

Jakten på ny fysikk

Heidi Sandaker

Standard modellen for fundamentale partikler og vekselvirkninger

Leptons	Strong	Electromagnetic
Tau -1 0 Tau Neutrino	Gluons (8)	Photon
Muon -1 0 Muon Neutrino	Quarks	Atoms Light Chemistry Electronics
Electron -1 0 Electron Neutrino	Mesons Baryons	Nuclei

Quarks

	Electric Charge		
Bottom	-1/3	2/3	Top
Strange	-1/3	2/3	Charm
Down	-1/3	2/3	Up

each quark: R, B, G 3 colors

Gravitational

Graviton ?

Solar system
Galaxies
Black holes

Weak

Bosons (W,Z)

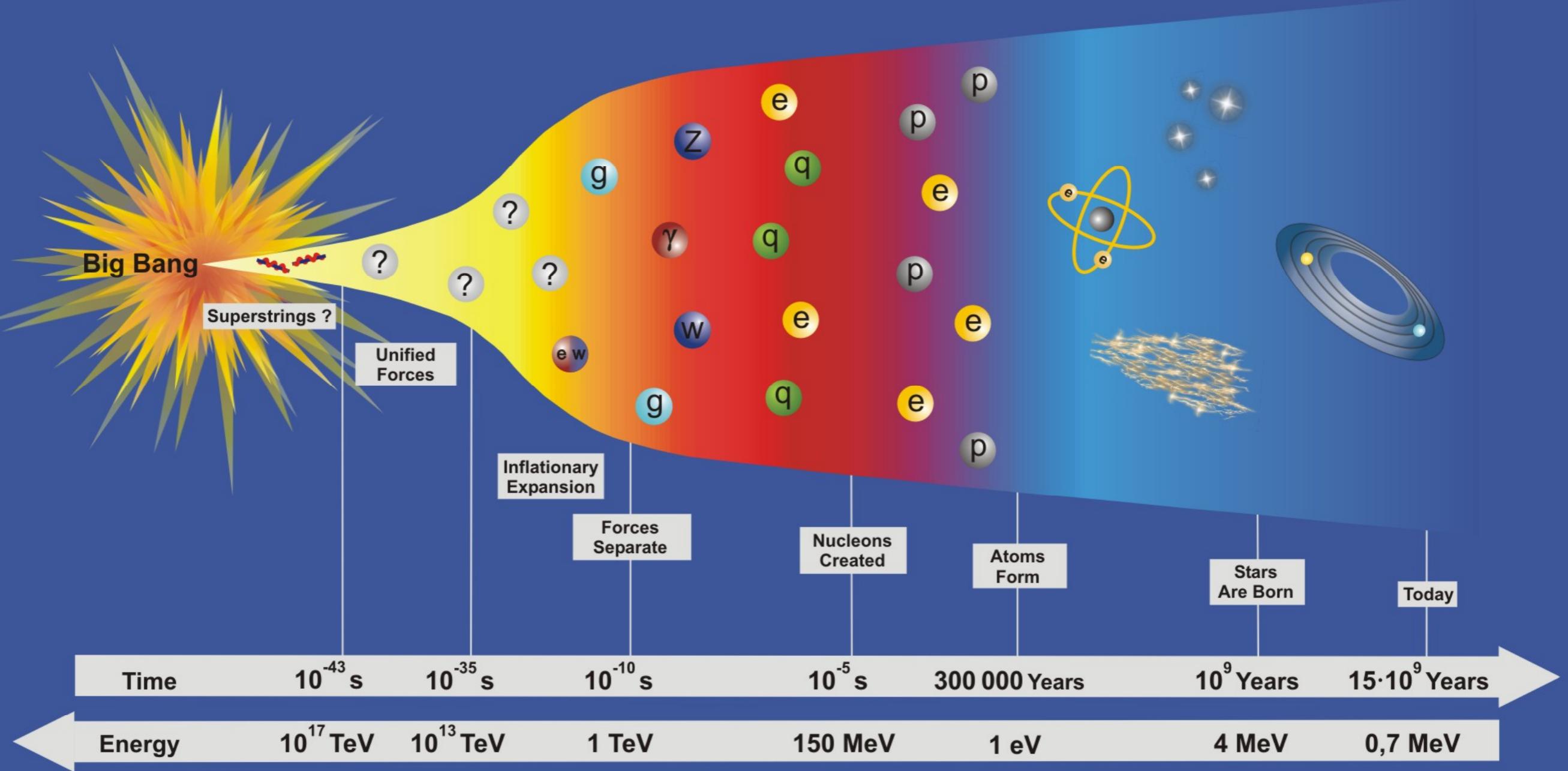
Neutron decay
Beta radioactivity
Neutrino interactions
Burning of the sun

En modell som er testet med veldig høy presisjon !

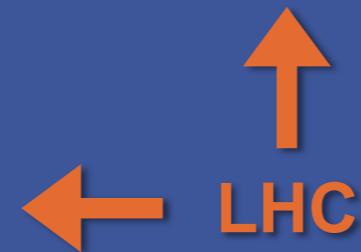
Men Standard Modellen har ikke svar på alt vi lurer på !



The Big Bang - Hva skjedde egentlig like etterpå ?

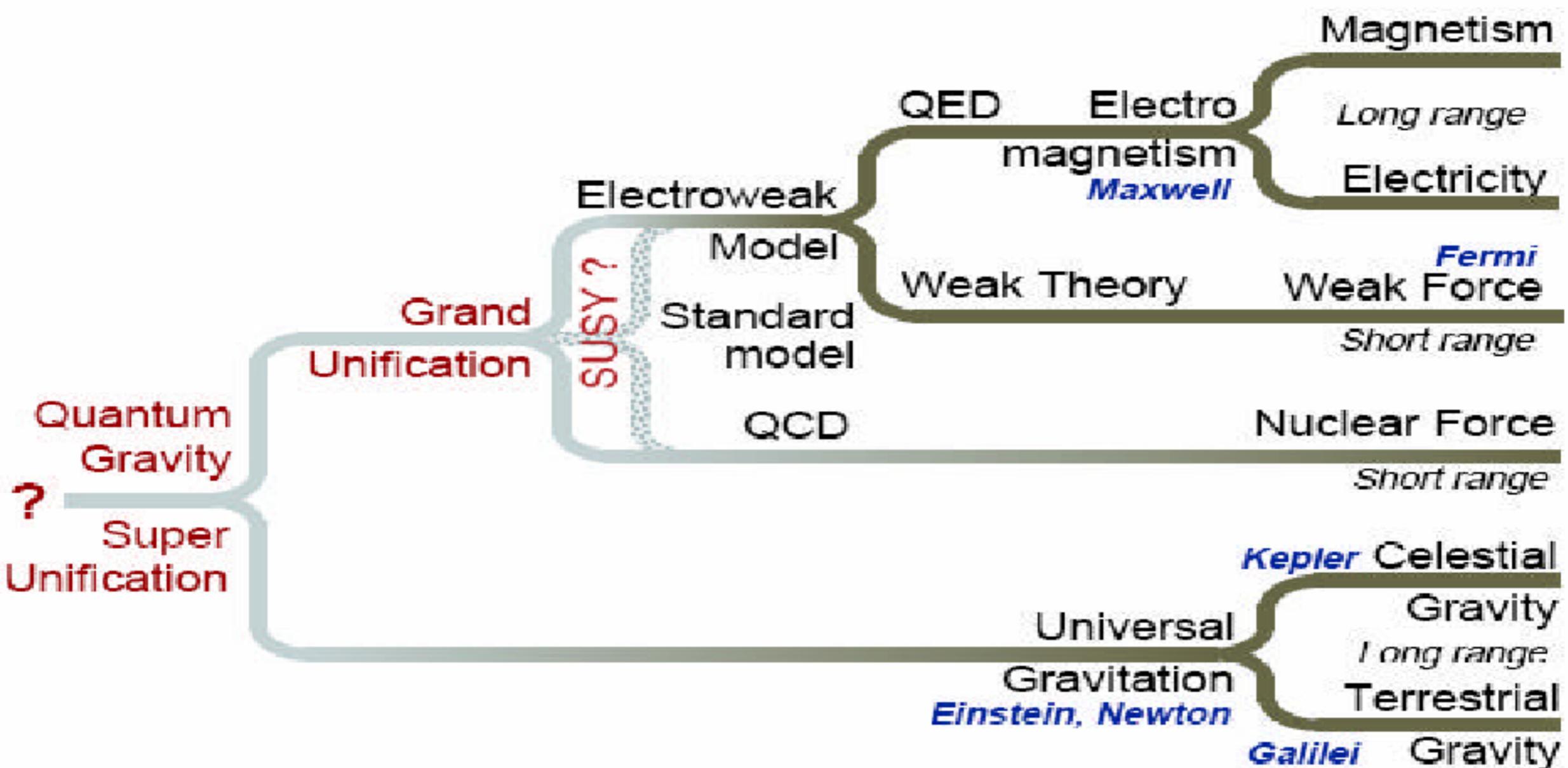
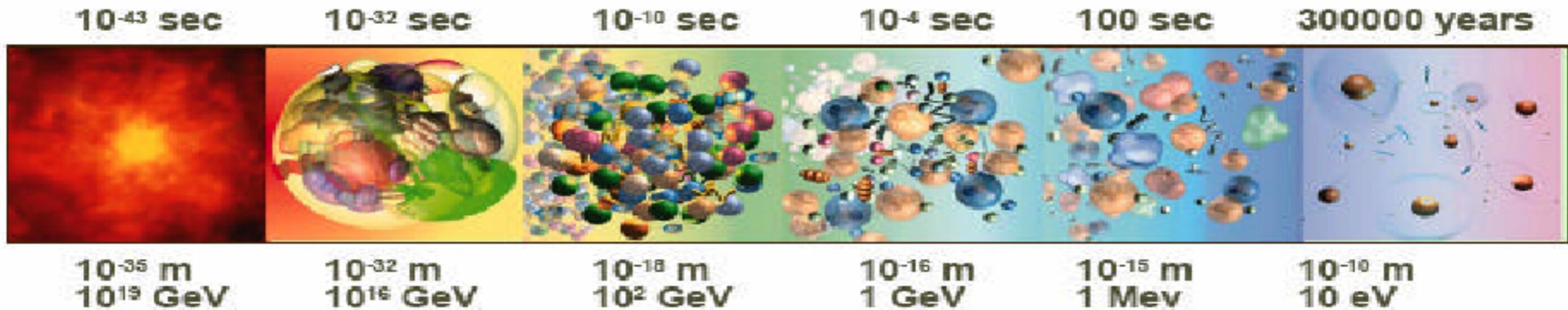


Ekstrapolering via
presisjonsmålinger



Hvorfor ble det ikke laget like
mye materie som antimaterie ?

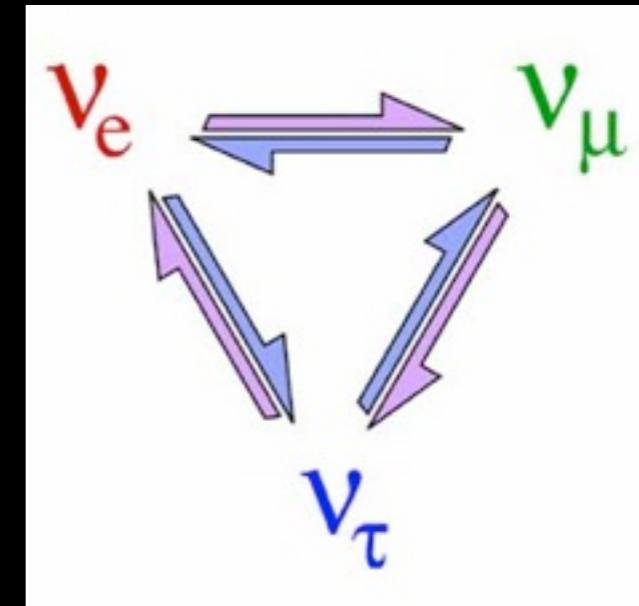
Hvordan ble kvarkene dannet ? ...



Forening av noen krefter mangler i Standard Modellen !

... Og Standard Modellen forklarer ikke hvorfor partikler har masse !

- Standard modellen forutsier at neutrinoer skal være masseløse - i virkeligheten har de masse !
- Faktisk kan ikke Standard modellen forutse noen av massene til de forskjellige partiklene !



Higgs mekanismen :

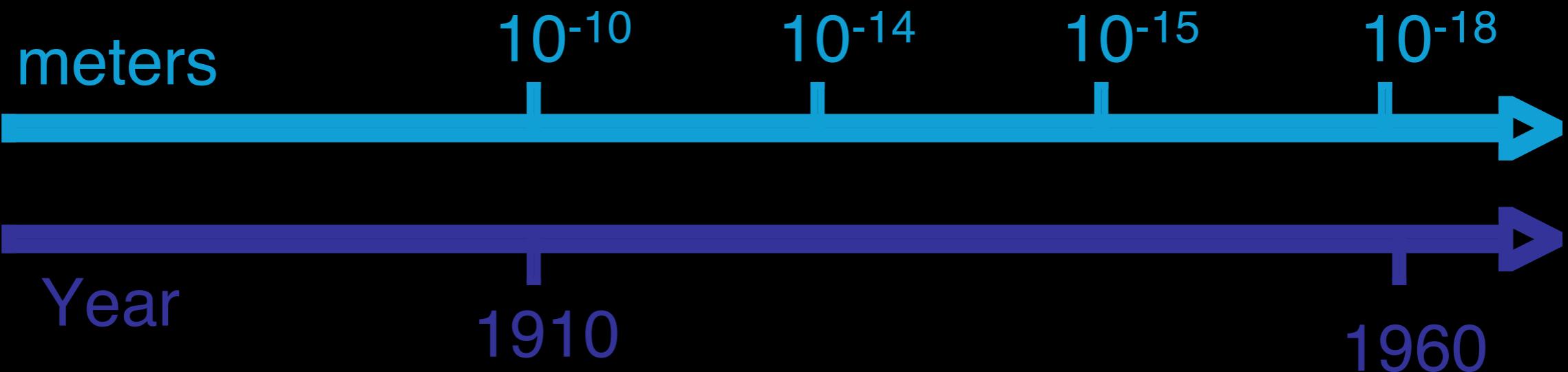
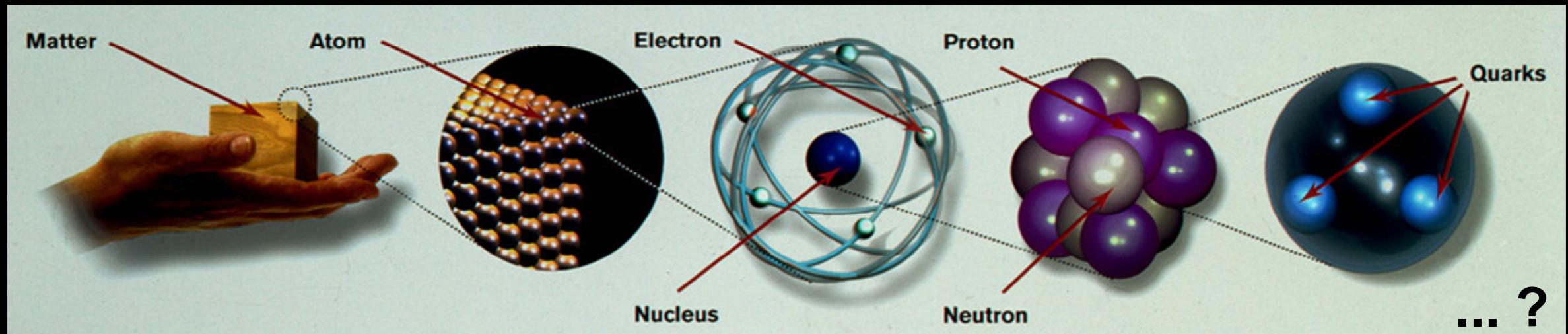
The result of Higgs field in the vacuum is that all particles get masses by gluing into it



Hvis modellen stemmer må det finnes en ny partikkel, **Higgs partikkelen !**

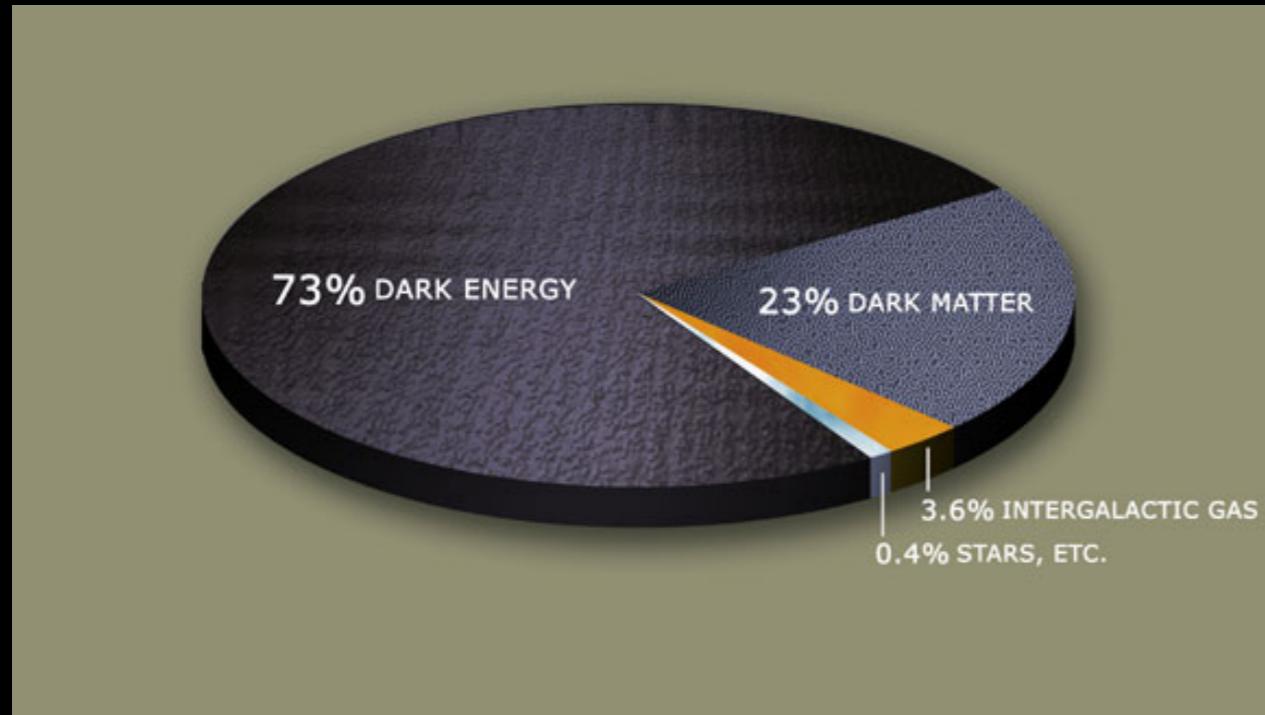
... Og vi vet heller ikke om kvarker er fundamentale !

$$10^{-18} = 0.0000000000000000001$$



... Og heller ikke Mørk Materie !

- Bare 4 % av det universet består av vet vi hva er !



- Kanskje det finnes nye partikler som er opphavet til mørk materie ?



Galaxy Cluster Abell 1689
Hubble Space Telescope • Advanced Camera for Surveys

NASA, N. Benitez (JHU), T. Broadhurst (The Hebrew University), H. Ford (JHU), M. Clampin (STScI), G. Hartig (STScI), G. Illingworth (UCO/Lick Observatory), the ACS Science Team and ESA
STScI-PRC03-01a

What is dark matter ?

Definition:

Dark Matter is matter which one can only observe through gravitational effects on visible matter

- does not interact with electromagnetic radiation

No doubt that dark matter exists - there is a multitude of direct observational evidence (since the 1930s):

- Galactic rotational curves
- Velocity dispersion of galaxies
- Galaxy clusters and gravitational lensing
- Cosmic microwave background
- Sky surveys and baryon acoustic oscillations



Fritz Zwicky



Vera Rubin

A couple of examples:

Observational evidence for dark matter



Picture of the galaxy cluster
ZwCl0024+1652, 5 billion
light years away, showing one
of the strongest evidence of
dark matter !



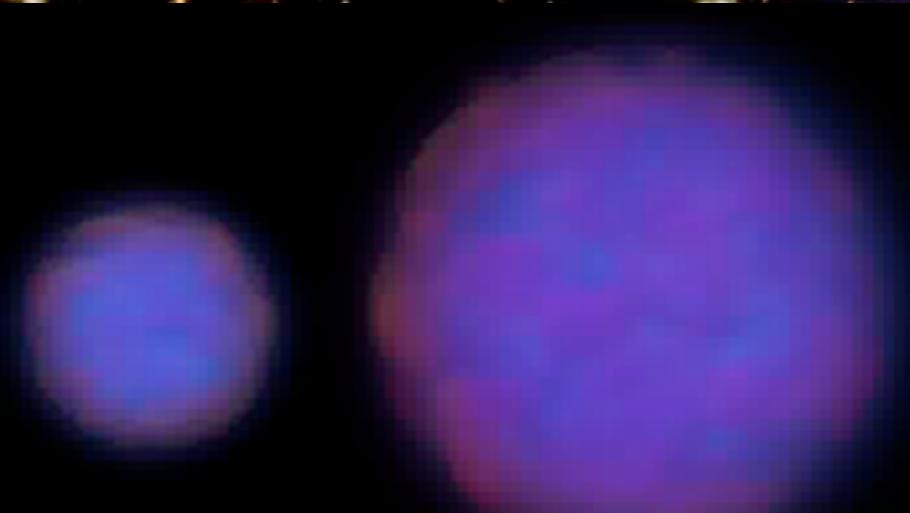
Pictures from the Hubble
telescope

Gravitational lensing
makes the galaxies appear
as disks

Observational evidence for dark matter



Hot gas (pink) detected in two galaxy clusters, one with a particular bullet shape. Other telescopes detected the bulk matter in the clusters which turns out to be dark matter (blue)

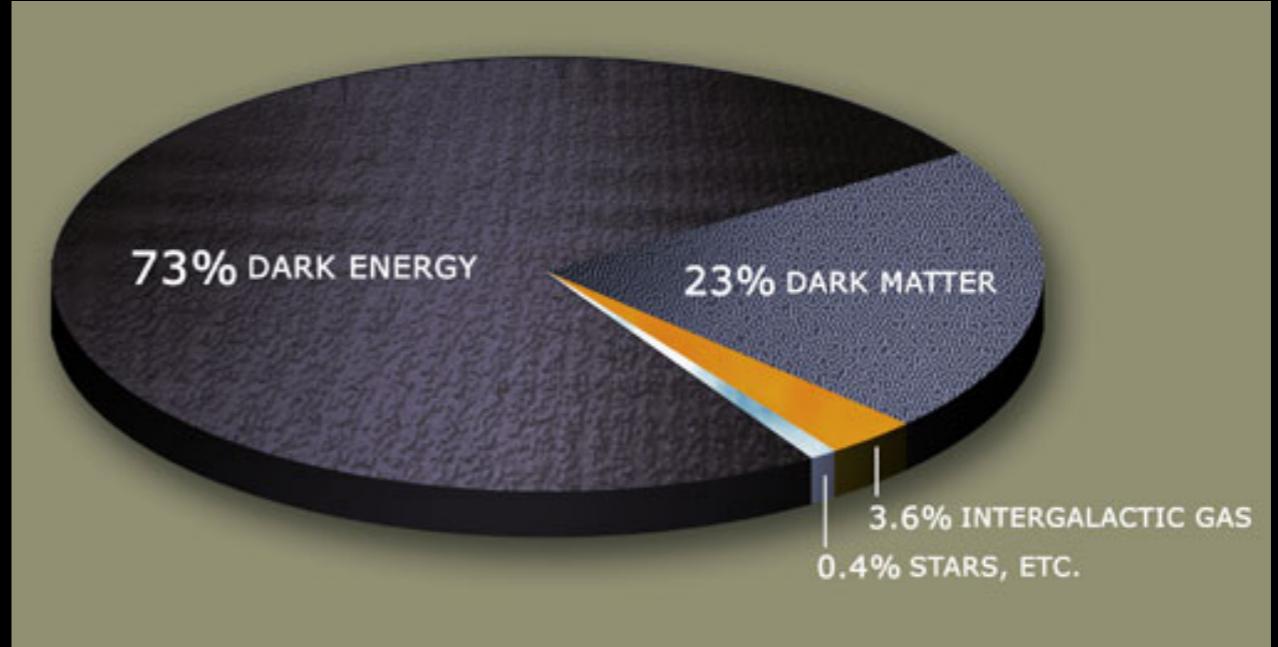


CREDIT: X-ray: NASA/CXC/CfA/M.Markevitch et al.; Optical: NASA/STScI; Magellan/U.Arizona/D.Clowe et al.; Lensing Map: NASA/STScI; ESO WFI; Magellan/U.Arizona/D.Clowe et al.

What is dark matter ?

These and similar studies of the universe show that the universe consists of :

- *4 % matter*
- *23 % dark matter*
- *73 % dark energy*



DARK MATTER IS PROPOSED TO EXPLAIN THE EXTRA MASS SEEN IN THE UNIVERSE ASSUMING THAT OUR ASSUMPTIONS ABOUT GRAVITY ARE CORRECT

WHAT DO WE KNOW ABOUT DARK MATTER ?

- It accounts for the additional gravitational effects observed in the universe
- It interacts only weakly with regular matter
- It also interact with other dark matter particles only through gravity

Three detection methods

THIS MAKES DARK MATTER VERY HARD TO DETECT AND STUDY !

-
-
-



- Indirect searches for the products of dark matter particle annihilation (e.g. Amanda, IceCube, GLAST-FERMI, EGRET)



- Direct searches for the atom recoil energy when a WIMP is passing (e.g. DAMA, CDMS, Xenon-10)



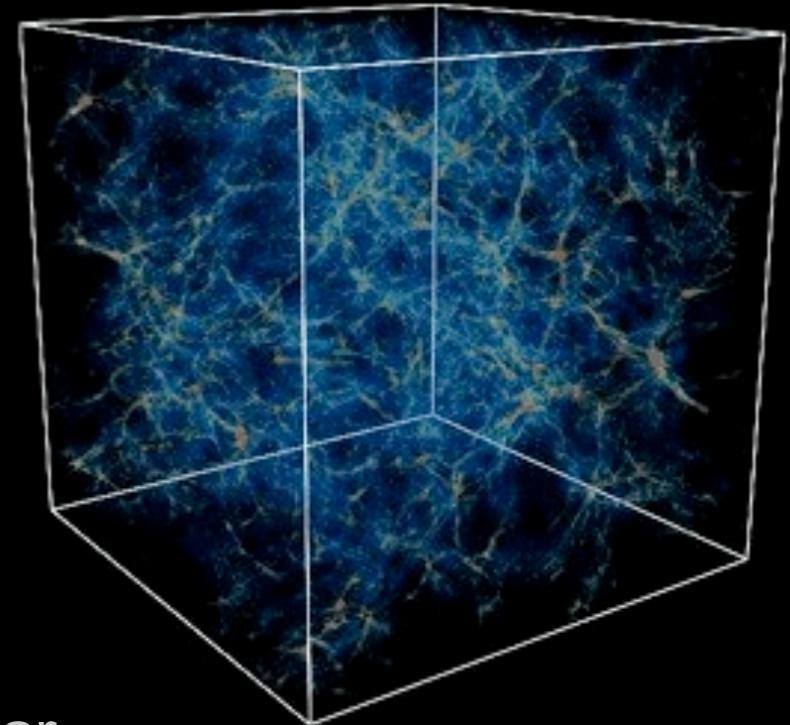
- Direct production made in high energy laboratories on earth (e.g. CERN, LHC, Tevatron)
Starting in 2009 LHC could be the first dark matter factory



What is dark matter ?

DIFFERENT THEORIES

- Multitude of models providing candidates as to what dark matter could be, from astrophysics, cosmology and particle physics
- **MACHOS** - *Massive, Compact Halo Objects*
 - Brown dwarfs, neutron stars or black holes, ...
→ not likely to account for more than a small fraction
- **WIMPS** - *Weakly Interactive Massive Particles*
 - **Known particle**
 - Neutrino → not likely to clump as DM do
 - **Unknown particle**
 - Supersymmetry → Massive neutralinos
(superpartner of neutral SM bosons)
 - Other exotic particles, e.g. gravitinos, axions, scalar dark matter

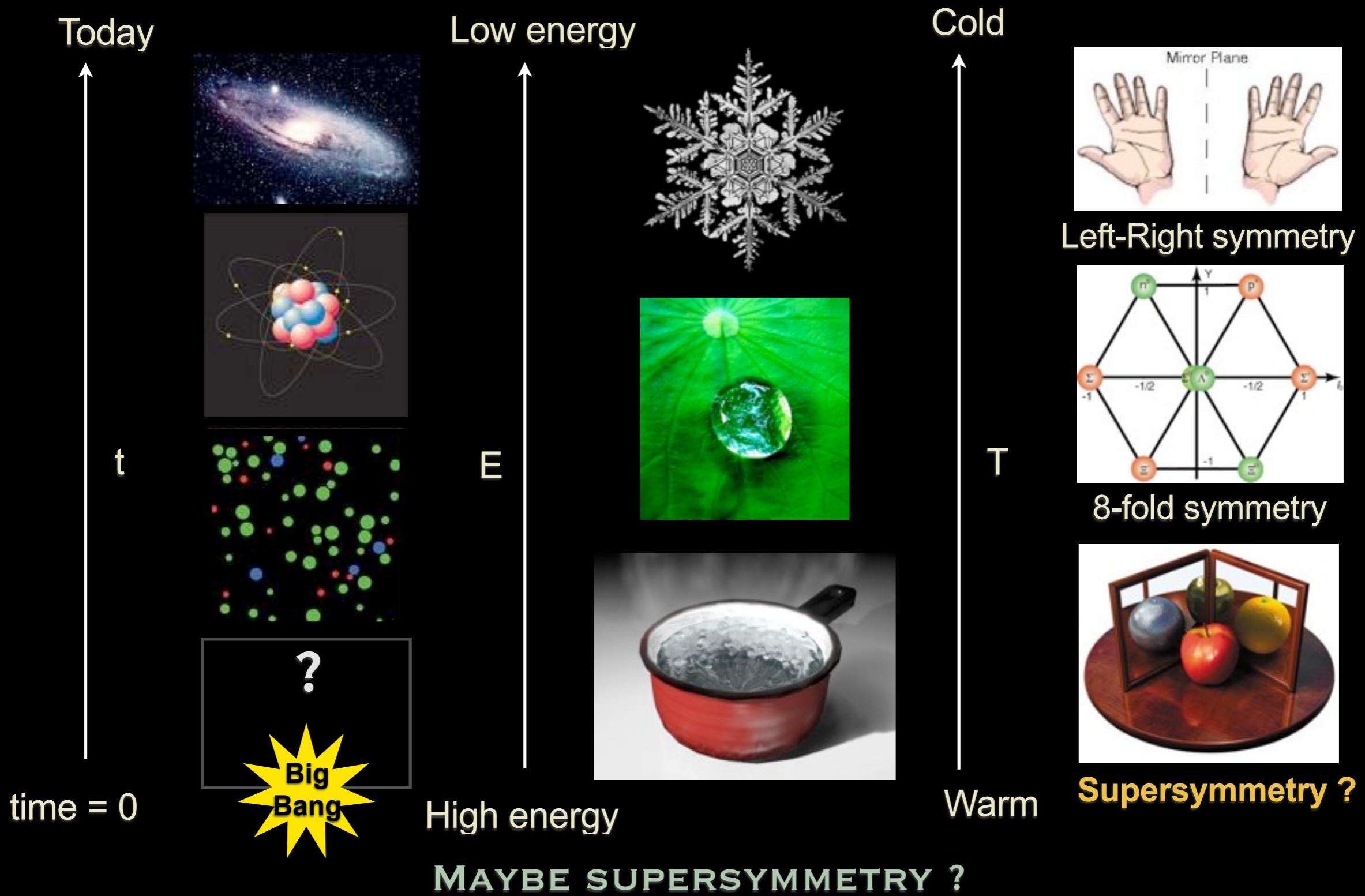


simulation represents the distribution of galaxy superclusters.

WE NEED TO LOOK FOR NEW PHYSICS BEYOND THE STANDARD MODEL !

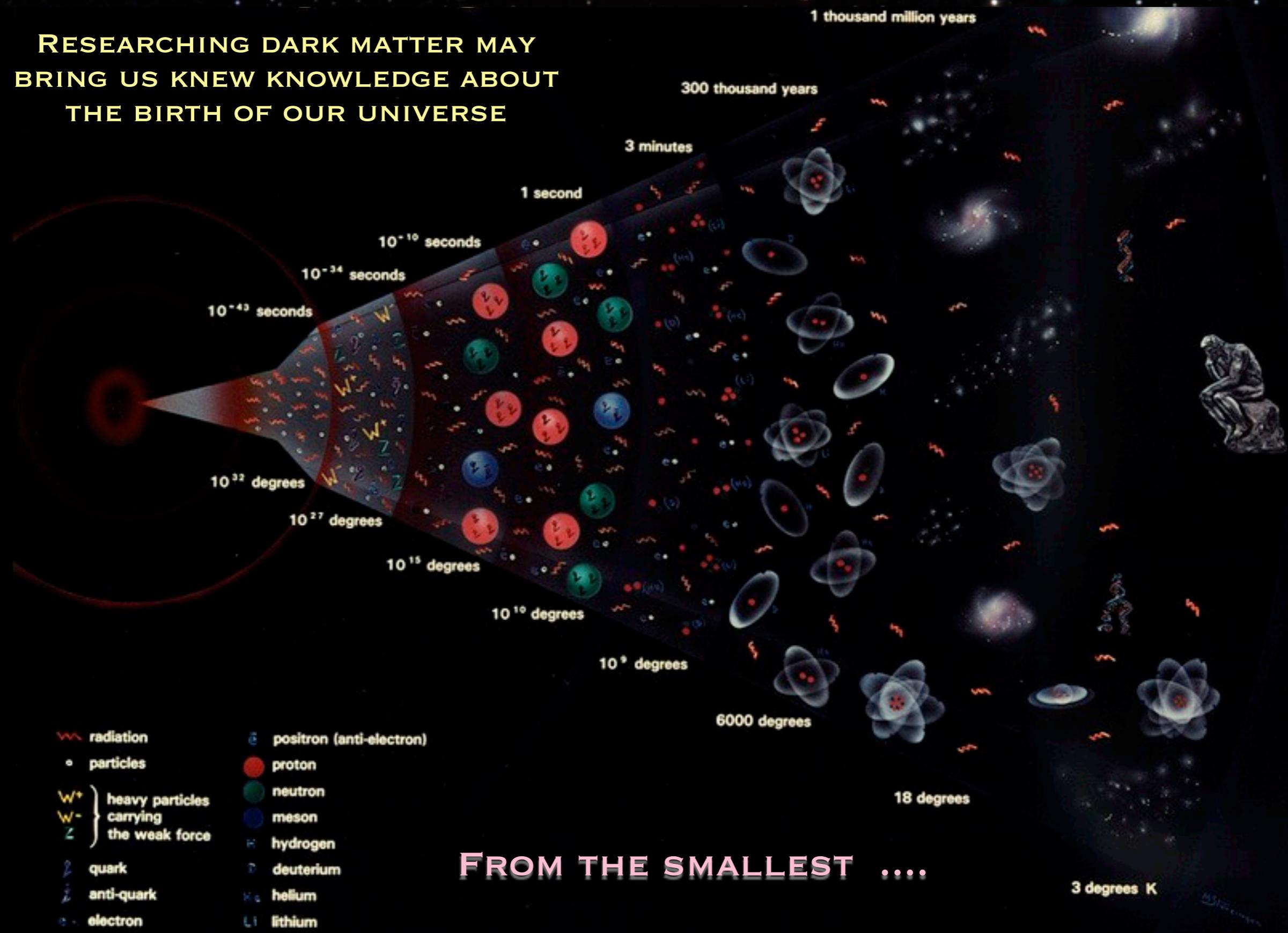
Possible new physics

**NEW STATES AND NEW SYMMETRIES COULD HAVE EXISTED JUST
AFTER THE BIG BANG**



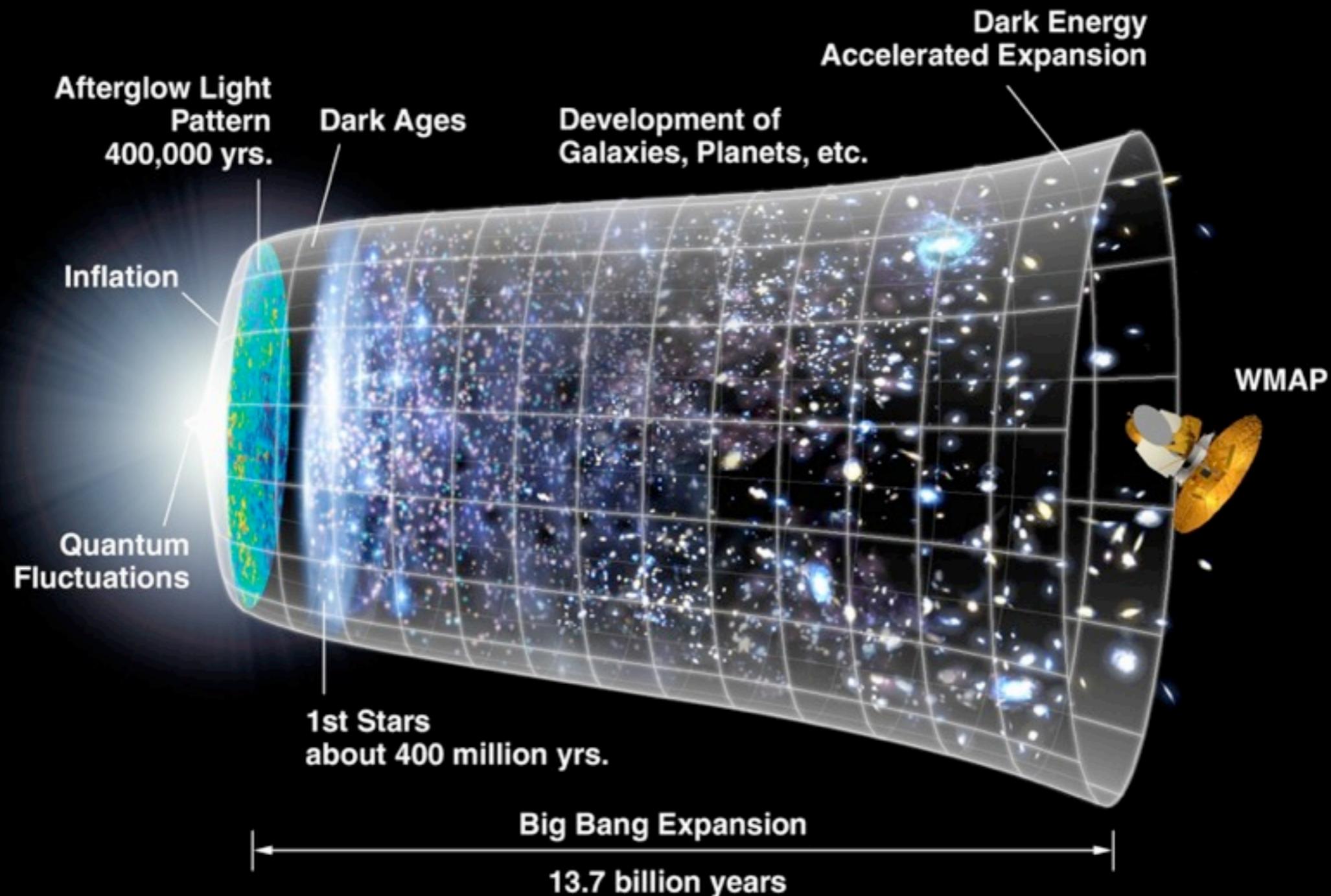
Particle physicists view of the Big Bang

RESEARCHING DARK MATTER MAY
BRING US NEW KNOWLEDGE ABOUT
THE BIRTH OF OUR UNIVERSE



Astrophysicist view of the Big Bang

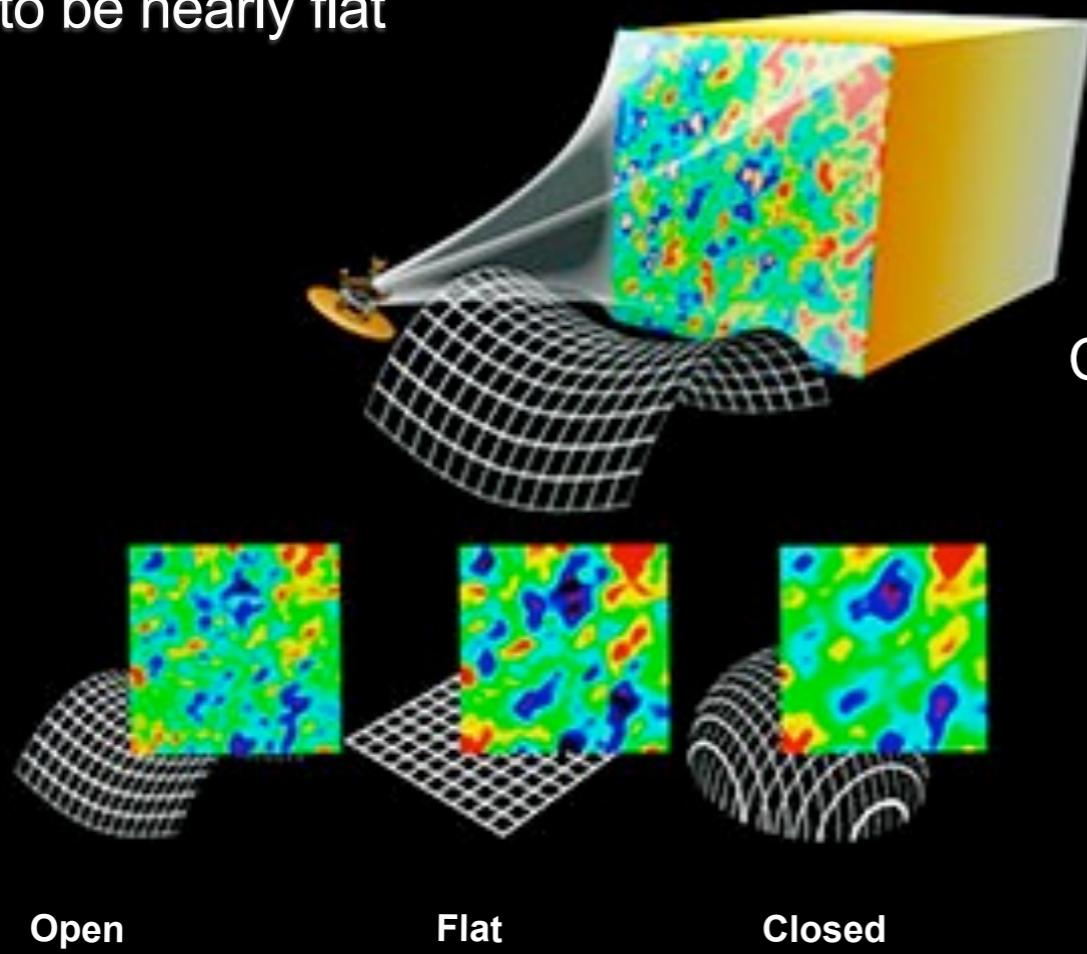
.... TO THE BIGGEST !



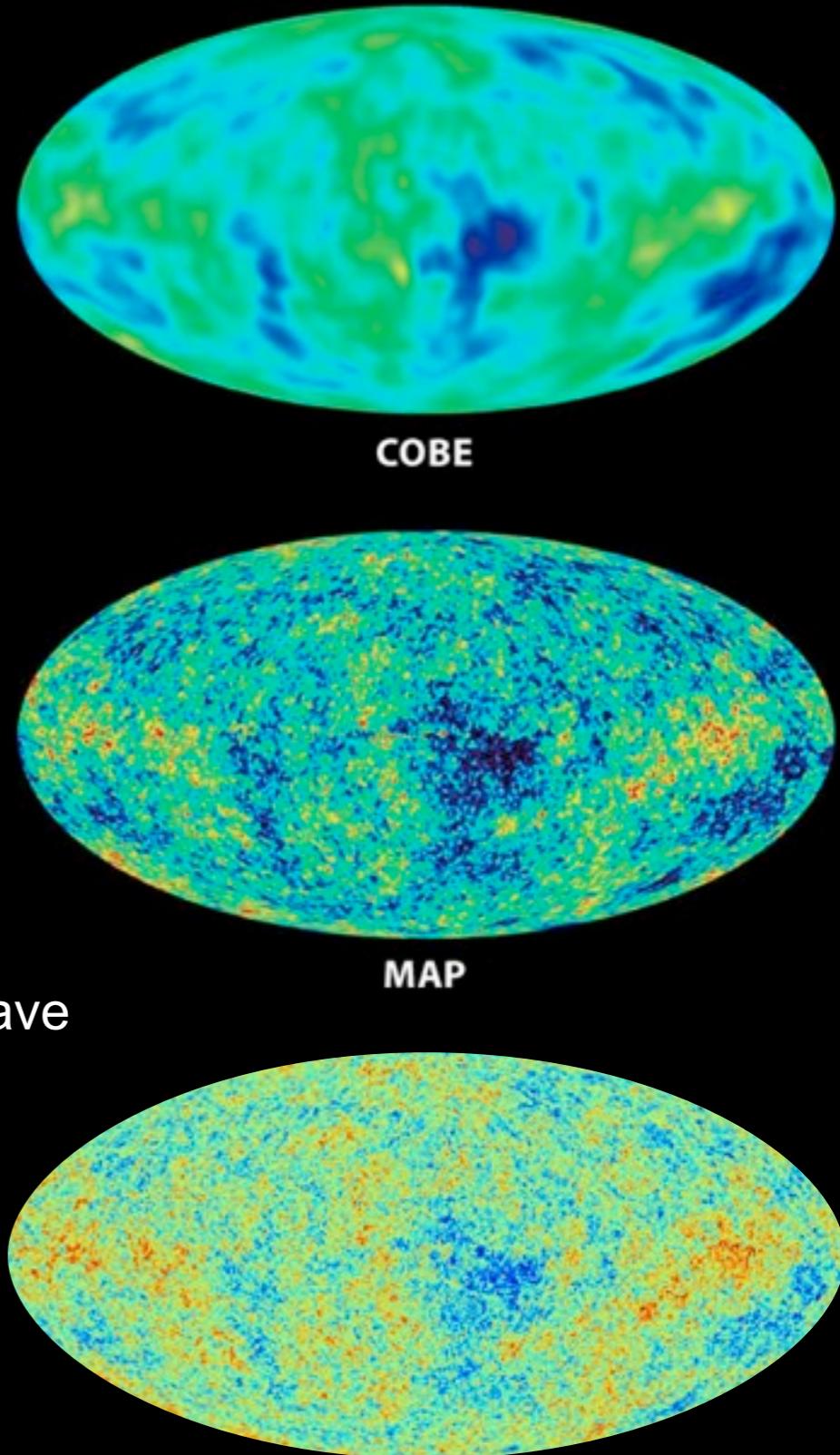
DARK MATTER PLAYS A CENTRAL ROLE IN THE MODELING OF STRUCTURE FORMATION AND THE EVOLUTION OF GALAXIES

New information of the geometry of the universe

- The amount of dark matter and dark energy in the universe is crucial to determine the geometry of space
 - Open : density less than critical density
 - Flat : density equal to critical density
 - Closed: density more than the critical density
- Gives information on the evolution of the universe (eternal expansion, in equilibrium, or stop and collapse)
- The spacial geometry have been measured by WMAP to be nearly flat



Cosmic microwave background

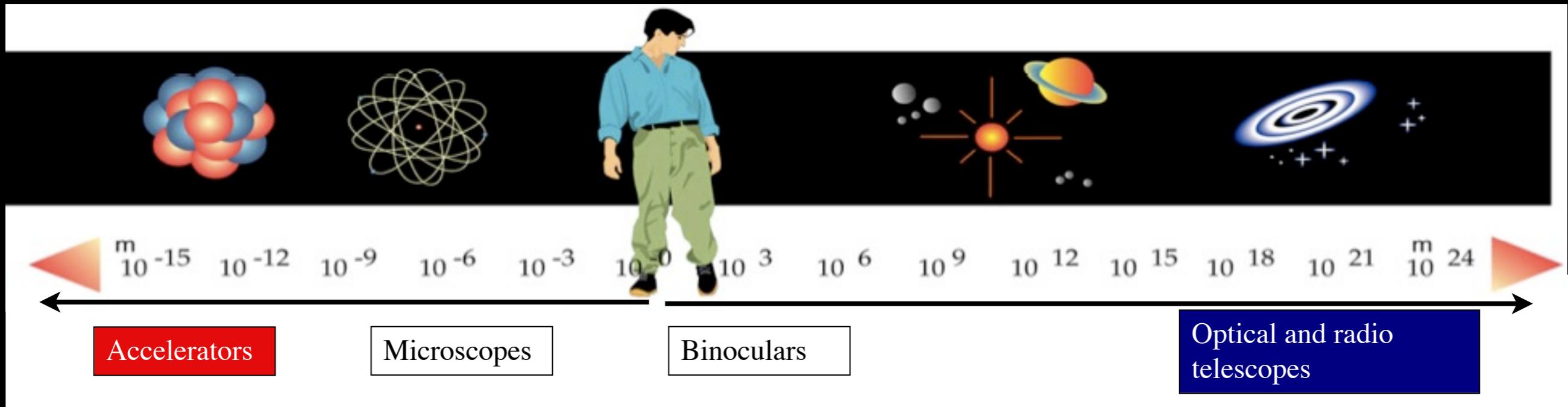


PLANCK (simulated)

ESA/NASA

Hva trenger vi for å undersøke dette nærmere ?

Eksempel: LHC



Et sterkt mikroskop - Akselerator !

Som også er en tidsmaskin

- Vi kan studere tilstander like etter big bang



Hva trenger vi for å undersøke dette nærmere ?



Et veldig god detektor ! - ATLAS -

Som fungerer som kamera som tar bilder av kollisjonene

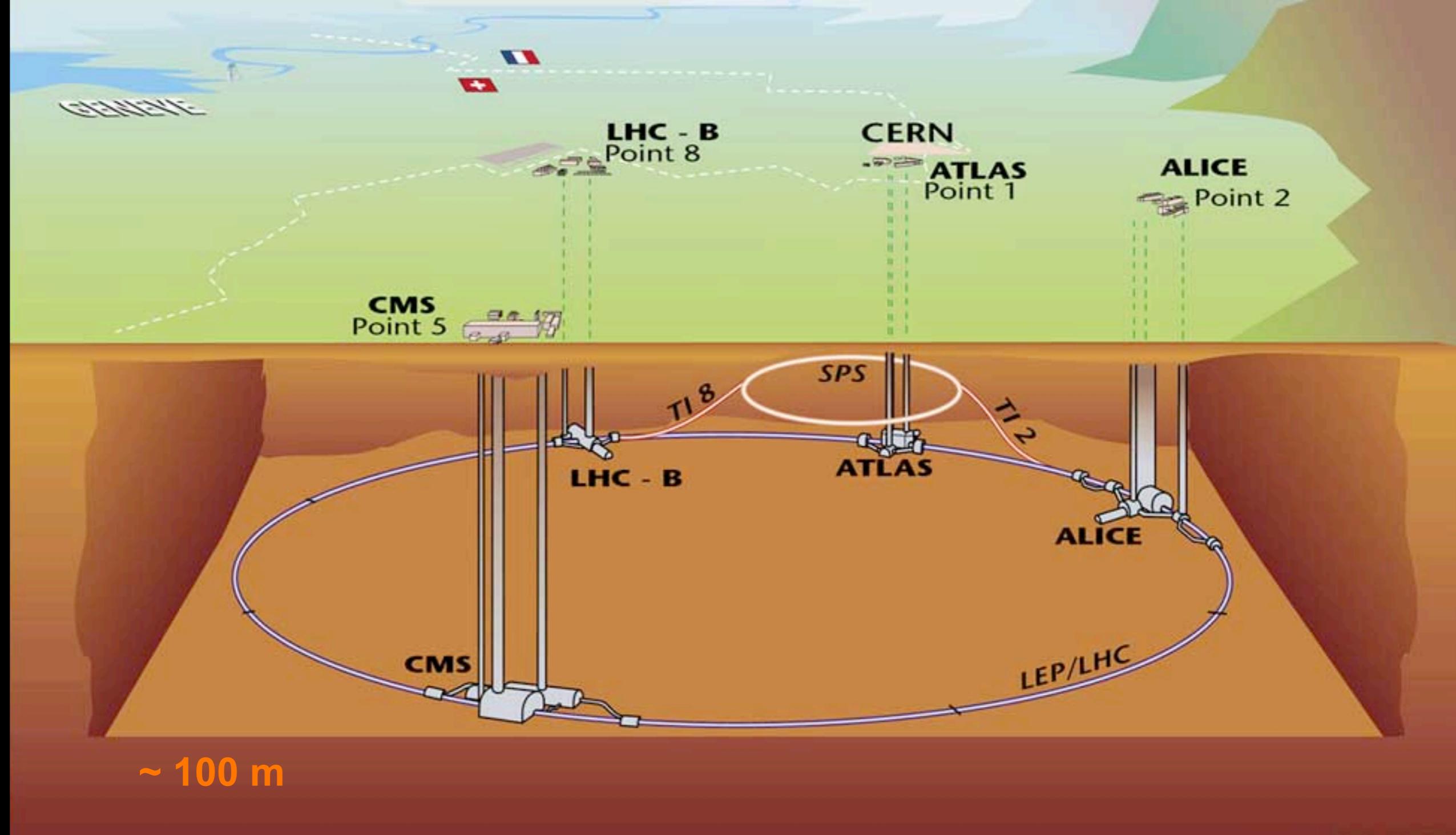
1. The challenge of LHC

Four gigantic underground caverns to host the huge detectors

The highest energy of any accelerator in the world

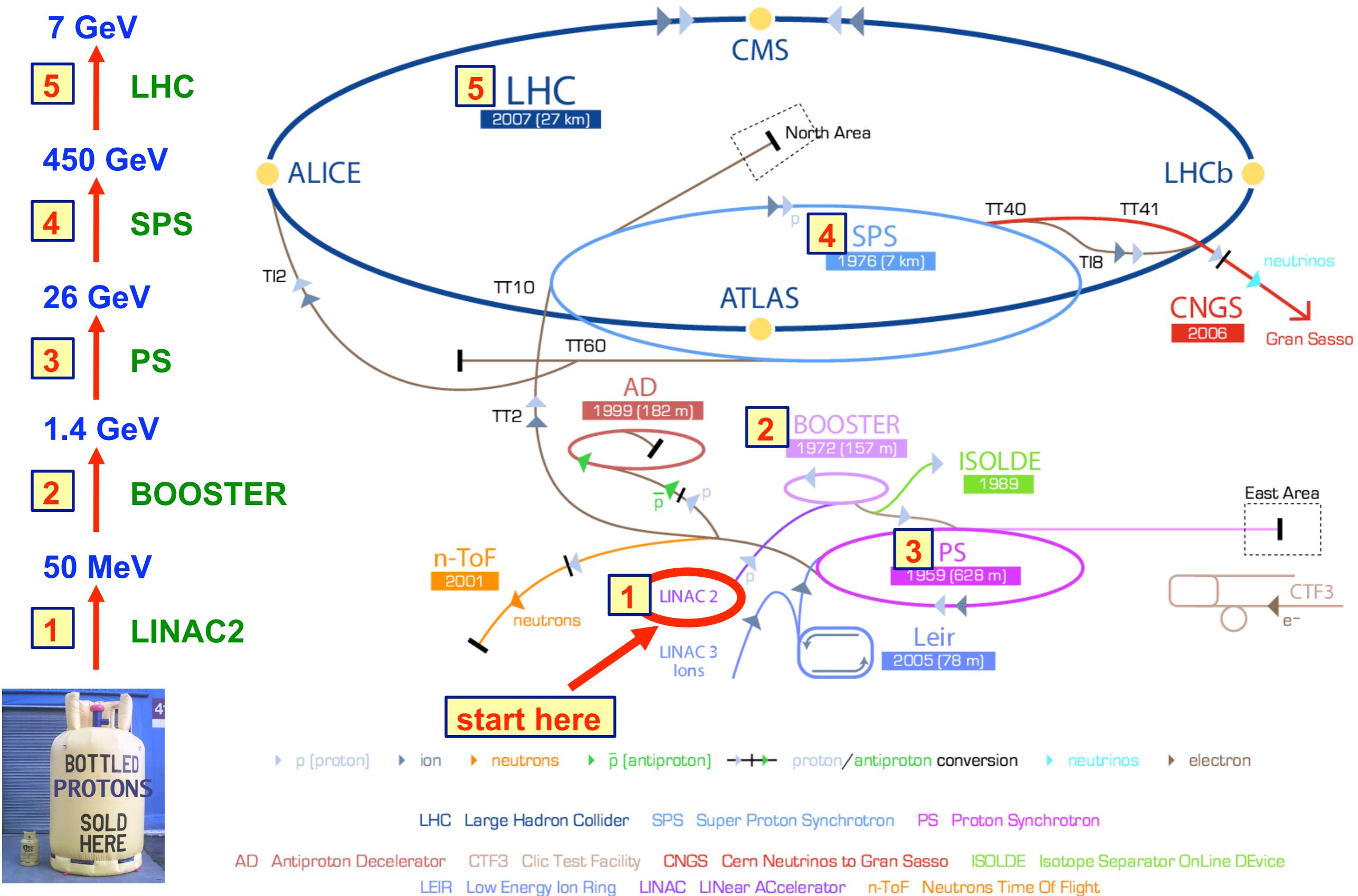
The most intense beams of colliding particles

It will operate at a temperature (1.9 K) colder than outer space



~ 100 m

1. CERN accelerator complex



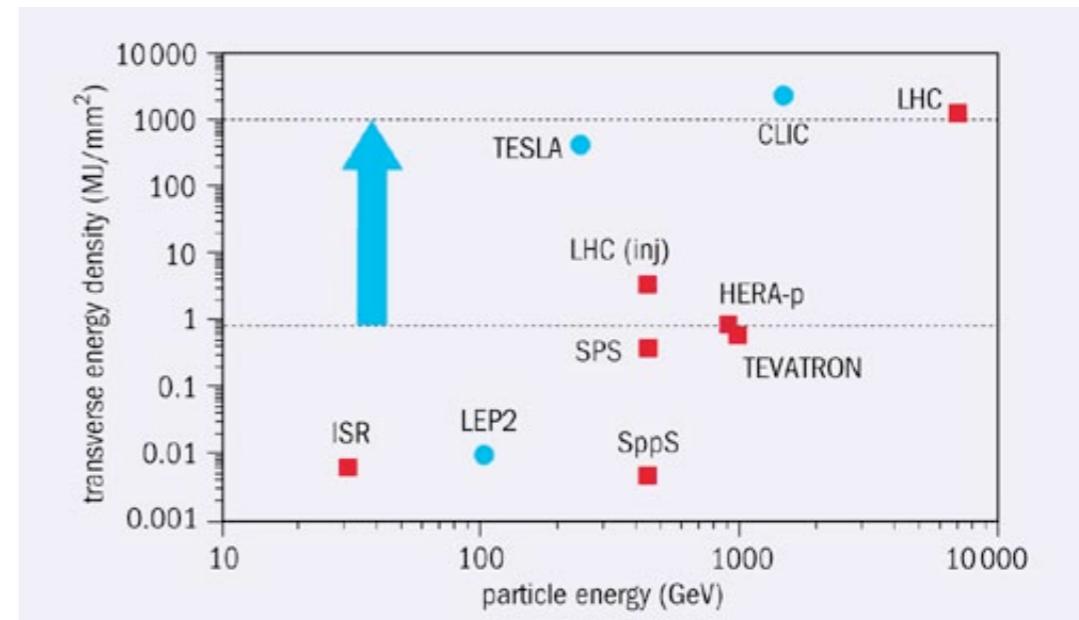
1. LHC parameters

	LHC (2009)	Tevatron (1987)	SppS (1981)
max. Energy (TeV)	7	1	0.450
circumference (km)	26.7	6.3	6.9
luminosity ($10^{30} \text{cm}^{-2}\text{s}^{-2}$)	10000	210	6
time between collisions (μs)	0.0025	0.396	3.8
crossing angle (μrad)	300	0	0
p/bunch (10^{10})	11	27/7.5	15/8
number of bunches	2808	36	6
beam size (μm)	16	34/29	36/27
filling time (min)	7.5	30	0.5
acceleration (s)	1200	86	10

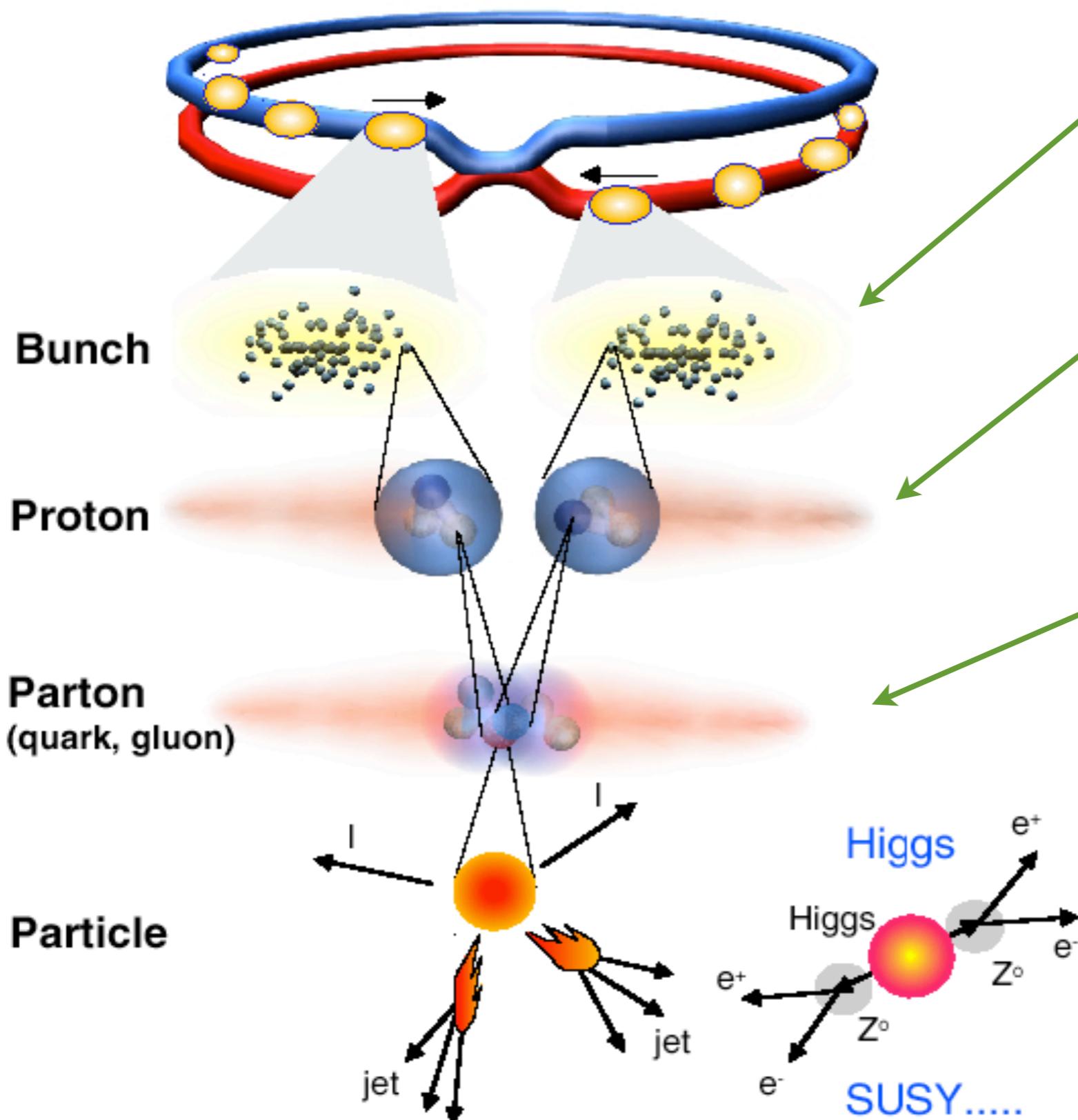
Compared with Tevatron:

- 7 times more energy
- ~ 50-100 times more luminosity

We need to build detectors for energies and luminosity not yet explored !



1. LHC Collisions



Proton/proton collisions

- 7 TeV beam energy
- 10^{11} protons per bunch

Event rate in detectors

- $N = L \times \sigma (pp)$
- $N \sim 1\ 000\ 000\ 000$ interactions/s

Event Type

- Mostly soft (low p_T) events
- Interesting hard (high p_T) events are rare

Interesting events are very rare !

- 1 in $10\ 000\ 000\ 000\ 000$ collisions !

We need detectors that can cope with high statistics !

1. LHC Beam Energy and Density

- 2808 bunches in the ring
- 1.1×10^{11} protons per bunch
- 7 TeV

= 350 MJ stored energy per proton beam
Same as colliding 2 x 120 elephants !

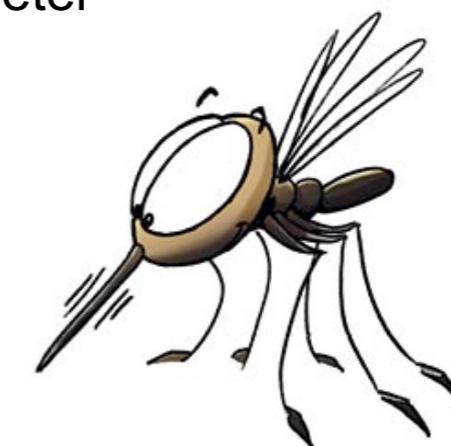


120 elephants running at 40 km / h



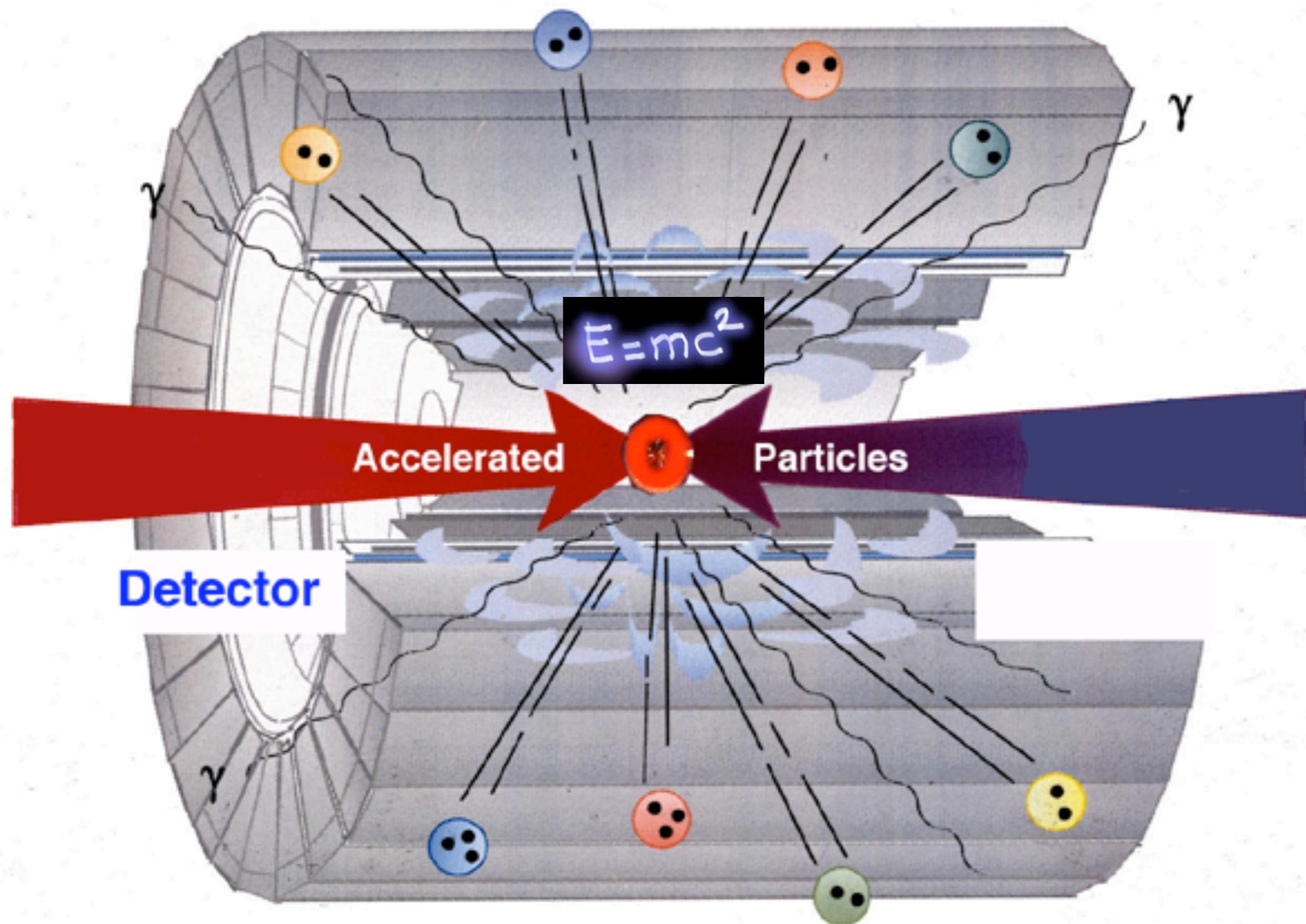
120 elephants running at 40 km / h

- The eye of a needle is 0.3 mm in diameter, the proton beams at the interaction point are 10 times smaller = 0.03 mm in diameter
- The energy of a single 7 TeV proton is equivalent to a flying mosquito 1 μJ

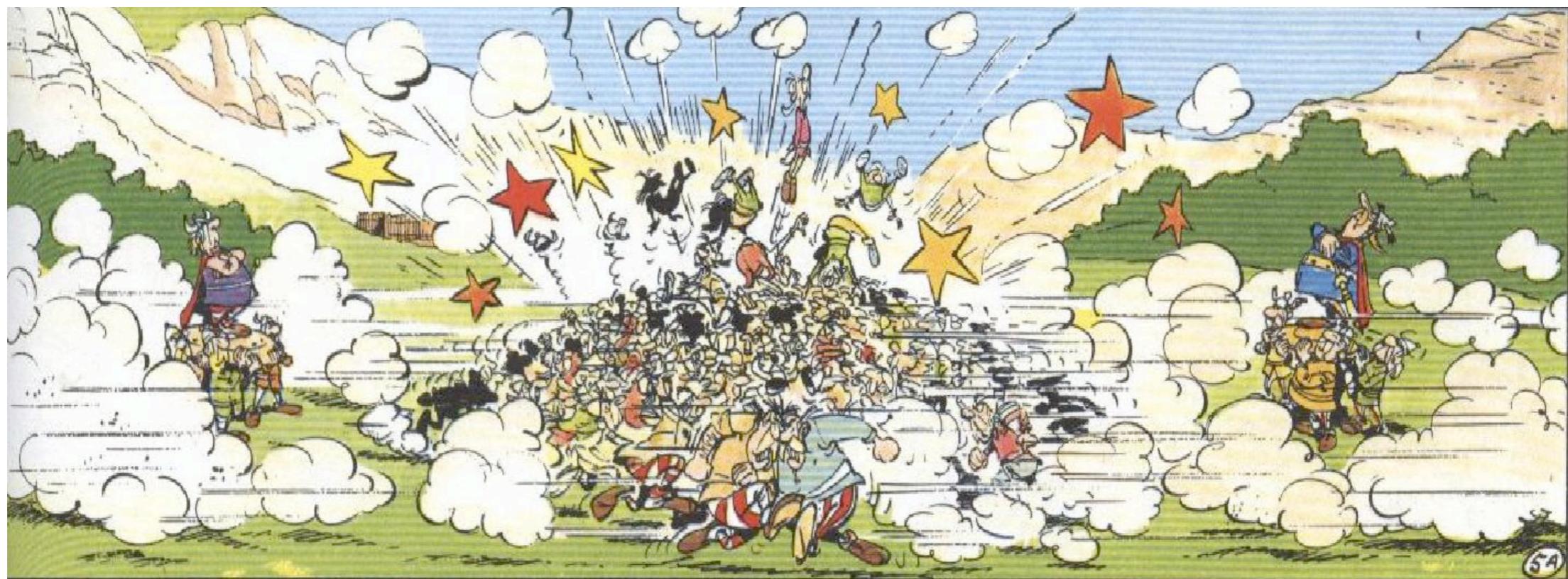


H. Pernegger

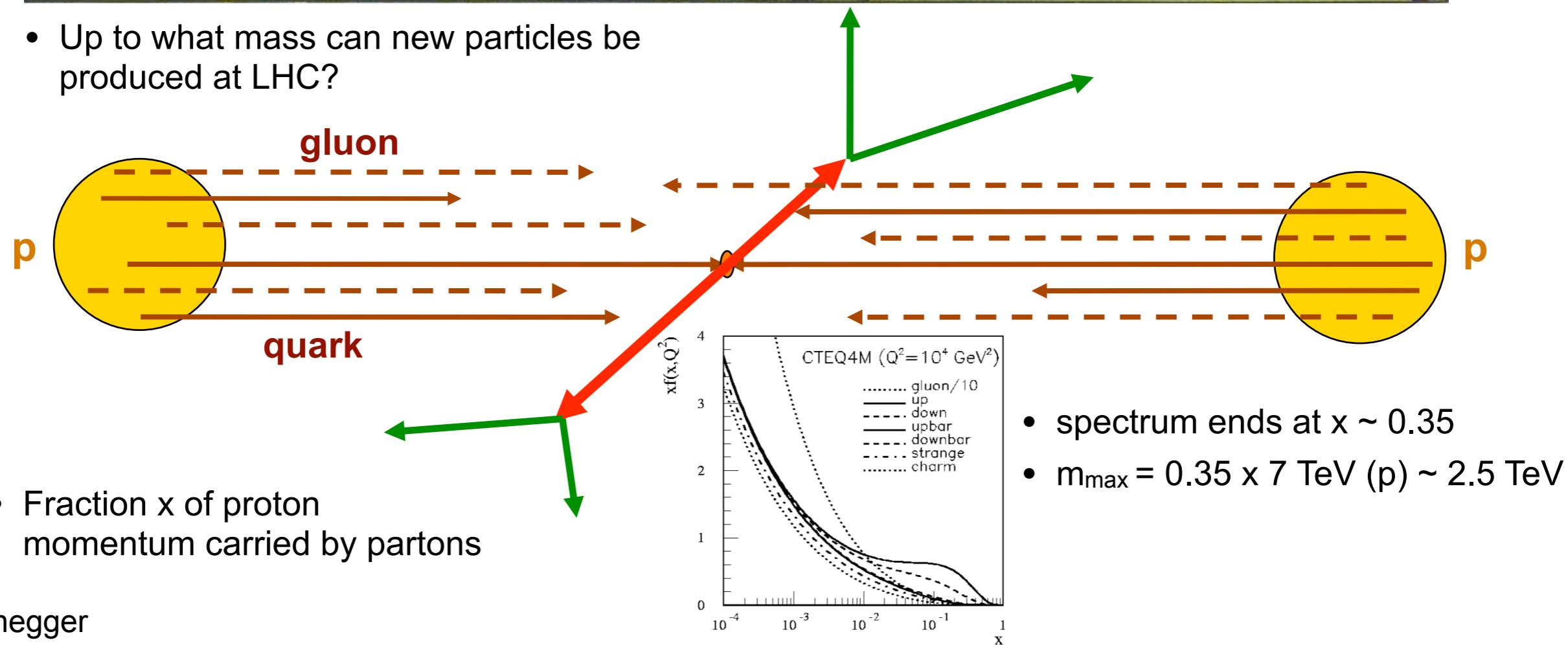
1. Particle Detection



1. Particle Production



- Up to what mass can new particles be produced at LHC?



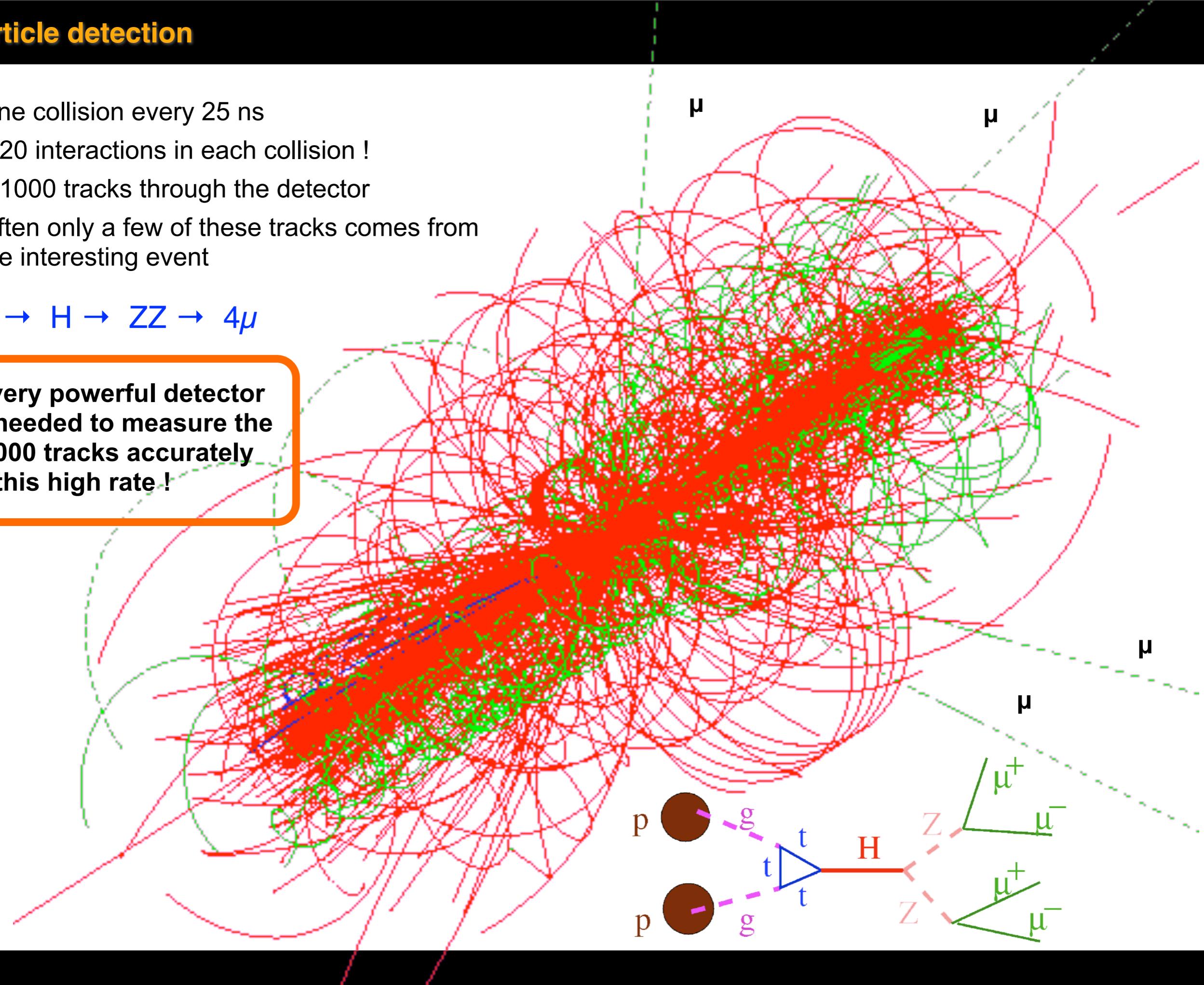
H. Pernegger

1. Particle detection

- One collision every 25 ns
- < 20 interactions in each collision !
- ~ 1000 tracks through the detector
- Often only a few of these tracks comes from the interesting event

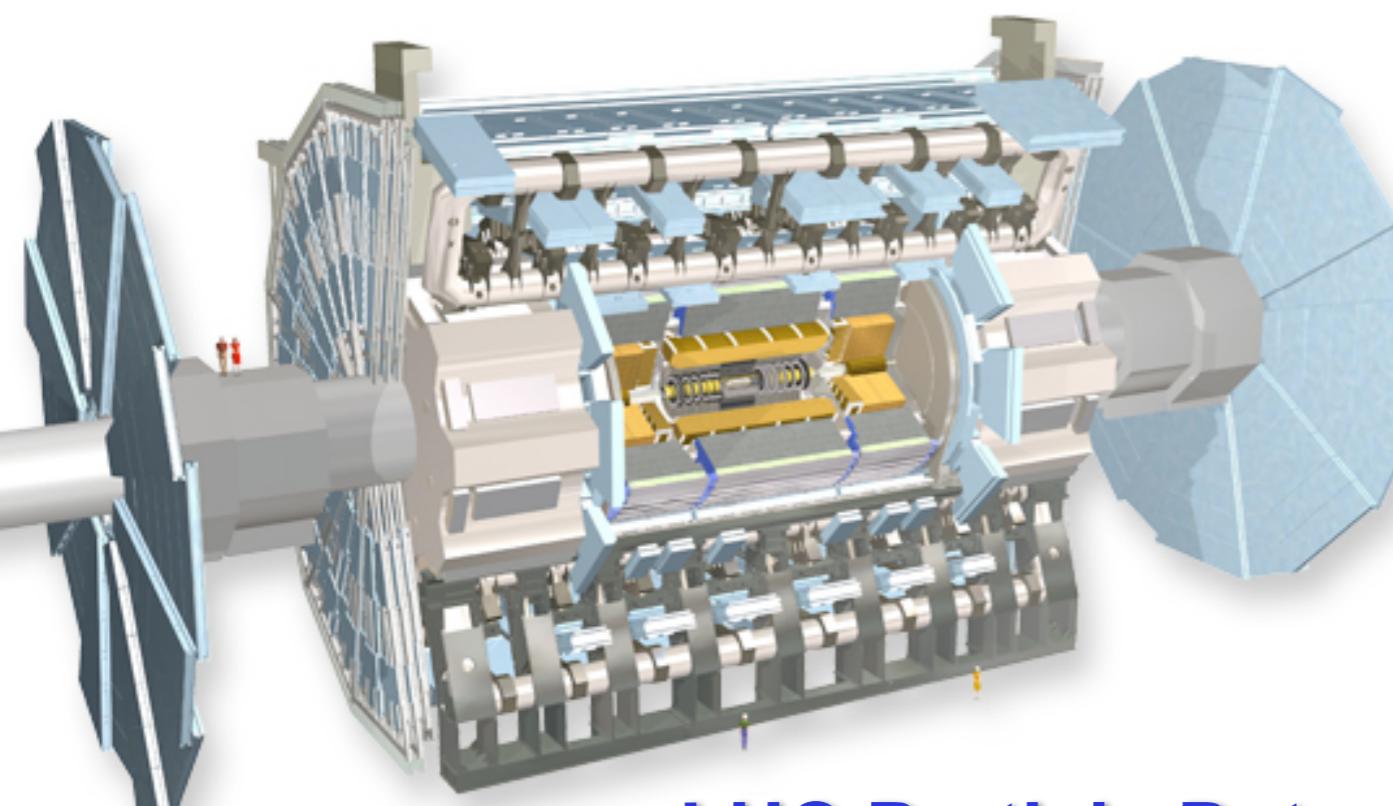
$pp \rightarrow H \rightarrow ZZ \rightarrow 4\mu$

A very powerful detector
is needed to measure the
~1000 tracks accurately
at this high rate !

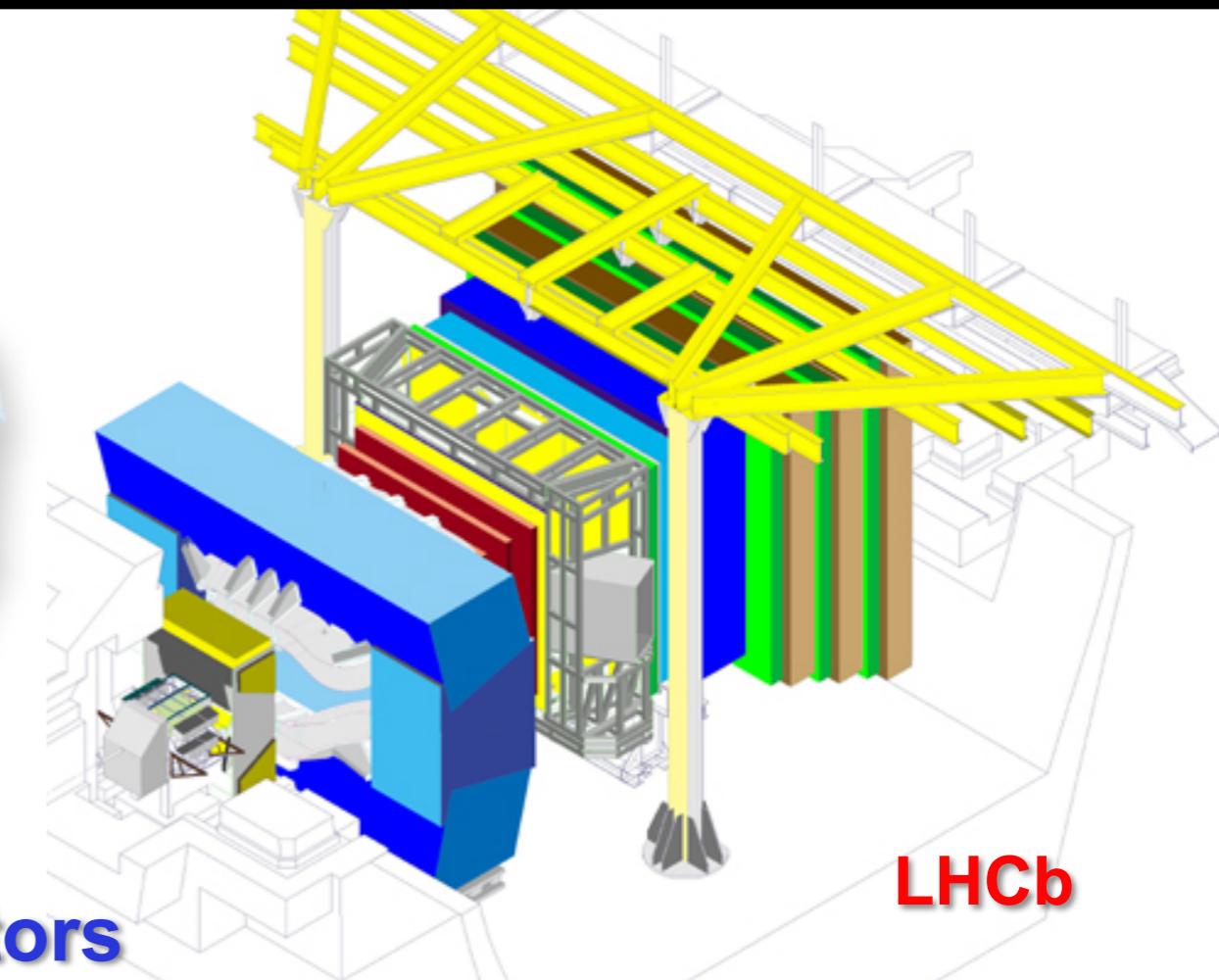


1. LHC detectors

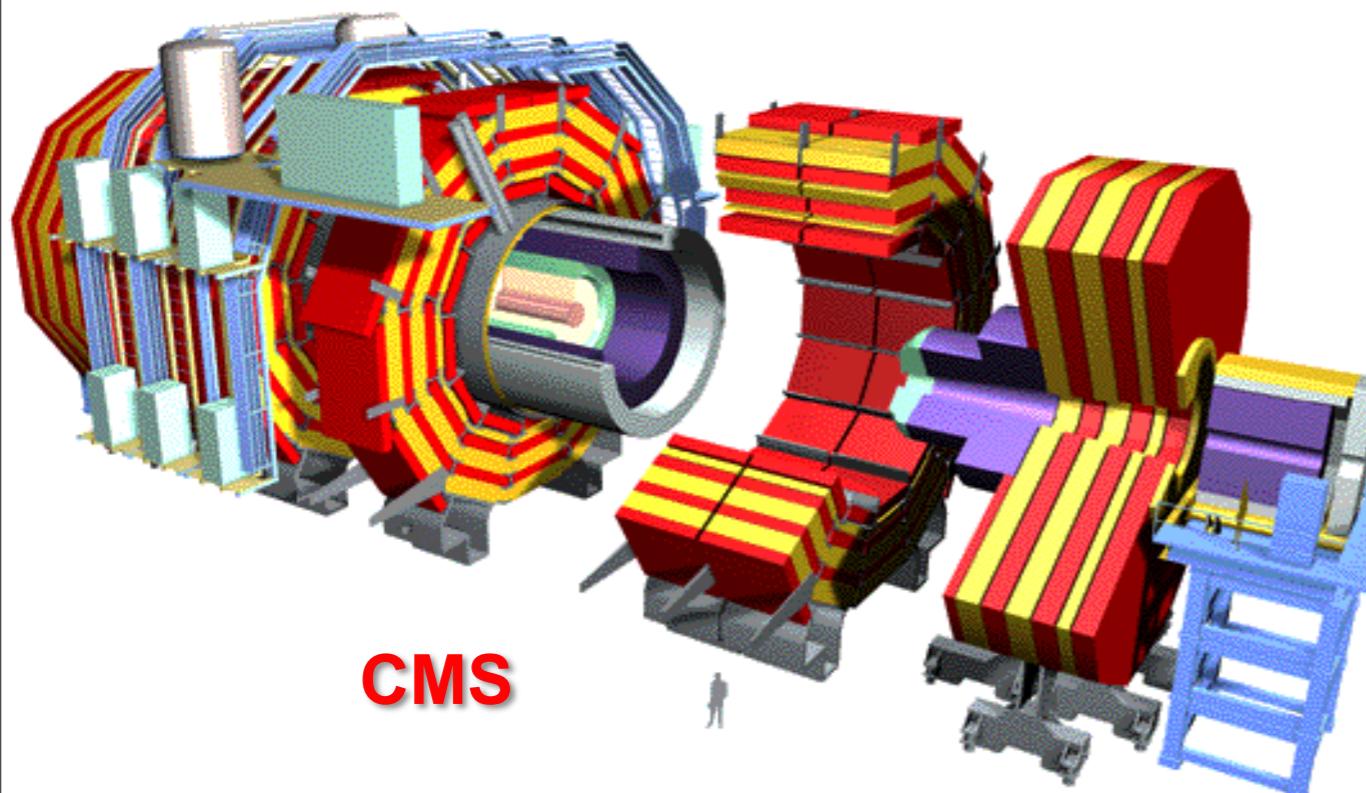
ATLAS



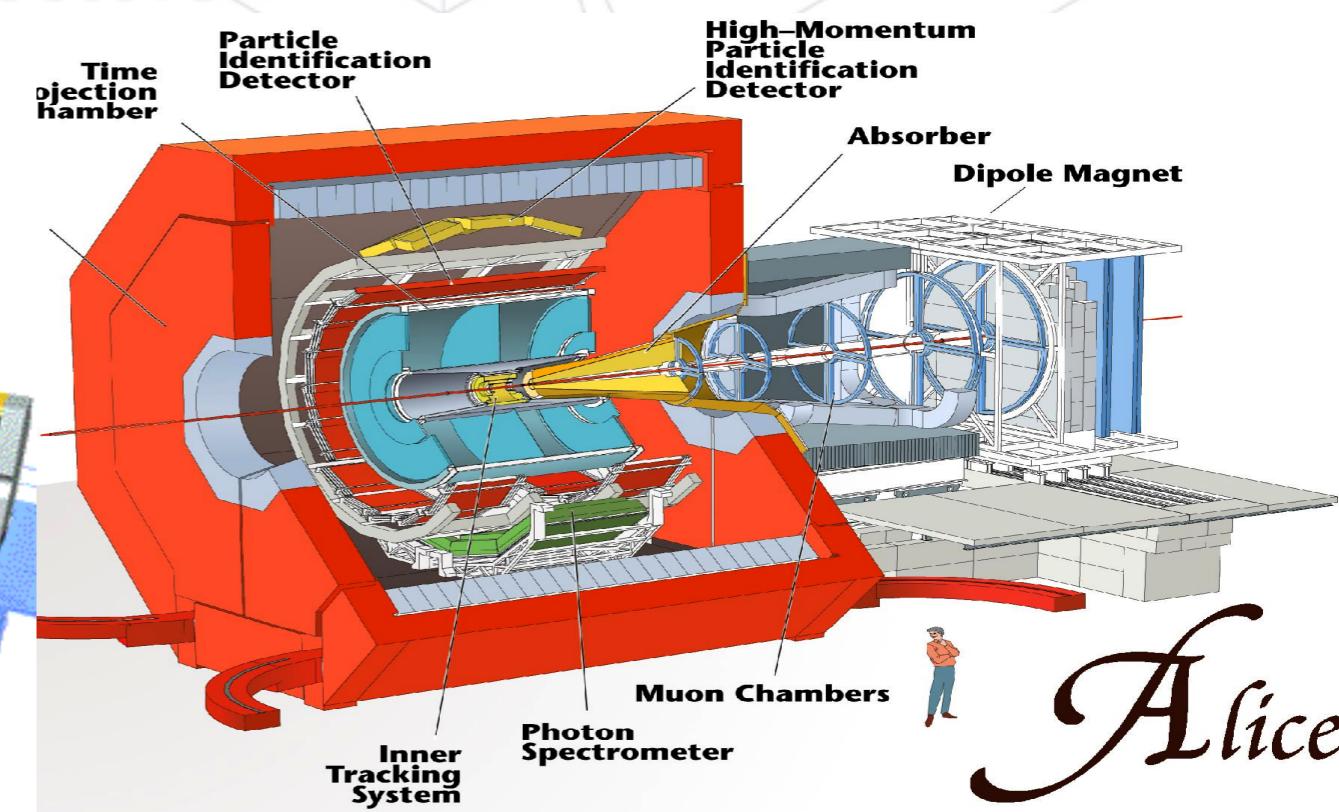
LHCb



LHC Particle Detectors



CMS



Alice

4. Constructing ATLAS



Argentina	Netherlands
Armenia	Norway
Australia	Poland
Austria	Portugal
Azerbaijan	Romania
Belarus	Russia
Brazil	Serbia
Canada	Slovakia
China	Slovenia
Czech Republic	Spain
Denmark	Sweden
France	Switzerland
Georgia	Taiwan

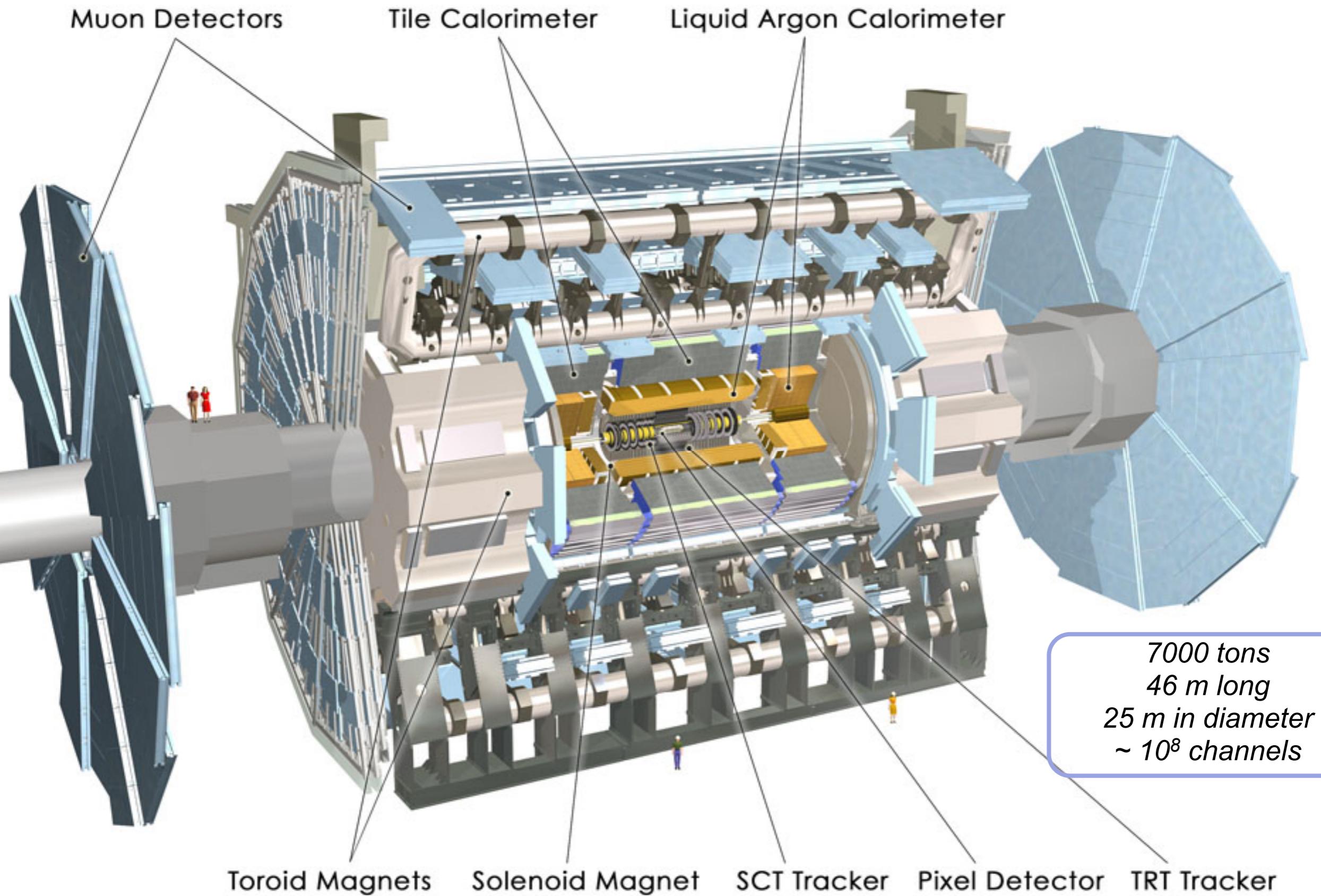
35 countries
158 institutions
1770 scientific authors
100s of people at CERN for ~ 3 years

Morocco

ATLAS
Collaboration



3. Example: The ATLAS detector



2. Detecting Particles

<http://pdg.lbl.gov>

~ 180 Selected Particles

$\pi, W^\pm, Z^0, g, e, \mu, \gamma, \nu_e, \nu_\mu, \nu_\tau, \pi^\pm, \pi^0, \eta, f_0(600), g(800),$
 $w(782), \eta'(858), f_0(980), a_0(980), \phi(1020), h_+(1170), b_+(1235),$
 $a_+(1260), f_+(1270), f_+(1285), \eta(1295), \pi(1300), a_2(1320),$
 $f_0(1370), f_+(1420), w(1420), \eta(1440), a_0(1450), g(1450),$
 $f_0(1500), f_+(1520), w(1650), w_3(1670), \pi_2(1670), \phi(1680),$
 $g_3(1690), g(1700), f_0(1710), \pi(1800), \phi_3(1850), f_+(2010),$
 $a_+(2040), f_+(2050), f_+(2300), f_+(2340), K^\pm, K^0, K_S^0, K_L^0, K^*(892),$
 $K_+(9270), K_+(1400), K^*(1410), K_0^*(1430), K_2^*(1430), K^*(1680),$
 $K_2(1770), K_3^*(1780), K_2(1820), K_4^*(2045), D^\pm, D^0, D^*(2007)^0,$
 $D^*(2010)^\pm, D_0(2420)^\pm, D_s^*(2460)^\pm, D_s^*(2460)^\pm, D_s^\pm, D_s^{*\pm},$
 $D_{s1}(2536)^\pm, D_{s1}(2573)^\pm, B^\pm, B^0, B^*, B_S^0, B_c^\pm, \eta_c(1S), J/\psi(1S),$
 $\chi_{c0}(1P), \chi_{c1}(1P), \chi_{c2}(1P), \psi(2S), \psi(3770), \psi(4040), \psi(4160),$
 $\psi(4415), \Upsilon(1S), \chi_{b0}(1P), \chi_{b1}(1P), \chi_{b2}(1P), \Upsilon(2S), \chi_{b3}(2P),$
 $\chi_{b3}(2P), \Upsilon(3S), \Upsilon(4S), \Upsilon(10860), \Upsilon(11020), p, n, N(1440),$
 $N(1520), N(1535), N(1650), N(1675), N(1680), N(1700), N(1710),$
 $N(1720), N(2190), N(2220), N(2250), N(2600), \Delta(1232), \Delta(1600),$
 $\Delta(1620), \Delta(1700), \Delta(1905), \Delta(1910), \Delta(1920), \Delta(1930), \Delta(1950),$
 $\Delta(2420), \Lambda, \Lambda(1405), \Lambda(1520), \Lambda(1600), \Lambda(1670), \Lambda(1690),$
 $\Lambda(1800), \Lambda(1810), \Lambda(1820), \Lambda(1830), \Lambda(1890), \Lambda(2100),$
 $\Lambda(2110), \Lambda(2350), \Sigma^+, \Sigma^0, \Sigma^-, \Sigma(1385), \Sigma(1660), \Sigma(1670),$
 $\Sigma(1750), \Sigma(1775), \Sigma(1915), \Sigma(1940), \Sigma(2030), \Sigma(2250), \Xi^0, \Xi^-,$
 $\Xi^-(1530), \Xi^-(1690), \Xi^-(1820), \Xi^-(1950), \Xi^-(2030), \Omega^-, \Omega(2250)^-,$
 $\Lambda_c^+, \Lambda_c^0, \Sigma_c(2455), \Sigma_c(2520), \Xi_c^+, \Xi_c^0, \Xi_c^+, \Xi_c^0, \Xi(2645)$
 $\Xi_c^-(2780), \Xi_c^-(2815), \Omega_c^0, \Lambda_b^0, \Xi_b^0, \Xi_b^-, t\bar{t}$

There are Many more

W. Riegler / CERN

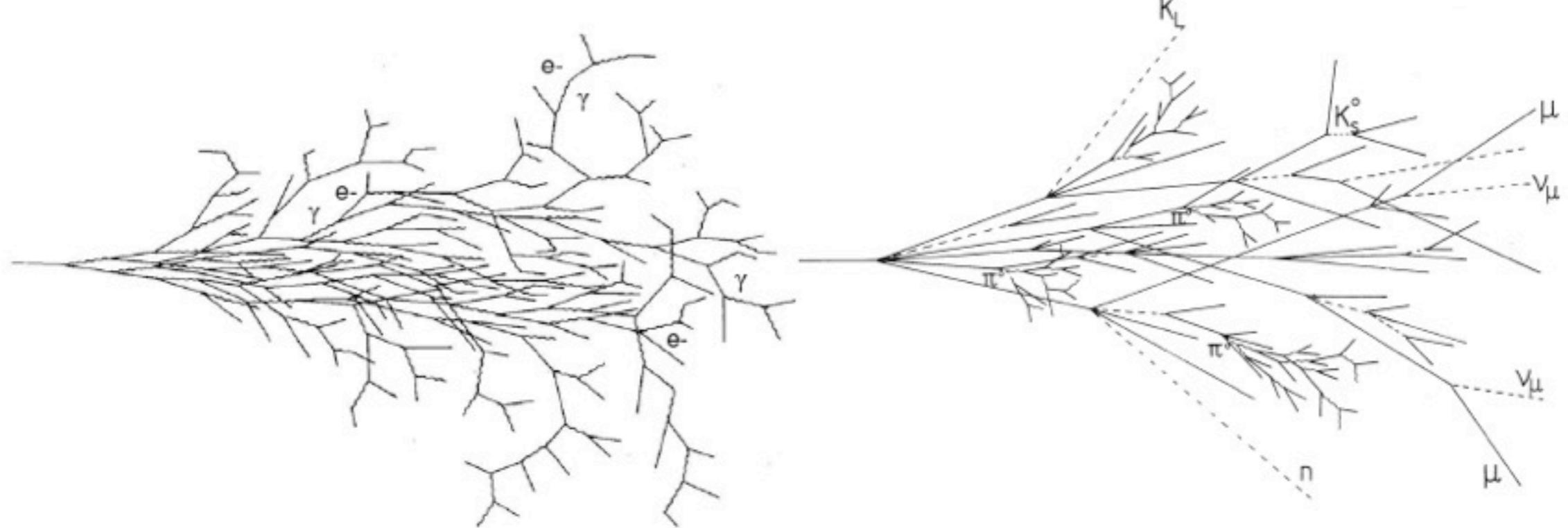
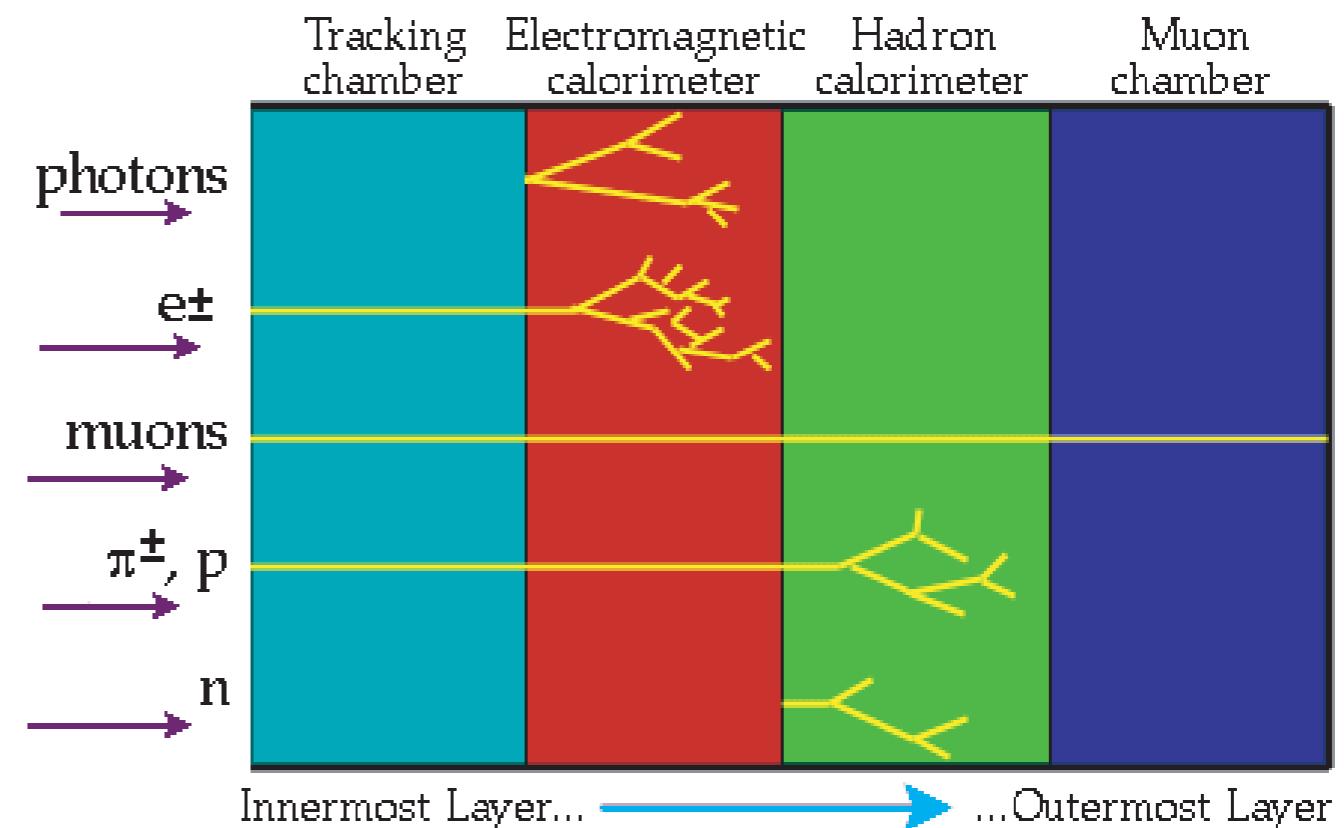
Particle lifetimes:

- Large number of particles !
- Only ~ 27 with a lifetime $c\tau > 1 \mu\text{m}$, which we can see as tracks in the detector
- ~ 13 of these have $c\tau < 500 \mu\text{m}$, very short tracks
- From the ~14 remaining particles only 8 are very frequent:
 - $e^{+/-}, \mu^{+/-}, \gamma, \pi^{+/-}, K^{+/-}, K^0, p^{+/-}, n$
 - A particle detector must be able to identify and measure energy and momentum of these 8 particles !

Although many particles are produced in the collisions only a few live long enough to move through the detector !

2. Particle Interactions

- Electrons ionise and show **bremsstrahlung** due to the small mass
- Photons don't ionise but show **pair production** in high Z material
- Charged hadrons ionise and produce a shower in dense materials
- Neutral hadrons don't ionise and produce a **shower** in dense material
- Muons ionise and **don't** produce a shower
- Magnetic fields to bend the tracks !



2. Detecting Particles

To collect all information from all particles we build a 4π multipurpose detector

... and use different layer to find different parameters

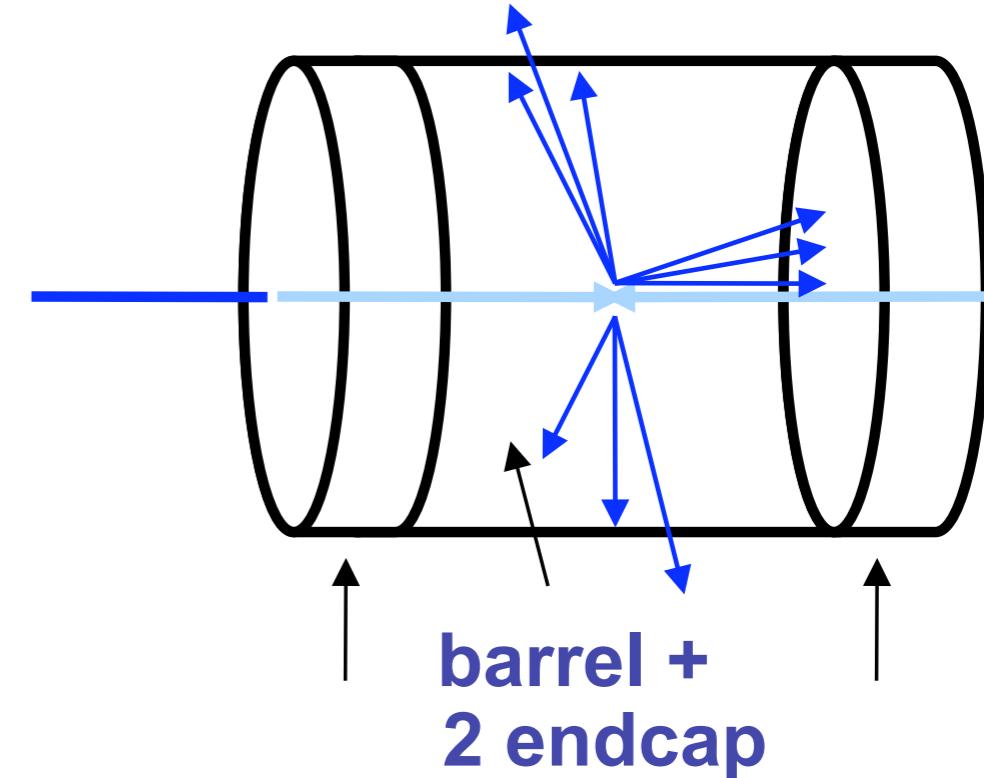
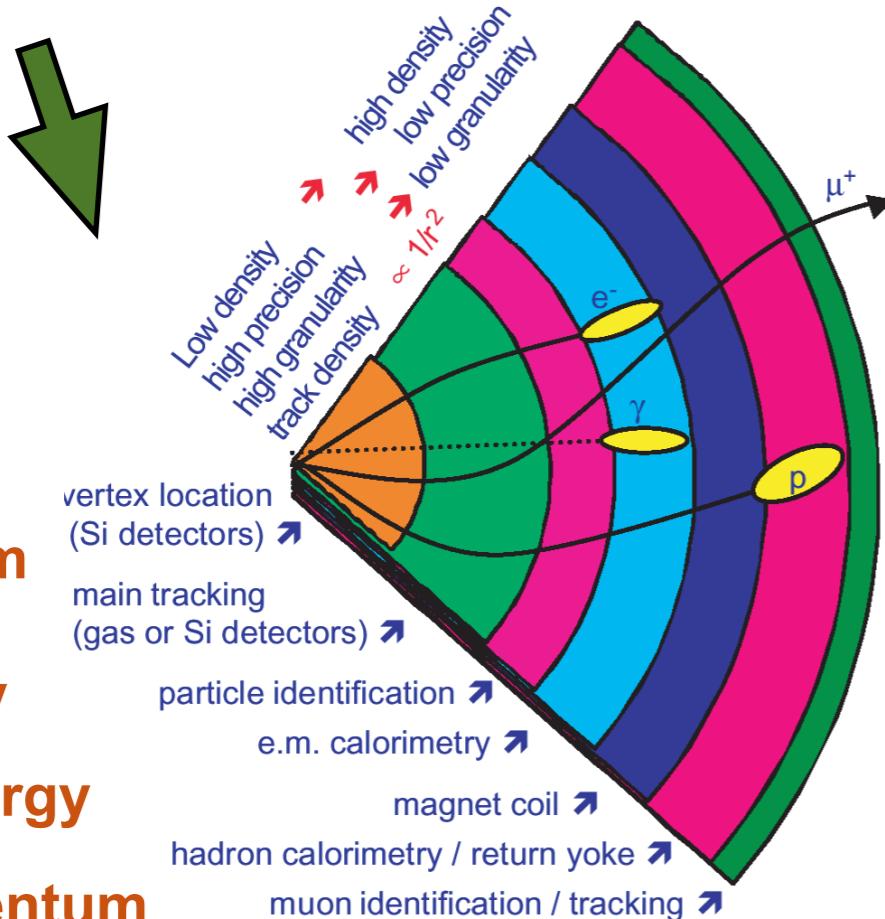
Momentum

Identity

Energy

Momentum

A perfect detector would reconstruct any type of interaction with 100 % efficiency and unlimited resolution !



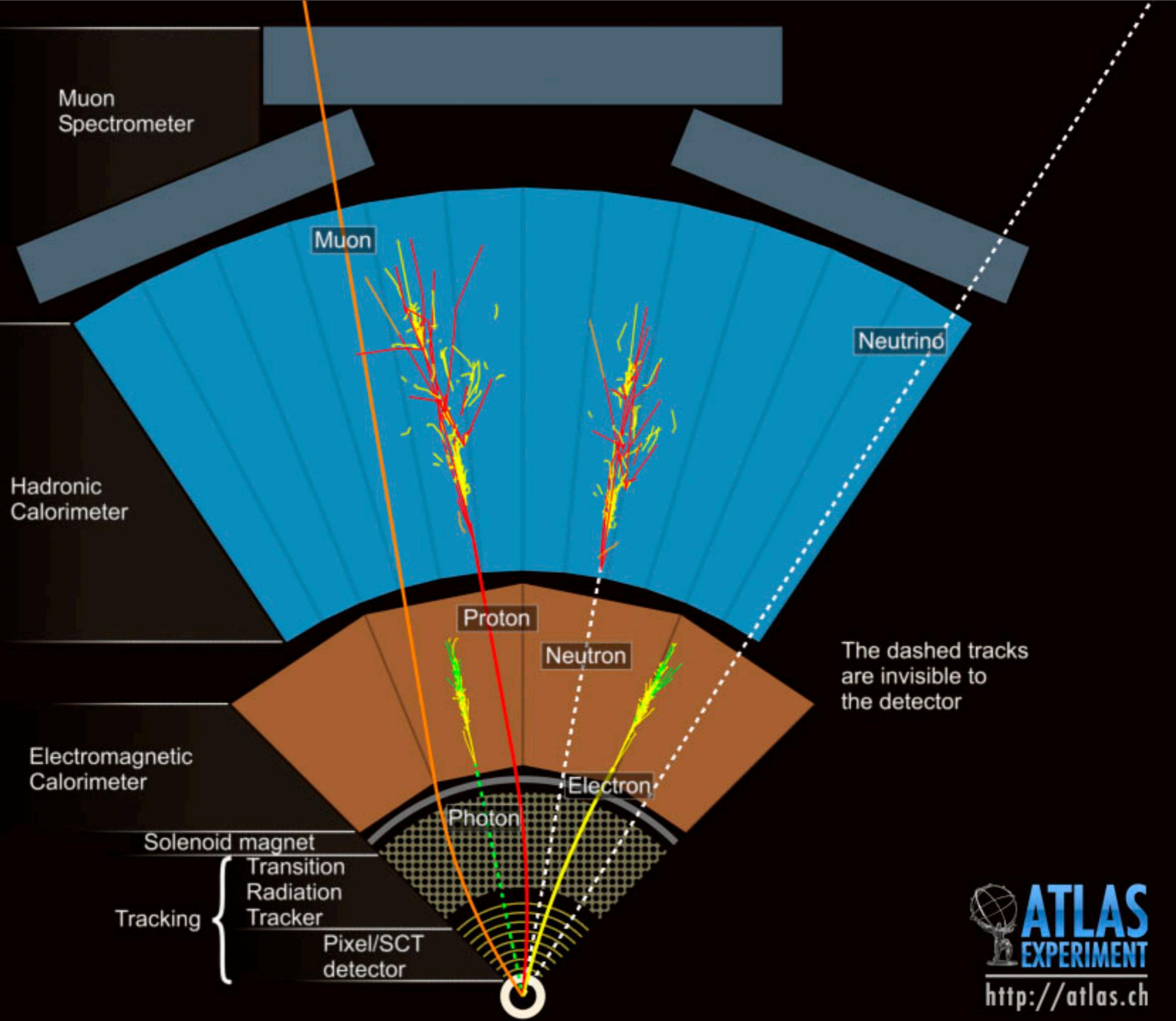
barrel +
2 endcap

What do we need to know about each particle in a collision?

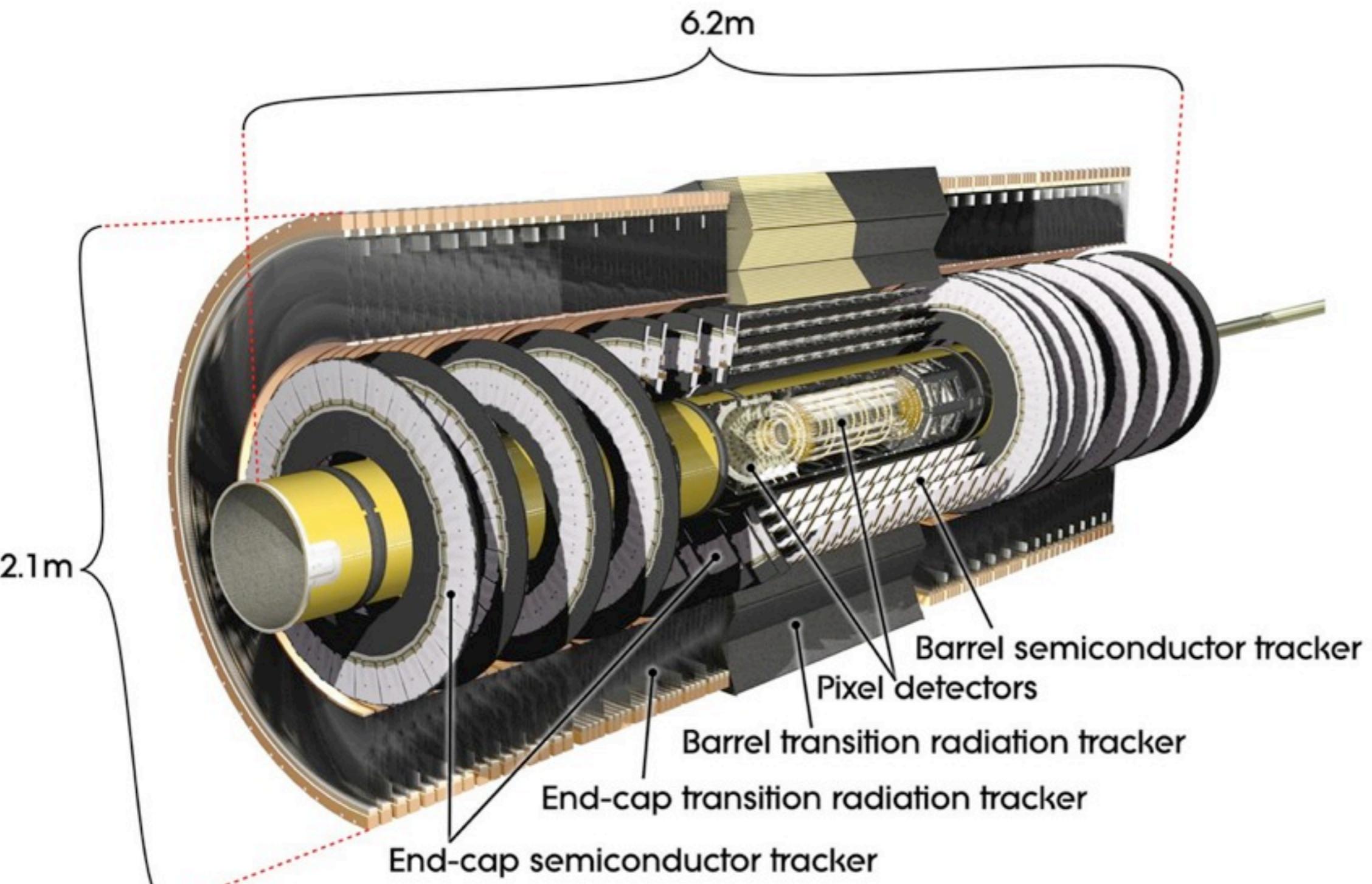
- Charge
- Mass
- Momentum-vector

- But not all particles are detected
 - Neutrinos escape (missing energy) !
 - Some particles goes through non sensitive detector areas (pipes, cables, electronics, cooling, mechanics)





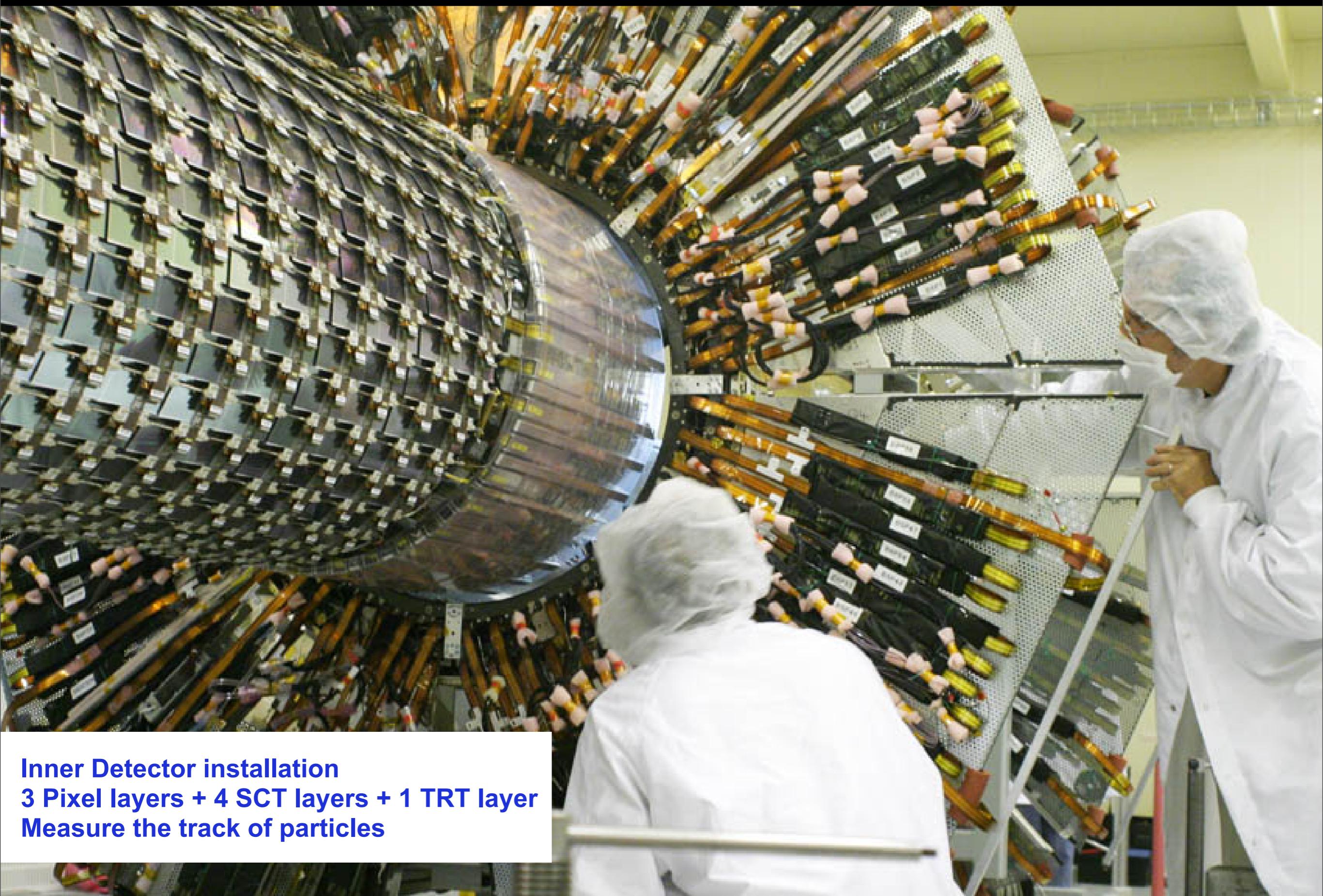
Inner Detector



2T solenoid field

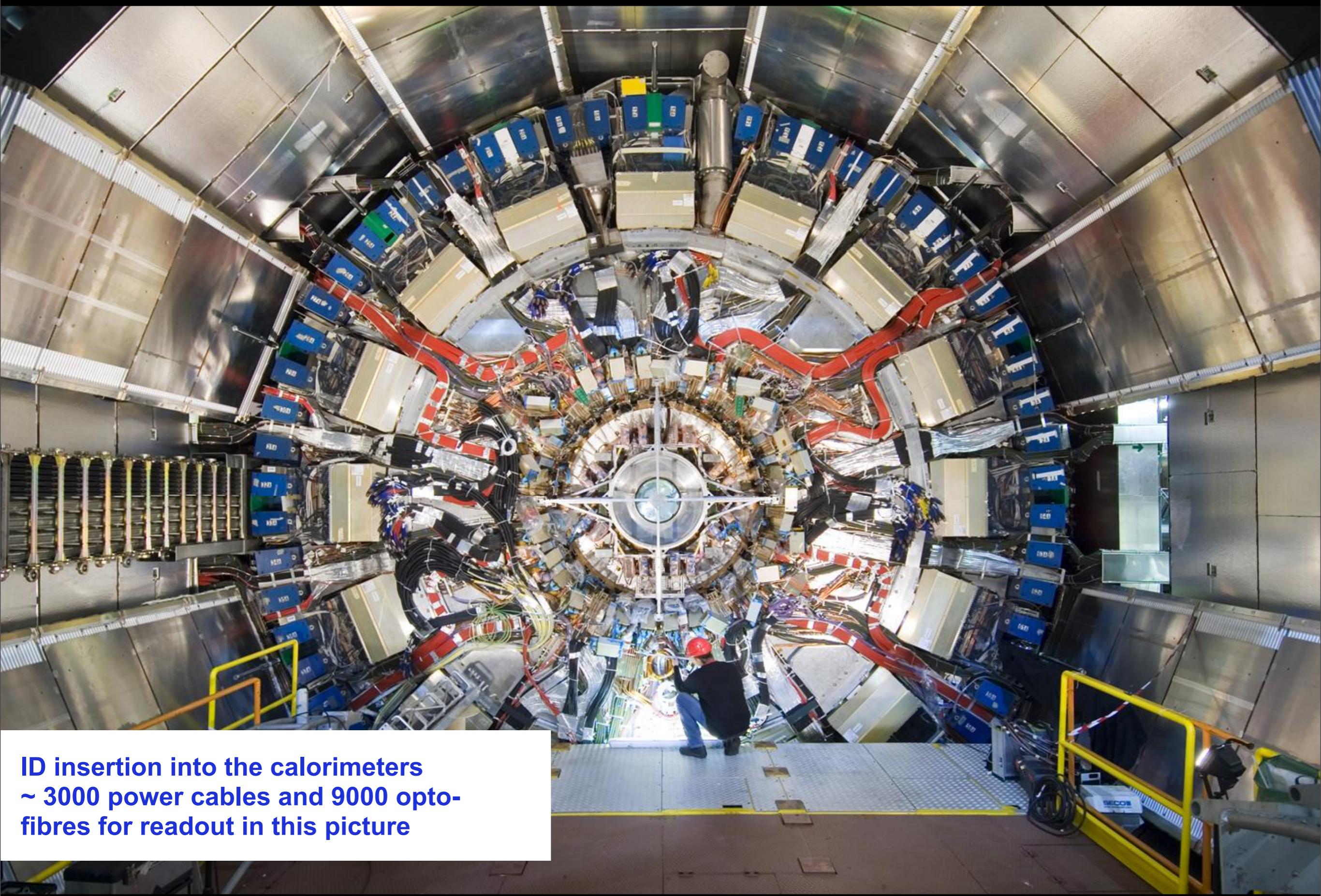


4. Constructing ATLAS



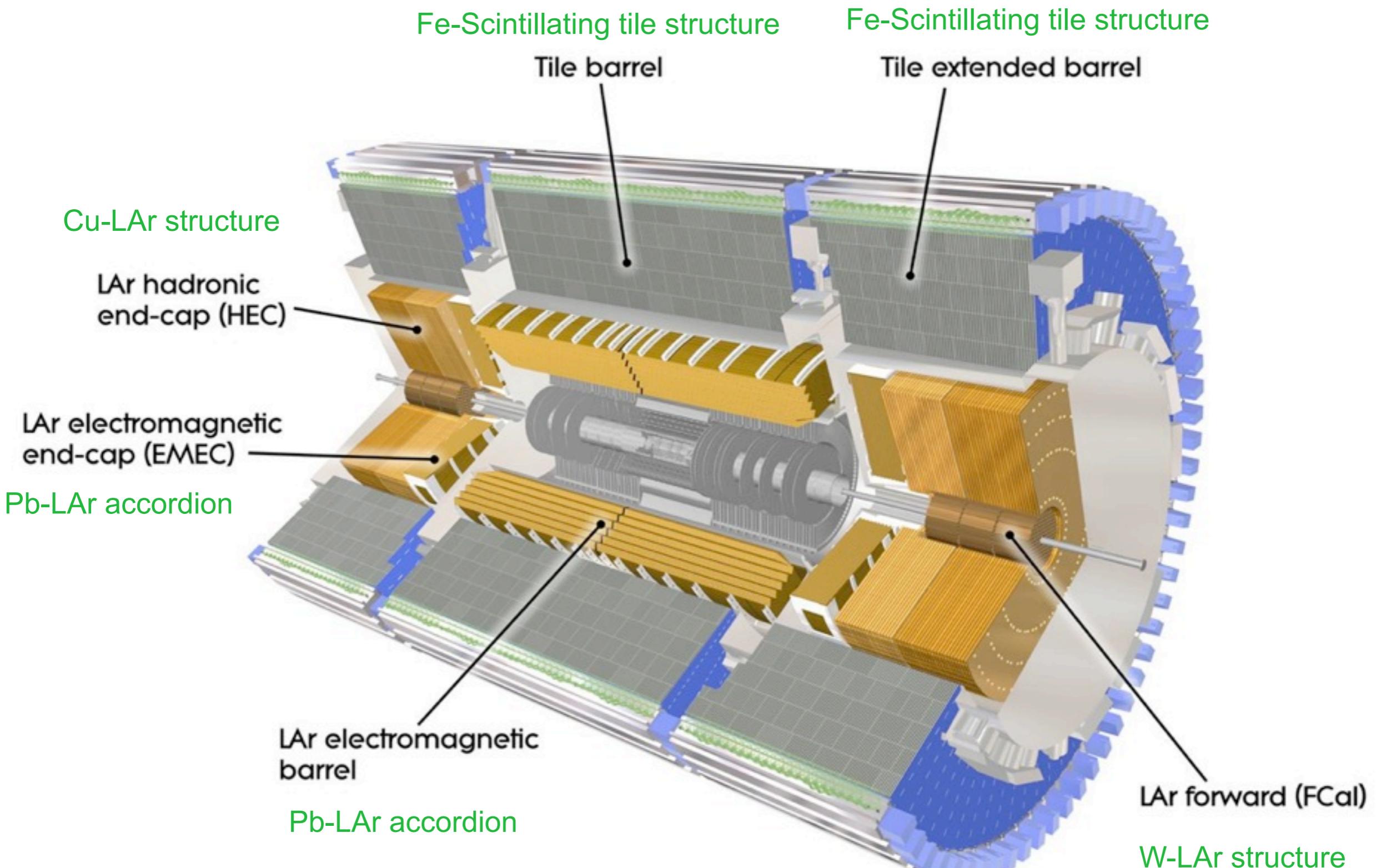
Inner Detector installation
3 Pixel layers + 4 SCT layers + 1 TRT layer
Measure the track of particles

4. Constructing ATLAS

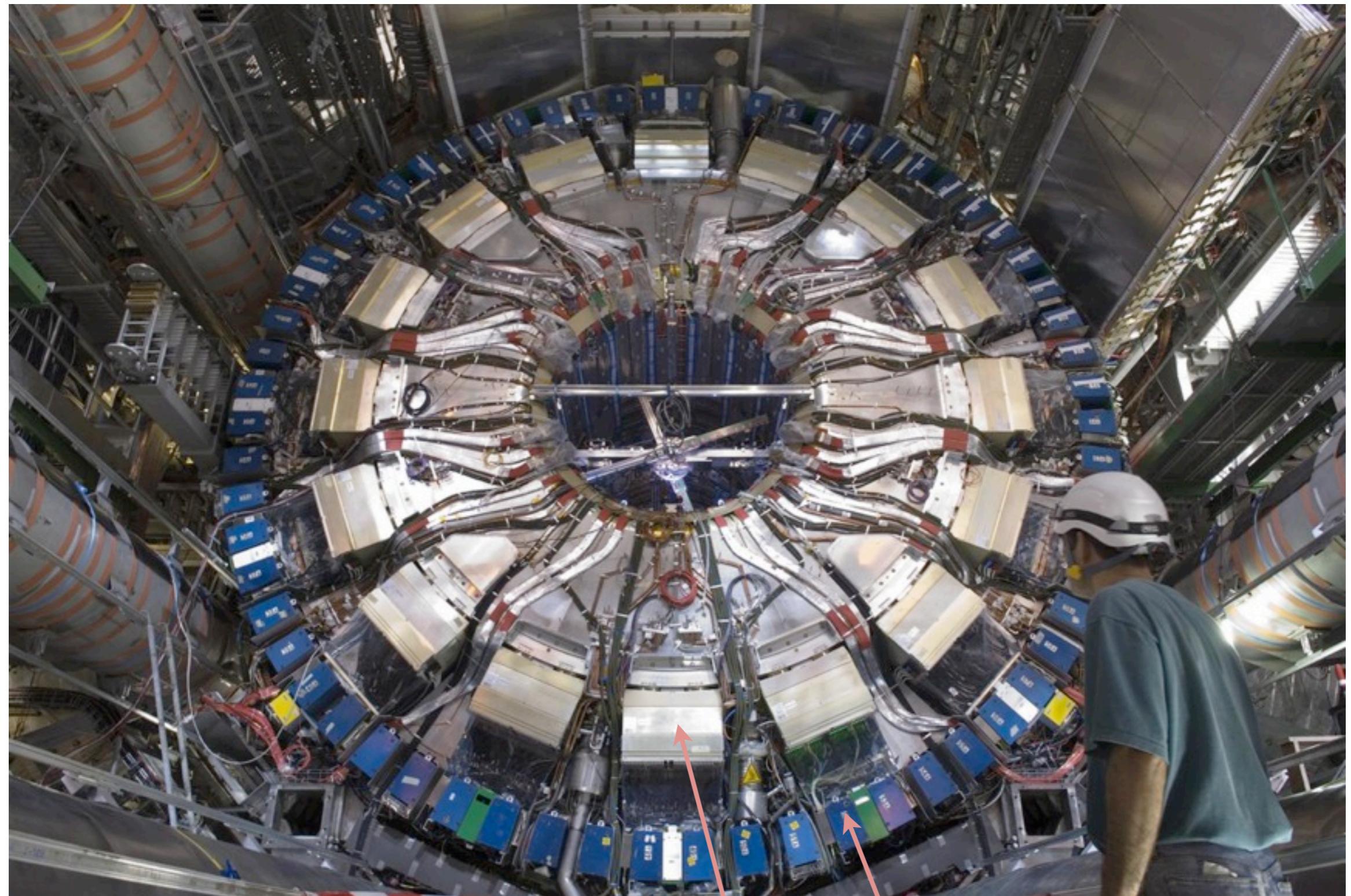


ID insertion into the calorimeters
~ 3000 power cables and 9000 opto-fibres for readout in this picture

Calorimeters



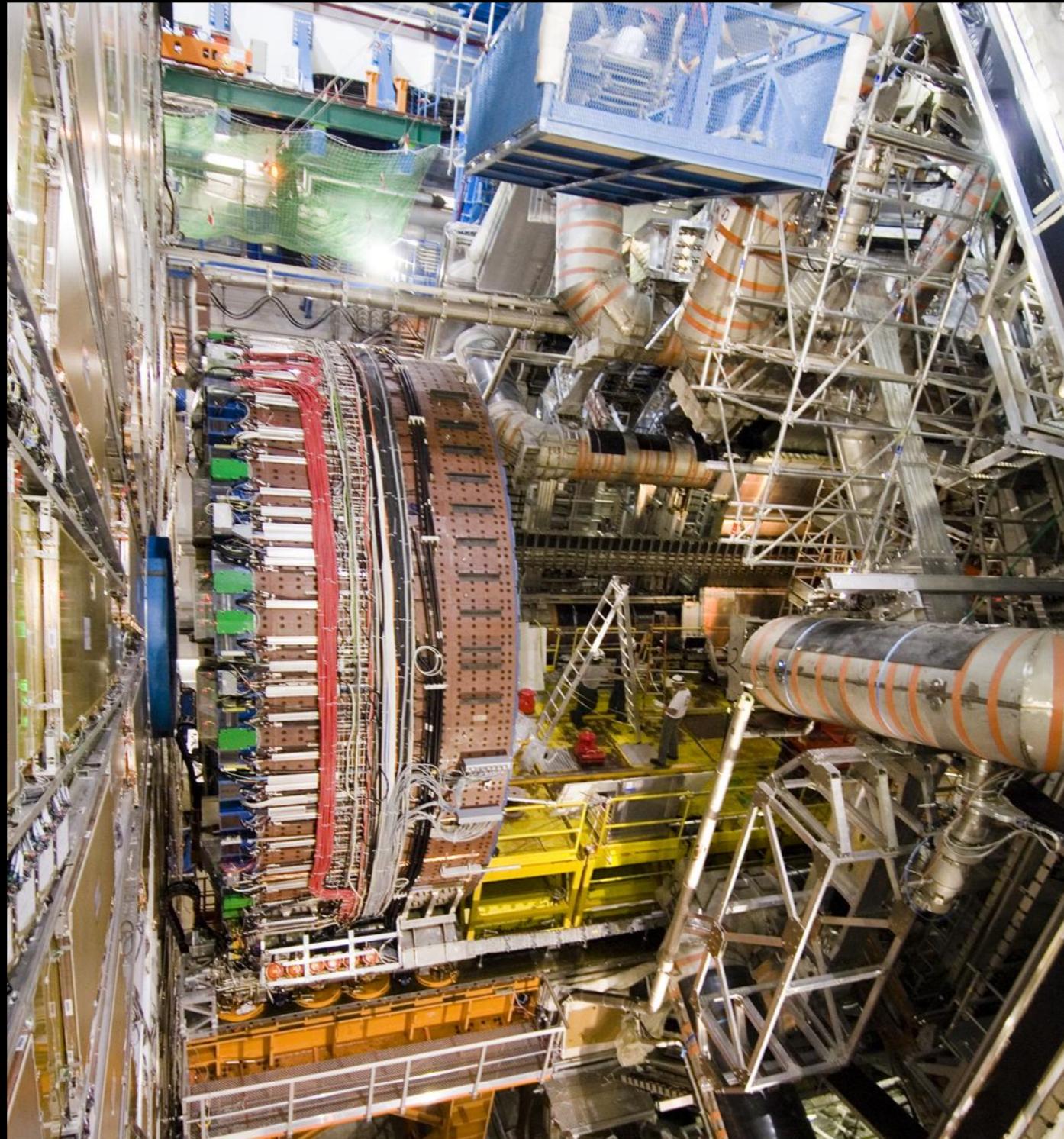
Calorimeters



LAr TileCal



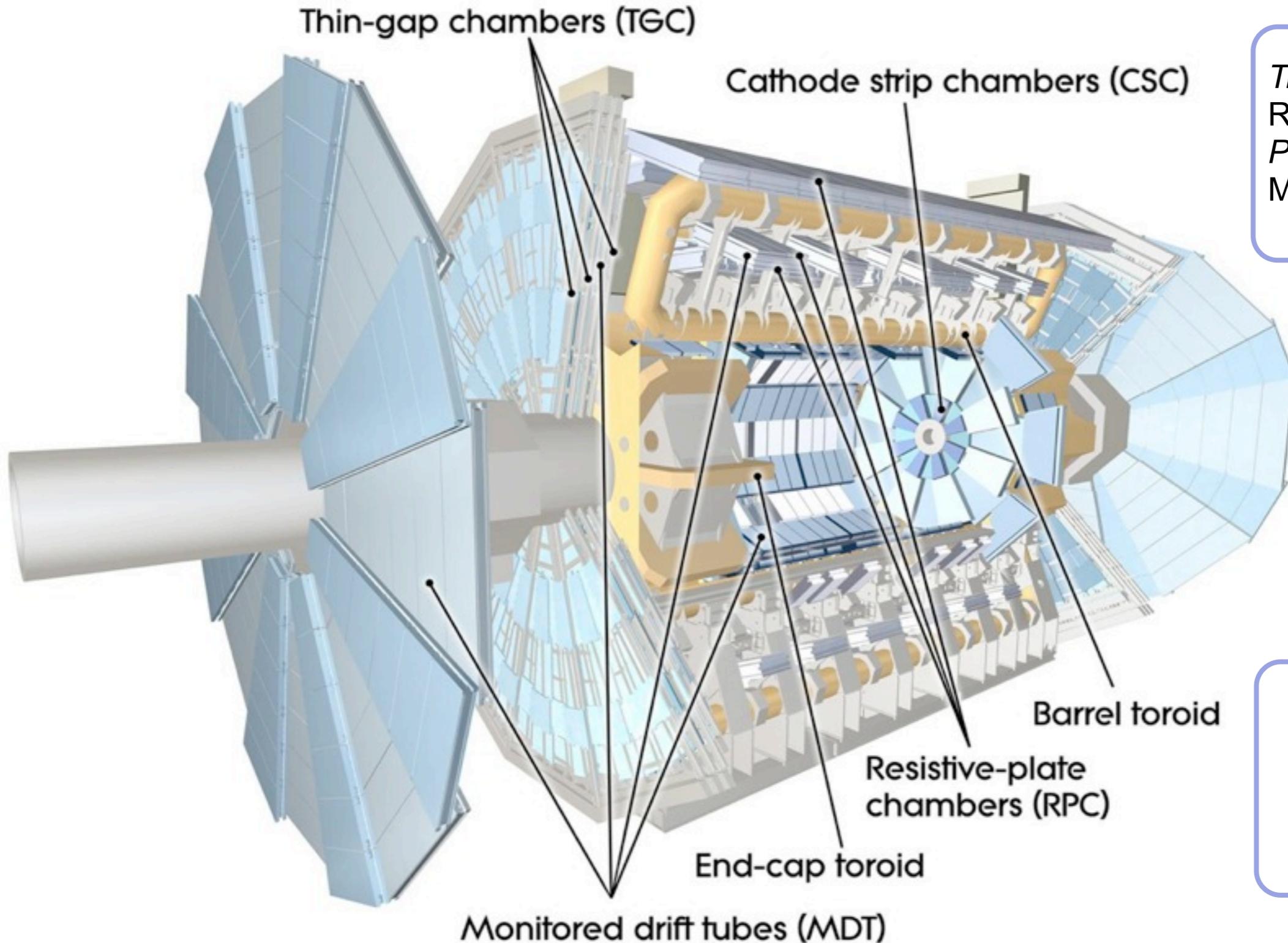
4. Constructing ATLAS



**Calorimeter installation
Barrel + 2 endaps, movable on rails
Measure the particle energies**



Muon Spectrometer

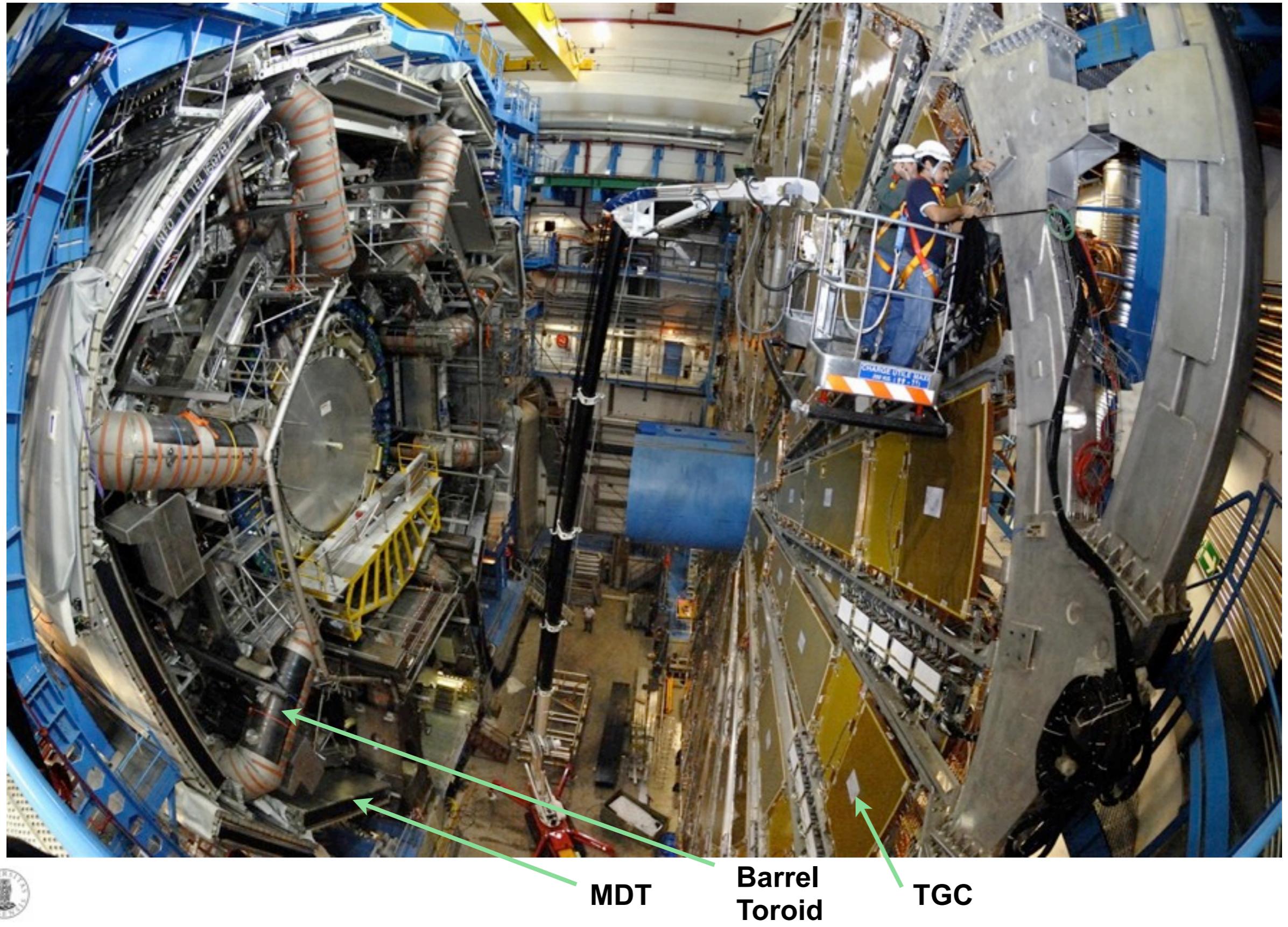


Triggers:
RPC and TGC
Precision:
MDT and CSC

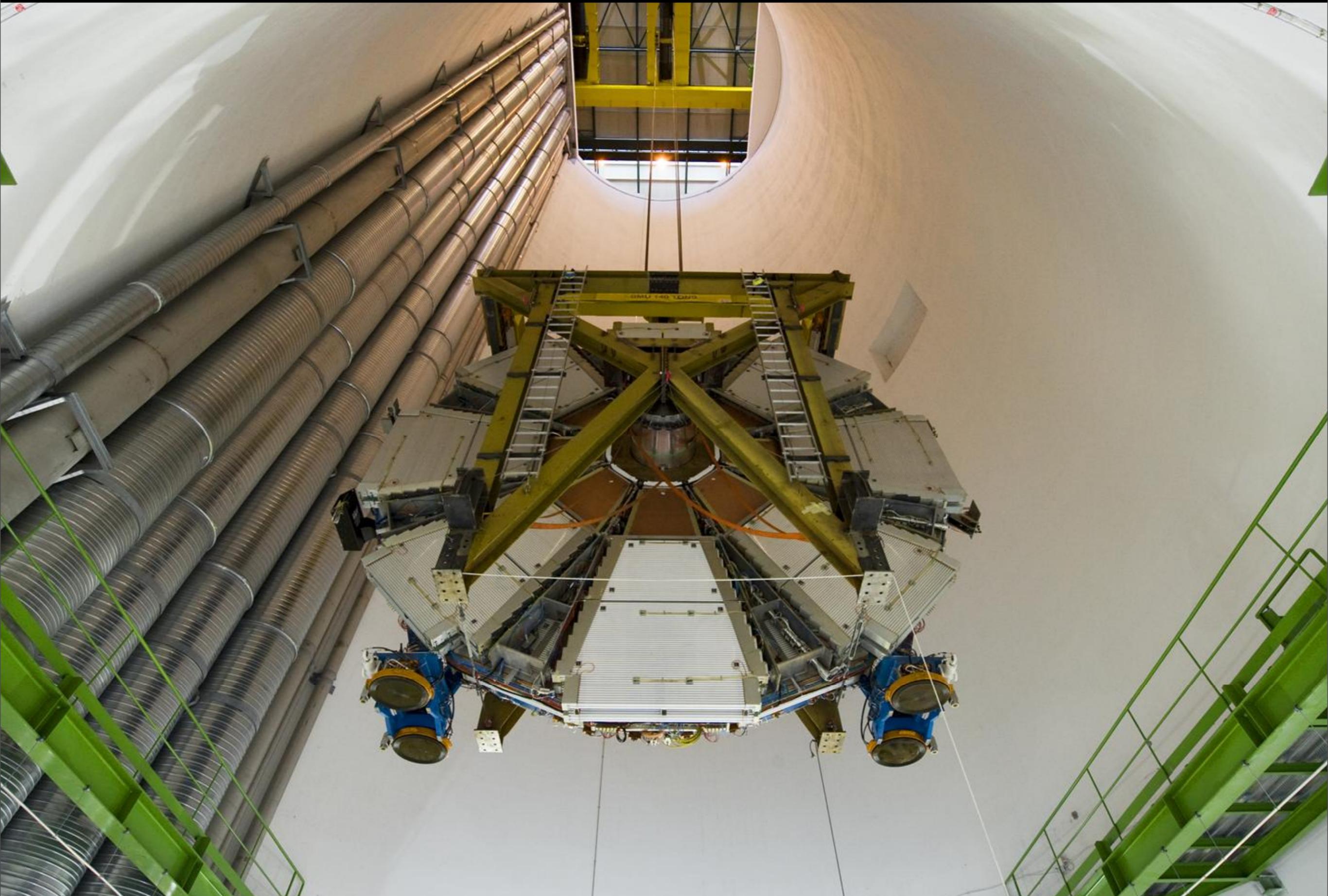
Toroid field:
Peak field 4 T
Bending power
~ 2-5 Tm



Muon Spectrometer



4. Constructing ATLAS

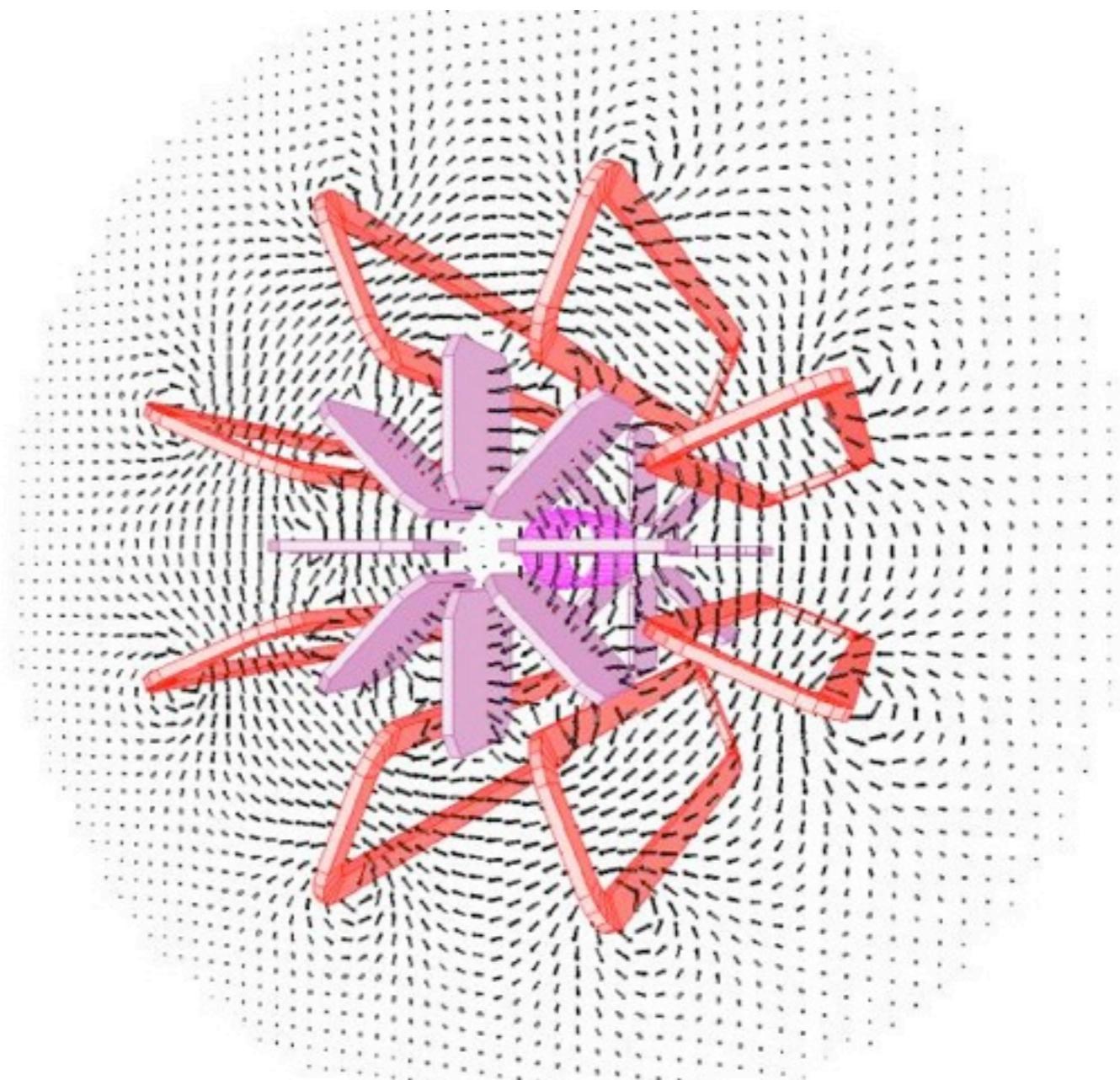
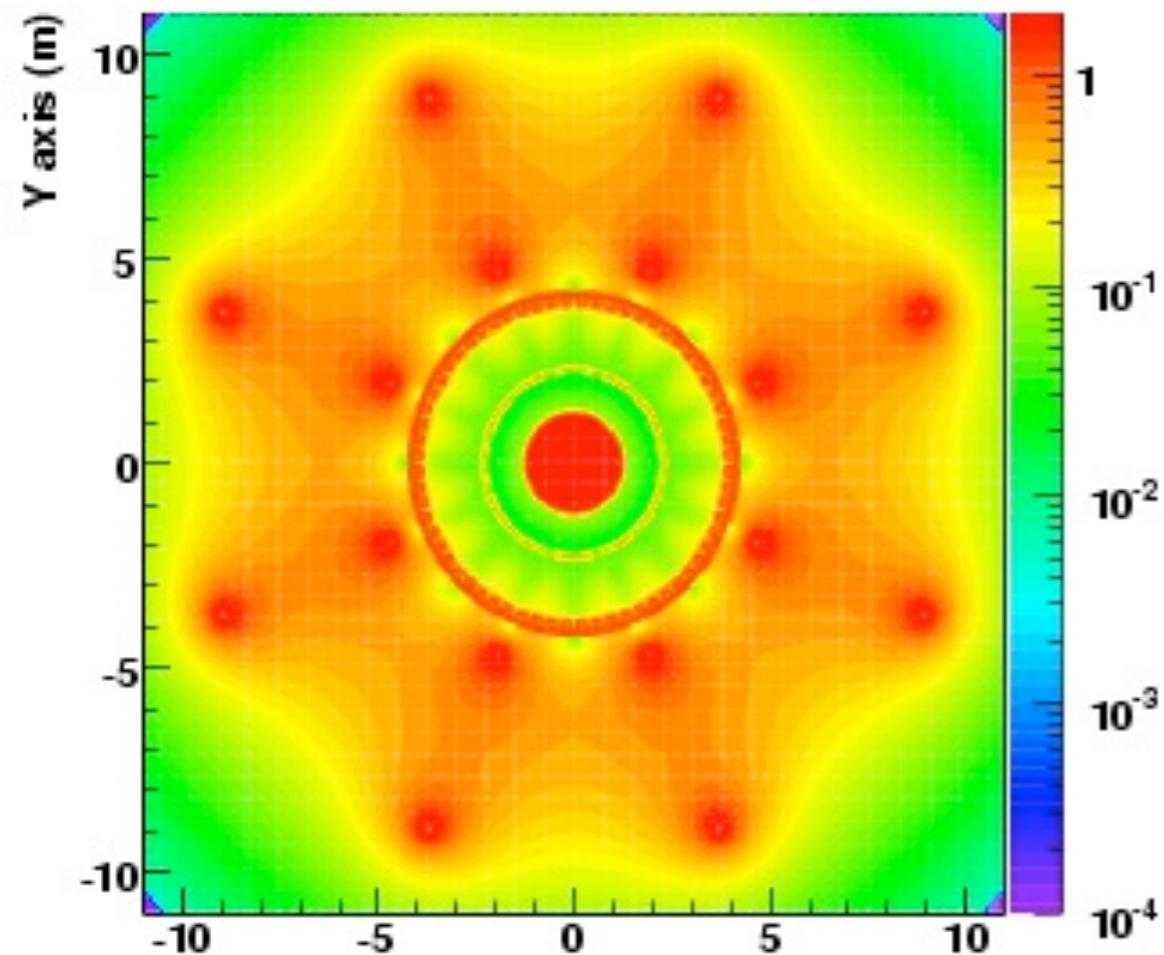


Magnet system

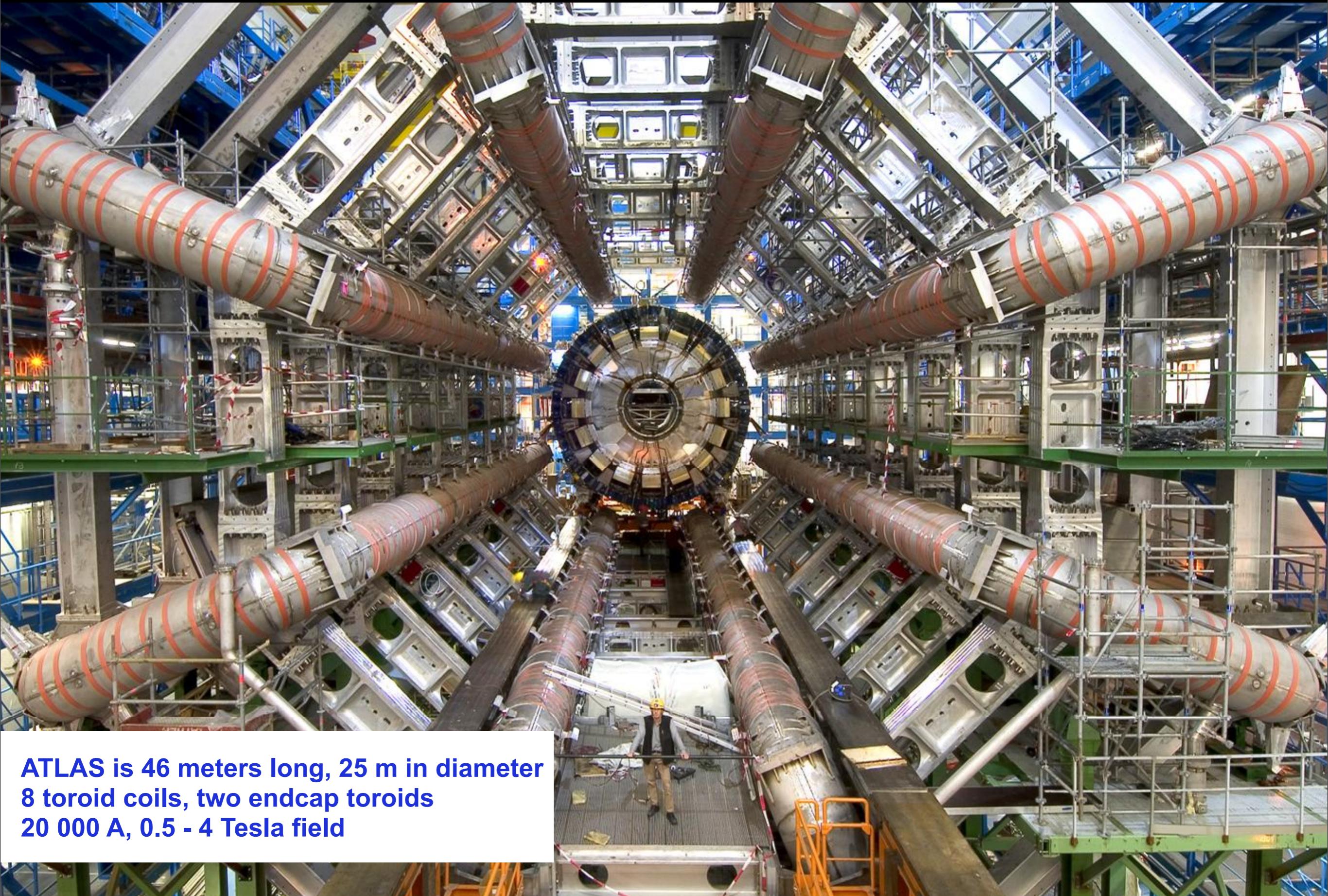
Solenoid field: 2T

Toroid field: Peak field 4 T, Bending power ~ 2-5 Tm

$z = -20\text{cm}, \phi l = 2\pi l$



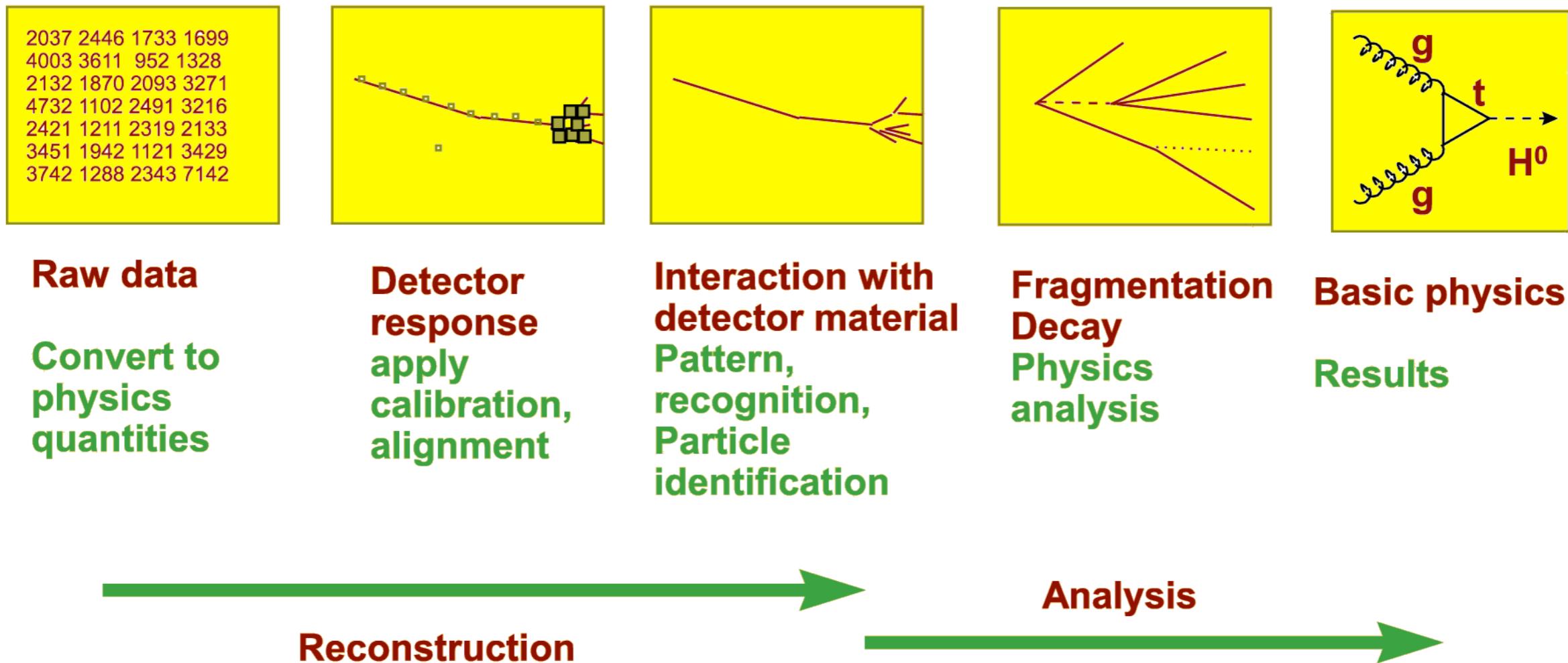
4. Constructing ATLAS



**ATLAS is 46 meters long, 25 m in diameter
8 toroid coils, two endcap toroids
20 000 A, 0.5 - 4 Tesla field**

5. Data readout

From raw data to physics

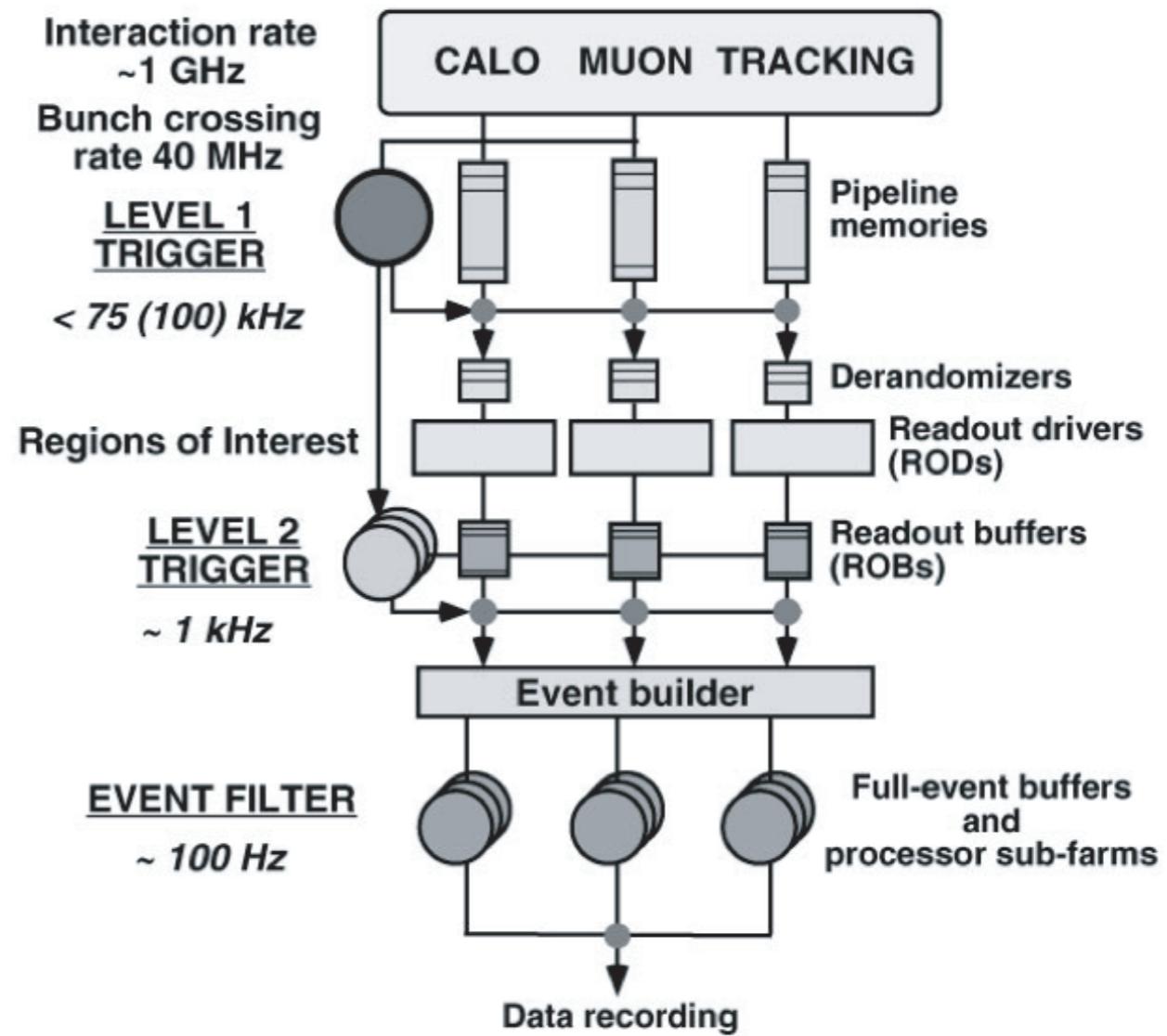
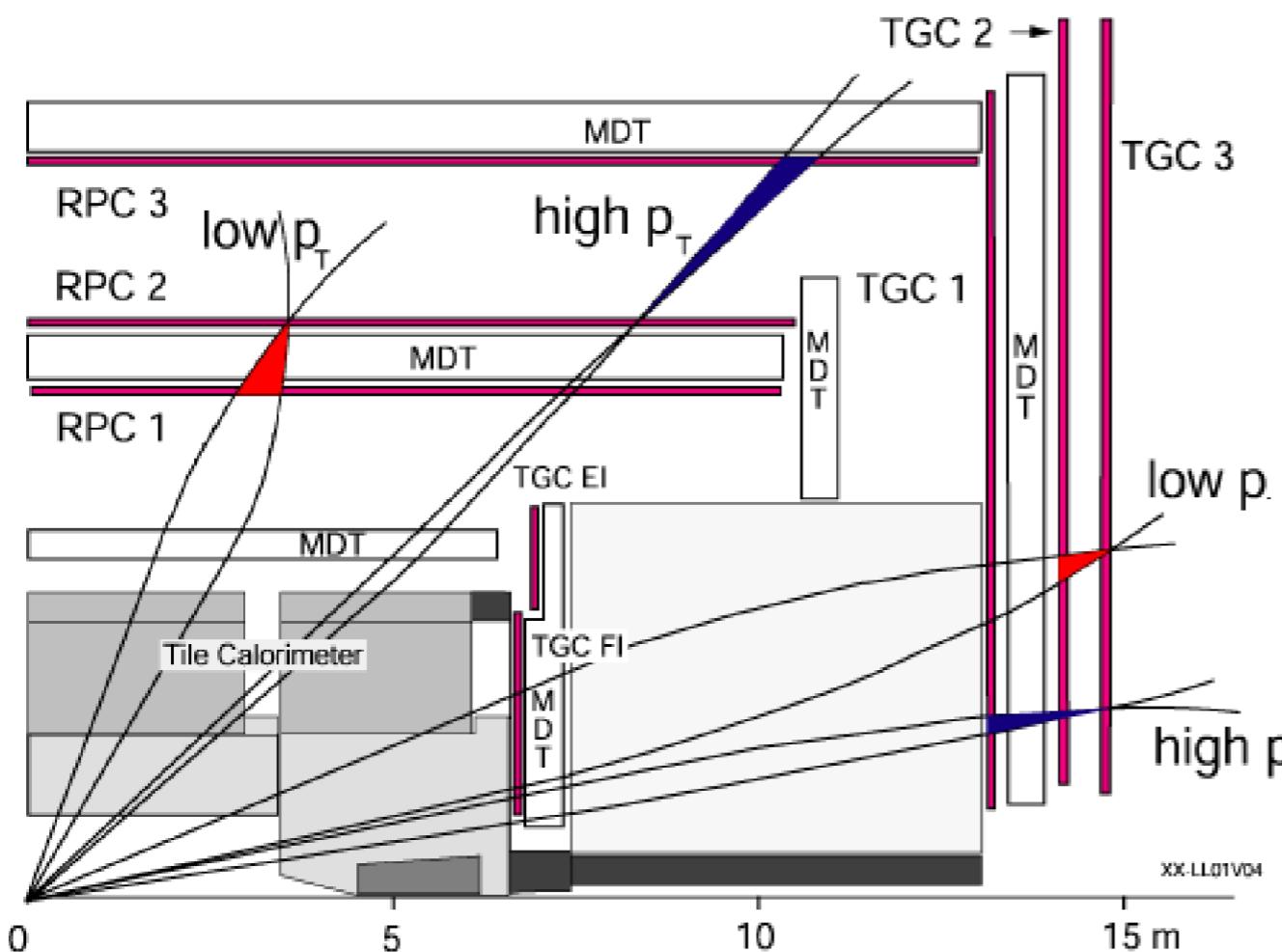


- Raw data strongly reduced by trigger and event selection (~40MHz to ~50kHz)
- Will record raw data at a rate of 400MB/s for Atlas & CMS
- Need a lot (!) of computing power to reconstruct and analyse data

5. Trigger system

Dedicated fast detectors provides fast readout of the collisions !

- Muon system and Calorimeter have fast detectors
- This information is fed back to the LVL1 trigger who removes all non-interesting events based on a set of trigger criterias



The trigger system reduces the data from ~ 1 GHz to 100 Hz !

5. Data readout



**ATLAS - ~ 100 racks / 2500 highest performance
mult-core PCs in the final system**

Computing centre

5. Data readout

Simulations:

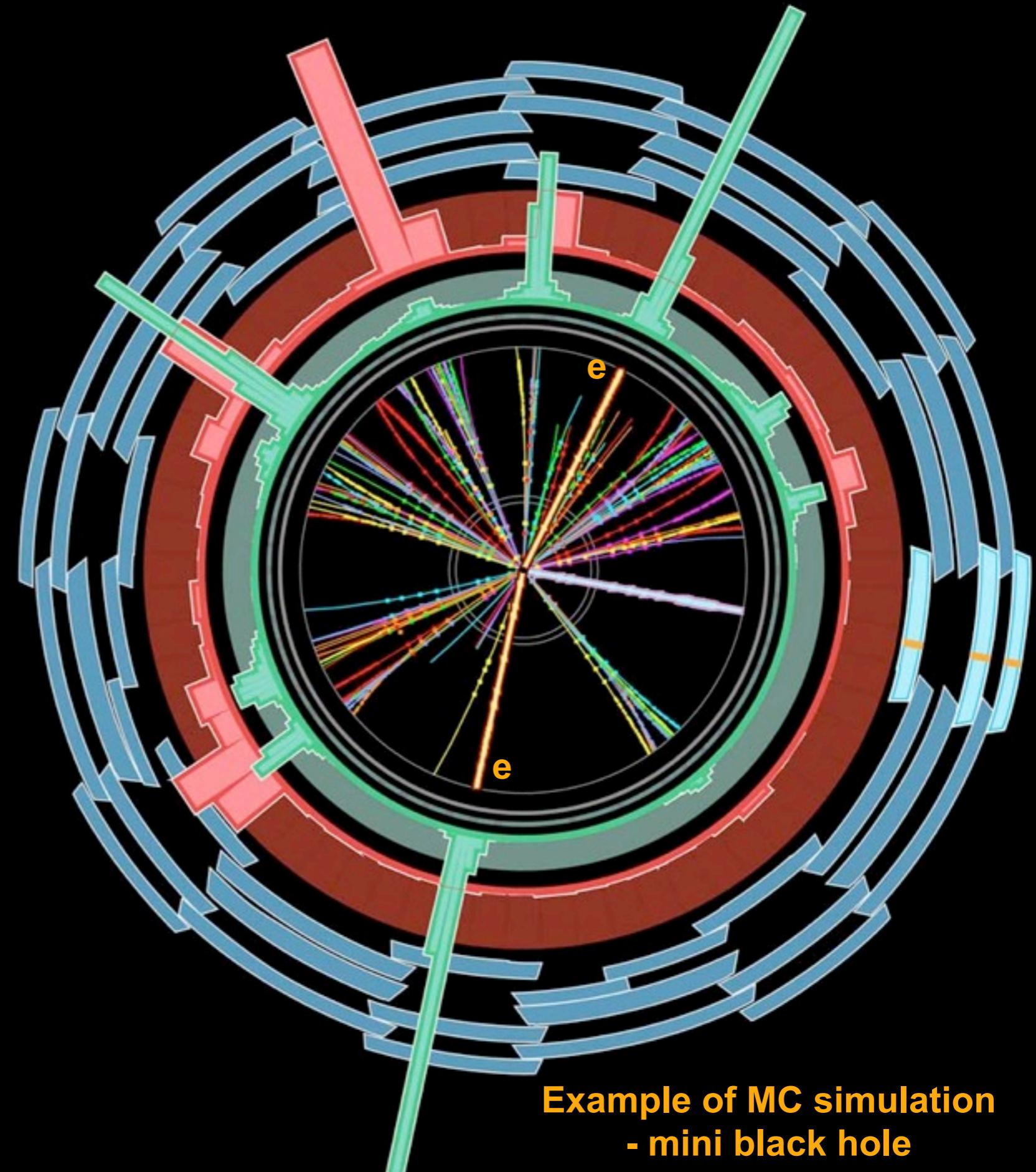
- We use Monte-Carlo simulations to study the behaviour of particles when they pass through ATLAS
- Signal + Background
- Standard Model signals & new physics

Cosmic radiation:

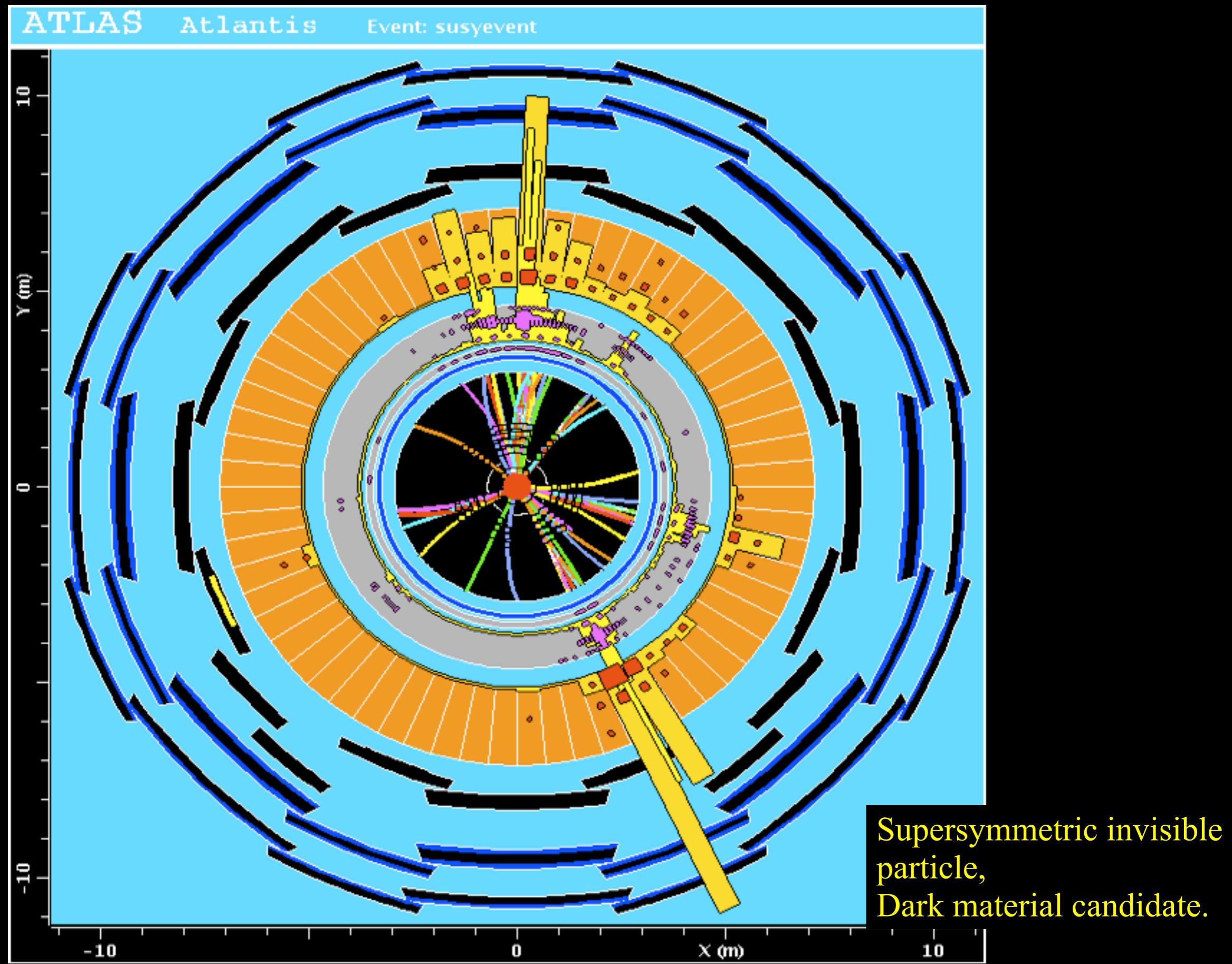
- To verify and improve the detector performance before LHC start

Data from LHC:

- Some data from LHC are being looked at to learn about the beam and the detector response to single beam operation

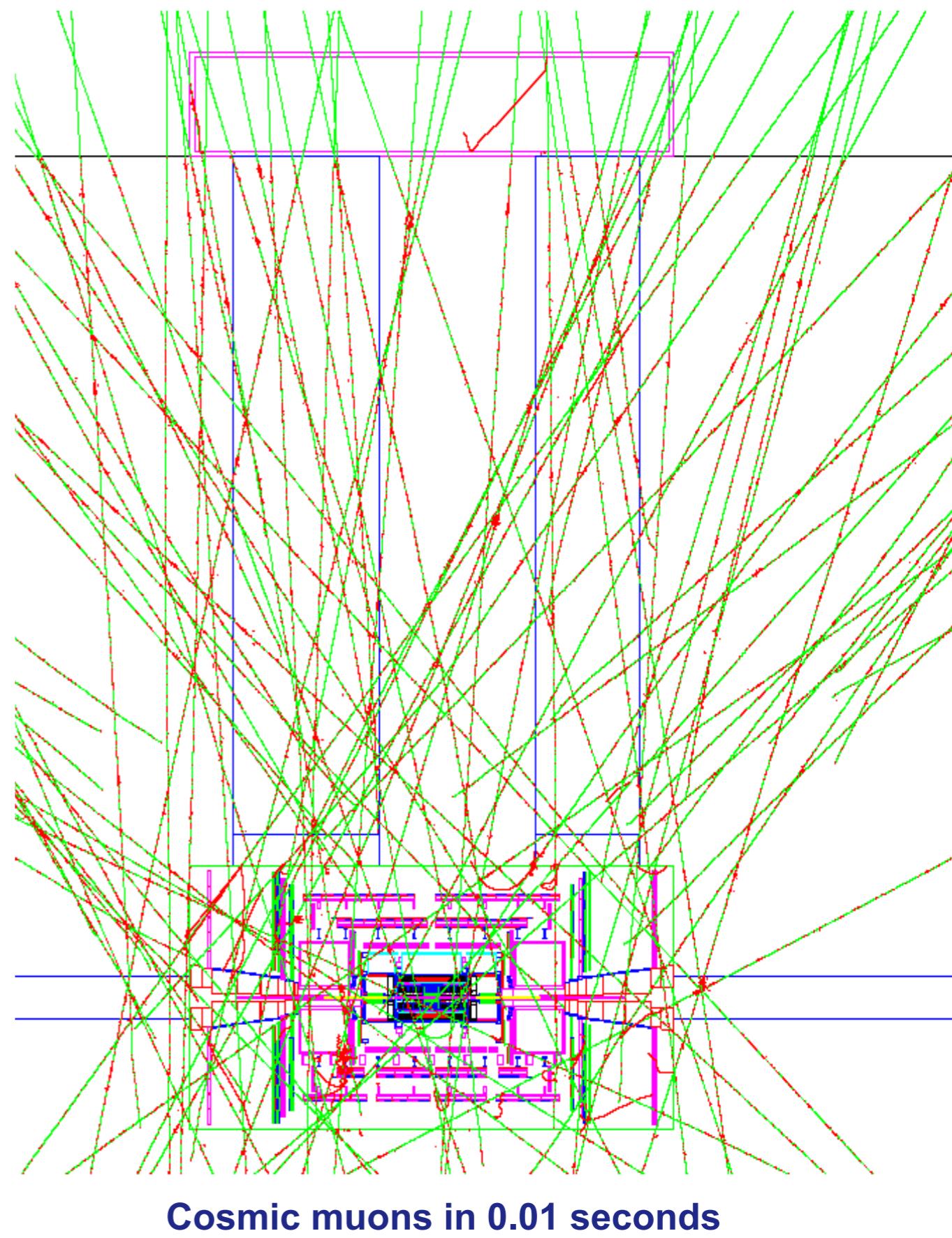


5. Simulated Data

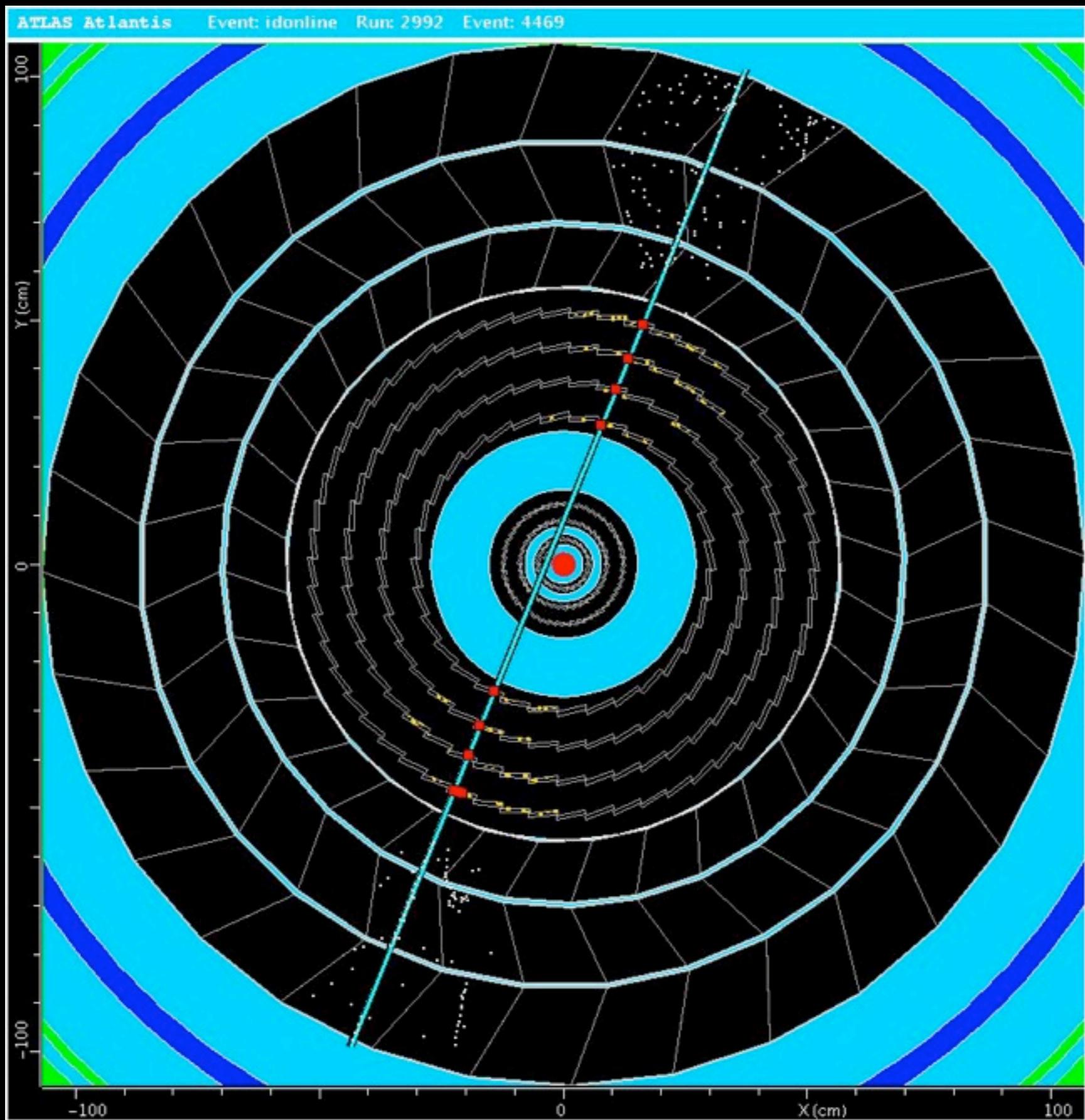


5. Cosmic radiation

- **Combined Inner Detector runs**
 - DAQ, Online, Offline software integration
 - Final tuning of readout parameters
 - Detector calibration, finding dead or noisy channels
 - Detector synchronisation
 - Track reconstruction and Alignment
 - Physics performance studies
- **Combined ATLAS runs**
 - Using the ATLAS trigger system
 - Integrated ATLAS software
 - Combined runs with the Calorimeter and the Muon systems
 - Building ATLAS events



5. Cosmic Ray Data

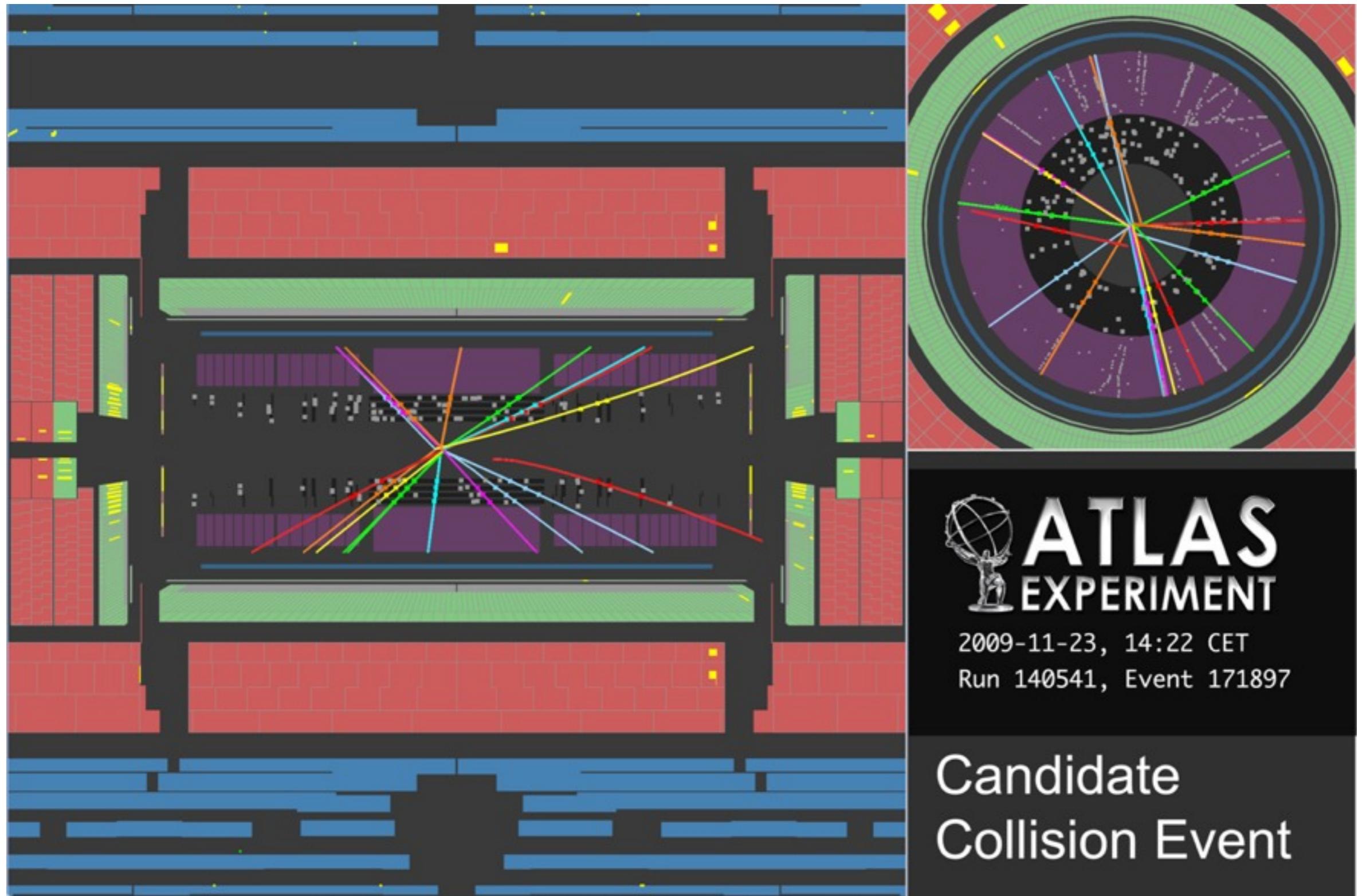


Cosmic ray
through the ID
Offline data

First LHC beam - 20.11.2009



First beam, 900 GeV collisions - 20.11.2009



<http://atlas.web.cern.ch/Atlas/public/EVTDISPLAY/events.html>

SCT in standby, Pixel off, no solenoid field

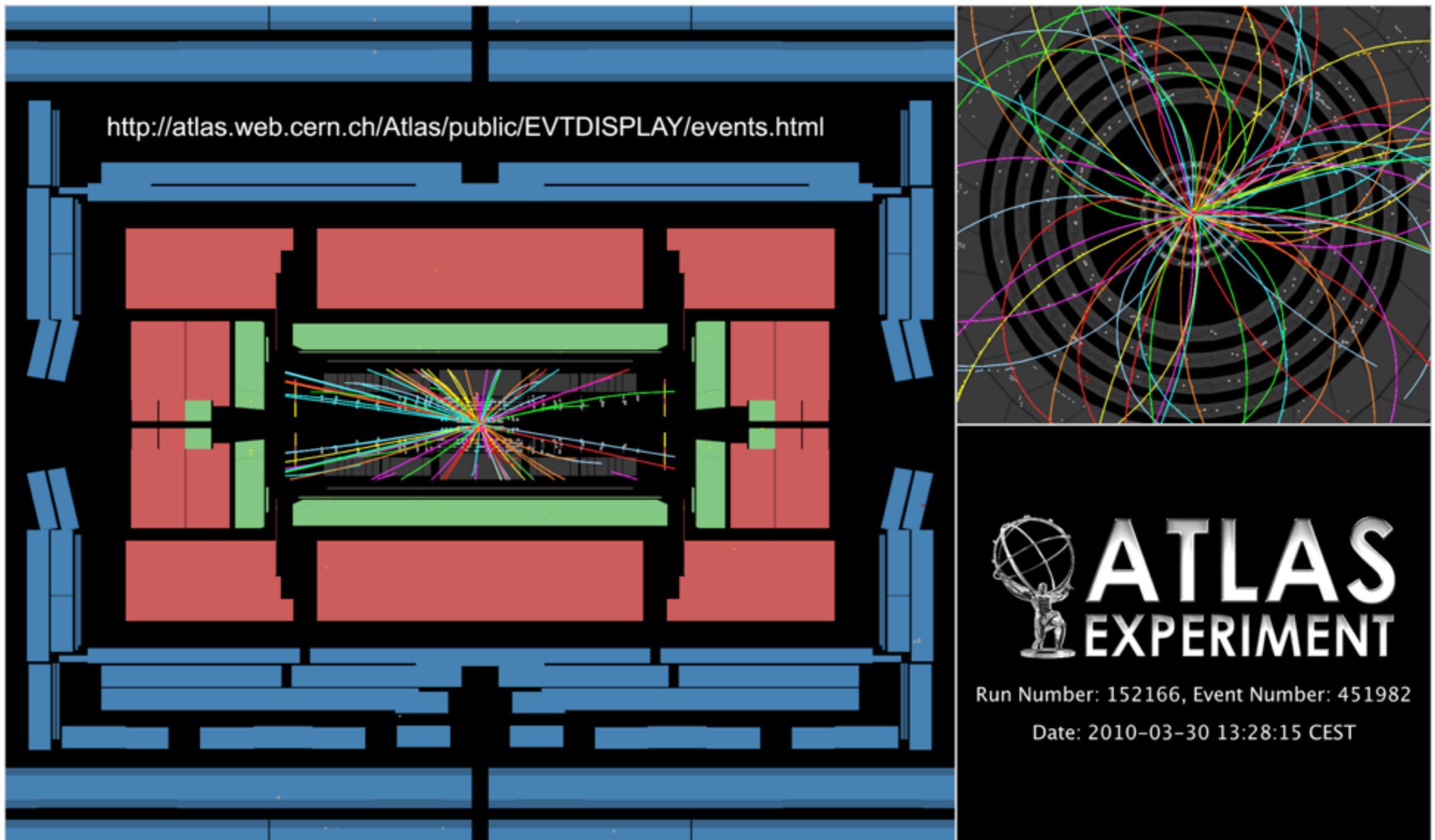


First 7 TeV collisions - 30.03.2010



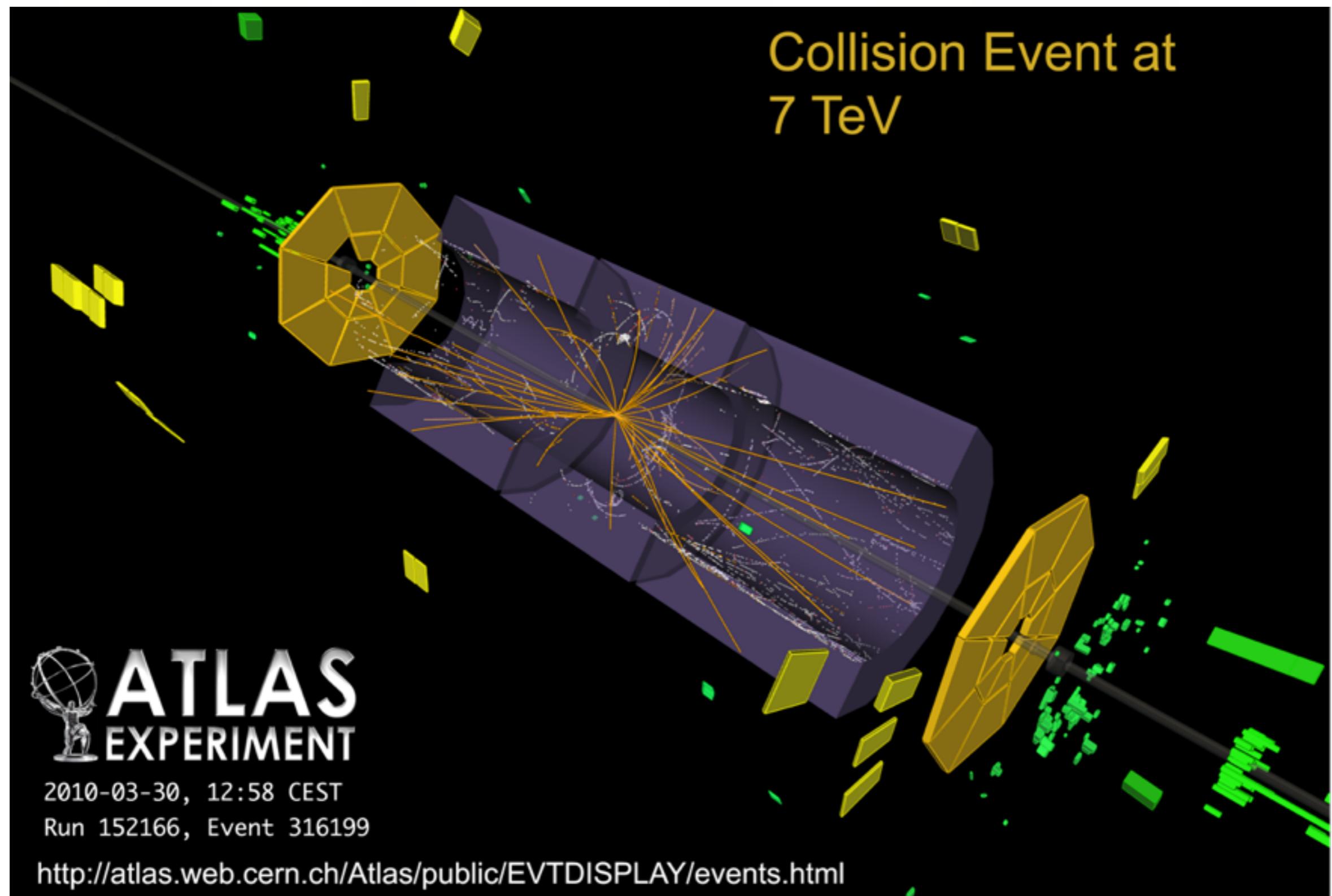
"LHC and Beyond" - 08.06.2010 - Heidi Sandaker

First 7 TeV collisions - 30.03.2010

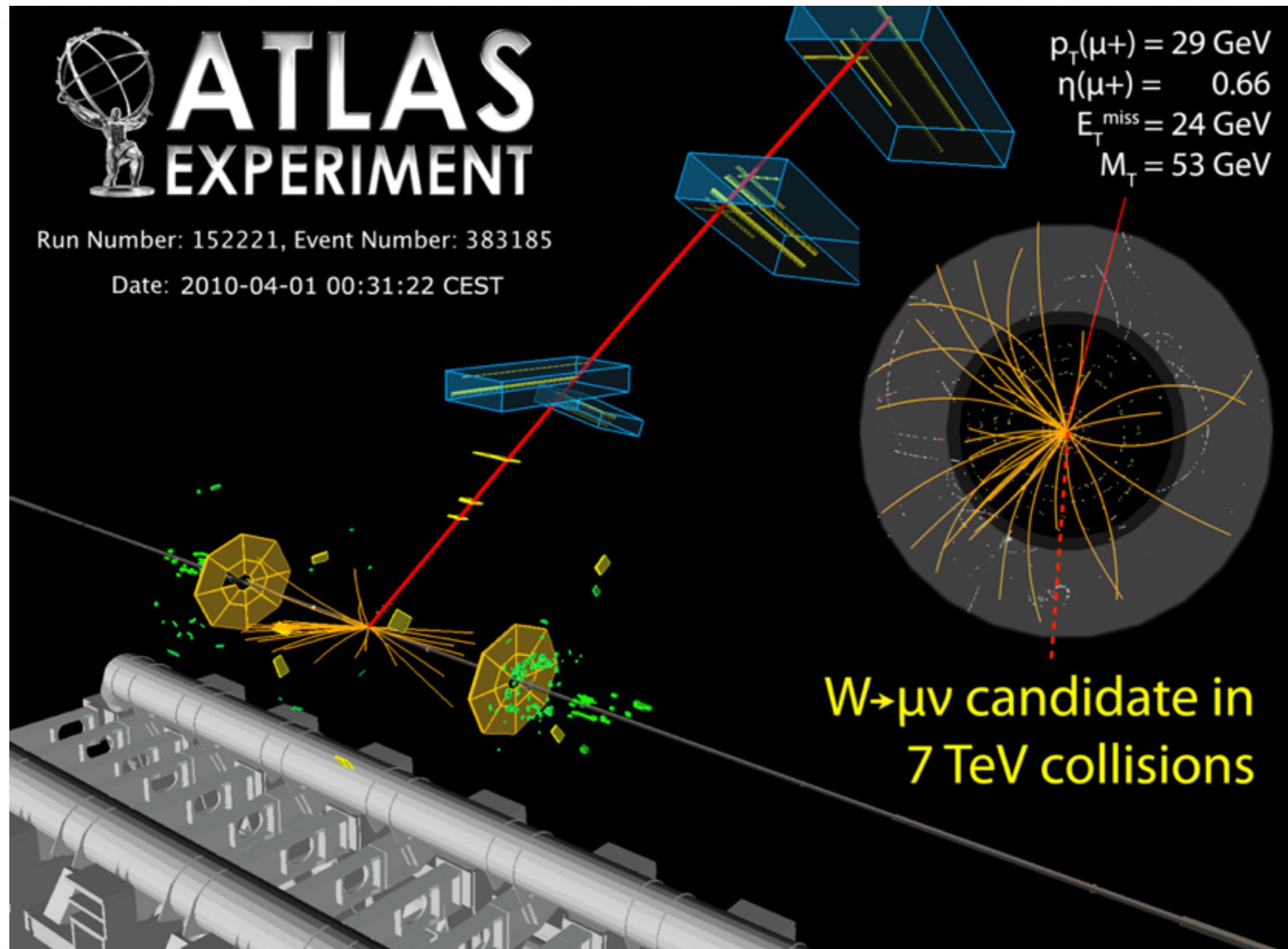


All Inner Detectors and solenoid on !

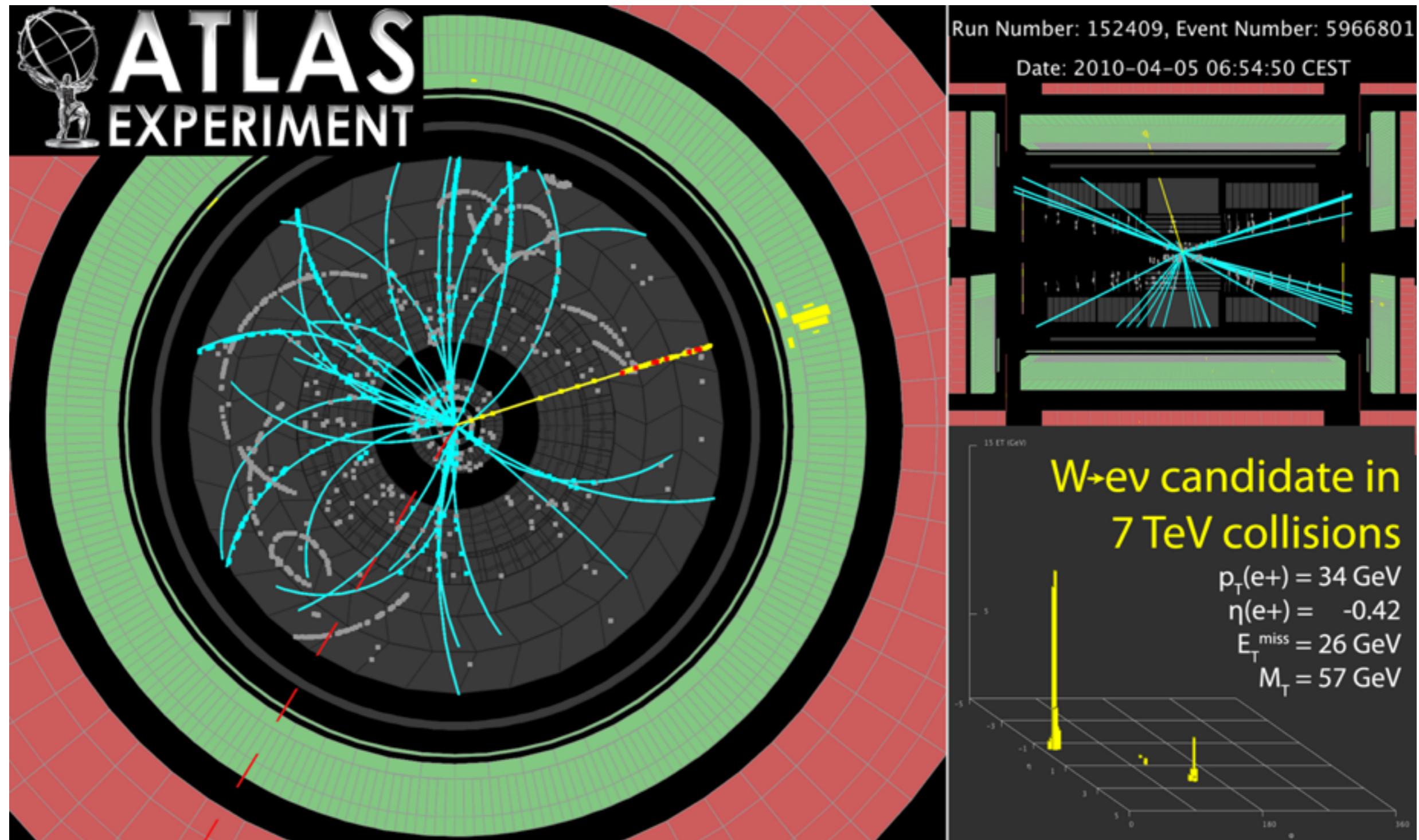
Trigger System



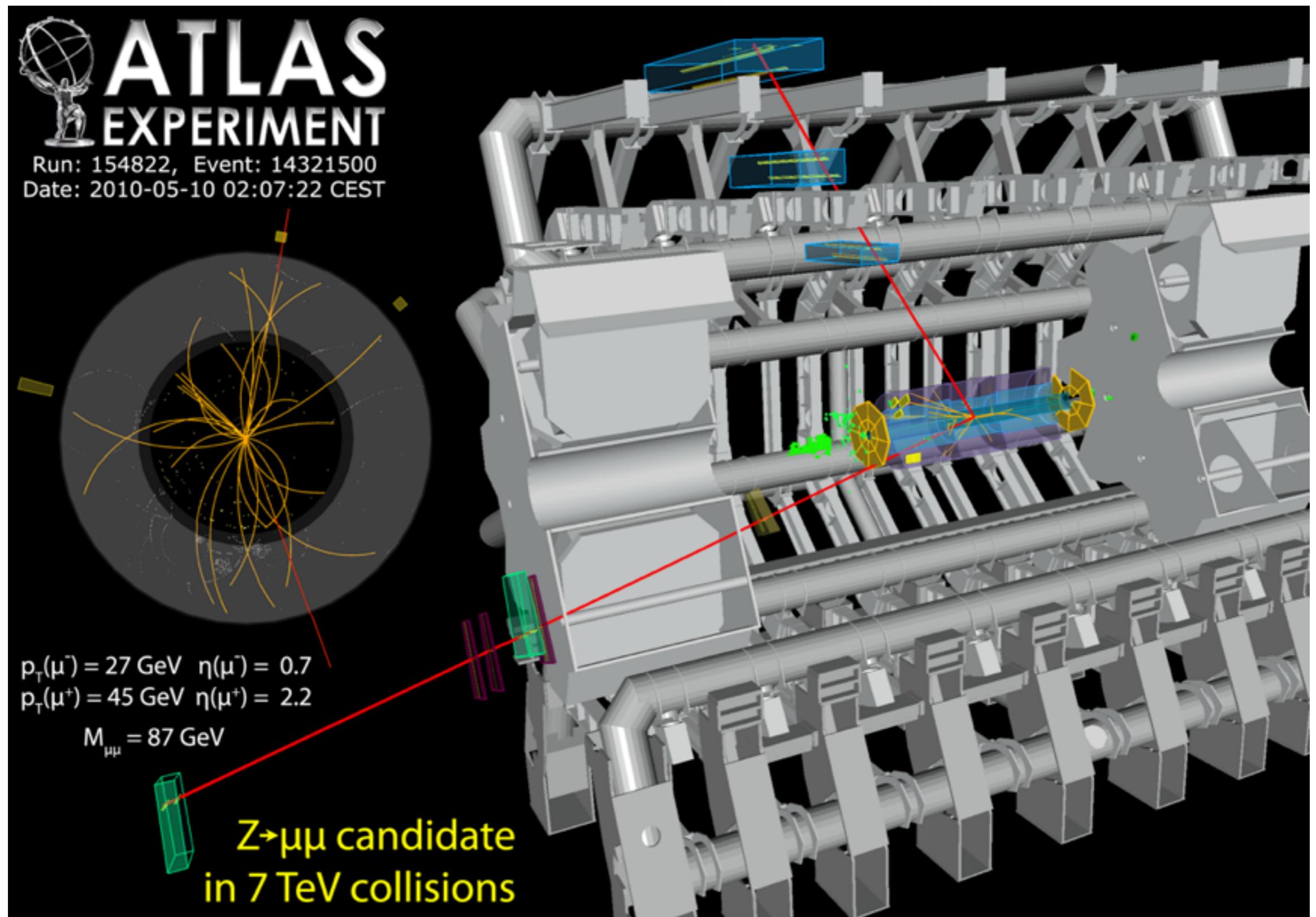
$W \rightarrow \mu\nu$ candidate in 7 TeV



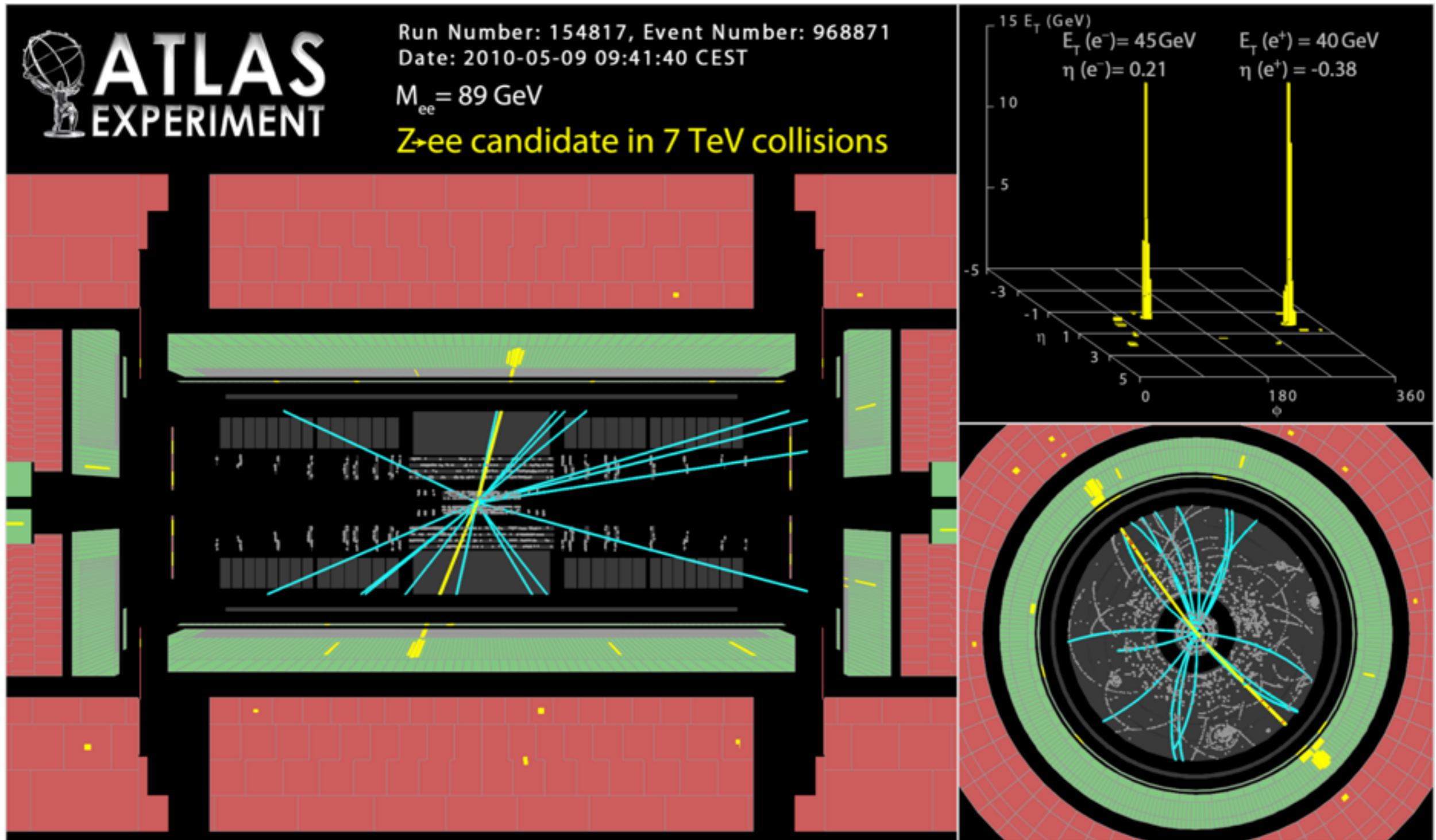
$W \rightarrow e\nu$ candidate in 7 TeV



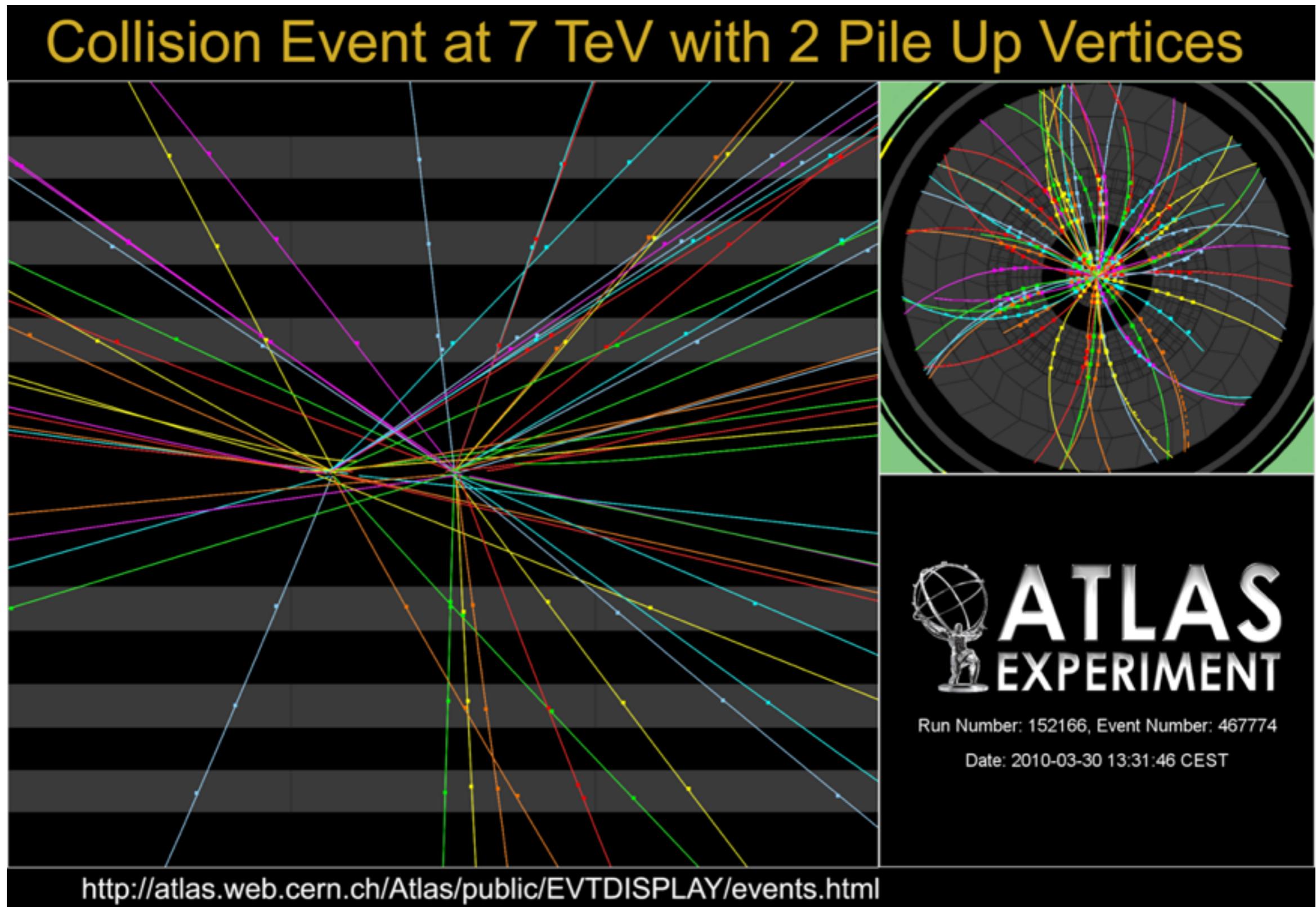
$Z \rightarrow \mu^+\mu^-$ candidate 7 TeV



$Z \rightarrow e^+e^-$ candidate 7 TeV

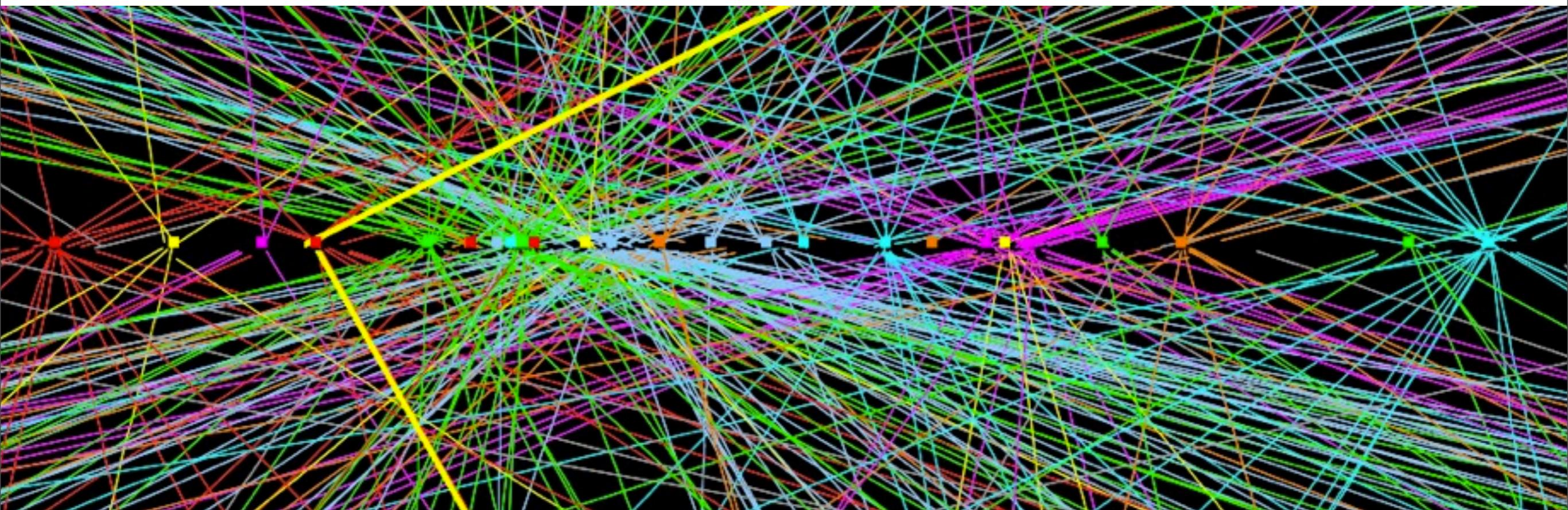


Pile up 7 TeV



Multivertex events have been reconstructed

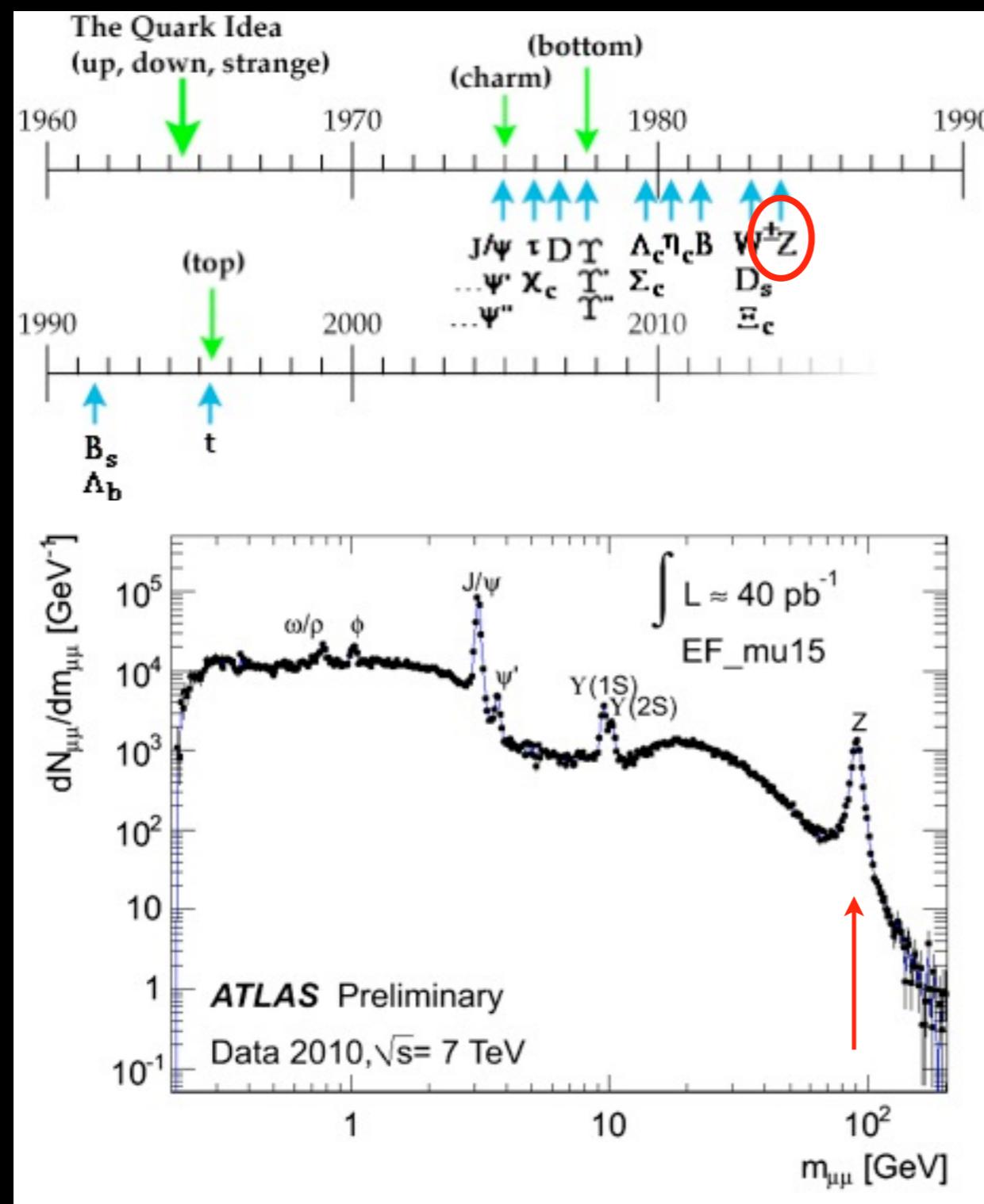
Pile up 2012 - 8 TeV



Lykke til med å finne SUSY!

6. Results from 2010 - mid 2012

REMEASURING THE STANDARD MODEL



Quarks				Leptons				Bosons (Forces)			
2.4 MeV $2/3$ $1/2$ u up	1.27 GeV $2/3$ $1/2$ c charm	171.2 GeV $2/3$ $1/2$ t top	0 0 1 γ photon								
4.8 MeV $-1/3$ $1/2$ d down	104 MeV $-1/3$ $1/2$ s strange	4.2 GeV $-1/3$ $1/2$ b bottom	0 0 1 g gluon								
<2.2 eV 0 $1/2$ νe electron neutrino	<0.17 MeV 0 $1/2$ νμ muon neutrino	<15.5 MeV 0 $1/2$ ντ tau neutrino	91.2 GeV 0 1 Z ⁰ weak force								
0.511 MeV -1 $1/2$ e electron	105.7 MeV -1 $1/2$ μ muon	1.777 GeV -1 $1/2$ τ tau	80.4 GeV ± 1 W [±] weak force								
ALL OK ...											

6. Results from summer 2012



.... Higgs in ATLAS ?

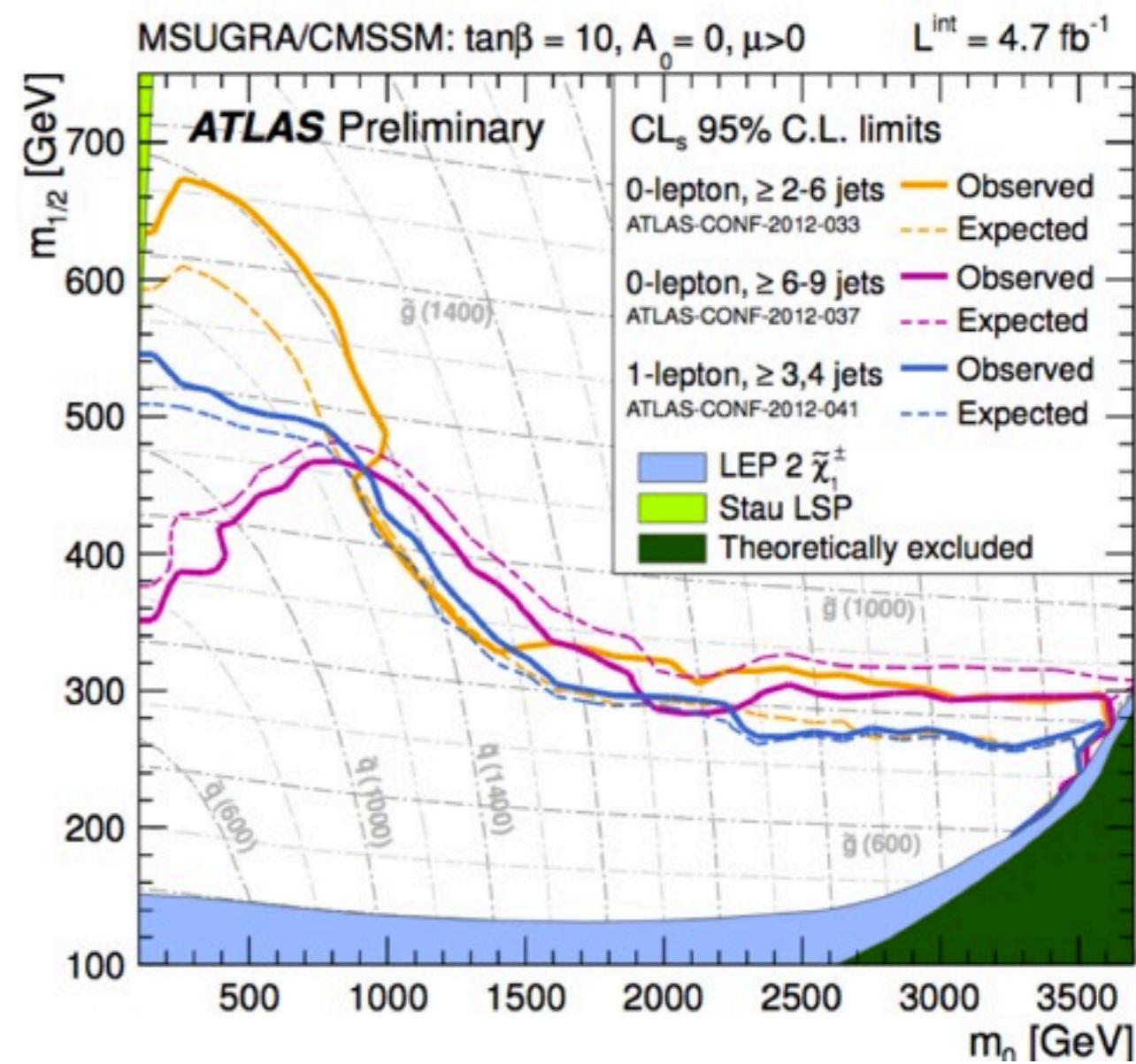
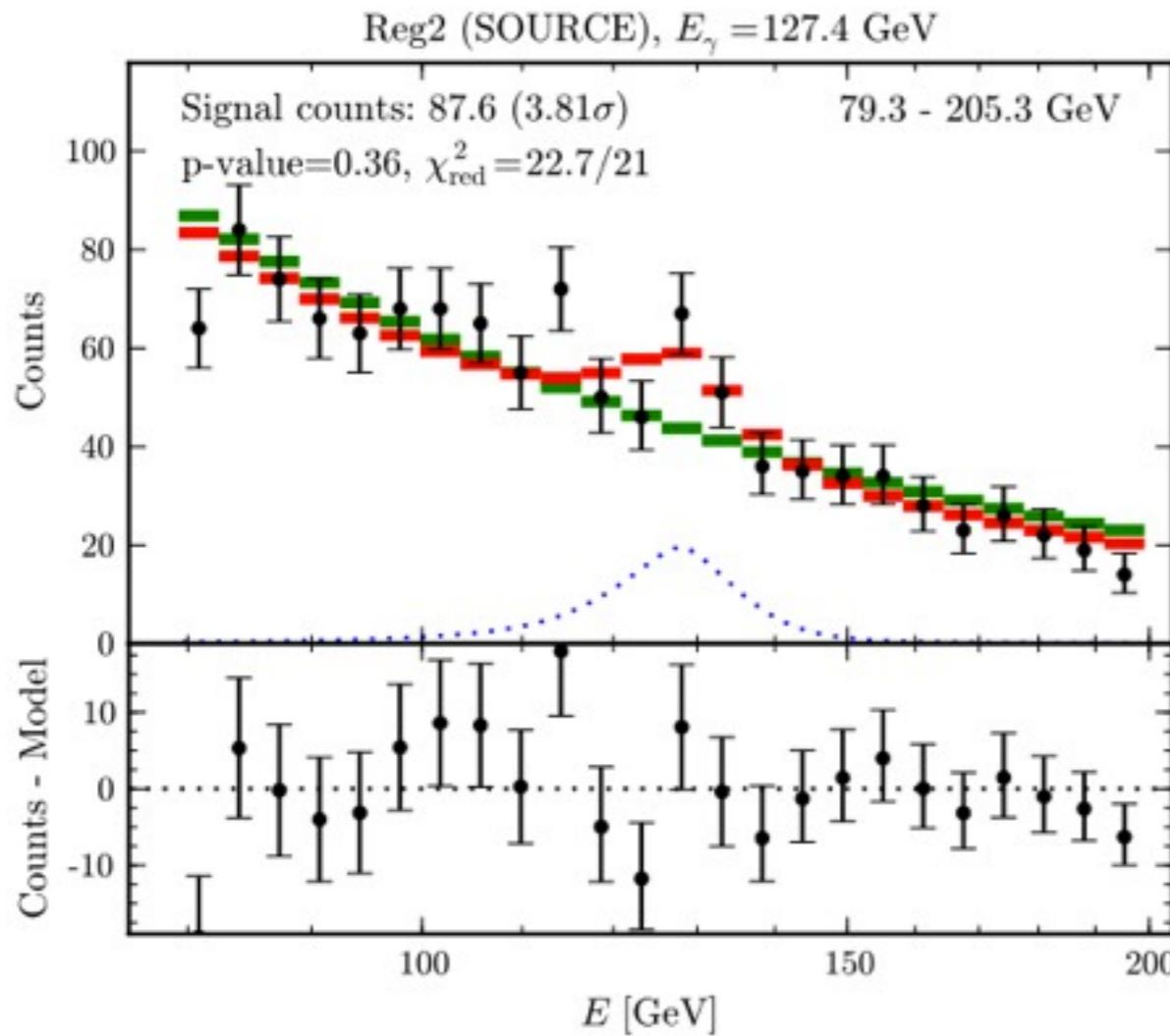
Latest news about Dark Matter

News from LHC

A new boson, probably Higgs is found

Supersymmetry is increasingly excluded in various limits in various models

No sign of Dark Matter signals in model “independent searches” like monojets and monophotons



News from Fermi

Intriguing signal at 130 GeV found by several teams, Fermi will recalibrate and perform a new search for this gamma emission line



❖ Noen store oppdagelser og oppfinnelser (CERN)

- ✓ ISR produserte i **1971** verdens første proton - proton kollisjoner !
- ✓ Oppdagelsen av Neutral Currents **1973** ved Gargamelle boble kammer
- ✓ I **1979** bestemte CERN å gjøre om SPS til verdens første proton-antiproton collider
- ✓ Oppdagelsen av W og Z bosonene i **1983** ved UA1 og UA2 eksperimentene. Dette resultatet bekreftet foreningen av svake og elektro-magnetiske krefter som beskrevet i standard modellen.
 - *Nobelprisen i fysikk i 1984 ble gitt til Carlo Rubbia og Simon van der Meer for oppdagelsen av W og Z*
 - *Nobelprisen i fysikk i 1992 ble gitt til Georges Charpak for utviklingen av MultiWire Proportional Chamber*
- ✓ Oppfinnelsen av World Wide Web av Tom Berners-Lee. Allerede i 1991 var et tidlig www-system tilgjengelig på CERN.



❖ Noen historiske Nordmenn på CERN

- ✓ **Normenn har vært involvert på CERN siden begynnelsen og er det fremdeles....**
- ✓ **Odd Dahl** ble med i en nimanns ekspertgruppe som ble dannet i 1951. Han ble snart leder for byggingen av CERNs første akselerator (en synkrosyklotron – SC) som sto ferdig i 1957.
 - Odd Dahl ble tilbudt stillingen som CERNs første generalsekretær men takket nei
- ✓ **Rolf Widerøe**, en annen norsk fysiker, var også veldig aktiv fra begynnelsen. Han konstruerte blant annet den første **linear akseleratoren** for positive ioner. Han regnes også som en av oppfinnerene av **betatronen**
 - Widerøe overbeviste Dahl i 1952 om nødvendigheten av en protonsykroton (PS) noe Widerøe hadde tatt patent på allerede i 1949, og Dahl var veldig aktiv med å fremme byggingen av en slik akselerator. Stod ferdig i 1959 som verdenes største akselerator
- ✓ **Kjell Johnsen**, som var sentral i byggingen av PS, ledet byggingen av ISR (Intersecting storage rings) for å bedre partiklenes kinetiske energi



Hvordan delta i CERN's aktiviteter

❖ Mange relevante studier

- ✓ Partikkel fysikk, material fysikk, detektor teknologi, elektronikk, mekanikk, informatikk, instrumentering
- ✓ Universiter (UiB, UiO...), Siv-Ing. Studier (NTNU..), Ing. Studier, Høyskoler etc...
- ✓ Utenlands studier (INSA, ...)



Summer Students enjoying a lecture on particle physics by Ronald Kleis

H. Sandaker



CERN summer school

❖ Sommer skole

- ✓ Allerede under studiet
er det mulig å delta på
CERN !
- ✓ 2-3 måneder om
sommeren
- ✓ Undervisning i
akselerator fysikk,
detektorer og
partikkel fysikk
- ✓ Prosjekt oppgave
- ✓ Møt unge forskere fra
hele verden !



Wolfgang Pauli at the Summer School
for theoretical physics Les Houches

H. Sandaker



CERN Technical Student

❖ Teknisk Student

- ✓ Diplomarbeid eller Master
- ✓ ~ 6-12 måneder
- ✓ Prosjektarbeid
- ✓ Mulighet til å følge seminarer ved CERN
- ✓ Internasjonalt samarbeid !



The Technical Student Programme draws Norwegians

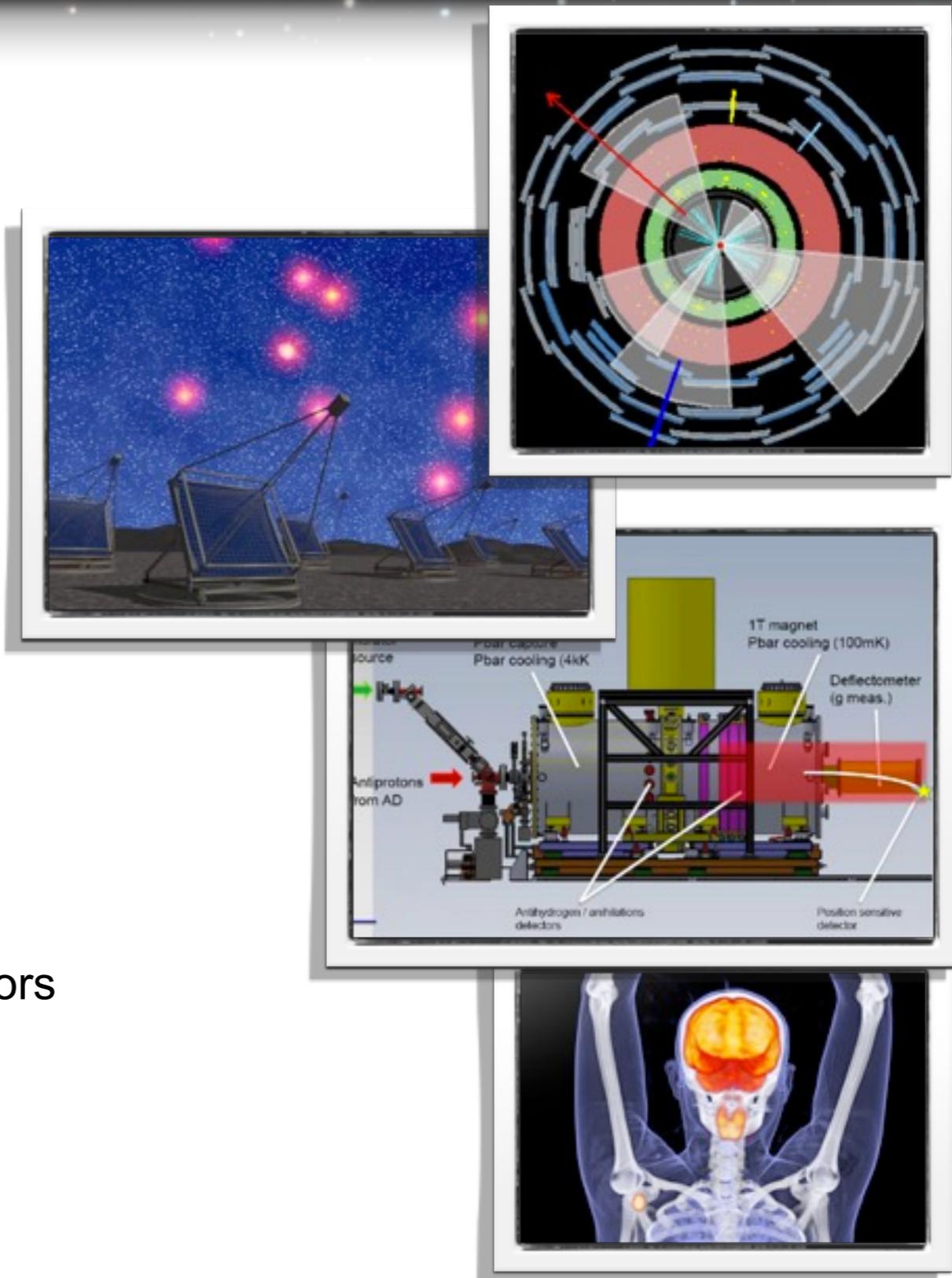
Erik Hejne, second from left, Chairman of the Technical Students Committee, and Jens Vigen, who is concerned specifically with Norwegian students at CERN, with some of the Norwegian technical students who arrived at CERN in spring 2005, together with their teachers.

H. Sandaker

Typical Master project work

Example 1:

LHC study of Dark Matter candidates (ATLAS)
(1 PhD + 1/2 postdoctor)



Example 2:

Astroparticle study of Dark Matter candidates (CTA)
(1 PhD + 1/2 postdoctor)

Example 3:

Particle detector upgrade for dark matter detection
(ATLAS upgrade, AEgIS and CTA)
(1 PhD + 1 postdoctor)

Example 4:

New applications for particle and astroparticle detectors
(Masters)



- ❖ **TAKK FOR BESØKET !**

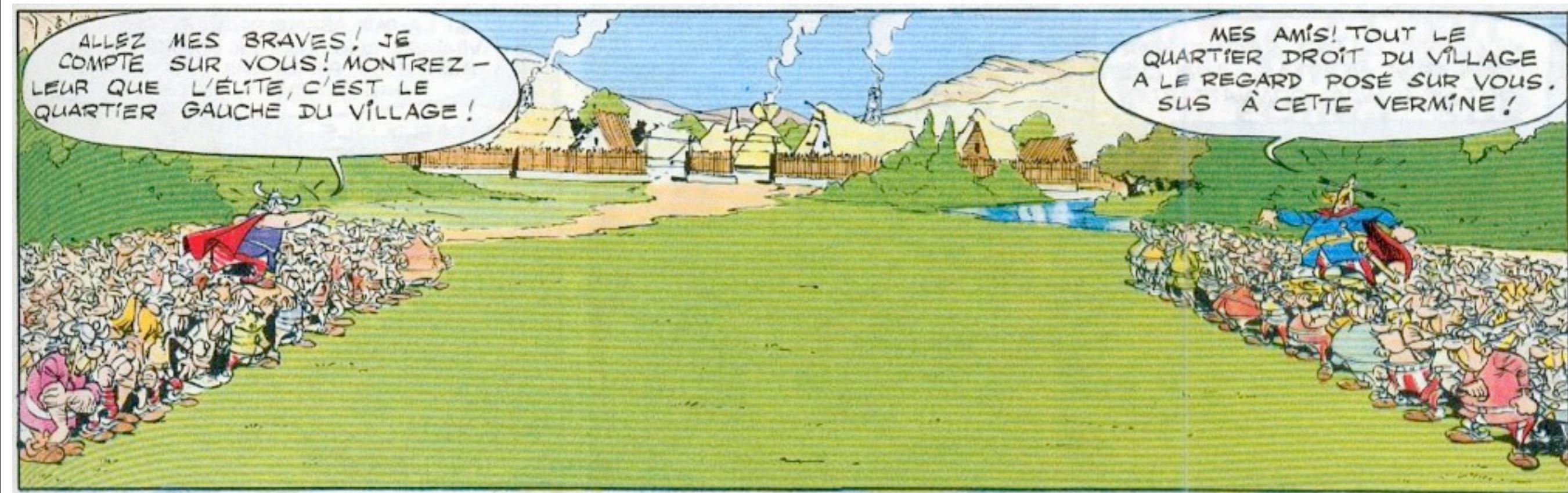
- ❖ **VELKOMMEN TILBAKE !**

H. Sandaker



Nær framtid !!

En LHC kollisjon?



H. Sandaker



Nær framtid !!

En LHC kollisjon?

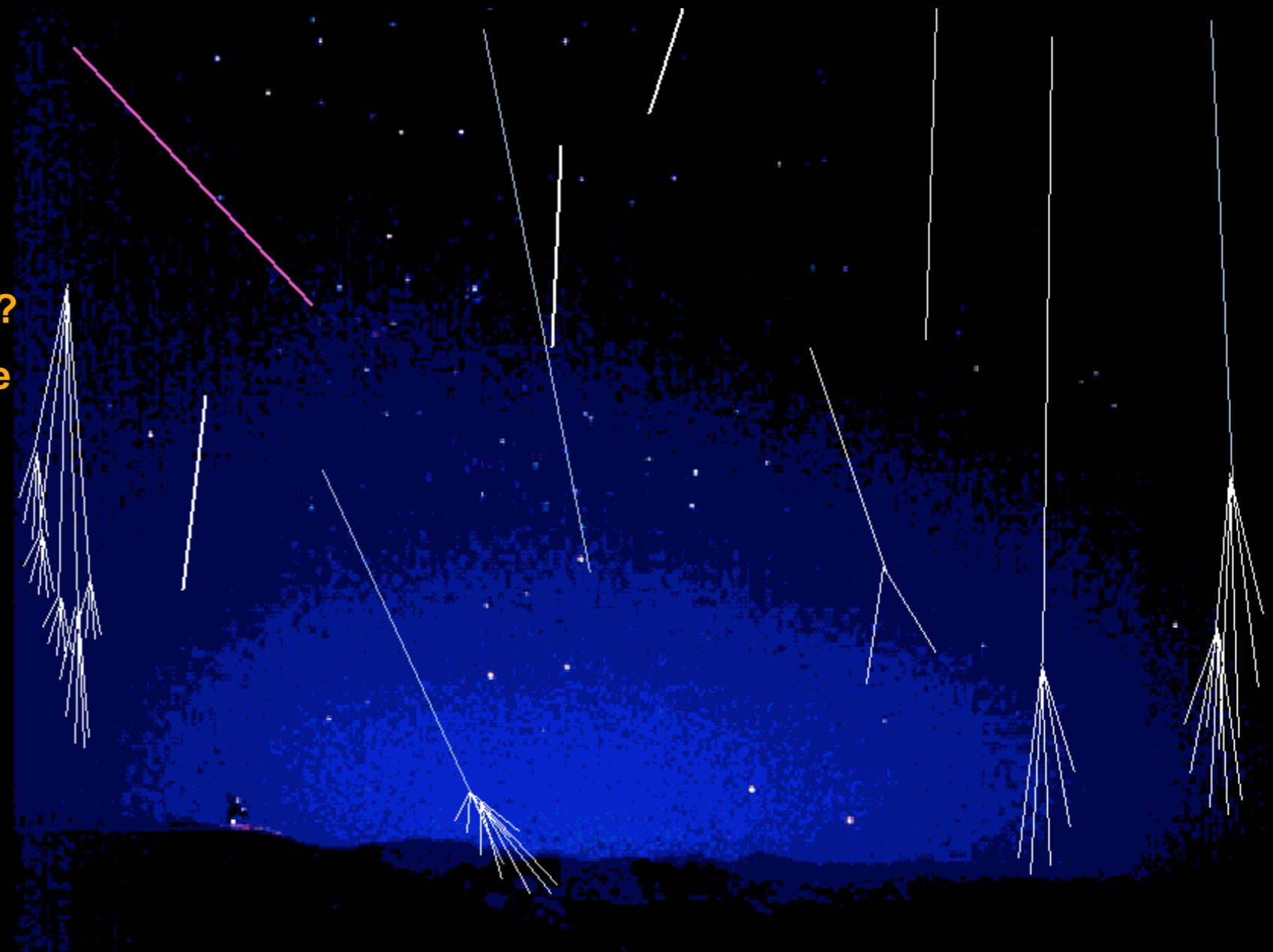


H. Sandaker

Takk for oppmerksomheten

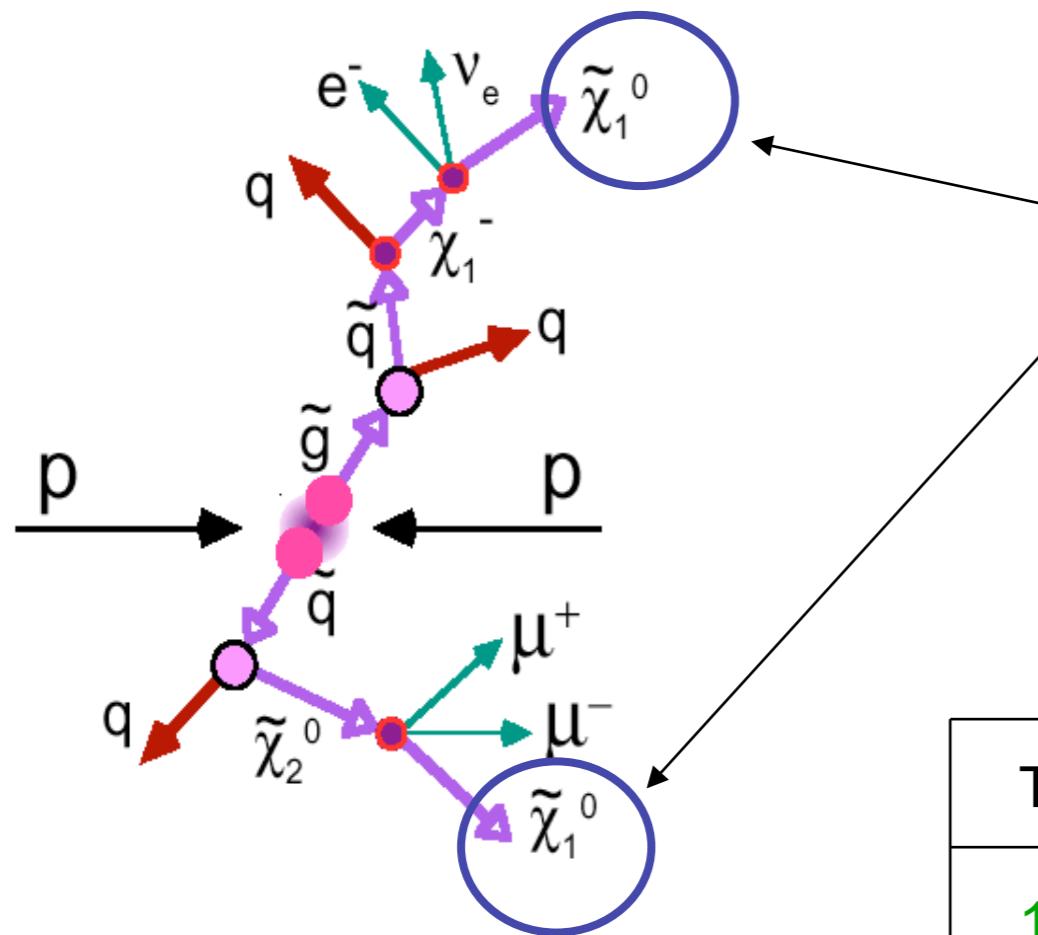
Ekstra

1. Hva skal ATLAS forske på?
2. Hvordan skjer kollisjonene
3. Om ATLAS detektoren
4. Data fra ATLAS 2010
5. Planer for 2011



Partikkel kollisjoner i atmosfæren

Supersymmetry & Dark Matter



This particle (neutralino) is a good candidate for the universe dark matter

ATLAS discovery reach

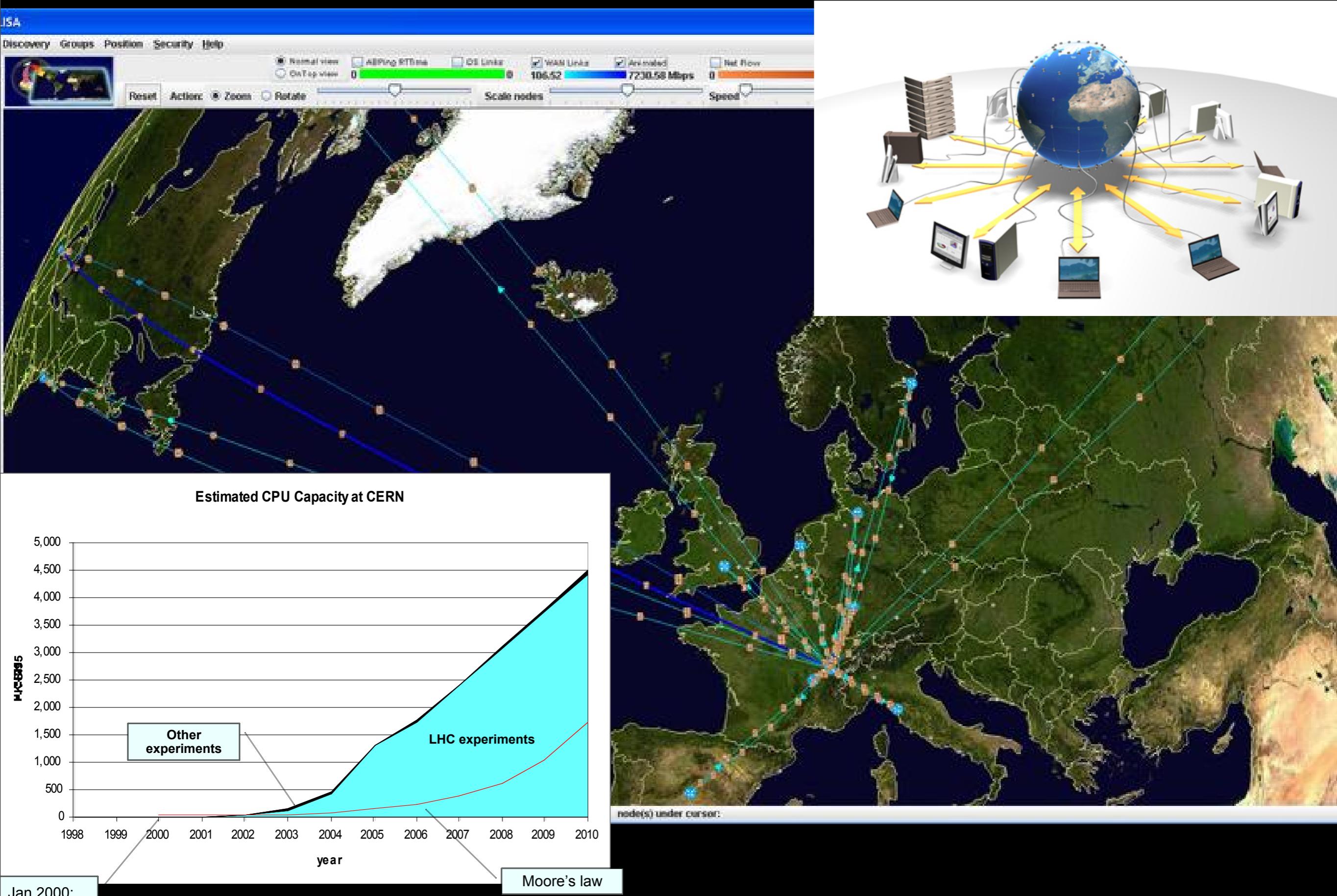
Time	reach in squark/gluino mass
1 month	~ 1.3 TeV
1 year	~ 1.8 TeV
3 years	~ 2.5 TeV
ultimate	up to ~ 3 TeV

New particle discoveries depends
on mass of the particles !
.... may take some time ...

Neutralino mass can be measured to 10% -> SUSY discovery and neutralino mass measurement at LHC can solve problem of universe cold dark matter

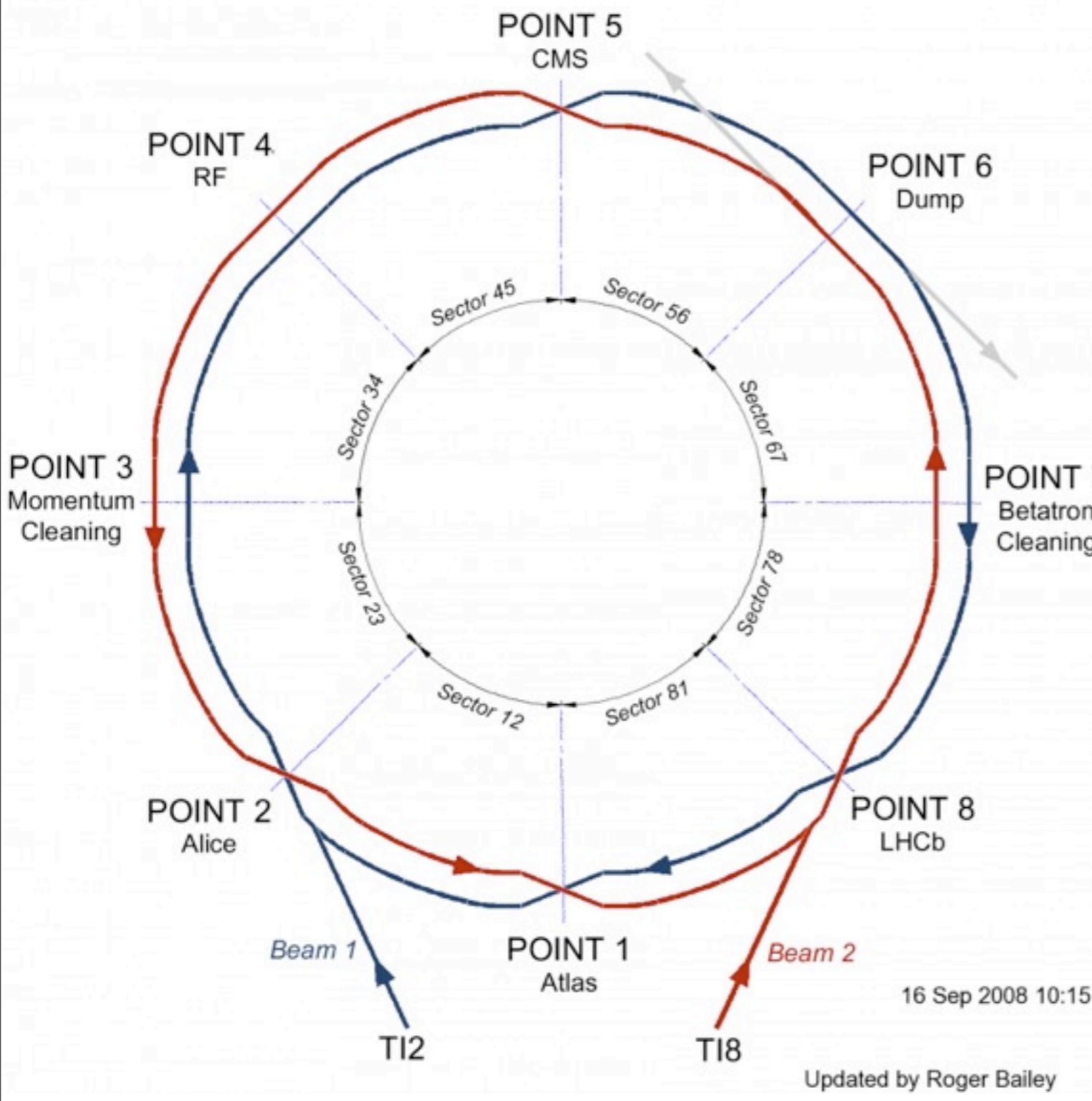
H. Pernegger

5. Grid - World wide analysis



6. First results from LHC

10. September - historical LHC start !

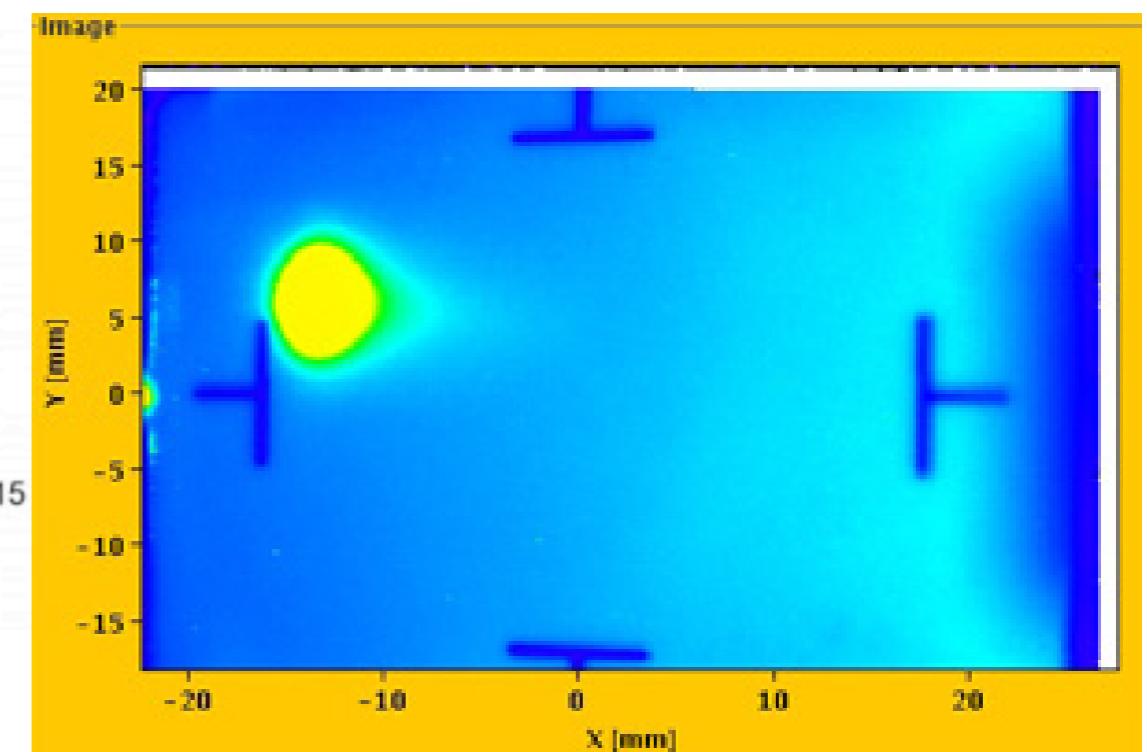


16 Sep 2008 10:15

Updated by Roger Bailey

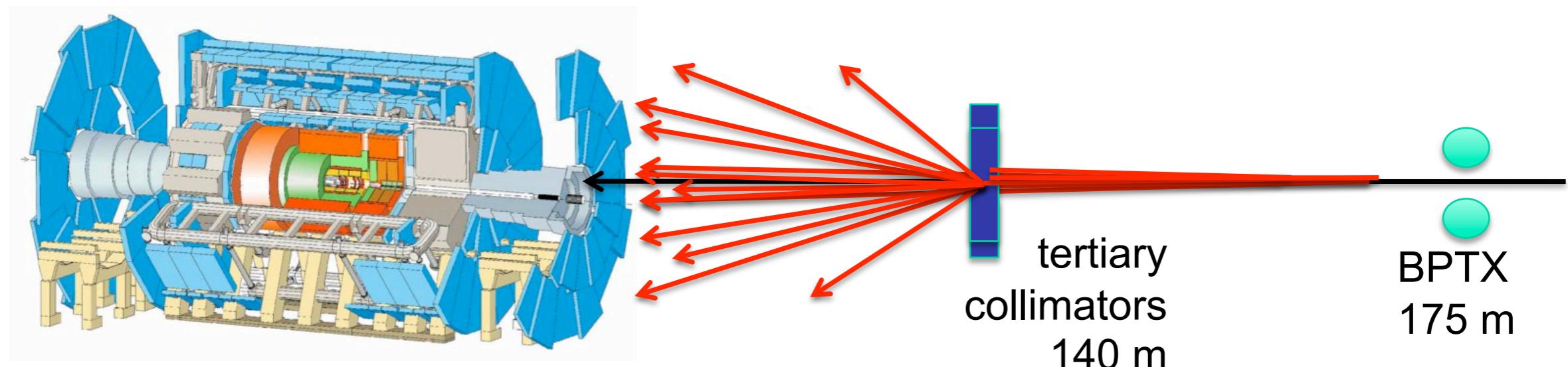
“Geneva, 10 September 2008. The first beam in the Large Hadron Collider at CERN was successfully steered around the full 27 kilometres of the world’s most powerful particle accelerator at 10h28 this morning”

“On Thursday night, 11 September, beam two, the anti-clockwise beam, was captured and circulated for over half an hour before being safely extracted from the LHC... The LHC is on course for first collisions in a matter of weeks”



6. First results from LHC

What particles did we see in ATLAS ?



ATLAS was ready for first beam:

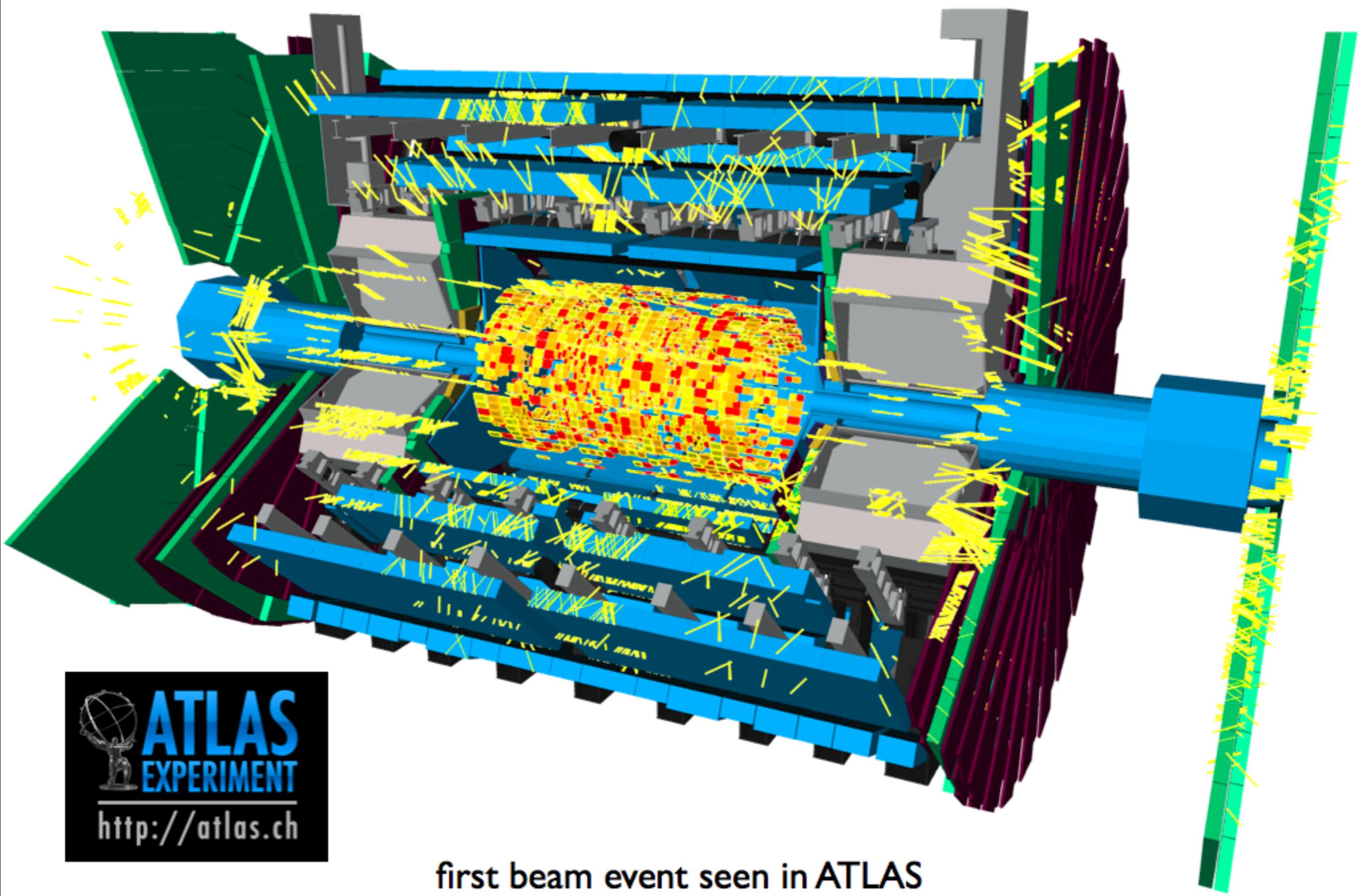
- Muon system (MDT, RPC, TGC) on at reduced HV
- LAr (-FCAL HV), Tile on
- TRT on, SCT reduced HV, Pixel off
- BCM, LUCID, MinBias Scint. (MBTS), Beam pickups (BPTX)
- L1 trigger processor, DAQ up and running, HLT available (but used for streaming only)

Two LHC start-up scenarios:

1. Open all collimators, go around as far as beam goes, correct as needed
 - Little activity expected except for accidents
2. Go step-by-step, stopping beam on collimators, re-align with centre, open collimator, keep going
 - Splash event from collimators for each beam shot

B. Stugu

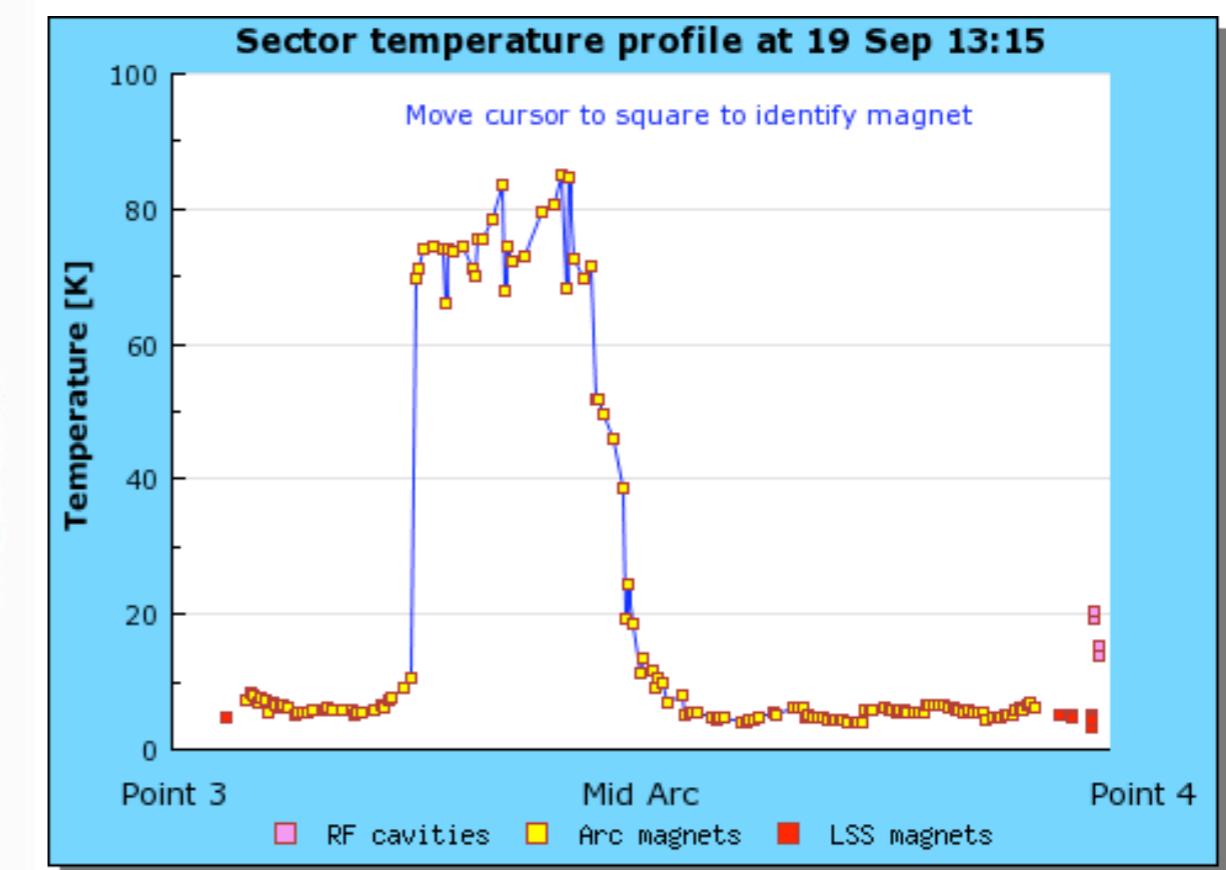
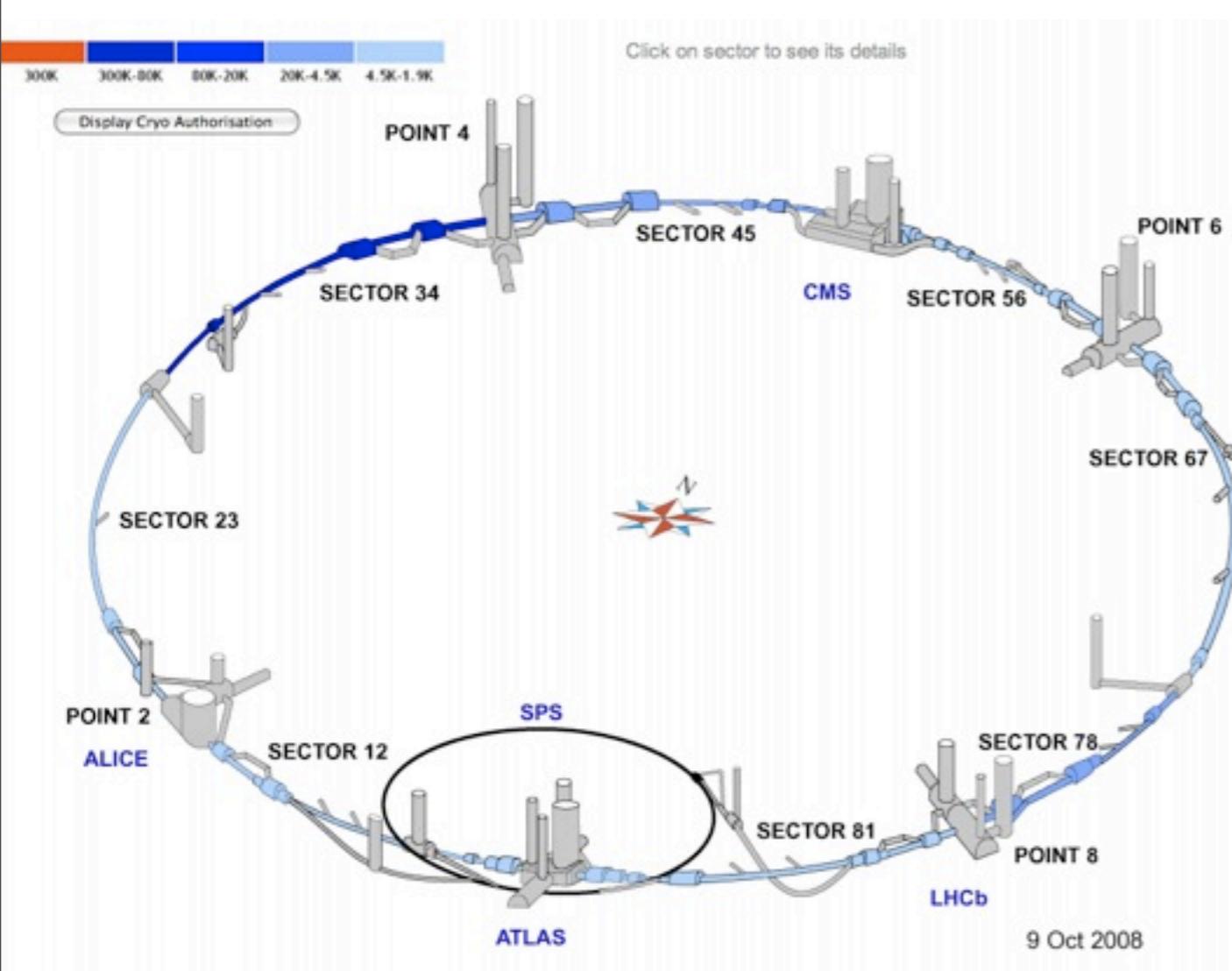
6. First results from LHC



6. First results from LHC

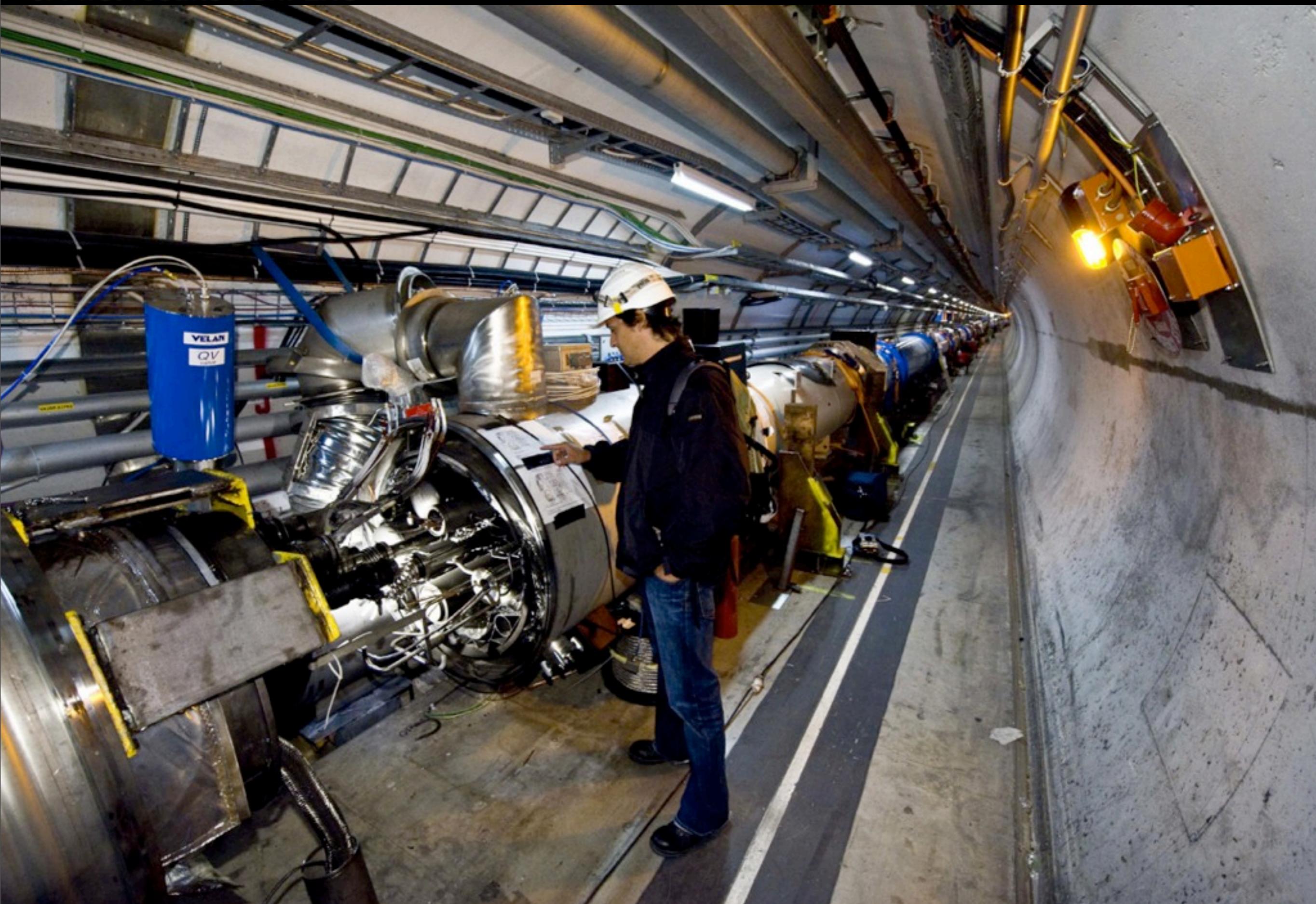
19. September - historical LHC stop !

"During commissioning (without beam) of the final LHC sector (sector 3-4) at high current for operation at 5 TeV, an incident occurred at mid-day on Friday 19 September resulting in a large helium leak into the tunnel. ... the initial malfunction was caused by a faulty electrical connection between two magnets, which resulted in mechanical damage and release of helium from the cold mass into the tunnel"



But the excitement is not over ... new startup in 2009 !

6. First results from LHC



6. First results from LHC

Plan for 2009

Reparasjon og installering

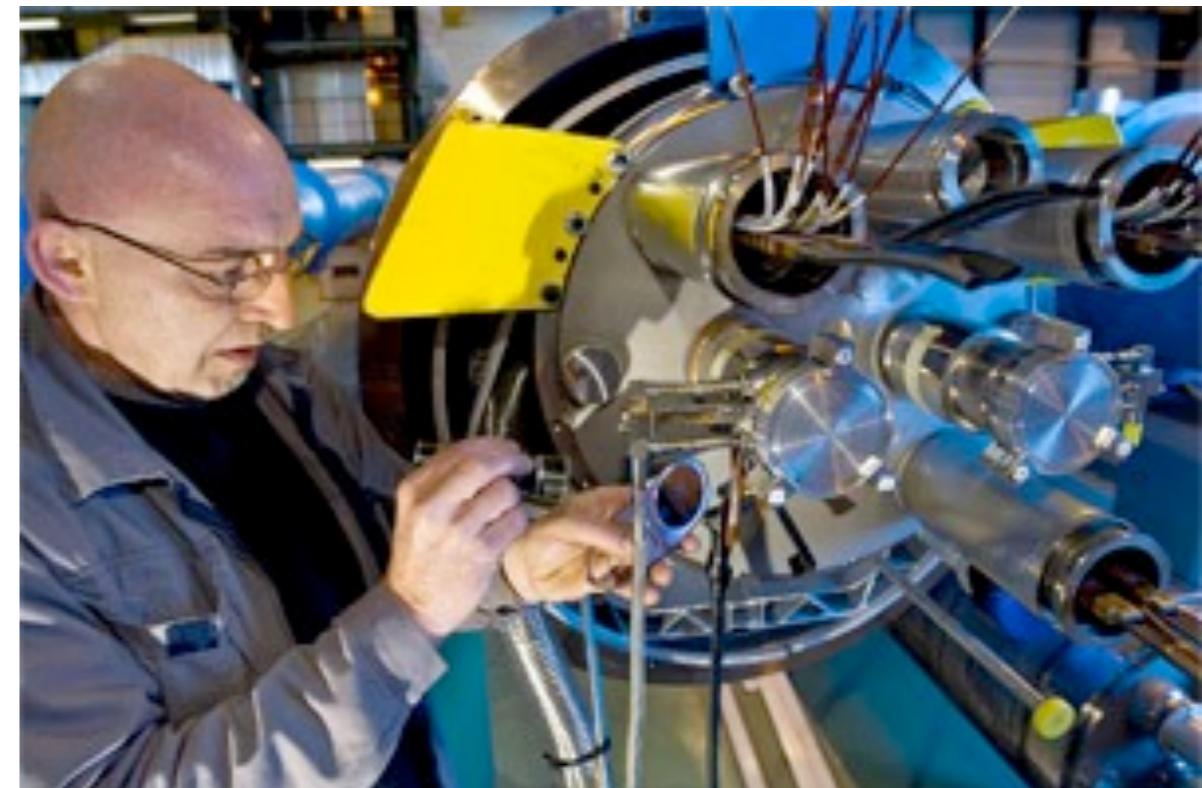
- 53 magneter måtte ut av tunnellen for rensing eller reparasjon
- Planen er at den siste magneten er tilbake i tunnellen i slutten av mars

Testing

- Nedkjøling av magnetene tar flere måneder
- Planen er at det er klart for strømtester ved slutten av juni

Kollisjoner

- I løpet av høsten
- En konferanse i Chamonix akkurat nå hvor dette blir diskutert samt en plan for hvilke kollisjoner man skal ha til hvilken energi



Final preparations on a replacement magnet ready to be lowered into sector 3-4

"The top priority for CERN today is to provide collision data for the experiments as soon as reasonably possible"

Mini svarte hull på CERN

Universet har laget høy energetiske proton-proton kollisjoner i løpet av milliarder av år. Vi er fremdeles her ...

Universet

- Det skjer det mer en 10 millioner millioner LHC-lignende eksperiment hvert sekund !

Cosmisk stråling

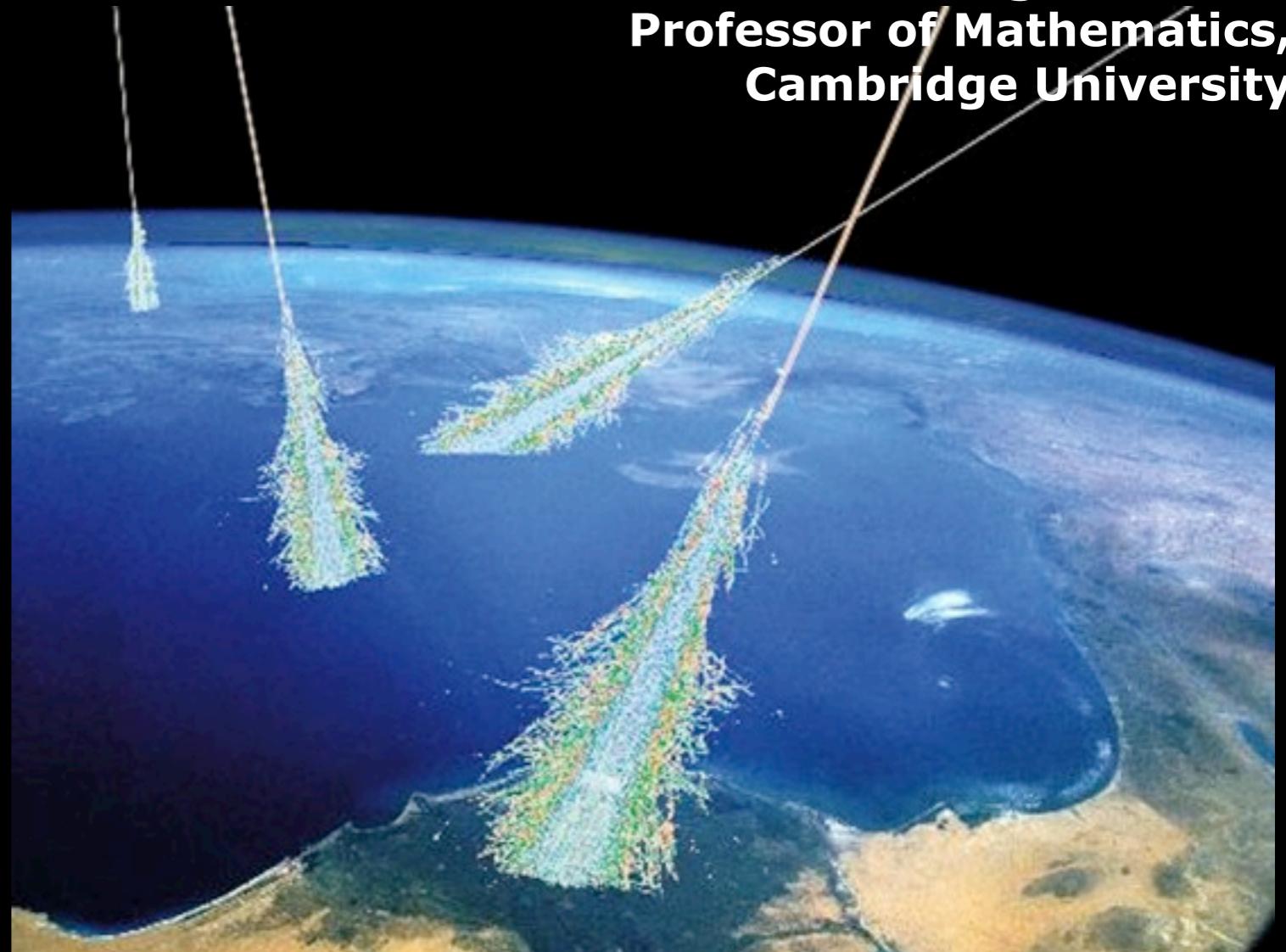
- Hele tiden skjer det at kosmiske stråler kolliderer med vår atmosfære, dette skjer ofte med mye høyere energi enn det vi kan lage ved LHC

Svarte hull

- En mulig teoretisk modell !
- Astronomiske svarte hull kan ikke lages på CERN, det som kanskje! kan lages er mikroskopiske hull når to protoner kolliderer
- Disse er ikke levedyktige og dør med en gang

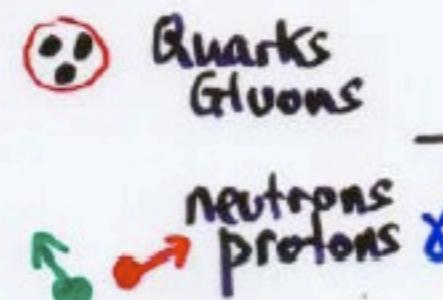
"The world will not come to an end when the LHC turns on. The LHC is absolutely safe. ... Collisions releasing greater energy occur millions of times a day in the earth's atmosphere and nothing terrible happens."

Prof. Steven Hawking, Lucasian Professor of Mathematics, Cambridge University



...and patterns (that change)

QG Plasma



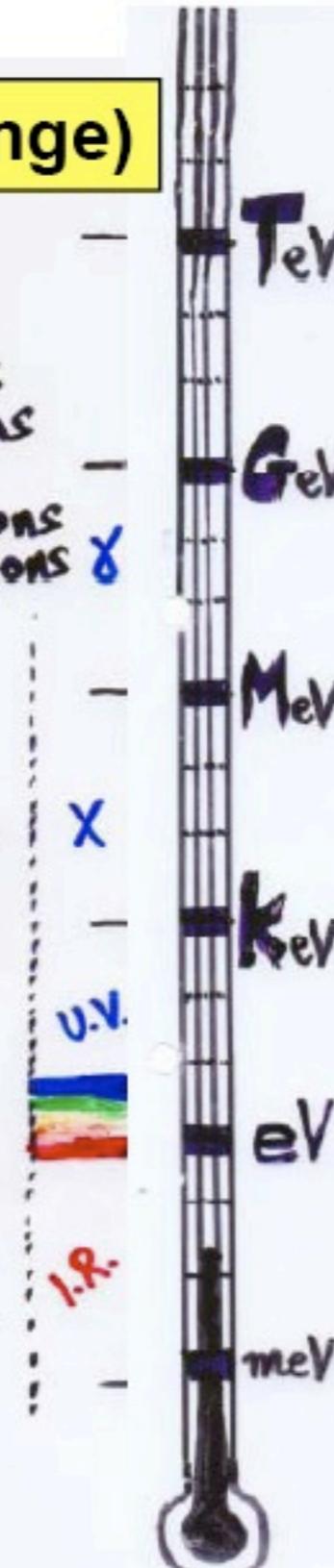
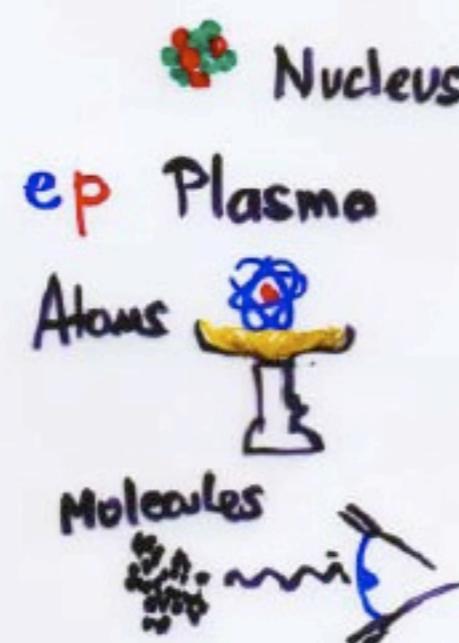
Nuclei melt

↓ exist

H melt: plasma
↓ exist

Ice melt
↓ exist

F. Close

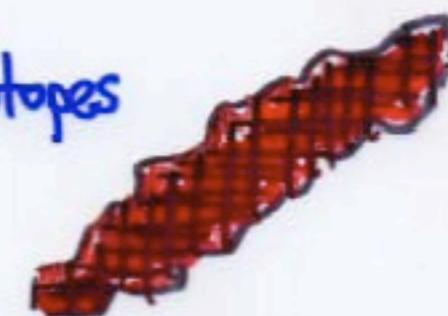


No mass. Unified Theory

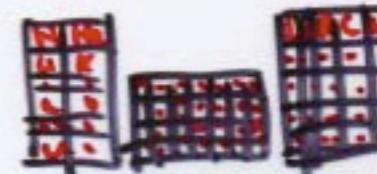
Standard Model MASS

t	b	T	v	w
c	s	p	v	z
u	d	e	v	γg

Nuclear Isotopes



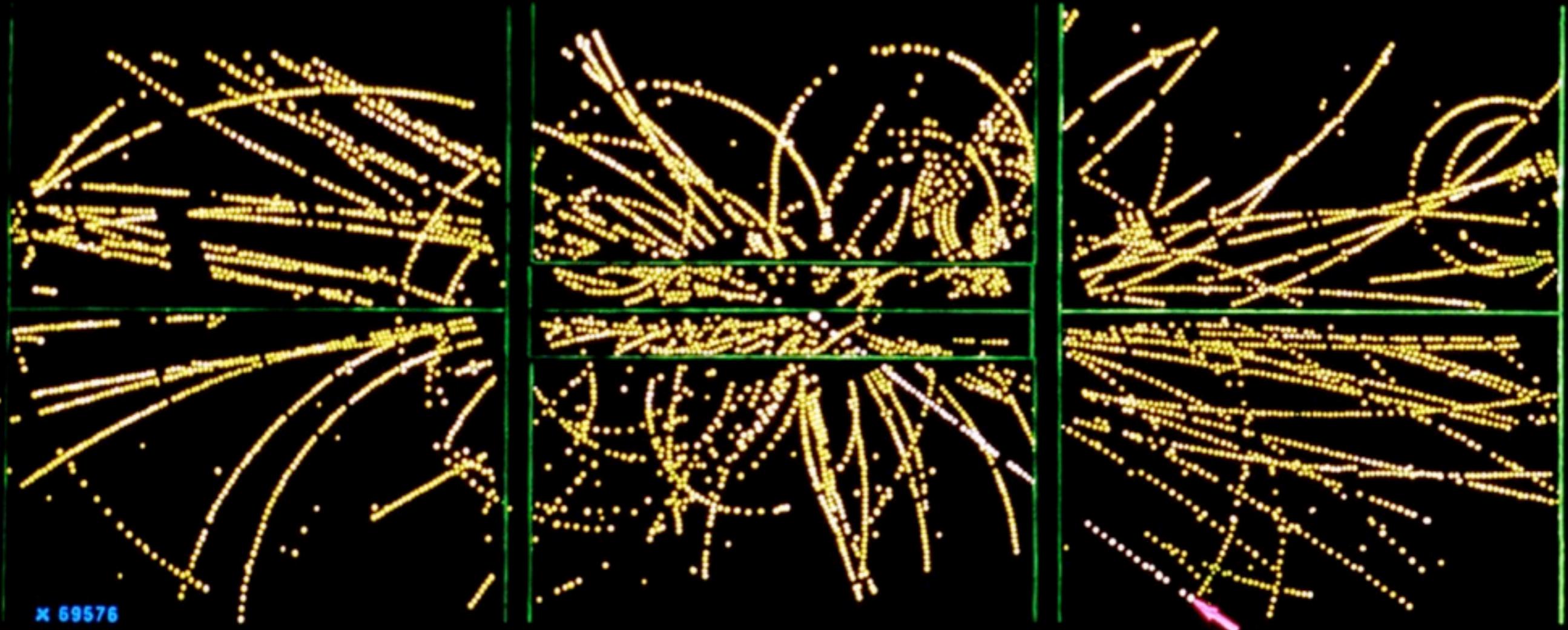
Mendeleev



Snowflake pattern

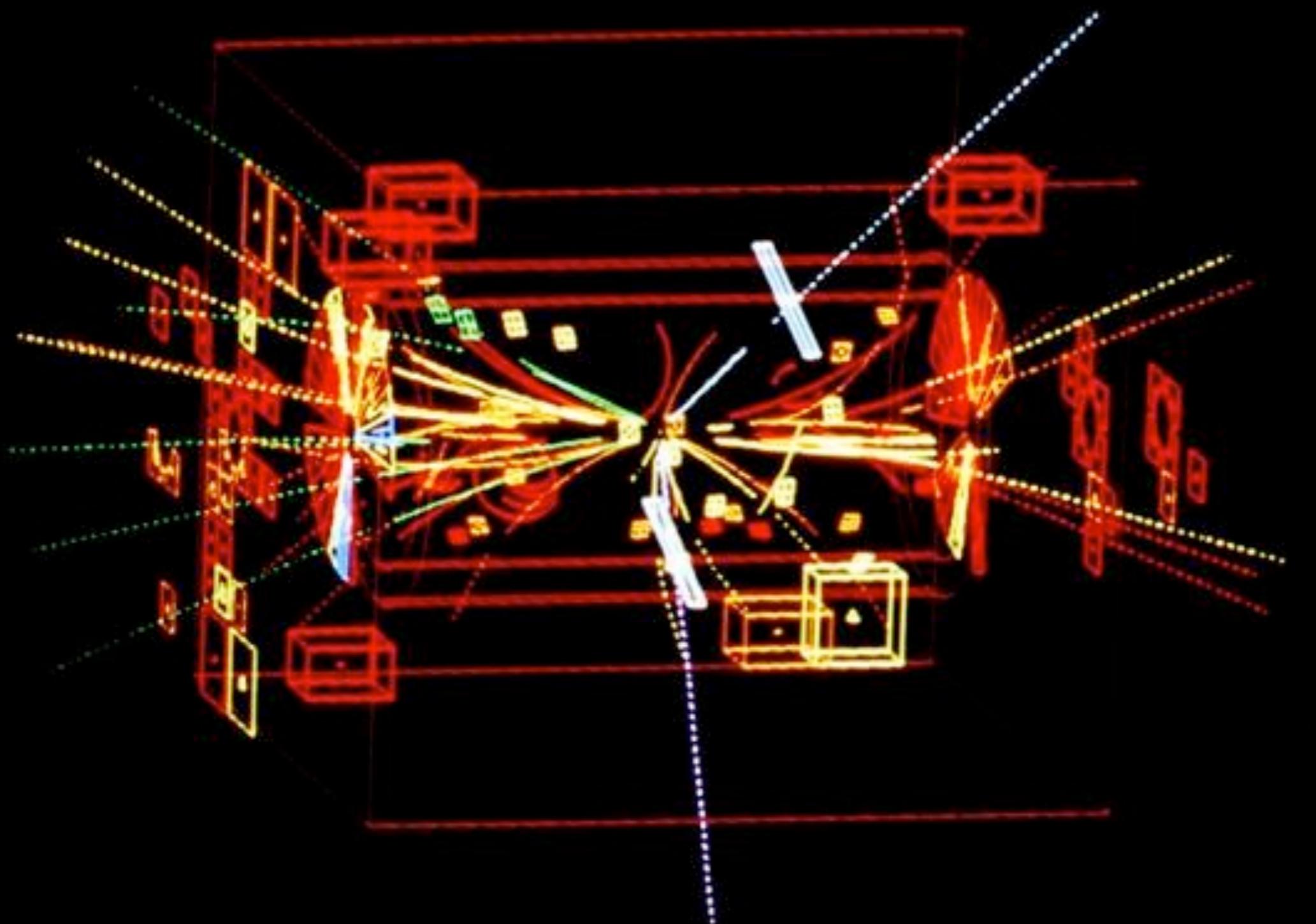


Oppdagelsen av W bosonet ved UA1 (SppS)



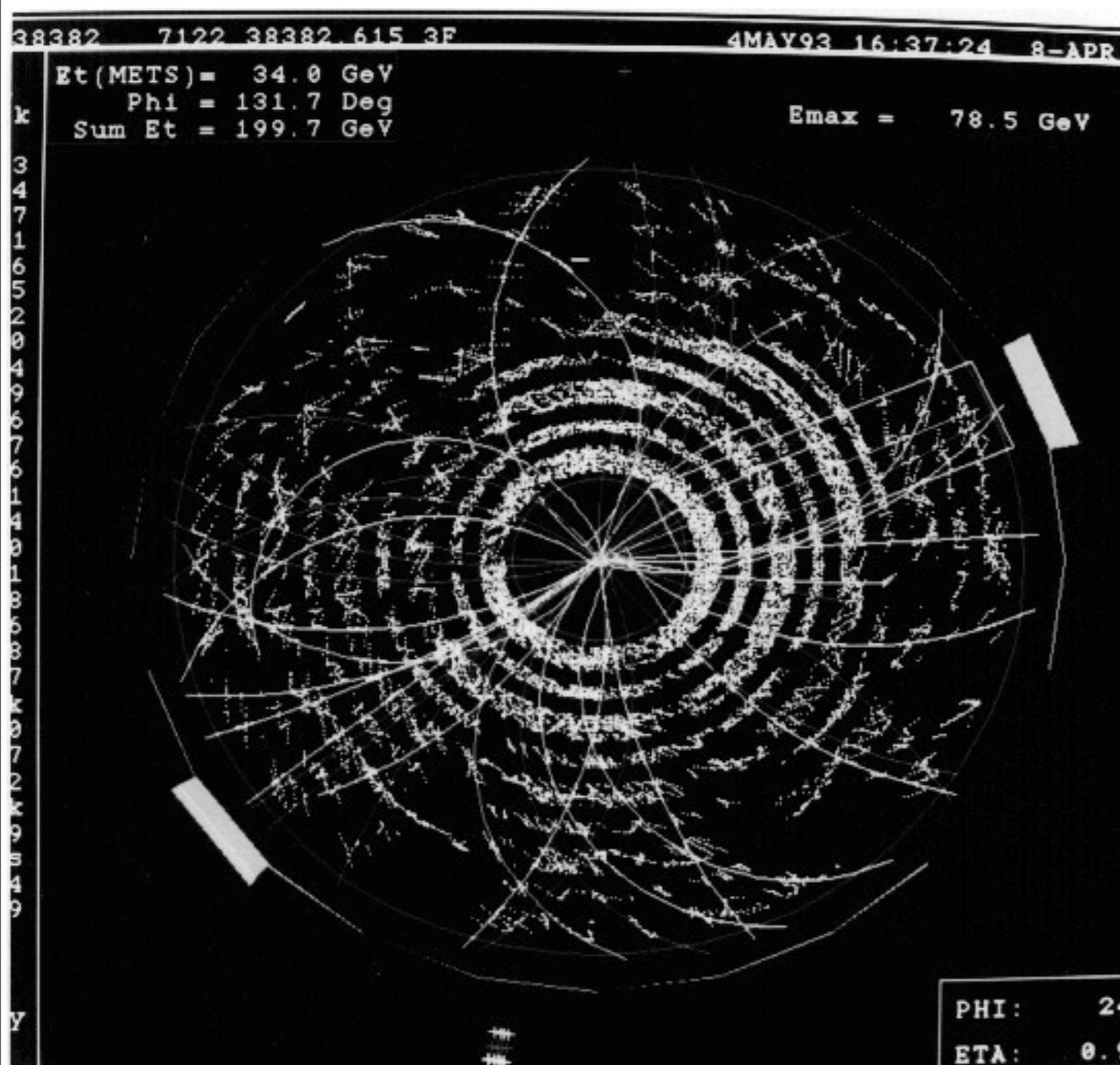
- Discovery of the W particle in 1982, producing a high transverse energy electron back to back with missing energy (neutrino)

Oppdagelsen av Z bosonet ved UA1 (SppS)



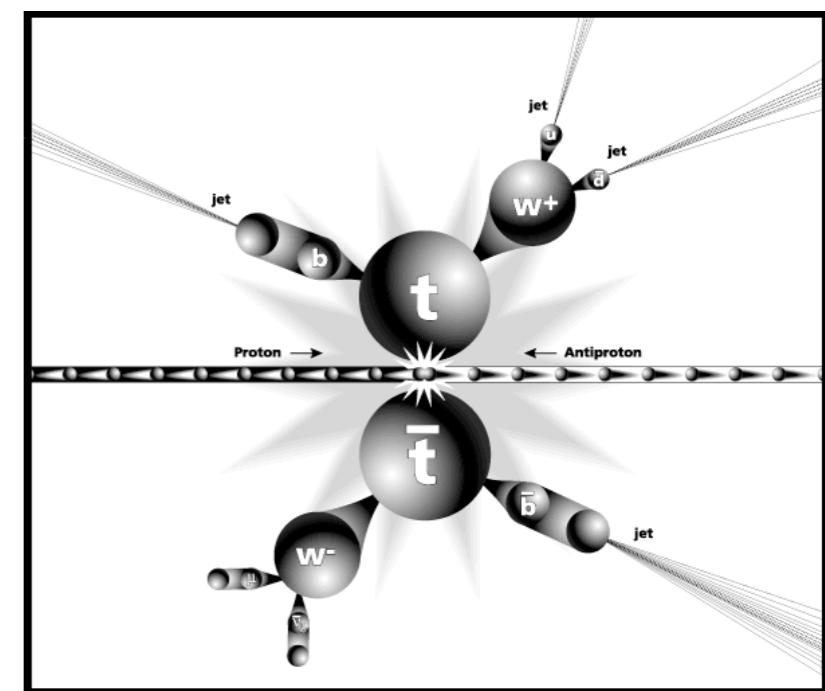
→ Discovery of the Z particle in 1983. In this the neutral Z decays into a high-energy electron and a positron (antielectron) carrying the mass that once was the Z particle, converted according to $E = mc^2$.

4. Standard Modellen - En familie til !



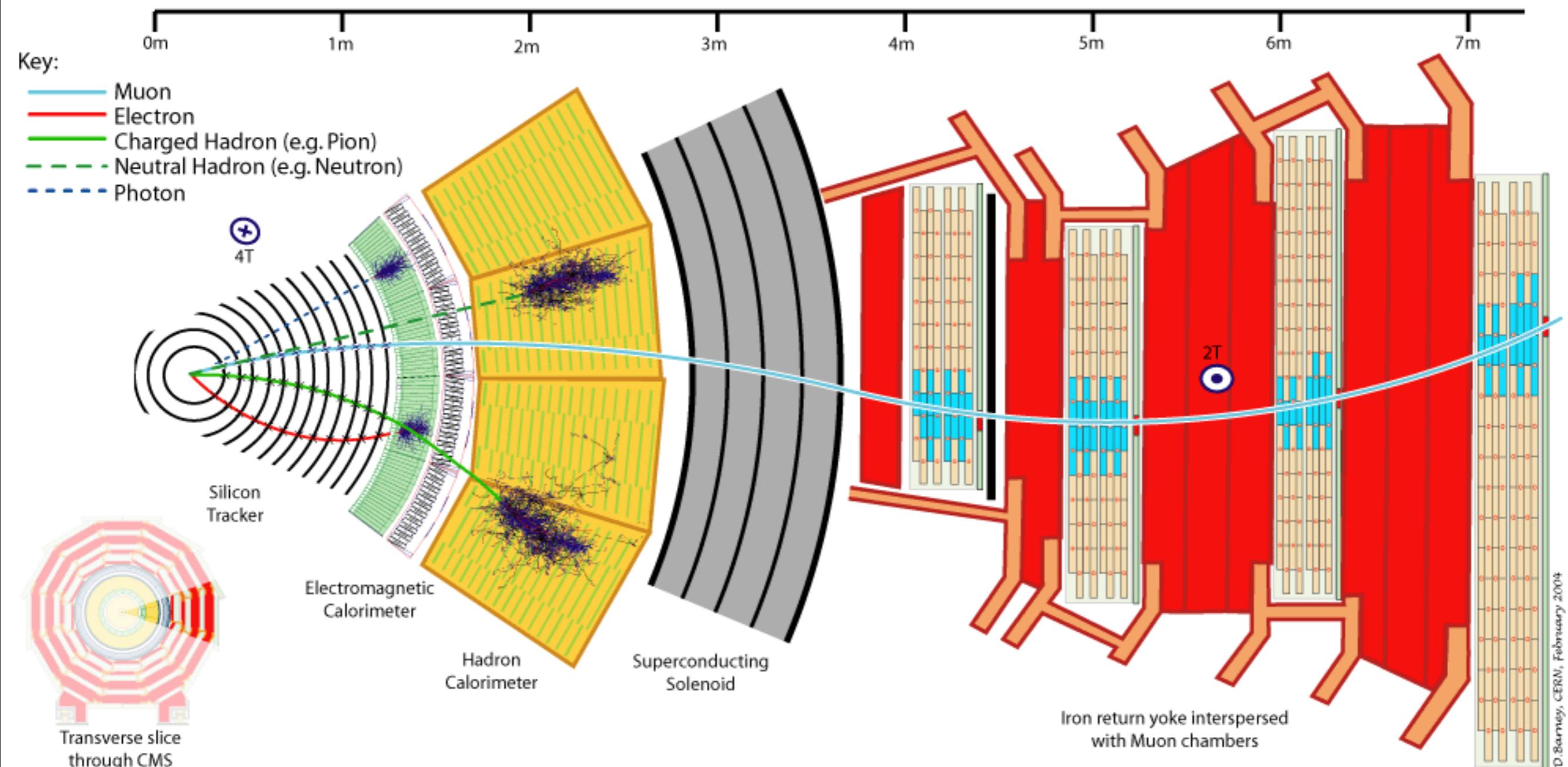
→ Top kvarken ble oppdaget i 1995 ved Fermilab, Chicago

→ Masse 172 GeV (proton 0.94 GeV)



Physicists recognize particles produced in collisions by their electronic signatures, shown graphically by computers. The circle shows a computer-generated view of a potential top quark signature, with particle tracks emerging from the center of a collision.

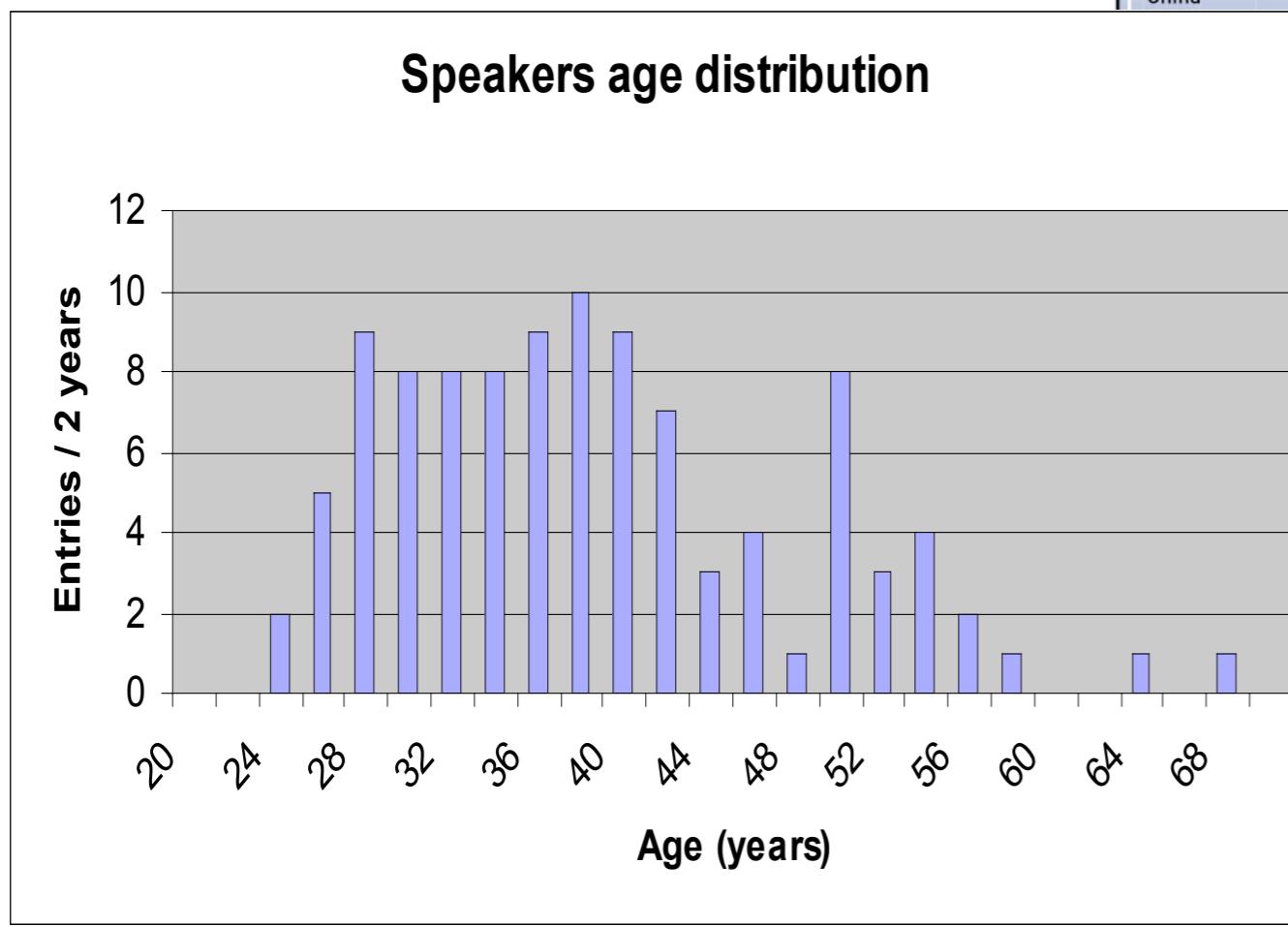
Partiklenes bane gjennom en detektor



> 100 Million Electronics Channels, 40 MHz ---> TRIGGER

The Atlas collaboration

A truly worldwide collaboration:
35 Countries
158 Institutions
1770 Scientific Authors



e.g 4th ATLAS Physics Workshop
Athens, May 2003

Speakers age distribution
of 103 (of the 104) talks

28 female and 76 male
speakers

/ CERN

35

Black holes in LHC?

- Black holes could, in principle, be arbitrarily small. However, according to standard General Relativity, there is no chance to produce black holes at the LHC, since conventional gravitational forces between fundamental particles are too weak.
- There is no established quantum theory for gravitation (certainly needed for small ones).
- Some quantum gravity proposals (involving more than 3 spacial dimensions) make speculative predictions on production of black holes in proton.proton collisions at LHC
- *But in these models they are always unstable, both because of Hawking radiation, and because they always can decay back into the particles that produced them.*
- (I'm of course brainwashed by the Cern/LHC Safety Assessment Group, Ellis,Guidice,Mangano,Tkachev,Wiedemann, CERN-PH-TH/2008-136)