

QCD+EW corrections to DY in POWHEG

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Recent efforts in QCD \oplus EW N⁽ⁿ⁾LO for DY

$$d\sigma = d\sigma_0 + d\sigma_{\alpha_s} + d\sigma_{\alpha} \\ + d\sigma_{\alpha_s^2} + d\sigma_{\alpha\alpha_s} + \dots$$

Fixed order MC:

- FEWZ: α_s^2 (+ α for neutral DY)

Melnikov & Petriello, PRL 96 (2006) 231803
Li & Petriello, arXiv:1208.5967

- SANC: $\alpha_s + \alpha$

Bardin & al., arXiv:1207.4400

Matching realized using different generators:

- PYTHIA8/HERWIG++ \oplus SANC

Richardson & al., JHEP 1206 (2012) 090

- MC@NLO \oplus HORACE

Balossini & al., JHEP 1001 (2010) 013

NLO matched to PS (POWHEG):

- W_EW-BMNNP

L. B. & al., JHEP 1204 (2012) 037

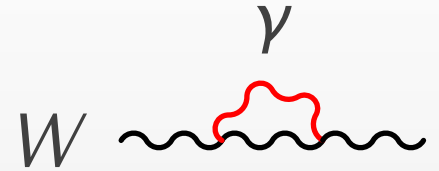
- W_EW-BW

Bernaciack & Wackerroth, PRD 85 (2012) 093003

- Z in progress (preliminary results)

General features of EW radiative corrections

- Unstable particles ($\Gamma \neq 0$) in the loops
→ **avoid gauge invariance violation!**
→ Complex Mass (for DY simpler schemes)



- $m_\gamma = m_q = 0 \Rightarrow$ IR singularities (regularized):

Soft → cancel between R and V

IS collinear → redefinition of PDFs ~~$\log(Q^2/m_q^2)$~~

FS collinear → $\log(Q^2/m_\ell^2)$ are **physical**:

- μ^\pm “always” separated from γ

corrections enhanced

- $e^\pm - \gamma$ recombined (“electromagnetic jet”)

General features of EW radiative corrections

- at $\sqrt{\hat{s}} \sim M_{W/Z}$ QED corrections important

→ QED structure functions

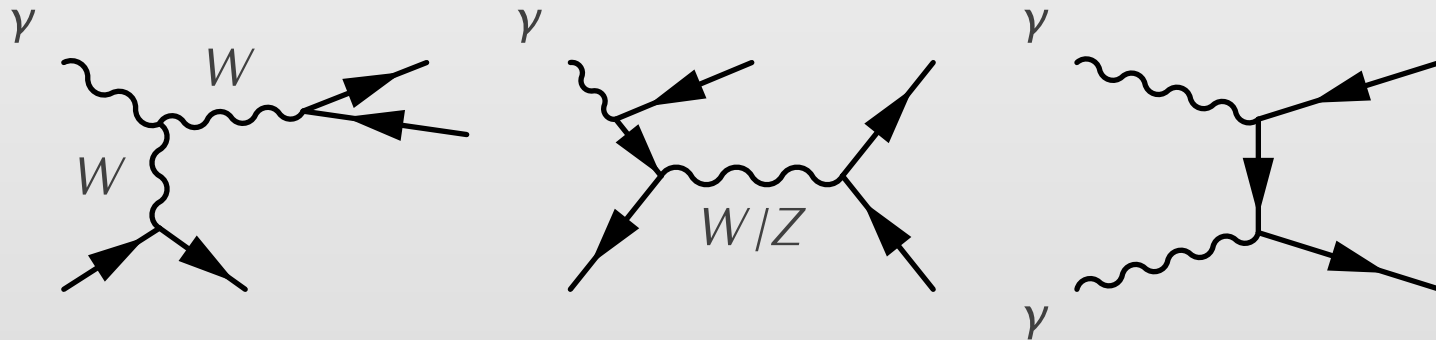
→ QED PS → HORACE / PHOTOS

→ YFS → WINHAC + SANC

Golonka & Was, EPJC45 (2006) 97
Carloni & al., JHEP 0612 (2006) 016

Jadach & Placzek, EPJC29 (2003) 325

- γ induced processes




→ only MRSTQED2004 provides photon distribution function

QCD/EW NLO Cross section

$$\begin{aligned}
 d\sigma_{NLO} = & \left\{ B(\Phi_n) + \overbrace{V^b(\Phi_n)}^{-\infty} \right. \\
 & + \int \overbrace{R(\Phi_n, \Phi_{rad})}^{\infty} d\Phi_{rad} \\
 & \left. + \int \frac{dz}{z} G^{\oplus}(\Phi_{n,\oplus}) + \int \frac{dz}{z} G^{\ominus}(\Phi_{n,\ominus}) \right\} d\Phi_n \\
 & \equiv \boxed{\bar{B} d\Phi_n} \\
 V^b = & V_{EW}^b + V_{QCD}^b \quad R = R_{EW} + R_{QCD} \quad G = G_{EW} + G_{QCD}
 \end{aligned}$$

QCD/EW NLO Cross section

$$d\sigma_{NLO} = \underbrace{\left\{ B(\Phi_n) + \overbrace{V^b(\Phi_n)}^{-\infty} + \int \overbrace{C(\Phi_n, \Phi_{rad})}^{\infty} d\Phi_{rad} \right\}}_{\text{finite}} + \sum_{\alpha_r} \int \underbrace{\left[\overbrace{R^{\alpha_r}(\Phi_n, \Phi_{rad})}^{\infty} - \overbrace{C^{\alpha_r}(\Phi_n, \Phi_{rad})}^{\infty} \right]}_{\text{finite}} d\Phi_{rad}$$



$$+ \left. \left\{ \int \frac{dz}{z} G^{\oplus}(\Phi_{n,\oplus}) + \int \frac{dz}{z} G^{\ominus}(\Phi_{n,\ominus}) \right\} d\Phi_n \right\} \equiv \bar{B} d\Phi_n$$

$$V^b = V_{EW}^b + V_{QCD}^b \quad R = R_{EW} + R_{QCD} \quad G = G_{EW} + G_{QCD}$$

• Subtraction procedure $\rightarrow C = C_{QCD} + C_{QED}$

Virtual part (W_EW-BMNNP)

- V_{EW}^b calculated (**cross-checked**) in different ways

- We chose to use:

- Dittmaier & Krämer - 2002 for W PRD65 073007
- Dittmaier & Huber - 2010 for Z/γ^* JHEP 1001 060

- finite part of dimensional regularization of IR divergent scalar functions

$$(m_q^{in} = m_\gamma = 0, m_\ell \neq 0) \quad \begin{array}{l} \text{Denner, Dittmaier NPB 844:199-242, 2011} \\ \text{Dittmaier NPB 565:69-122, 2000} \end{array}$$

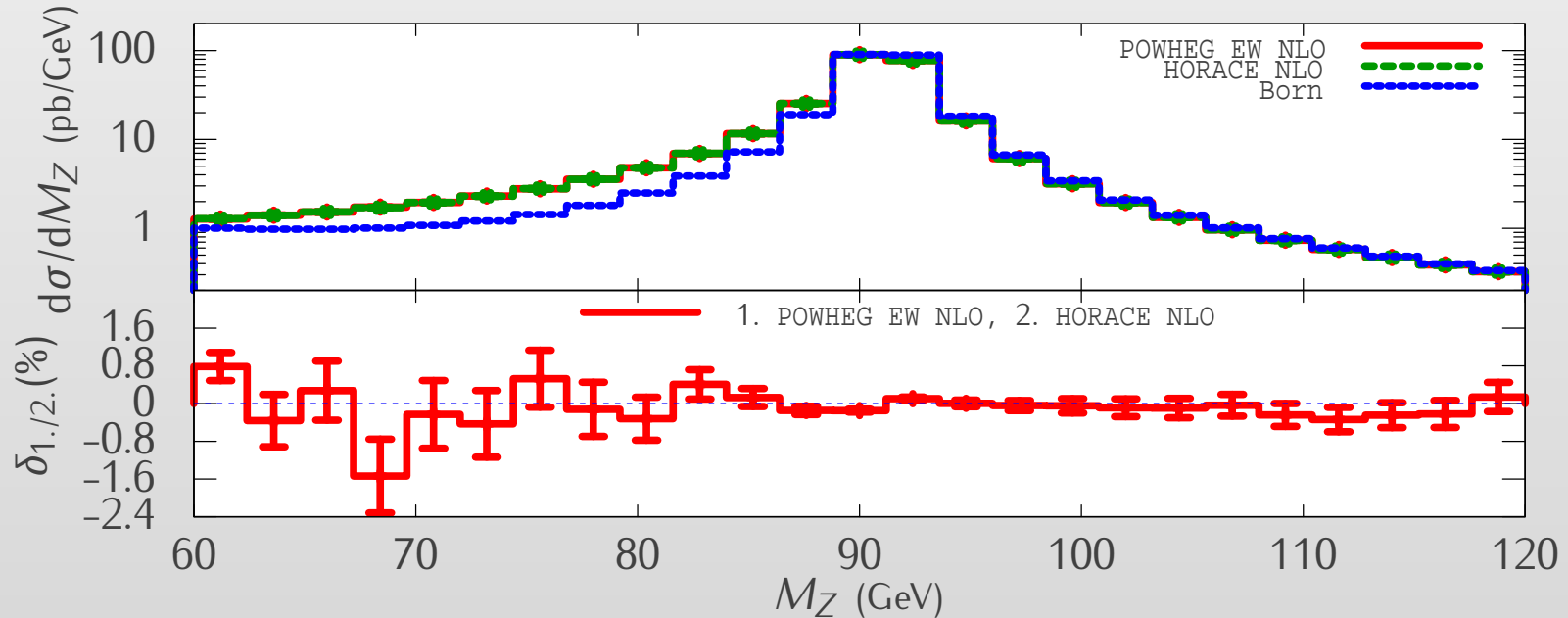
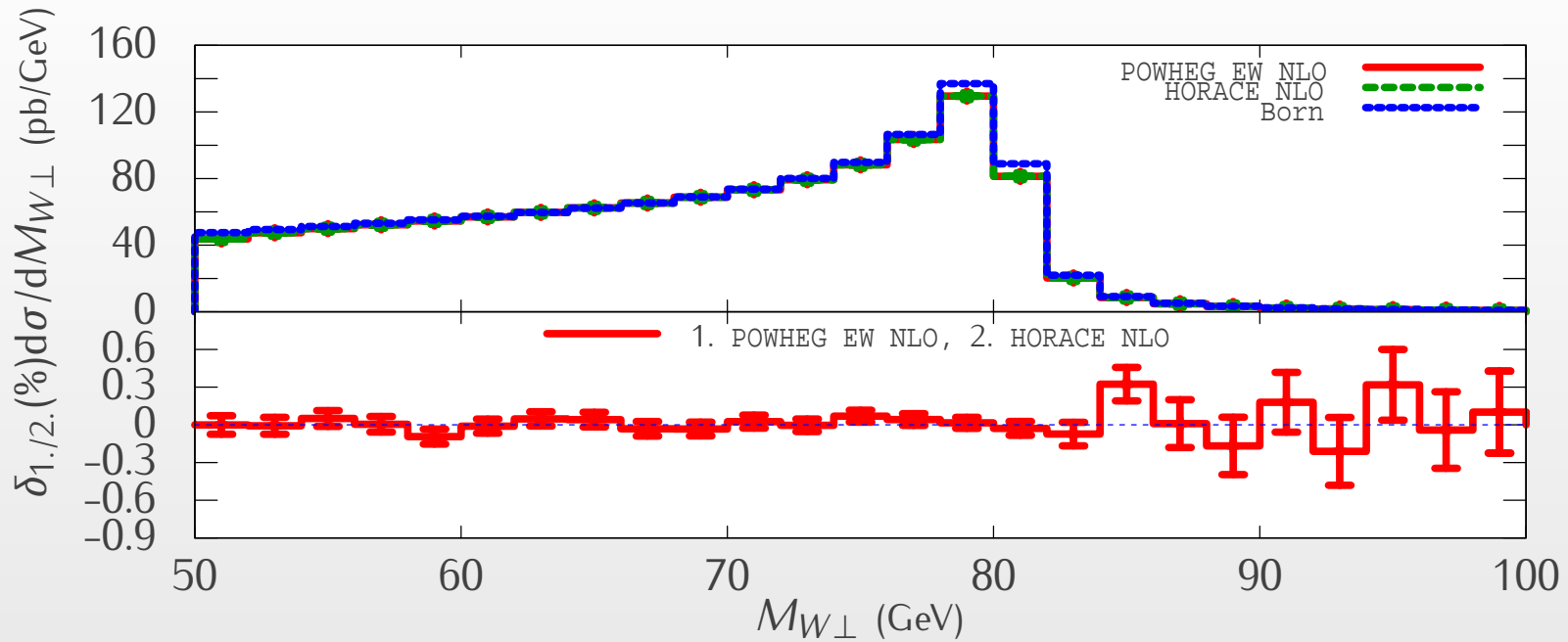
- factorizing out $\mathcal{N} = \frac{(4\pi)^\epsilon}{\Gamma(1-\epsilon)} \left(\frac{\mu_R^2}{Q^2}\right)^\epsilon$

⇒ direct extension of subtraction procedure

C/G_{QCD} in POWHEG with FKS algorithm → C/G_{QED}

NLO Checked with HORACE

Carloni & al., JHEP 10 (2007) 190
Carloni & al., JHEP 12 (2006) 016



Virtual part II (W_EW-BW)

- V_{EW}^b taken from WGRAD2 Baur & Wackerth, PRD 70, 073015, 2004
 - Mass regularization ($m_q \neq 0, m_\gamma \neq 0$)
 - 2-cutoff (δ_c, δ_s) phase space slicing
 $m_q, m_\gamma, \delta_c, \delta_s$ must cancel in physical result

$$d\sigma_{NLO}^{EW} = \left\{ B(\Phi_n) + \underbrace{(V + s)(\Phi_n)}_{\sim \delta_s} + \int_{1,2} \frac{dz}{z} \underbrace{HC(\Phi_n, z)}_{\sim \delta_c} + \int_{\delta_s, \delta_c} d\Phi_{rad} H\bar{C}(\Phi_n, \Phi_{rad}) \right\} d\Phi_n$$

Within POWHEG-BOX:

$$\bar{B} = B + V_{QCD} + V_{s,EW} + R_{QCD} + R_{EW} + G_{QCD} + G_{EW}$$

NLO Checks

	Tevatron, W^+	LHC, W^+	LHC, W^-
WGRAD2	362.55(2) pb	1059.6(1) pb	759.26(3) pb
W_EW-BW	362.4(2) pb	1059.0(5) pb	758.7(8) pb

(δ_s, δ_c)	Tevat, W^+	LHC, W^+	LHC, W^-
0.01, 0.005	362.4(2) pb	1059.0(5) pb	758.7(8) pb
0.01, 0.001	362.4(2) pb	1059.1(7) pb	759.2(5) pb
0.001, 0.0005	362.3(2) pb	1059.4(9) pb	759.4(5) pb
0.001, 0.0001	362.3(2) pb	1059.2(8) pb	759.3(5) pb

Mapping $(\Phi_n, \Phi_{rad}) \leftrightarrow \Phi_{n+1} \xrightarrow{\quad} \text{Radiated particle}$

p_{\perp} of hardest radiation generated according to:

$$\Delta = \exp \left\{ - \sum_{\alpha_r} \int \frac{R^{\alpha_r}(\Phi_{n+1}) \vartheta(p_{\perp}^{\alpha_r}(\Phi_{n+1}) - p_{\perp}^{min}) d\Phi_{rad}}{B(\Phi_n)} \right\}$$

Cross section after the first emission:

Singular zones of R

$$d\sigma = \bar{B}(\Phi_n) d\Phi_n \left[\Delta(\Phi_n, p_{\perp}^{min}) + \sum_{\alpha_r} \Delta(\Phi_n, p_{\perp}^{\alpha_r}) \frac{R^{\alpha_r}(\Phi_n, \Phi_{rad})}{B(\Phi_n)} d\Phi_{rad} \right].$$

p_{\perp} veto on shower emissions \Rightarrow no double-counting

QCD/EW NLO + PS in W_EW-BMNNP

$$\{\alpha_r\} = \{\alpha_r^{QED}, \alpha_r^{QCD}\}$$

$$\Delta_{\max(k_\perp, p_\perp)} \equiv \Delta_{k_\perp}^{EW} \Delta_{p_\perp}^{QCD}$$

2 different p_\perp scales: $\begin{cases} m_\ell & \text{for leptons} \\ \Lambda_{QCD} & \text{for quarks} \end{cases}$

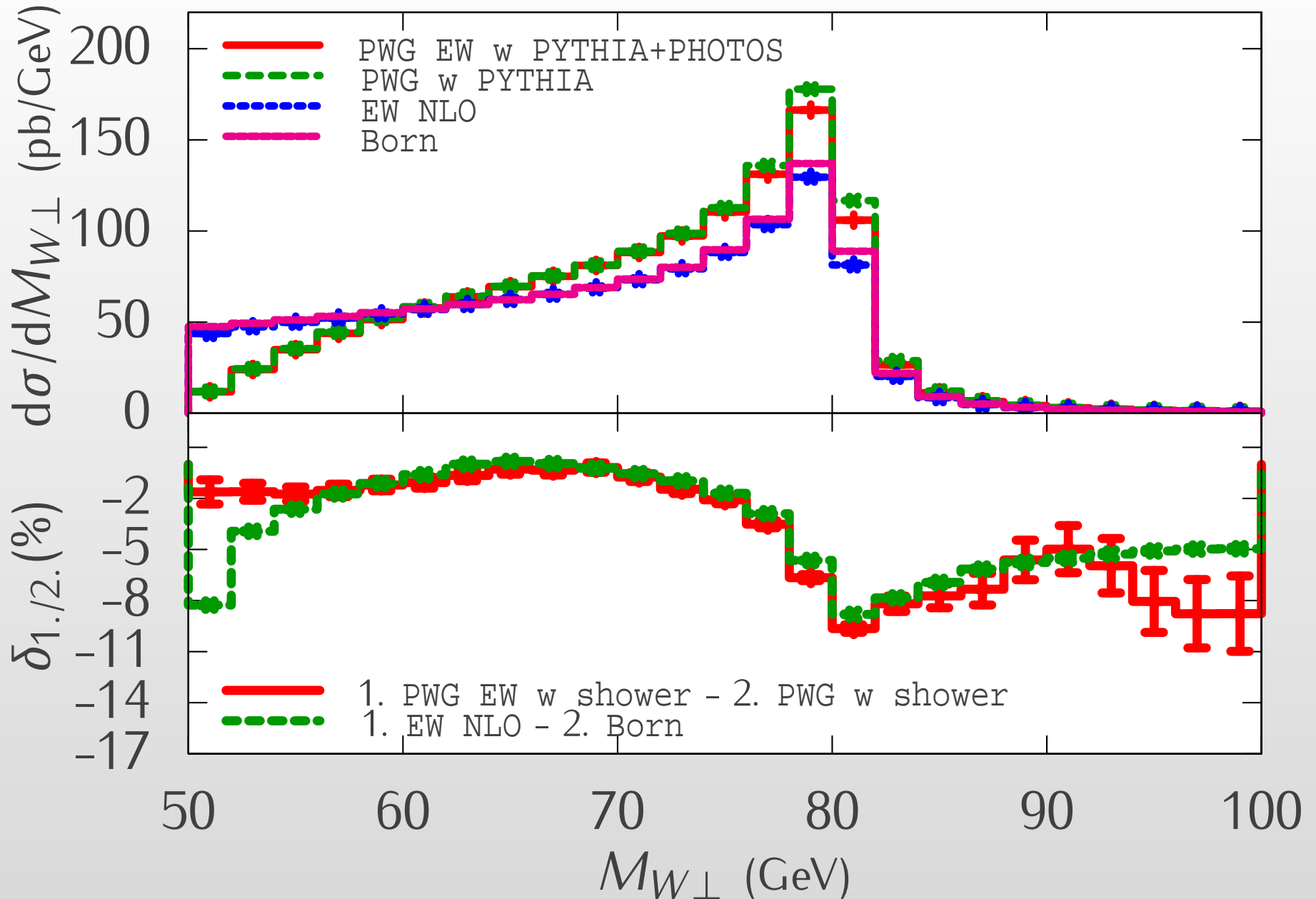
$$\Delta_{p_\perp}^{QCD} = \exp \left\{ - \int \frac{R_{QCD}}{B} \vartheta(p'_\perp - p_\perp) d\Phi'_{rad} \right\}$$

$$\Delta_{k_\perp}^{EW} = \exp \left\{ - \int \frac{R_{EW}}{B} \vartheta(k'_\perp - k_\perp) d\Phi''_{rad} \right\}$$

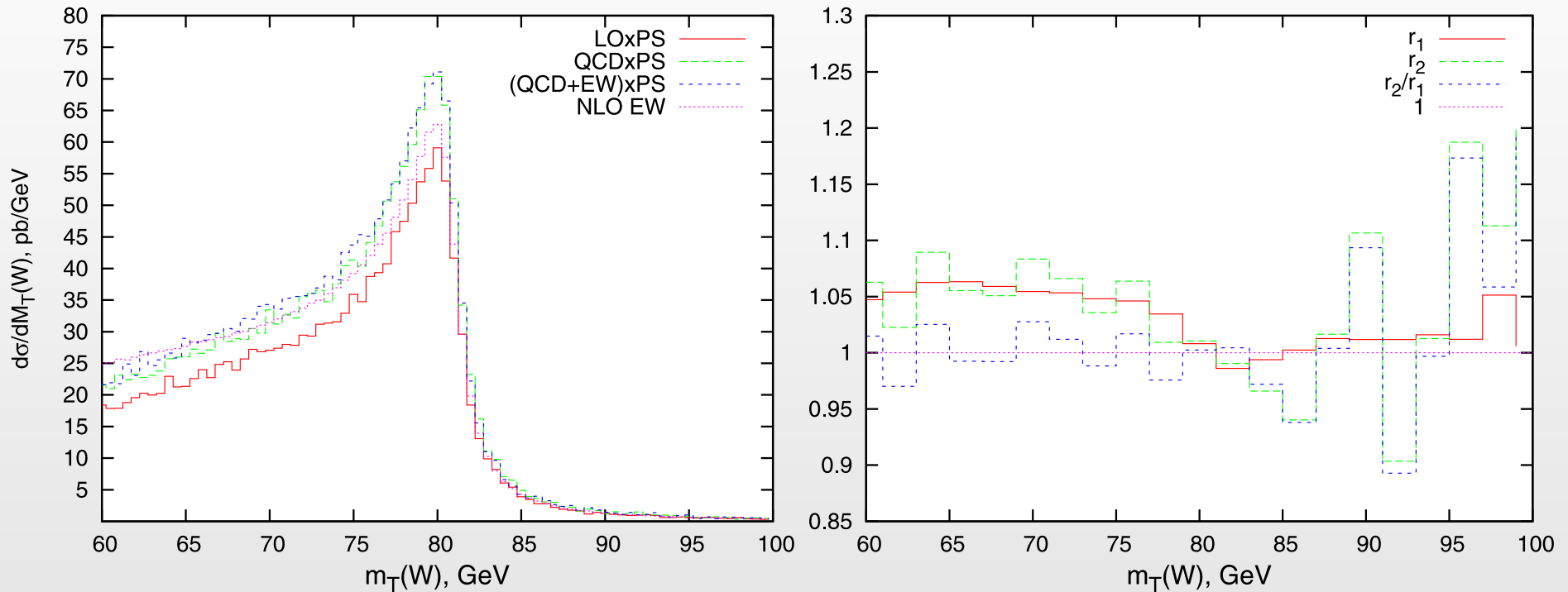
if $\begin{cases} \text{emitted } \gamma/g/q \ (p_\perp \gtrsim \Lambda_{QCD}) \Rightarrow \text{veto } \gamma/g/q \text{ with } p'_\perp > p_\perp \\ \text{emitted } \gamma \quad (p_\perp \lesssim \Lambda_{QCD}) \Rightarrow \text{veto } \begin{cases} g/q \\ \gamma \text{ with } p'_\perp > p_\perp \end{cases} \end{cases}$

Interfaced to PYTHIA / (HERWIG) \oplus PHOTOS

W^+ Transverse Mass at LHC $w_{EW-BMNNP}$

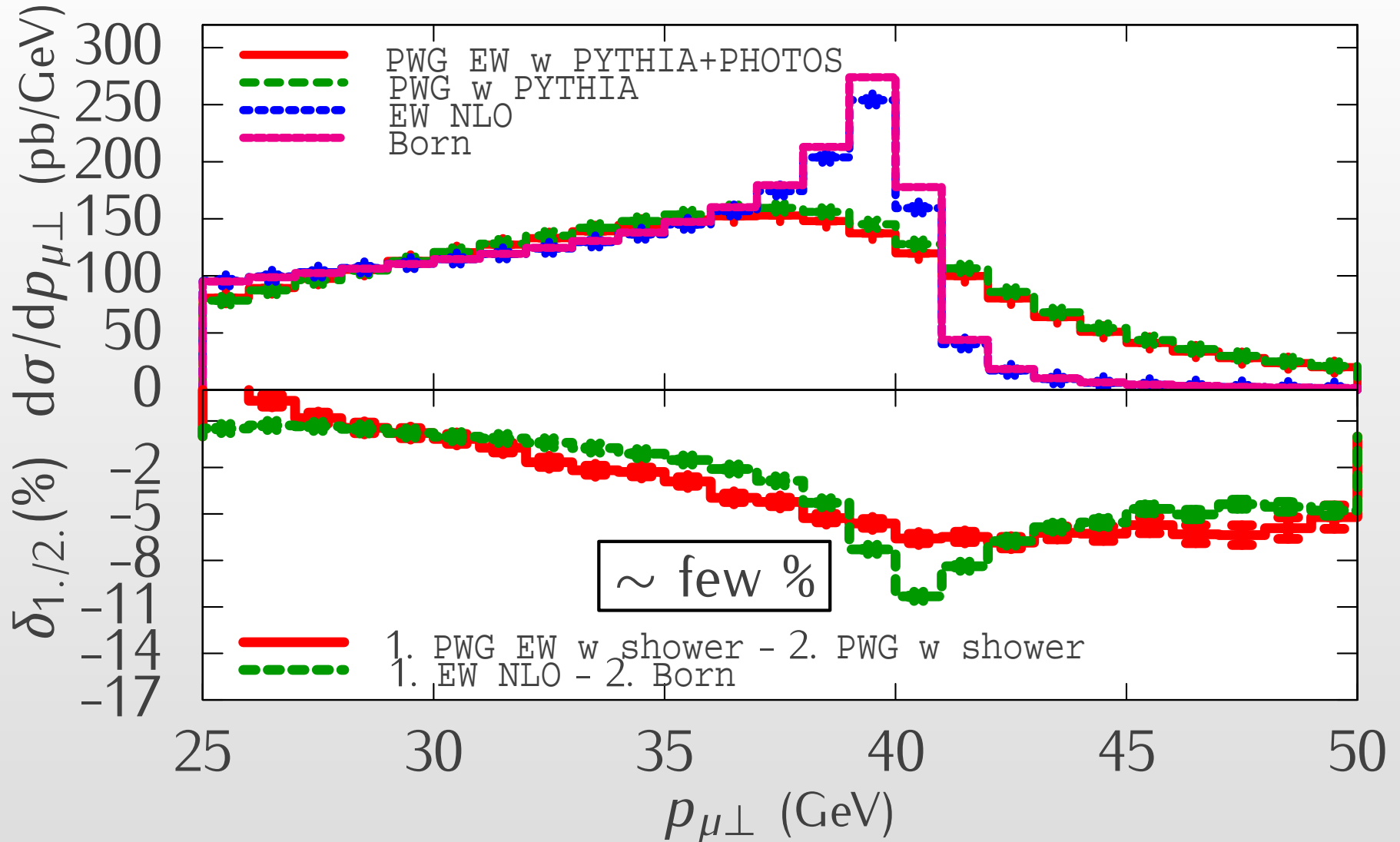


W^+ Transverse Mass at LHC w_{EW-BW}



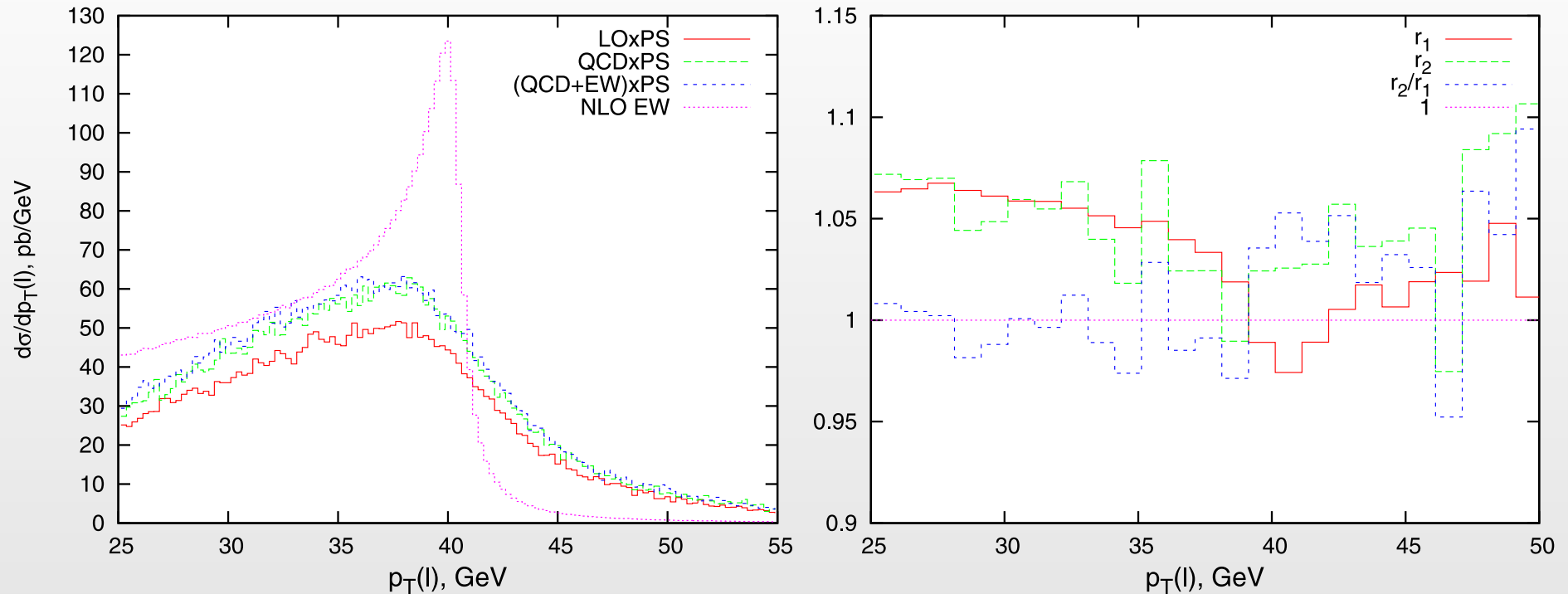
W transverse mass is an observable inclusive
over QCD radiation
 $\Rightarrow \mathcal{O}(\alpha\alpha_s)$ leading logs \sim cancel.

$\mu^+ \perp$ Momentum at LHC w_EW-BMNNP



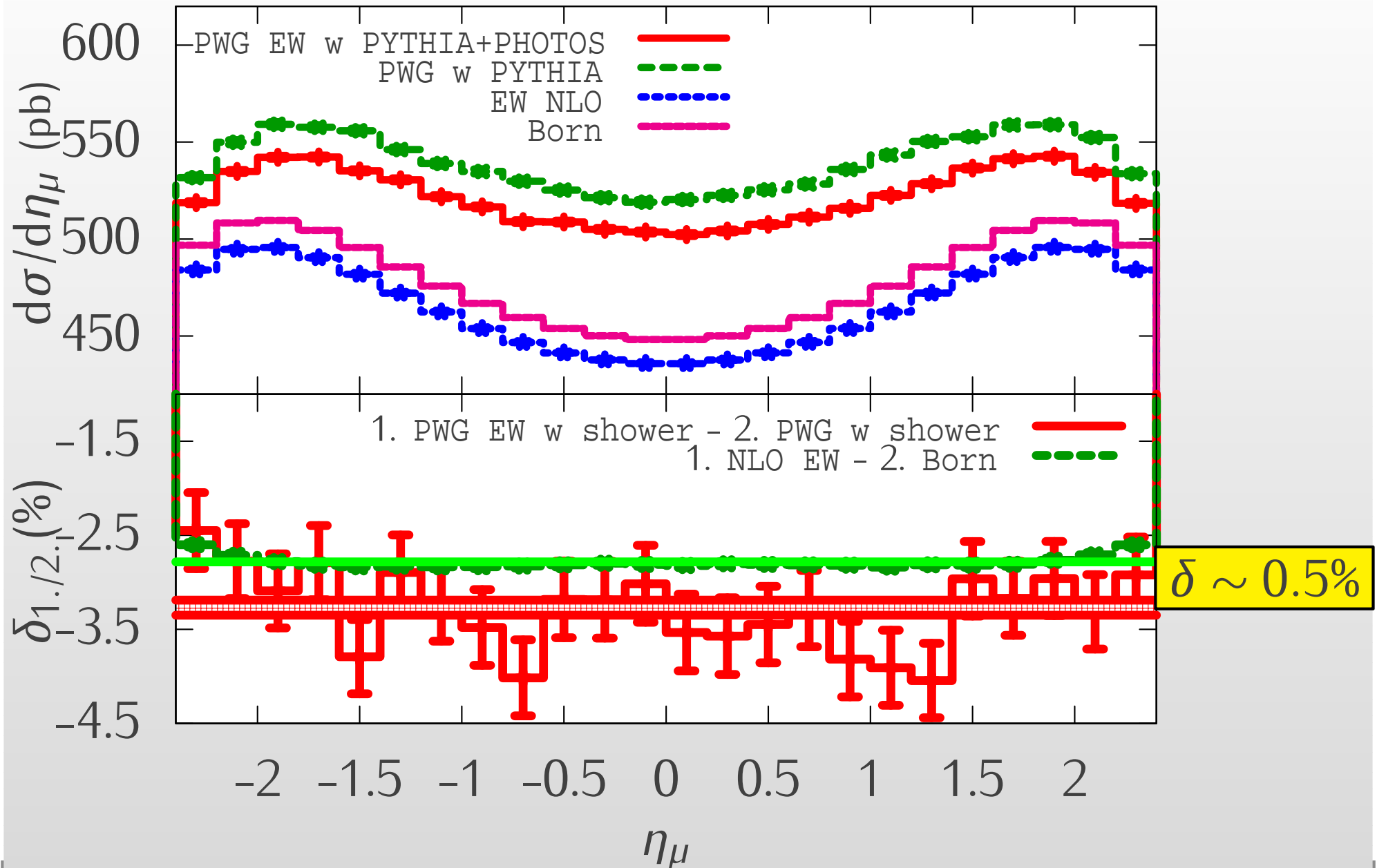
Effects of $\mathcal{O}(\alpha\alpha_s)$ on M_W measurement?

$\mu^+ \perp$ Momentum at LHC w_{EW-BW}



μ transverse momentum isn't an observable inclusive
over QCD radiation
 $\Rightarrow \mathcal{O}(\alpha\alpha_s)$ leading logs important.

W^+ Total Cross Section at LHC $w_{EW-BMNNP}$



W^+ Total Cross Section at LHC w_{EW-BW}

σ_{LO}	1024.0(1) pb	
σ_{NLOEW}	1059.0(5) pb	
	PYTHIA	HERWIG
$\sigma_{QCD \otimes PS}$	1014(3) pb	1027(3) pb
$\sigma_{(QCD+EW) \otimes PS}$	1052(3) pb	1066(3) pb

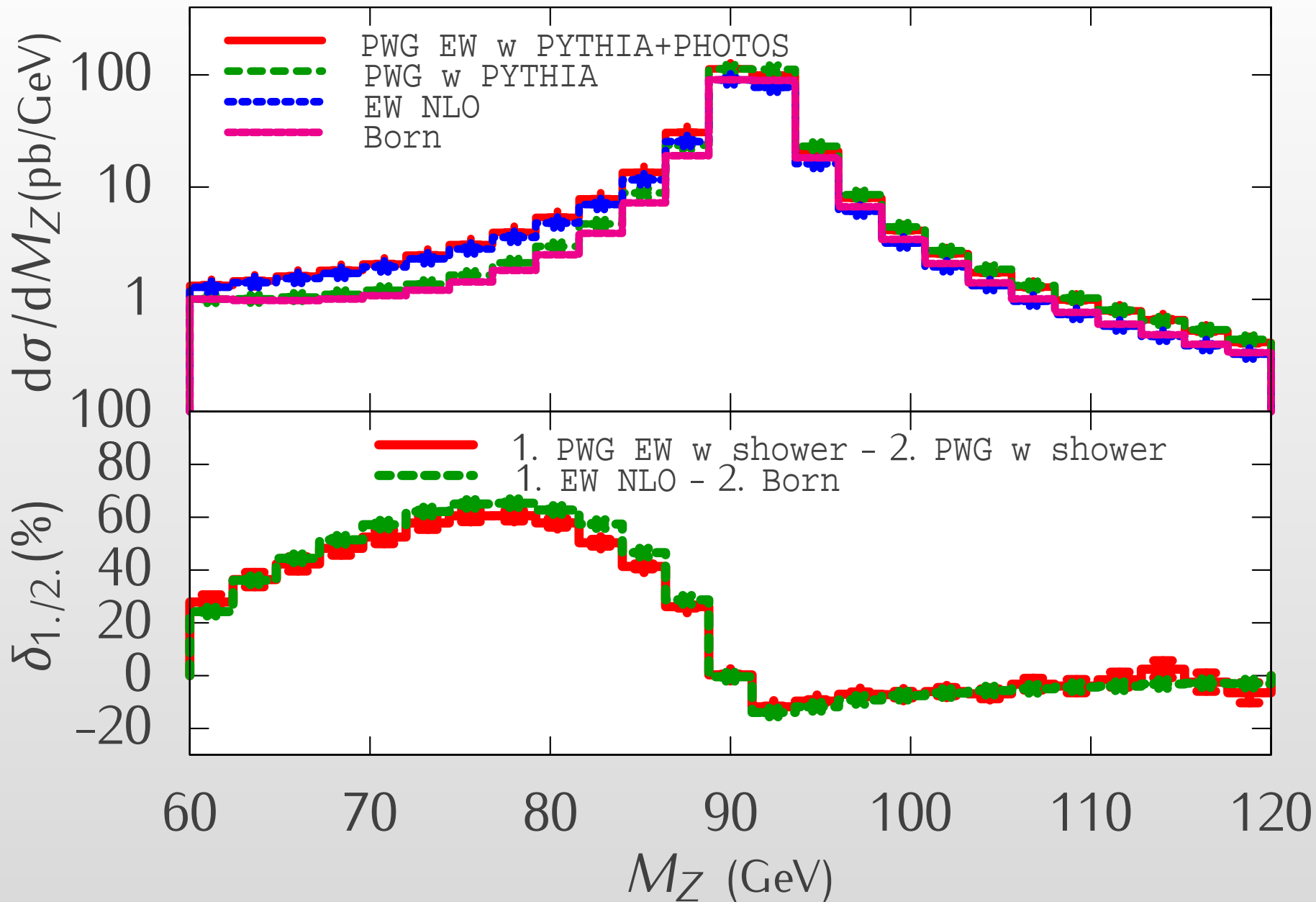
$$\left| \frac{\sigma_{NLOEW} - \sigma_{LO}}{\sigma_{LO}} - \frac{\sigma_{(QCD+EW) \otimes PS} - \sigma_{QCD \otimes PS}}{\sigma_{QCD \otimes PS}} \right| \sim 0.5\%$$

Same order of Balossini & al., JHEP 1001 (2010) 013.

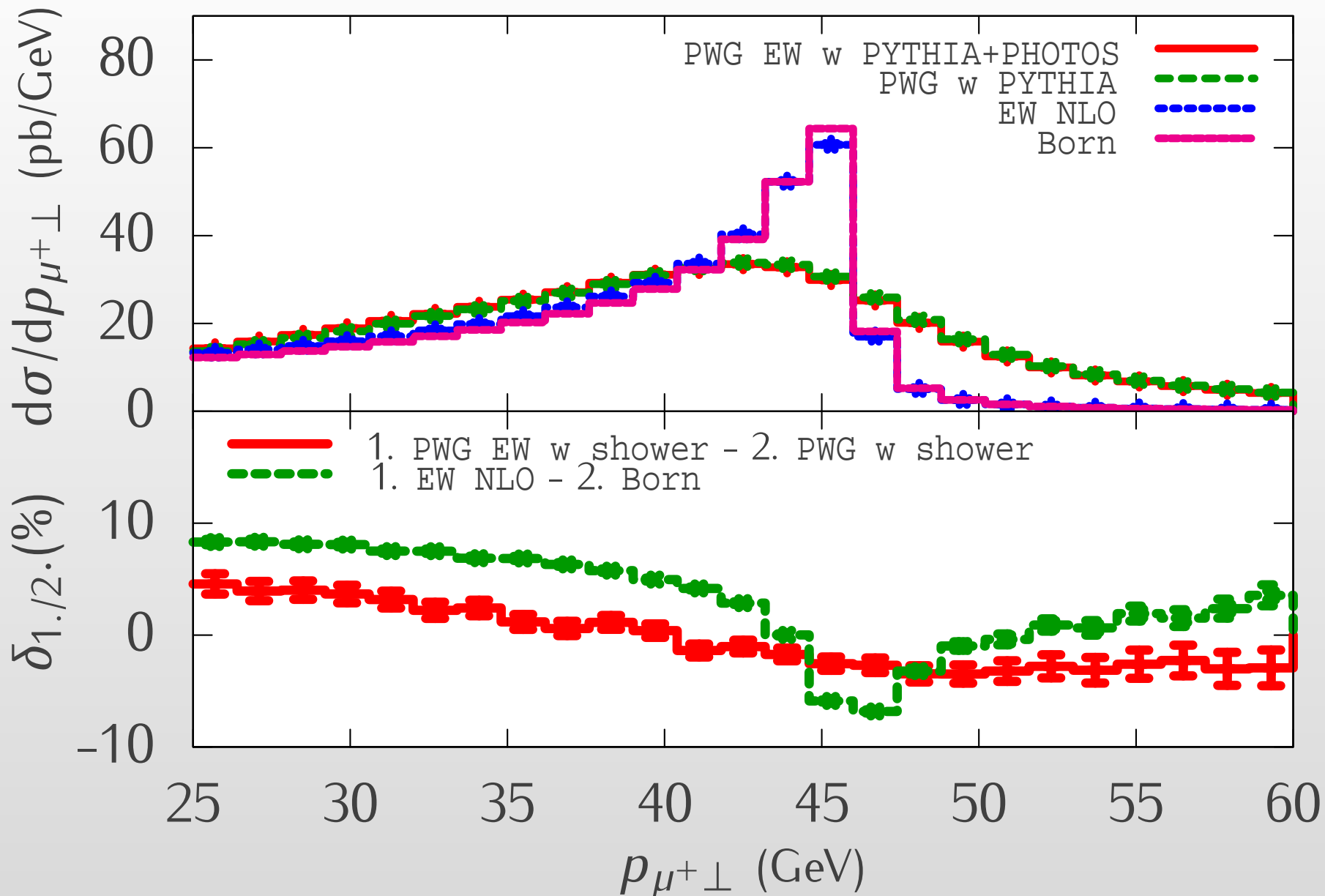
Cuts & parameters different

⇒ Comparison needed.

Z preliminary results W_EW-BMNNP



Z (very) preliminary results W_EW-BMNNP



Summary

POWHEG-BOX for CC DY with QCD & EW at NLO

- normalization with NLO QCD \oplus EW accuracy;
- matched mixed QCD \otimes QED PS;
- leading part of mixed $\mathcal{O}(\alpha\alpha_s)$ corrections
→ important contribution for some observables;
- similar results from different approaches
→ comparison ongoing.

For NC DY ongoing work

Summary

POWHEG-BOX for CC DY with QCD & EW at NLO

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For NC DY ongoing work

THANK YOU!!

Differences? Multiple photon radiation

$$\bar{B}^{W_EW-BMNNP} = \bar{B}^{W_EW-BW} \Rightarrow \text{Normalization OK}$$

$$d\sigma = \bar{B}(\Phi_n) d\Phi_n \left[\Delta(p_{\perp}^{min}) + \sum_{\alpha_r} \Delta(p_{\perp}^{\alpha_r}) \frac{R^{\alpha_r}(\Phi_{rad})}{B} d\Phi_{rad} \right]$$

$$\{\alpha_r\}_{BMNNP} = \left\{ \alpha_r^{QED}, \alpha_R^{QCD} \right\} \quad \{\alpha_r\}_{BW} = \left\{ \alpha_R^{QCD} \right\}$$

W_EW-BMNNP: try to generate γ s as hardest particle
→ consistent to use QED shower

W_EW-BW: do not try to generate γ s
→ double counting if QED shower

- It is much more likely to generate g / q instead of γ

Small effects