

Avoiding Death by Vacuum

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A. Barroso, P.F. , I. Ivanov, R. Santos, J.P. Silva, arXiv:1211.6119

- **LHC hints of the presence of a scalar particle with mass ~ 125 GeV.**

- **Up to now, all is compatible with the Standard Model (SM) Higgs particle.**

BORING!

Two-Higgs Doublet model, 2HDM (Lee, 1973) : one of the easiest extensions of the SM, with a richer scalar sector. Can help explain the matter-antimatter asymmetry of the universe, provide dark matter candidates, ...

**G.C. Branco, P.M. Ferreira, L. Lavoura, M. Rebelo, R. Santos, M. Sher, J.P Silva,
Physics Reports 716, 1 (2012)**

Scalar sector of the 2HDM

Two doublets \Rightarrow 4 neutral scalars (h, H, A) + 1 charged scalar (H^\pm).

h }
H } CP-even scalars

$h, H \rightarrow \gamma \gamma$
 $h, H \rightarrow ZZ, WW$ (real or off-shell)
 $h, H \rightarrow ff$
 $H \rightarrow hh$ (if $m_H > 2m_h$)
...

A - CP-odd scalar
(pseudoscalar)

$A \rightarrow \gamma \gamma$
 ~~$A \rightarrow ZZ, WW$~~
 $A \rightarrow ff$
...

Vacuum structure more rich => different types of minima possible!

The **NORMAL** minimum,

$$\langle \Phi_1 \rangle = \begin{pmatrix} 0 \\ v_1 \end{pmatrix} \quad e \quad \langle \Phi_2 \rangle = \begin{pmatrix} 0 \\ v_2 \end{pmatrix}$$

The **CHARGE BREAKING (CB)** minimum, with

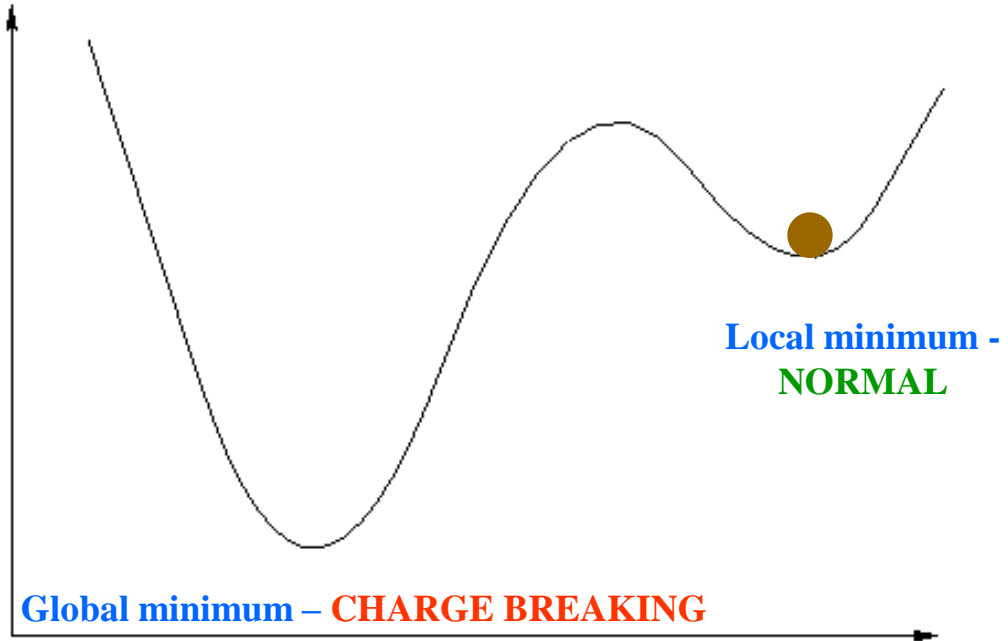
$$\langle \Phi_1 \rangle = \begin{pmatrix} 0 \\ v'_1 \end{pmatrix} \quad e \quad \langle \Phi_2 \rangle = \begin{pmatrix} \alpha \\ v'_2 \end{pmatrix} \quad \alpha \text{ has electric charge!}$$

The **CP BREAKING** minimum, with

$$\langle \Phi_1 \rangle = \begin{pmatrix} 0 \\ v'_1 + i\delta \end{pmatrix} \quad e \quad \langle \Phi_2 \rangle = \begin{pmatrix} 0 \\ v'_2 \end{pmatrix} \quad \delta \text{ breaks CP!}$$

Would there be any problem if the potential had two of these minima simultaneously?

Answer: there might be, if the CB minimum, for instance, were “deeper” than the normal one (metastable).



~~$m_\gamma = 0$~~

$m_\gamma \neq 0 !$

THEOREM: if a Normal Minimum exists, the Global minimum of the theory is Normal - the photon is guaranteed to be massless.

(not so in SUSY, for instance)

Barroso,
Ferreira,
Santos

$$\begin{aligned}
 V_{CB} - V_N &= \left(\frac{M_{H^\pm}^2}{4v^2} \right)_N [(v'_1 v_2 - v'_2 v_1)^2 + \alpha^2 v_2^2] \\
 V_{CP} - V_N &= \left(\frac{M_A^2}{4v^2} \right)_N [(\bar{v}_1 v_2 \cos \theta - \bar{v}_2 v_1)^2 + \bar{v}_1^2 v_2^2 \sin^2 \theta]
 \end{aligned}
 \left. \vphantom{\begin{aligned} V_{CB} - V_N \\ V_{CP} - V_N \end{aligned}} \right\}$$

If N is a minimum, it is the deepest one, and stable against **CB** or **CP**

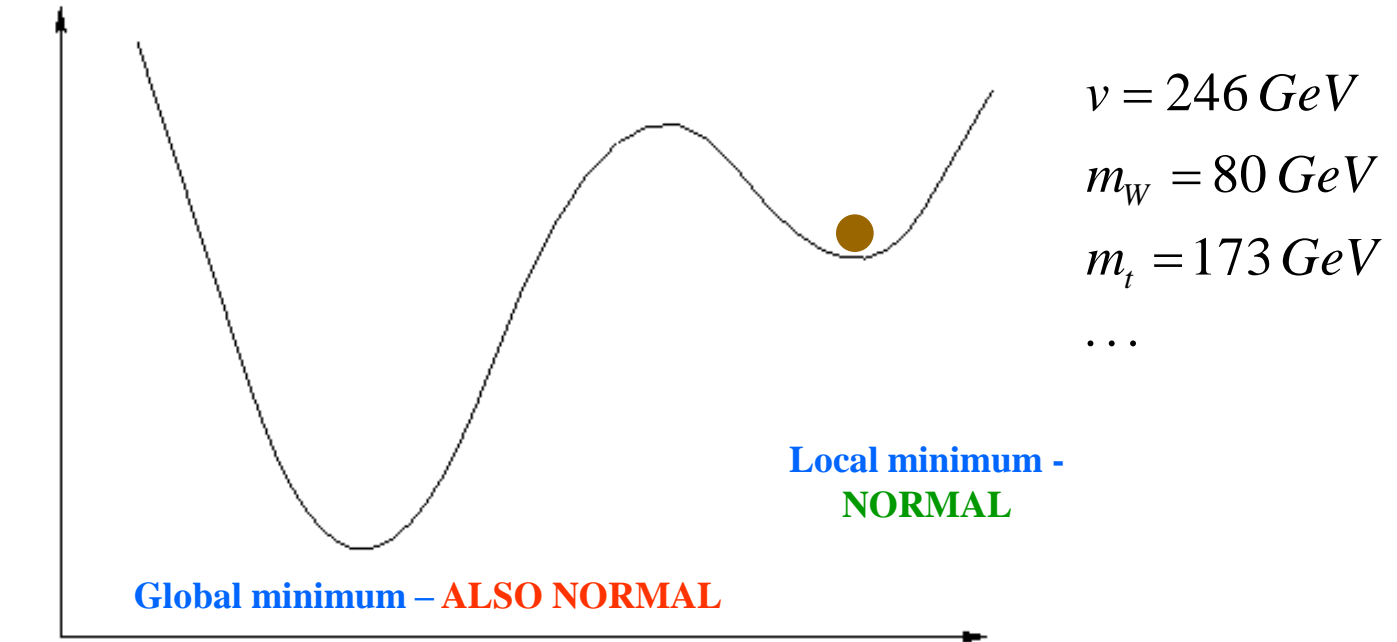
$$V_{N_2} - V_{N_1} = \frac{1}{4} \left[\left(\frac{M_{H^\pm}^2}{v^2} \right)_{N_1} - \left(\frac{M_{H^\pm}^2}{v^2} \right)_{N_2} \right] (v_{1,1} v_{2,2} - v_{2,1} v_{1,2})^2$$

But there is another possibility –

AT MOST TWO NORMAL MINIMA COEXISTING...

Ivanov

So, though our vacuum cannot tunnel to a deeper CB or CP minimum, there is another scary prospect...



$v \neq 246 \text{ GeV}$
 $m_W \neq 80 \text{ GeV}$
 $m_t \neq 173 \text{ GeV}$
...

PANIC VACUUM!!

We consider a simple version of the 2HDM: a softly-broken Peccei-Quinn symmetry

$$\begin{aligned} V_H = & m_{11}^2 \Phi_1^\dagger \Phi_1 + m_{22}^2 \Phi_2^\dagger \Phi_2 - m_{12}^2 \left[\Phi_1^\dagger \Phi_2 + \Phi_2^\dagger \Phi_1 \right] \\ & + \frac{1}{2} \lambda_1 (\Phi_1^\dagger \Phi_1)^2 + \frac{1}{2} \lambda_2 (\Phi_2^\dagger \Phi_2)^2 \\ & + \lambda_3 (\Phi_1^\dagger \Phi_1) (\Phi_2^\dagger \Phi_2) + \lambda_4 (\Phi_1^\dagger \Phi_2) (\Phi_2^\dagger \Phi_1), \end{aligned}$$

Coupling to fermions

MODEL I: Only Φ_2 couples to fermions.

MODEL II: Φ_2 couples to up-quarks, Φ_1 to down quarks and leptons.

Theoretical bounds on 2HDM scalar parameters

Potential has to be bounded from below:

$$\lambda_1 \geq 0, \quad \lambda_2 \geq 0,$$

$$\lambda_3 \geq -\sqrt{\lambda_1 \lambda_2}, \quad \lambda_3 + \lambda_4 - |\lambda_5| \geq -\sqrt{\lambda_1 \lambda_2}$$

Theory must respect unitarity:

(for the case under study, $\lambda_5 = 0$)

$$a_{\pm} = \frac{3}{2} (\lambda_1 + \lambda_2) \pm \sqrt{\frac{9}{4} (\lambda_1 - \lambda_2)^2 + (2\lambda_3 + \lambda_4)^2},$$

$$b_{\pm} = \frac{1}{2} (\lambda_1 + \lambda_2) \pm \frac{1}{2} \sqrt{(\lambda_1 - \lambda_2)^2 + 4\lambda_4^2},$$

$$c_{\pm} = \frac{1}{2} (\lambda_1 + \lambda_2) \pm \frac{1}{2} \sqrt{(\lambda_1 - \lambda_2)^2 + 4\lambda_5^2},$$

$$e_1 = \lambda_3 + 2\lambda_4 - 3\lambda_5$$

$$e_2 = \lambda_3 - \lambda_5,$$

$$f_+ = \lambda_3 + 2\lambda_4 + 3\lambda_5,$$

$$f_- = \lambda_3 + \lambda_5,$$

$$f_1 = \lambda_3 + \lambda_4,$$

$$p_1 = \lambda_3 - \lambda_4.$$

$$|a_{\pm}|, |b_{\pm}|, |c_{\pm}|, |f_{\pm}|, |e_{1,2}|, |f_1|, |p_1| < 8\pi$$

**Under what conditions does the 2HDM scalar potential
have two normal minima?**

If, and only if,

$$m_{11}^2 + k^2 m_{22}^2 < 0$$

$$\sqrt[3]{x^2} + \sqrt[3]{y^2} \leq 1,$$

**Interior of
an astroid**



with

$$x = \frac{4 k m_{12}^2}{m_{11}^2 + k^2 m_{22}^2} \frac{\sqrt{\lambda_1 \lambda_2}}{\lambda_{34} - \sqrt{\lambda_1 \lambda_2}}$$

$$y = \frac{m_{11}^2 - k^2 m_{22}^2}{m_{11}^2 + k^2 m_{22}^2} \frac{\sqrt{\lambda_1 \lambda_2} + \lambda_{34}}{\sqrt{\lambda_1 \lambda_2} - \lambda_{34}}$$

and

$$\lambda_{34} = \lambda_3 + \lambda_4$$

$$k = \sqrt[4]{\frac{\lambda_1}{\lambda_2}}$$

I. P. Ivanov, Phys. Rev. D **75**, 035001 (2007) [Erratum-
ibid. D **76**, 039902 (2007)] [hep-ph/0609018]; *ibid*, **77**,
015017 (2008).

**And out of those two minima,
how can you know whether you are in a panic vacuum?**

Let $D = (m_{11}^2 - k^2 m_{22}^2) (\tan \beta - k)$. **IF D < 0 PANIC!!!**

. Notice that these discriminants which specify the existence of a second normal minimum N'

ARE ONLY BUILT WITH QUANTITIES OBTAINED IN “OUR” MINIMUM.

**Is this at all relevant for phenomenology of the 2HDM?
Must verify what the current data tell us...**

- Generate random values for all potential's parameters, such that $m_h = 125 \text{ GeV}$ (all masses $> 90 \text{ GeV}$, $< 800 \text{ GeV}$, $1 < \tan \beta < 30$).

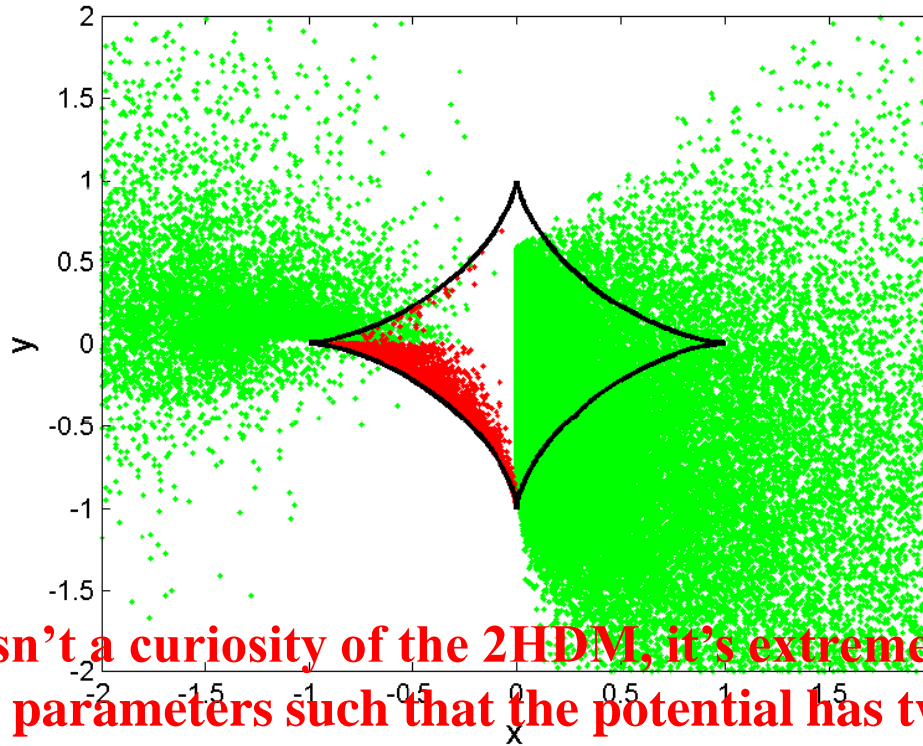
- Ensure the]

- Impose *curi* bounds).

- Calculate a

- Compare w

- Serve shake



lity, etc).

parameter

This isn't a curiosity of the 2HDM, it's extremely simple to choose parameters such that the potential has two minima!!

Inside the astroid: two minima

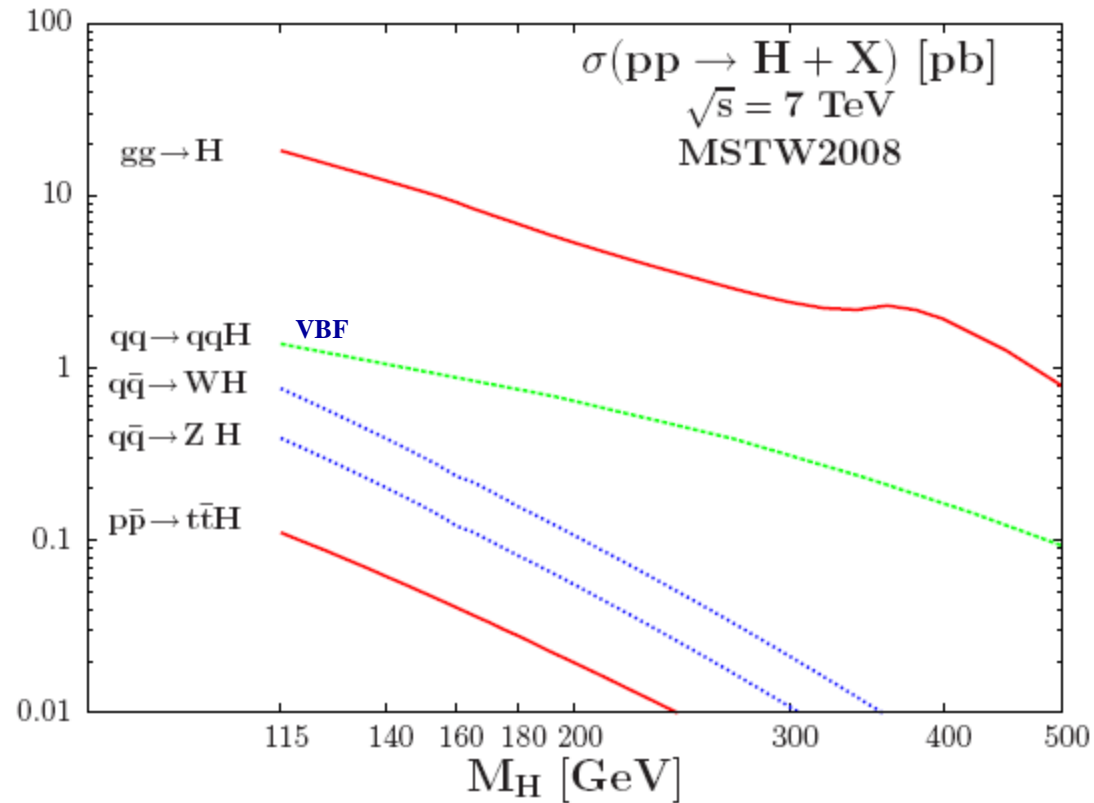
In red: panic vacua points

What we compare to data:

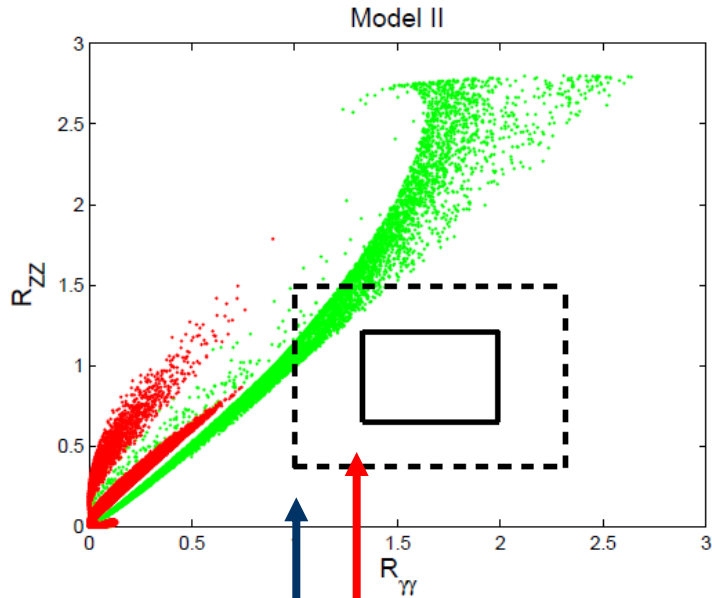
$$R_{XX} = \frac{\sigma^{2HDM} (pp \rightarrow h) \times BR^{2HDM} (h \rightarrow XX)}{\sigma^{SM} (pp \rightarrow h) \times BR^{SM} (h \rightarrow XX)}$$

Plenty of different
production processes
possible at the LHC:

J. Baglio and A. Djouadi, JHEP 03
(2011) 055

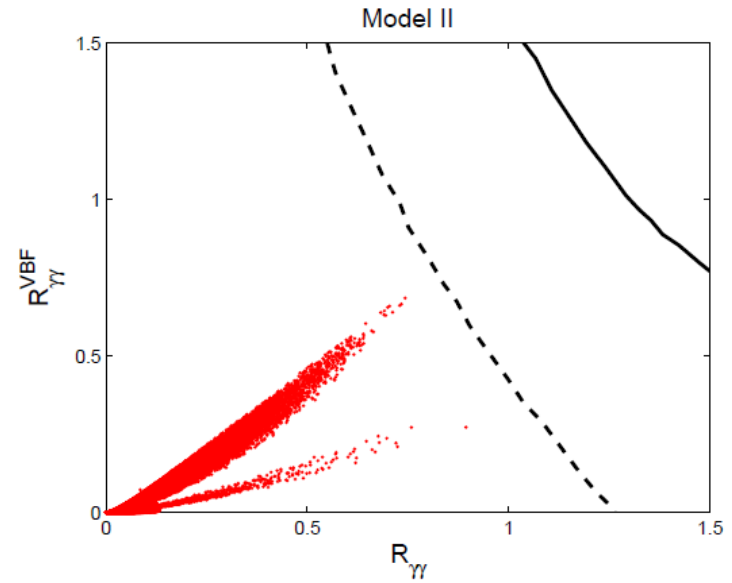


So, what does the LHC tell us? Can we sleep at night?



$1-\sigma$ LHC bounds

$2-\sigma$ LHC bounds

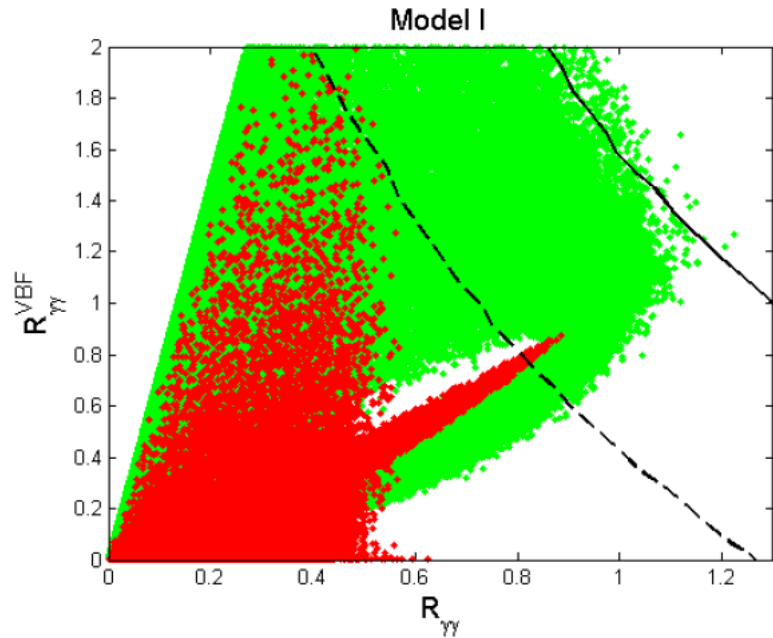
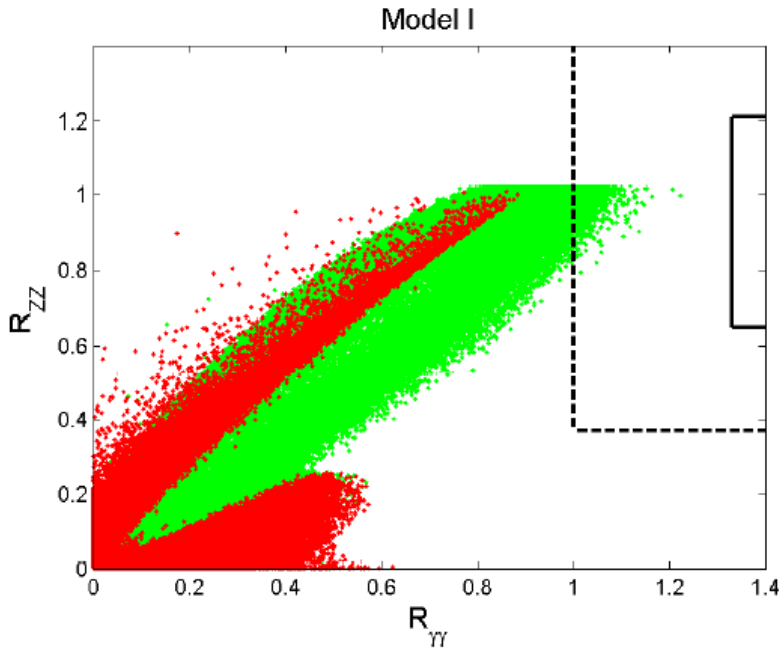


Bounds from: Arbey *et al*, JHEP 1209 (2012) 107;
ATLAS Coll., PLB 716 (2012) 1

No panic in Model II!!!



What about Model I?



Seems to be doing well in R_{ZZ} and $R_{\gamma\gamma}$, but the 2- σ bounds on the VBF rate still include panic vacua.

So maybe the Mayans were on to something after all...



CONCLUSIONS

- **The 2HDM can have two Normal minima.**
- **This situation occurs for MANY choices of parameters, it is not rare.**
- **There is the possibility that our current vacuum is not the global minimum of the potential.**
- **We have developed extremely simple analytical criteria to determine the nature of our vacuum.**
- **Remarkably, the LHC can already tell us a great deal about this situation, and it seems we can sleep safe...**
- **These bounds should be taken into account in the study of the 2HDM, just as much as the bounded-from-below ones are.**

$$\begin{aligned}
m_{12}^2 &= m_A^2 s_\beta c_\beta, \\
\lambda_1 &= \frac{-s_\beta^2 m_A^2 + c_\alpha^2 m_H^2 + s_\alpha^2 m_h^2}{v^2 c_\beta^2}, \\
\lambda_2 &= \frac{-c_\beta^2 m_A^2 + s_\alpha^2 m_H^2 + c_\alpha^2 m_h^2}{v^2 s_\beta^2}, \\
\lambda_3 &= \frac{2m_{H^\pm}^2 - m_A^2}{v^2} + \frac{s_{2\alpha}(m_H^2 - m_h^2)}{v^2 s_{2\beta}}, \\
\lambda_4 &= \frac{2(m_A^2 - m_{H^\pm}^2)}{v^2},
\end{aligned}$$