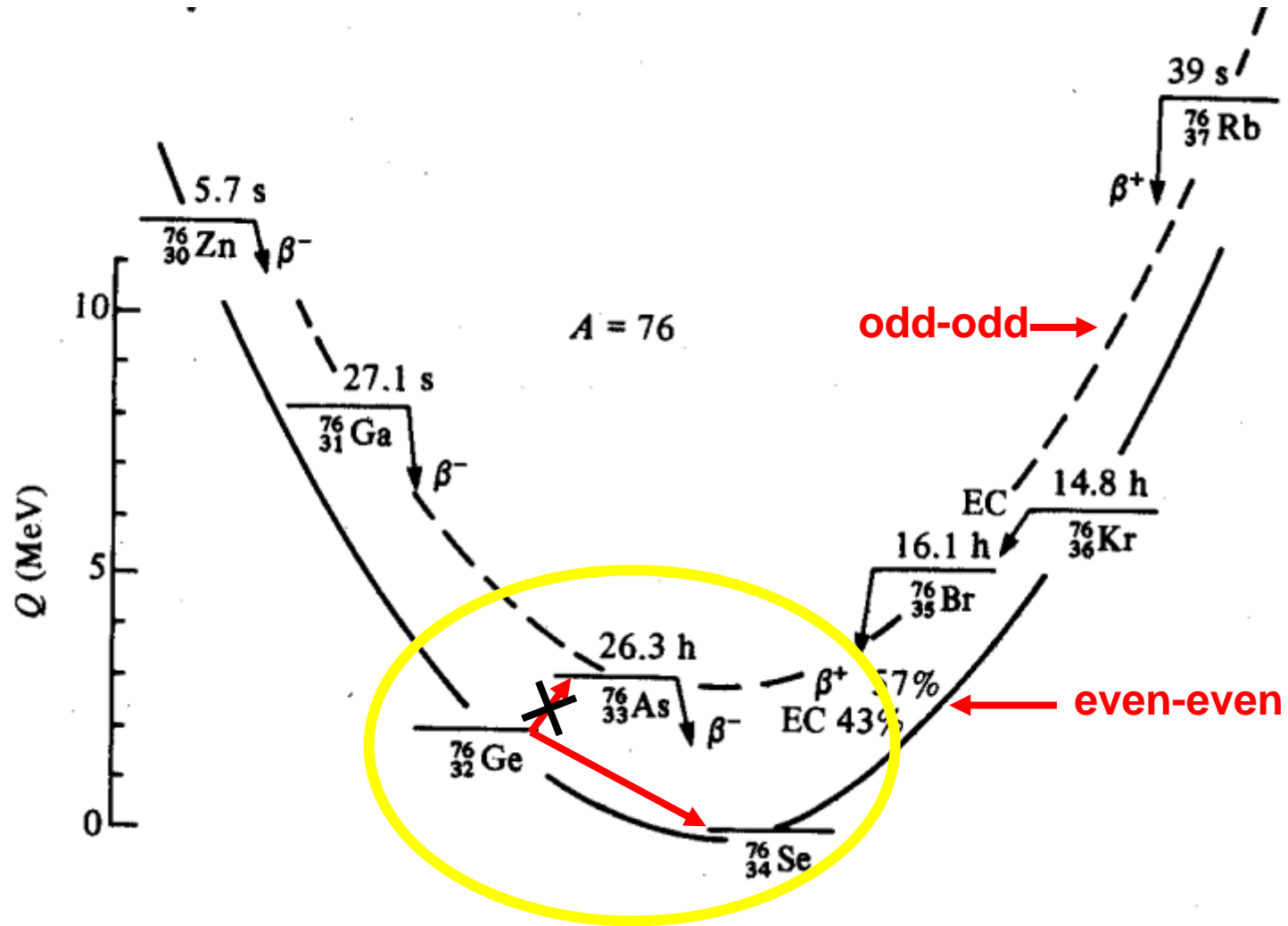


# Neutrinoless double beta decay search

GDR Neutrino Caen 2012, 30-31 October 2012

Stefan Schönert, Physik Department, TU München

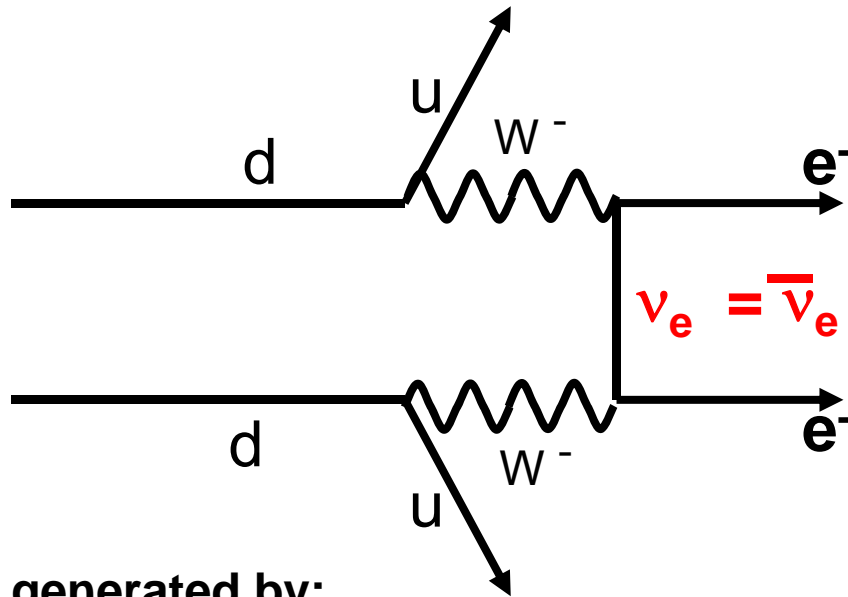
# Mass parabolas



Ground states of even-even nuclei:  $0^+$

# $0\nu\text{-}\beta\beta$ Decay

$$(A, Z) \rightarrow (A, Z + 2) + e_1^- + e_2^- \quad \Delta L=2$$



**$0\nu\beta\beta$  can be generated by:**

- exchange of light Majorana neutrinos (active/sterile)
- SUSY
- LR
- .....

Schechter & Valle:  
if  $0\nu\beta\beta$  observed  $\Rightarrow \nu$  is Majorana particle!

# Why is it interesting?

1) Violation of lepton number conservation ( $\Delta L=2$ )

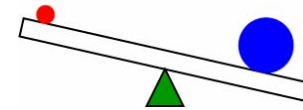
Dirac vs. Majorana particle: (i.e. its own anti-particle)?

$0\nu\beta\beta \Rightarrow$  Majorana nature

Majorana  $\Rightarrow$  See-Saw mechanism

$$m_\nu = \frac{m_D^2}{M_R} \ll m_D$$

For  $m_3 \sim (\Delta m_{\text{atm}}^2)^{1/2}$ ,  $m_D \sim m_t \Rightarrow M_R \sim 10^{15} \text{ GeV}$



2) Absolute mass scale:

(effective) neutrino mass

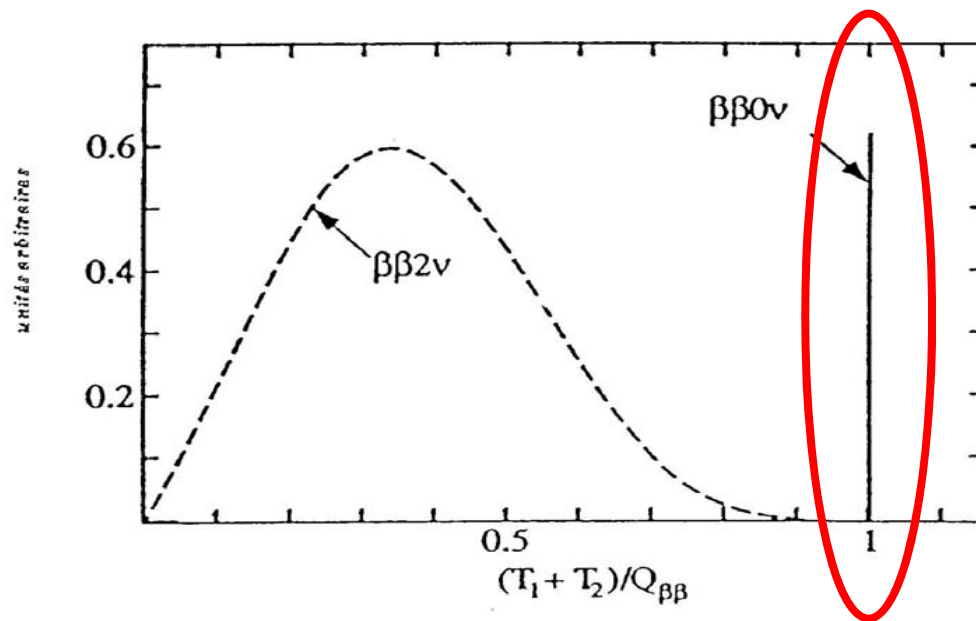
Hierarchy: degenerate, inverted or normal

sensitive to Majorana CP phases



# Main $0\nu\beta\beta$ experimental observables:

- Sum energy of 2 electrons equal to Q-value



- 2 electrons from one vertex
  - ➔ point-like in high density detectors
  - ➔ Two tracks from single vertex in tracking detector

# Life time and effective Majorana mass

$$1 / t_{1/2} (0\nu) = G M^2 m_{ee}^2$$

Phase space      Nuclear matrix element      Effective neutrino mass

$$m_{ee} = \left| \sum_i U_{ei}^2 m_i \right|$$

$U_{ei}$  complex:

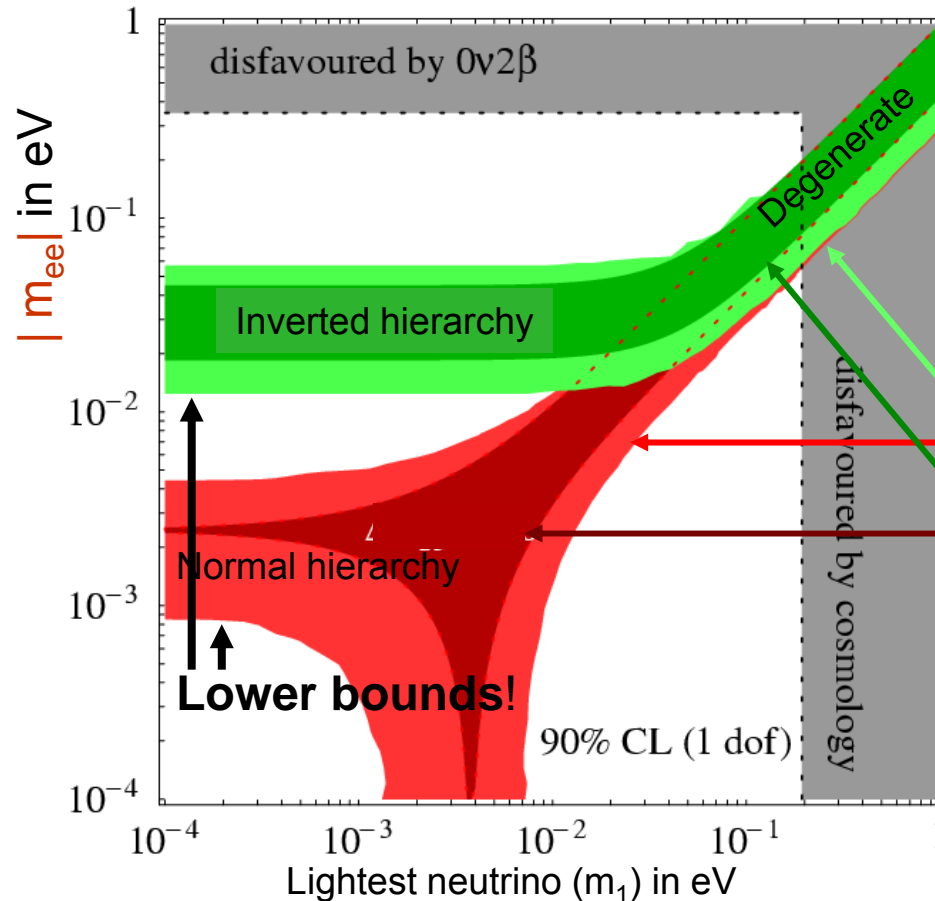
⇒ sensitive to CP violating phases (optimist☺)

⇒ cancellation possible (pessimist)

# $0\nu\beta\beta$ : Range of $m_{ee}$ derived from solar and atmospheric oscillation experiments

$$m_{ee} = f(m_1, \underbrace{\Delta m_{sol}^2, \Delta m_{atm}^2, \theta_{12}, \theta_{13}}_{\text{from oscillation experiments}}, \alpha-\beta)$$

$$\langle m_{ee} \rangle = \left| \sum_i U_{ei}^2 m_i \right|$$



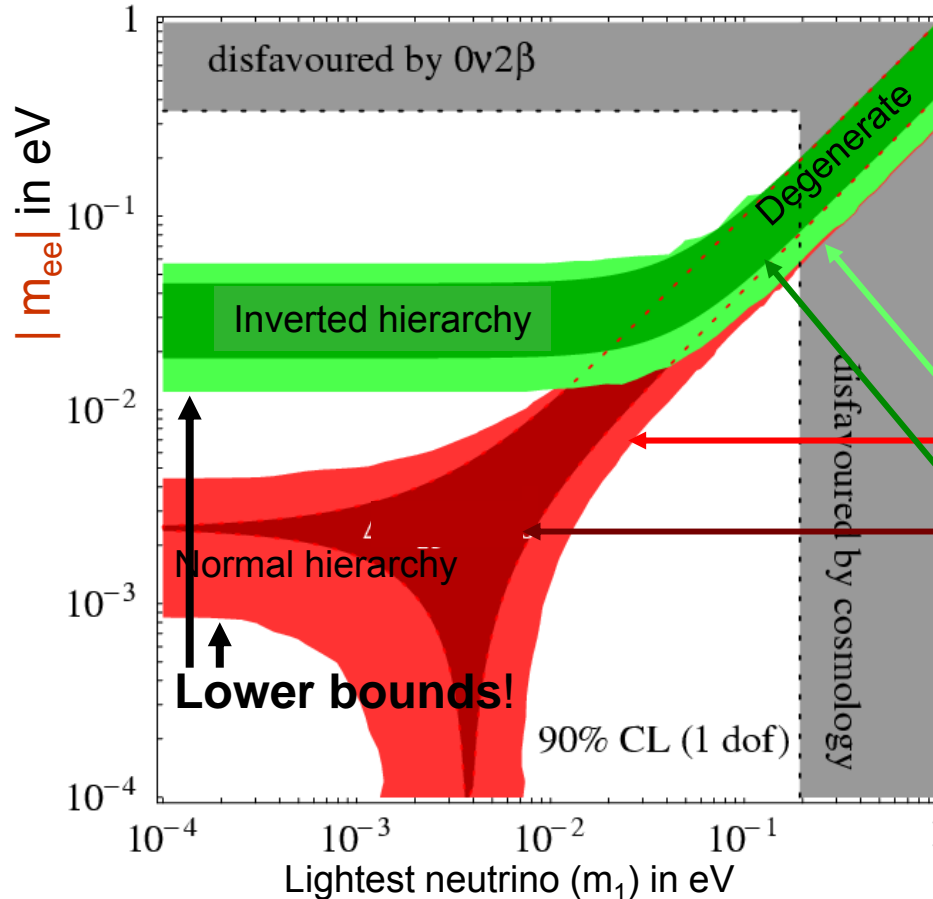
F.Feruglio,  
A. Strumia,  
F. Vissani,  
NPB 637

# $0\nu\beta\beta$ : Range of $m_{ee}$ derived from solar and atmospheric oscillation experiments

$$m_{ee} = f(m_1, \underbrace{\Delta m_{sol}^2, \Delta m_{atm}^2, \theta_{12}, \theta_{13}}_{\text{from oscillation experiments}}, \alpha-\beta)$$

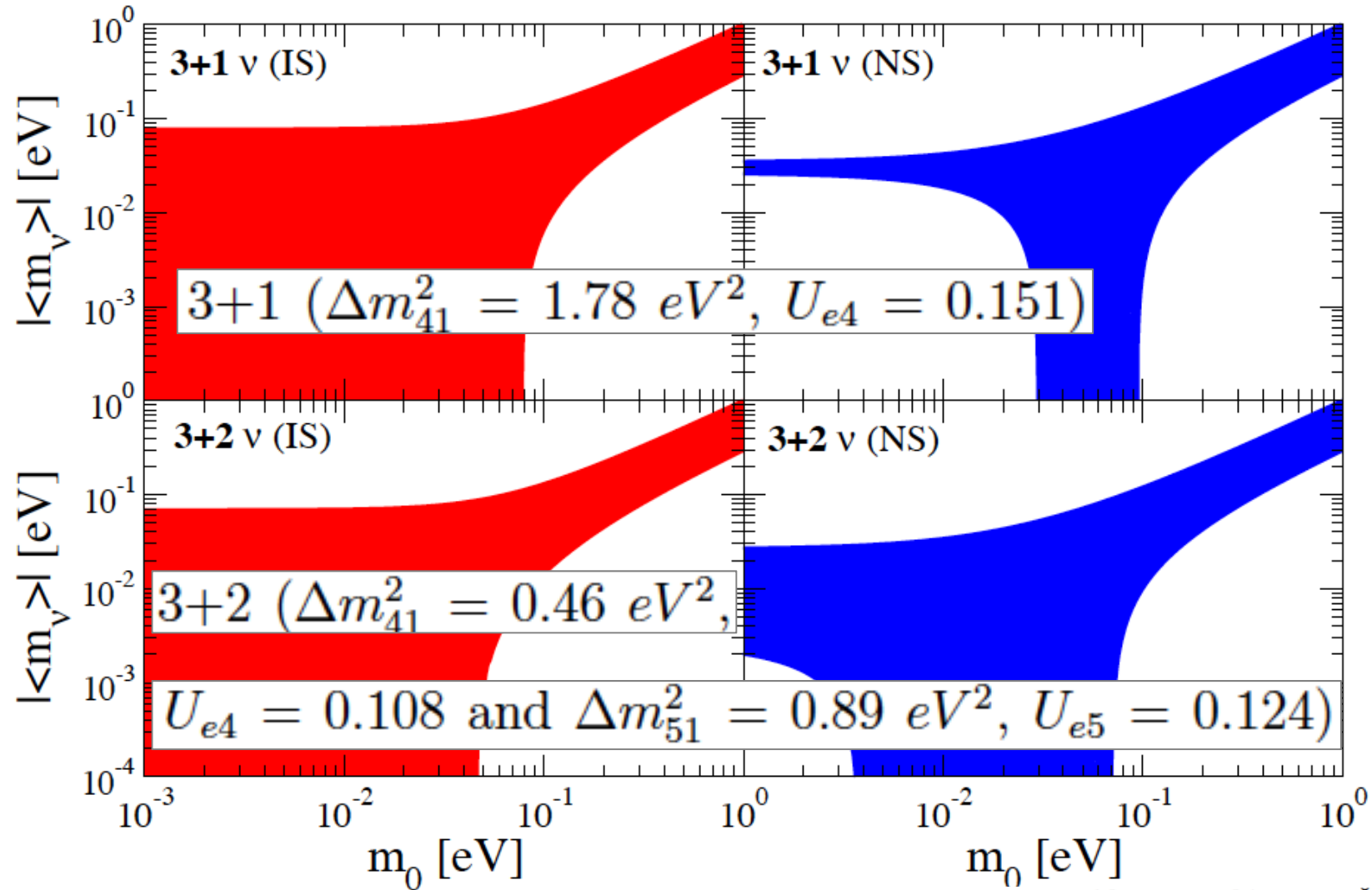
KDKC claim:  
0.44 eV

Goal of next  
generation  
experiments:  
~10 meV

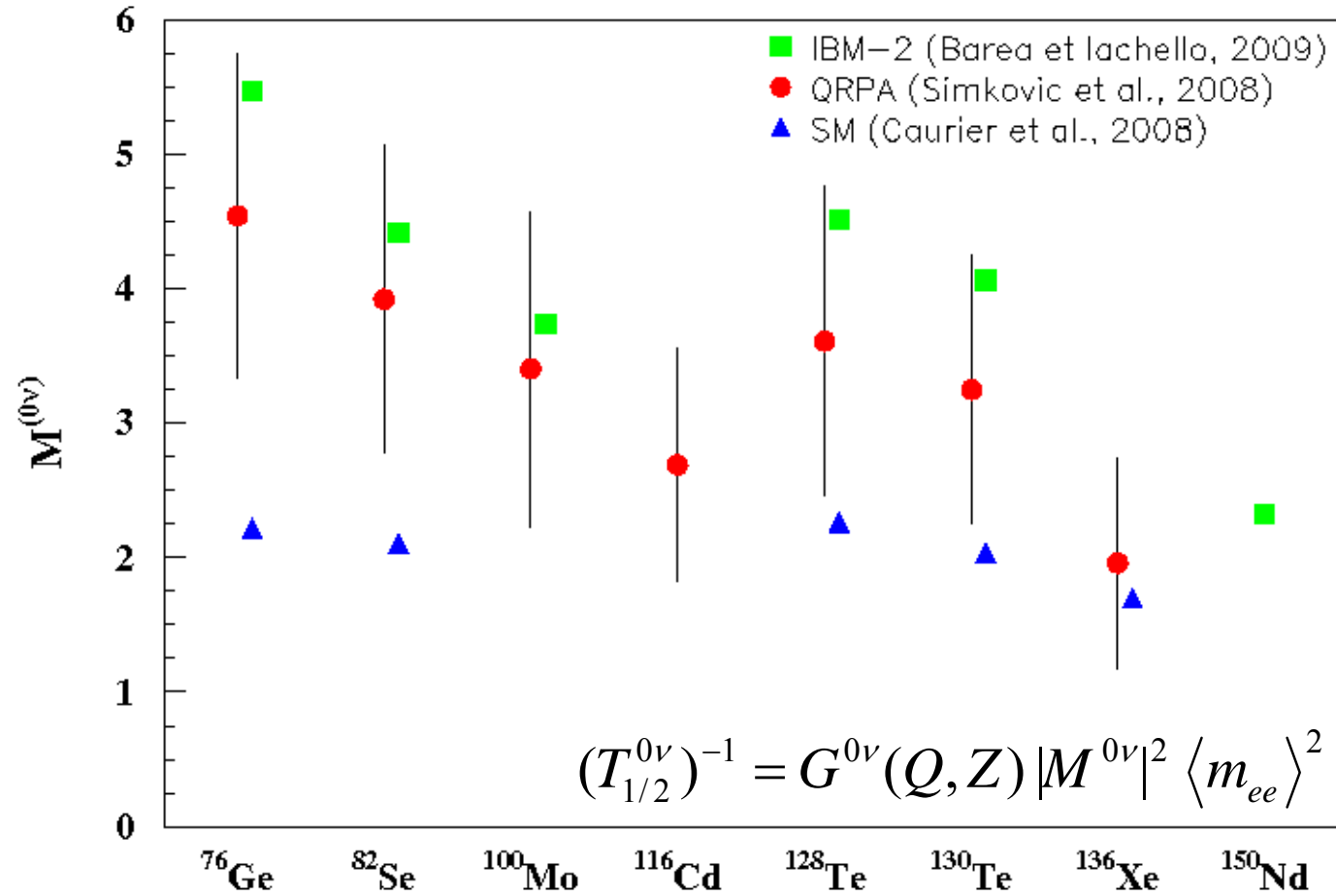


F.Feruglio,  
A. Strumia,  
F. Vissani,  
NPB 637

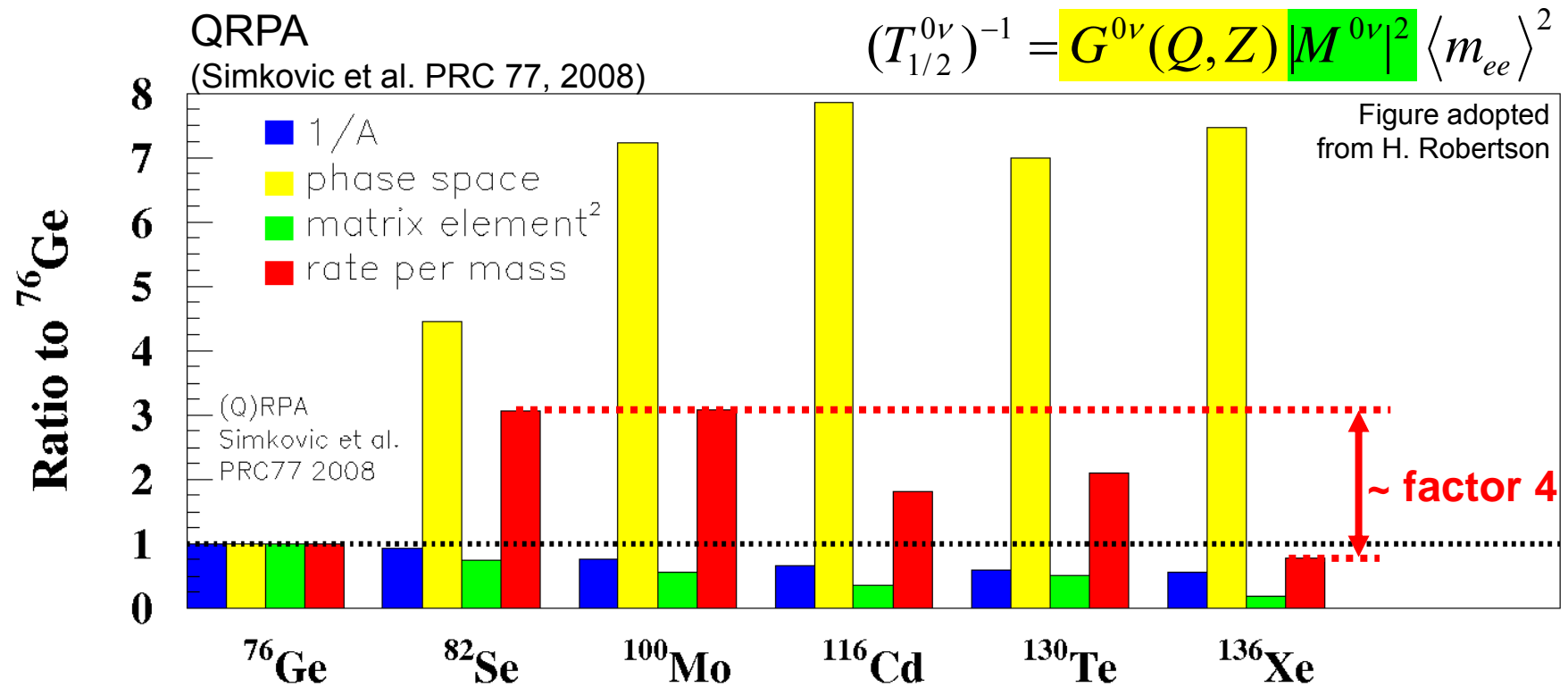
$0\nu\beta\beta$  including sterile neutrinos: range of  $m_{ee}$  derived from solar/atm experiments & reactor/Ga anomalies



# Comparison of DBD isotopes: NME's

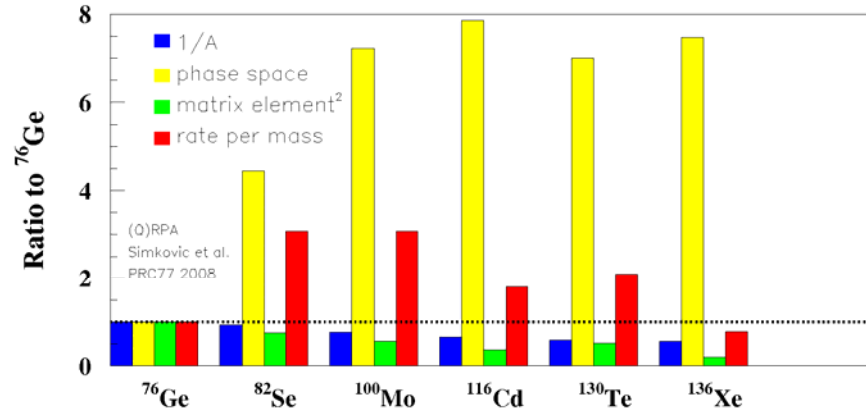


# Comparison of isotopes: Is there a *super-DBD-isotope* ?



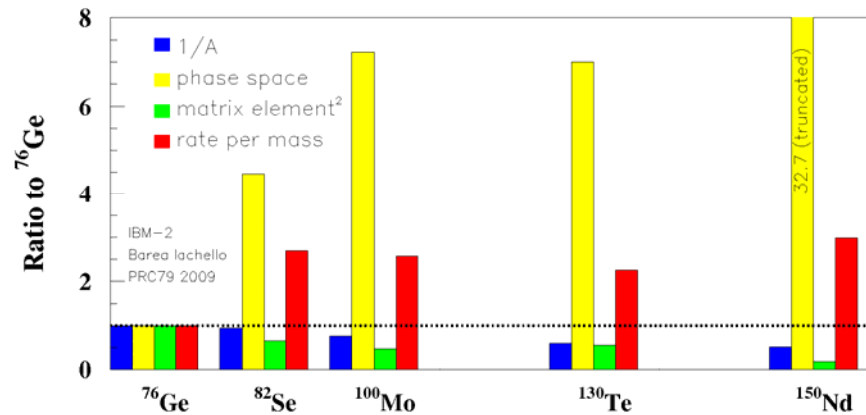
Expected  $0\nu\beta\beta$  **rates per mass** vary within a factor  $\sim 4$  !

QRPA  
(Simkovic et al.  
PRC 77, 2008)



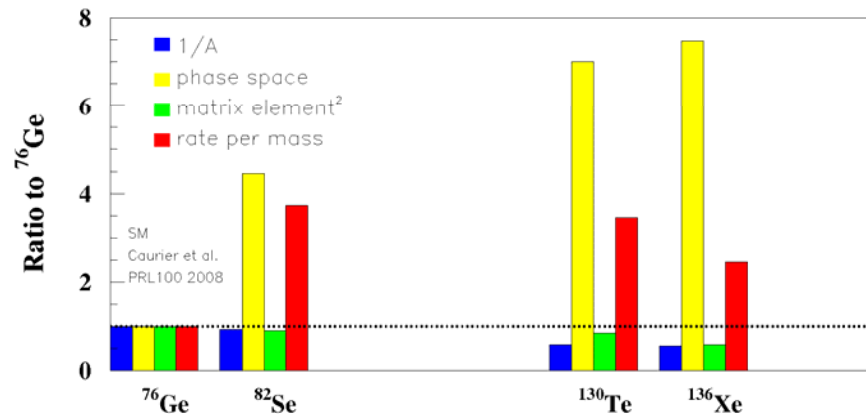
for  $\langle m \rangle = 50$  meV :  
9.1 cts/(ton year)

IBM2  
(Barea and  
Iachello, PRC  
79, 2009)



13.2 cts/(ton year)

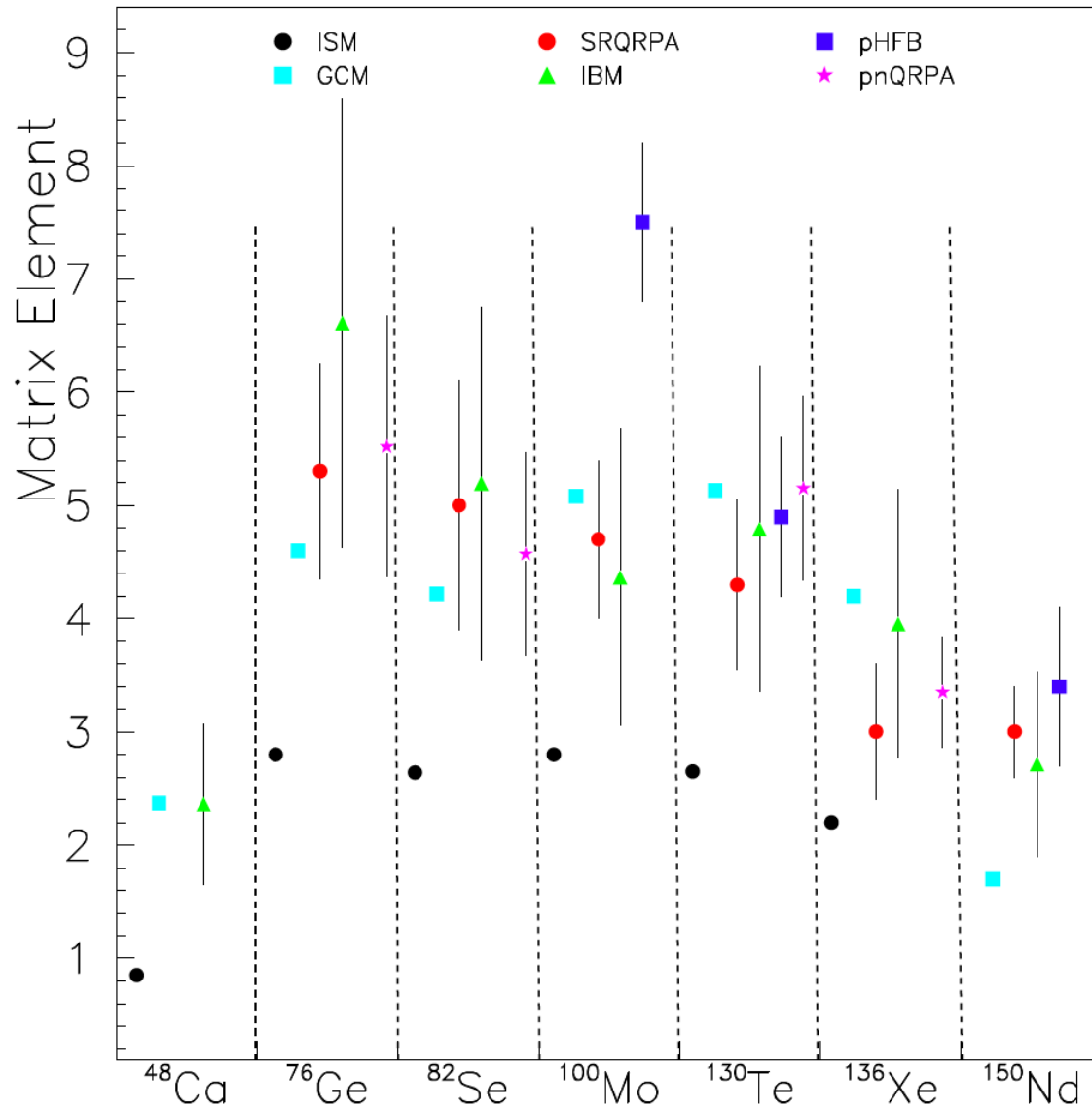
SM  
(Caurier et al.,  
PRL 100, 2008)



2.2 cts/(ton year)

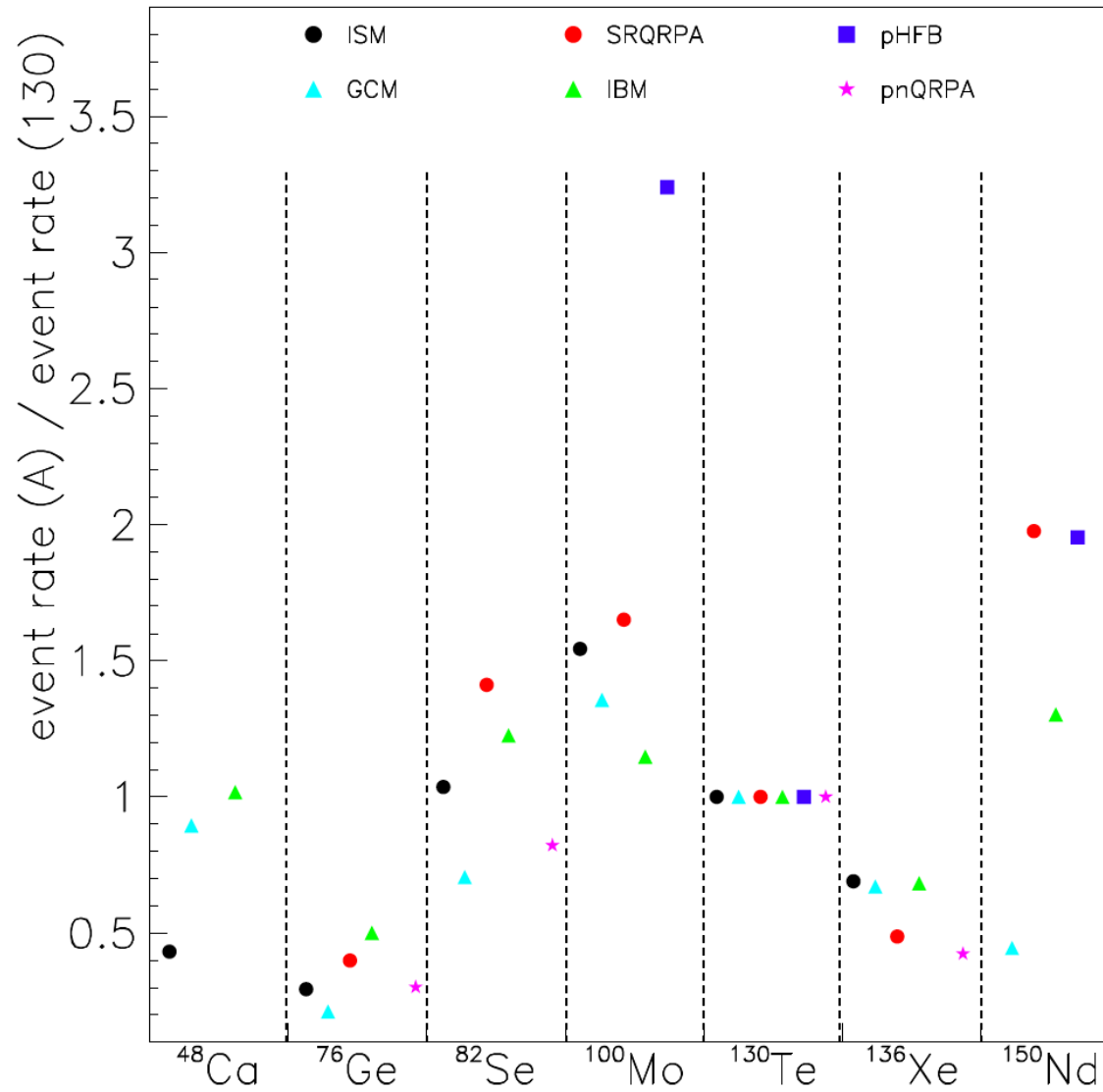


# Recent up date: NME's



B. Schwingerheuer, Annalen der Physik, August 22, 2012

# Recent up date: relative event rate per kg of isotopes



B. Schwingenheuer, Annalen der Physik, August 22, 2012

# Experimental sensitivity: w/o background

Experimental life time

$$\tau = \frac{N_N T}{N_S}$$

number of nuclides under control  $\propto M$

Time of measurement

number of detected decays

Background free limit:

0 cnts in the analysis energy window  $\Rightarrow$  Poisson upper limit:  $N_P$

Remember:  $1/t_{1/2}(0\nu) = G M^2 m_{ee}^2$

$$\tau \geq \frac{N_N T}{N_P} \propto M \cdot T \Rightarrow \langle m \rangle < \frac{\text{const}}{(M T)^{1/2}}$$

# Sensitivity: with background

If no decay is observed in presence of  $N_B$  background events in an energy window  $\Delta E$ :

$$N_S < (N_B)^{1/2} \quad \longrightarrow \quad \tau > \frac{N_N T}{(N_B)^{1/2}}$$

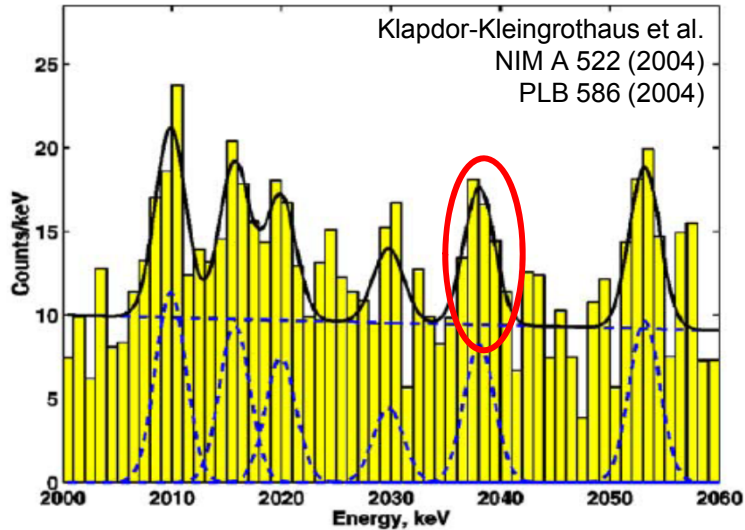
↑  
detector  
energy  
resolution

$$N_B = b M T \Delta E \quad \mathbf{b: \text{background index}} [1/(\text{kg} \cdot \text{year} \cdot \text{keV})]$$

$$\Rightarrow \tau > \frac{N_N T}{(b M T \Delta E)^{1/2}} \propto \left( \frac{M T}{b \Delta E} \right)^{1/2}$$

$$\Rightarrow \langle m \rangle < \text{const.} \cdot \left( \frac{b \Delta E}{M T} \right)^{1/4}$$

# As of early 2012: limits & claim

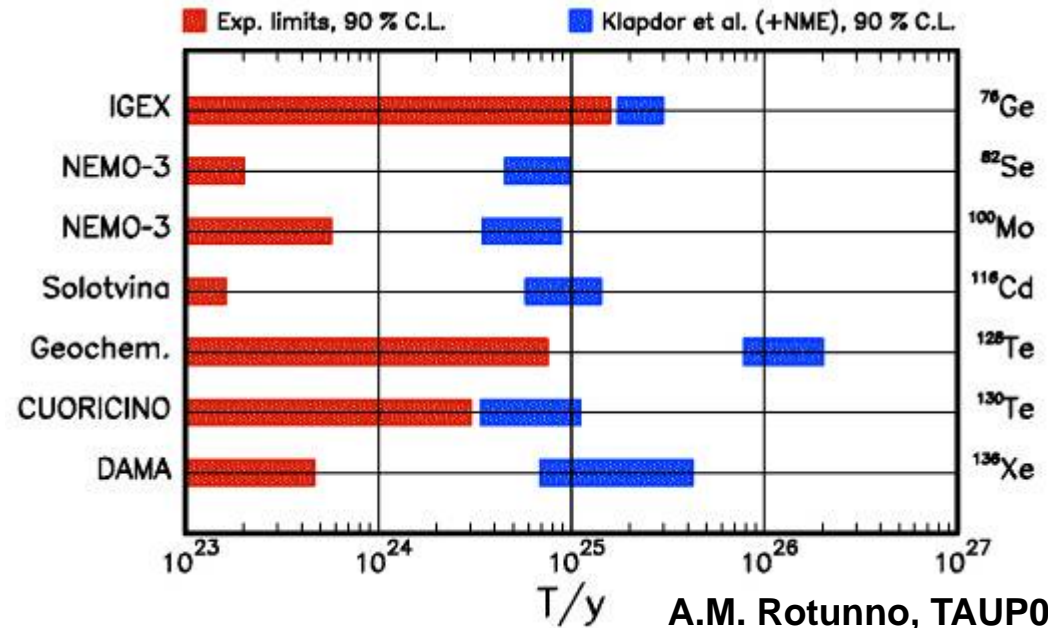


- 71.7 kg year - Bgd 0.11 / (kg y keV)
- $28.75 \pm 6.87$  events (bgd: ~60)
- Claim:  $4.2\sigma$  evidence for  $0\nu\beta\beta$
- $(0.69-4.18) \times 10^{25}$  y ( $3\sigma$ )
- Best fit:  $1.19 \times 10^{25}$  y (NIMA 522/PLB 586)
- PSA analysis (Mod. Phys. Lett. A21):  $(2.23 + 0.44 - 0.31) \times 10^{25}$  y ( $6\sigma$ ) (but analysis & results flawed....)
- Tuebingen/Bari group (PRD79):  $m_{ee} / eV = 0.28$  [0.17-0.45] 90%CL

Significance and  $T_{1/2}$  depend on bgd discription:

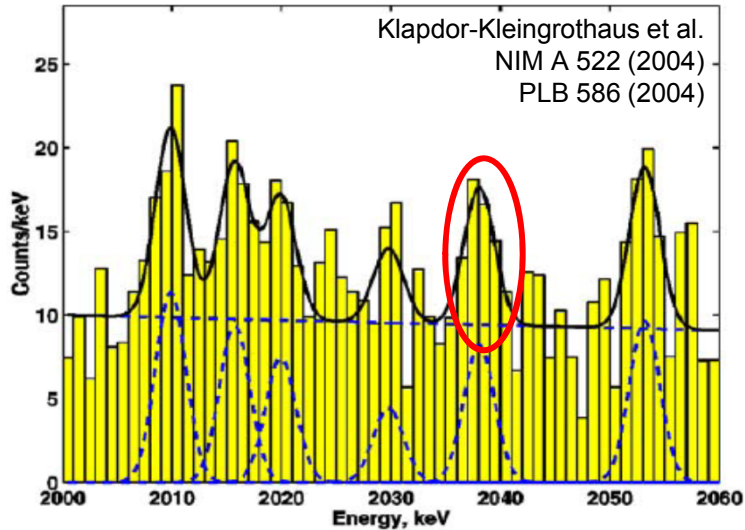
- Strumia & Vissani Nucl.Phys. B726 (2005)
- Chkvorets, PhD dissertation Univ. HD, (2008): using realistic background model

$\Rightarrow$  peak significance:  $1.3\sigma$ ,  
 $\Rightarrow T_{1/2} = 2.2 \times 10^{25}$  y



A.M. Rotunno, TAUP09 (PRD 79)

# As of early 2012: limits & claim

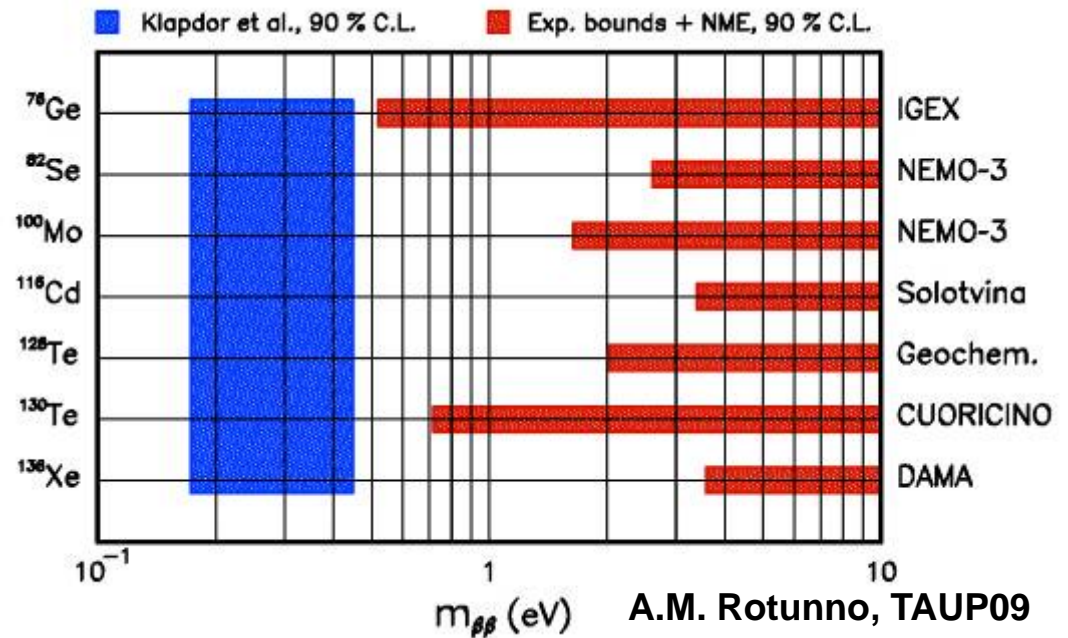


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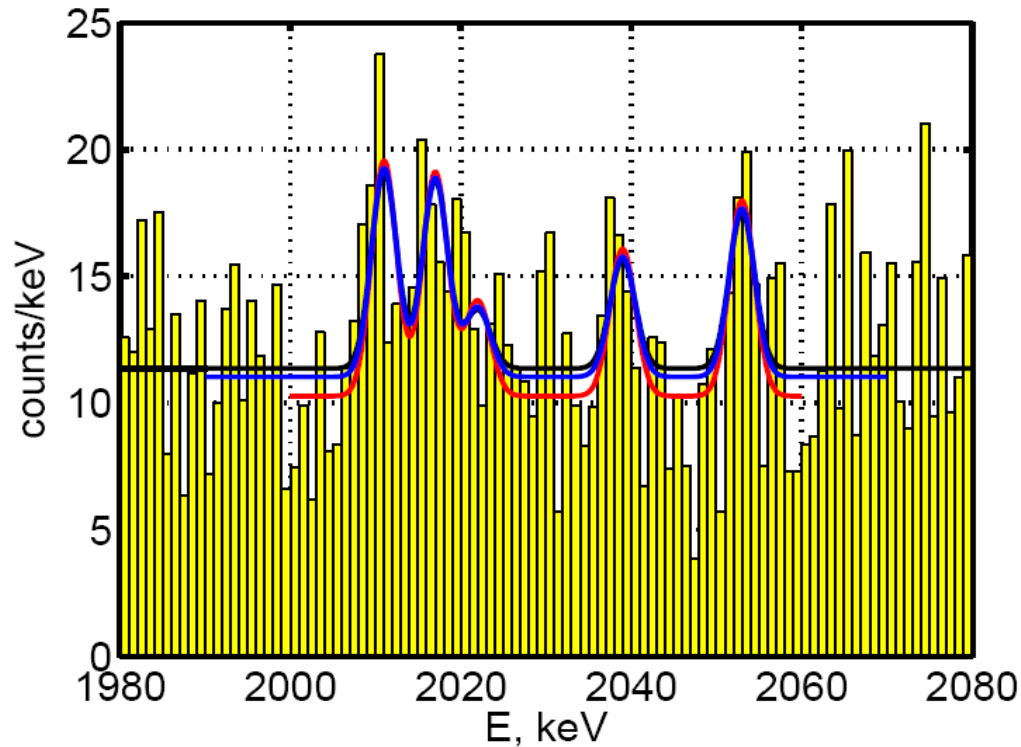
- Strumia & Vissani Nucl.Phys. B726 (2005)
- Chkvorets, PhD dissertation Univ. HD, (2008): using realistic background model

$\Rightarrow$  peak significance reduced to  $1.3\sigma$ ,  
 $\Rightarrow T_{1/2} = 2.2 \times 10^{25}$  y

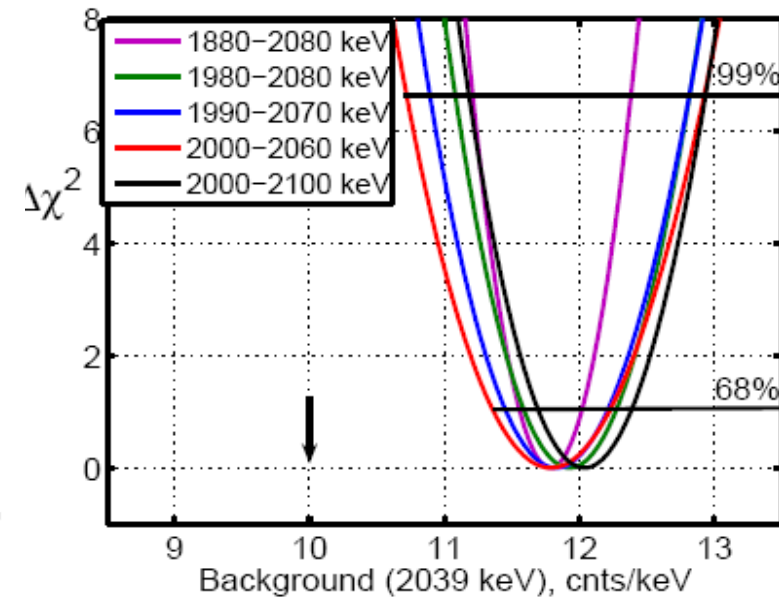


$\Rightarrow$  Claim must be scrutinized with  $^{76}\text{Ge}$  AND other isotopes

# Significance of claim weaker if background uncertainties included



O.Chkvorets, Dissertation  
 Univ. Heidelberg 2008;  
 see also Vissani Feruglio, Strumia)



Method	$E_{peak}$ [keV]	Peak area [counts]
$S_{Feldman-Cousins}$	2039.0	$15 \pm 12$
$S_{Fit}$	$2038.1 \pm 1.3$	$13.0 \pm 8.5$

**Significance:  
 ~1.3-1.5  $\sigma$  !**

Table 4.7: The 2039 keV peak intensity determined with Feldman-Cousins and LSQ methods using background  $11.9 \pm 0.5$  counts/keV.

**Table 3** Selection of  $0\nu\beta\beta$  experiments.

experiment	isotope	mass [kg]	method	start / end	ref.
past experiments					
Heidelberg-Moscow	$^{76}\text{Ge}$	11	ionization	-2003	[6]
Cuoricino	$^{130}\text{Te}$	11	bolometer	-2008	[40]
NEMO-3	$^{100}\text{Mo}$ , $^{82}\text{Se}$	7,1	track. +calorim.	-2011	[41]
current experiments					
EXO-200	$^{136}\text{Xe}$	175	liquid TPC	2011-	[2]
Kamland-Zen	$^{136}\text{Xe}$	330	liquid scintil.	2011-	[3]
GERDA-I/ GERDA-II	$^{76}\text{Ge}$	15/35	ionization	2011-/ 2013-	[45]
CANDLES	$^{48}\text{Ca}$	0.35	scint. crystal	2011-	[46]
funded experiments					
NEXT	$^{136}\text{Xe}$	100	gas TPC	2015	[47]
Cuore0/ Cuore	$^{130}\text{Te}$	10/200	bolometer	2012- 2015-	[48]
Majorana Demo.	$^{76}\text{Ge}$	30	ionization	2013	[49]
SuperNEMO demo./total	$^{82}\text{Se}$	7/100	track.+calorim	2014- /??	[50]
SNO+	$^{150}\text{Nd}$	44	liquid scint.	2013	[51]

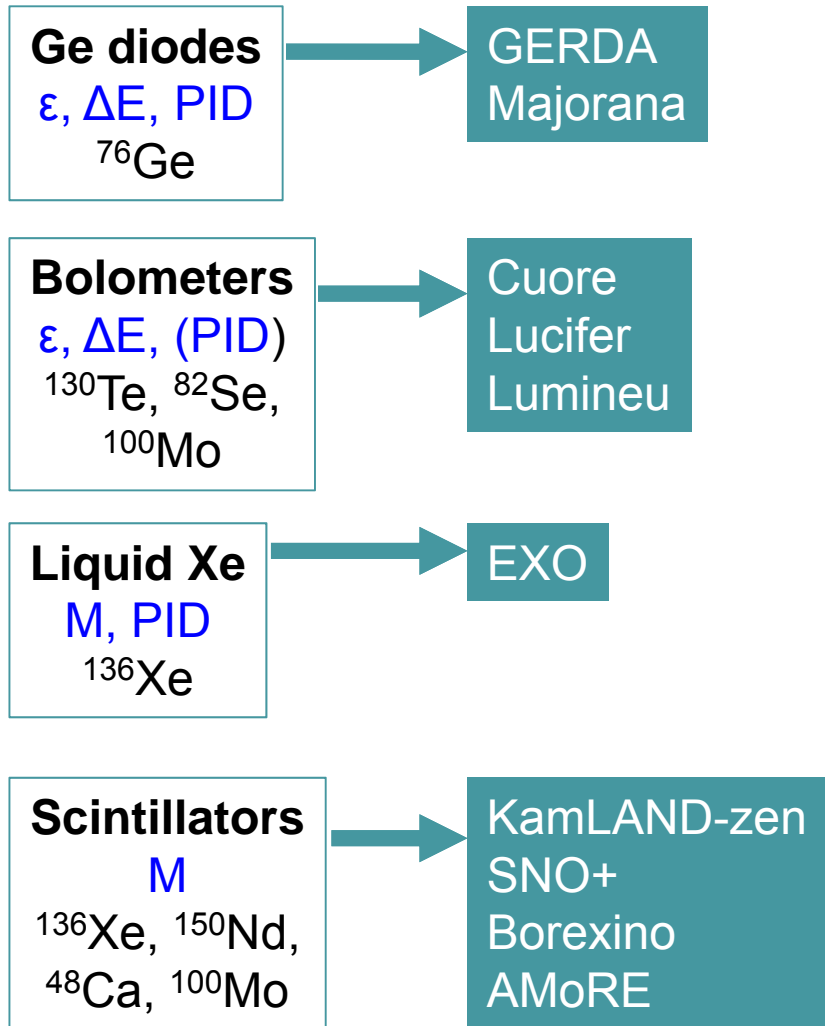
proposal, proto-typing				
Cobra	$^{116}\text{Cd}$		solid TPC	[52]
Lucifer	$^{82}\text{Se}$		bolom. +scint.	[53]
DCBA/MTD	$^{150}\text{Nd}$	32	tracking	[54]
MOON	$^{82}\text{Se}$ , $^{100}\text{Mo}$	30- 480	track. +scint.	[55]
AMoRE	$^{100}\text{Mo}$	100	bolom. +scint.	[56]
Cd exp.	$^{116}\text{Cd}$		scint.	[57]

B. Schwingenheuer, Annalen der Physik,  
August 22, 2012

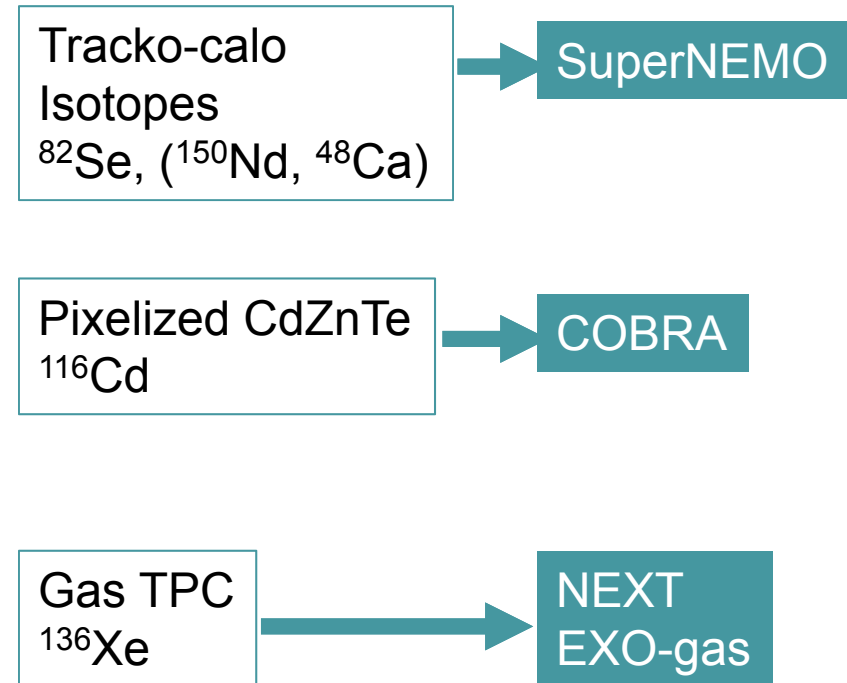


# Next generation experiments

## Calorimeters



## Electron tracking



Adopted from F. Piquemal

# Next generation experiments

## Calorimeters

**Ge diodes**  
 $\epsilon$ ,  $\Delta E$ , PID  
 $^{76}\text{Ge}$

→ **GERDA**  
Majorana

**Bolometers**  
 $\epsilon$ ,  $\Delta E$ , (PID)  
 $^{130}\text{Te}$ ,  $^{82}\text{Se}$ ,  
 $^{100}\text{Mo}$

→ **Cuore/**  
**Coureo**  
Lucifer  
Lumineu

**Liquid Xe**  
M, PID  
 $^{136}\text{Xe}$

→ **EXO**

**Scintillators**  
M  
 $^{136}\text{Xe}$ ,  $^{150}\text{Nd}$ ,  
 $^{48}\text{Ca}$ ,  $^{100}\text{Mo}$

→ **KamLAND-zen**  
SNO+  
Borexino  
AMoRE

## Electron tracking

Tracko-calor  
Isotopes  
 $^{82}\text{Se}$ , ( $^{150}\text{Nd}$ ,  $^{48}\text{Ca}$ )

→ **SuperNEMO**

Pixelized CdZnTe  
 $^{116}\text{Cd}$

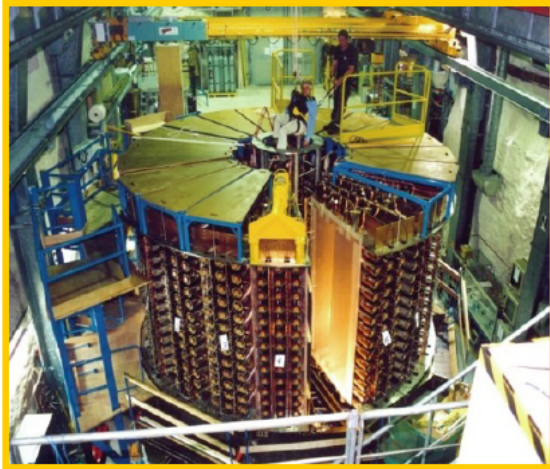
→ **COBRA**

Gas TPC  
 $^{136}\text{Xe}$

→ **NEXT**  
EXO-gas

**Experiments started  
operations in 2011/12**

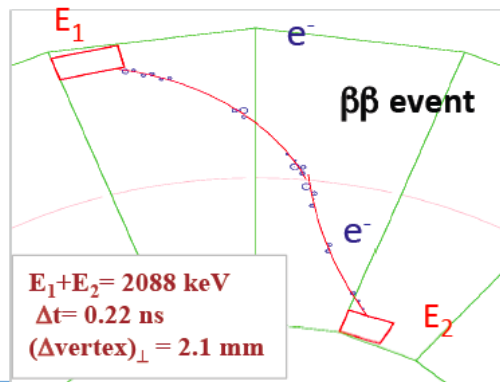
# NEMO 3



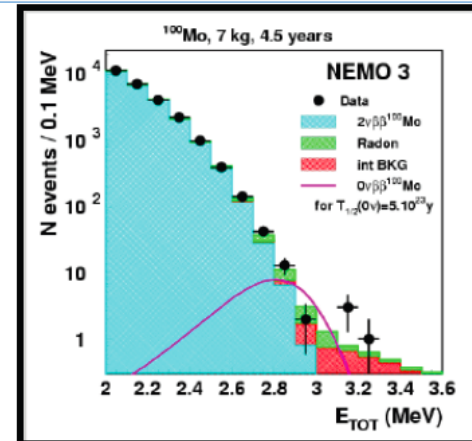
Tracking detector: drift chambers (6180 Geiger cells)  
 $\sigma_t = 5 \text{ mm}, \sigma_z = 1 \text{ cm}$  (vertex)  
 Calorimeter (1940 plastic scintillators and PMTs)  
 Energy Resolution FWHM=8 % (3 MeV)  
 Identification  $e^-, e^+, \gamma, \alpha$   
 Very high efficiency for background rejection  
 Background level @  $Q_{\beta\beta}$  [2.8 – 3.2 MeV] :  $1.2 \cdot 10^{-3}$  cts/keV/kg/y  
 Multi-isotope (7 measured at the same time)  
 Running at Modane underground laboratory (2003 - 2011)

## Unique feature

Measurement of all kinematic parameters:  
 individual energies and angular distribution



Measurement of 7 isotopes  $\beta\beta(2\nu)$  half-lives  
 Excited states, Majoron limits for  $\beta\beta(0\nu)$



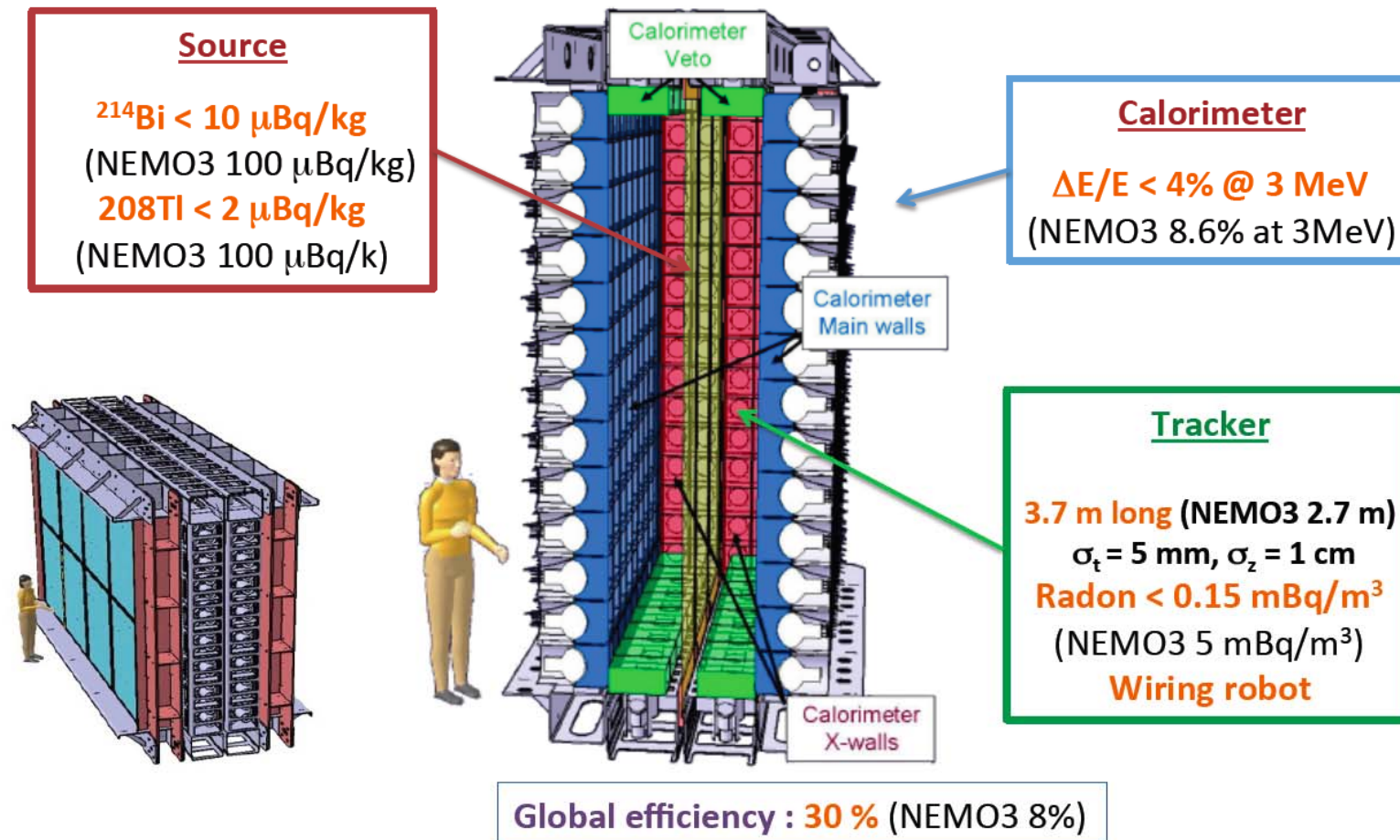
[2.8 – 3.2] MeV 18 observed events,  $16.4 \pm 1.3$  expected

$^{100}\text{Mo}$   $T_{1/2}(\beta\beta 0\nu) > 1.0 \cdot 10^{24} \text{ y}$  (90% C.L.)  
 $\langle m_{\nu} \rangle < 0.31 - 0.79 \text{ eV}$

From F. Piquemal, Neutrino 2012

# SuperNEMO Demonstrator

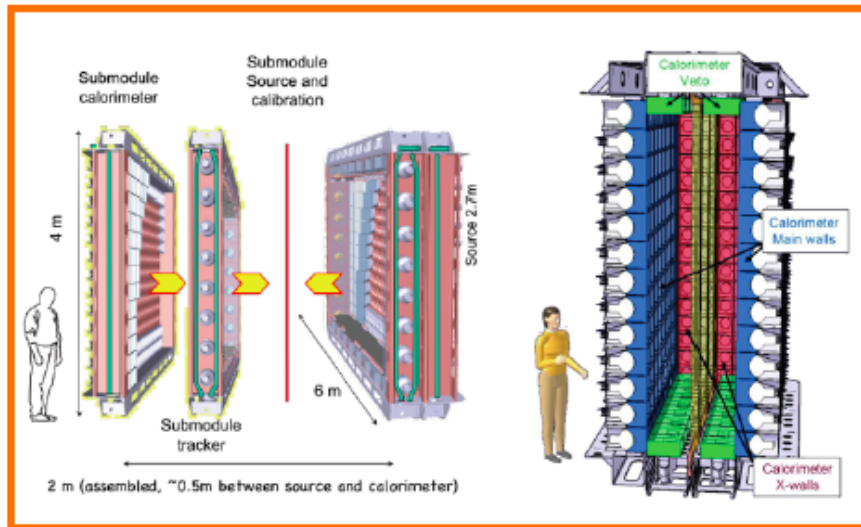
**Objective:** to reach the background level for 100 kg  
to perform a no background experiment with 7 kg isotope of  $^{82}\text{Se}$  in 2 yr



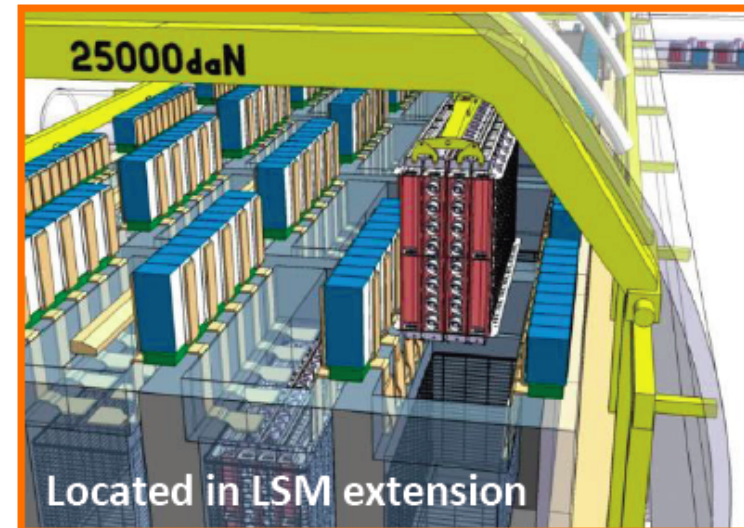
From F. Piquemal, Neutrino 2012

# SuperNEMO

## A module



## 20 modules



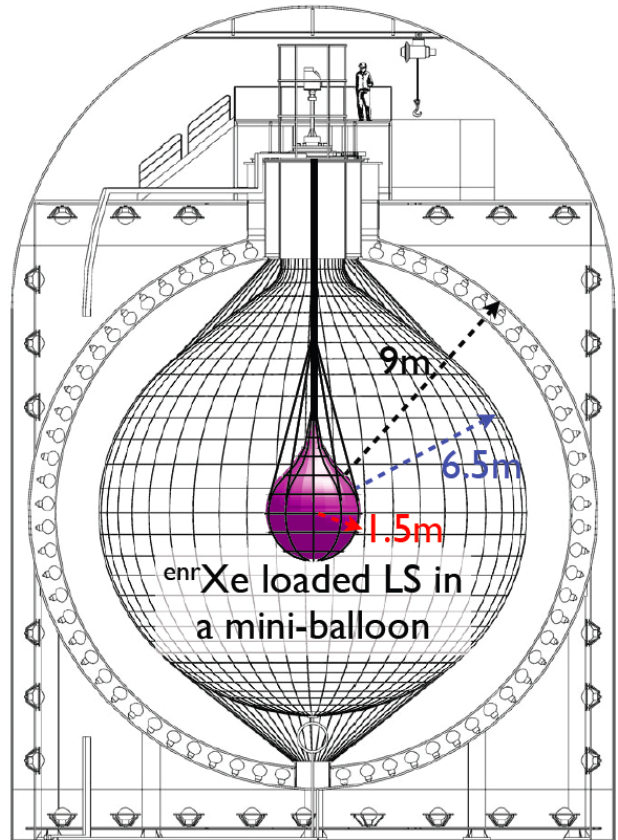
	Demonstrator module	20 Modules
Source : $^{82}\text{Se}$	7 kg	100 kg
Drift chambers for tracking	2 0000	40 000
Electron calorimeter	500	10 000
$\gamma$ veto (up and down)	100	2 000
$T_{1/2}$ sensitivity	$6.6 \cdot 10^{24}$ y (No background)	$1 \cdot 10^{26}$ y
$\langle m_{\nu} \rangle$ sensitivity	200 – 400 meV	40 – 100 meV

From F. Piquemal, Neutrino 2012



# KamLAND-Zen

Zero Neutrino  
double beta decay search



idea to load Xe into LS is from Raju PRL72,1411(1994)

~320kg 90% enriched  $^{136}\text{Xe}$  installed so far  
total 600+ kg in the mine  
production reaches 700kg in this year

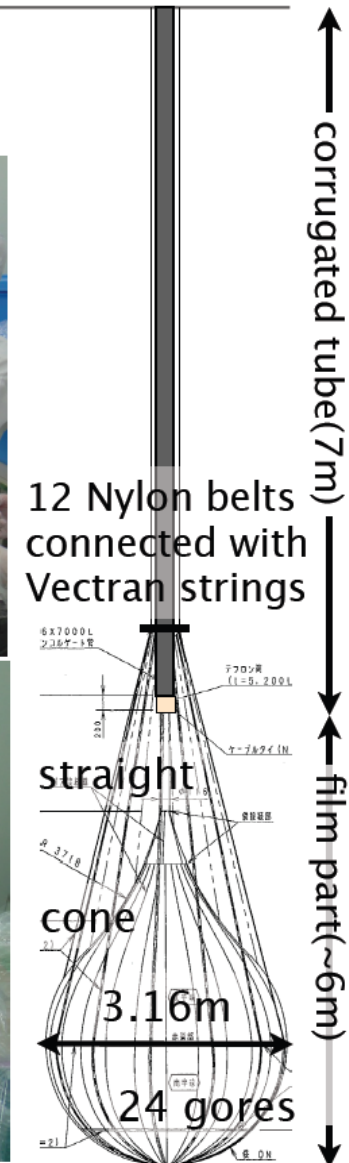
From K. Inoue, Neutrino 2012

# mini balloon fabrication

produced in a class 1 super-clean-room  
(class 1 = less than 1 0.5-micron-particle in cubic feet)

less material → 25 μm Nylon6  
transparency 99.4% @400nm  
strength 19.4 N/cm  
Xe barrier < 220 g/year

low radioactivity  
→ specially made no filler film  
U : 150 → 2x10<sup>-12</sup>g/g  
Th : 59 → 3x10<sup>-12</sup>g/g  
<sup>40</sup>K : 140 → 2x10<sup>-12</sup>g/g



From K. Inoue, Neutrino 2012



Installation in a class 10~100 clean room built at the top of KamLAND



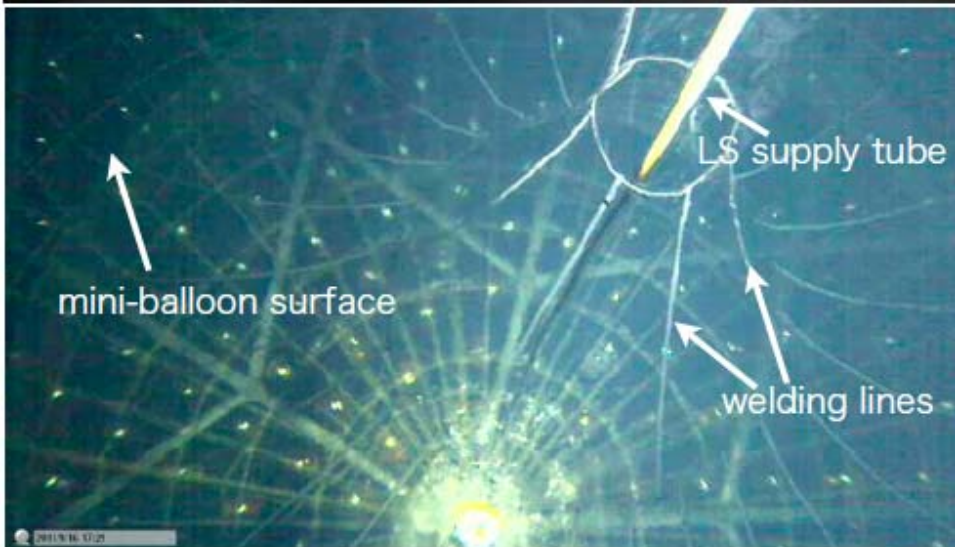
balloon and corrugated tube deployment



balloon went through the black sheet



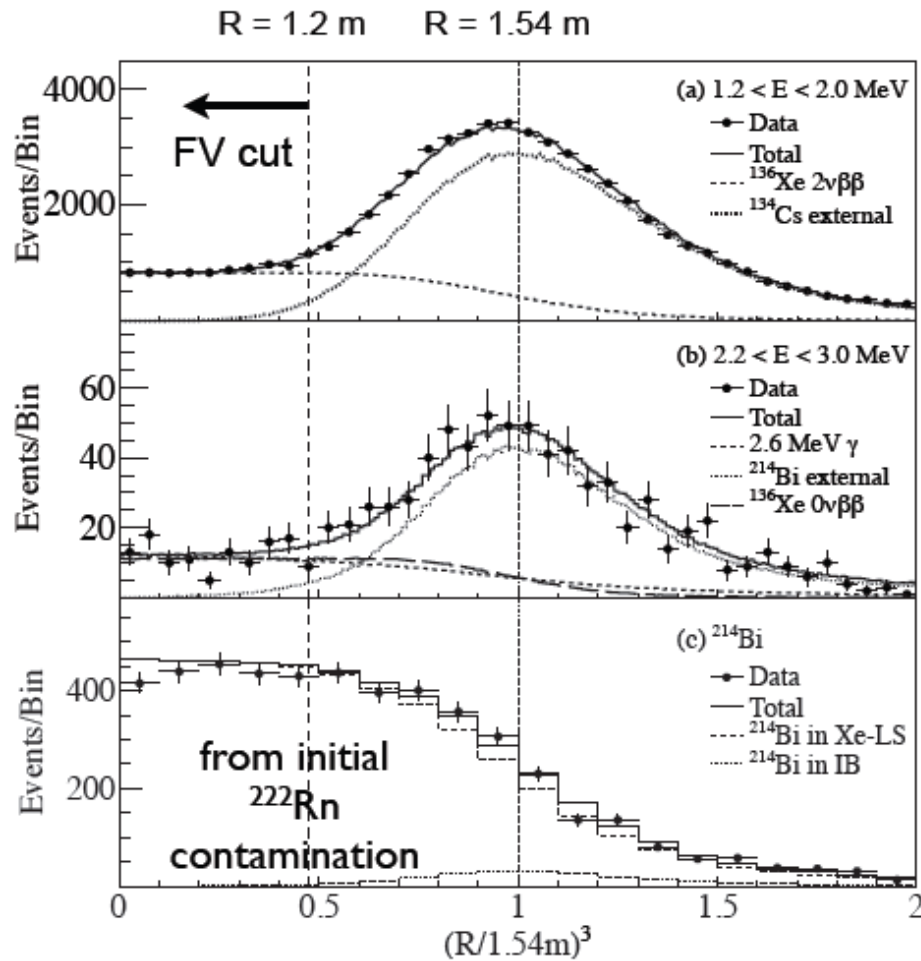
installation completed



mini-balloon inflated with dummy LS and then replaced with Xe-loaded LS  
density tuning finished and tubes to be extracted<sub>5</sub>



## R<sup>3</sup> vertex distribution



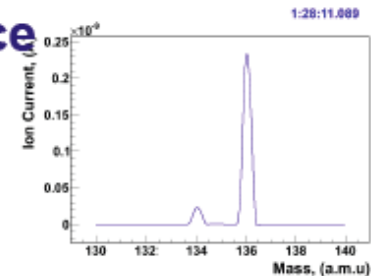
vertex resolution from balloon events  
 $\sim 15\text{cm}/\sqrt{E}$

## Fiducial volume of Xe-LS

$$R < 1.2 \text{ m} \quad V = 7.24 \text{ m}^3$$

## Xe isotopic abundance

$^{136}\text{Xe}$  : 90.93%,  
 $^{134}\text{Xe}$  : 8.89%



## Xe amount by weight measurement

(supplied)-(returned)

- Xe concentration =  $(2.44 \pm 0.01)$  wt%

## consistent direct gas chromatography analysis

- Xe concentration =  $(2.52 \pm 0.07)$  wt%

## Uncertainty of the fiducial volume

$^{214}\text{Bi}$  vertex uniformity check

$$V_{1.2\text{m}} / V_{\text{total}} = 0.438 \pm 0.005$$

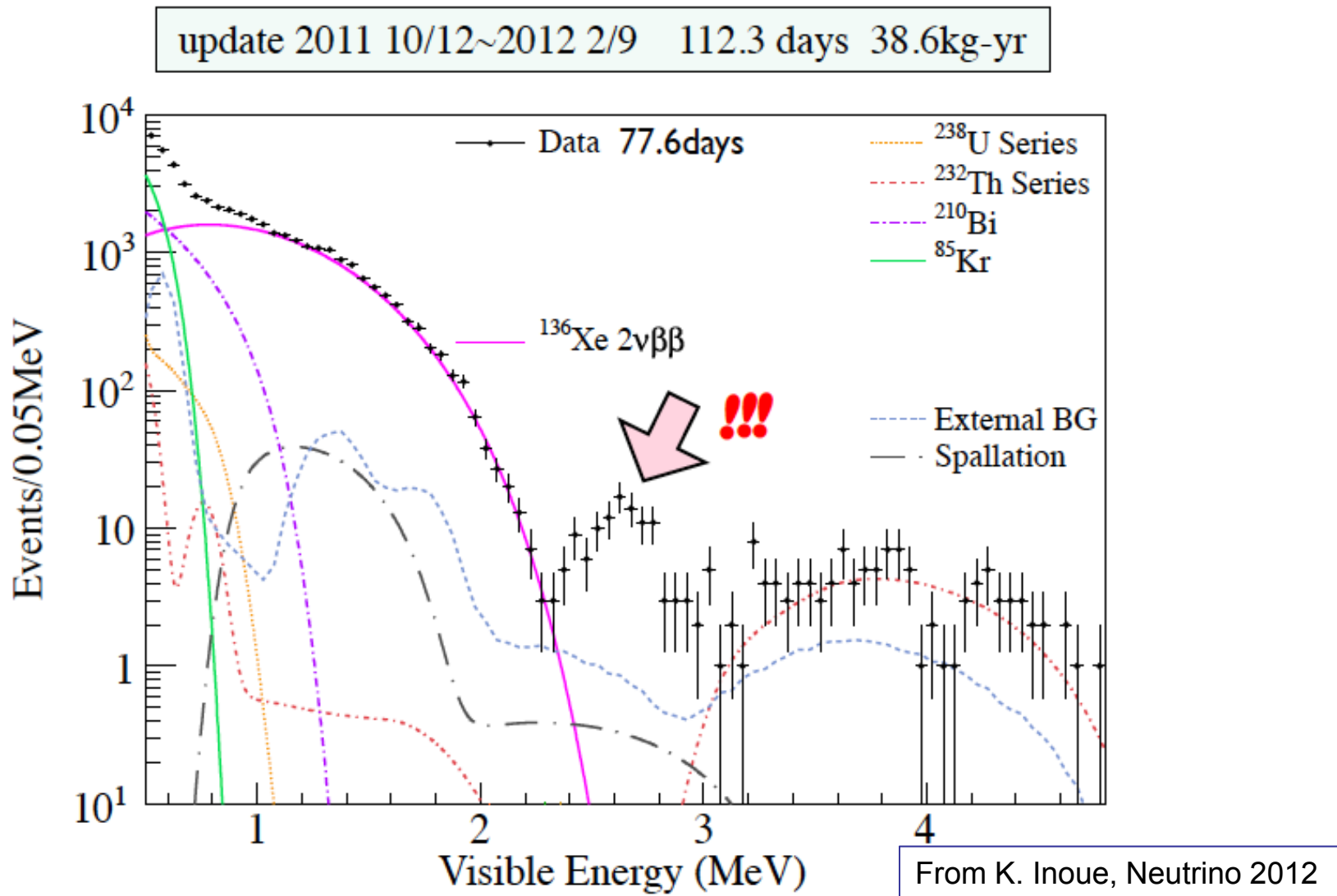
$$N_{1.2\text{m}} / N_{\text{total}} = 0.423 \pm 0.007(\text{stat}) \pm 0.004(\text{syst})$$

FV uncertainty  $\rightarrow$  5.2% dominant error

## Xe amount in the fiducial volume

125 $\pm$ 7 kg of  $^{136}\text{Xe}$  in the FV

# KamLAND-Zen: energy spectrum



# Measurement of the $2\nu 2\beta$ half life

DAMA (2002)

Liquid Xe scintillator

$$T_{1/2}^{2\nu} > 1.0 \times 10^{22} \text{ years at 90\% CL} \quad \text{Phys.Lett.B546,23(2002)}$$

factor 5 contradiction

EXO-200 (2011)

Liquid Xe TPC + scintillator

$$T_{1/2}^{2\nu} = 2.11 \pm 0.04(\text{stat}) \pm 0.21(\text{syst}) \times 10^{21} \text{ years}$$

Phys.Rev.Lett.107,212501(2011)

update

$$T_{1/2}^{2\nu} = 2.23 \pm 0.017(\text{stat}) \pm 0.22(\text{syst}) \times 10^{21} \text{ years}$$

arXiv:1205.5608

KamLAND-Zen (2012)

Xe loaded liquid scintillator

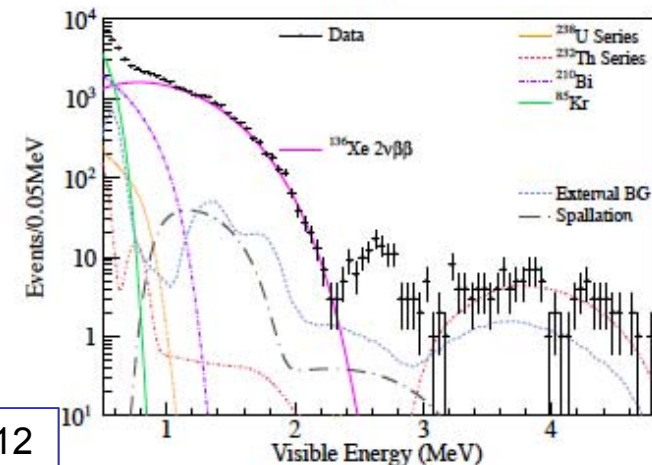
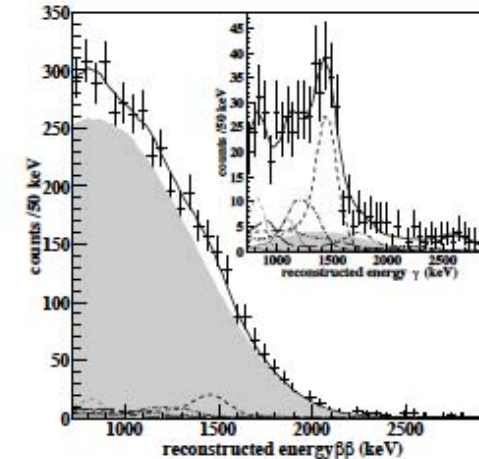
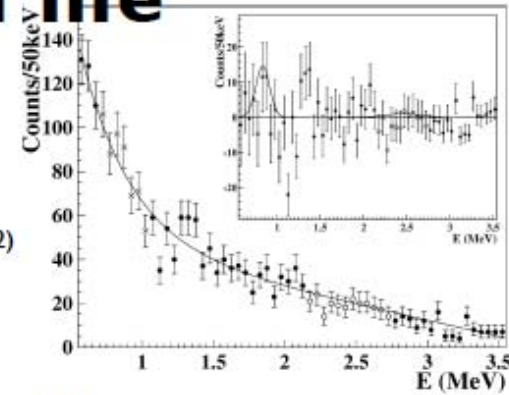
$$T_{1/2}^{2\nu} = 2.38 \pm 0.02(\text{stat}) \pm 0.14(\text{syst}) \times 10^{21} \text{ years}$$

Phys.Rev.C85,045504(2012)

update

$$T_{1/2}^{2\nu} = 2.30 \pm 0.02(\text{stat}) \pm 0.12(\text{syst}) \times 10^{21} \text{ years}$$

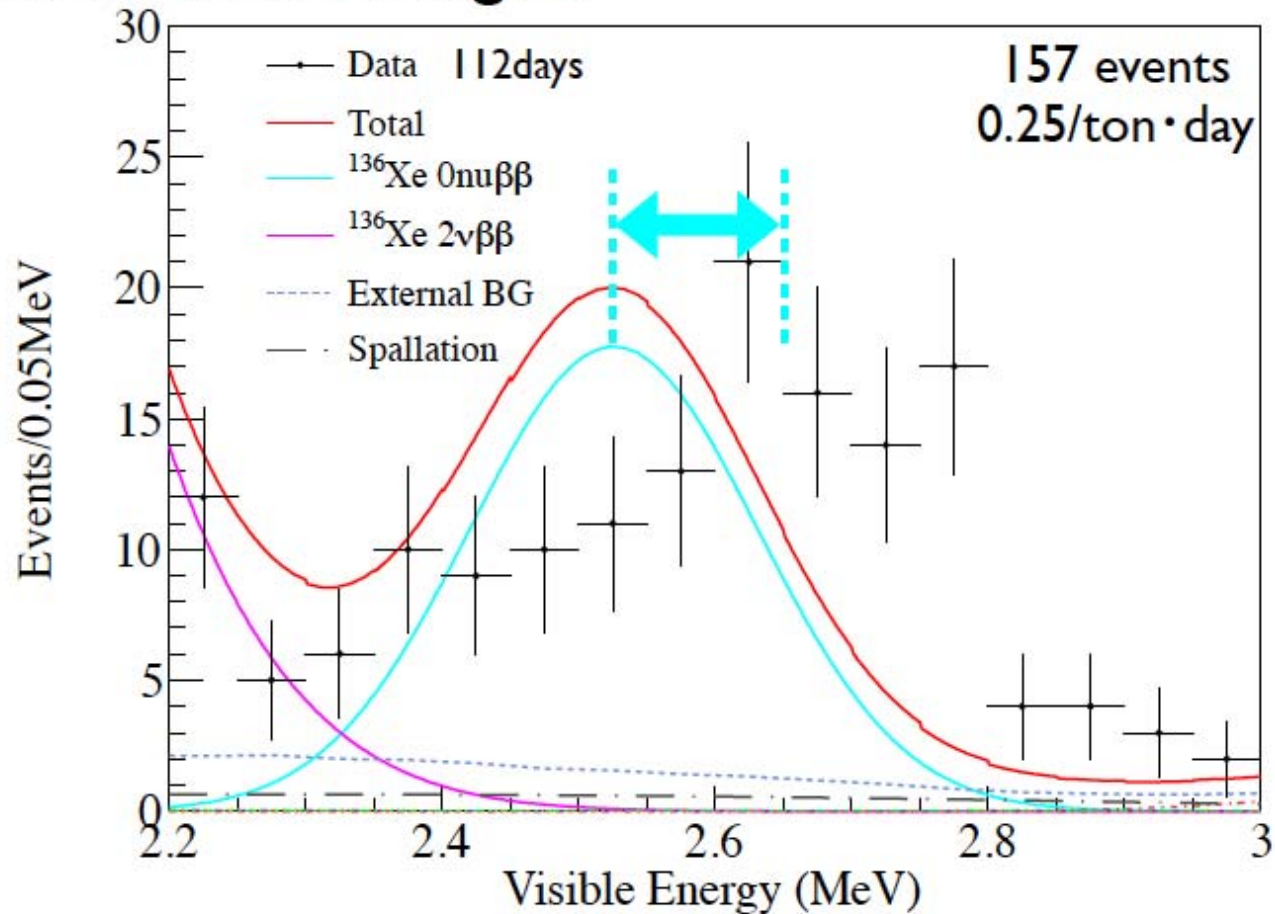
arXiv:1205.6372



From K. Inoue, Neutrino 2012

# Background situation

## Peak fit with 0V signal

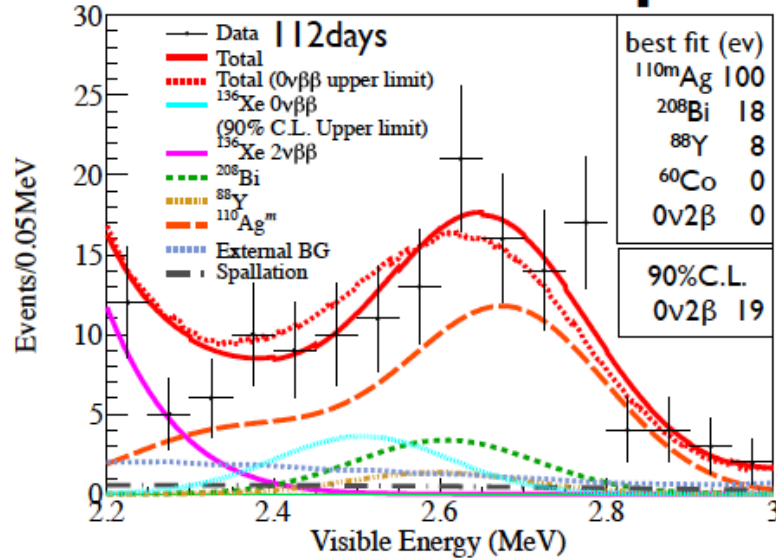


Peak position is different from that of expected 0V.

0V only is rejected at more than  $8\sigma$  level.



# Limit on the $0\nu 2\beta$ half life

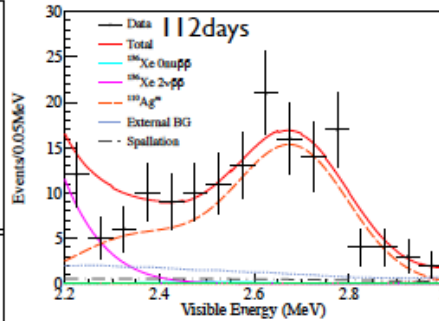


simultaneous fit and 90% CL upper limit for  $0\nu 2\beta$

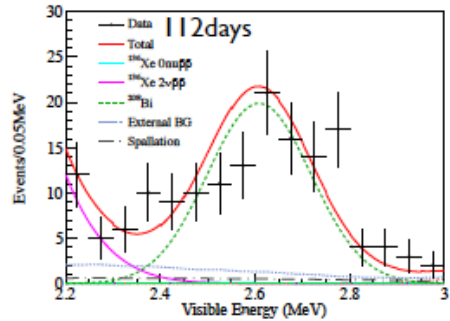
( $\chi^2$  at 2.2~3.0MeV)

	$\chi^2$ 112days	
simul. fit	11.6	
$0\nu + {}^{110m}\text{Ag}$	13.1	
$0\nu + {}^{208}\text{Bi}$	22.7	$\triangle$
$0\nu + {}^{88}\text{Y}$	22.2	$\triangle$
$0\nu + {}^{60}\text{Co}$	82.9	$\times$
$0\nu$ only	85.0	$\times$

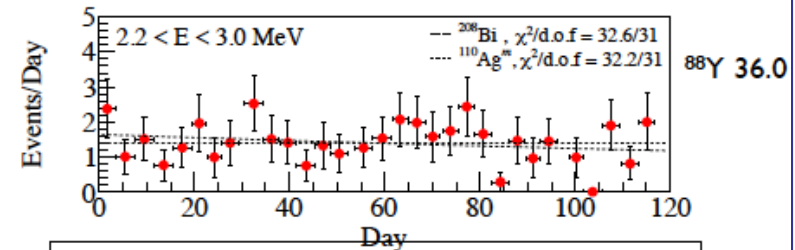
BG is likely to be  ${}^{110m}\text{Ag}$



110mAg fits well



208Bi doesn't fit well



stable in time, but no strong discrimination yet

$T^{0\nu}_{1/2} > 5.7 \times 10^{24}$  years at 90% C.L. (78days)

factor 5 improvement from DAMA

$T^{0\nu}_{1/2} > 6.2 \times 10^{24}$  years (KL-Zen 112days)

(ref. current best is  $16 \times 10^{24}$  years from EXO-200)



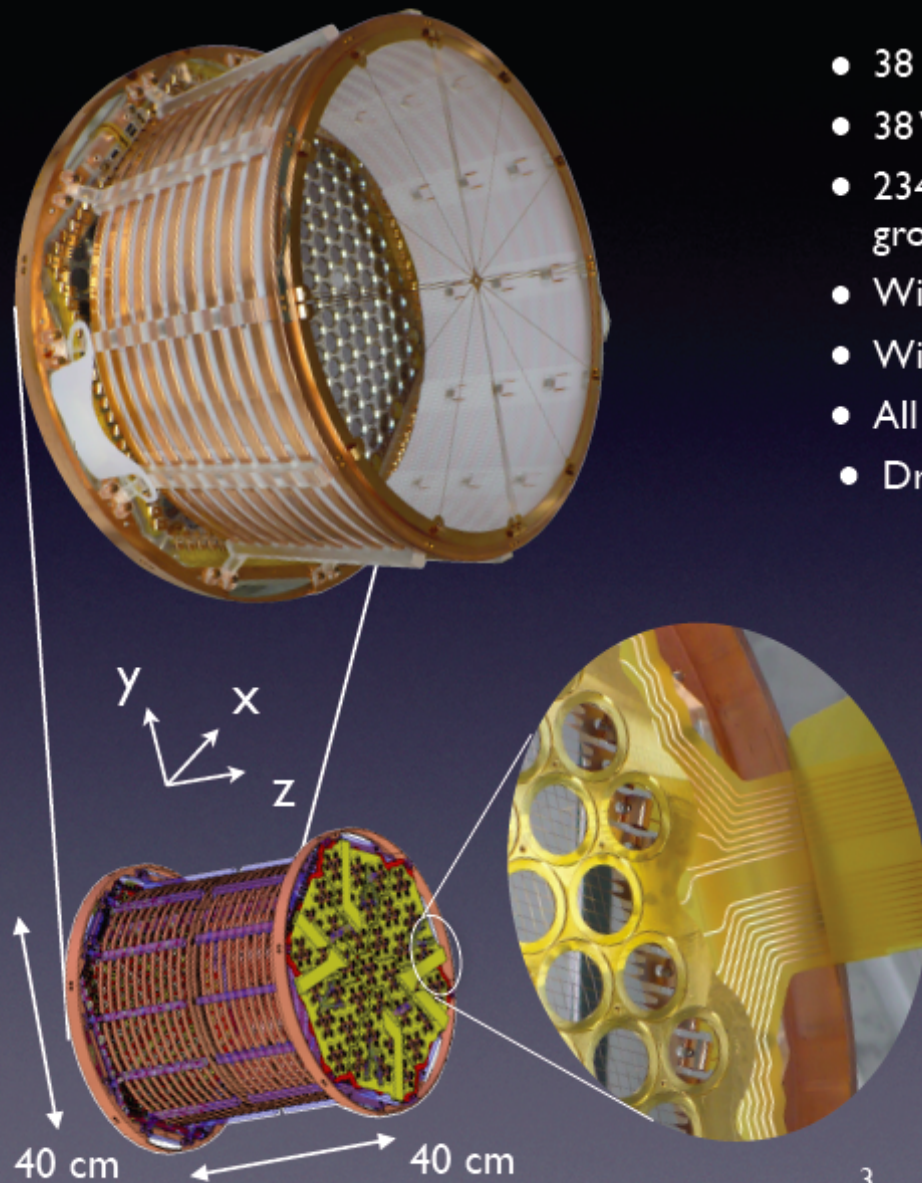
(R)QRPA (CCM SRC)  
Phys.Rev.C79,055501(2009)

$\langle m_{\beta\beta} \rangle < 0.26 \sim 0.54$  eV @90% C.L.<sub>15</sub>

From K. Inoue, Neutrino 2012

N.B.: limit depends on correctness of background model

# The EXO-200 TPC



Two almost identical halves reading **ionization** and 178 nm **scintillation**, each with:

- 38 U triplet wire channels (charge)
- 38 V triplet wire channels, crossed at  $60^\circ$  (induction)
- 234 large area avalanche photodiodes (APDs, light in groups of 7)
- Wire pitch 3 mm (9 mm per channel)
- Wire planes 6 mm apart and 6 mm from APD plane
- All signals digitized at 1 MS/s,  $\pm 1024S$  around trigger
- Drift field 376 V/cm

- Field shaping rings: copper
- Supports: acrylic
- Light reflectors/diffusers: Teflon
- APD support plane: copper; Au (Al) coated for contact (light reflection)
- Central cathode, U+V wires: photo-etched phosphor bronze
- Flex cables for bias/readout: copper on kapton, no glue

Comprehensive material screening program

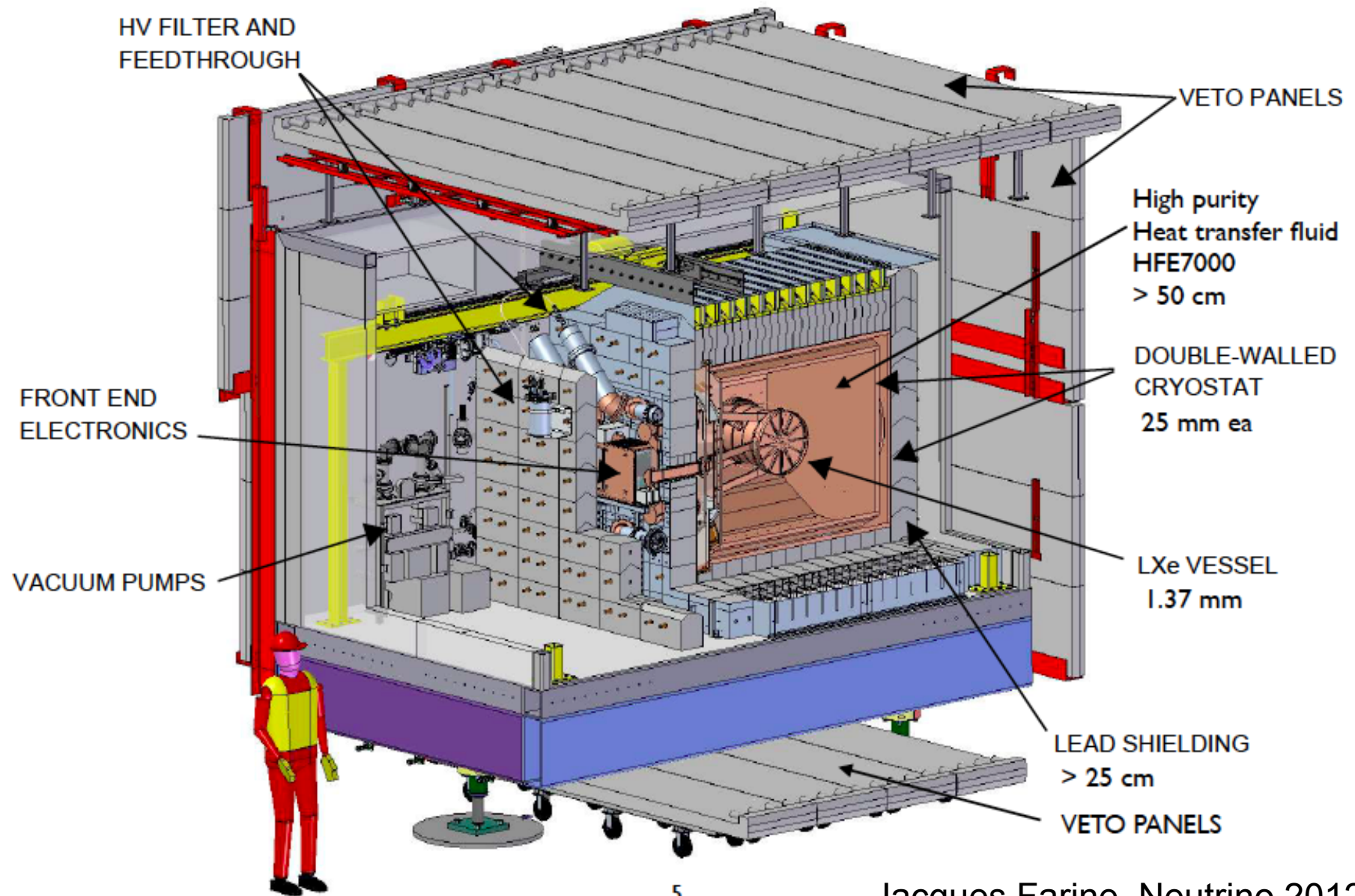
**Goal: 40 cnts/2y in  $0\nu\beta\beta \pm 2\sigma$  ROI, 140 kg LXe**

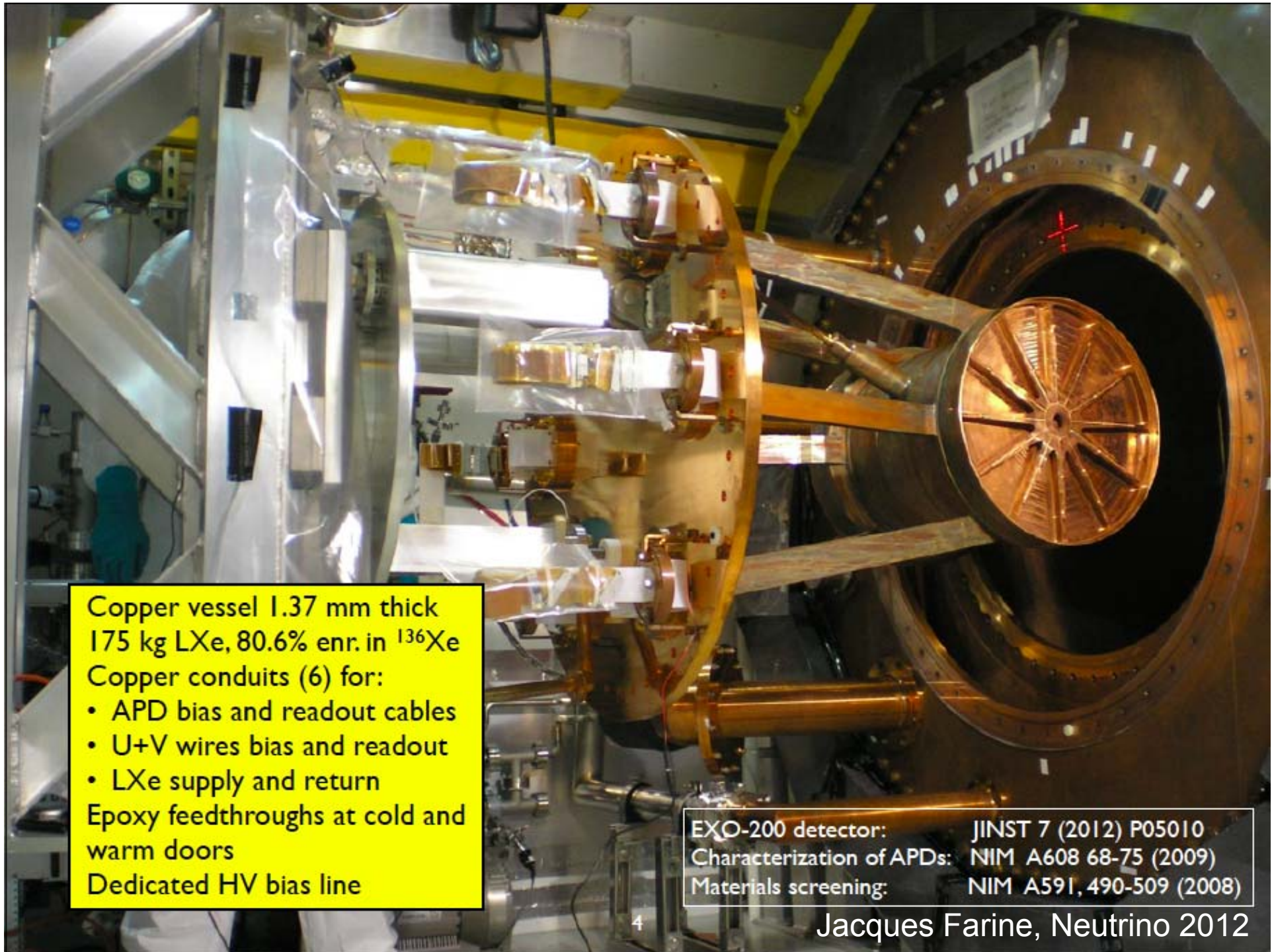
Jacques Farine, Neutrino 2012

3



# The EXO-200 Detector





Copper vessel 1.37 mm thick  
175 kg LXe, 80.6% enr. in  $^{136}\text{Xe}$   
Copper conduits (6) for:

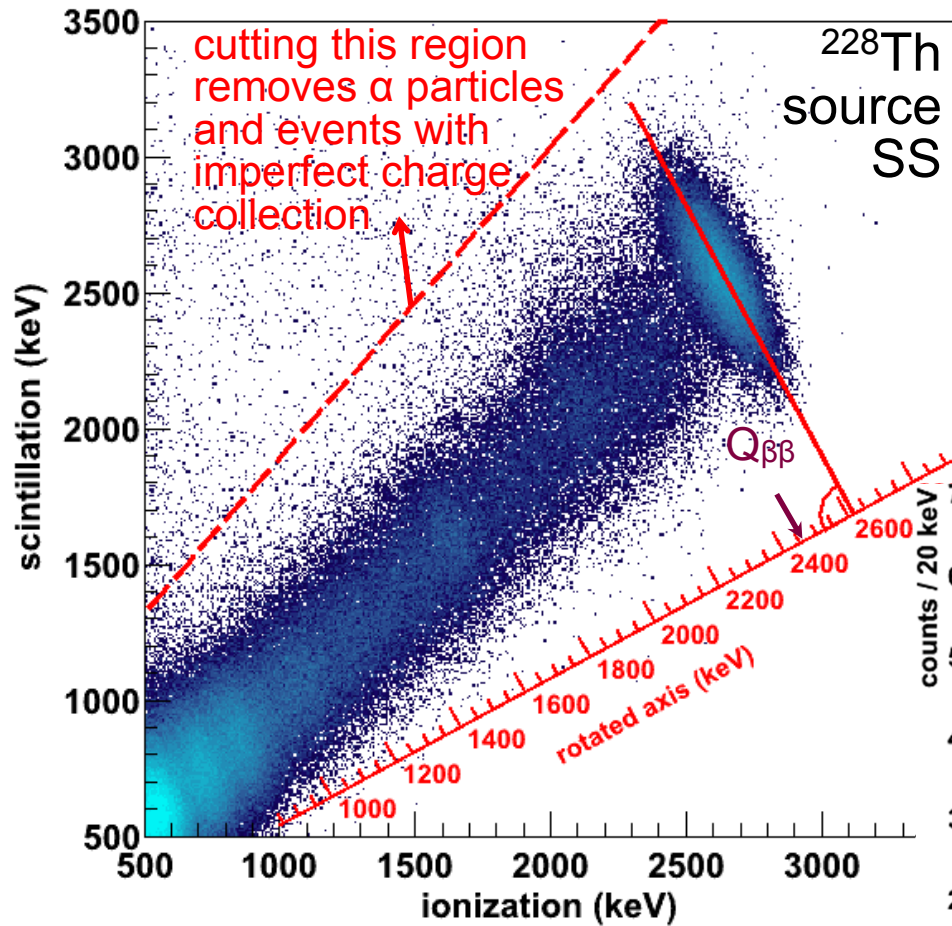
- APD bias and readout cables
- U+V wires bias and readout
- LXe supply and return

Epoxy feedthroughs at cold and warm doors  
Dedicated HV bias line

EXO-200 detector: JINST 7 (2012) P05010  
Characterization of APDs: NIM A608 68-75 (2009)  
Materials screening: NIM A591, 490-509 (2008)



# Combining Ionization and Scintillation

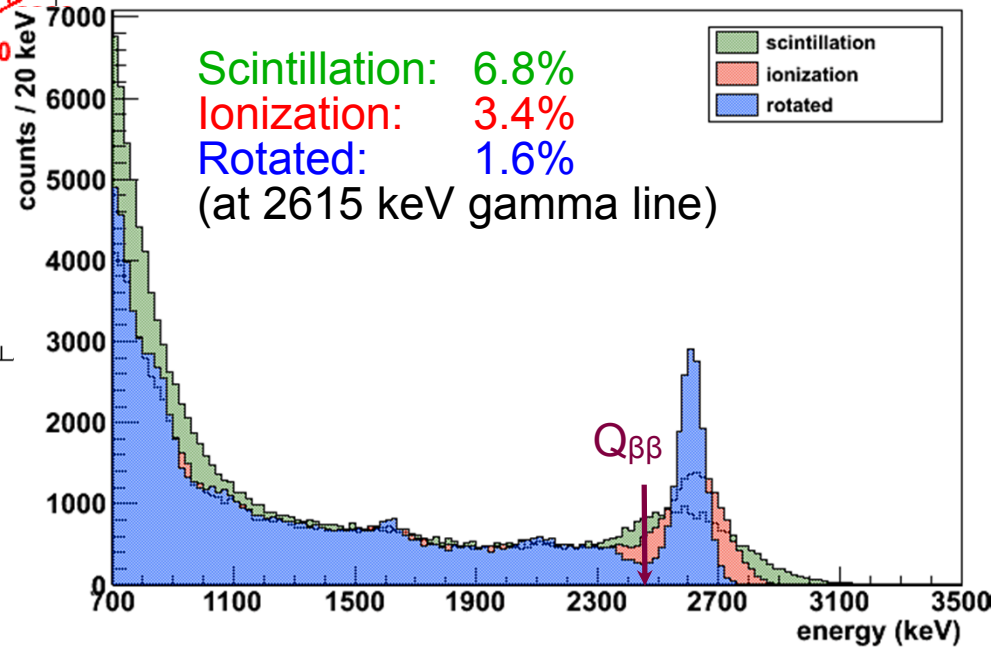


Rotation angle chosen to optimize energy resolution at 2615 keV

Properties of xenon cause increased scintillation to be associated with decreased ionization (and vice-versa)

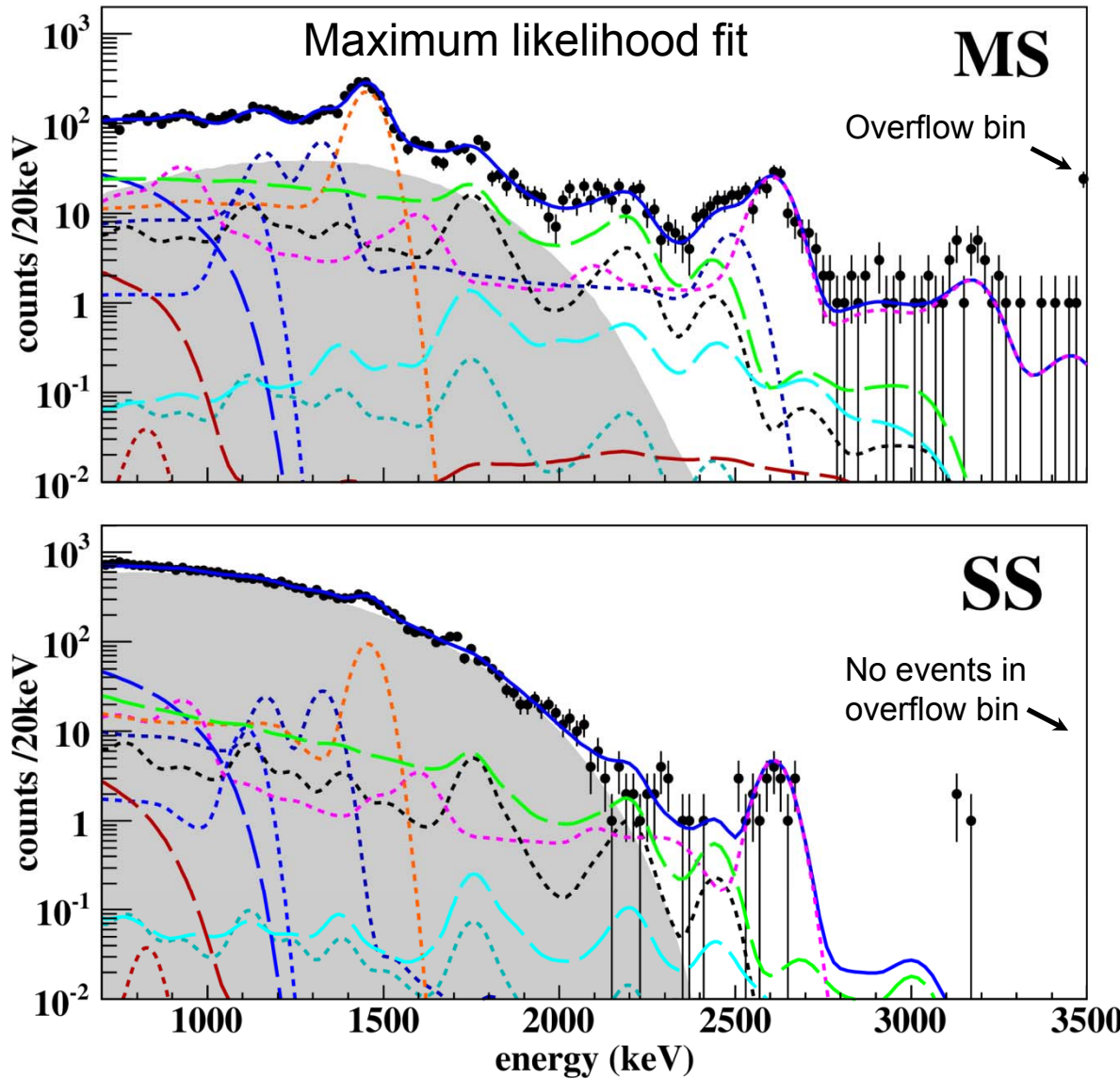
E. Conti et al. Phys. Rev. B 68 (2003) 054201

Use projection onto a rotated axis to determine event energy

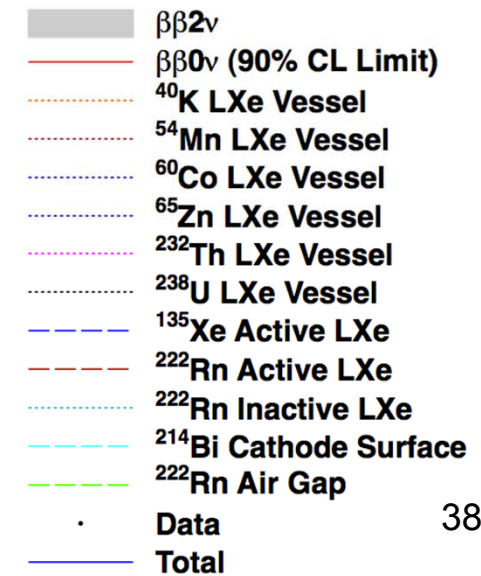


Jacques Farine, Neutrino 2012 <sup>37</sup>

# Low Background Spectrum



- Trigger fully efficient above 700 keV
- Low background run livetime: **120.7 days**
- Active mass: **98.5 kg LXe (79.4 kg <sup>136</sup>LXe)**
- Exposure: **32.5 kg-yr**
- Total dead time (vetos): 8.6%
- Various background PDFs fitted along with  $2\nu\beta\beta$  and  $0\nu\beta\beta$  PDFs



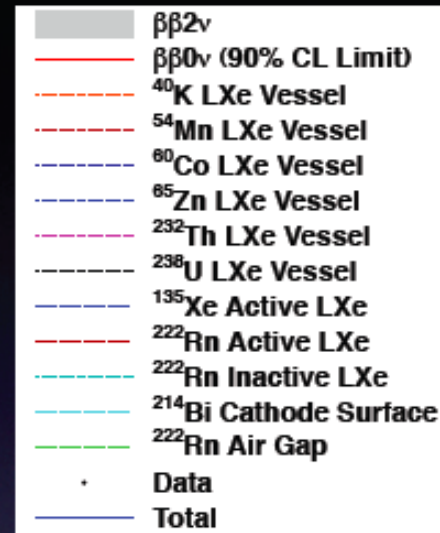
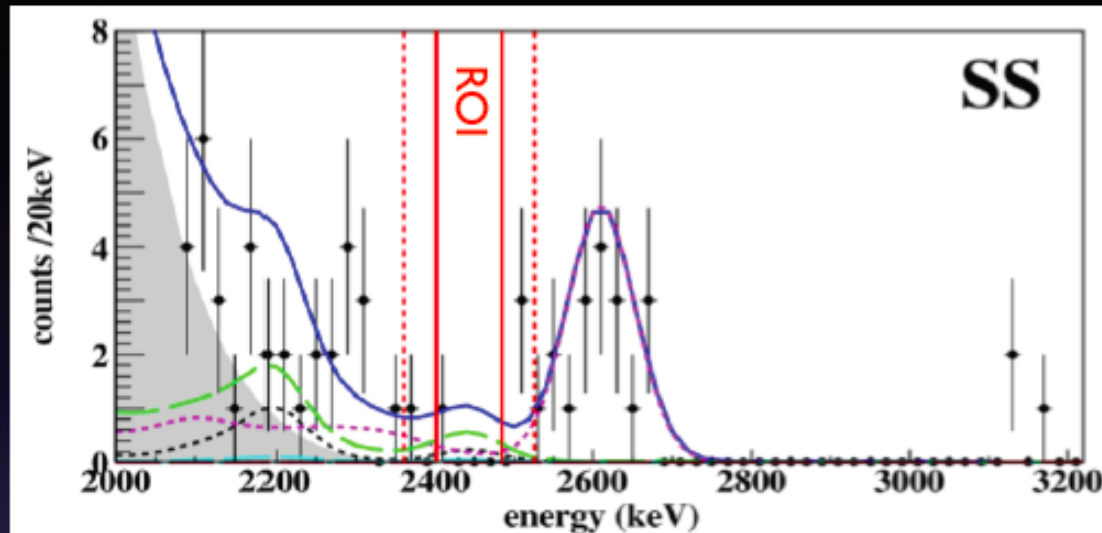
$T_{1/2}^{2\nu\beta\beta} (^{136}\text{Xe}) = (2.23 \pm 0.017 \text{ stat} \pm 0.22 \text{ sys}) \cdot 10^{21} \text{ yr}$

(In agreement with previously reported measurements)

Neutrino-less double beta decay

From J. Farine, Neutrino 2012, 30-31 October 2012, S. SCHOTT, TU MÜNCHEN

# Background counts in $\pm 1, 2 \sigma$ ROI



	Expected events from fit			
	$\pm 1 \sigma$		$\pm 2 \sigma$	
$^{222}\text{Rn}$ in cryostat air-gap	1.9	$\pm 0.2$	2.9	$\pm 0.3$
$^{238}\text{U}$ in LXe Vessel	0.9	$\pm 0.2$	1.3	$\pm 0.3$
$^{232}\text{Th}$ in LXe Vessel	0.9	$\pm 0.1$	2.9	$\pm 0.3$
$^{214}\text{Bi}$ on Cathode	0.2	$\pm 0.01$	0.3	$\pm 0.02$
All Others	~0.2		~0.2	
<b>Total</b>	<b>4.1</b>	<b><math>\pm 0.3</math></b>	<b>7.5</b>	<b><math>\pm 0.5</math></b>
Observed	1		5	
Background index $b$ ( $\text{kg}^{-1}\text{yr}^{-1}\text{keV}^{-1}$ )	$1.5 \cdot 10^{-3} \pm 0.1$		$1.4 \cdot 10^{-3} \pm 0.1$	

EXO-200 goal (slide 3):

40 cnts/2y in  $\pm 2\sigma$  ROI,  
140 kg LXe

In this data 120 days, 98.5 kg, this would be: 4.6

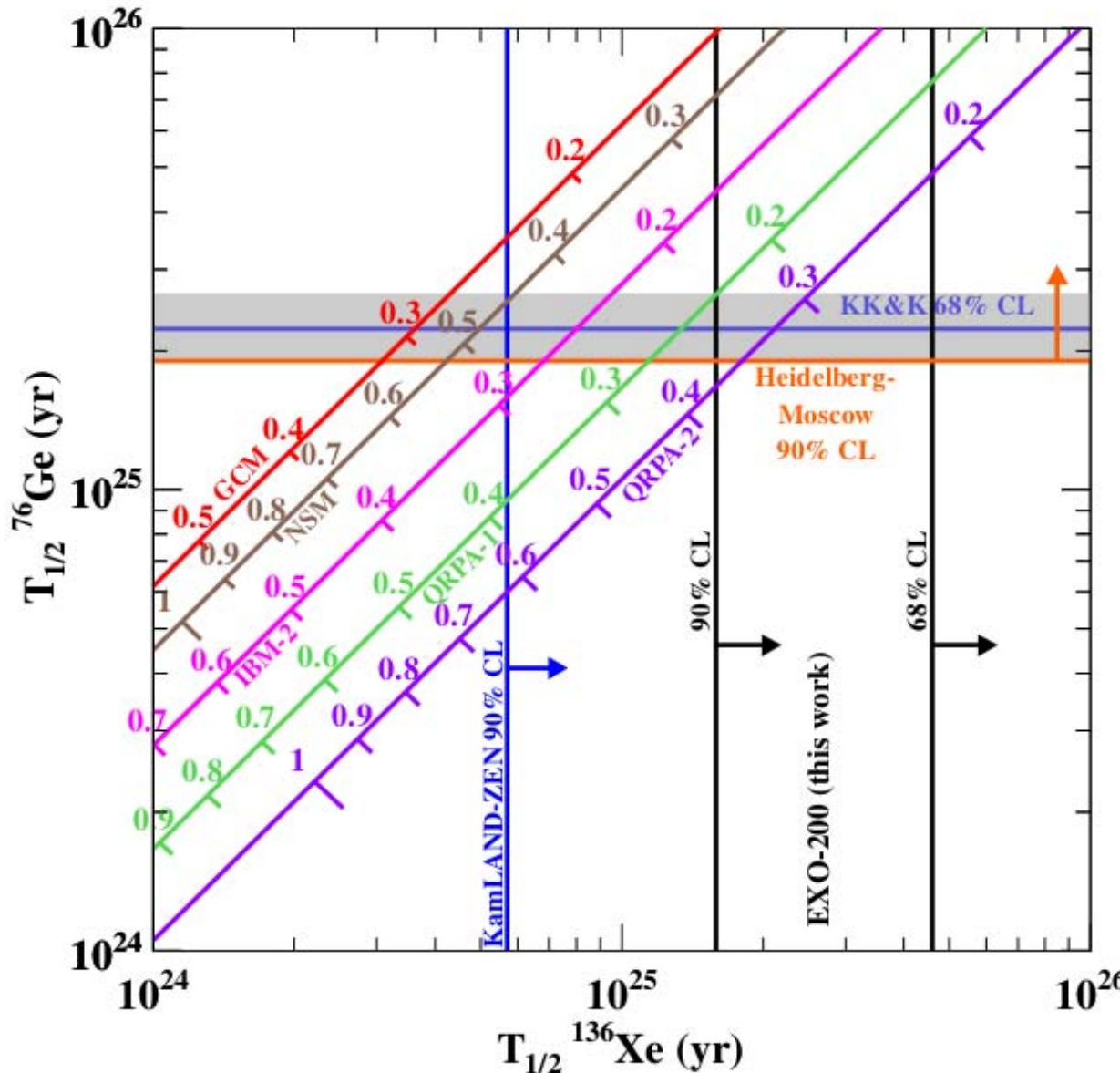
Expected from the fit: 7.5

Observed: 5

**Background within expectation**



# Limits on $T_{1/2}^{0\nu\beta\beta}$ and $\langle m_{\beta\beta} \rangle$



90% C.L. limit compared with Recent  $^{136}\text{Xe}$  constraints (KamLAND-ZEN) >2.5 factor improvement.

$T_{1/2}^{0\nu\beta\beta} > 1.6 \cdot 10^{25} \text{ yr}$   
 $\langle m_{\beta\beta} \rangle < 140 - 380 \text{ mV}$   
**(90% C.L.)**

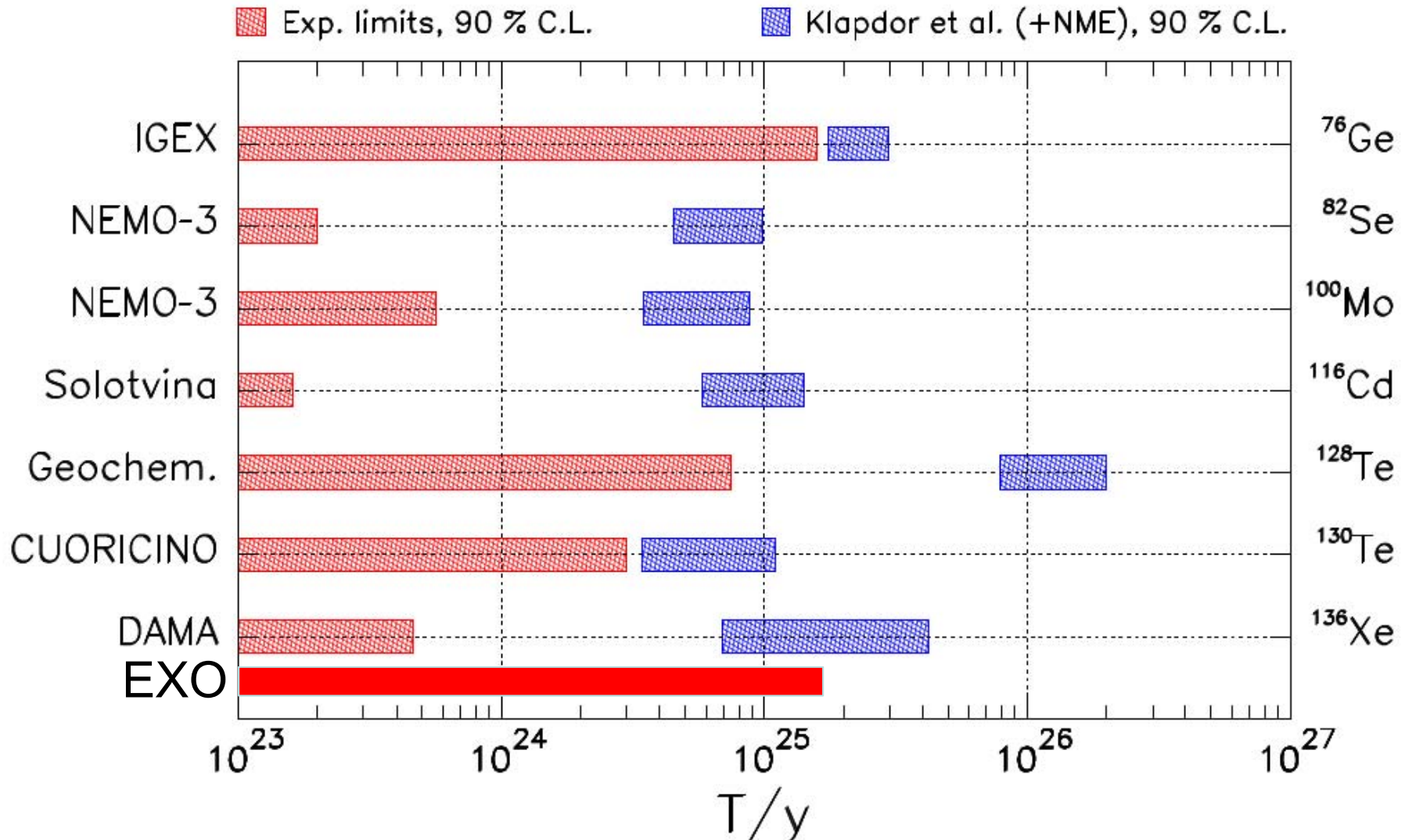
[arXiv:1205.5608]

Tension with discovery claim in Ge.

KamLAND-Zen Collaboration  
 Phys. Rev. C 85 (2012) 045504  
 [H.V. Klapdor-Kleingrothaus et al.  
 Eur. Phys. J. A12 (2001) 147]  
 [H.V. Klapdor-Kleingrothaus and I.V. Krivosheina  
 Mod. Phys. Lett., A21 (2006) 1547]

# Comparison of upper limits (90 % C.L.) with claim [16] for QRPA NME

A.Faessler, G.L. Fogli, E. Lisi, V. Rodin, A.M. Rotunno, F. Simkovic, PhysRevD.79.053001  
 arXiv:0810.5733v2 (EXO result included by 'hand')



[16] H. V. Klapdor-Kleingrothaus and I. V. Krivosheina, "The Evidence For The Observation Of  $0\nu\beta\beta$  Decay: The Identification Of  $0\nu\beta\beta$  Events From The Full Spectra," Mod. Phys. Lett. A **21**, 1547 (2006).  
neutrino-less double beta decay

**Table 2** Bayesian posterior probabilities  $p(\bar{H})$  using EXO-200 data for the hypothesis that the  $0\nu\beta\beta$  signal of Heidelberg-Moscow is correct. Probabilities are given for different matrix element calculations and for the  $\pm 1\sigma$  and  $\pm 2\sigma$  energy windows.

method	expected signal events	$p(\bar{H})$ in %	expected signal events	$p(\bar{H})$ in %
	in $\pm 1\sigma$ window		in $\pm 2\sigma$ window	
QRPA max	$4.4 \pm 1.1$	4	$6.1 \pm 1.5$	6
QRPA min	$2.8 \pm 0.7$	11	$3.9 \pm 0.9$	16
ISM	$10.6 \pm 2.5$	0.1	$14.8 \pm 3.5$	0.2
GCM	$14.3 \pm 3.4$	0.03	$19.9 \pm 4.8$	0.05
pnQRPA	$6.3 \pm 1.5$	1	$8.8 \pm 2.1$	2
IBM	$6.1 \pm 1.5$	1	$8.6 \pm 2.1$	2

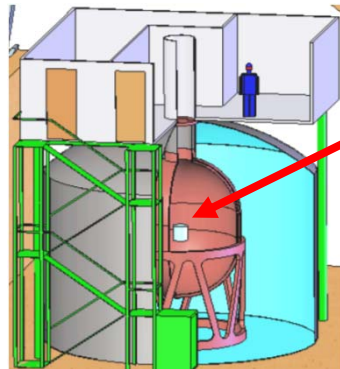
N.B. comparison with HdM claim ( $28 \pm 6.86$ ) cts in 71.7 kg yr

B. Schwingenheuer, Annalen der Physik, August 22, 2012

# Two $^{76}\text{Ge}$ projects:



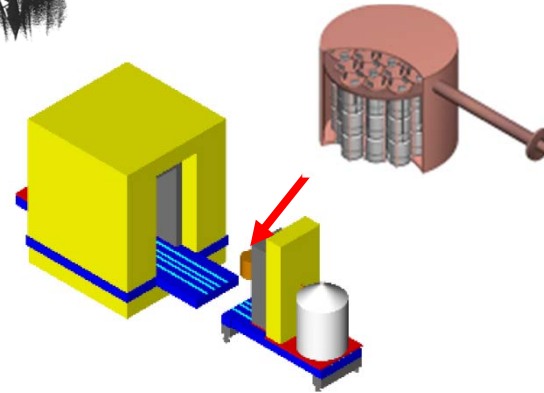
## GERDA



- 'Bare'  $^{76}\text{Ge}$  array in liquid argon
- Shield: high-purity liquid Argon /  $\text{H}_2\text{O}$
- Phase I: 18 kg (HdM/IGEX) / 15 kg nat.
- Phase II: add  $\sim 20$  kg new enr. detectors; total  $\sim 40$  kg



## Majorana



- Array(s) of  $^{76}\text{Ge}$  housed in high-purity electroformed copper cryostat
- Shield: electroformed copper / lead
- Initial phase: R&D demonstrator module: Total  $\sim 60$  kg (30 kg enr.)

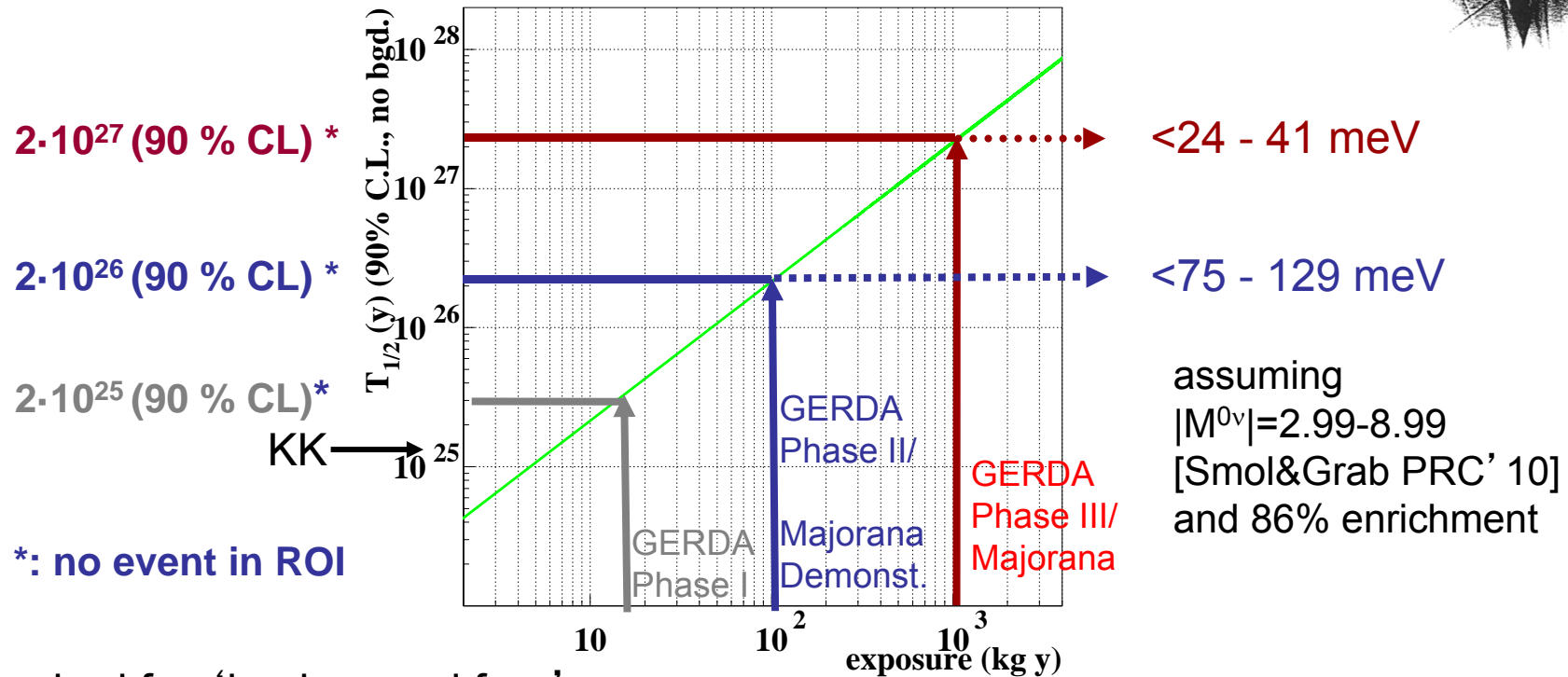
**Physics goals:** degenerate mass range  
**Technology:** study of bgds. and exp. techniques

**LoI**

- open exchange of knowledge & technologies (e.g. MaGe MC)
- intention to merge for O(1 ton) exp. (inv. Hierarchy) selecting the best technologies tested in GERDA and Majorana



# Phases and physics reach



required for 'background free' exp. with  $\Delta E \sim 3.3$  keV (FWHM):  $O(10^{-3})$   $O(10^{-4})$  counts/(kg·y·keV)

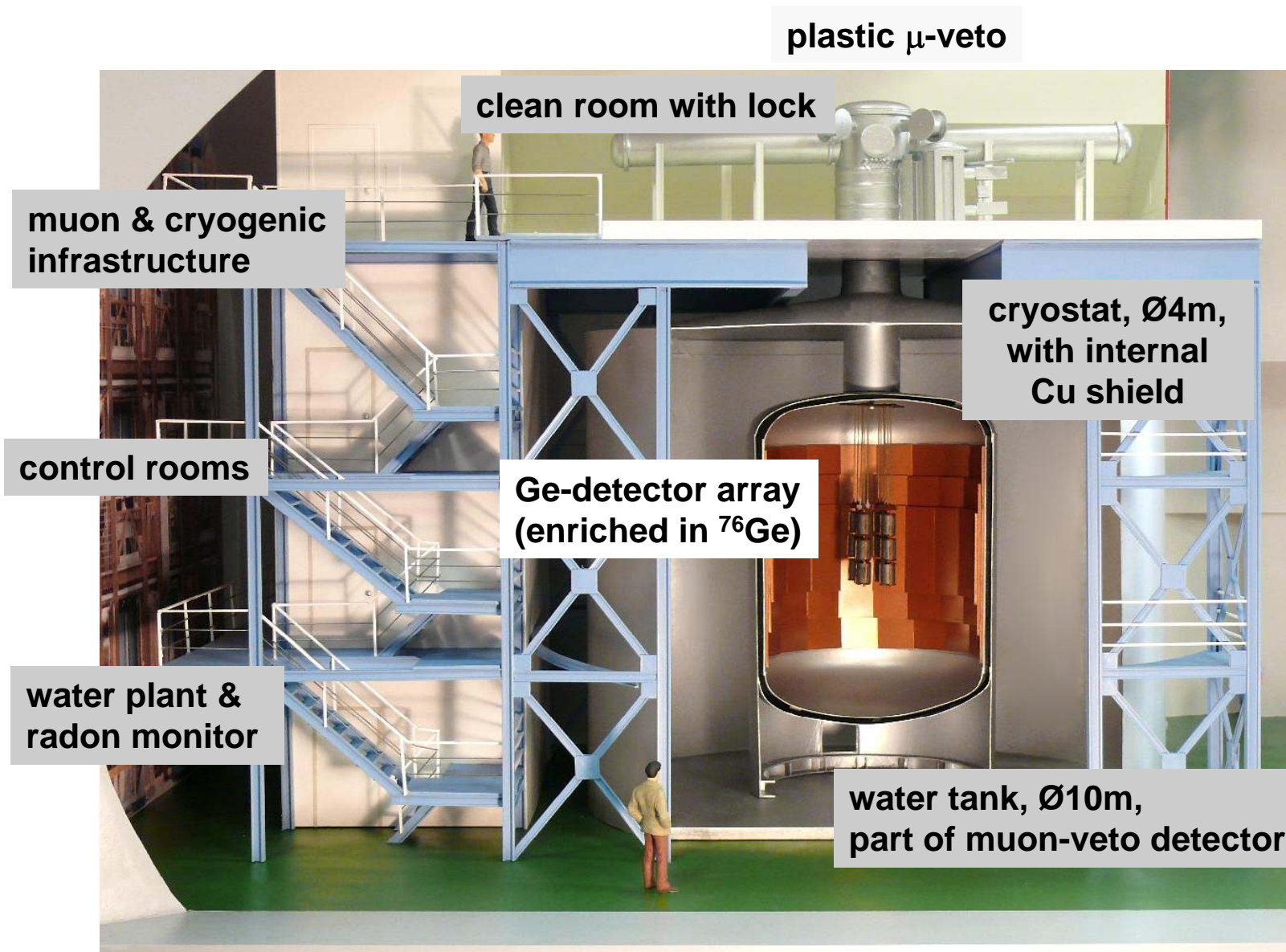
## Background requirement for GERDA/Majorana:

- $\Rightarrow$  Background reduction by factor  $10^2 - 10^3$  required w.r. to precursor exps.
- $\Rightarrow$  Degenerate mass scale  $O(10^2 \text{ kg}\cdot\text{y}) \Rightarrow$  Inverted mass scale  $O(10^3 \text{ kg}\cdot\text{y})$





# GERDA @ LNGS







# GERDA construction phase 2008-2010



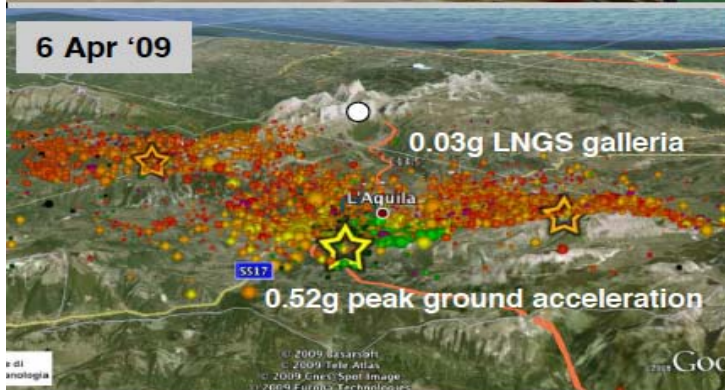
6 Mar '08



5 May '08



29 feb '09



Aug '09

view inside water tank



active cooling system inst.

18 Jul '09



18 May '10

glove box

Neutrino-less double beta decay



inauguration  
9 Nov 2010

Cryostat filled since  
December 2009





# GERDA phase I detectors

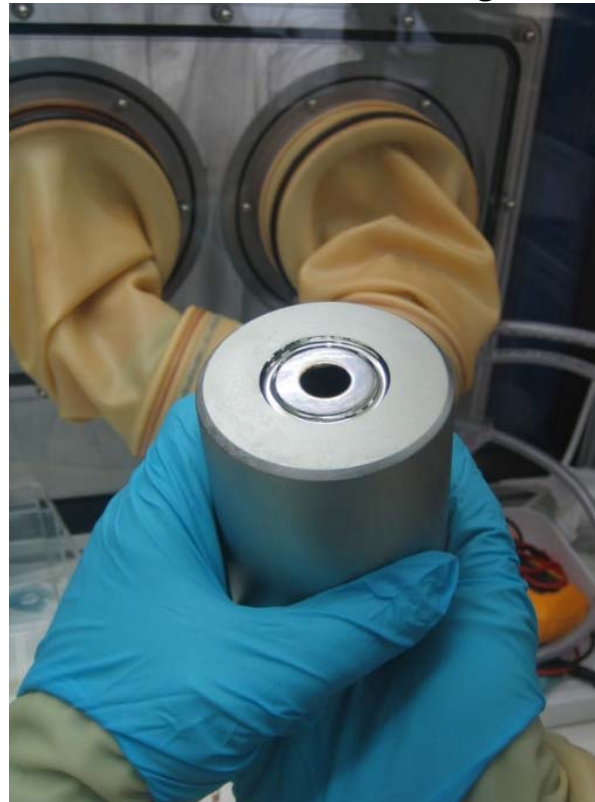


- All diodes reprocessed and optimized for LAr
- Well tested procedure for detector handling
- Long term stability in LAr established
- Energy resolution in LAr:  $\sim 2.5$  keV (FWHM) @1.3 MeV

Neutrino-less double beta decay

## 8 diodes (from HdM, IGEX):

- Enriched 86% in  $^{76}\text{Ge}$
- Total mass 17.66 kg



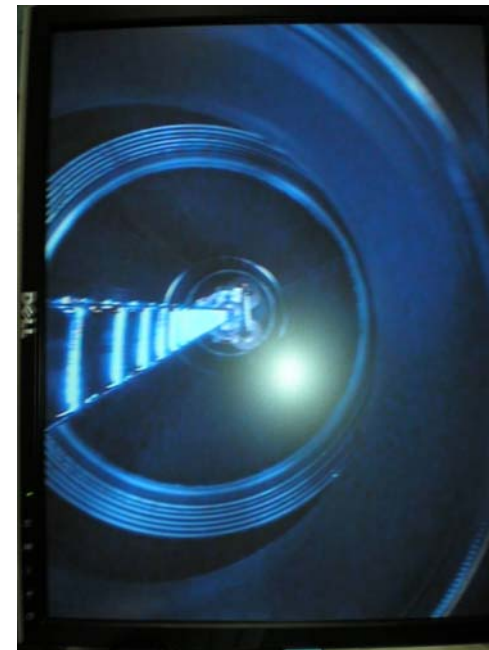
## 6 diodes from Genius-TF:

- $\text{natGe}$
- Total mass: 15.60 kg



# string assembly for phase I run

started Nov  
2011

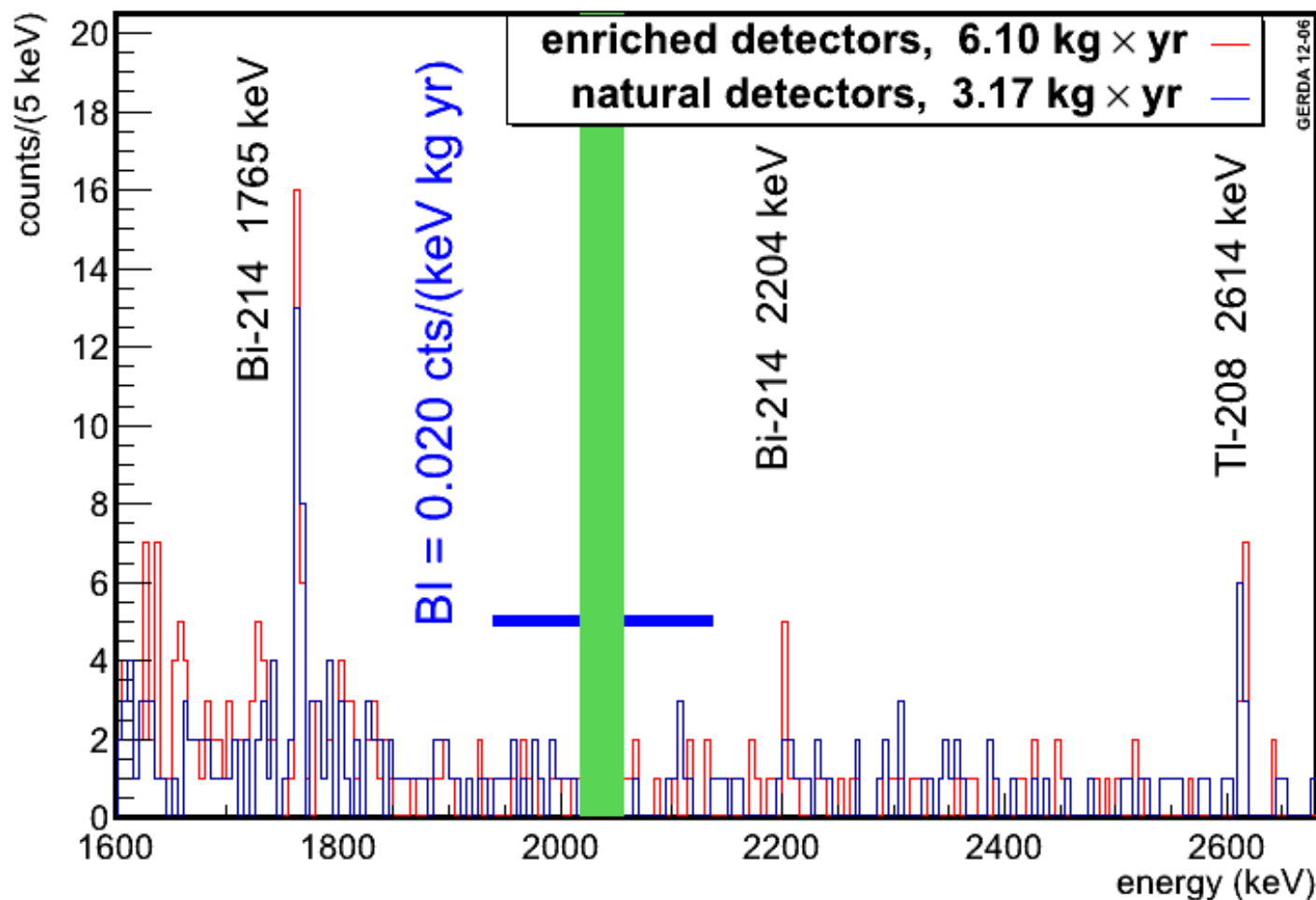


8 refurbished enriched diodes from HdM & IGEX

- 86% isotopically enriched in Ge-76
- 17.66 kg total mass
- plus 1 natural Ge diode from GTF

2 diodes shut off because leakage current high:

- total enriched enriched detector mass 14.6 kg

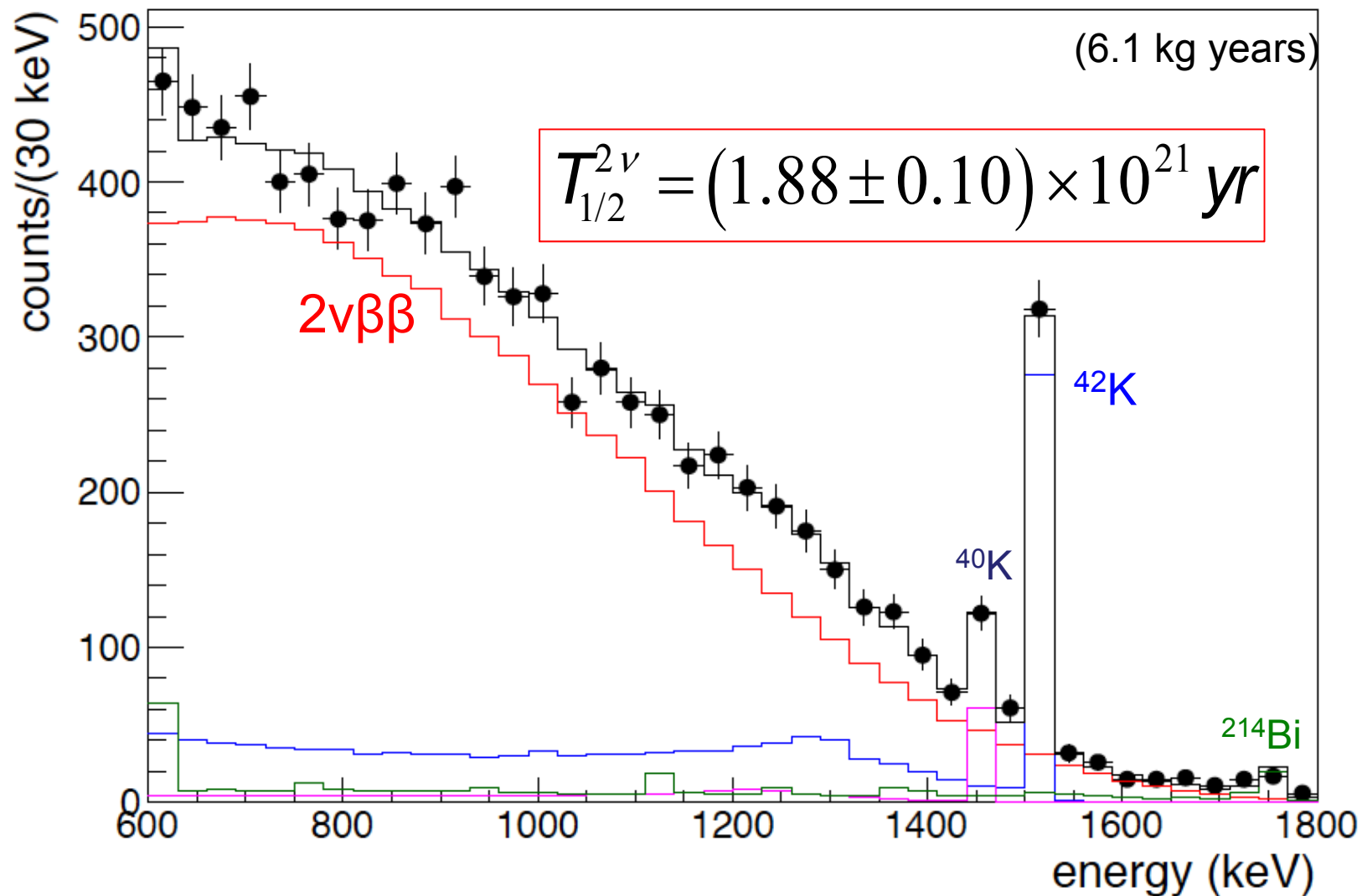


Background index usually evaluated in  $(Q_{\beta\beta} \pm 100)$  keV  
(excluding blinded region of  $(Q_{\beta\beta} \pm 20)$  keV)



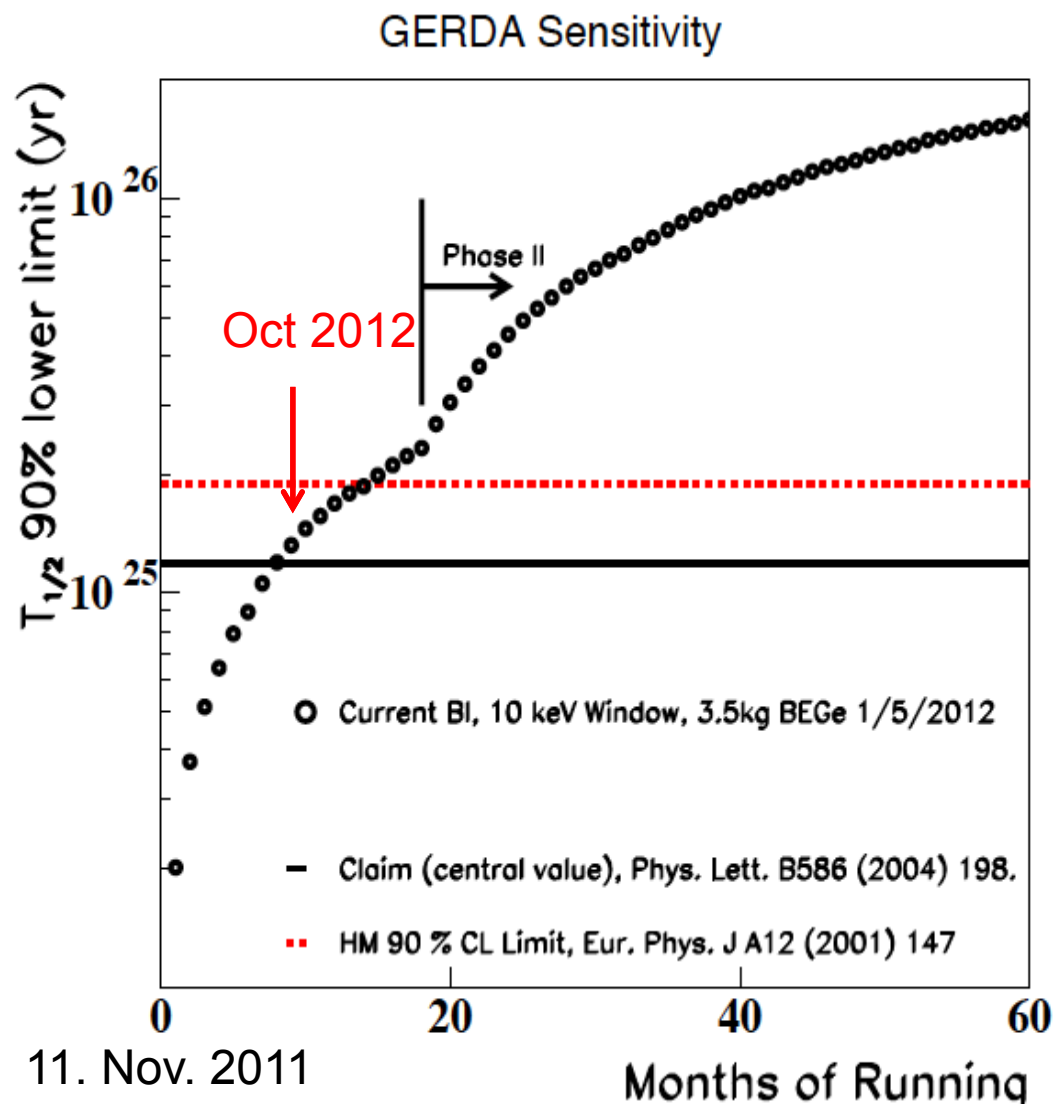
# preliminary results

$2\nu\beta\beta$ -  
spectrum



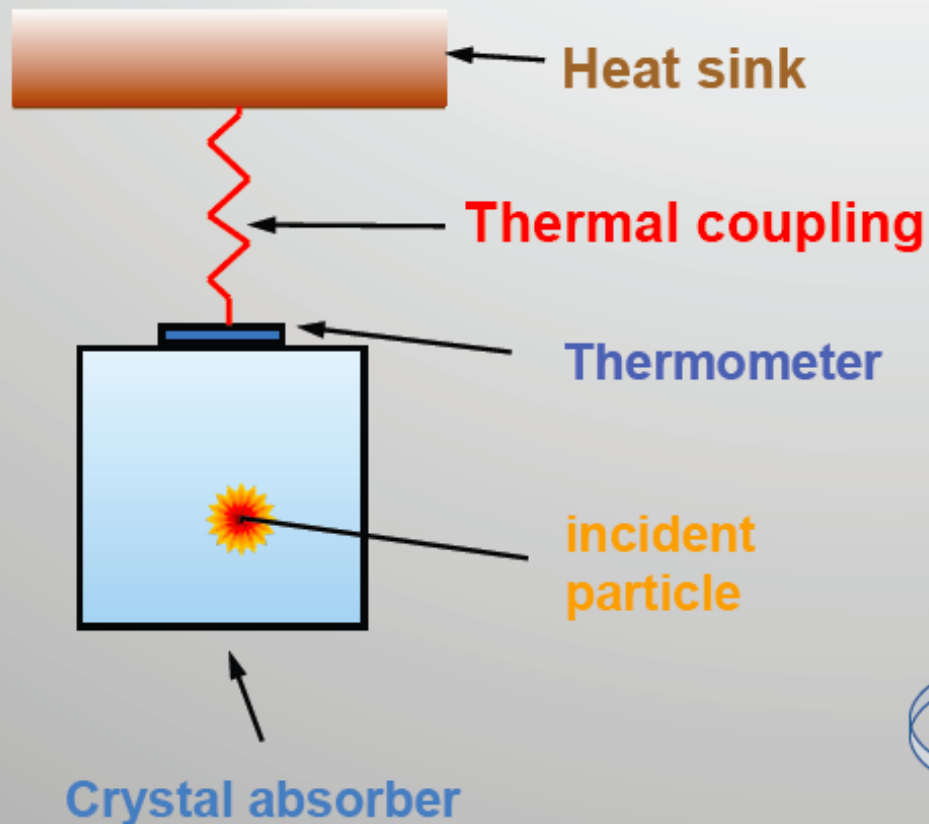


# GERDA sensitivity



# Bolometric technique

The working principle is very simple:



This technique measures **all** the energy deposited by a particle in form of increase of temperature in the absorber

Absorber  $\equiv$  DBD source

$$\text{Signal: } \Delta T = E/C$$
$$\text{Time constant} = C/G$$

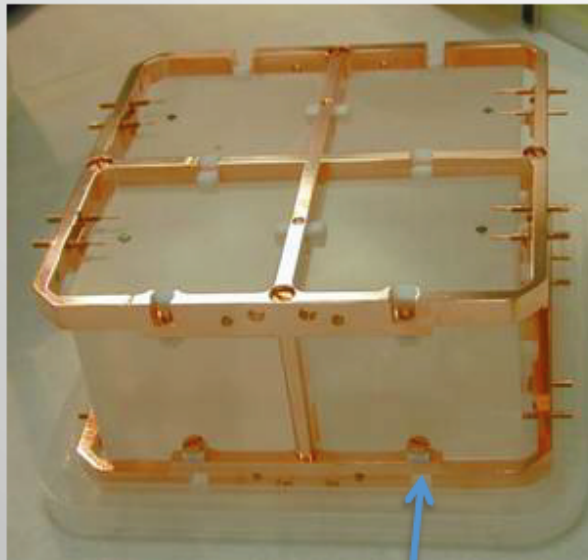
Low heat capacity

- **Low temperatures** ( $\sim 10\text{mK}$ )
- Dielectric diamagnetic materials

From M. Pedretti, Neutrino 2012



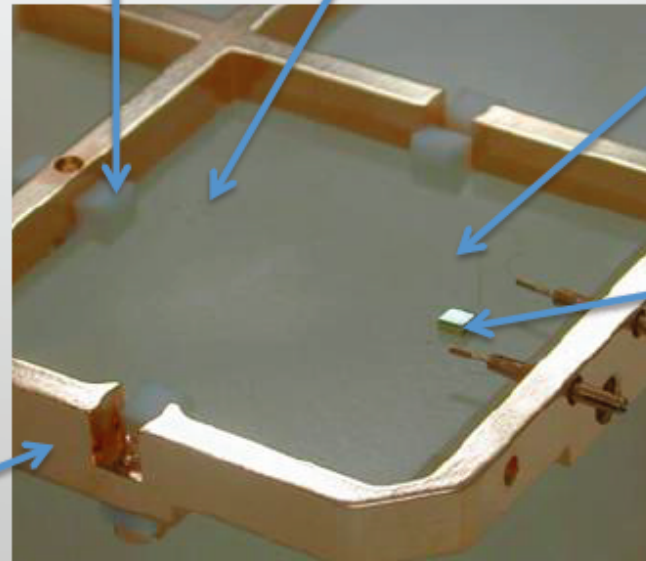
# TeO<sub>2</sub> Bolometers



Copper holder

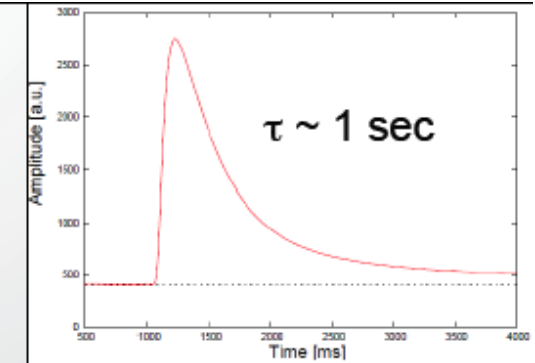
PTFE pieces

TeO<sub>2</sub> crystal



25 μm gold wire connection

Neutron Transmutation Doped Ge sensor



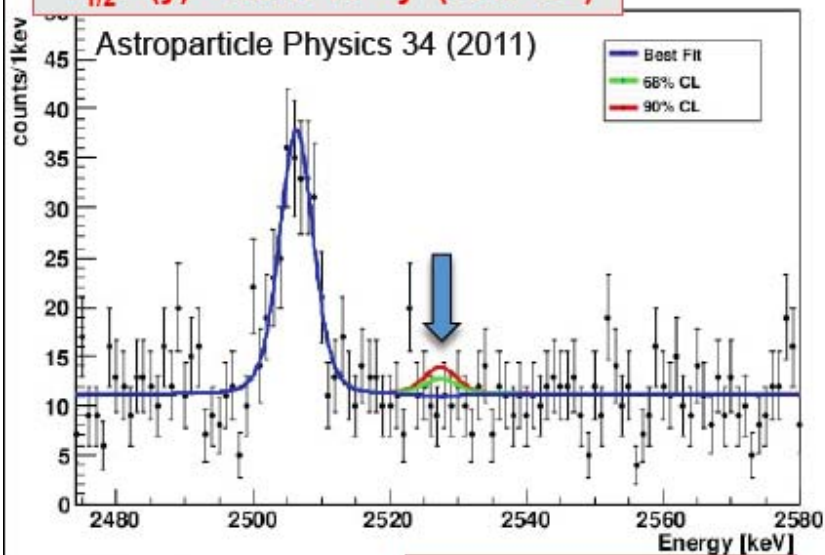
From M. Pedretti, Neutrino 2012

# Bolometric $0\nu\text{DBD}$ experiment evolution

$$\Delta E = 6.2 \pm 2.5 \text{ keV} \quad (\sim 0.3\% \text{ FWHM})$$

$$\text{Bkg} = 0.169 \pm 0.006 \text{ c/keV/kg/y}$$

$$T_{1/2}^{0\nu} (\text{y}) > 2.8 \times 10^{24} \text{ y} \quad (90\% \text{ CL})$$

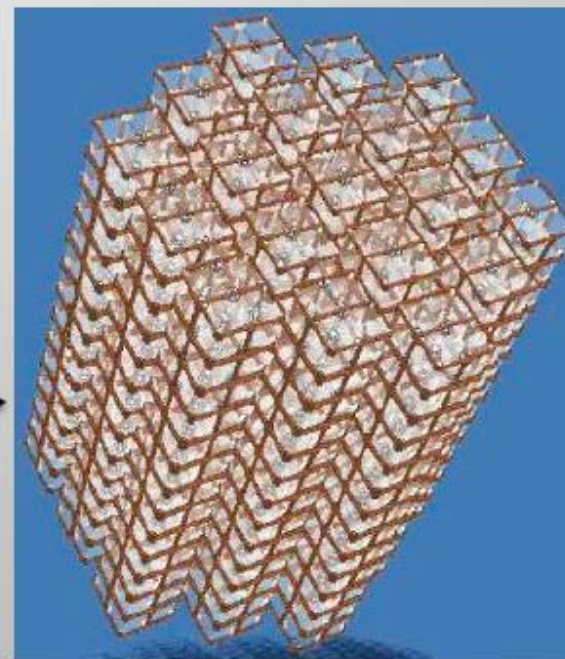


$$\langle M_{bb} \rangle < 0.3 - 0.7 \text{ eV}$$

CUORICINO  
40 kg  
(2003-2008)

CUORE-0  
(2012)

CUORE  
1 ton  
(~2014)

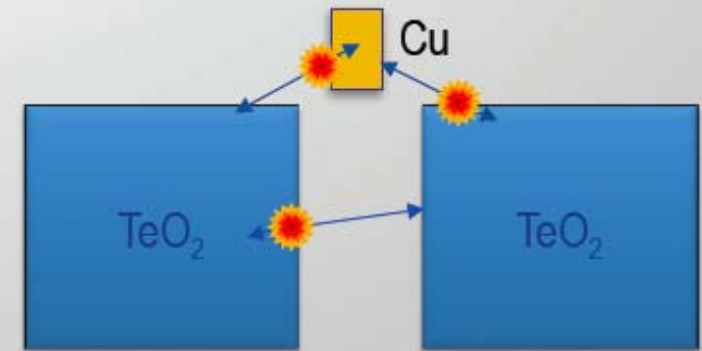
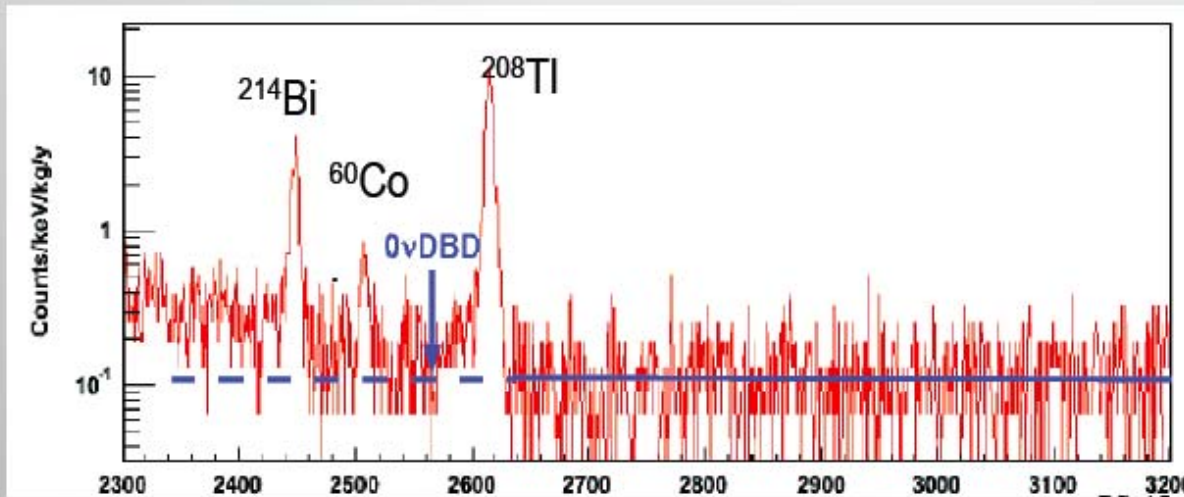


From M. Pedretti, Neutrino 2012

# CUORICINO lesson: background

Sensitivity of current generation bolometric DBD experiment is limited by bkg.

MC: the background in CUORICINO is due to degraded alpha particles which release only part of their energy in the detector (surface contamination)



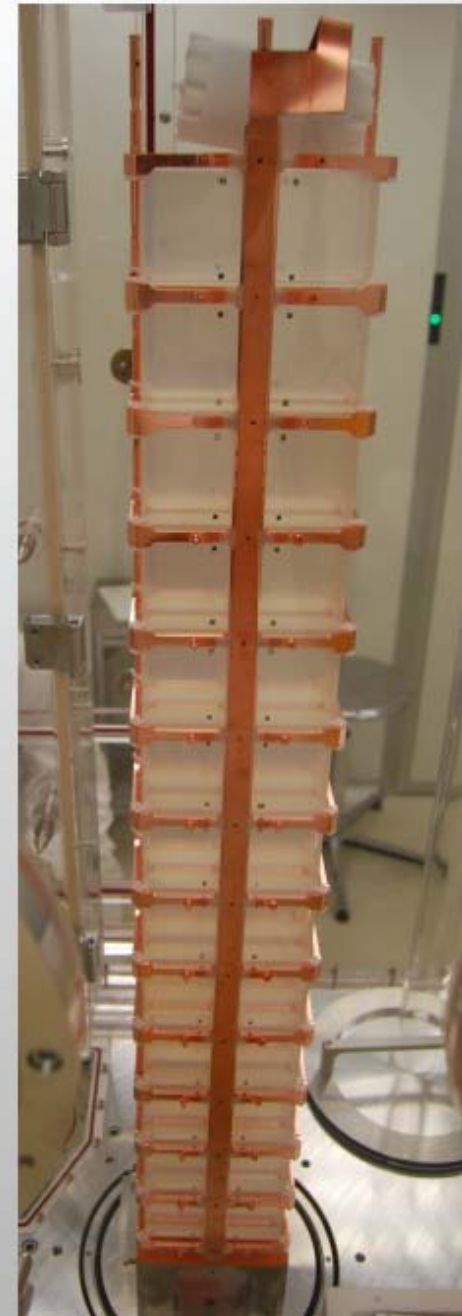
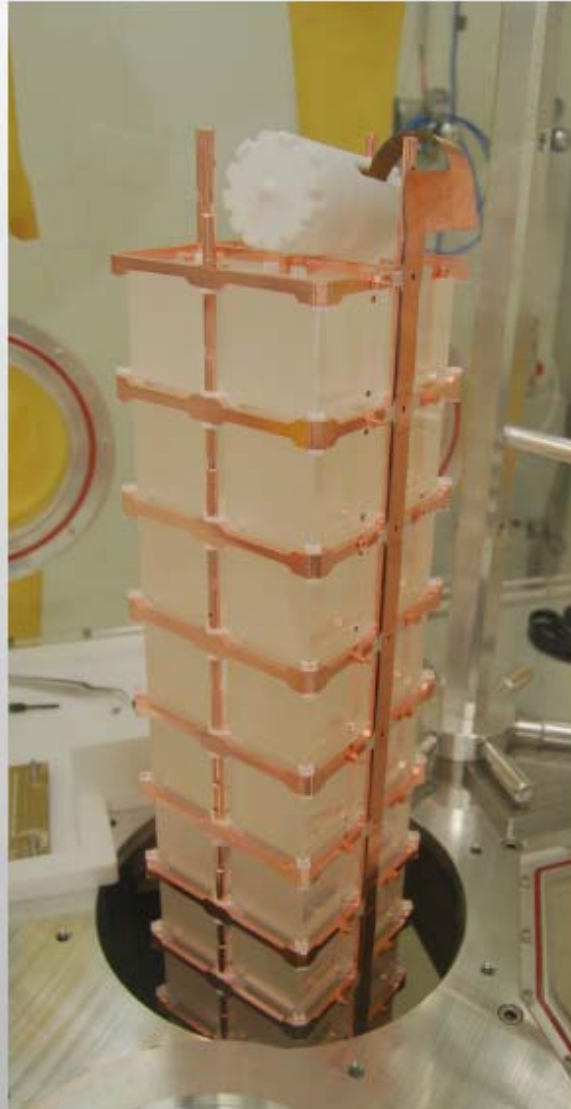
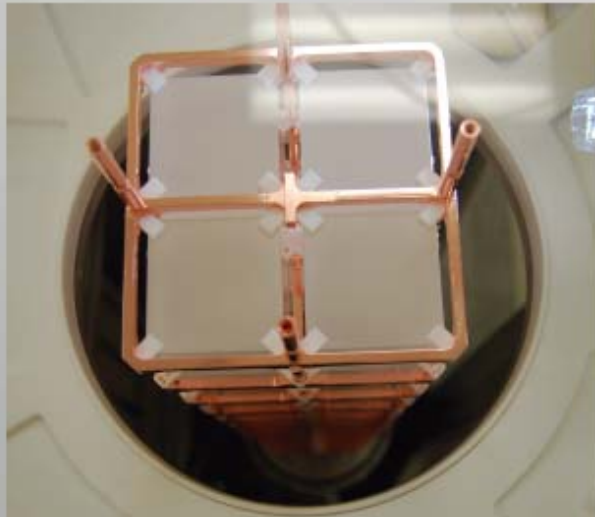
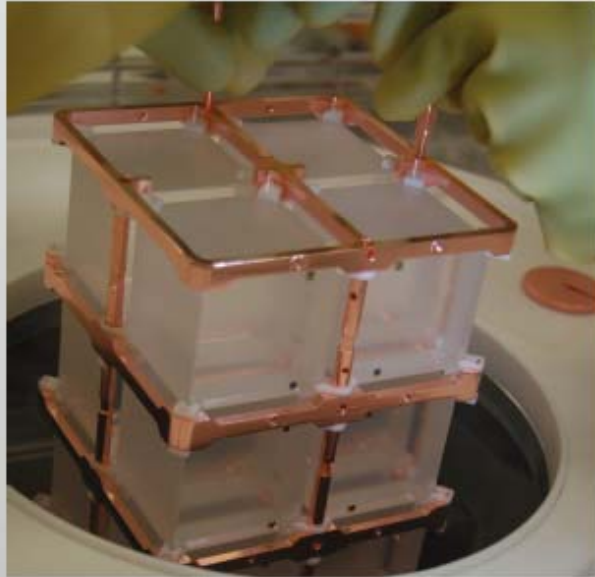
$b_{\text{CUORICINO}} = 0.169 \text{ c/keV/kg/y}$  due to:

- <sup>232</sup>Th in cryostat (30 ± 10%) →  $\gamma$
  - TeO<sub>2</sub> surfaces (10 ± 5%)
  - Surfaces facing detectors (50 ± 20%)
- } degraded  $\alpha$  particles

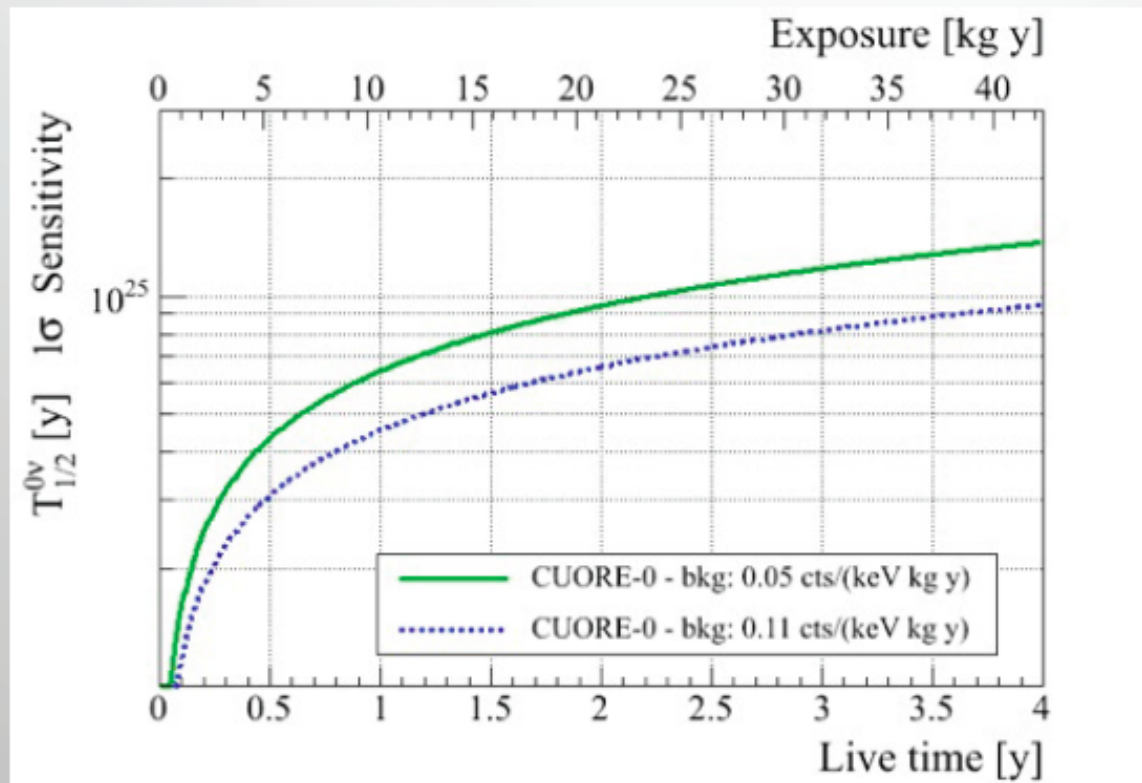
From M. Pedretti, Neutrino 2012



# CUORE-0: tower assembly



# CUORE-0: sensitivity



Limited by the cryostat contamination bkg

Background: 0.05-0.11 c/keV/kg/y range

If 0.05 c/(keV kg y), expected 2-year sensitivity

is

$$T_{1/2} = 5.9 \times 10^{24} \text{ y @ 90\% CL}$$

(CUORICINO:  $T_{0\nu} > 2.8 \times 10^{24} \text{ y}$ )

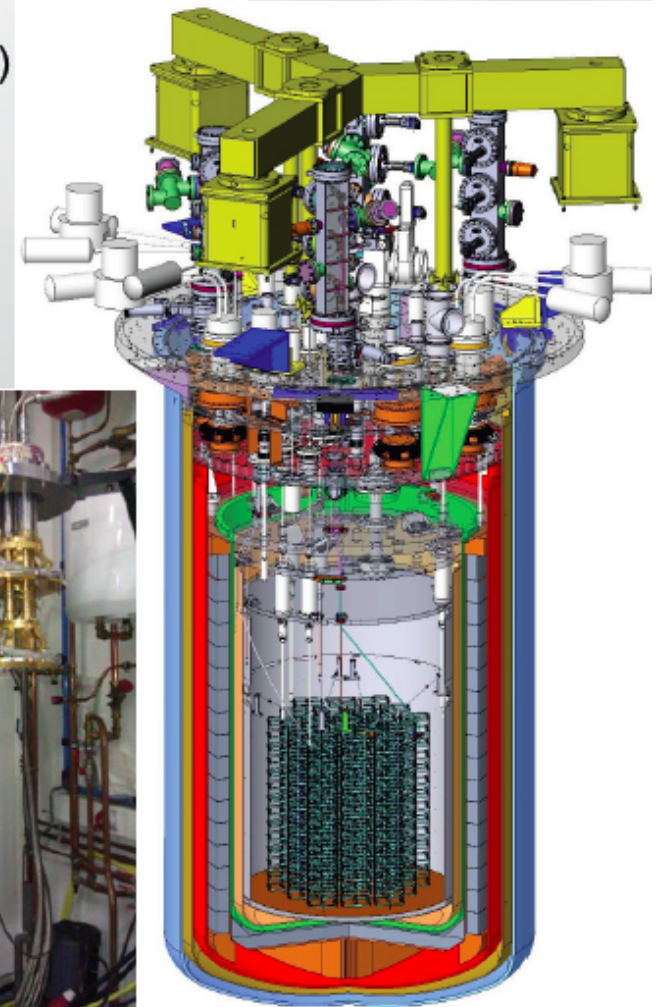
$$m_{\beta\beta} = 170 - 390 \text{ meV}$$

From M. Pedretti, Neutrino 2012



# CUORE status

- Crystals, almost all arrived (all at LNGS by the end of 2012)
- Copper parts are being machined and cleaned
- Dilution unit delivered to LNGS (though some repairs needed)
- CUORE Hut, and most of all the infrastructures, ready
- Detector assembly line, ready (small modifications)
- Radon abatement system installed
- 3 (of 6) cryostat vessels delivered soon at LNGS
- Commissioning of the cryostat second half of 2012

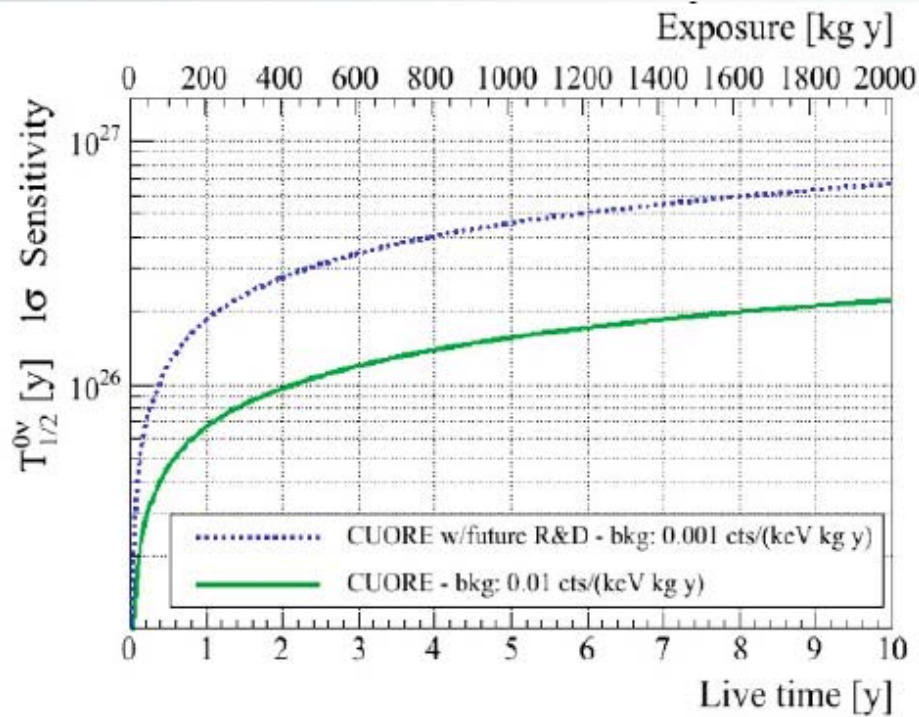


Crystals	12/12
Thermistors	13/03
Cleaned Cu parts	13/12
Cryogenic	13/12
Tower Assembly	14/04
Detector insertion	14/07
Cool Down	14/11

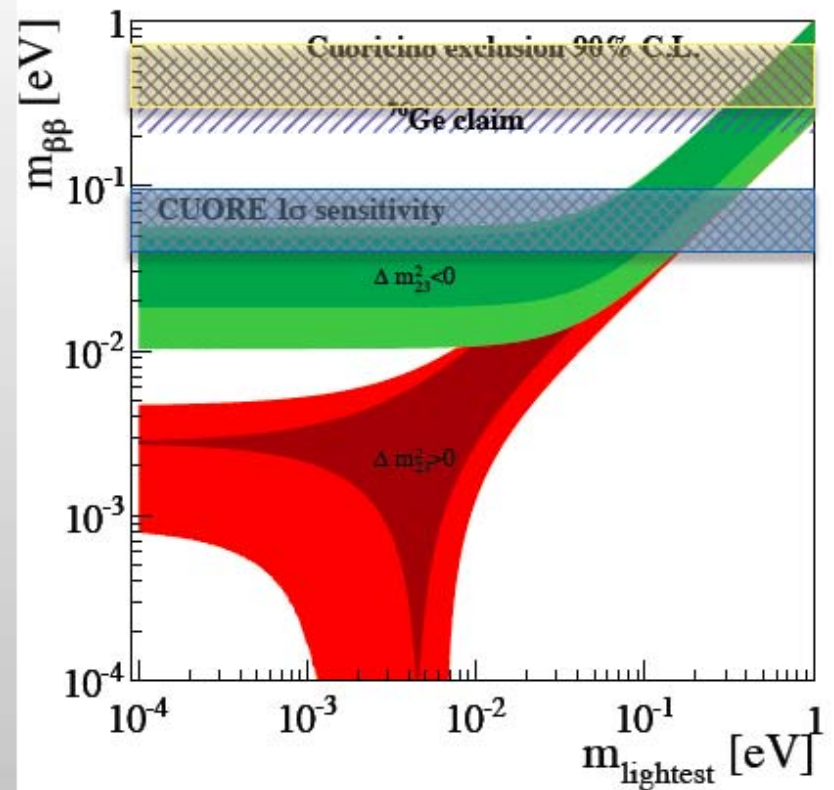


From M. Pedretti, Neutrino 2012

# CUORE sensitivity



F. Alessandria et al. [CUORE coll.]  
<http://arxiv.org/abs/1109.0494v1>



$$T_{1/2}^{0\nu} ({}^{130}\text{Te}) > 1.6 \times 10^{26} \text{ y (1}\sigma)$$

$$m_{\beta\beta} < 40 - 90 \text{ meV}$$

From M. Pedretti, Neutrino 2012

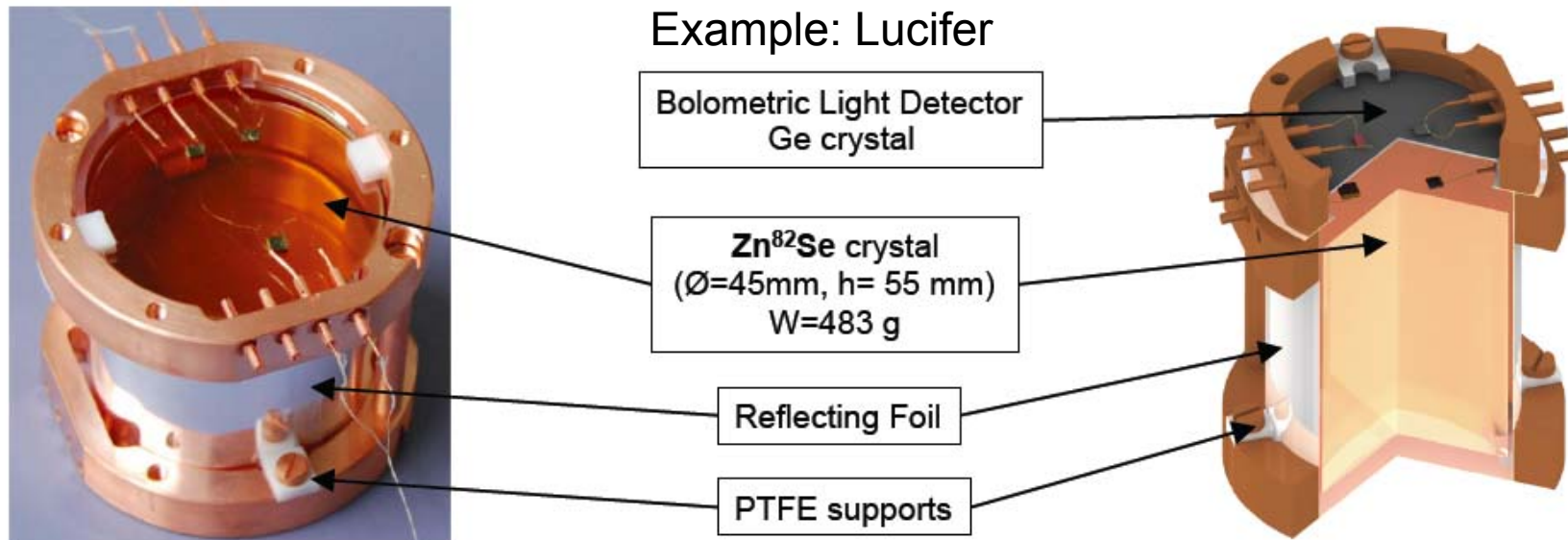


# Scintillating bolometers: Lucifer/Lumineu

Lucifer (Italy): ZnSe

Lumineu (France): ZnMoO<sub>4</sub>

Scintillating bolometers to recognize the  $\alpha$ -induced background thanks to the readout of the scintillation light



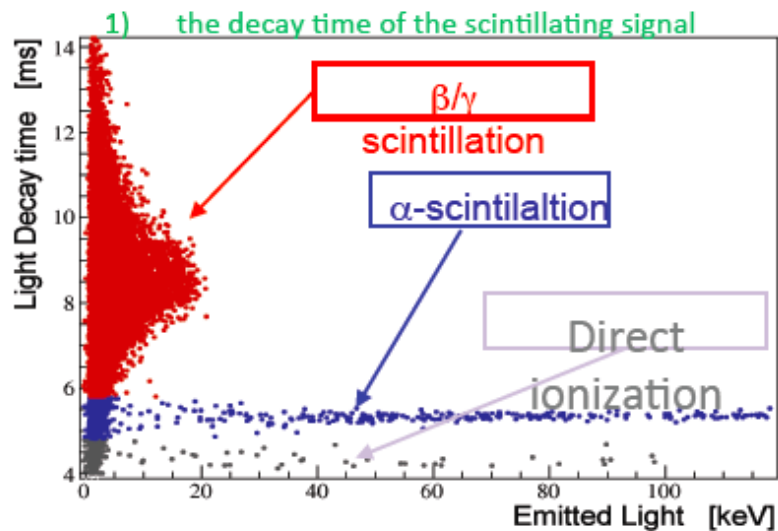
Array of 36÷44 enriched (95%) Zn<sup>82</sup>Se crystals.

Expected background in the ROI (2995 keV) is  $\sim 3\div 6 \cdot 10^{-3}$  c/keV/kg/y

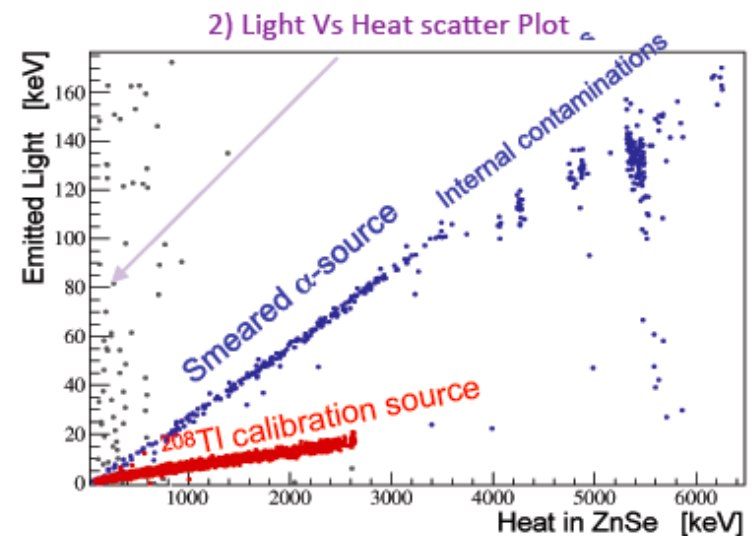
Energy resolution  $\sim 10$  keV FWHM

The  $\alpha$ -induced background is recognized:

- 1) the decay time of the scintillating signal
- 2) the different scintillation yield between  $\alpha$  and  $\gamma/\beta$  particles (the “usual” light Vs Heat scatter plot)



12.5 days measurement



2012

2013

2014

2015

R&D on light detectors

15 kg  $^{82}\text{Se}$  production

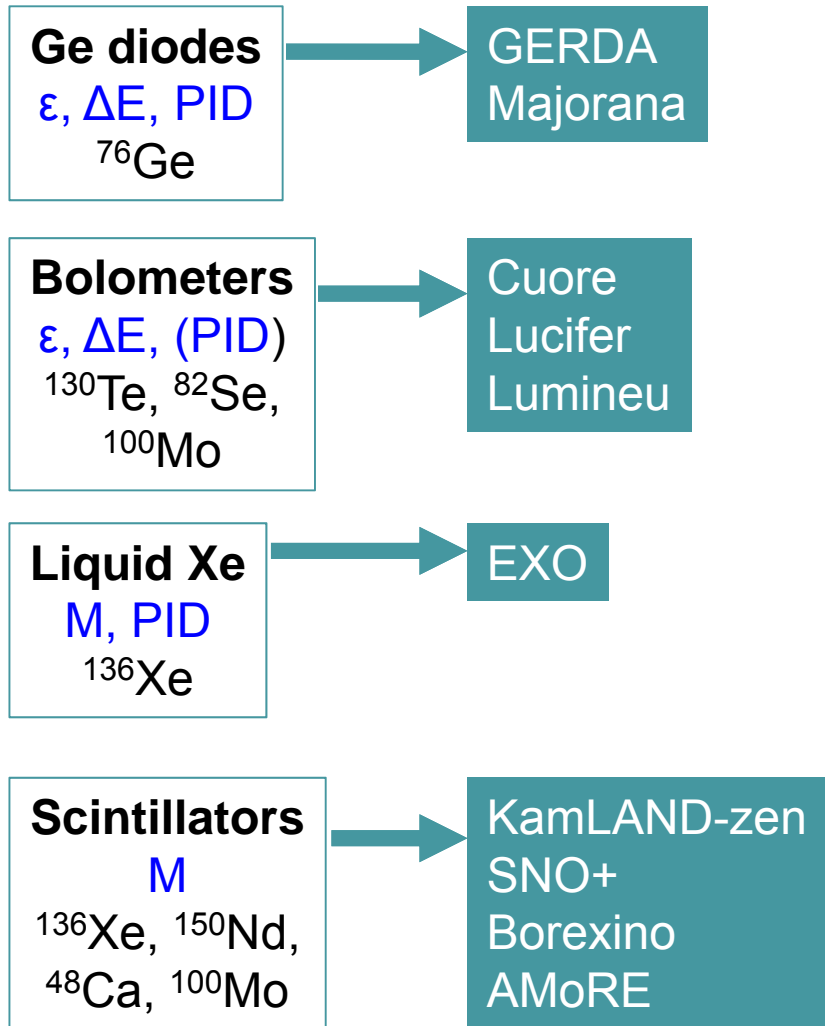
Enriched crystal growth

Detector assembling

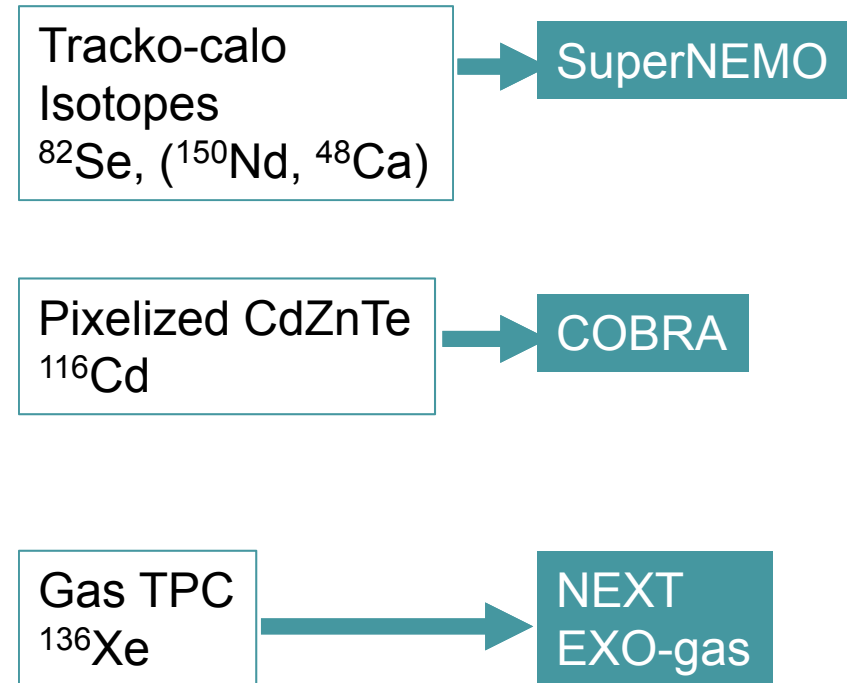
LUCIFER will be located in CUORICINO (now CUORE-0) cryostat, once CUORE-0 will finish his data taking (2015)

# Next generation experiments

## Calorimeters



## Electron tracking



Adopted from F. Piquemal