^{0\nu}_{1/2} > 1.6 \times 10^{26}$ years
 $\langle m_{\beta\beta} \rangle < 40 \sim 94$ meV
- Present: **Cuore-0**, a CUORE-like tower.
 - ▶ same mass of Cuoricino, 0.05 cpy/keV/kg.



CUORE: 988 bolometers

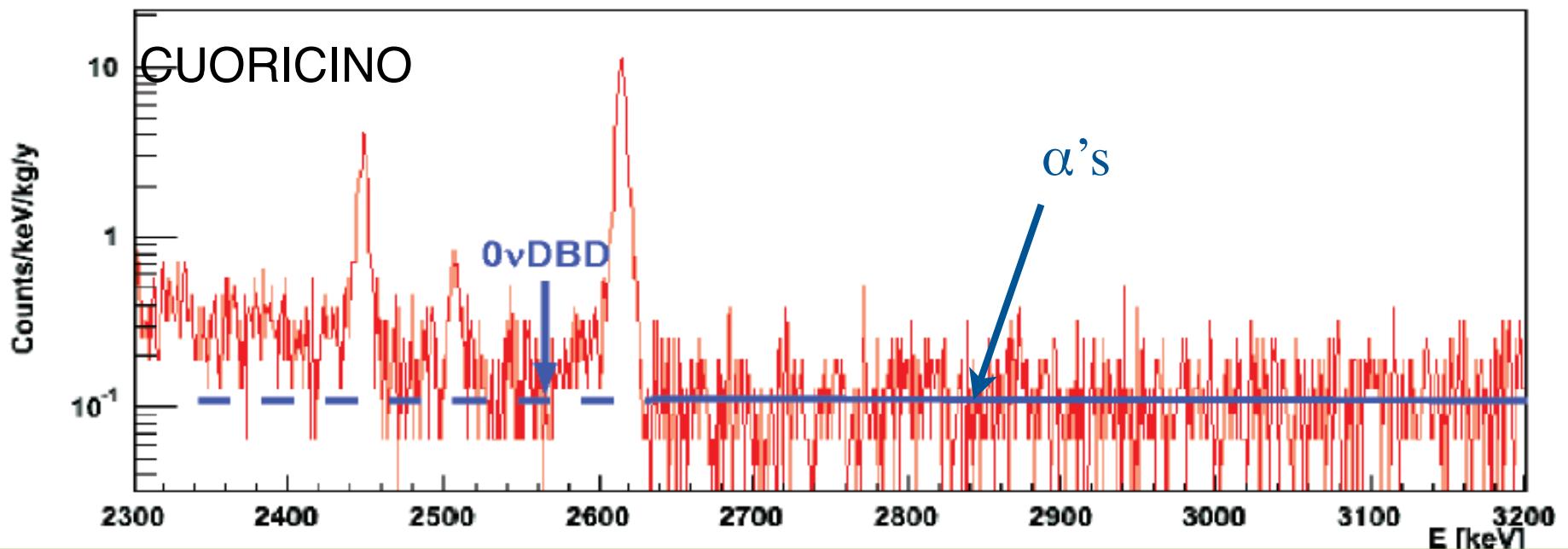
750 kg TeO_2

200 kg ^{130}Te



CUORE: the α nightmare

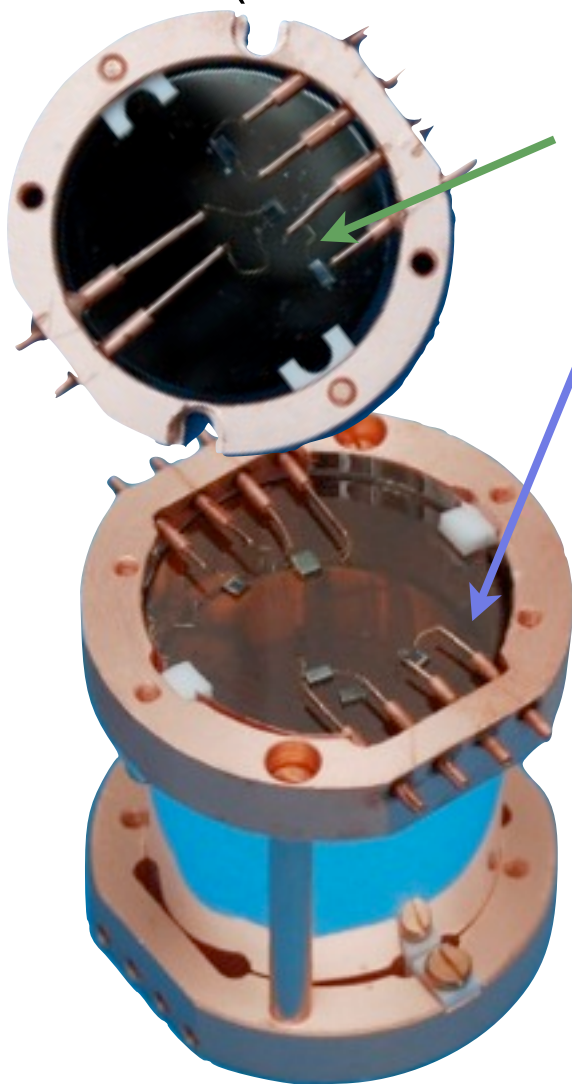
- MC: the background in CUORICINO is due to degraded α particles which release only a part of their energy in the detector (surface contaminations, mainly in copper).



- TeO_2 bolometers, per se, do not allow to discriminate β and α particles.
 - ▶ α bkg partially reduced by cleaning the detector parts.

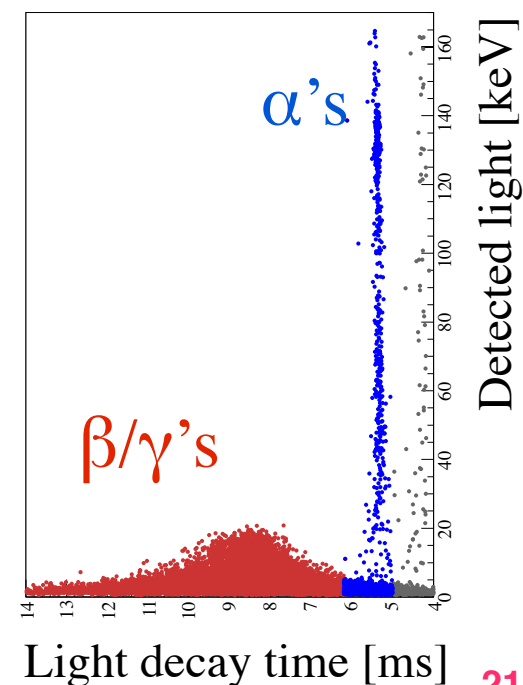
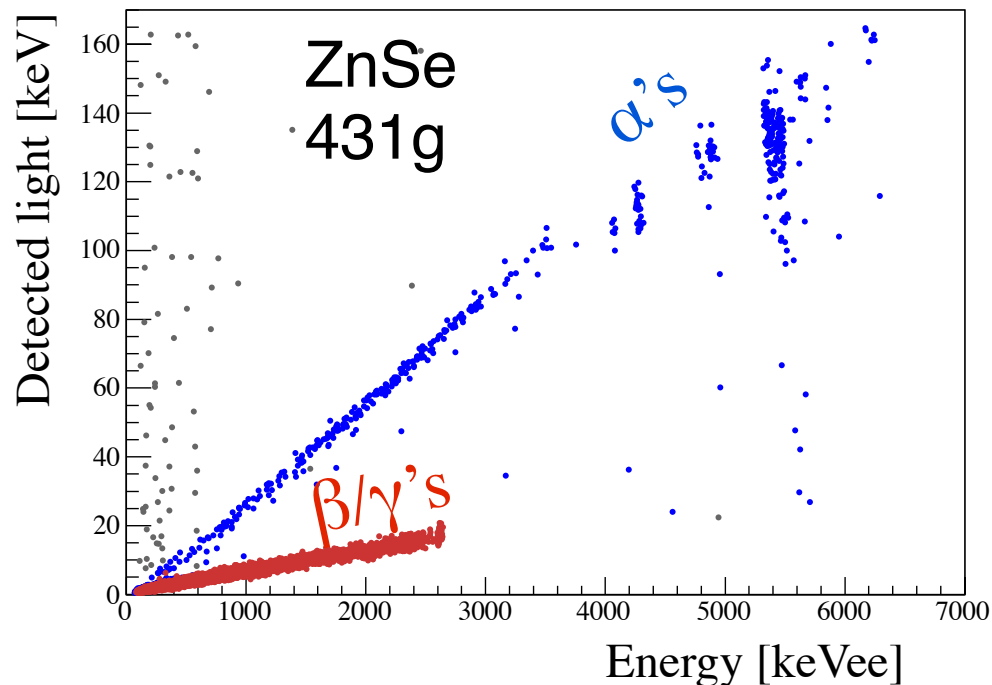
LUCIFER - @ LNGS

- Scintillating bolometers to discriminate the α background, enriched in ^{82}Se or ^{100}Mo .
 - ▶ Target: define the technology for a ZERO background (<1 count/ton/year), ~1-ton isotope experiment after CUORE.



Light detector:
Ge bolometer

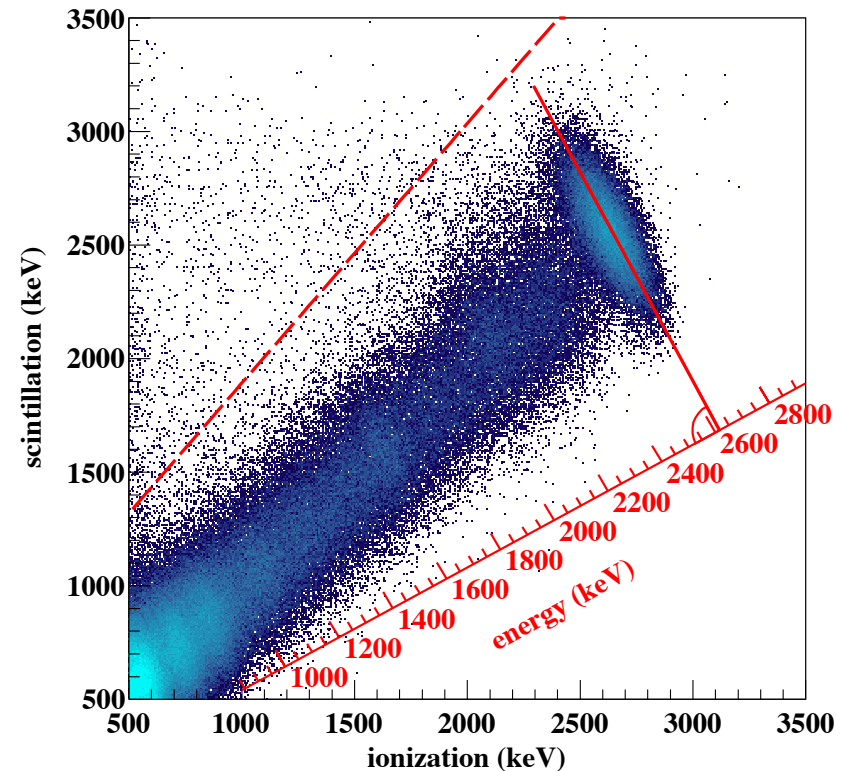
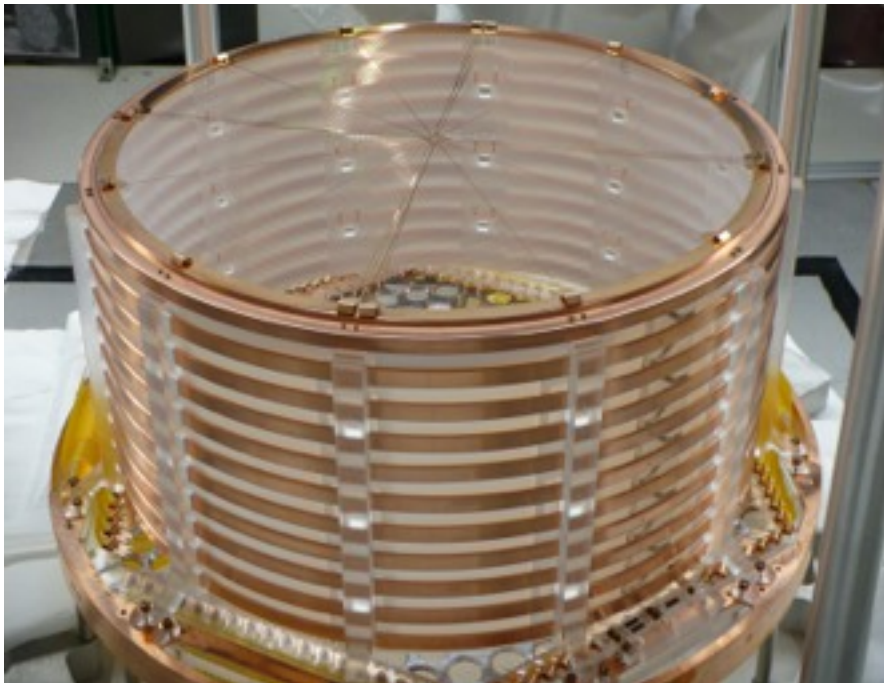
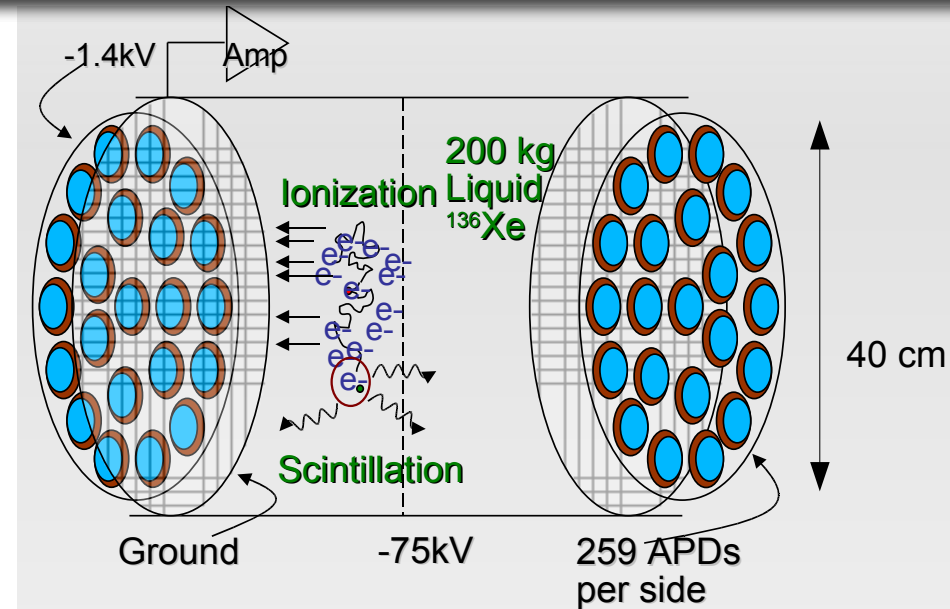
Absorber bolometer: Zn^{82}Se or $\text{Zn}^{100}\text{MoO}_4$



Liquid scintillators

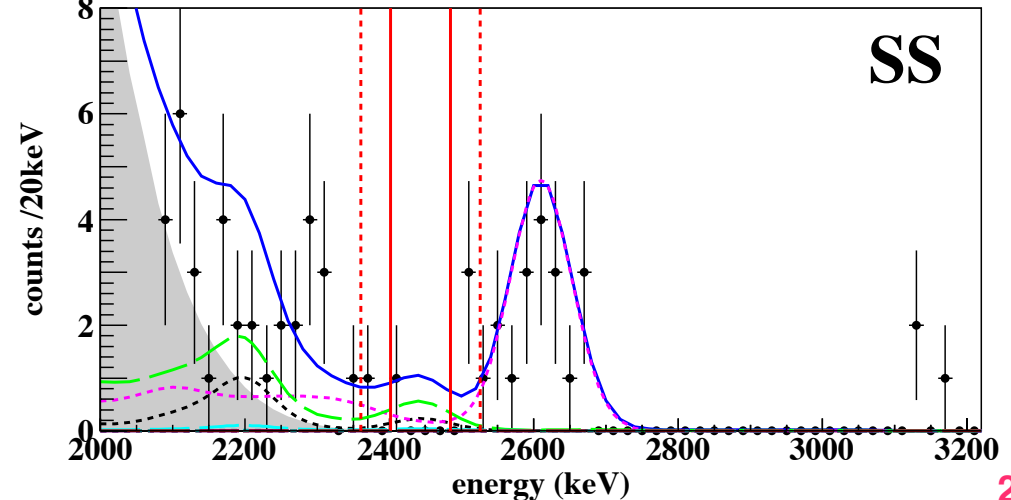
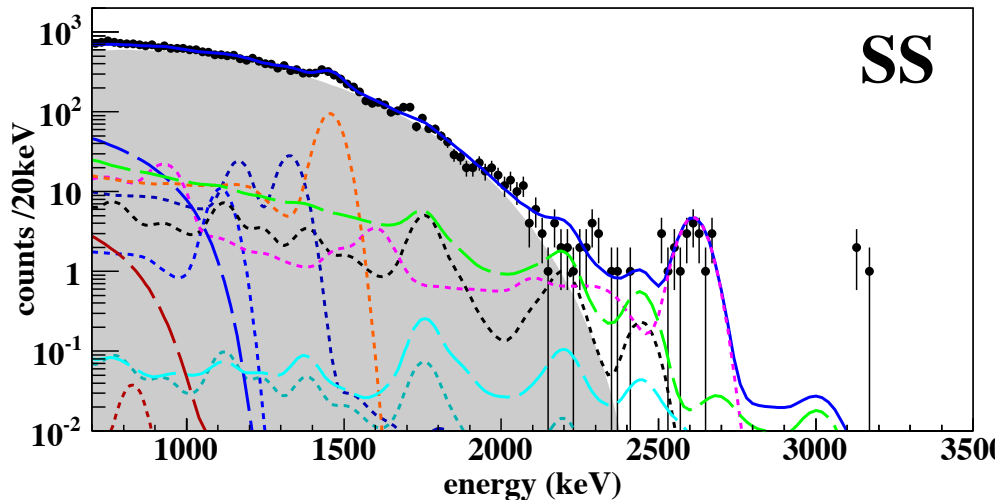
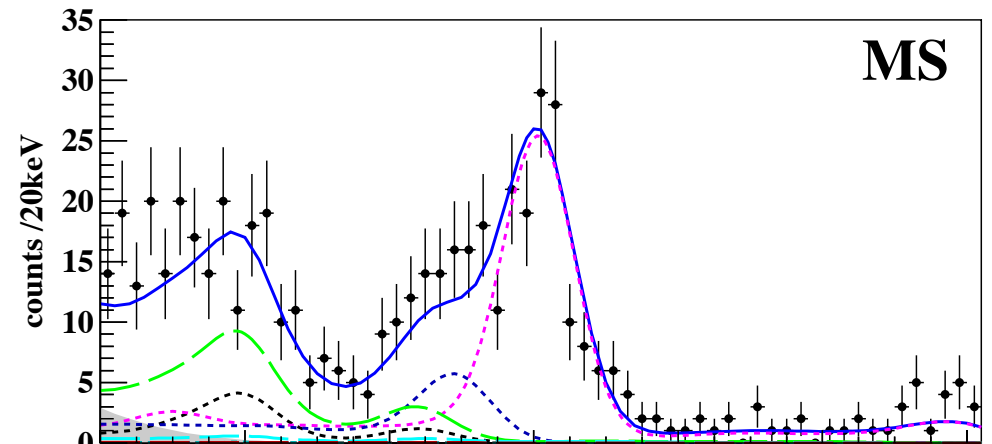
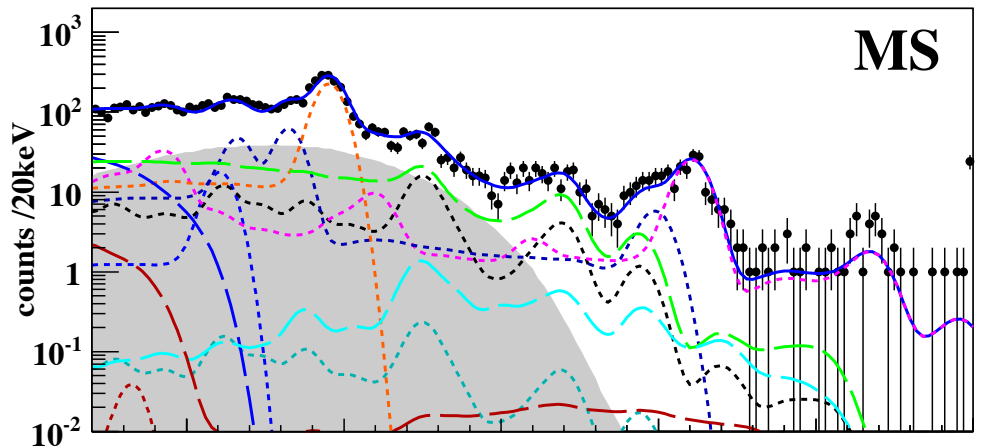
Exo-200 - @WIPP

- Double sided TPC
- 200 kg Liquid ^{136}Xe
- $\Delta E = 100 \text{ keV FWHM}$
- Multi/Single site discrimination
- Bkg: 0.0015 cpy/keV/kg
- Data taking started in Spring 2011

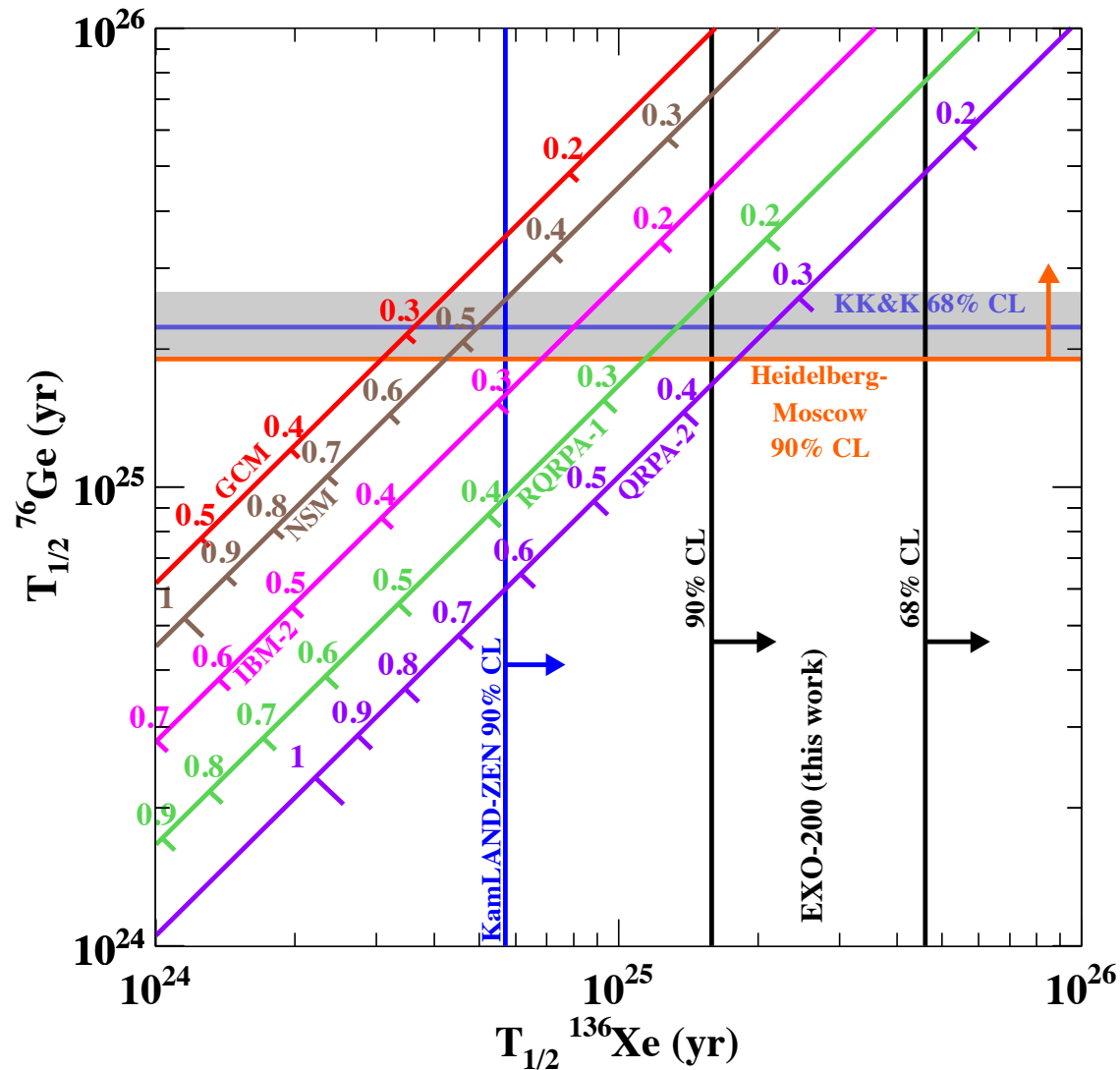


Exo-200 results (32.5 kg×yr)

- First observation of $2\nu\beta\beta$ decay of ^{136}Xe
 - ▶ $T^{2\nu}_{1/2} = (2.23 \pm 0.017 \text{ stat} \pm 0.22 \text{ sys}) \cdot 10^{21} \text{ yr}$ (6x faster than previous limit)
- $T^{0\nu}_{1/2} > 1.6 \cdot 10^{25} \text{ yr}$ \longrightarrow $\langle m\beta\beta \rangle < 140\text{--}380 \text{ meV}$ (90% C.L.) [arXiv:1205.5608](https://arxiv.org/abs/1205.5608)

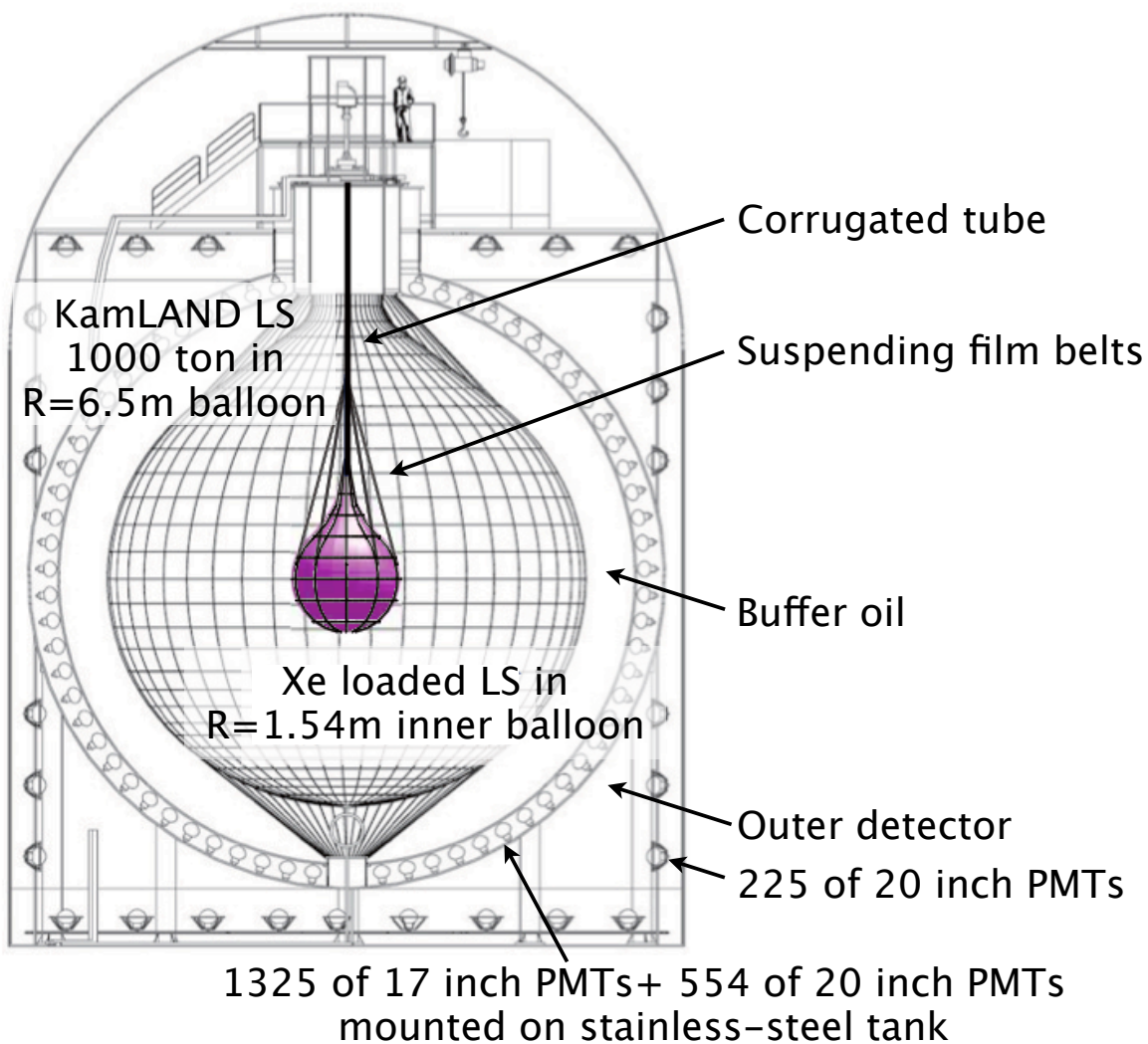


Exo vs KK claim



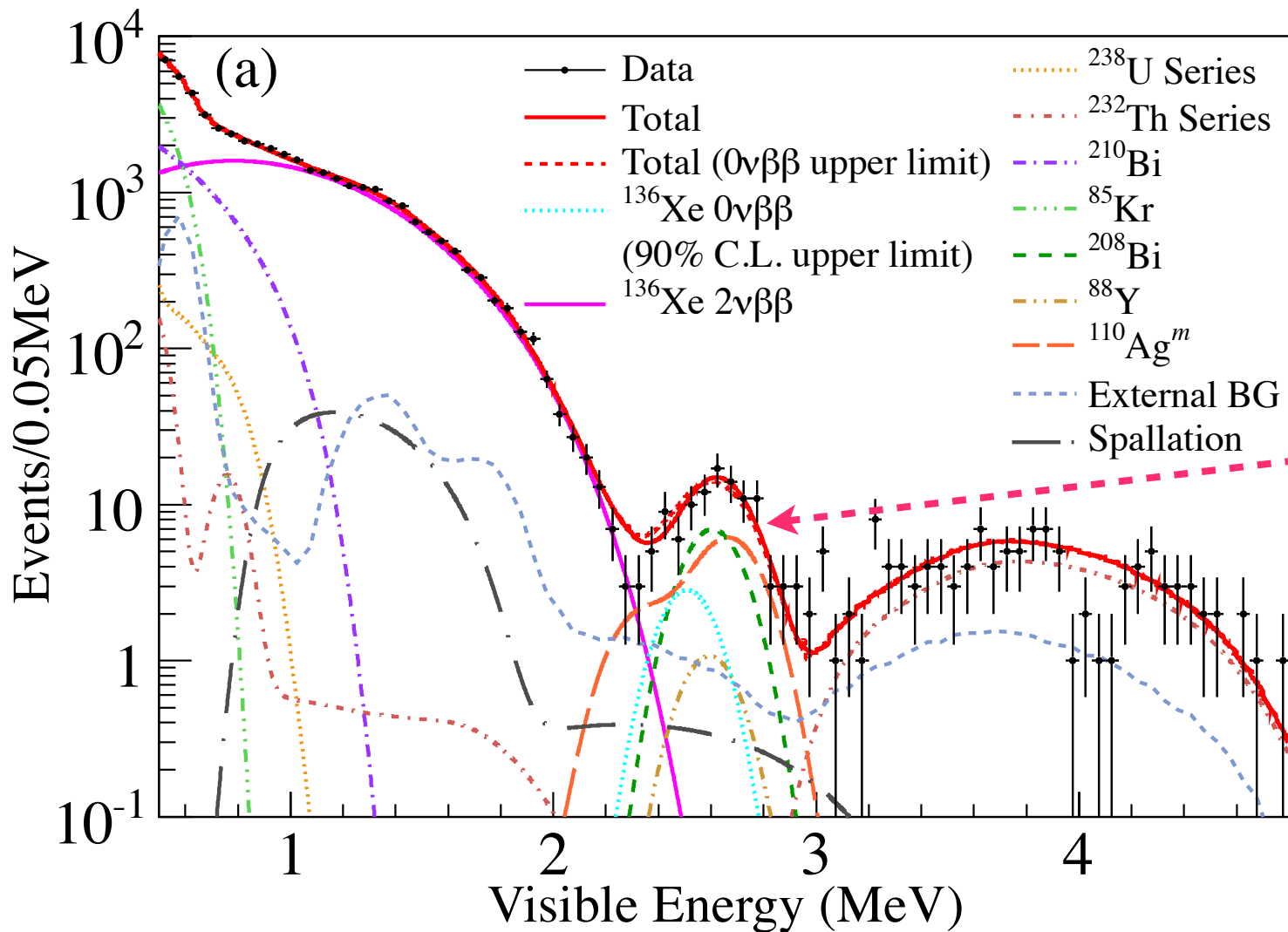
- Now: Improving energy resolution and pattern recognition
- Future: 1T with tagging of $^{136}\text{Ba}^{++}$ ion (^{136}Xe daughter).

Kamland-Zen



- Total: ~320 kg 90% enriched ^{136}Xe (2.4 wt%)
- Fiducial ($\sim R/2$): **125 ± 7 kg ^{136}Xe**
- U : 3.5×10^{-18} g/g
Th: 5.2×10^{-17} g/g
- Res@2615 keV: **250 keV FWHM**
- Taking data since September 24, 2011

Kamland-Zen: results (112 days)



Background peak very close to the $0\nu\text{DBD}$: $^{110\text{m}}\text{Ag}$ or ^{208}Bi , ^{88}Y from Spallation or from Fukushima.

$0\nu\text{DBD}$ hypothesis rejected at 8σ

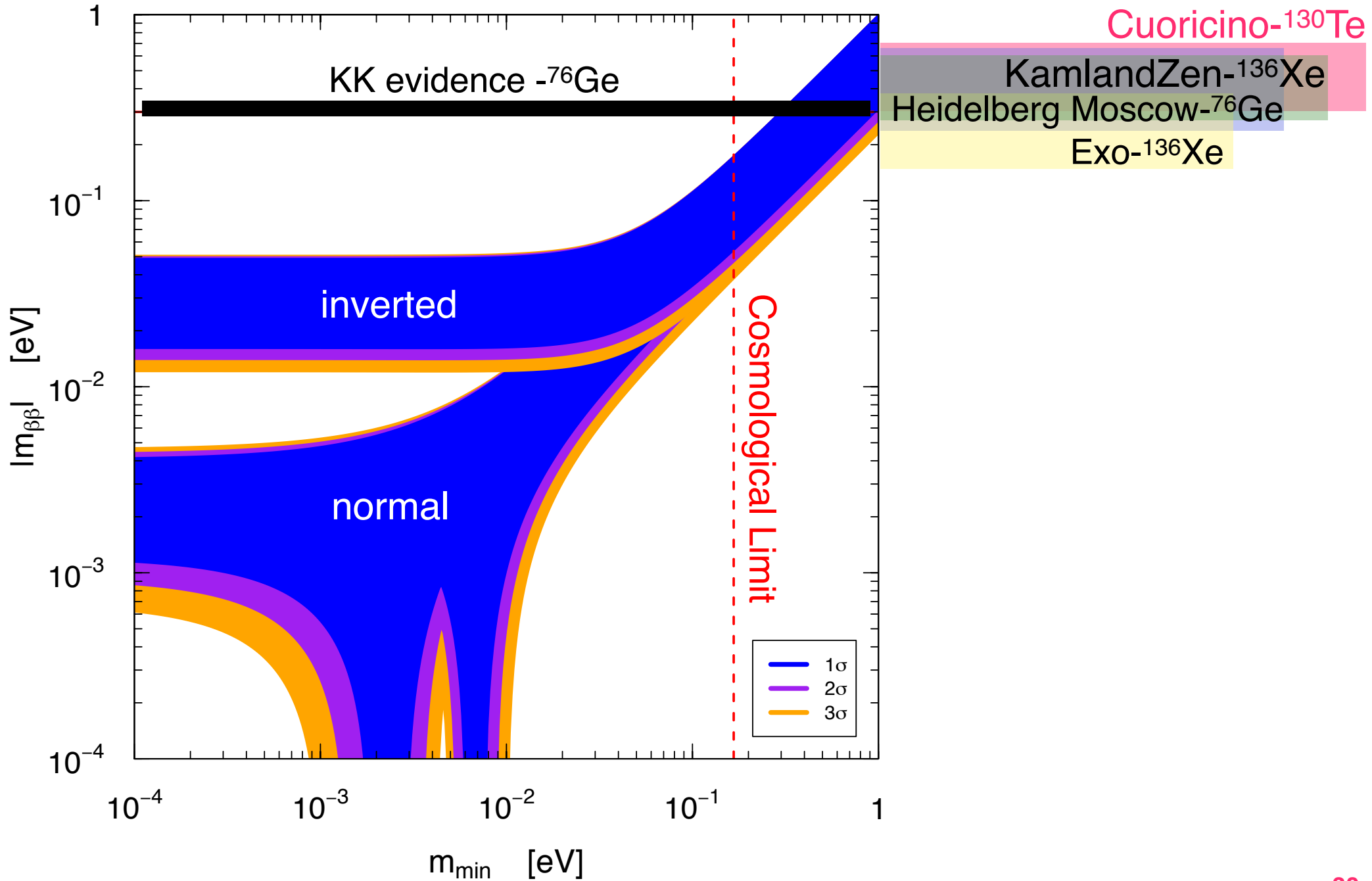
- $T^{0\nu}_{1/2} > 6.2 \times 10^{24}$ yr (90% CL) \longrightarrow $\langle m_{\beta\beta} \rangle < 260 \sim 540$ meV
- $T^{2\nu}_{1/2} = 2.30 \pm 0.02(\text{stat}) \pm 0.12(\text{syst}) \times 10^{21}$ yr (arXiv:1205.6372)

- **Soon (target sensitivity: ~ 80 meV):**
 - ▶ improved purification of Xe-LS to remove peaking contaminants ^{110m}Ag , ^{208}Bi , ^{88}Y -> Target: more than a factor 100 reduction.
- **Near future (target sensitivity ~ 40 meV):**
 - ▶ increased Xe amount and cleaner balloon
- **Future (target sensitivity ~ 20 meV):**
 - ▶ improve resolution with new LS and more efficient light collection.

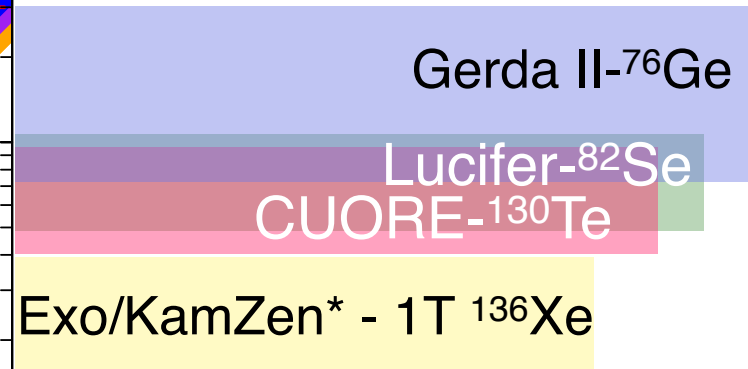
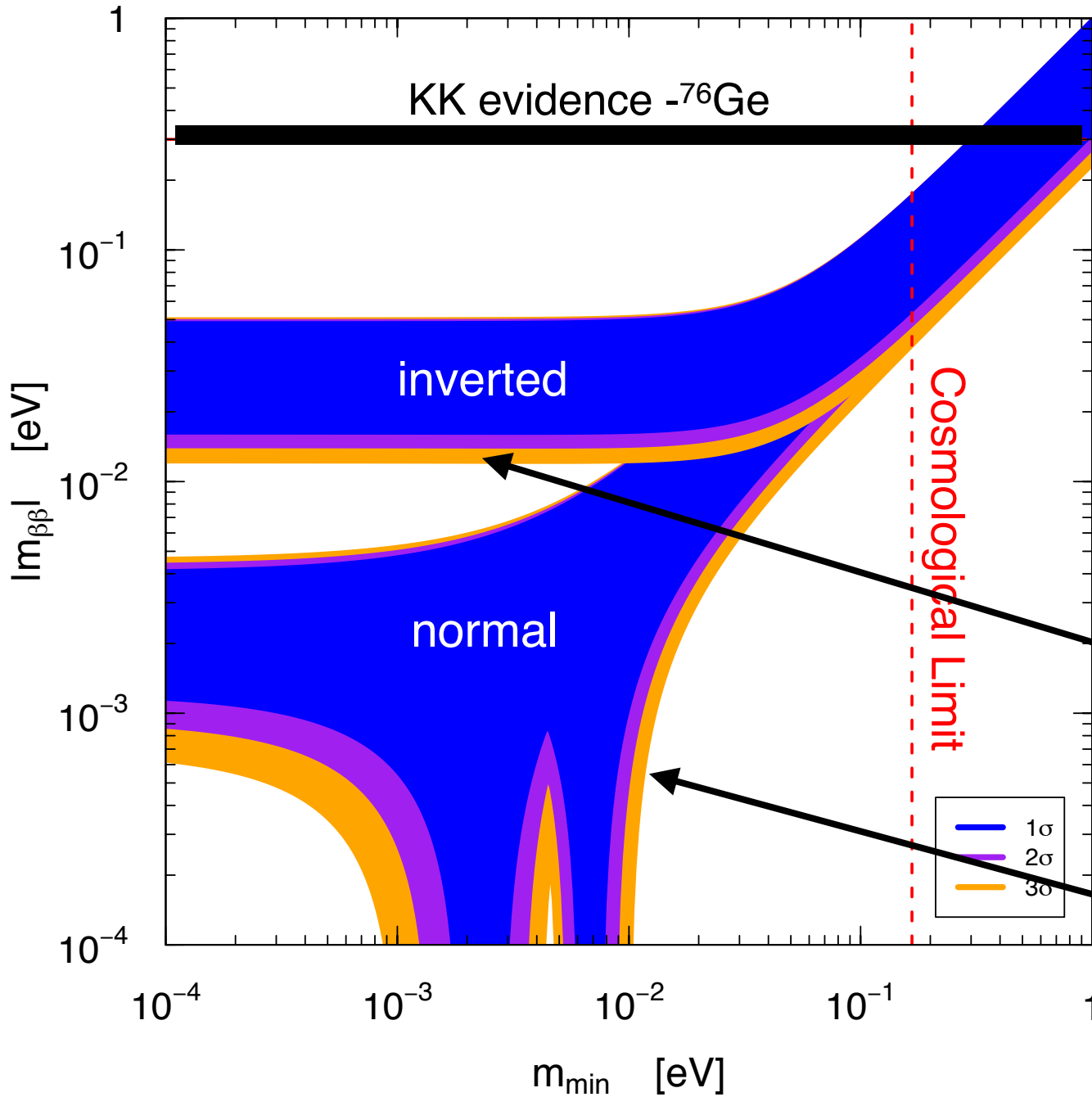
Other projects

- **Amore**: $\text{Ca}^{100}\text{MoO}_4$ bolometers (development)
- **Next**: 100kg, high pressure, ^{136}Xe -gas TPC (development)
- **Sno+**: 44kg ^{150}Nd in liquid scintillator (development)
- **SuperNemo**: 100 kg of various isotopes, tracking (development)
- ... and others.

Conclusion: present limits



Conclusion: next reaches



Many ideas to reach this point: a slow but intriguing race....
New players may overcome old ones...

Our grandsons will think about

* with detector upgrade