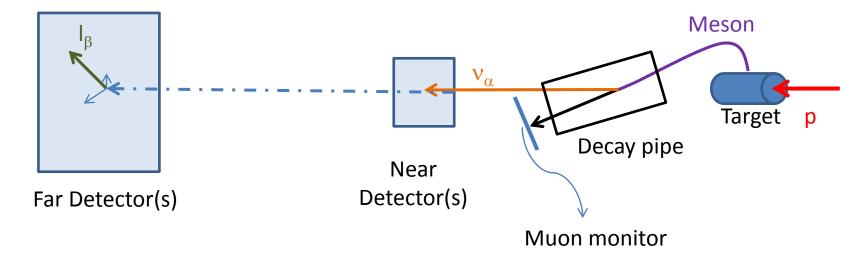
#### Cross Sections: The Beginning And End A Talk In Two Parts

Ryan Terri (QMUL) IoP Half-Day Meeting 7 November 2012

# Generic Long-Baseline Experiment Schematic (Not To Scale)



Protons from beam collide with target

Mesons produced from this collision

Mesons are focused into a decay pipe producing neutrinos & their lepton pair Near detector(s) characterize the beam & try to measure relevant cross sections Far detector(s) detect interactions from beam neutrinos after oscillations have occurred Use all of this information to extract oscillation parameters (e.g.  $\theta_{13}$ , mass hierarchy)

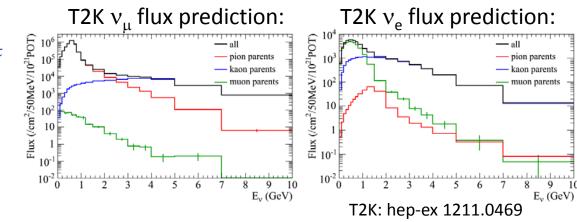
#### HADRON PRODUCTION

# **Meson Production**

Normally want a pure  $v_{\mu}$  beam in long- baseline (LBL) experiments to study  $v_{\mu} \rightarrow v_{e,\mu,\tau}$ 

Many contributions to the flux:

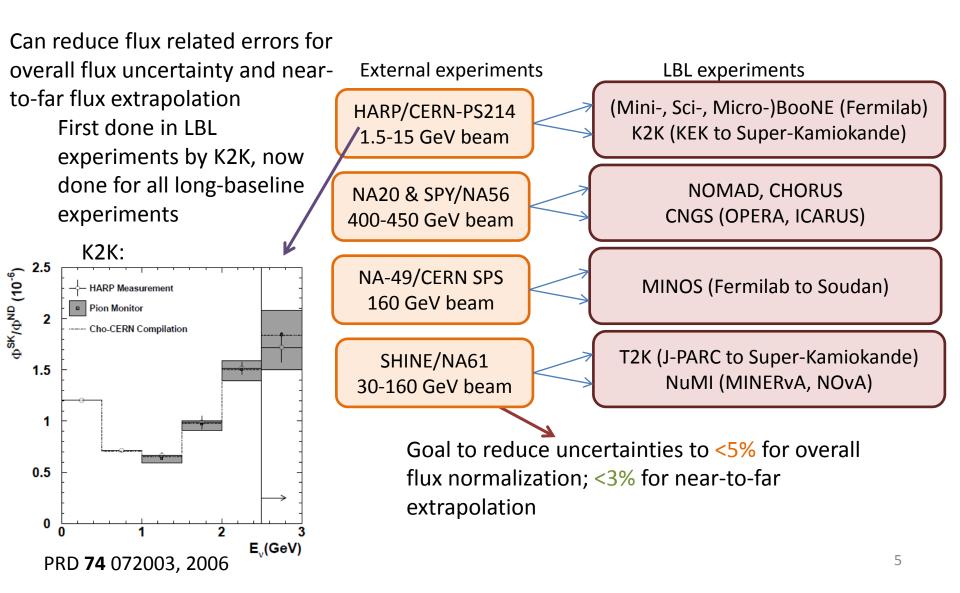
```
\pi \rightarrow \mu \nu_{\mu}
K^{\pm} \rightarrow \mu \nu_{\mu} \text{ or } K^{\pm}_{(e \text{ or } \mu)3}
K^{0} \rightarrow \pi \mu \nu_{\mu} \text{ or } \pi e \nu_{e}
\mu \rightarrow e \nu_{\mu} \nu_{e}
```



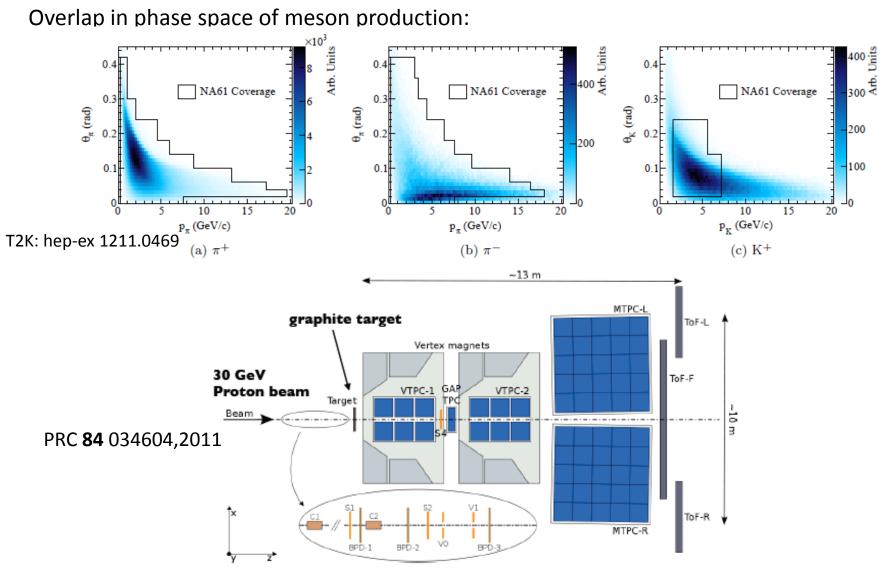
How do we measure these contributions to the flux?

- Muon monitors at end of decay pipe
  - Energy-dependent (µs must make it to the monitor)
- Near detector
  - Want them to be capable of detecting more than one neutrino flavour and, preferably, the antiparticles
- External Experiment
  - Does not need to worry about effects from target station, horn, and decay pipe
  - Focuses only on hadron production

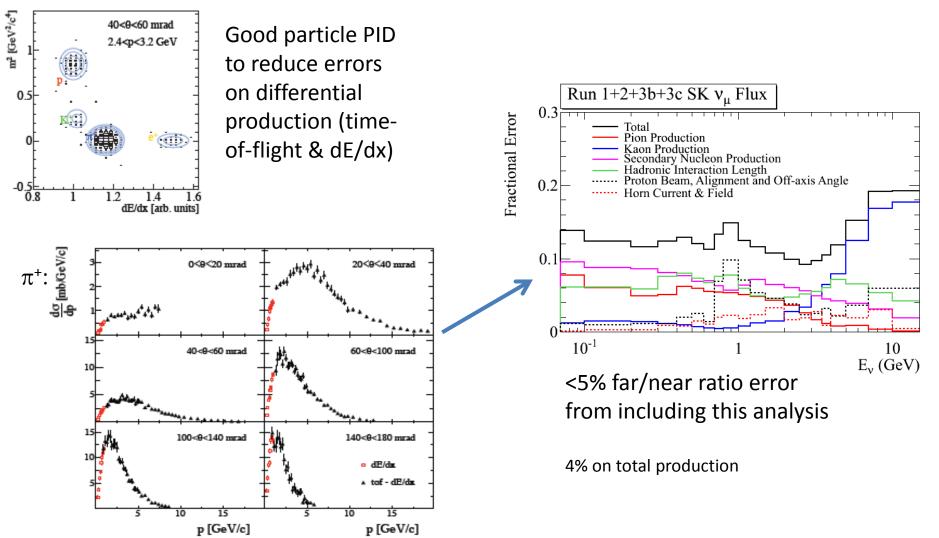
# **External Experiments**



### Example: NA61/SHINE



### Example: NA61/SHINE

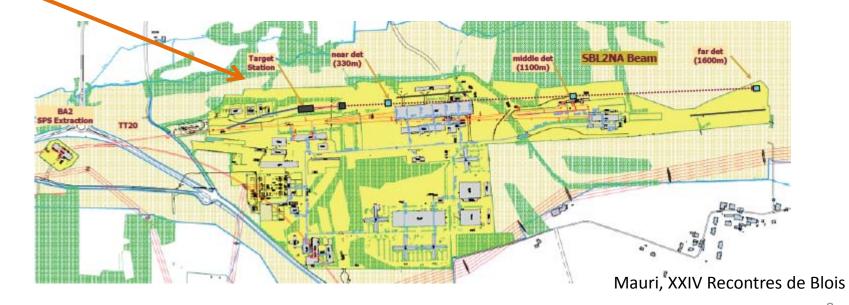


S. Murphy, NuFact 2012

# Who Will Need To Have An External Hadron Production Experiment?

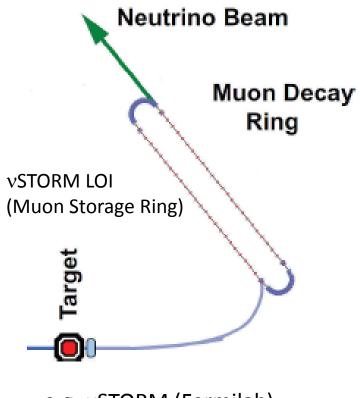
Or: who will be using a conventional beam? T2HK (extension of T2K in Japan w/ Hyper-Kamiokande (HK) as the far detector) (Nakaya for HK) LBNE (USA) (Evans) GLADE (USA) (Evans) LBNO (Europe) (McCauley) NESSIE (Short-baseline sterile v search at CERN)

\*Remember: T2K, MINOS, NOvA, MiniBooNE, ArgoNeuT and others already are using external experiments



## Is There A Way to Avoid External Experiments? Well, yes.

#### Neutrino factories & $\beta$ -beams:



Provide very pure neutrino beams

Little flux error since the decay processes that produce the neutrinos are well-known

Due to high boost (e.g.  $\gamma \approx 1500$ ), highly collimated beam

See talk later on today from K. Long

e.g. vSTORM (Fermilab)

### PART II: NEUTRINO CROSS SECTIONS

# Why is the neutrino-nucleon cross section important?

Far detector: reconstruct neutrino energy to extract oscillation parameters via either outgoing lepton kinematics (assuming charged current quasi-elastic (CCQE)):

$$E_{\nu}^{rec} = \frac{m_{N}E_{\mu} - \frac{1}{2}m_{\mu}^{2}}{m_{N} - E_{\mu} + p_{\mu}\cos\theta_{\mu}}$$

or calorimetry, though moving into using only lepton kinematics for analyses

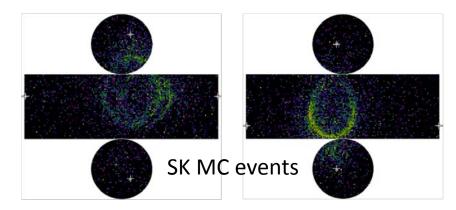
Near detector: measure processes contributing to signal and/or background, or just to figure out what's happening in this energy range ignoring neutrino oscillations

#### Signal & background processes both need to be well-understood

 $v_{\mu} \rightarrow v_{\mu}$ : Signal:  $v_{\mu}$ CCQE interaction Some backgrounds:  $v_{\mu}$ CC1 $\pi^{\pm}$ , NC1 $\pi^{\pm}$ , anti $v_{\mu}$ CC

 $\nu_{\mu} \rightarrow \nu_{e}$ : Signal:  $\nu_{e}$ CC

Some backgrounds: beam  $v_e$ CC, NC1 $\pi^0$ 

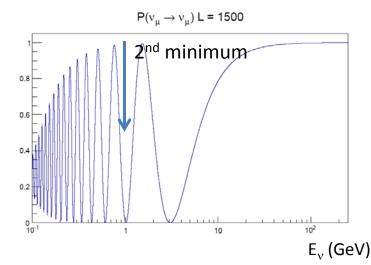


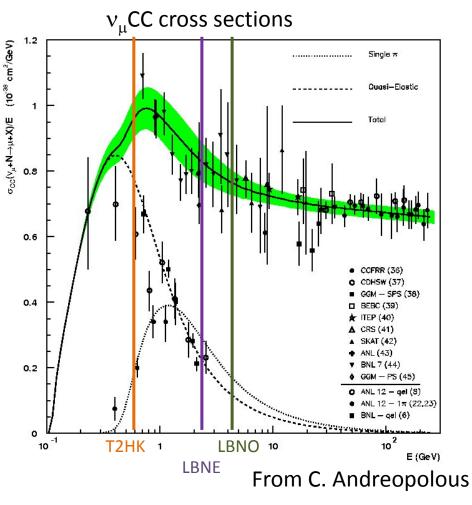
CC = charged current NC = neutral current

### Which Cross Sections Matter?

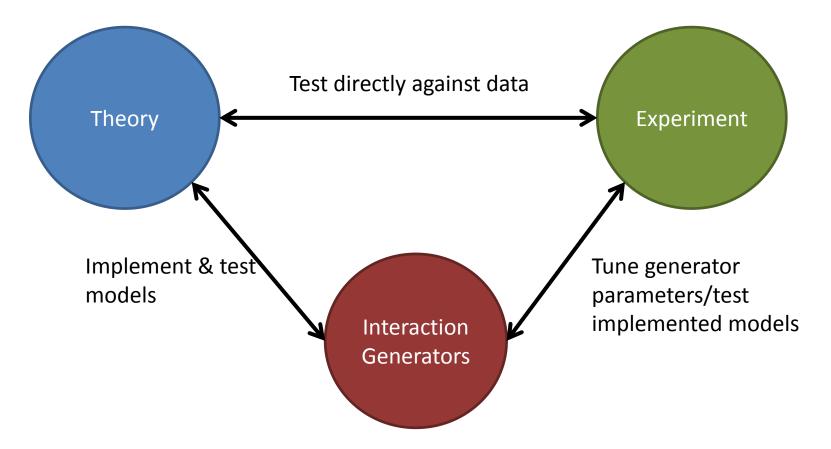
Depends on the experiment: T2HK ~300 km  $\rightarrow E_v \approx 0.6 \text{ GeV}$  (CCQE is dominant) LBNE ~1500 km  $\rightarrow E_v \approx 2.5 \text{ GeV}$ LBNO ~2300 km  $\rightarrow E_v \approx 4.5 \text{ GeV}$ These are for the first  $v_{\mu}$  disappearance minimum

The desire to see the second minimum just means we need to understand the lower energy cross sections as well





# How Is The Data Understood?



Lots of discussion amongst various members of the community needed

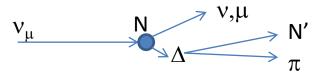
#### Models For A Generic Event Generator

Have base set of models from which we draw our understanding of the physics Most have a variation on this theme:

(Quasi-)elastic scattering: Llewellyn Smith + relativistic Fermi gas (Smith Moniz, Nucl.Phys. B43, 605 (1972))



Resonant production: Rein & Sehgal (Ann.Phys.133, 79 (1981))

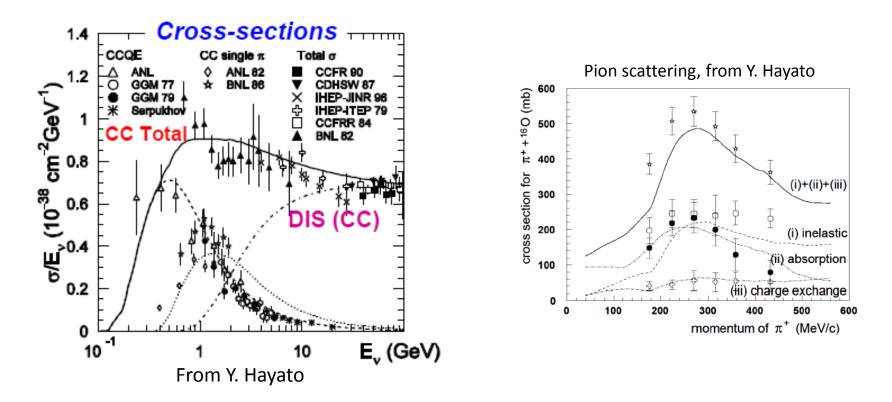


Coherent pion production: Rein & Sehgal (Phys.Lett.B657:207-209,2007)



DIS: GRV98 PDFs w/ Bodek-Yang scaling for x (GRV: Eur.Phys.J. C5, 1998, BY multiple, e.g. hep-ph 1012.0261) Intranuclear effects: cascade model

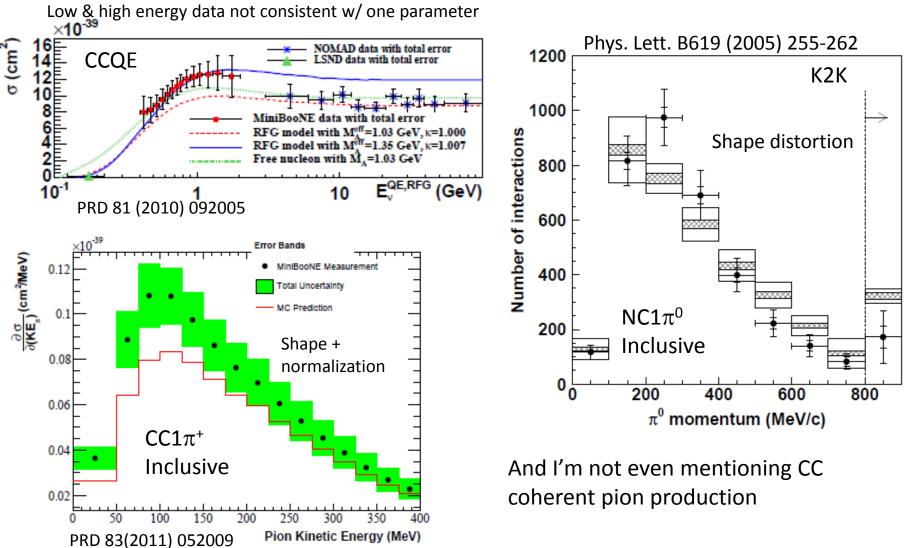
#### Inputs For A Generic Event Generator



These models are not produced in a vacuum

Various parameters tuned to older xsec data sets, electron scattering data, and pion-nucleon scattering

### Problem At Lower Energies ( $E_v \approx 1 \text{ GeV}$ )



# Why The Discrepancy?

Low energy region affected more by nuclear environment Can't get away with idea that neutrino interacts w/ independent nucleon in this energy regime, need better model for this Problem of definition: What do we mean when we describe a certain interaction type? How does that affect our interpretation of the data? Are these models actually complete? Are we missing some type of interaction that can reduce/explain the discrepancy?

Are the models tuned properly?

# Example: CCQE (1/4)

Llewellyn Smith Model (target nucleon rest frame):

$$\begin{split} \frac{d\sigma^{\nu,\ \overline{\nu}}}{dQ^2} &= \frac{M^2 G_F^2 cos^2 \theta_c}{8\pi E_\nu^2} \times \\ \left[ A(Q^2) \mp \frac{(s-u)B(Q^2)}{M^2} + \frac{C(Q^2)(s-u)^2}{M^4} \right] \end{split}$$

s,u = Mandelstam variables A,B,C contain vector & axial form factors

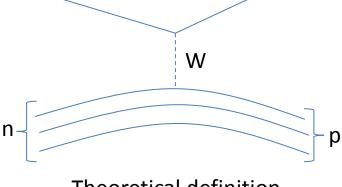
$$\mathcal{F}_A(q^2) = \frac{g_A}{\left(1 + \frac{Q^2}{M_A^2}\right)^2}$$

Dipole approximation, not from first principles

To simulate nuclear targets, use relativistic Fermi gas (RFG) model

 $\nu_{\mu}$ 

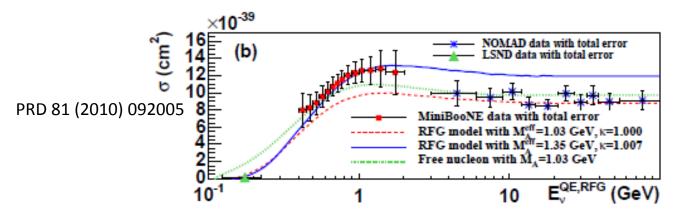
Outgoing proton undergoes nuclear effects in nuclear environment via cascade model, resulting in some events having a pion in the final state



 $\mu^{-}$ 

Theoretical definition

# Example: CCQE (2/4)



Problem is, experiments have to use a different definition of CCQE than theorists  $v_{\mu}+n \rightarrow \mu+p$  (see Feynman diagram)  $v_{\mu}+X \rightarrow \mu+X'+0\pi$  (MB)  $v_{\mu}+X \rightarrow \mu+X'+0\pi + no vertex activity$   $v_{\mu}+X \rightarrow \mu+X'+0\pi+0\gamma+no vertex activity$   $v_{\mu}+X \rightarrow \mu+p+X'+0\pi$  (NOMAD) Etc. These are all based on what is observed by your detector in the final state of the interaction

Possible background comes from  $v_{\mu}+n \rightarrow \mu+\Delta$ ,  $\Delta \rightarrow p+\pi$ ,  $\pi$  is absorbed in the nucleus, p is not observed

MiniBooNE unfolded its data after subtracting backgrounds (some data driven), so it should be closer to the theory definition of CCQE  $$_{
m 19}$$ 

# Example: CCQE (3/4)

Experiment	M <sub>A</sub> Measured (GeV/c <sup>2</sup> )		
World Average (p,n)	1.03±0.03		
K2K SciFi (O)	1.20±0.12		
K2K SciBar (C)	1.14±0.10		
MiniBooNE (C)	1.35±0.17		
MINOS (Fe)	1.19±0.17		
NOMAD (C)	1.05±0.06		

Table from Y. Hayato, NuInt 2012

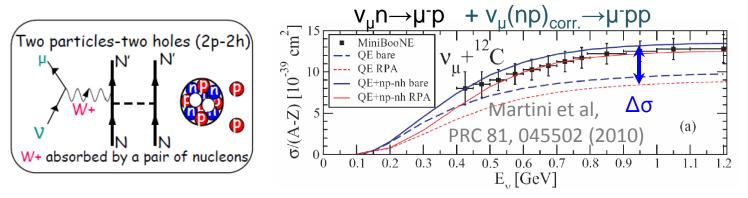
Nuclear environment plays a larger role in cross sections <2 GeV

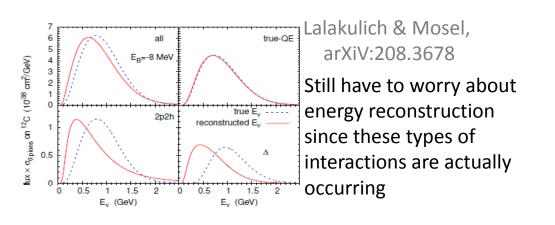
Is M<sub>A</sub> sacred? (Depends on who you ask)

In the **bold** measurements,  $M_A$  becomes a rug with which to sweep our ignorance under and is more of an effective parameter than a fundamental one (that doesn't mean  $M_A$  is physical, though)

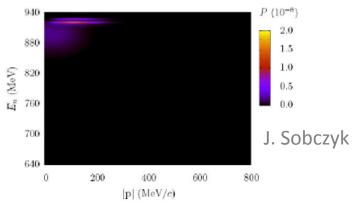
# Example: CCQE (4/4)

Plenty of models have arisen to explain the MiniBooNE CCQE data Most popular is np-nh (lots of work on this in the last few years)



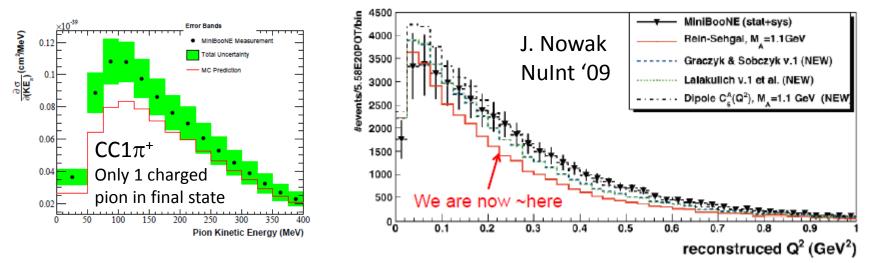


Spectral Function rather than Fermi Gas for nuclear environment

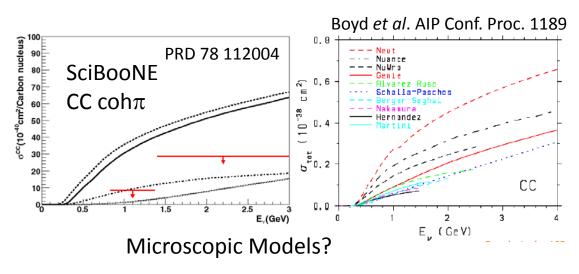


Calculated from first principles ("no free parameters") 21

#### **Other Possible Places For Improvement**

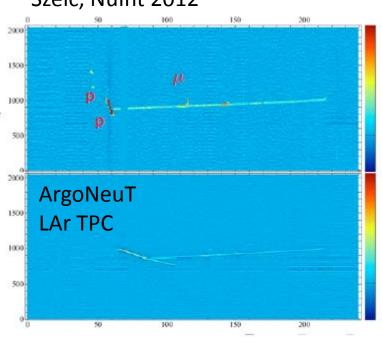


Retuning of the Rein & Sehgal model's parameters should bring better agreement to this (and other  $1\pi$ ) picture(s)



Also use more recent  $\pi A$ scattering data for additional tuning of intranuclear effects

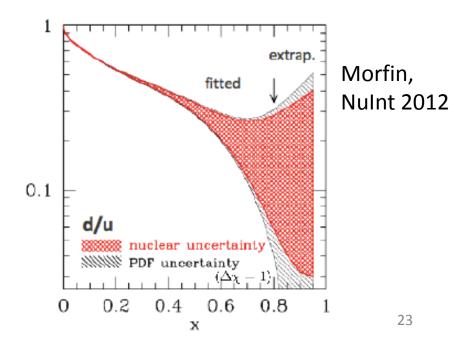
# This Is All Right Now, What About In Szelc, Nulnt 2012 the Future?



2p?!?

With some future experiments in region where DIS turns on, need to understand PDFs in high-x, low Q<sup>2</sup> region

Start looking at a greater number of exclusive final states experimentally Can start to resolve nuclear effects and continue testing various models



# A Path

#### Theorists:

Not only develop model, but provide way to implement for use in experiments Either in generator or vectors that can be put directly into detector simulator Need to move beyond investigating outgoing lepton kinematics (i.e. nuclear effects)

#### Generator providers:

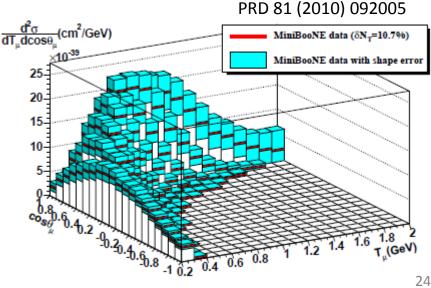
Figure out how to implement models w/ proper outgoing nucleon kinematics before & after final state effects

Validate that it is reasonable

#### **Experimentalists**:

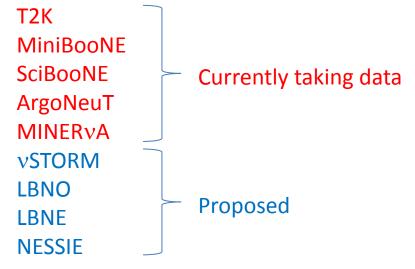
Do best to provide model-independent measurements (or be explicit on the model) Data releases of not only cross section measurements, but also fluxes and complete errors also needed (including correlations between datasets)

i.e. follow & improve on lessons from MB



### Some Current & Future Xsec Experiments

Plenty of experiments will help increase our understanding now and in the near future:



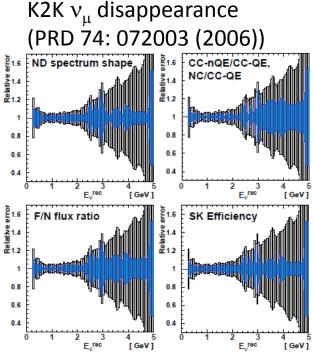
And as you've seen, there's plenty of work for everyone to do, especially when you go back to considering possible impact on oscillation analyses

# Summary

- Greater understanding of the flux through cooperation with external hadron production experiments needed to nail down the neutrino flux
  - True for current method of neutrino beam production, not so much once only one particle type is decaying
- Neutrino-nucleus interactions need more study
  - Lots of data-model discrepancies, some of which directly affects how we understand oscillation signals & backgrounds
  - Lots of ideas on how to better understand them for the next generation of experiments
    - Still need to measure them for each experiment to cancel some systematics as well as add to overall body of knowledge we need for understanding these interactions

#### BACKUPS

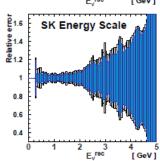
# Which Systematic Errors Am I Concerned About?



MINOS (PRL 106:181801 (2011))

Source of	$\delta(\Delta m^2)$	$\delta(\sin^2(2\theta))$
systematic uncertainty	$(10^{-3}{\rm eV}^2)$	
(a) Hadronic energy	0.051	< 0.001
(b) $\mu$ energy (range 2%, curv. 3%)	0.047	0.001
(c) Relative normalization (1.6%)	0.042	< 0.001
(d) NC contamination (20%)	0.005	0.009
(e) Relative hadronic energy (2.2%)	0.006	0.004
(f) $\sigma_{\nu}(E_{\nu} < 10 \text{ GeV})$	0.020	0.007
(g) Beam flux	0.011	0.001
(h) Neutrino-antineutrino separation	0.002	0.002
(i) Partially reconstructed events	0.004	0.003
Total systematic uncertainty	0.085	0.013
Expected statistical uncertainty	0.124	0.060

#### T2K $v_e$ appearance (PRL 107:041801 (2011))



Error Source	Flux	Near Detector	Near Detector Stats.	Xsec	SK	Total
Size (%); $sin^{2}2\theta_{13} = 0.1$	8.5	+5.6/-5.2	2.7	10.5	9.4	+17.6/ -17.5