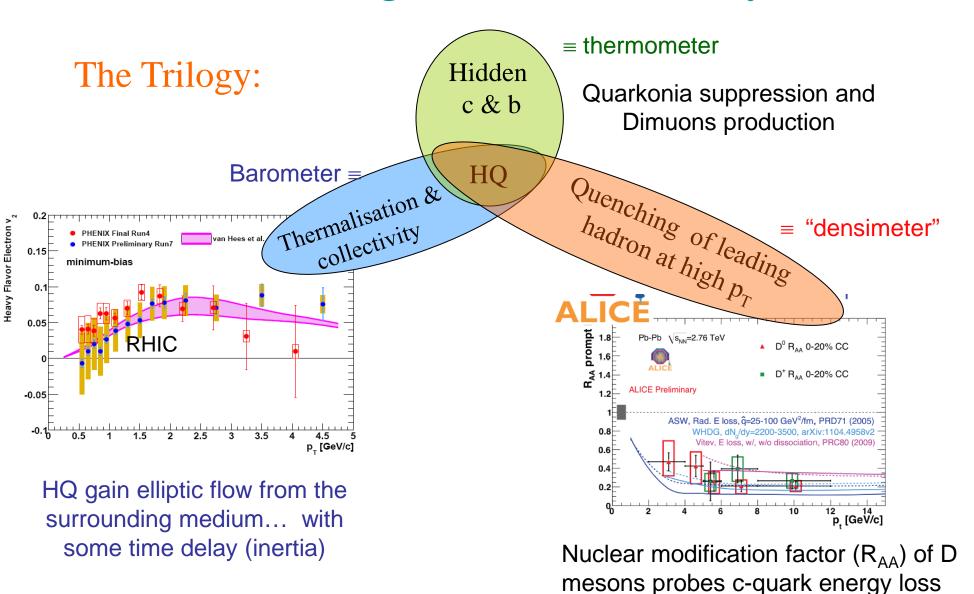
Coherence effects In heavy quark and quarkonium production in ultrarelativistic heavy ion collisions

P.B. Gossiaux (SUBATECH, UMR 6457)

Thanks to J Aichelin, H. Berrehrah, M. Bluhm, Th. Gousset, R Katz, V Marin, M. Narhgang, S. Vogel, K. Werner

2nd International Conference on New Frontiers in Physics (Kolymbari, Greece)

Hard Probing QGP with heavy flavors

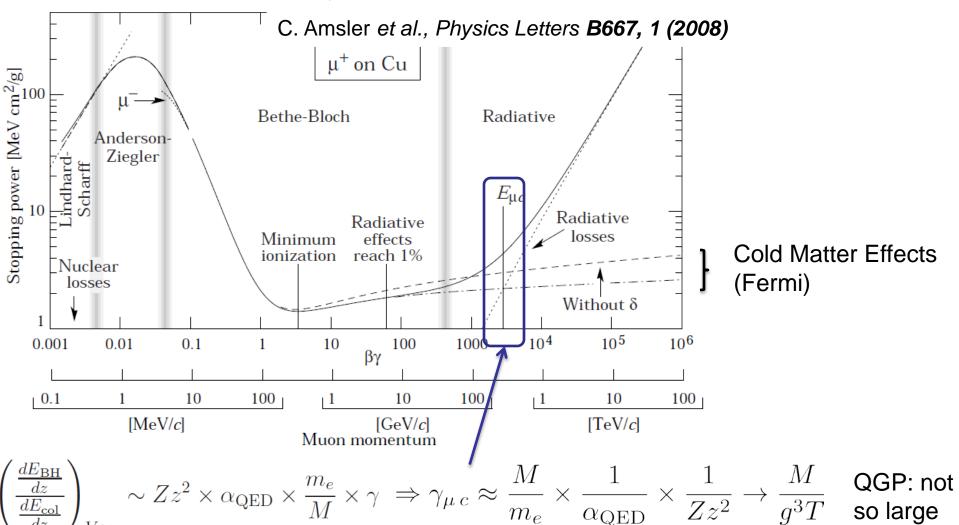


in QGP (not seen in pA)

Heavy flavor quenching

Quenching – Energy loss in cold atomic matter

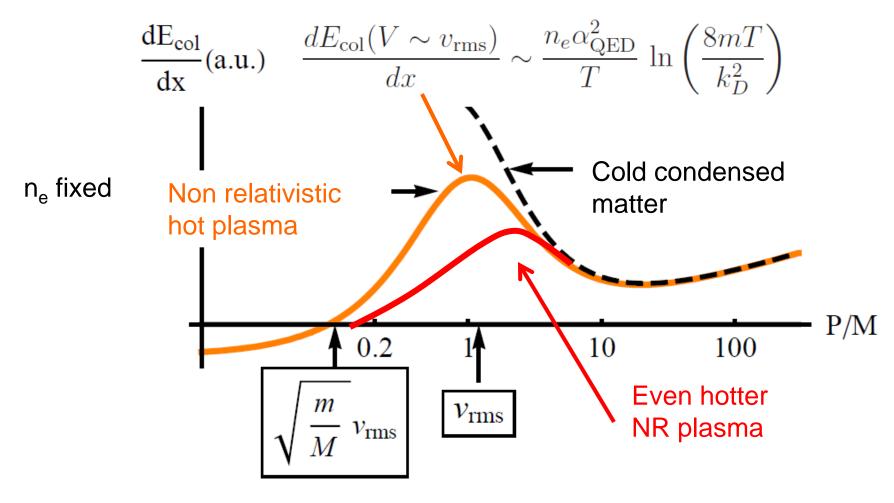
Energy loss of a charged particles passing through cold atomic matter: extensive field of research in the XXth century



4

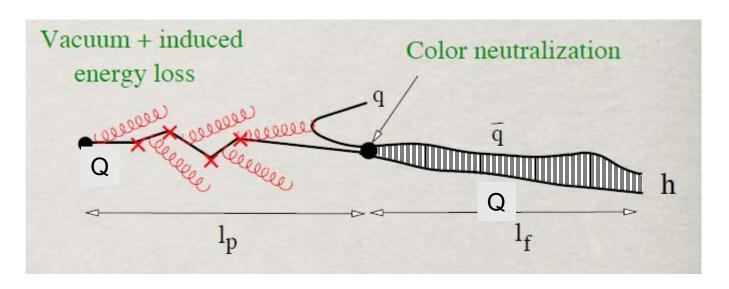
fermion Energy loss in a NR (Q)ED plasma

Reduction of the collisional energy loss! (need to "touch" the plasmon pole: v≈v_{rms} α T^{1/2})



What if T still increases (until m_e) and $v_{rms} \approx 1$?

Partons in QCD plasma



From B. Kopeliovich (this conf)

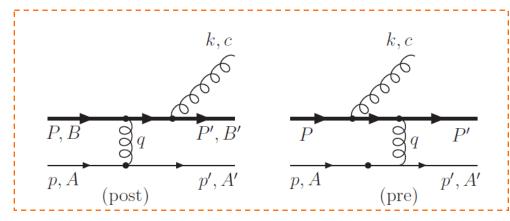
We will concentrate here on the *radiative induced* energy loss, which is the key ingredient of most of the models... (assuming the interactions with the QGP are strong enough to weaken / break the QQbar resonance)

Basic of induced radiation (Gunion-Bertch)

Radiation α deflection of current (semi-classical picture)

Eikonal limit (large E, moderate q)

QED-like



 $\begin{array}{c} k,c \\ \hline \\ P \\ \hline \\ p,A \\ \hline \\ \end{array}$

ω: energy of radiated gluon; x=ω/E

$$\omega \frac{d^3 \sigma_{\rm rad}^{x \ll 1}}{d\omega d^2 k_{\perp} dq_{\perp}^2} = \frac{N_c \alpha_s}{\pi^2} (1 - x) \times \frac{J_{\rm QCD}^2}{\omega^2} \times \frac{d\sigma_{\rm el}^{Qq}}{dq_{\perp}^2}$$

Dominates as small x as one "just" has to scatter off the virtual gluon k'

with

$$\frac{J_{\text{QCD}}^2}{\omega^2} = \left(\frac{\vec{k}_{\perp}}{k_{\perp}^2 + x^2 M^2 + (1 - x) m_g^2} - \frac{\vec{k}_{\perp} - \vec{q}_{\perp}}{\left(\vec{k}_{\perp} - \vec{q}_{\perp}\right)^2 + x_{\perp}^2 M^2 + (1 - x) m_g^2}\right)^2$$
Gluon thermal mass ~2T Quark mass

Both cures the collinear divergences, and have large impact on the radiation spectra

Radiation spectra (incoherent)

$$\omega \frac{d^2 \sigma_{\mathrm{rad}''\mathrm{QCD''}}^{x\ll 1}}{d\omega dq_\perp^2} \approx \frac{2N_c \alpha_s}{\pi} \ln \left(1 + \frac{q_\perp^2}{3\tilde{m}_g^2}\right) \times \frac{d\sigma_{\mathrm{el}}^{Qq}}{dq_\perp^2} \qquad \qquad \text{in to convolute with your favorite elastic cross section}$$

For Coulomb screened (µ) scattering:

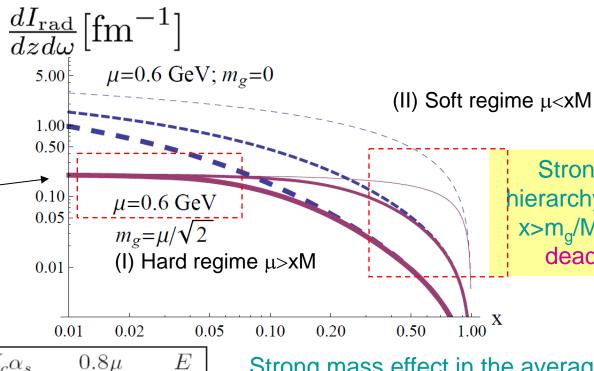
$$\tilde{m}_g^2 = (1 - x)m_g^2 + x^2M^2$$

Light quark

c-quark

b-quark

Little mass dependence for finite "gluon mass" (especially from $q\rightarrow c$)



Strong mass hierarchy effect for $x>m_q/M_Q$ (but no dead cone)

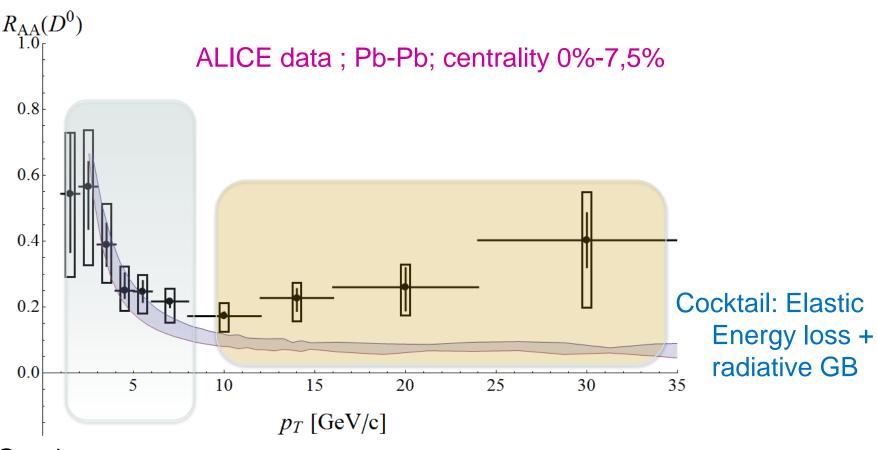


$$\frac{dE_{\rm GB}(Q)}{dz} \approx \frac{4N_c\alpha_s}{\pi} \times \frac{0.8\mu}{M+\mu} \times \frac{E}{\lambda_Q}$$

Strong mass effect in the average Eloss (mostly dominated by region II)

Easily implemented in some MC codes like URQMD, pHSD, BAMPS...

Gunion-Bertch radiation vs data

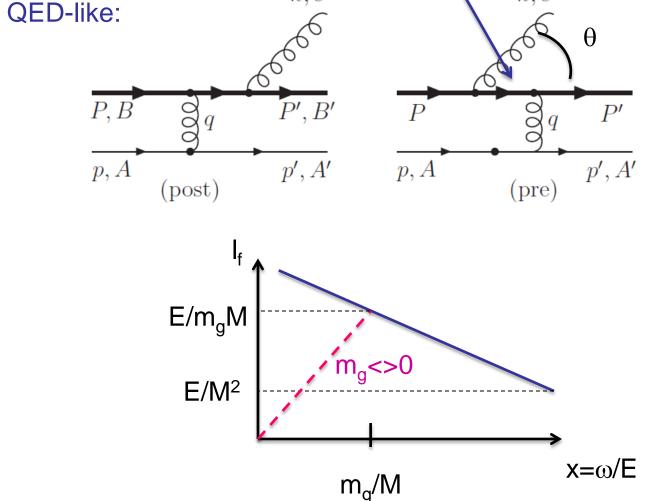


Good agreement at intermediate pT

Increasing disagreement with increasing p_T

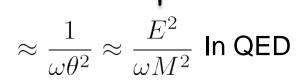
Formation time for a single collision

Formation time extracted from the virtuality of the off shell Heavy Quark



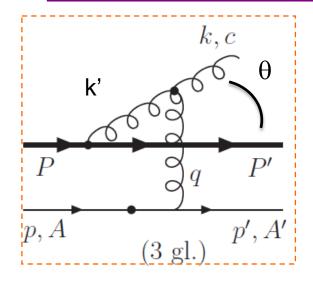
$$l_{f,\mathrm{sing}}^{r} \approx \frac{2(1-x)}{\omega \left(\frac{M^2}{E^2} + (1-x)\frac{m_g^2}{\omega^2} + \theta^2\right)}$$

$$l_{f,\text{sing}}^{post} \approx \frac{2(1-x)}{\omega \left(\frac{M^2}{E^2} + (1-x)\frac{m_g^2}{\omega^2} + \left(\vec{\theta} - \frac{\vec{q}_{\perp}}{E}\right)^2\right)}$$

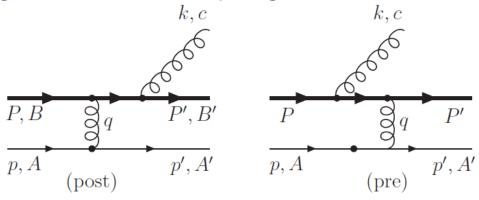


QED: Long formation times for small radiation angles and small frequencies

Formation time for a single coll.



In the genuine QCD, the pre-gluon k' is struck



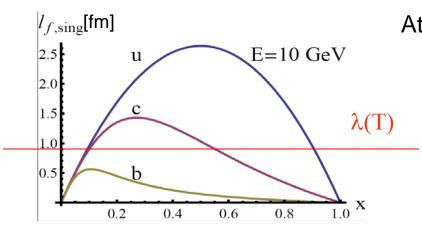
$$l_{f,\text{sing}}^{3 gl} \approx t_f \approx \frac{2(1-x)\omega}{(\vec{k}_{\perp} - \vec{q}_{\perp})^2 + x^2 M^2 + (1-x)m_q^2}$$

Radiation at wider angle; smaller formation times than for the QED-like

For 0 masses: still
$$\approx \frac{1}{\omega \theta^2}$$
 but $\theta = \frac{k_t}{\omega} \approx \frac{\mu}{\omega}$

QCD: Longer and longer formation times for increasing frequencies

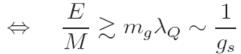
Formation time for a single coll.

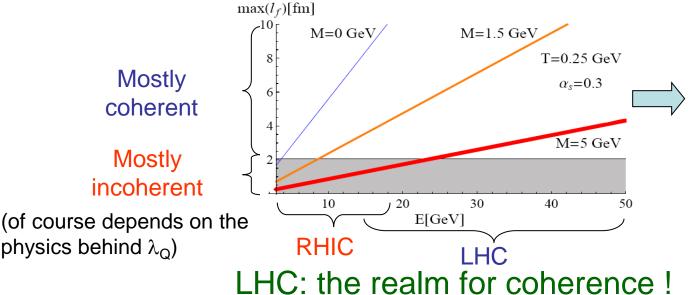


At 0 deflection:

Comparing the formation time (on a single scatterer) with the mean free path:

Coherence effect for HQ gluon radiation :





Coherence effect (equiv. LPM in QED) mandatory for high p_T HQ.

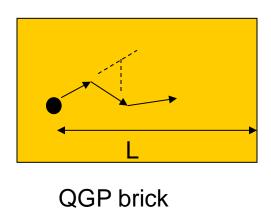
(and even more for high p_T light quark)...

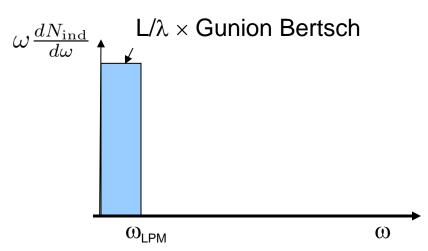
That will mostly affect the radiation pattern at intermediate x

(light q)

Application for radiative energy loss in the eikonal limit

various regimes:



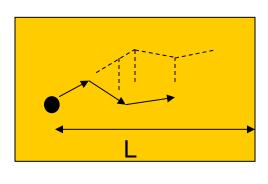


ightarrow a) Low energy gluons: Typical formation time ω/k_t^2 is smaller than mean free path λ : $\omega<\omega_{\rm LPM}:=\frac{\hat{q}\lambda^2}{2}$ Incoherent Gunion-Bertsch radiation

Where $\hat{q}=\frac{\langle \delta q_\perp^2 \rangle}{\lambda}$ (transport coefficient) is the average square momentum increase of the partons per unit time... Very important quantity, in principle calculable from lattice QCD

(light q)

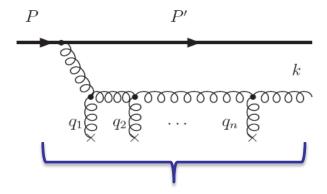
various regimes:



QGP brick

→ b) Inter. energy gluons:

Application for radiative energy loss in the eikonal limit



Production on $N_{\rm coh}$ scatterings => reduction of the GB radiation by a factor 1/ $N_{\rm coh}$

Produced coherenty on $N_{\rm coh}$ centers after typical formation time ${\rm t_f}$ such

 $t_f=rac{\omega}{k_t^2}$ (as usual) but also $k_t^2=\hat{q}t_f$ (stochastic propagation of the gluon)

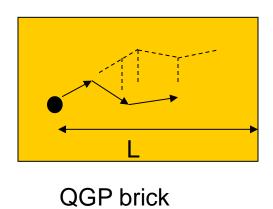
$$\Rightarrow t_f = \sqrt{\frac{\omega}{\hat{q}}} \Rightarrow N_{\rm coh} = \frac{t_f}{\lambda} = \sqrt{\frac{\omega}{\omega_{\rm LPM}}}$$

Multiple formation time

(light q)

Application for radiative energy loss in the eikonal limit

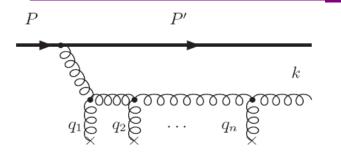
various regimes:



 $\omega \frac{dN_{\rm ind}}{d\omega} \text{ L/λ} \times \text{Gunion Bertsch}$ $\times \sqrt{\frac{\omega_{\rm LPM}}{\omega}} \propto \sqrt{\frac{\hat{q}L^2}{\omega}}$ BDMPS (96-00); pattern inverted wrt LPM (gluon charged) $\omega_{\rm LPM}$

- a) Low energy gluons: Typical formation time ω/k_t^2 is smaller than mean free path λ : $\omega < \omega_{\text{LPM}} := \frac{\hat{q}\lambda^2}{2}$ Incoherent Gunion-Bertsch radiation
- ightarrow b) Inter. energy gluons: Produced **coherenty** on N_{coh} centers after typical formation time $t_f = \sqrt{\frac{\omega}{\hat{q}}} \Rightarrow N_{\mathrm{coh}} = \frac{t_f}{\lambda} = \sqrt{\frac{\omega}{\omega_{\mathrm{LPM}}}}$ leading to an effective reduction of the GB radiation spectrum by a factor $1/N_{\mathrm{coh}}$

Formation time in a random walk



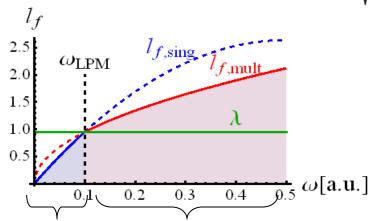


Phase shift at each collision

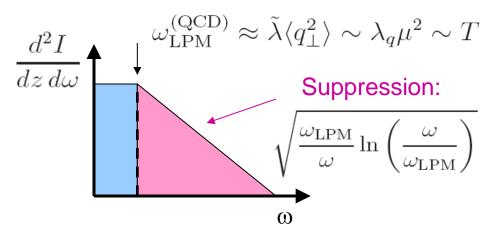
Following Landau-Pommeranchuk: one obtains an effective formation time by imposing the cumulative phase shift to be $\Phi_{\rm dec}$ of the order of unity

For light quark (infinite matter):

$$l_{f,\mathrm{mult}}(q+g) = l_{f,\mathrm{scat}}(q+g) \approx 2\sqrt{\frac{\omega\Phi_{\mathrm{dec}}}{\hat{q}}} \implies$$
 3 scales: $l_{f,\mathrm{mult}}$, $l_{f,\mathrm{sing}}$ & λ



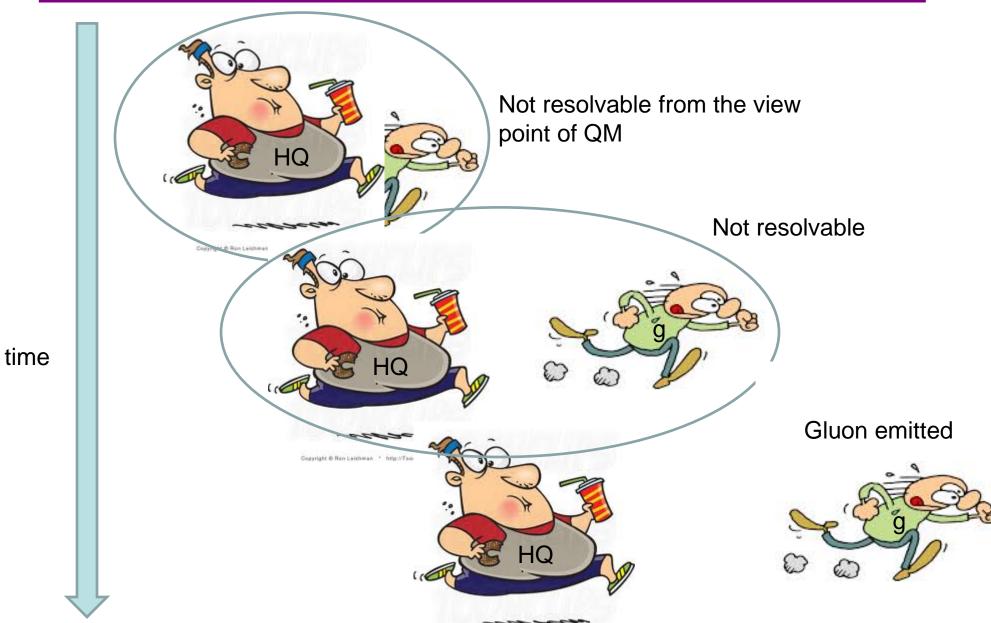




Especially important for av. energy loss

$$\frac{dE_{\rm BDMPS}(q)}{dz} \sim \sqrt{\frac{\omega_{\rm LPM}}{E}} \times \frac{dE_{\rm GB}(q)}{dz}$$

Gluon emission from HQ



Formation time and decoherence for HQ

$$l_{f,\text{mult}}(Q+g) = \frac{2\omega\Phi_{\text{dec}}}{\sqrt{\omega\hat{q}\Phi_{\text{dec}} + \left(\frac{M^2\omega^2}{2E^2}\right)^2 + \frac{M^2\omega^2}{2E^2}}}$$

"Competition" between

$$ullet$$
 decoherence" due to the masses: $m_g^2 + x^2 M^2$

• decoherence due to the transverse kicks $\langle Q_{\perp}^2 \rangle = l_{f, \mathrm{mult}} \, \hat{q}$

Special case:
$$\lambda < l_{f, \text{mult}} < L_{\text{QCD}}^{\star\star} := \frac{m_g^2 + x^2 M^2}{\hat{q}}$$

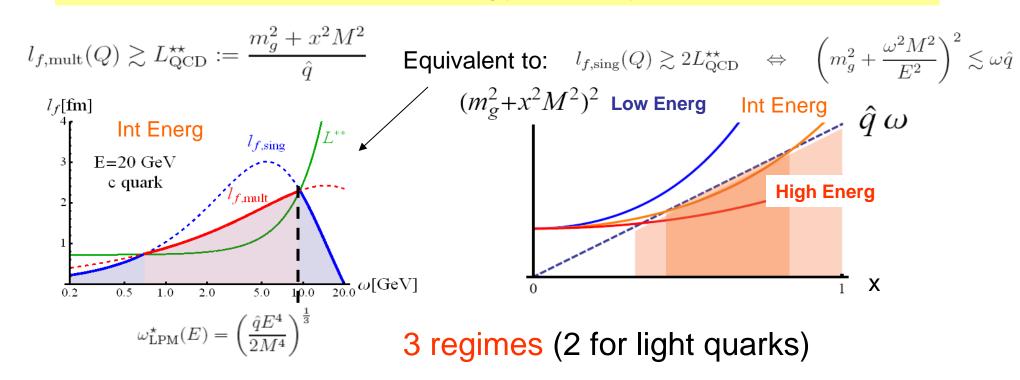
One has a possibly large coherence number $N_{coh} := I_{f,mult}/\lambda$ but the radiation spectrum per unit length stays mostly unaffected:

Radiation on an effective center of length
$$I_{f,mult} = N_{coh} \lambda$$
 Radiation at small angle $\alpha \langle Q_{\perp}^2 \rangle$ i.e. αN_{coh} Compensation at leading order!

LESSON: HQ radiate less, on shorter times scales and are less affected by coherence effects than light ones!!! (dominance of 1rst order in opacity expansion)

Formation time and decoherence for HQ

Criteria: HQ radiative E loss strongly affected by coherence provided:



Low energy: radiation from HQ unaffected by coherence

Intermediate energy:
coherence affects radiation on
an increasing part of the
spectrum (up to ω_{LPM}*)

$$E_{\text{NO-LPM}}^{\star} := 3 \frac{M m_g^3}{\hat{q}} \sim \frac{M}{g_s}$$

High energy: HQ behaves like a light one; coherence affects radiation from ω_{LPM} on.

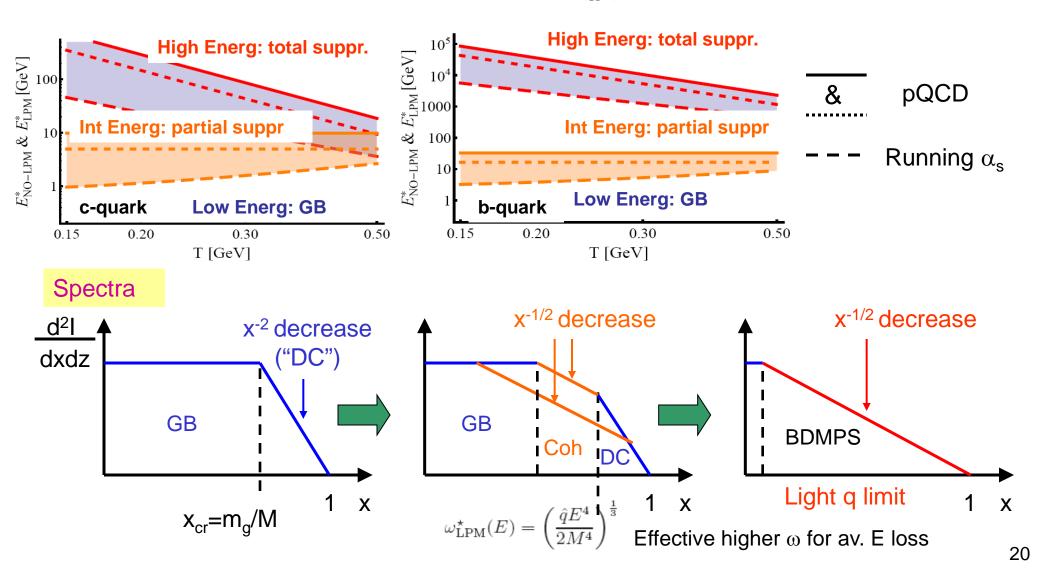
$$E_{\text{LPM}}^{\star} := \frac{M^4}{\hat{q}}$$

Regimes and radiation spectra

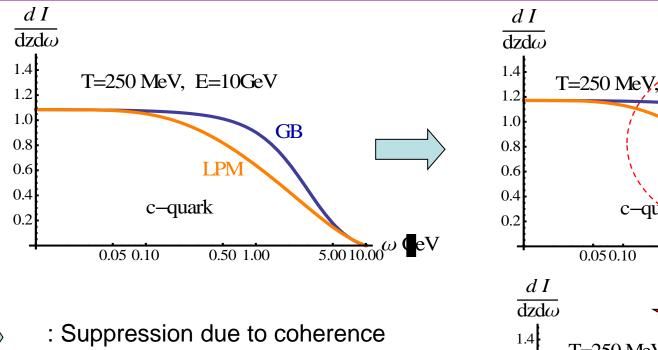
Hierarchy of scales:

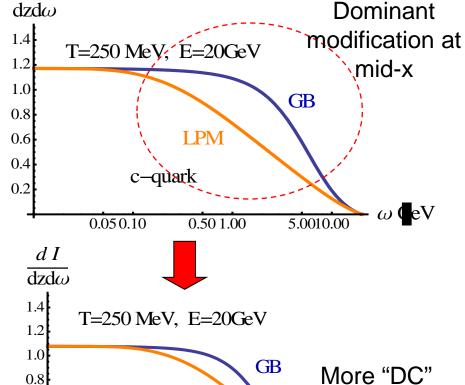
$$\underbrace{E_{\mathrm{LPM}}(q)}_{T} \ll \underbrace{E_{\mathrm{NO-LPM}}^{\star}(Q)}_{\frac{M}{g_s T} \times T} \ll \underbrace{E_{\mathrm{LPM}}^{\star}(Q)}_{\left(\frac{M}{g_s T}\right)^4 \times T}$$

larger coupling ⇒ Larger coherence effects



Reduced spectra from coherence in particular model





LPM

0.50 1.00

b-quark

0.050.10

increases with increasing energy



: Suppression due to coherence decreases with increasing mass

Quantum coherence: very difficult to implement in MC/hydro codes

In (first) Monte Carlo implementation: we quench the probability of gluon radiation by the ratio of coherent spectrum / GB spectrum

0.6

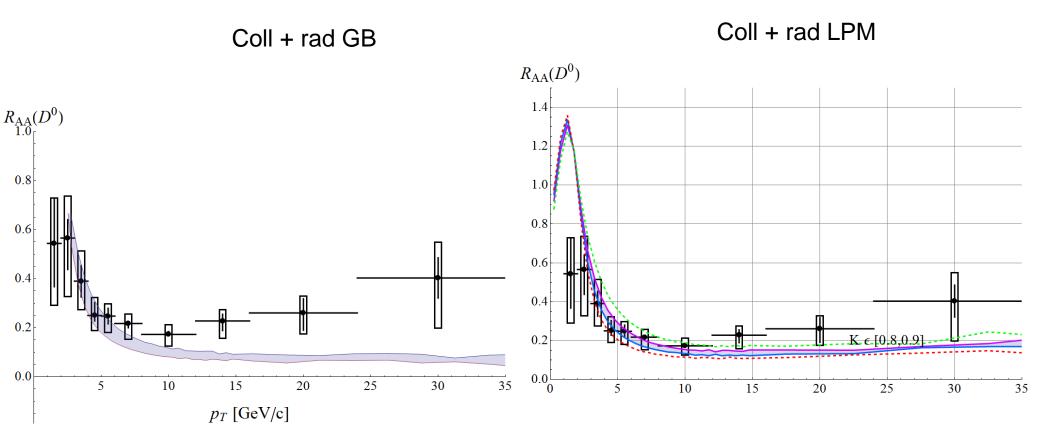
0.4

0.2

effect

5.0010.00

D mesons at LHC (vs ALICE 0%-7.5%)



Part of the disagreement cured by the introduction of such coherence effects... still some room for improvement:

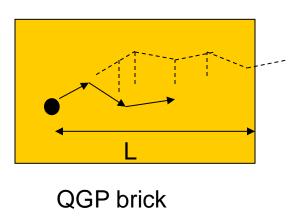
A) Finite Path length effects

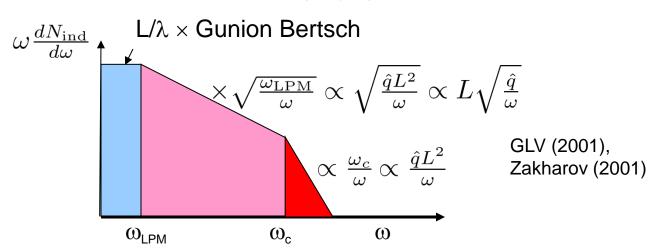
B) Other effects

(light q)

Application for radiative energy loss in the eikonal limit

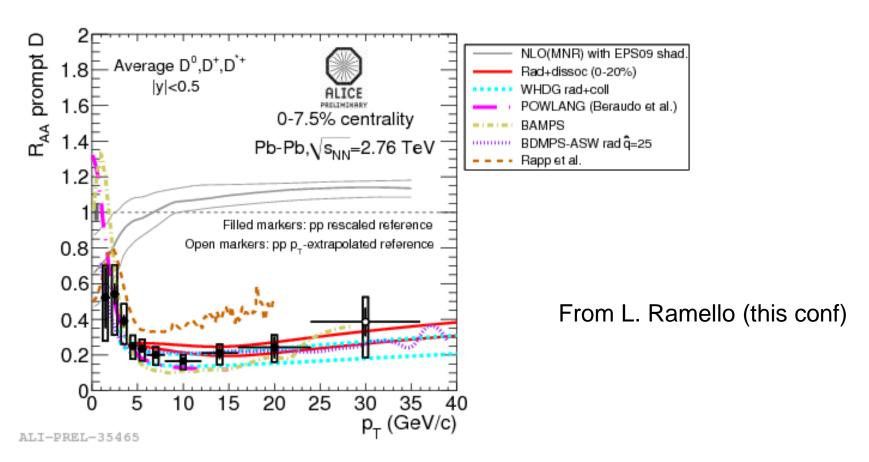
finite path length:





- a) Low energy gluons: Incoherent Gunion-Bertsch radiation
- b) Inter. energy gluons: Produced **coherenty** on $N_{\rm coh}$ centers after typical formation time $t_f=\sqrt{\frac{\omega}{\hat{q}}}$
- ightharpoonup c) High energy gluons: Produced mostly outside the QGP... nearly as in vacuum do $\sqrt{\frac{\omega}{\hat{q}}} > L \Rightarrow \omega > \omega_c := \frac{\hat{q}L^2}{2}$ not contribute significantly to the induced energy loss
 - => Average Energy loss along the path way: $\langle \Delta E \rangle \sim \hat{q} L^2 \ln \left(\frac{E}{\omega_c} \right)$ often the only result retained₂₃

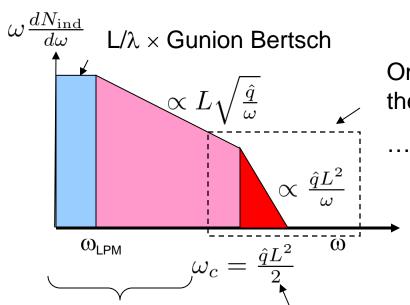
Model vs Experiment (3rd round)



Most of the models based on energy loss mechanism which explain the quenching reduction at large p_T include those finite path length effects... but the counter part is that they do not include proper medium evolution

(light q)

Application for radiative energy loss in the eikonal limit



Only this tail makes the L² dependence in the average Eloss integral ...

...provided the higher boundary $\omega = E > \omega_c$.

Otherwise, everything α L

Bulk part of the spectrum still scales like path length L

Concrete values @ LHC $\frac{\hat{q} \sim 25 {
m GeV}^2/{
m fm}}{L \sim 2 {
m fm}}$

 $\omega_c \sim 500 {
m GeV}$ Huge value !

Personal opinion: before looking on coherence effects on large distances

(5-10 fm/c) let us make sure nothing was left over!

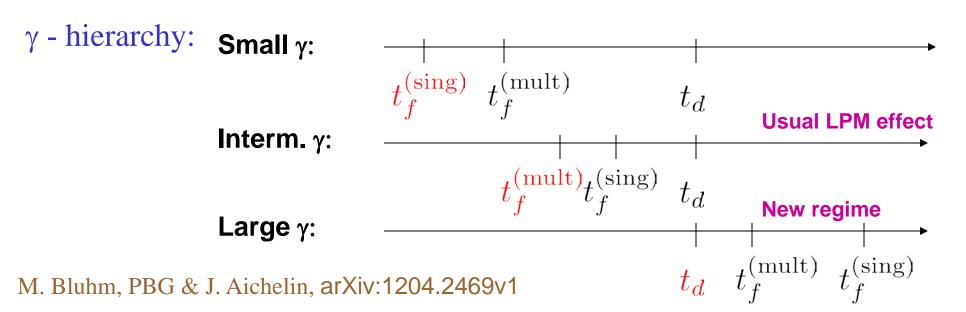
Consequences of radiation damping on energy loss

Basic question: Implications of a finite lifetime of the radiated gluon?

Litttle attention in the litterature (V. M. Galitsky and I. I. Gurevich, Il Nuovo Cimento **32** (1964) 396 for classical electrodynamics).

Concepts

- In QED or pQCD, damping is a NLO process (damping time $t_d >> \lambda$); neglected up to now.
- \blacktriangleright However: formation time of radiation t_f increases with boost factor γ of the charge
- Expected effects when $t_f \approx t_d$ or $t_f > t_d$: in this regime, t_d should become the relevant scale (gluons absorbed while being formed)



Modification of the LPM effect due to radiation damping?

Naïve thoughts (bets) about the consequences of photon damping

a) Relaxed attitude: "Nothing special happens to the Work, as photons are absorbed after being emitted"



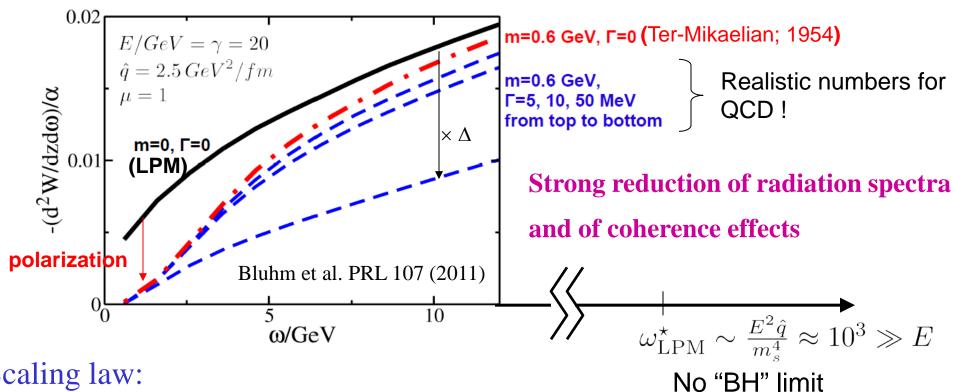
- b) Vampirish though: as the medium "sucks" the emitted photons, the charge will have a tendency to emit more of them => increased energy loss
- c) Less energy loss (find the argument)

Beware: we are not speaking of the radiated energy in the far distance (always reduced) but on the impact on the radiating parton



Consequences of radiation damping on energy loss

PRL 107 (2011): Revisiting LPM effect in ED using complex index of refraction, focussing on the radiation at time of formation $n^2(\omega) = 1 - m^2/\omega^2 + 2i\Gamma/\omega$



Scaling law:

$$\frac{\frac{dN}{d\omega}}{\frac{dN_{\mathrm{sing}}}{d\omega}} pprox \frac{\min(t_d, t_f^{\mathrm{(sing)}}, t_f^{\mathrm{(mult)}})}{t_f^{\mathrm{sing}}}$$

Allows for first phenomenological study in QCD case

Formation time of radiated gluon

$$(1-x, \vec{P}_{\perp}, m_s)$$

$$(E, \vec{P}_{\perp} + \vec{k}_{\perp})$$
 Final Emitter State
$$(x, \vec{k}_{\perp}, m_g)$$
 Emitter gluon

Final HQ Emitted

$$t_f \left[\frac{\langle p_B^2 \rangle + x^2 m_s^2 + (1 - x) m_g^2}{2x(1 - x)E} \right] \simeq 1$$

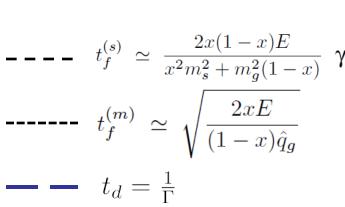
$$p_B^2 := \left((1-x) \vec{k}_\perp + x \vec{P}_\perp \right)^2 \quad \Rightarrow \quad \langle p_B^2 \rangle \approx (1-x)^2 \hat{q}_g t_f \stackrel{\text{In QCD: mostly gluon}}{\text{rescattering}}$$

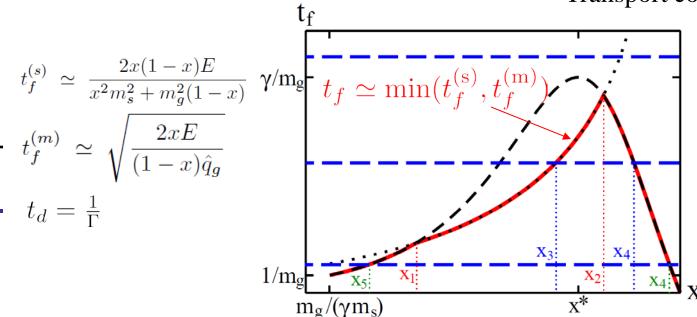
$$\Rightarrow$$

$$\langle p_B^2 \rangle \approx (1-x)^2$$

=> Self consistent expression for t_f

Transport coefficient: [GeV²/fm]



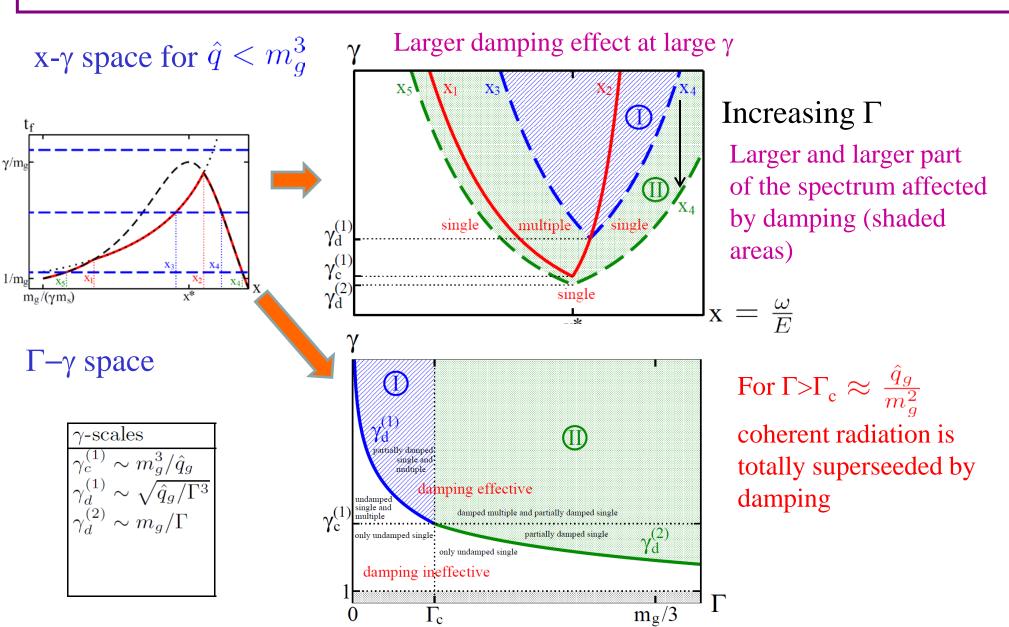


Small Γ

Interm. Γ

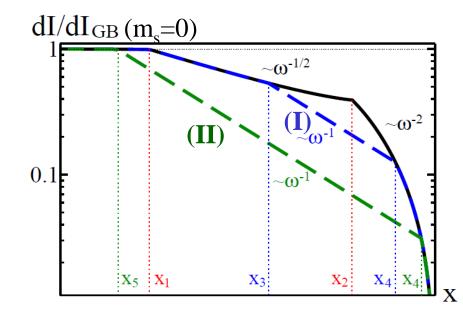
Large Γ

New regimes when including gluon damping



Consequences on the power spectra

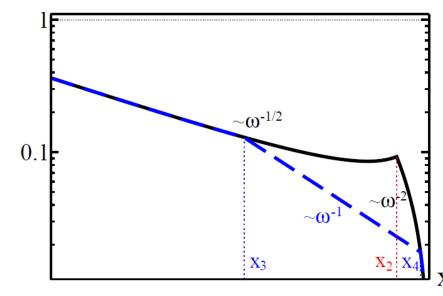
$$\hat{q} < m_g^3$$



(I) and (II): moderate and large damping (see previous slide)

$$\begin{split} &E{=}~45~\text{GeV},\,m_s{=}1.5~\text{GeV}\\ &m_g{=}0.6~\text{GeV},\,\hat{q}\,=\,0.1\text{GeV}^2/\text{fm}\\ &\Gamma{=}0.05~\text{GeV}~\text{(I)}~\&~0.15~\text{GeV}~\text{(II)} \end{split}$$



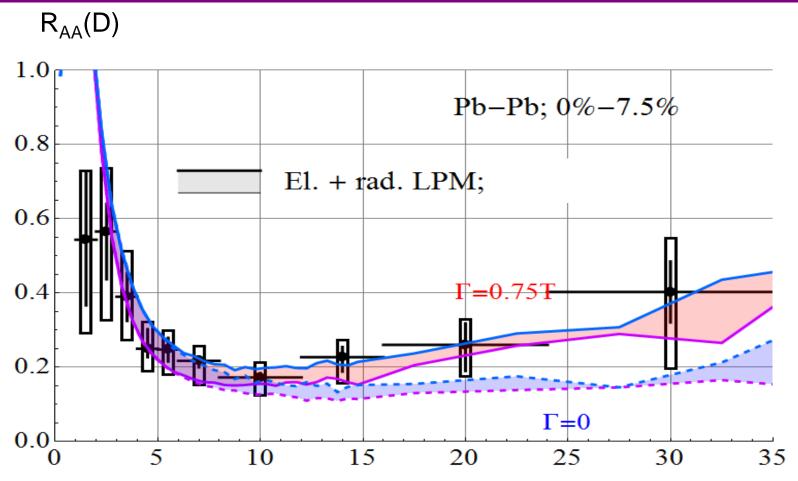


Same but

$$\hat{q} = 2 \text{GeV}^2 / \text{fm}$$

Γ=0.25 GeV

Consequences on the HQ observables



Damping of radiated gluons reduces the quenching of D mesons and allows reproducing their R_{AA}

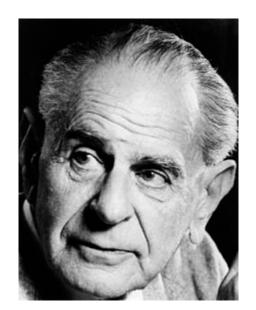
Damping vs Finite Path Length



Quite generically, damping effects dominate over path length effects if

$$\Gamma_d > \frac{\hbar c}{L} \approx 50 \text{ MeV}$$

Realistic scale in strongly coupled système ($\Gamma_d = O(g^n) T$)



Falsifiability: a)path length dependence still α L with damping effects, while α L² with usual BDMPS argument or b) γ -D/B correlations

« turn LHC into a precision tool »... not only for Higgs and SUSY

Quarkonia production in dynamical QGP

Work in progress

Probing deconfinement?

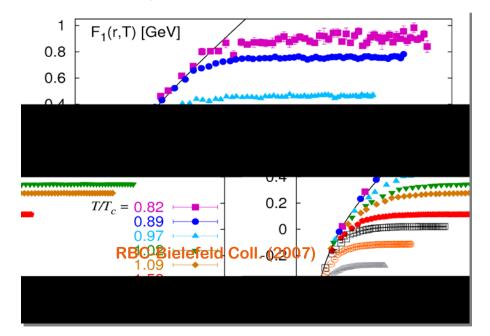
How can we prove that we have really achieved a *deconfined* state of matter in ultra-relativistic heavy ions collisions?

Challenge

"deconfinometer" =

- Color fluctuations
- Propagation of individual quarks over large distances

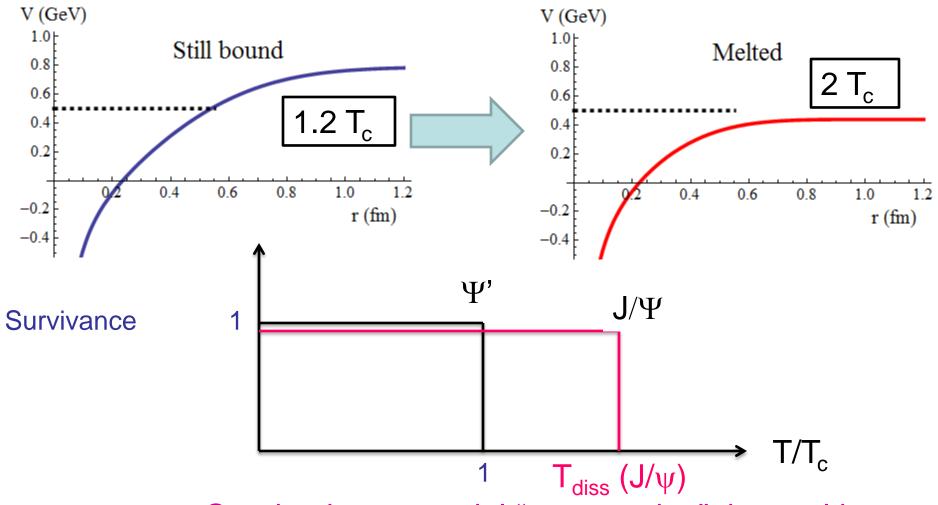
Looking at the QQbar potential on the lattice



Increased screening at larger temperatures

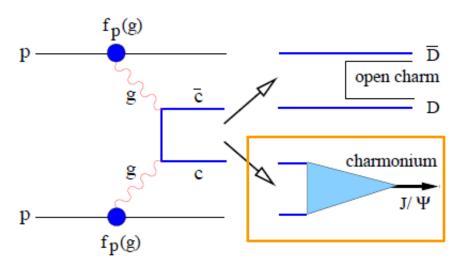
Quarkonia in Stationary QGP

Consequence for Q-Qbar states (Q: heavy quark):



Best candidate: Quarkonia sequential "suppression", i.e. melting and/or dissociation (Matsui & Satz 86)

Dynamical version of the sequential suppression scenario



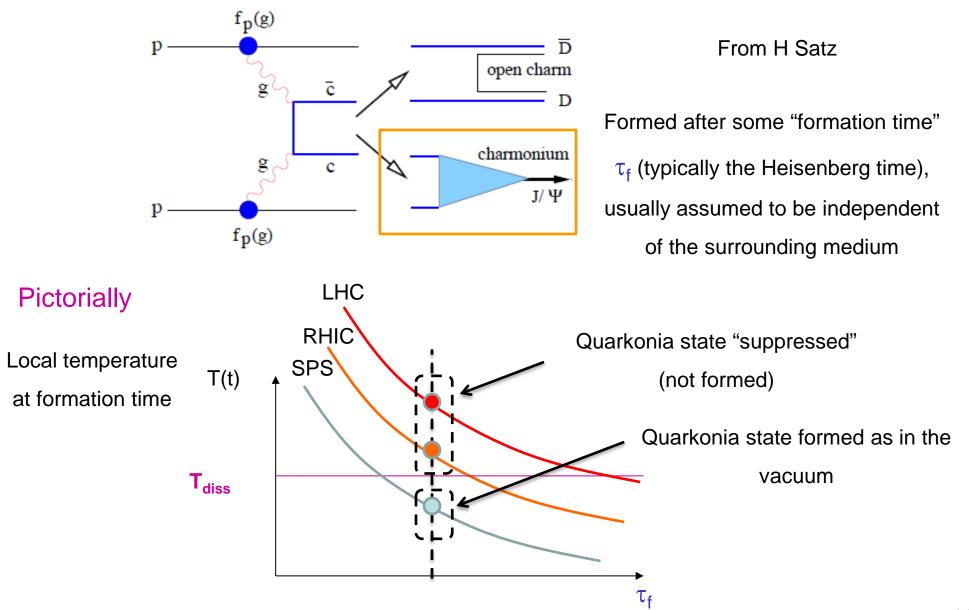
From H Satz

Formed after some "formation time" $\tau_f \mbox{ (typically the Heisenberg time),} \\ \mbox{usually assumed to be independent} \\ \mbox{ of the surrounding medium}$

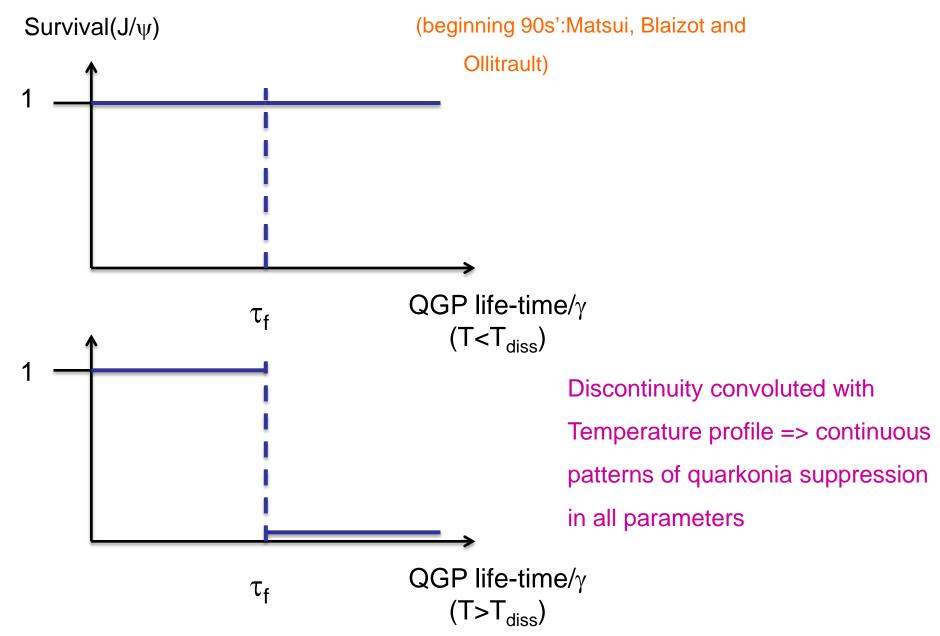
Standard folklore of sequential suppression: The quarkonia which should be formed at (τ_f, x_0) is not if $T(\tau_f, x_0) > T_{diss} = > Q$ -Qbar pair is "lost" for quarkonia formation

Need to know the *formation times* as well in order to have a predictive scheme (not so obvious, especially for the upsilons, which are produced during the very early stage of the nucleus-nucleus collision)

Dynamical version of the sequential suppression scenario



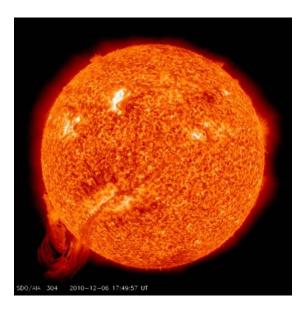
Dynamical version of the sequential suppression scenario



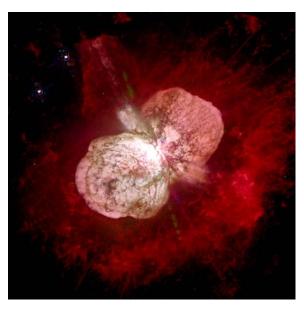
Caviats & Uncertainties

What does the sequential suppression in a stationary QGP has to do with reality anyhow?





Picture

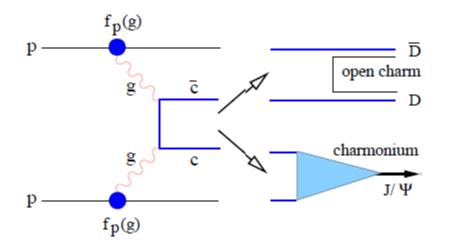


Reality

Need for a genuine time-dependent scenario

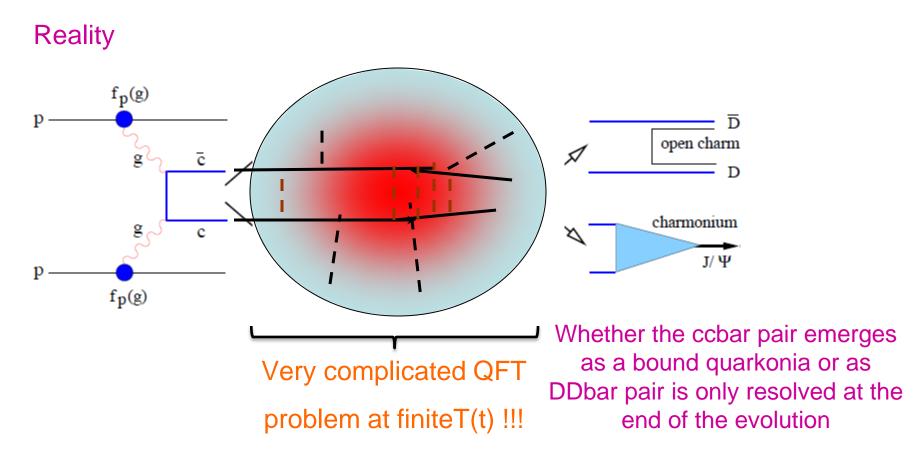
Beyond the (quasi-stationnary) sequential suppression scenario

Picture



Early decoupling between various states

Beyond the (quasi-stationnary) sequential suppression scenario



But one should aim at solving it, especially as the quarkonia content of a QQbar quantum state is at most of the order of a few % (continuous transitions under external perturbations)

1rst Quantum approach

• Time-dependent Schrödinger equation for the QQ pair

Where

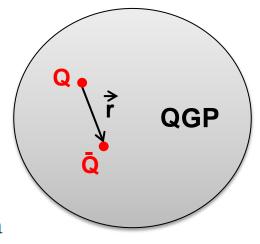
$$\widehat{H} = 2m_q - \frac{(\hbar c)^2}{m_q} \nabla^2 + V(r, T_{\text{red}})$$

$$\Psi_{Q\bar{Q}}(\mathbf{r},t) = R_{Q\bar{Q}}(r,t) \times Y_{Q\bar{Q}}(\theta,\phi)$$

Initial

Initial wavefunction:
$$R_{Q\bar{Q}}\left(r,t=0\right) = \left(\frac{1}{\pi a^2}\right)^{3/4} e^{-\frac{r^2}{2a^2}}$$

where $a_{c\bar{c}}=0.165~\mathrm{fm}$ and $a_{b\bar{b}}=0.045~\mathrm{fm}$

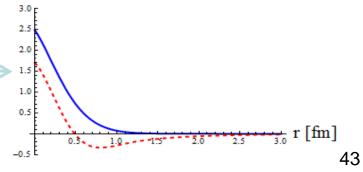


Projection onto the S states: the S weights

$$W_S(t) = \left(4\pi \operatorname{Abs}\left[\int_0^\infty R_{Q\bar{Q}}\left(r, t, T_{red}\right) \times R_S(r, T_{red}^{had}) r^2 dr\right]\right)^2$$

Charmonium radial S states

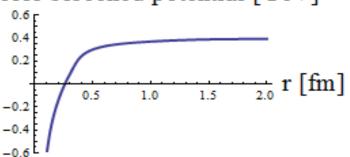
Radial eigenstates of the hamiltonian



Additional ingredients

The color potentials V(Tred, r)

Color screened potential [GeV]



"Weak potential F<V<U"

$$V_{\text{close}}(r) = -\frac{\alpha}{r} + \sigma r - \frac{0.8\sigma}{m^2 r},$$

$$V_{\text{int}}(r) = \frac{V_0 + g_1(r - r_0) + g_2(r - r_0)^2}{1 + g_3(r - r_0) + g_4(r - r_0)^2},$$

$$V_{\text{far}}(r) = V_{\infty} - \frac{4}{3} \frac{\alpha_1}{r} e^{-\sqrt{4\pi} \tilde{\alpha}_1} Tr$$

"Strong potential V=U"

$$U = \left(-\frac{\alpha}{r} + \sigma r - \frac{0.8\sigma}{m^2 r}\right) \times e^{-(\mu r)^2}$$
$$+V_0 \times \left(1 - e^{-(\mu r)^2}\right)$$

Evaluated by Mócsy & Petreczky* and Kaczmarek & Zantow** from IQCD

Additional ingredients

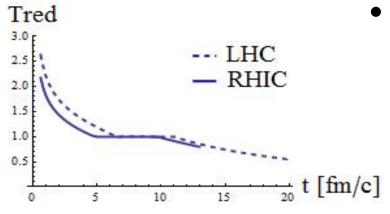
The temperature scenarios

• At fixed temperatures

$$T_{red} = T/T_c,$$
 where $T_c = 0.165~{
m Gev}$

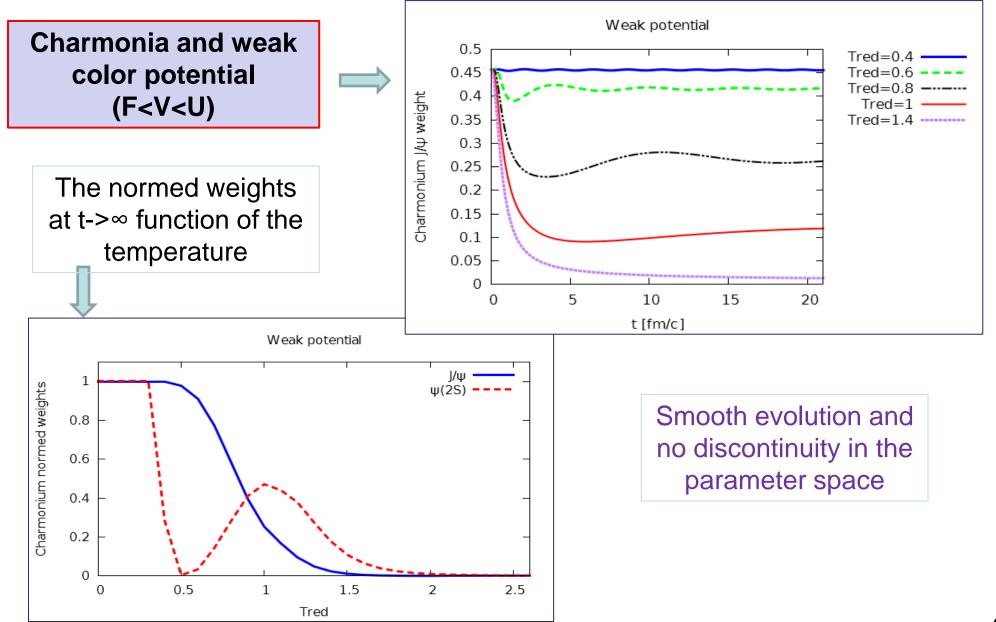


Instantaneous transition from QGP (T_{red}) to hadronisation phase $T_{red}^{had} \leq 0.4$

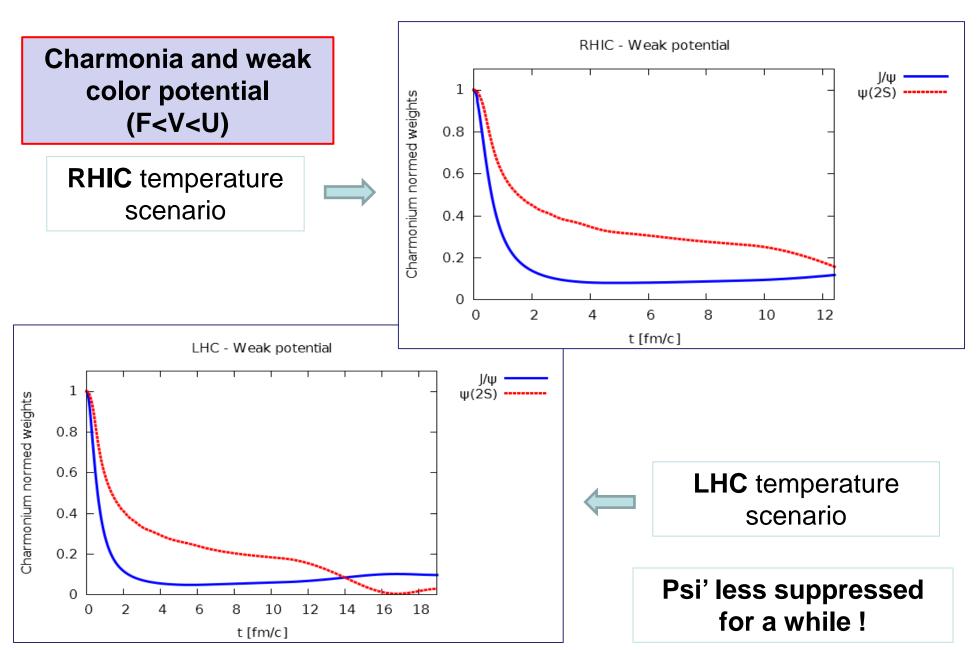


- Time dependent temperature
 - Cooling of the QGP over time by Kolb and Heinz* (hydrodynamic evolution and entropy conservation)
 - > At LHC ($\sqrt{s_{NN}}=2.76~{\rm TeV}$) and RHIC ($\sqrt{s_{NN}}=200~{\rm GeV}$) energies

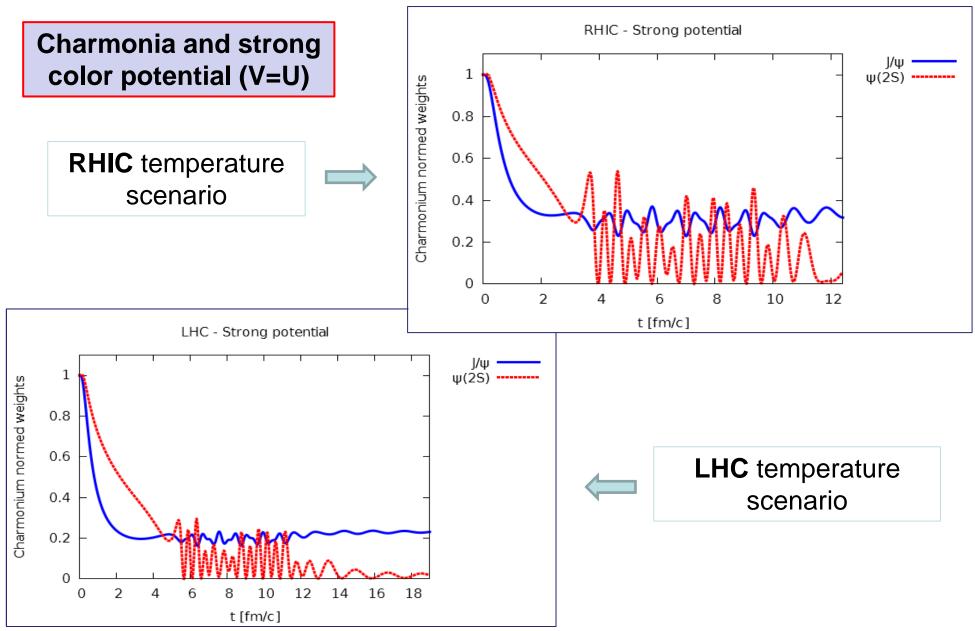
Evolution at fixed temperature



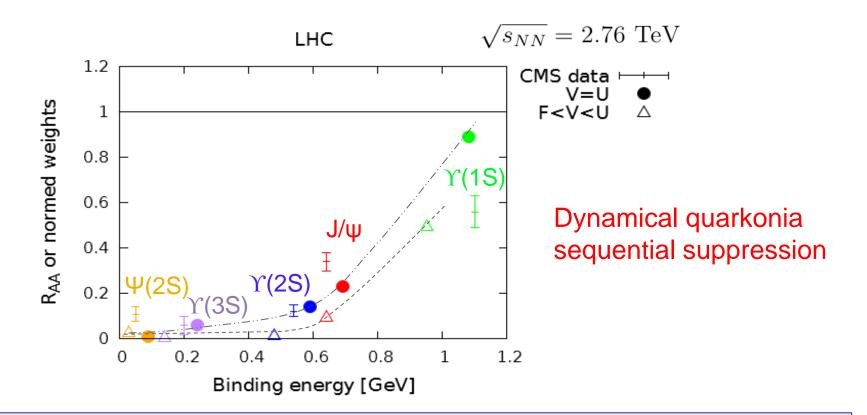
Evolution in realistic T scenarios



Evolution in realistic T scenarios

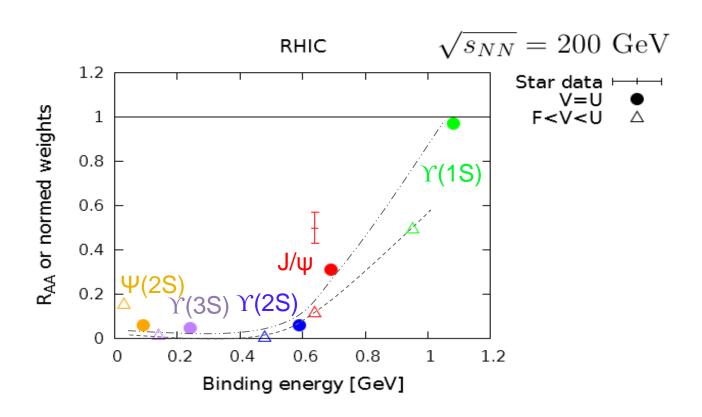


Sum up of LHC results



- The results are quite encouraging for such a simple scenario!
- J/ ψ and ψ (2S) are underestimated (room for regeneration) and Υ (1S) overestimated
- Feed downs from exited states and CNM to be implemented

Sum up of RHIC results



- Similar suppression trends obtained for both RHIC and LHC.
- Less J/ψ suppression at RHIC than at LHC.
- $\Upsilon(1S+2S+3S)$ suppression can be estimated with Star data to ~ 0.55±0.10, we obtain ~ 0.48 for V=U and ~ 0.24 for F<V<U.

A taste of quantum thermalisation

Background?

- RHIC and LHC experimental results => quarkonia thermalise partially in the QGP
- But how to thermalise our wavefunction? Quantum friction/stochastic effects have been a long standing problem because of their irreversible nature.

The open quantum approach: Considering the whole system, quarkonia and environment, the latter

being finally integrating out

Y. Akamatsu [arXiv:1209.5068] Laine et al. JHEP 0703 (2007) 054 2nd possible approach:



Unravel the open quantum approach by using a stochastic operator and a dissipative non-linear potential

- A. Rothkopf et al. Phys. Rev. D 85, 105011 (2012)
- N. Borghini et al. Eur. Phys. J. C 72 (2012)
- S. Garashchuk et al. Jou. of Chem. Phys. 138, 054107 (2013)

New Schrödinger equation

$$i\hbar\frac{\partial\Psi_{Q\bar{Q}}(\mathbf{r},t)}{\partial t} = \left(\widehat{H}(\mathbf{r}) - \underline{\mathbf{F}(t).\mathbf{r}} + \underline{A\big(S(\mathbf{r},t) - \langle S(\mathbf{r},t)\rangle_{\mathbf{r}}\big)}\right)\Psi_{Q\bar{Q}}(\mathbf{r},t)$$
 Friction

Where: $S(\mathbf{r},t) = \arg(\Psi_{Q\bar{Q}}(\mathbf{r},t))$ and $\langle \mathbf{F}(t) \rangle = 0$, $\langle \mathbf{F}(t)\mathbf{F}(t') \rangle = \Gamma(t,t')$

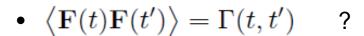
Model for a stochastic operator

• The hierarchy
$$m\gg T \ \Rightarrow \ \sigma\ll au_{\mathrm{relax}}$$
 (adiabatic invariance)

QGP

where

- $\checkmark \sigma$ is the quarkonia <u>autocorrelation time</u> with the gluonic fields (if $\sigma = 0$ the fluctuations are uncorrelated)
- \checkmark $\tau_{\rm relax}$ is the quarkonia <u>relaxation time</u>



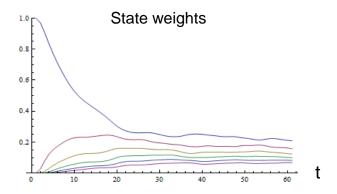


$$\Gamma(t,t') = B \; \frac{e^{-\frac{(t-t')^2}{2\sigma^2}}}{\sqrt{2\pi} \; \sigma} \; \xrightarrow[\sigma \to 0]{} \; B \, \delta(t-t')$$

 One has finally 3 parameters: A (the Drag coefficient), B (the diffusion coefficient) and σ.

First tests of stochastic Schroedinger equation Towards asymptotic distribution?

• Tested in an <u>harmonic potential:</u>

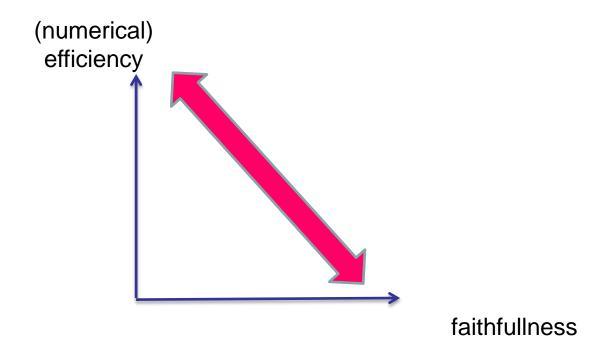


- One gets Boltzmann distributed state weights! Independently of σ and with the Einstein relation B \approx 2mT A between the diffusion coefficient and the Drag coefficient. At a finite time:
- \rightarrow high pt => high velocity => smaller σ => more excited states => more suppression
- low pt => small velocity => higher σ => less excited states => less suppression (=> no need for regeneration ?)



Will be generalized and used to our quarkonia thermalisation in the near future!

Conclusion: The new frontiers of my small world

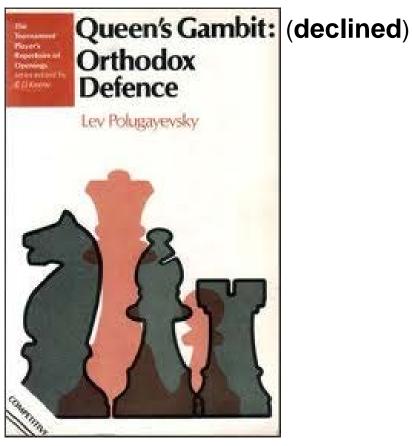


- How to implement reliable energy loss modeling that respect quantum coherence on large scales as well as medium evolution?
- How to implement the quantum evolution of a 2-body system in a dense colored stochastic environment where the concept of cross-section is meaningless



When I was (a lot younger)

Which (for the knowledgeable) is the



The most boring defence ever! ... and I must confess I developed bad feelings with the word "orthodox"

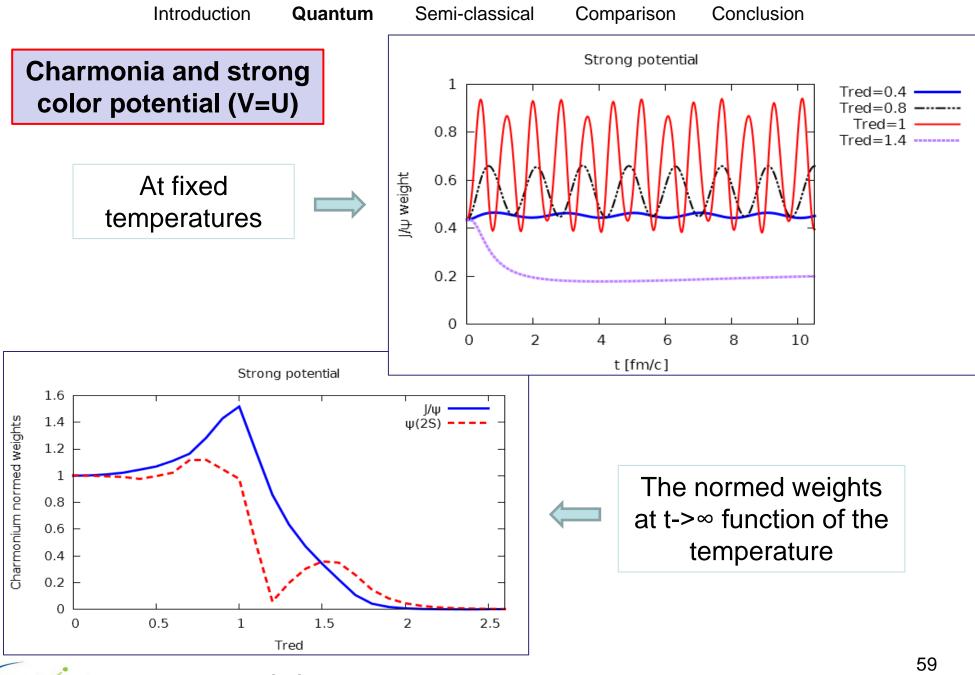
Today

I was pleased to accept the invitation of 3 queens in this orthodox academy of Crete



... and it was much more interesting than playing the orthodox defence!

Back up



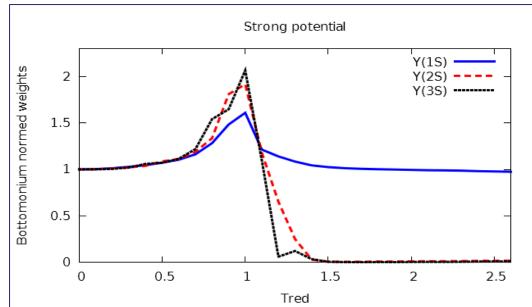


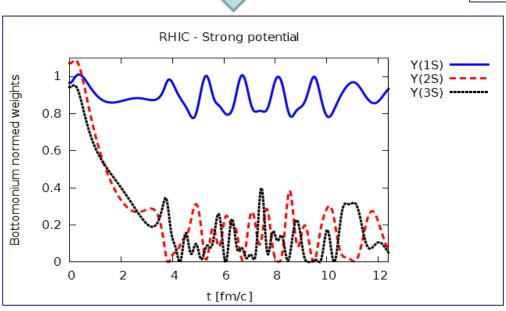
Bottomonia and strong color potential (V=U)

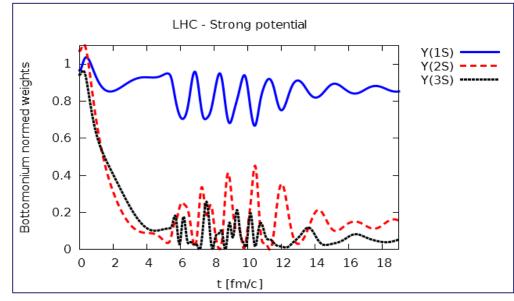
The normed weights at t->∞ function of the temperature

Temperature scenarios





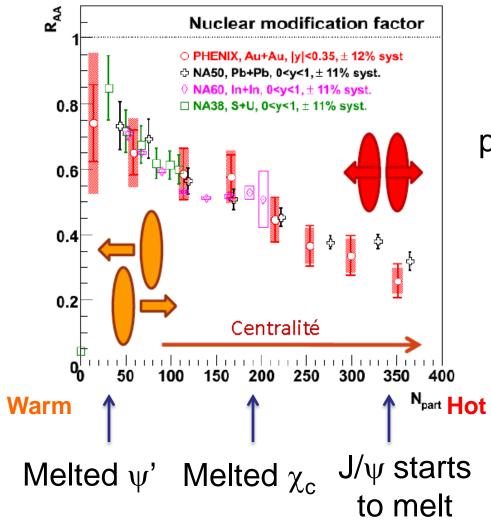






Quarkonia in Stationary QGP





Observed J/ ψ = prompt J/ ψ + 30% χ_c + 10 % ψ

No further suppression at RHIC (as compared to SPS)

=> Claim that T_{diss} (J/ ψ) is pretty high (strongly bound)

Quarkonia in Stationary QGP

 $\begin{array}{c|c}
2 & - \\
- & \chi_b(1P) \\
\hline
1.2 & J/\psi(1S) \\
\Upsilon'(2S)
\end{array}$ "robust" states Thermometer

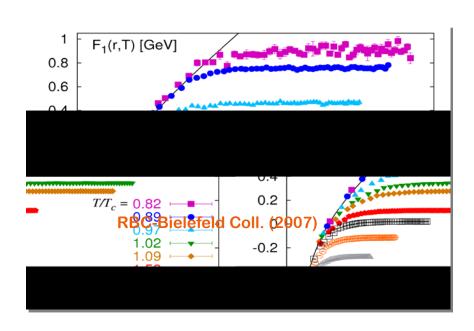
Indeed observed at SPS (CERN) and RHIC (BNL) experiments. However:

- alternative explanations, lots of unknown (also from theory side)
- less suppression at LHC

Caviats & Uncertainties

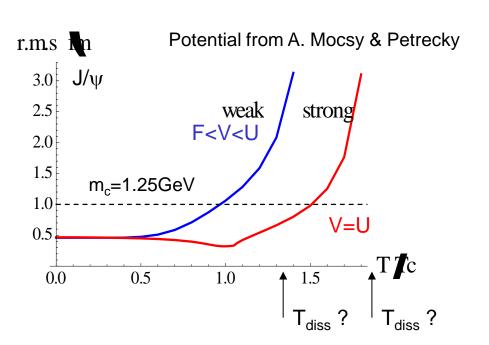
I. Quarkonia in *stationnary* medium are not well understood from the fundamental finite-T LQCD





From free energy $\Rightarrow V(r,T)$?

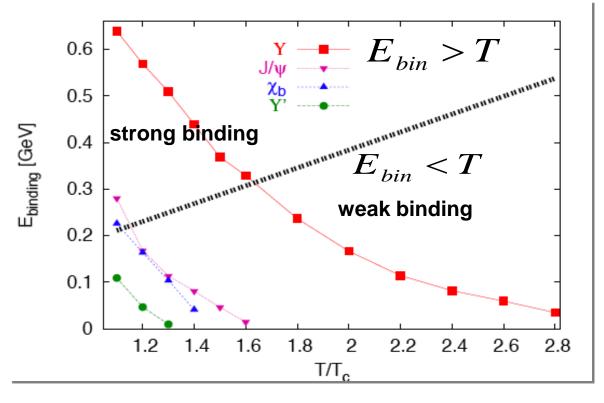
Several prescriptions in litterature



Caviats & Uncertainties

II. Criteria for quarkonia "existence" (as an effective degree of freedom) in *stationnary* medium is even less understood



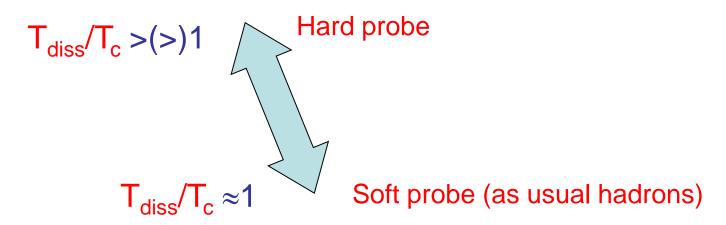


From A. Mocsy (Bad Honnef 2008)

Semi-Qualitative questions

The *main* object of interest here: T_{diss}, one of the fundamental quantities of statistical QCD.

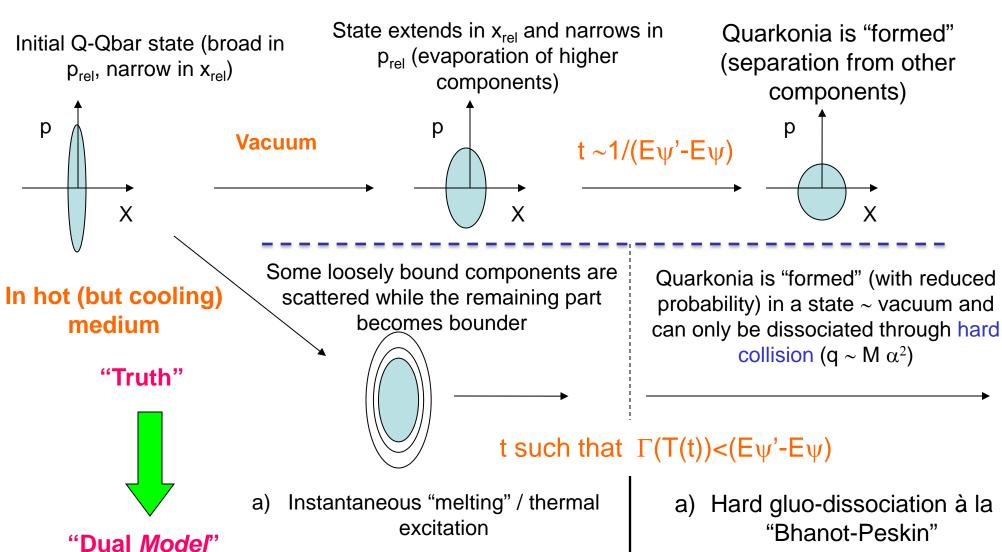
- 1. Can we try to extract the dissociation temperature from the data?
- 2. Are the data compatible with the picture of a strongly bound J/ψ (sequential suppression) ?



3. Can we challenge the picture of statistical recombination?

(A. Andronic, PBM, J. Stachel)

Quarkonia fate along decreasing T(t)



No "Q-Qbar→Quarkonia" fusion

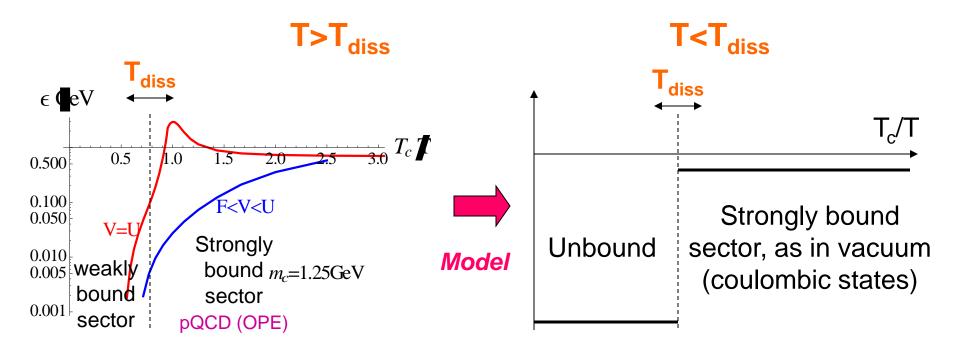
- "Q-Qbar → Quarkonia" fusion
 - allowed (+g) 67

Quarkonia fate along decreasing T(t)

"Dual Model"

- a) Instantaneous melting / thermal excitation
- o) No "Q-Qbar→Quarkonia" fusion

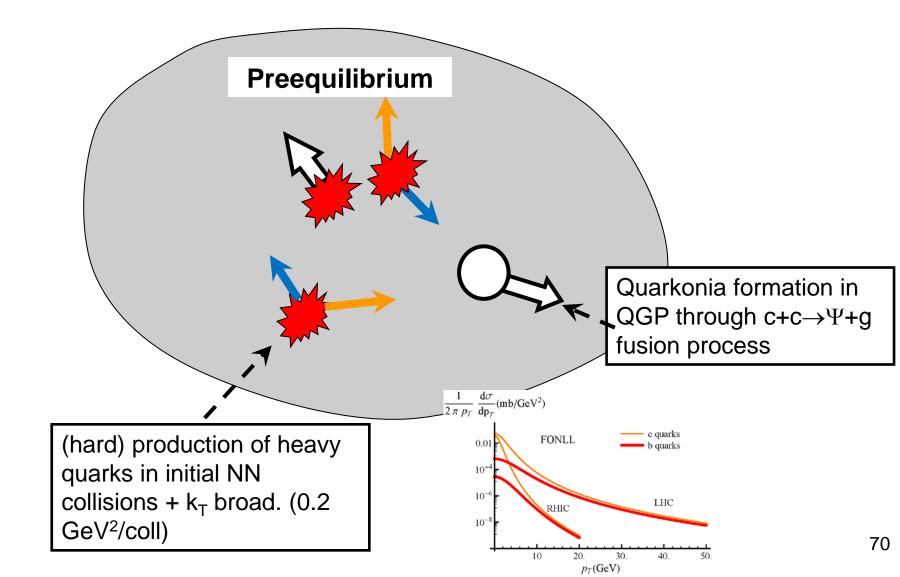
- a) Hard gluo-dissociation à la "Bhanot-Peskin"
- b) "Q-Qbar → Quarkonia" fusion allowed



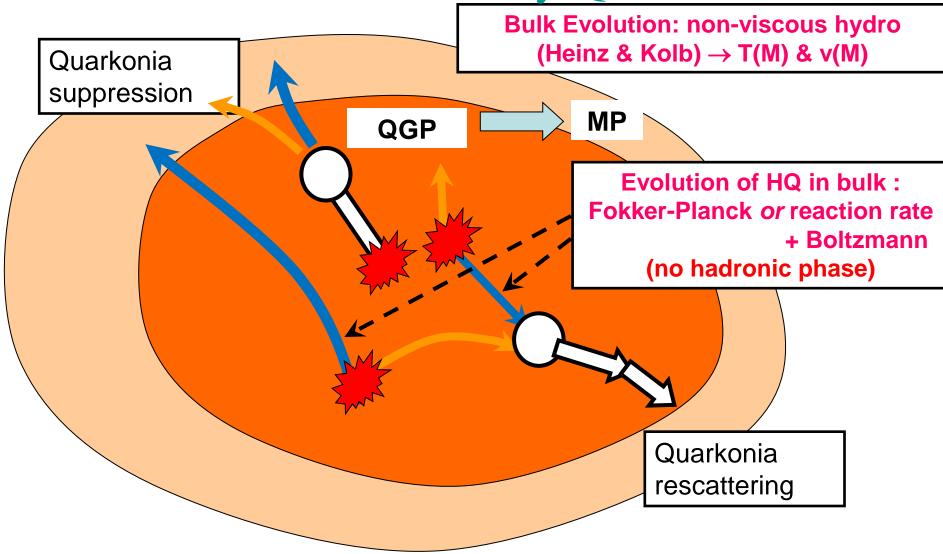
The key idea: AS THE LATTICE and POTENTIAL MODELS are inconclusive, let T_{diss} as a *free parameter* and see if this can be constrained by the data.

"Stationnary" quarkonia in evolving QGP

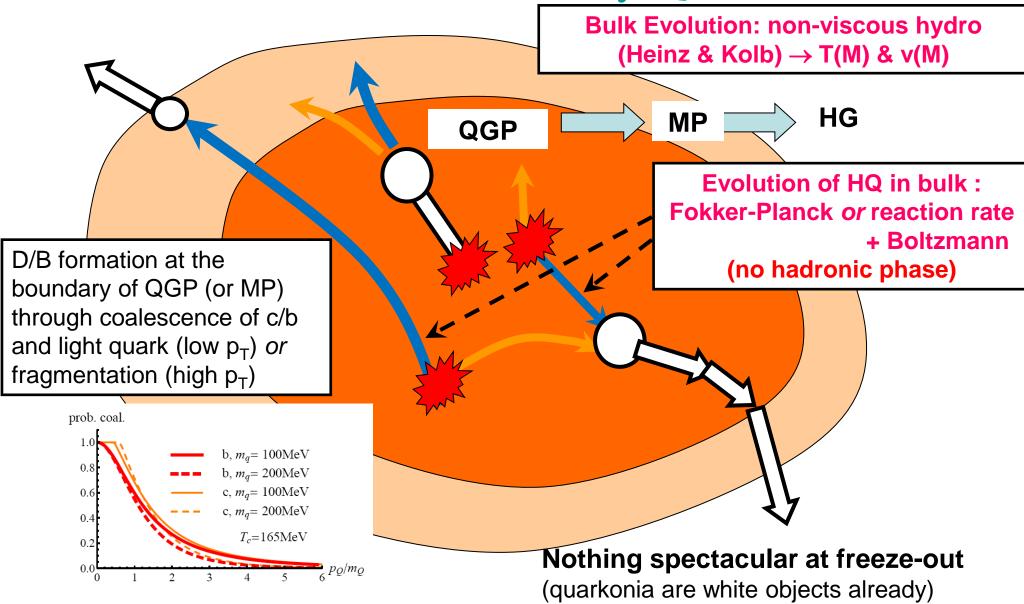
The Monte Carlo @ Heavy Quark Generator



The Monte Carlo @ Heavy Quark Generator

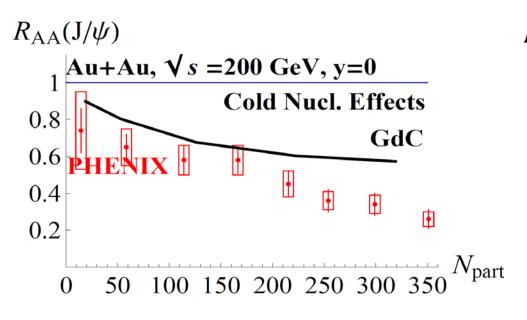


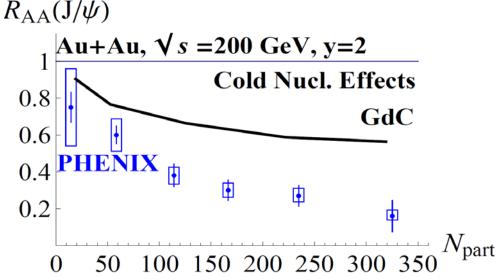
The Monte Carlo @ Heavy Quark Generator



Integrated J/Y numbers @ RHIC

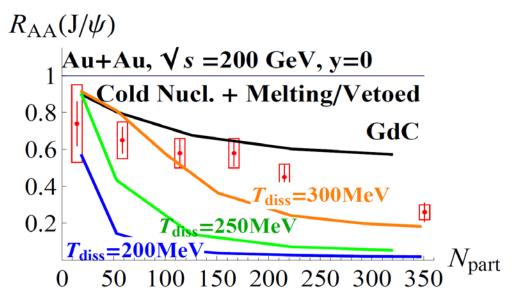
First, we need a baseline taking into account the cold nuclear matter effects (Shadowing, Cronin,..); we take the picture of R. Granier de Cassagnac (2007)

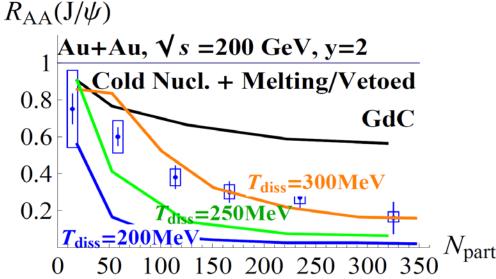




Integrated J/Y numbers @ RHIC

Next, the (instantaneous) vetoing of quarkonia formation due to melting:





Good agreement obtained with a rather large value of $T_{diss} \approx 2 T_c$.

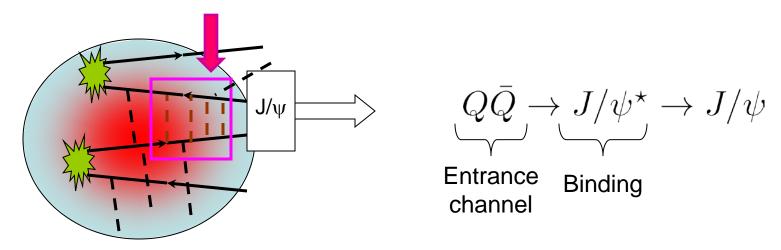
Some claims of "sequencial suppression" with a very bound J/ψ were indeed made by several physicists

""" We do not need recombination !"""...

except that Q and Qbar may be close in phase space

Turning on (re)combination + hard dissociation

(Re)combination (could be major process at LHC):

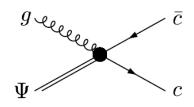


Often treated as a quasi-instantaneous fusion process

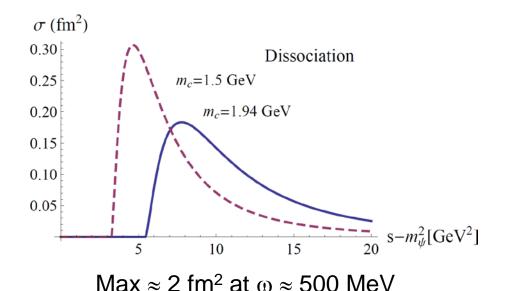
Basic Ingredients

Dissociation

hard dissociation taken according to Bhanot and Peskin + recoil correction (Arleo et al 2001)

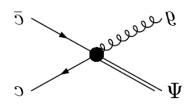


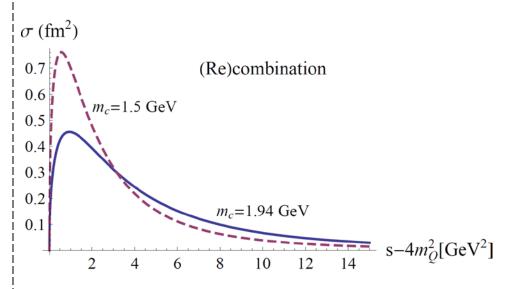
$$\sigma_{(Q\overline{Q})g}(\omega) = \frac{2^{11}}{3^4} \alpha_s \pi a_0^2 \frac{(\omega/\varepsilon(0) - 1)^{3/2}}{(\omega/\varepsilon(0))^5} \Theta(\omega - \varepsilon(0))$$



Recombination

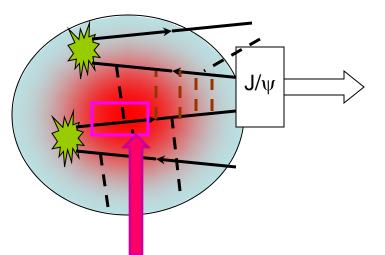
Cross section obtained from σ_{diss} via detailed balance





Turning on (re)combination + hard dissociation

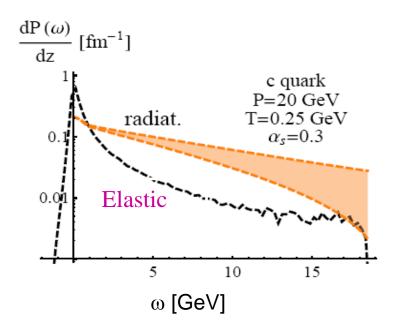
(Re)combination (could be major process at LHC):



Even if binding process is fast and mediumindependent (quarkonia are small bound states), the distributions of Q and Qbar in the entrance channel depend on the past history

(transport theory)

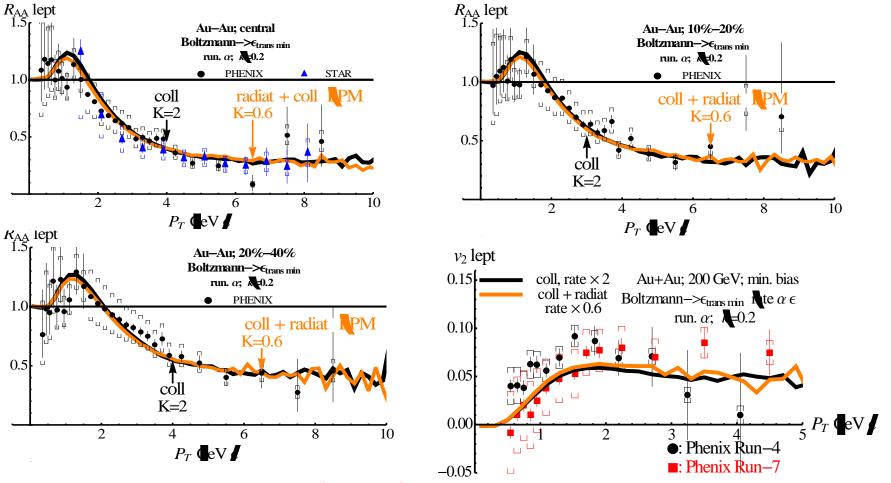
What is the dominant E loss mechanism @ RHIC and LHC? Does its detailed origin influences the fate of quarkonia's?



{Radiative + Elastic} vs Elastic for leptons @ RHIC

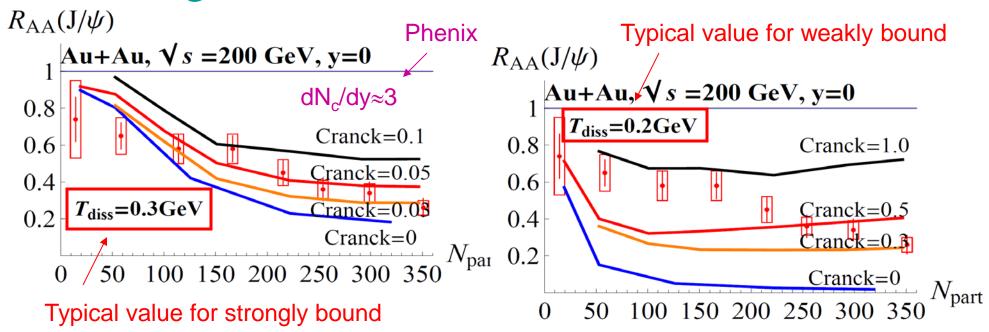
El. and rad. Eloss exhibit very different energy and mass dependences. However...

 σ_{el} & σ_{rad} cocktail: rescaling by K=0.6 σ_{el} alone rescaling: K=2



One "explains" it all with $\Delta E \alpha L$ (for HQ)

Turning on (re)combination + hard dissociation



Problem: One has to reduce the fusion probability by a factor ~ 10 to reproduce the data (if recomb. cross section taken at face value, one arrives at R_{AA} (most central > 2!).

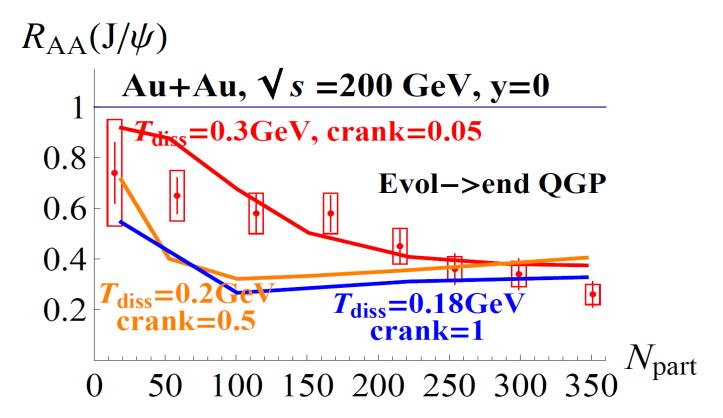
Problem never comes alone:

Strongly bound quarkonia are the ones for which the Bhanot-Peskin approach should be legitimate. Φ states exist early => lot of HQ pairs present in pahse space

Absolute numbers are better reproduced (if one believes in mostly canonical – cranck=0.5-1 – recombination), although the R_{AA} dependence on N_{part} is not as satisfying

Best parameters from R_{AA}

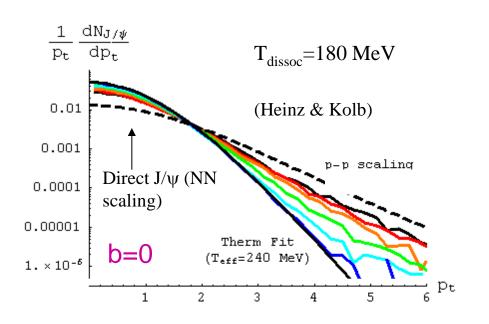
"Optimal" choices in the (T_{diss}, σ_{fus}) parameter plane

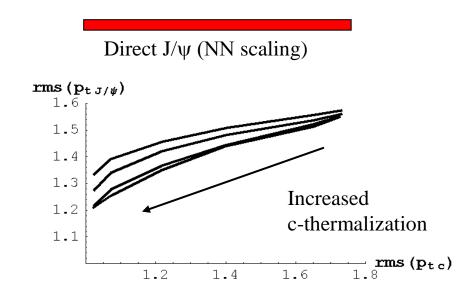


T_{diss} ∈ [0.2 GeV,0.3 GeV]... but difficult to go beyond

The p_t world

Differential production might reveal more physics





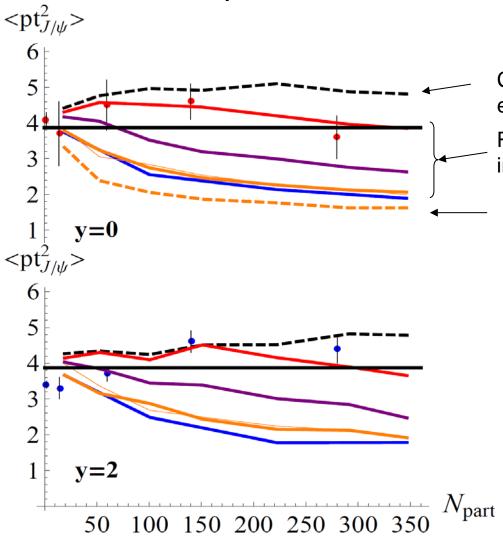
QGP "cools" the charms, even with the radial flow

Prediction for b=0 and just recombination

Softer p_t spectrum as for direct production. Possible " p_t shrinking" in A-A. But first, understand the k_t broadening in d+Au (recently observed by PHENIX)

The p_t world

... and now compared with the data:



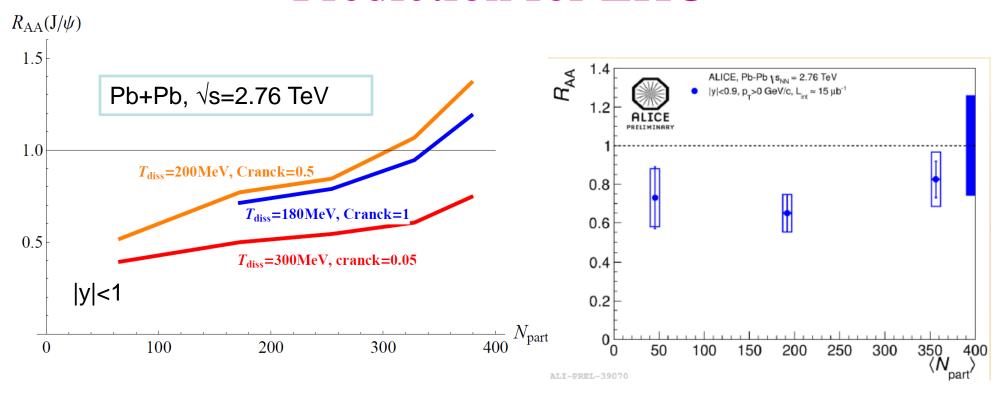
Cronin effect at initial stage (and no further effect)

Results for $T_{diss} = 0.3$, 0.25, 0.2 and 0.18 (with initial Cronin effect).

 $T_{diss} = 0.2$ and NO Cronin effect.

T_{diss}= 0.3 GeV should be favored

Prediction for LHC



Hydro Parameters:

$$s_0 = 268 \text{ fm}^{-3}$$

 $dN_{ch}/d\eta \approx 2300$ in PbPb, b=0

HQ Parameters:

dN_c/dy≈30 in PbPb

 $d\sigma_{\text{w}}/dy=2\mu b$ in pp

Fusion of c-quarks at LHC: 15-25 x more probable that at RHIC, but strong increase of the prompt J/ψ as well....

Preliminary conclusions

Reasonnable agreement with RHIC data for J/ψ , but difficulties to tame the recombination down

1. Can we try to *extract* the dissociation temperature from the data?

A rather large effective dissociation temperature ($T_{diss}\approx 0.25$ -0.3 GeV) seems to be favored by the data, provided one has a good quantitative argument to explain why the recombination of HQ should be reduced by a factor 10 w.r.t. the naive Bhanot - Peskin cross section (gluon mass ? $J/\psi(T)$ in BP ?)

Otherwise, low dissociation ($T_{diss} \approx 0.2$ GeV) are unavoidable... supported by finite J/ ψ v2 seen by ALICE

2. Are the data compatible with the picture of a strongly bound J/ψ (sequential suppression) ?

Not clear to us... questions the OPE approach

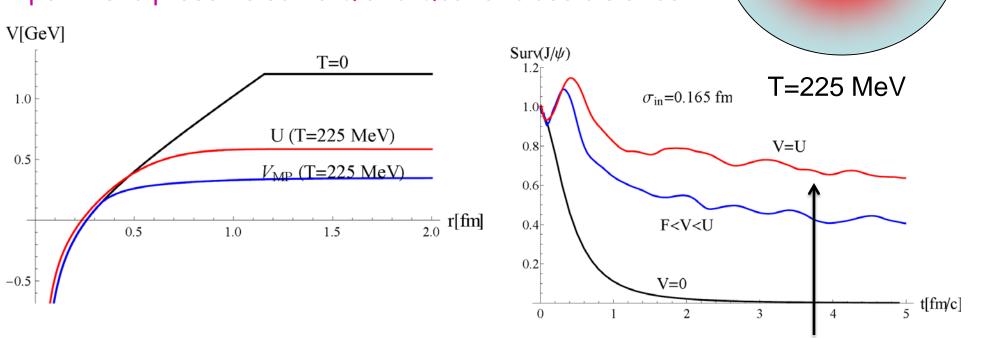
Need for a better description of Qqbar states in QGP

J/Ψ suppression (dynamical)

BUT: 2 missing ingredients

1. Q-Qbar forces (beginning 90s':Thews, Gossiaux and Cugnon,...):

permits to preserve some Q and Qbar at close distance



Indeed, the "residual" potential permits to slow down the suppression along time! We converge towards asymptotic survival probabilities \in [0,1]

J/w

J/Ψ suppression (dynamical)

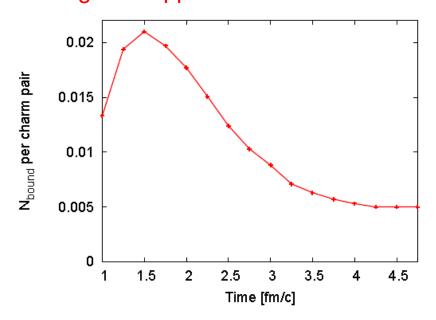
BUT: 2 missing ingredients

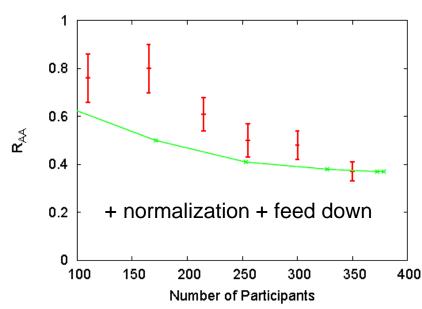
2. Stochastic q-Q, g-Q forces

For a long while: interactions with QGP/hot medium constituents only thought as the source for quarkonia dissociation (Bhanot – Peskin) and treated through inelastic cross-sections... True for dilute media

Shuryak & Young (08):

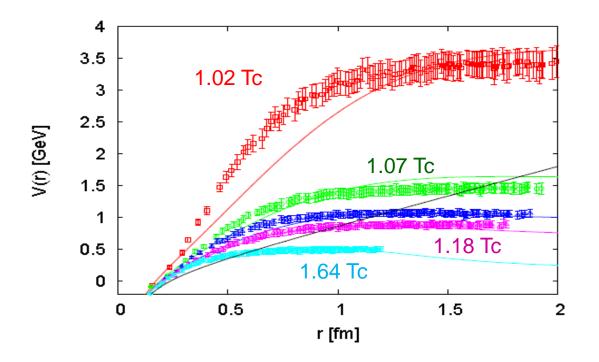
In strong QGP, diffusion of HQ slow down their separation (<r $^2> <math>\alpha$ D $_s$ t) and helps in reducing the suppression !!!





Shuryak & Young (08): some ingredients

✓ U as a potential.



The most "binding" choice; Around T_c: String tension up to 3 times string tension in vacuum !!!

Shuryak & Young (08): some ingredients

✓ Dealing both with quantum evolution and stochastic forces:

Wigner Moyal distribution (density operator):

$$F(\mathbf{x}^N, \mathbf{p}^N, t) = \left(\frac{1}{\pi \hbar}\right)^{3N} \int e^{2i\mathbf{p}^N \cdot \mathbf{y}^N l \hbar} \rho(\mathbf{x}_-^N, \mathbf{x}_+^N, t) d\mathbf{y}^N$$

Right concept for non pure quantum system (statistical average), but also to make contact with semi-classical interpretations

Wigner-Moyal equation in relative coordinates:

$$\left(\frac{\partial}{\partial t} + \frac{\vec{p}}{\mu} \cdot \frac{\partial}{\partial \vec{x}}\right) f(\vec{x}, \vec{p}; t) = \frac{2}{\hbar} \sin\left(\frac{\hbar}{2} \frac{\partial}{\partial \vec{p}} \cdot \frac{\partial}{\partial \vec{x}}\right) V(\vec{x}) f(\vec{x}, \vec{p}; t) + I_{\text{col}}$$

with
$$\vec{x} = \vec{x}_Q - \vec{x}_{\bar{Q}}$$
 and $\vec{p} = \frac{\vec{p}_Q - \vec{p}_{\bar{Q}}}{2}$

Exact equation, but difficult to solve due to sign problem

Shuryak & Young (08): some ingredients

✓ Dealing both with quantum evolution and stochastic forces:

Semi-classical expansion => 1 body Liouville equation:

$$\left(\frac{\partial}{\partial t} + \frac{\vec{p}}{\mu} \cdot \frac{\partial}{\partial \vec{x}} - \frac{\partial V}{\partial \vec{x}} \cdot \frac{\partial}{\partial \vec{p}}\right) f(\vec{x}, \vec{p}; t) = I_{\text{col}}$$

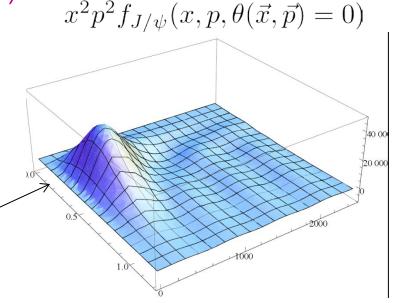
Test particles method, submitted to the QQbar force + stochastic external forces

Langevin evolution with binding force (♥ fast !!! ♥)

Prob J/
$$\psi$$
(t): $P_{J/\psi}(t) = \frac{1}{N} \sum_{i=1}^{N} f_{J/\psi}(\vec{x}_i(t), \vec{p}_i(t))$

Caviat: f is not a density (not defined positive) semi-classical approx justified?

Notice however that $f_{J/\psi}$ is mostly positive (but not a full justification)



Shuryak & Young (08): some ingredients

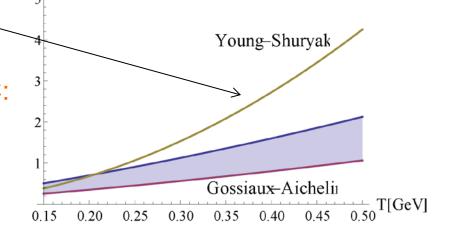
- ✓ Stochastic force on Q and Qbar are uncorrelated
 - ... although QQbar is seen as a dipole at short distances
 - ...but most of Q-Qbar pairs are not at close distance already after short time => probably ok !
- ✓ Hydro evolution and HQ dynamics from Moore and Teaney (2005). In particular D_c x 2πT=1.5-3 =>

$$A_c = \frac{T}{MD_c} = \frac{2\pi T^2}{1.5M}$$

Our model + detailed comparison to RHIC:

$$A_c[c/fm] = K (1.5T[GeV] + 1.25T^2)$$

Effective linear rise: $\alpha_s(T)$



Test of robustness

Goal of our contribution:

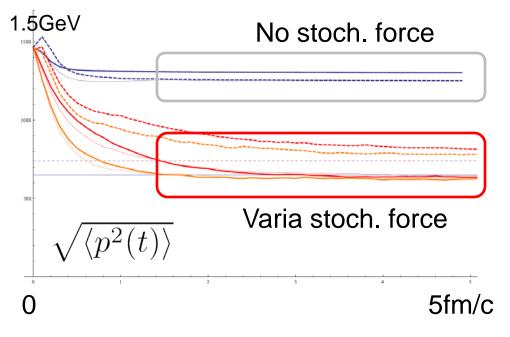
- Get acquainted with the impact of stochastic forces on quarkonia suppression
- ✓ Test the robustness of the results obtained by Young and Shuryak, modifying
 - a) the V(T)
 - b) the drag coefficient A(T)
 - c) the semi-classical treatment of the c-cbar evolution (tougher, not today)

Test of robustness for stationnary QGP

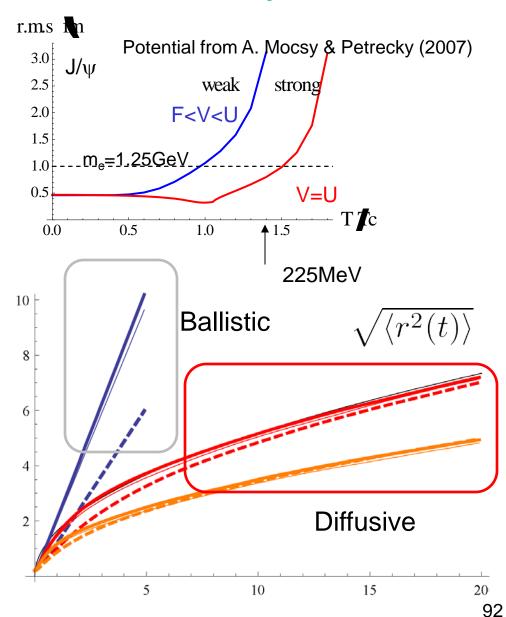
T=225 MeV (T/Tc ≈ 1.4):

Nearly unbound if one takes V=V_{PM}, still strongly bound if one takes V=U

$$\sqrt{\langle r^2(t=0)\rangle} = 0.2 \text{ fm}$$



Stochastic cooling of c-cbar state



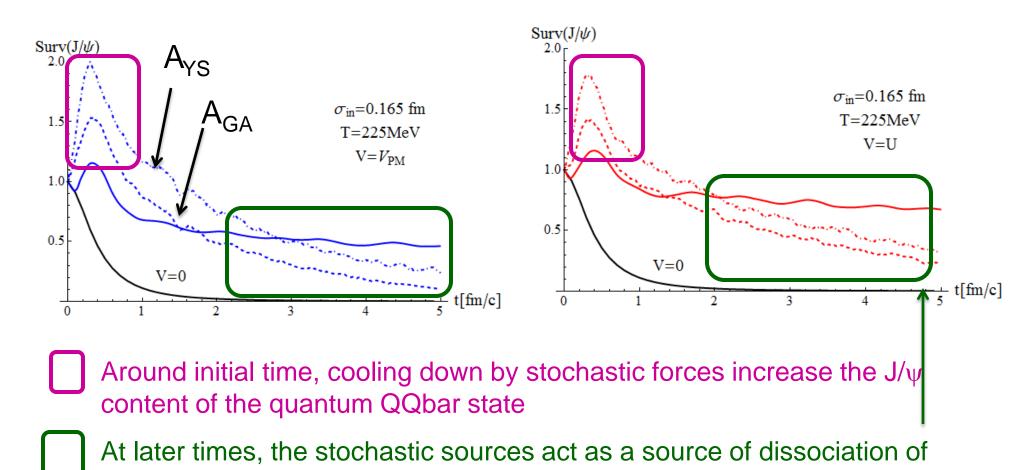
Test of robustness for stationnary QGP

T=225 MeV (T/Tc \approx 1.4):

the remaining state

V=V_{PM} (weakly bound)

V=U (strongly bound)

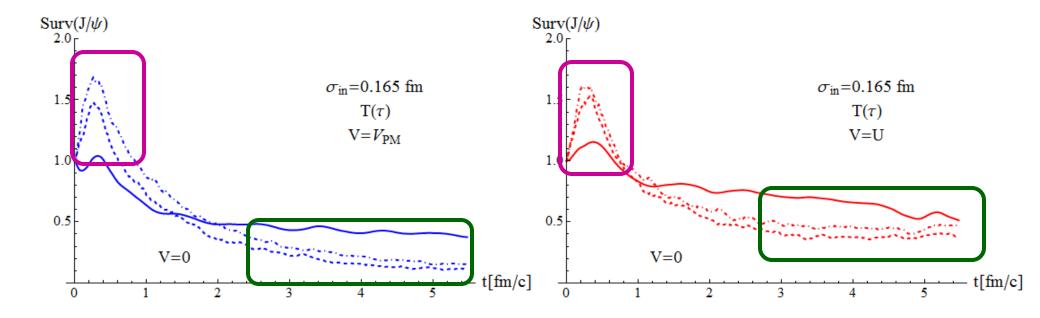


Test of robustness for evolving QGP

 $T(\tau)$, central Au-Au @ RHIC, $\vec{x}_{\perp} = \vec{0}$

V=V_{PM} (weakly bound)

V=U (strongly bound)



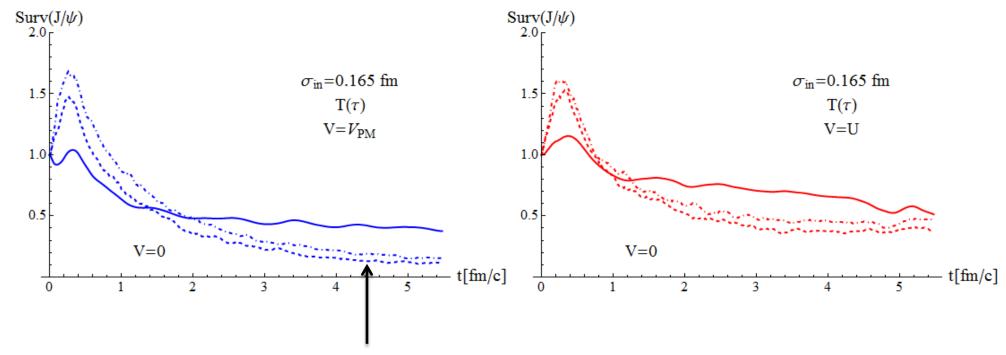
- ➤ Similar features as for T=225: rapid thermalization in p-space (-> quasi equilibrium), followed by induced leakage in r space
- ➤ For potential chosen as V=U, survival compatible to 0.5, as claimed by Young and Shuryak

Test of robustness for evolving QGP

T(τ), central Au-Au @ RHIC, $\vec{x}_{\perp} = \vec{0}$

V=V_{PM} (weakly bound)

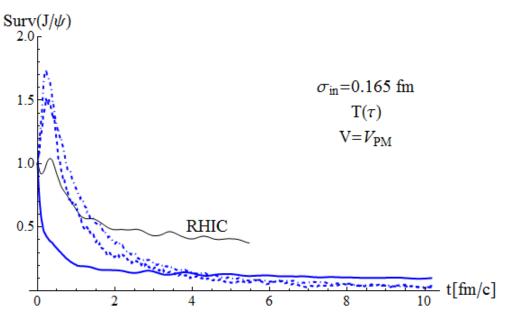
V=U (strongly bound)



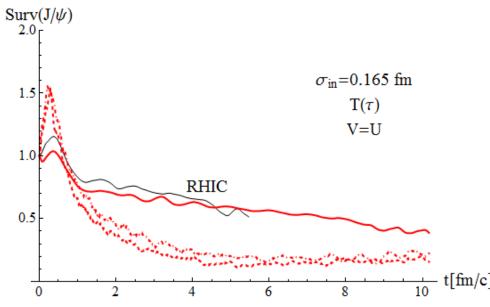
- ➤ No large dependence vs precise choice for drag coefficient...
- ➤ But large dependence vs choice of potential, especially if one includes the stochastic forces (can dissociate weakly bound states, but rather inefficient to dissociate strongly bound states).

Survival @ LHC

T(τ), central Pb-Pb @ LHC, $\vec{x}_{\perp} = \vec{0}$



Preliminary



Even at LHC, up to 25% survival if V=U; should not be neglected

Conclusion & Prospects

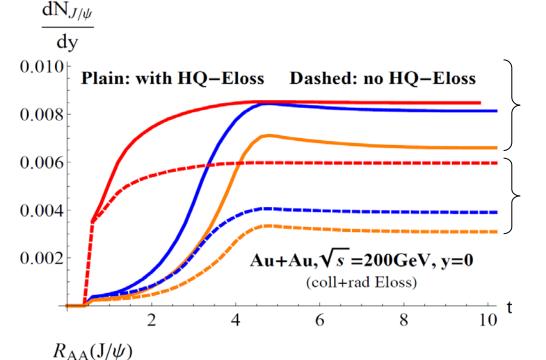
- 1. Important to include a time-dependent microscopic description of Q-Qbar states in the transport codes... to be pursued
- 2. We confirm the claim of Shuryak and Young of large J/ψ survival... for V chosen to be the total energy U...
- 3. However, their choice of parameters probably correspond to the most favorable case!

Possible way to make progress on this point: evaluate $\Gamma_{J/\psi}(T)$ for both types of potentials and compare with lattice

4. I am very excited(QCD) about all of this

Back Up

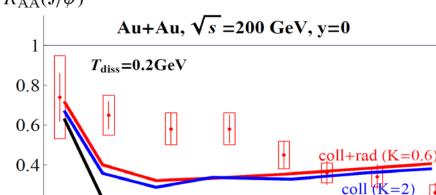
Finer analysis: role of HQ energy loss



Eloss

No Eloss

Energy loss favors the coalescence of J/ψ (brings the c quarks together in phase space)



150

200

No Eloss

300

350

250

0.2

0

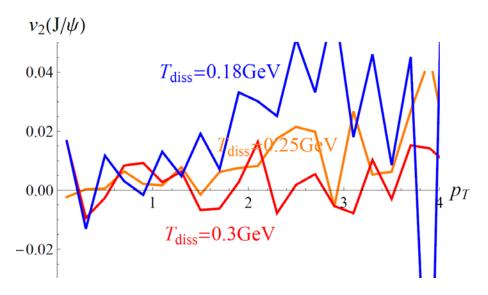
50

100

However: Once the Energy loss has been "properly" calibrated on non-photonic single-e R_{AA}, then the production rates do not depend too much on the detailed phenomena

The keystone (?): v₂

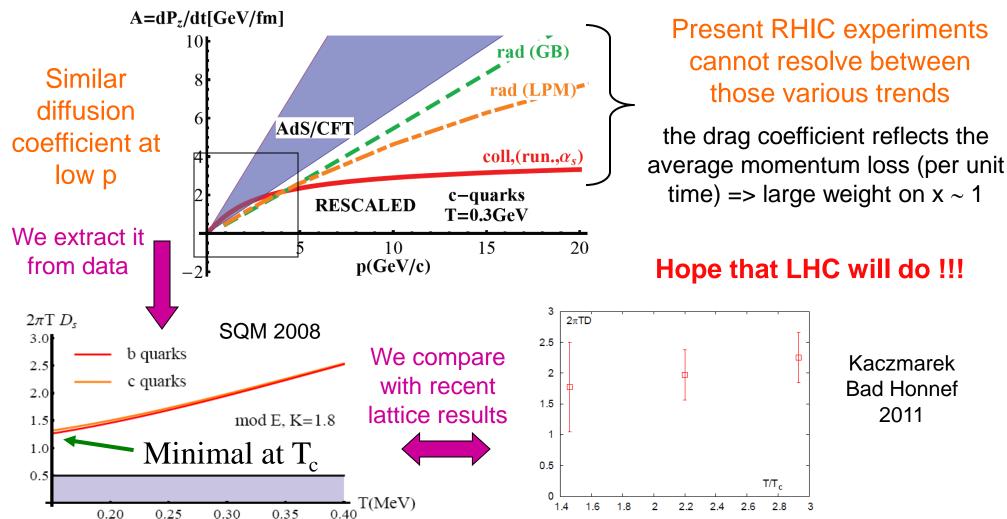
RHIC



Beware of the possible role of elastic cross section of J/ψ in the experimental v2

QGP properties from HQ probe

Gathering all rescaled models (coll. and radiative) compatible with RHIC R_{AA}:

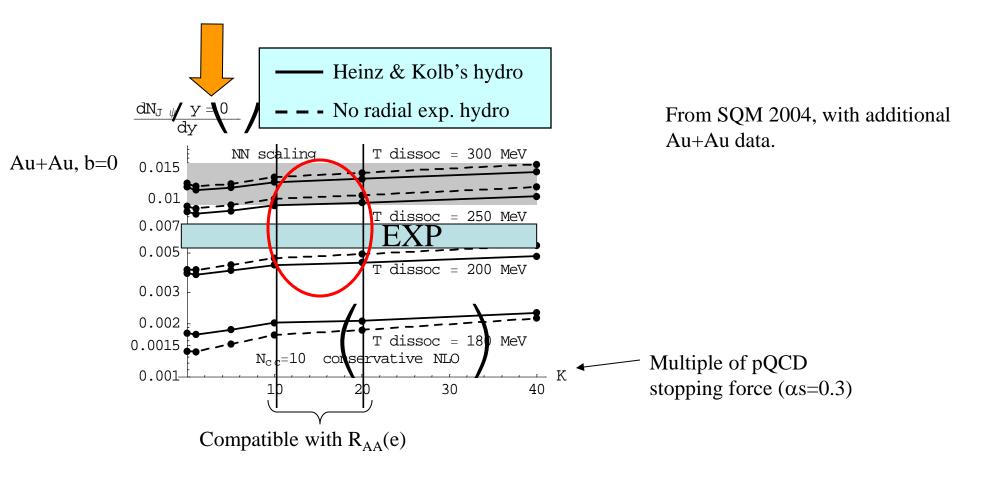


Lesson

Yes, it seems possible to reveal some fundamental property of QGP using HQ probes

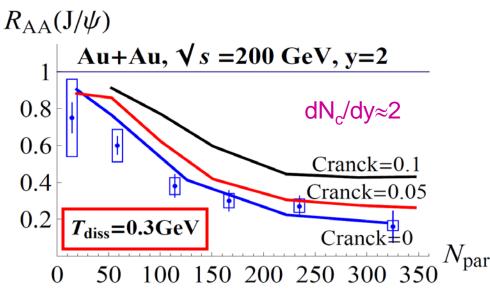
The Landscape

Degree of thermalization of heavy quarks will not affect "too much" the integrated production rates; T_{diss} is the driving parameter for "recombined" J/ψ :



Turning on (re)combination at y=2

 $R_{\rm AA}({\rm J}/\psi)$



Au+Au, $\sqrt{s} = 200 \text{ GeV}$, y= 2 $0.8 \\ 0.6 \\ 0.4 \\ 0.2$ Cranck=1.0

Cranck=0.5

Cranck=0

Npart

No room left for coalescence at y=2. What are the physical mechanisms for taming the fusion?

Moreover: The pQCD Bhanot and Peskin result is usually considered to be small w.r.t. other effective approaches at small s-M²

Good agreement with the same σ_{fus} band (Cranck. \in [0.5,1])

