

INFN-Theory Group- Astroparticle

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Outline

- Overall INFN structures
- Collaboration teams-*(IS Iniziative specifiche)*
- Description of the collaboration teams
- Funding -refereeing of the collaboration teams

INFN structures

- INFN has 5 groups, one of which is the Theory group (~ 750 members) - chaired by G. Marchesini
- Theory group is divided in 6 subgroups, Astroparticle is one of them, with ~100 members (100% working on this subject/full time equivalent **FTE**)
- I am referee in collaboration with Francesco Vissani of the Astroparticle -Theory group of INFN

Astroparticle funding from INFN Theory

60% collaborations (IS)

travel +foreign visitors

~ 310 K€

+

extra INFN funds

~ 140 K€

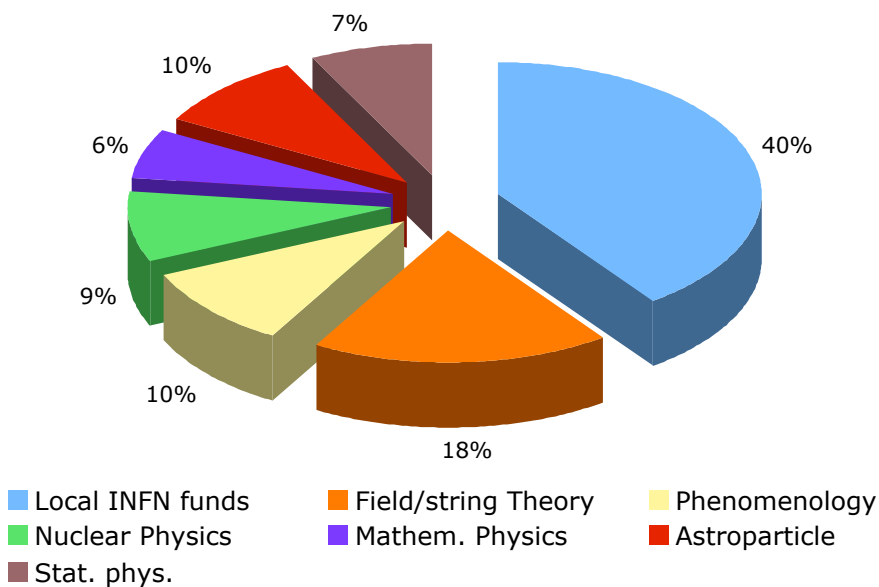
Astroparticle
⇒ 450K€

≥ 3 M€ for INFN TH

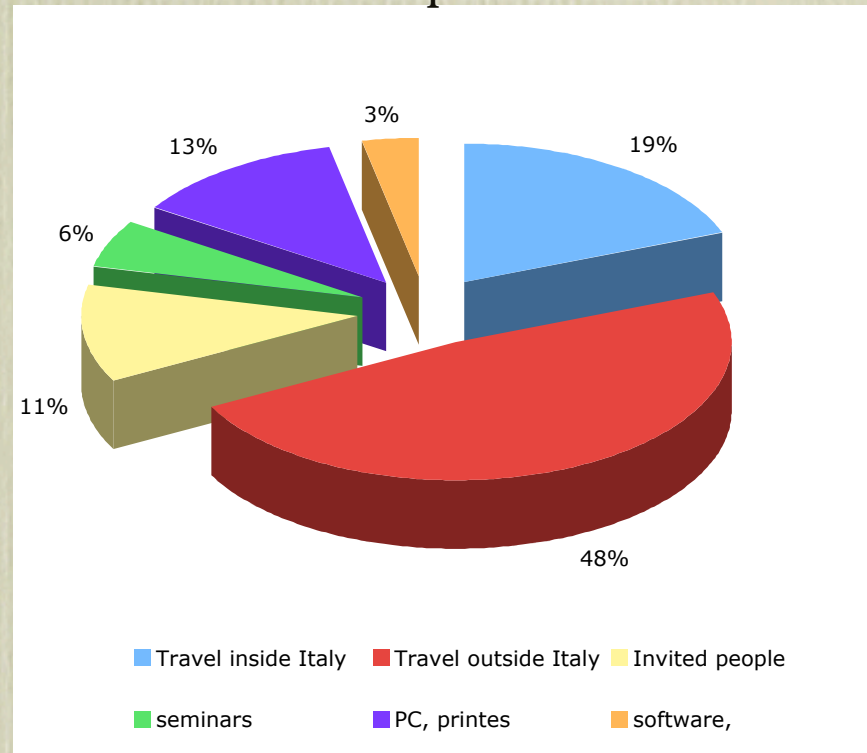
extra INFN funds

- travels
- seminars
- pc's, laptop, software
- foreign collaborators

Collaboration teams chart



INFN TH expenses chart



Astroparticle 450K€

Postdoc positions

- INFN funds ten foreign TH post-docs per year. Fellows can choose where in ITALY : 2~3 go to astroparticles
- INFN funds also some TH post-docs for italians abroad: some (about 5) go to astroparticles

5 Collaboration- teams

- **CT_{5I}** Compact stellar objects and dense hadronic matter (6 FTE)
- **GS_{5I}** Phenomenology of planck scale physics (5 FTE)
- **OG_{5I}** Modeling gravitational wave (GW) sources (11 FTE)
- **PD_{5I}** Inflation, dark matter, and the large-scale structure (16 FTE)
- **FA_{5I}** 50 FTE : Four subgroups

FA51 50 FTE: four subgroups

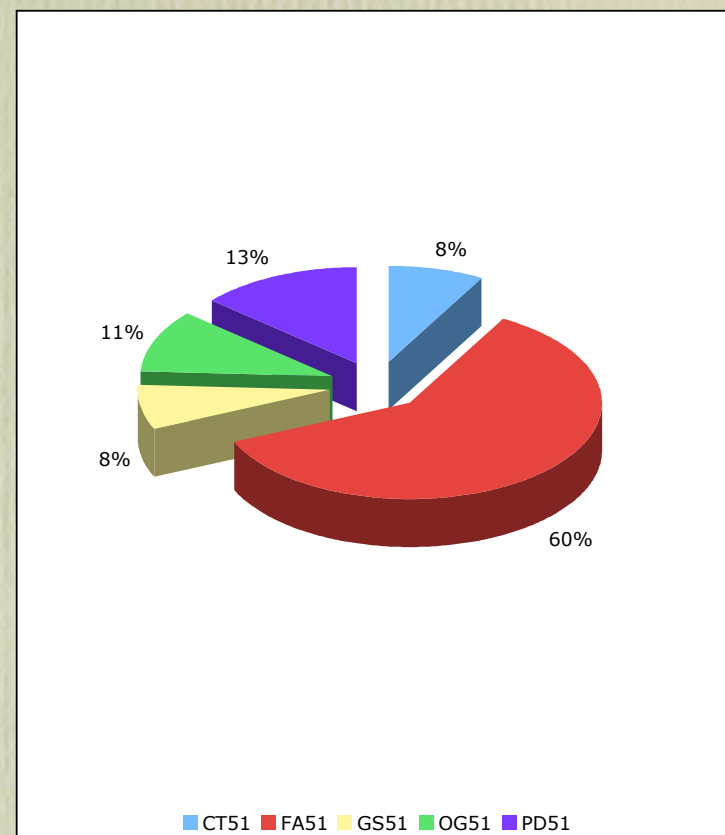
- A) Neutrinos in physics, astrophysics and cosmology (50%)
- B) Nuclear and subnuclear physics in the early universe (20%)
- C) Dark matter, dark energy and cosmic structures (15%)
- D) Astrophysical sources of radiation (15%)

Chart of Collaboration teams

- CT51 Nucl.-astroph. (resp. M. Baldo)
- GS51 Phen. of Planck scale (resp. A. Grillo)
- OG51 Grav. waves (resp. V. Ferrari)
- PD51 Inflation, DM, structures (resp. S. Matarrese)

FA51 (resp. G. Fogli)

- A) Neutrinos (50%)
- B) Early universe (20%)
- C) Dark matter/ energy (15%)
- D) Astrophysical /radiation (15%)



Astroparticles

Budget follows FTE's

Physics highlights

- **CT51** COMPACT STELLAR OBJECTS AND DENSE HADRONIC MATTER *European Project ESF (CompStar)*
- **GS51** PHENOMENOLOGY OF PLANCK SCALE PHYSICS (relation with **AGN**)
- Purpose of the **OG51** collaboration is *modeling gravitational wave (GW) sources* through theoretical studies and numerical simulations.
- **PD51** INFLATION: THEORETICAL PREDICTIONS AND OBSERVATIONAL TESTS. *Study of CMB anisotropies: BOOMERanG, MAXIMA-I, DASI, CBI, ARCHEOPS, WMAP .* DARK MATTER AND DARK ENERGY: ASTROPHYSICAL AND COSMOLOGICAL ASPECTS *PAMELA, AGILE and GLAST*

FA51: largest team

(for historical reasons: I coll. team)

- A) Neutrinos (50%)
- B) Early universe (20%)
- C) Dark matter/ energy (15%)
- D) Astrophysical /radiation (15%)

- **A) NEUTRINOS IN PHYSICS, ASTROPHYSICS AND COSMOLOGY** Solar neutrino physics, stellar astrophysics, Physics of atmospheric and (long-baseline) accelerator neutrinos, Properties of neutrinos of astrophysical origin (supernovae, nucleosynthesis). Absolute neutrino masses.
- **B) NUCLEAR AND SUBNUCLEAR PHYSICS IN THE EARLY UNIVERSE** Extended theories of gravitation and cosmology, CP violation, baryogenesis and leptogenesis, big bang nucleosynthesis
- **C) DARK MATTER, DARK ENERGY AND AND COSMIC STRUCTURES** Supersymmetric and mirror models Direct and indirect signatures of particle dark matter Cosmic background radiation (CMBR)
- **D) ASTROPHYSICAL SOURCES OF RADIATION** Nuclear reactions and stars evolution. Ultra high energy cosmic rays, Gamma ray bursts

European
networks:
ILIAS
ENTAPP
ISAPP

PhD Thesis

| | CT | GS | OG | PD | FA |
|----------|----|----|----|----|----|
| FTE | 6 | 5 | 11 | 16 | 50 |
| PhD 2004 | 1 | - | 2 | 1 | 9 |
| PhD 2005 | 1 | 0 | 0 | 1 | 5 |
| PhD 2006 | 0 | 0 | 1 | 4 | 5 |
| PhD 2007 | 1 | 2 | 5 | 3 | 4 |
| | 3 | 2 | 8 | 9 | 23 |

GS51 new

Popular papers in the last three years

| | CT | GS | OG | PD | FA |
|----------------|----|----|----|----|-----|
| >10 citations | 9 | 18 | 8 | 66 | 115 |
| >50 citations | 0 | 0 | 1 | 6 | 15 |
| >100 citations | 0 | 0 | 0 | 2 | 5 |

Popular authors: G.Fogli (BA), S. Matarrese(PD)
S.T. Petcov (TS) (seniors); E. Lisi (BA), A. Strumia
(PI),P.Ullio(TS) (mature); G.Bertone (PD),
S. Liberati (TS) (junior)

Some popular papers in the last four years

| # citations | Arguments | Authors | Collaboration |
|-------------|---|--|---------------|
| 245 | Global analysis of three-flavor neutrino masses and mixings. | G.Fogli-E.Lisi (BA), <i>et al.</i> Prog.Part.Nucl. Phys. 57:742, (06) | FA5I |
| 189 | Non-Gaussianity from inflation: Theory and observations | S. Matarrese (PD) <i>et al.</i> Phys.Rept.402:103 (04) | PD5I |
| 189 | Dark SUSY: Computing supersymmetric dark matter properties numerically | P. Ullio (TS) <i>et al.</i> JCAP 0407:008 (04) | FA5I |
| 145 | Theory of neutrinos: A White paper | S.T. Petcov (TS) <i>et al.</i> Rept.Prog.Phys. 70:1757,(07) | FA5I |
| 131 | Neutrino masses and mixings and... | A.Strumia (PI), F.Vissani (LNGS) hep-ph/0606054 | PD5I & FA5I |
| 134 | Implications of neutrino data circa 2005 | F. Vissani, (LNGS) <i>et al.</i> Nucl.Phys.B726:294-316,(05). | FA5I |

International refereeing every three years

- In order to finance our collaboration teams we ask several international referees to rate them (A,B,C)
- Local referees then finance accordingly