

Light sterile neutrinos and R_K

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Based on work in collaboration with
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ArXiv:1211.3052

Introduction

Violations of **lepton flavor universality** are among the most precise **tests** of the Standard Model.

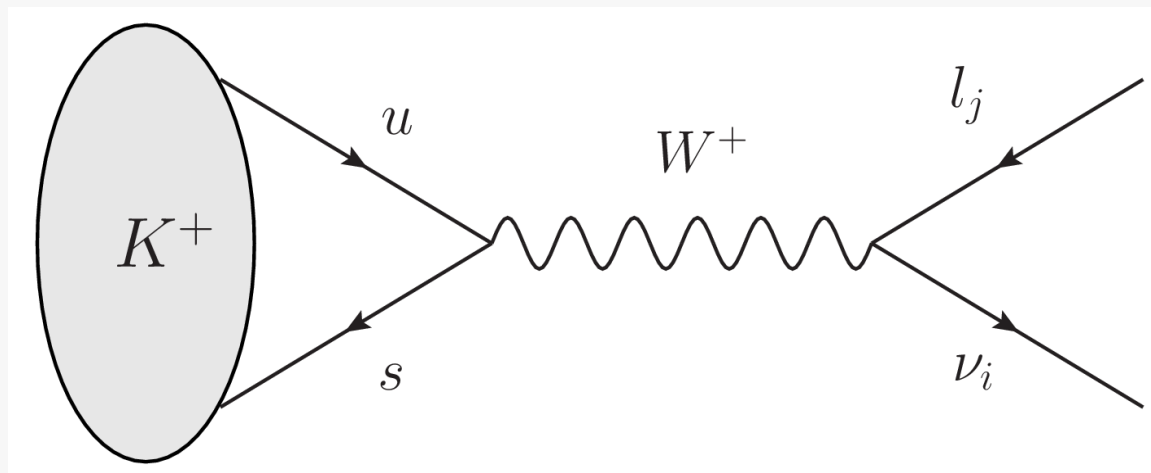
$$R_K = \frac{\Gamma(K^+ \rightarrow e^+ \nu)}{\Gamma(K^+ \rightarrow \mu^+ \nu)}$$

Measured in the **NA62 experiment** using a beam of kaons from the **CERN SPS**

Current experimental error: $\frac{\delta R_K}{R_K} \sim 0.4\%$

Expected sensitivity: $\frac{\delta R_K}{R_K} \sim 0.1\%$

R_K in the SM



V. Cirigliano, I. Rosell, PRL 99 (2007) 231801

$$R_K^{SM} = \left(\frac{m_e}{m_\mu}\right)^2 \left(\frac{m_K^2 - m_e^2}{m_K^2 - m_\mu^2}\right)^2 (1 + \delta R_{QED}) = (2.477 \pm 0.001) \times 10^{-5}$$

SM result

↑
Helicity suppression

↑
Small QED corrections

R_K in the SM

Experimental value:
(2011) $R_K^{exp} = (2.488 \pm 0.010) \times 10^{-5}$

NA62 Collaboration

$$R_K^{SM} = (2.477 \pm 0.001) \times 10^{-5}$$

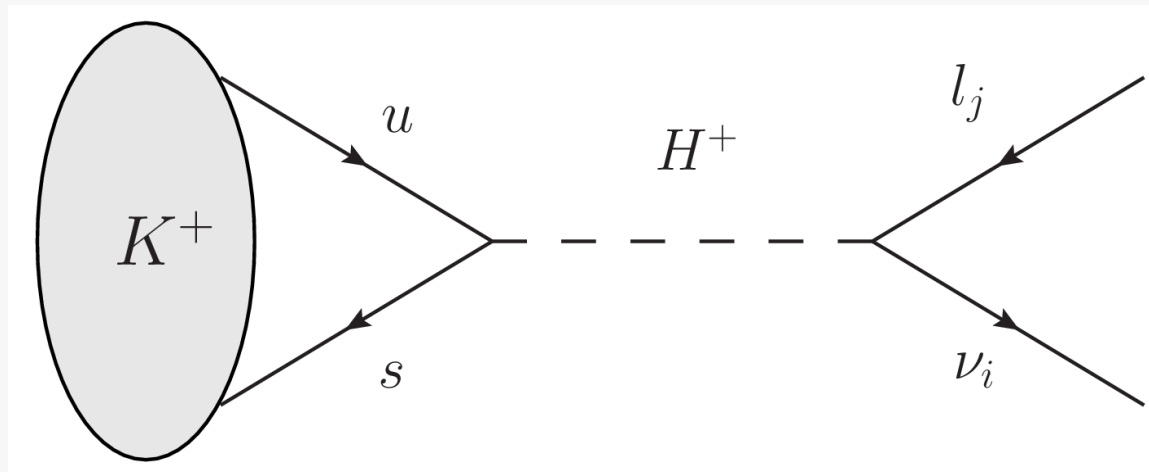
$$R_K = R_K^{SM} (1 + \Delta r)$$

$$\Delta r_K = (4 \pm 4) \times 10^{-3}$$

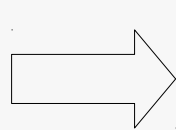
How to modify R_K ?

- In a 2HDM:

W.-S. Hou, PRD 48 (1993) 2342



$$\Gamma(K^\pm \rightarrow l^\pm \nu) = \Gamma^{\text{SM}}(K^\pm \rightarrow l^\pm \nu) \left(1 - \tan^2 \beta \frac{m_K^2}{m_H^2} \frac{m_s}{m_s + m_u} \right)^2$$



The correction is **lepton universal**

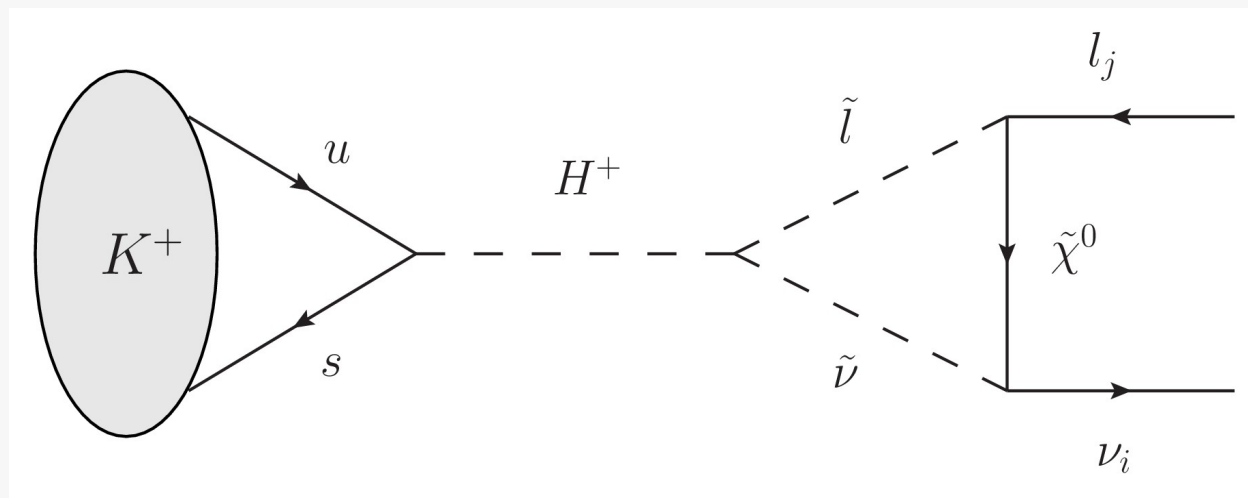
R_K is **not** modified

How to modify R_K ?

Higher order corrections are required

• In SUSY:

A. Masiero, P. Paradisi, R. Petronzio, PRD 74 (2006) 011701



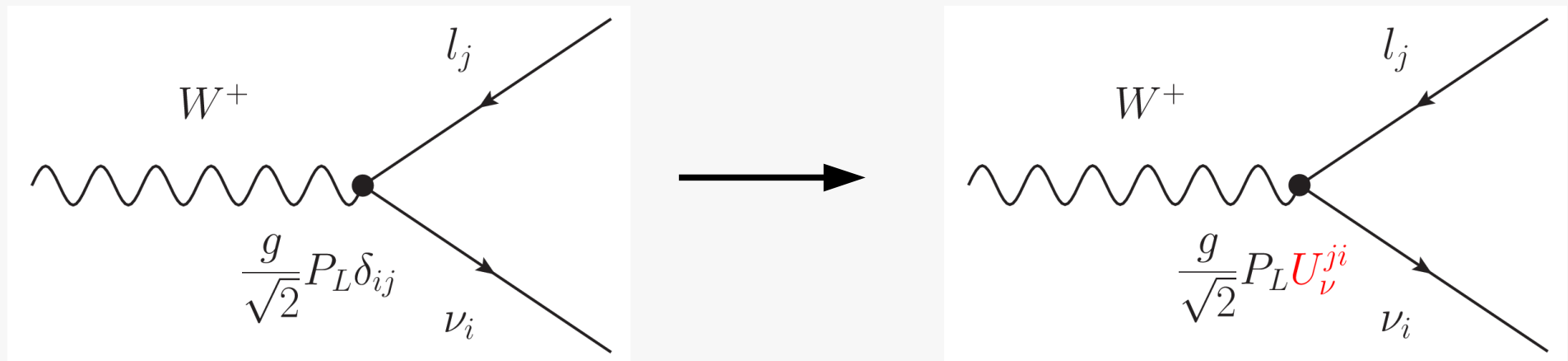
R. M. Fonseca, J. C. Romão, A. M. Teixeira, arXiv:1205.1411

Taking into account **current constraints** from $B \rightarrow \tau \nu$ one can go up to

$$\Delta r_K \lesssim 10^{-3}$$

A different scenario

Let us now consider **non-zero neutrino masses and mixings**...



$$\Delta r_K = \frac{m_\mu^2}{m_e^2} \frac{(m_K^2 - m_\mu^2)^2}{(m_K^2 - m_e^2)^2} \frac{\sum_m |U_\nu^{1m}|^2 G^{m1}}{\sum_n |U_\nu^{2n}|^2 G^{n2}} - 1$$

$$G^{ij} = \left[m_K^2 (m_{\nu_i}^2 + m_{l_j}^2) - (m_{\nu_i}^2 - m_{l_j}^2)^2 \right] \left[(m_K^2 - m_{l_j}^2 - m_{\nu_i}^2)^2 - 4m_{l_j}^2 m_{\nu_i}^2 \right]^{1/2}$$

A different scenario

However, under the **assumptions**...

$$1) \quad m_{\nu_i}^2 \ll m_K^2, m_{l_j}^2$$

$$\Rightarrow \quad G^{ij} \simeq m_{l_j}^2 (m_K^2 - m_{l_j}^2)^2$$

$$\Delta r_K \simeq \frac{\sum_m |U_\nu^{1m}|^2}{\sum_n |U_\nu^{2n}|^2} - 1$$

$$2) \quad \sum_i |U_\nu^{ji}|^2 = 1$$

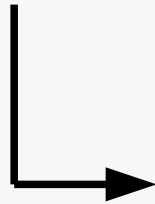
$$\Delta r_K \simeq 0$$

No deviation from
the SM result

Dropping the assumptions

1) ~~$m_{\nu_i}^2 \ll m_K^2, m_{l_j}^2$~~

Additional neutrino states can be produced in the kaon decay



Sterile neutrinos with masses $m_N \lesssim m_K$

2)
$$\sum_i |U_{\nu}^{ji}|^2 = 1$$

Dropping the assumptions

1) $m_{\nu_i}^2 \ll m_K^2, m_{l_j}^2$

Sterile neutrinos with masses $m_N > m_K$

2) ~~$\sum_i |U_{\nu_j i}|^2 = 1$~~

Deviation from **unitarity**
Large mixings are necessary

Viabile scenario?

This scenario was actually proposed **more than 30 years ago...**

R. E. Shrock, Phys. Lett. B 96 (1980) 159
R. E. Shrock, Phys. Rev. D 24 (1981) 1232

Things have changed a lot since then...

Is this scenario compatible with **current constraints?**

Constraints

There are, however, some **constraints**...

1) Lepton flavor violation

F. Deppisch, J. W. F. Valle, PRD 72 (2005) 036001

$$\text{Br}(\mu \rightarrow e\gamma) = \frac{\alpha_W^3 s_W^2 m_\mu^5}{256\pi^2 m_W^4 \Gamma_\mu} |H_{\mu e}|^2, \quad H_{\mu e} = \sum_i U_\nu^{2i} U_\nu^{1i*} G_\gamma \left(\frac{m_{\nu,i+3}^2}{m_W^2} \right)$$

$$\text{Br}(\mu \rightarrow e\gamma)_{\text{MEG}} = 2.4 \cdot 10^{-12}$$

2) Non-unitarity bounds

S. Antusch et al, Nucl. Phys. B 810 (2009) 369

$$U_{3 \times 3} = (1 - \eta) U_{PMNS}$$



Strongly **constrained**
if $m_N > m_{EW}$

Constraints

3) Laboratory bounds: direct searches for sterile neutrinos

Bounds from decays like

$$\pi^{\pm} \rightarrow \mu^{\pm} \nu$$

(they would lead to monochromatic lines in the muon spectrum)

A. Atre et al, JHEP 0905 (2009) 030

A. Kusenko, Phys.Rept. 481 (2009) 1

4) B physics

$$\text{Br}(B \rightarrow e\nu) < 9.8 \cdot 10^{-7}$$

$$\text{Br}(B \rightarrow \mu\nu) < 10^{-6}$$

$$\text{Br}(B \rightarrow \tau\nu) = (1.65 \pm 0.34) \cdot 10^{-4}$$

Constraints

5) LHC bounds

P. S. Bhupal Dev et al, arXiv:1207.2756

C. García Cely et al, arXiv:1208.3654

P. Bandyopadhyay et al, arXiv:1209.4803

Indirect bounds from **Higgs searches**

$$h \rightarrow \nu_R \nu_L$$

Relevant for sterile neutrino masses **around 100 GeV**

6) EW precision data

F. Del Aguila et al, PRD 78 (2008) 013010

Leptonic mixing affects **EW precision data fits**

Apply to sterile neutrino masses $\gtrsim 1$ **TeV**

Constraints

7) Cosmological bounds on sterile neutrinos

A. Y. Smirnov, R. Zukanovich Funchal, PRD 74 (2006) 013001

A. Kusenko, Phys.Rept. 481 (2009) 1

- Large scale structure
- Lyman- α limits
- BBN
- CMB
- X-ray constraints (from $\nu_i \rightarrow \nu_j \gamma$)

However, cosmological bounds can be evaded with a **non-standard cosmology** (for example, **low reheating temperature**)

G. Gelmini *et al*, JCAP 0810 (2009) 029

An example model

R. N. Mohapatra, J. W. F. Valle, PRD 34 (1986) 1642

An example model: Inverse Seesaw

$$-\mathcal{L}_{IS} \supset Y_\nu^{ij} \nu_i^c L_j \tilde{H} + M_{R_{ij}} \nu_i^c X_j + \frac{1}{2} \mu_{X_{ij}} X_i X_j$$

- 6 additional **singlet states**: 3 generations of ν^c and 3 generations of X
- **Non-zero neutrino masses**. In the limit $\mu_X \ll Y_\nu v \ll M_R$:

$$m_\nu \simeq \frac{v^2}{2} Y_\nu^T (M_R^T)^{-1} \mu_X M_R^{-1} Y_\nu$$

- The **suppression** by μ_X allows to have (in principle) $Y_\nu \sim \mathcal{O}(1)$

Large mixings and light sterile neutrinos are possible

An example model

The 9 neutrinos participate in the **charged current interaction** through their **LH admixture**

$$-\mathcal{L}_{cc} = \frac{g}{\sqrt{2}} U_{\nu}^{ji} \bar{l}_j \gamma^{\mu} P_L \nu_i W_{\mu}^{-} + \text{c.c.}$$

$$i = 1, \dots, 9$$

$$j = 1, \dots, 3$$

We have all the ingredients for a **large contribution to R_K**

Numerical results

Random scan of the parameter space of the inverse seesaw, filtered by all previous constraints.

We only fix active neutrino masses and mixings in order to reproduce **neutrino oscillation data**

Scenario 1

$$m_N \lesssim m_K$$

$$M_R \in [0.1, 200] \text{ MeV}$$

$$\mu_X \in [0.01 \text{ eV}, 1 \text{ MeV}]$$

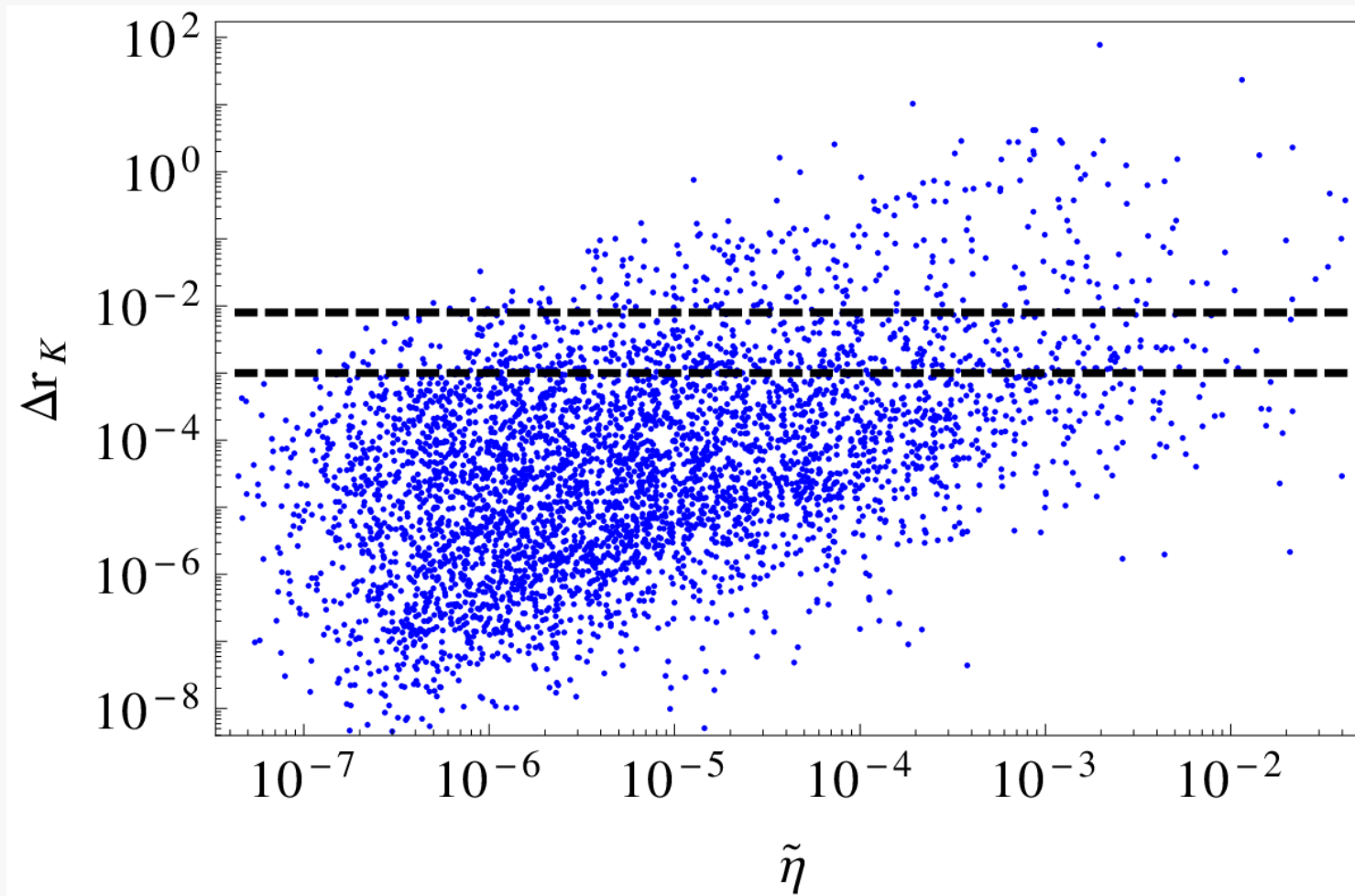
Scenario 2

$$m_N > m_K$$

$$M_R \in [1, 10^6] \text{ GeV}$$

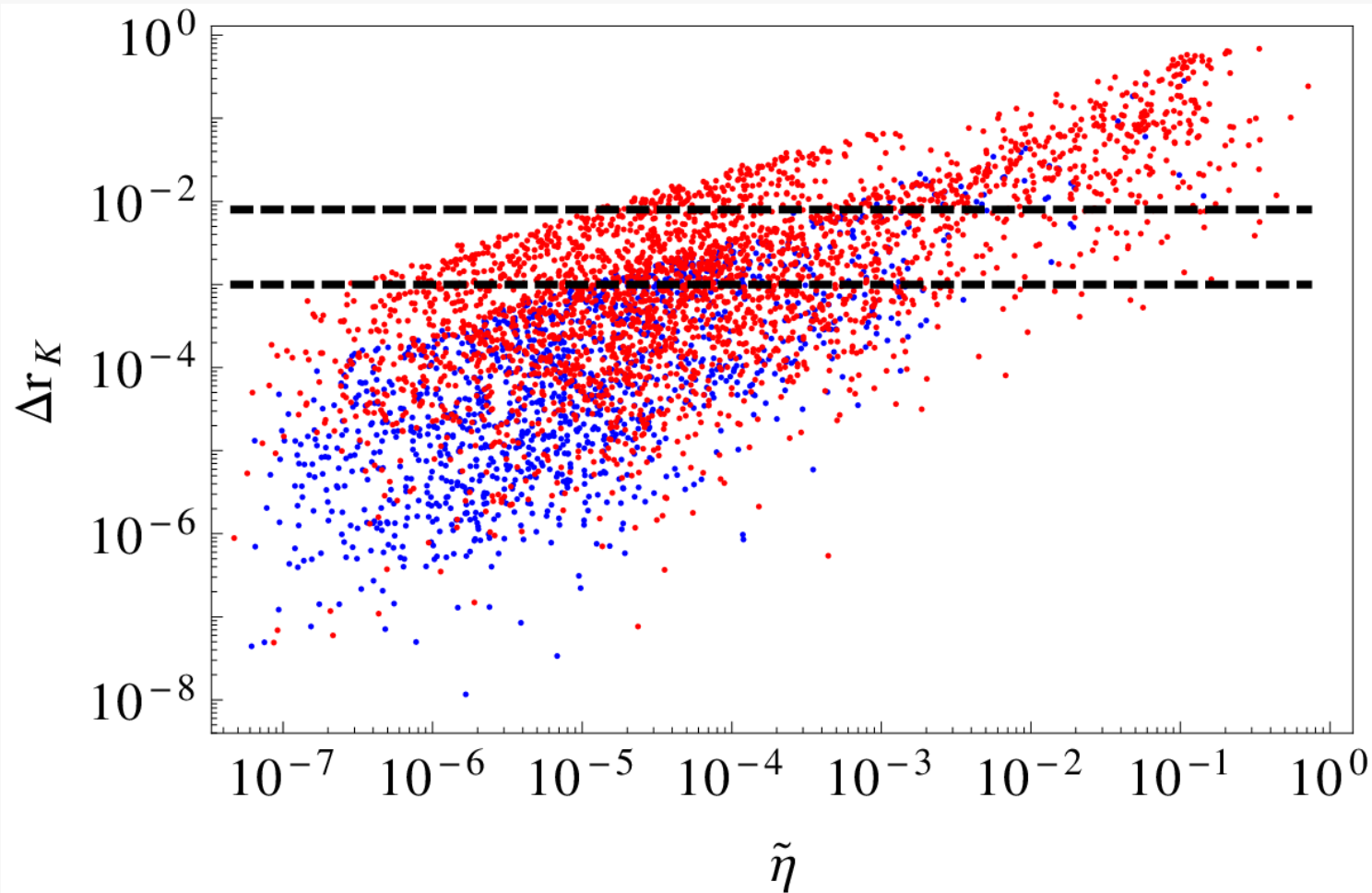
$$\mu_X \in [0.01 \text{ eV}, 1 \text{ MeV}]$$

Scenario 1 : $m_N \lesssim m_K$



- Here $\tilde{\eta} = 1 - \text{Det}(U_{3 \times 3})$ measures the deviation from unitarity.
- $\Delta r_K \sim \mathcal{O}(1)$ can be reached: **very good perspectives for NA62.**

Scenario 2 : $m_N > m_K$



$$Y_\nu > 10^{-2}$$

$$Y_\nu < 10^{-2}$$

- In this scenario one can also go up to $\Delta r_K \sim \mathcal{O}(1)$
- Strong limits from **non-unitarity** and **EW precision data** apply in this case.

Conclusions

- Lepton universality violating observables are very good tests of the SM: high precision (theory & experiment)
- The presence of “light” sterile neutrinos can modify R_K in a very substantial way.
- Two scenarios have been considered: good perspectives in both!
- Stay tuned to NA62...

Thank you!