

Neutrino oscillations: recent results and perspectives

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IHEP, Protvino, Russia, 27 June 2013



OUTLINE

- **neutrino oscillations**
- **long baseline accelerator experiments**
- **measurements of θ_{13}**
 - **accelerator experiments**
 - **reactor experiments**
- **near and far future perspectives**

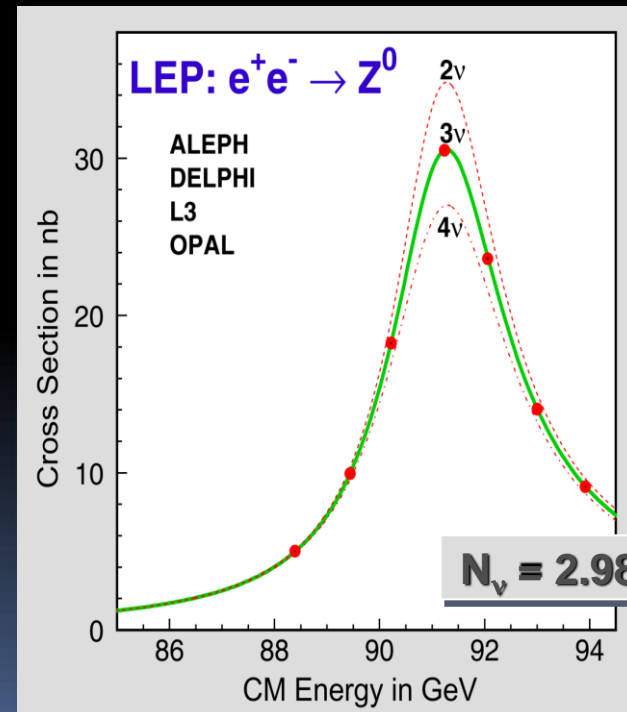


Standard Model

Three neutrino flavours: ν_e ν_μ ν_τ
Neutrino – partner of charged lepton: $W \rightarrow e\nu_e$ $W \rightarrow \mu\nu_\mu$ $W \rightarrow \tau\nu_\tau$
Neutrinos – massless particles
Lepton numbers L_e L_μ L_τ conservation
Neutrino oscillations forbidden
CP = 1 in lepton sector

LEP experiments:
from the width of Z

Three active neutrinos





Neutrino oscillation hypothesis

B. Pontecorvo: the idea of massive neutrino and oscillations – 1957



- neutrino – antineutrino transitions
- non-zero mass and mixture of neutrinos
- oscillation probability depends of neutrino **mass**, neutrino energy E_ν and path length L



Weak interaction eigenstates



$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$



m_1
 m_2
 m_3

Mass eigenstates

Weak (active) eigenstates are not equal to mass eigenstates



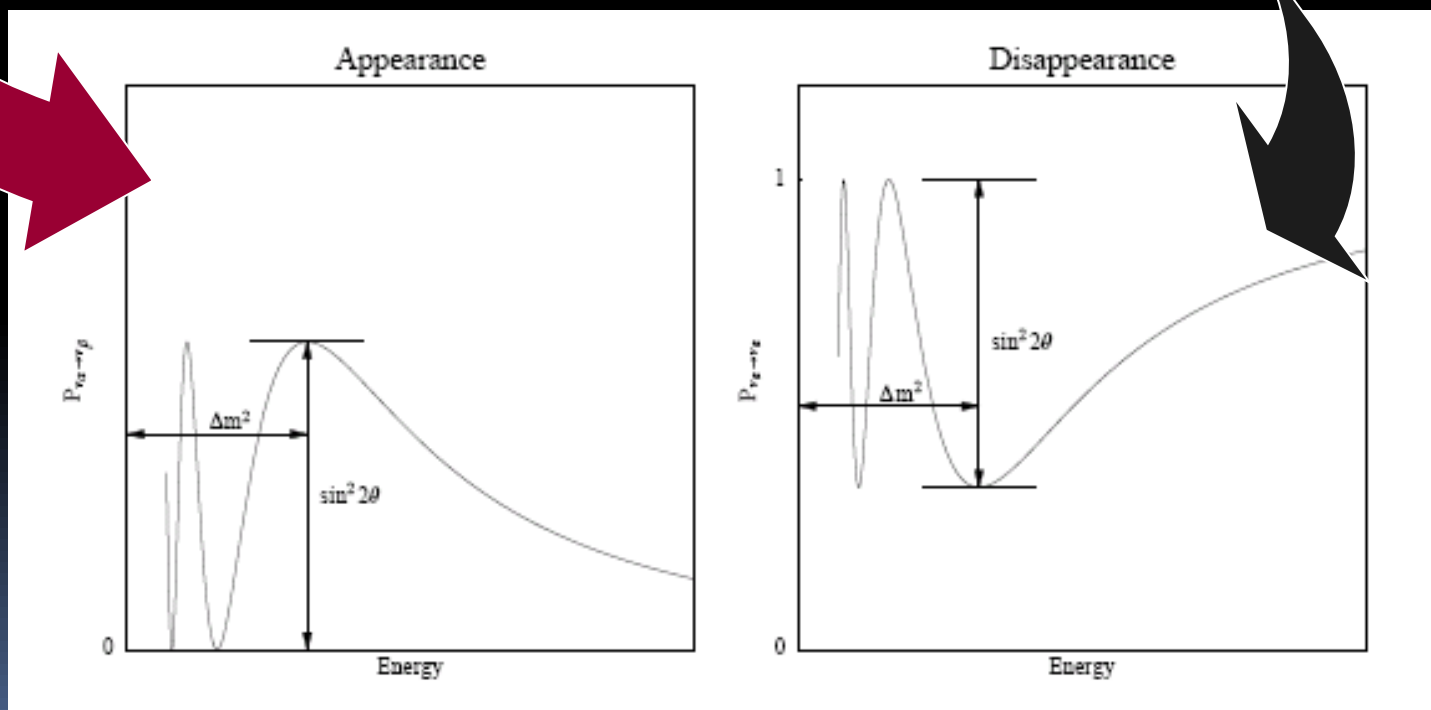
2 types of oscillation experiments

Oscillation experiments:
Appearance and Disappearance

Baseline, L
Neutrino energy, E_ν

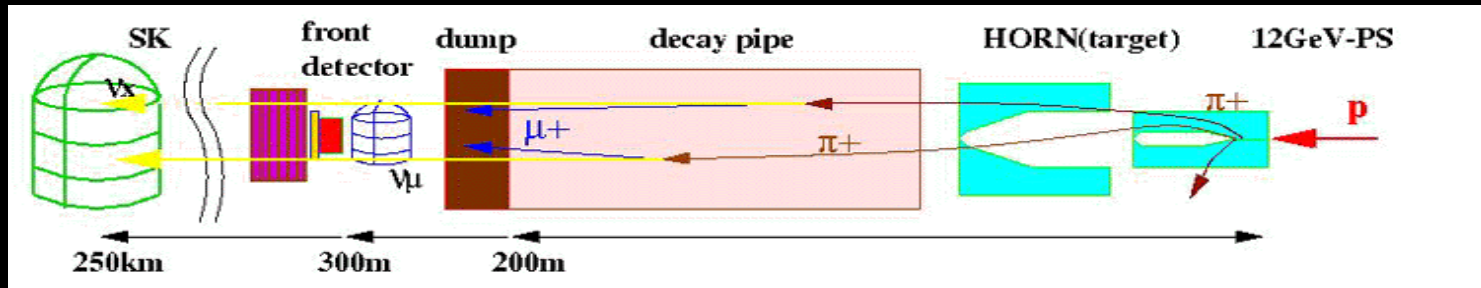
fixed
measured

$$P_{\nu_\alpha \rightarrow \nu_\beta} = \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4E} \right),$$
$$P_{\nu_\alpha \rightarrow \nu_\alpha} = 1 - \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4E} \right)$$



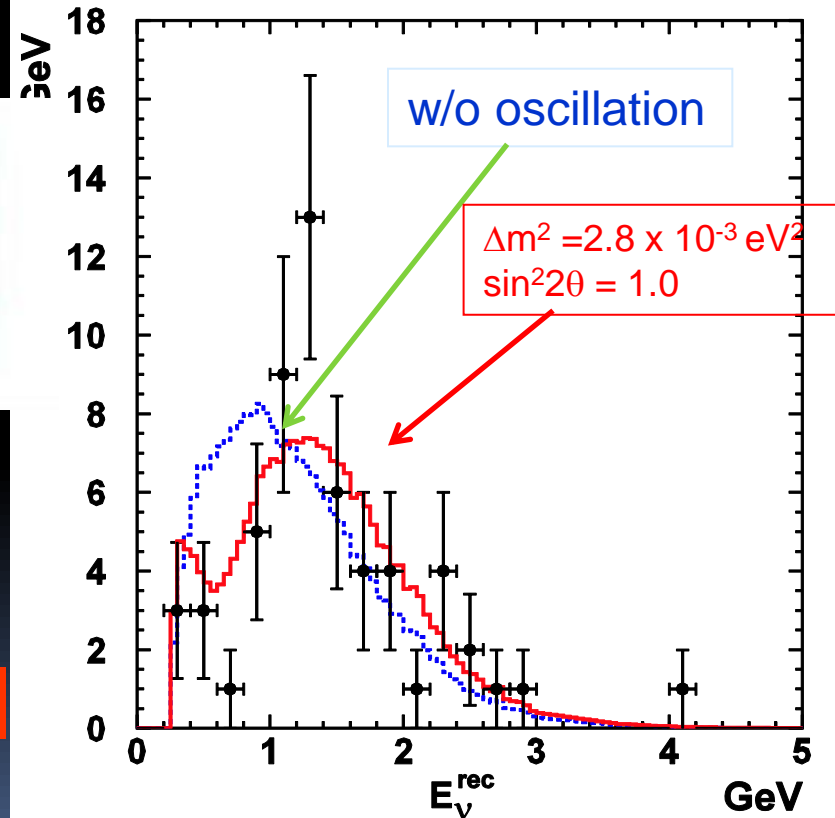


K2K: first LBL experiment



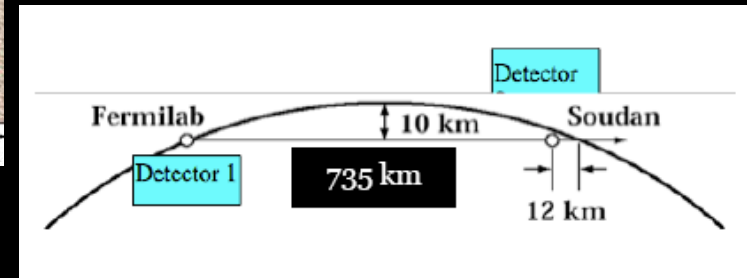
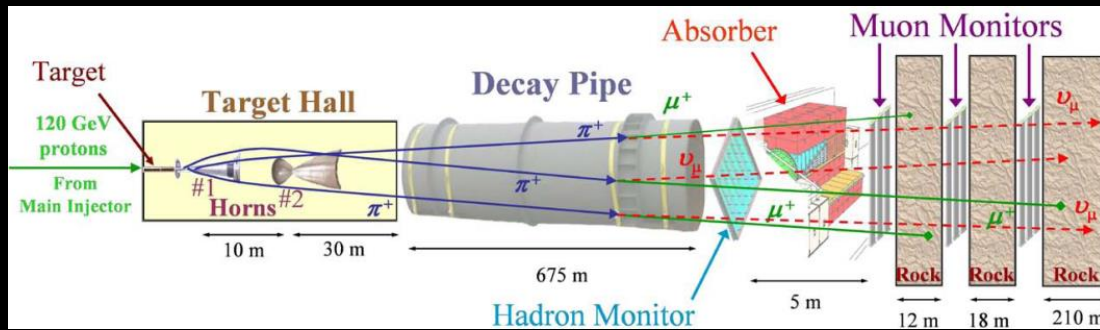
**K2K confirmed SK result:
oscillations of atmospheric neutrinos**

Null oscillation probability = 0.0050% (4.1 σ)





MINOS



Proton beam: 120 GeV protons
 ν - beam: ν_μ - 92.9%, anti- ν_μ - 5.8%, (ν_e + anti- ν_e)- 1.3%,
 peak energy \sim (3-9) GeV



~1 kT Near Detector

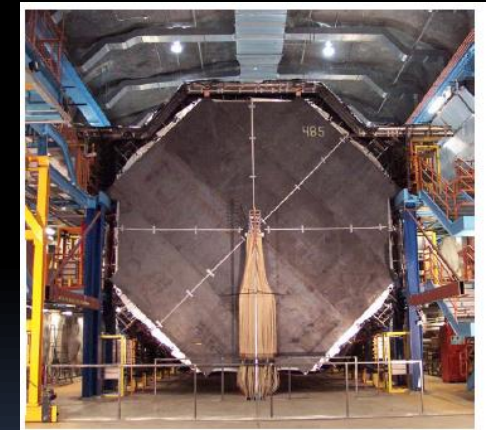
Far Detector

5.4 kton magnetized Fe/Sci Tracker/Calorimeter

Near Detector

980 ton at $L \approx 1$ km

the same technology in both detectors



5.4 kT Far Detector



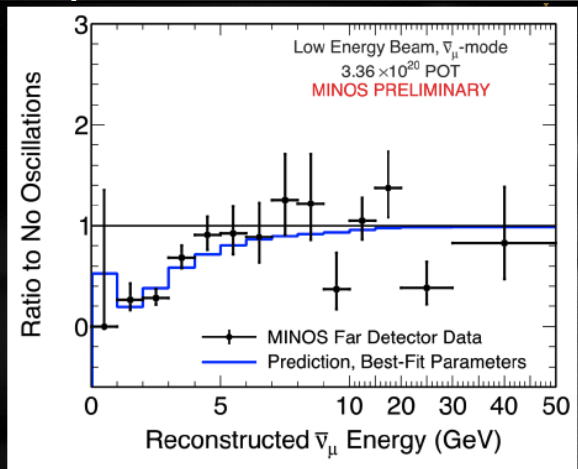
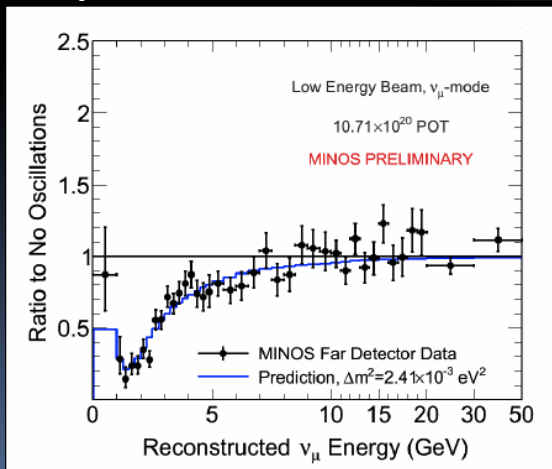
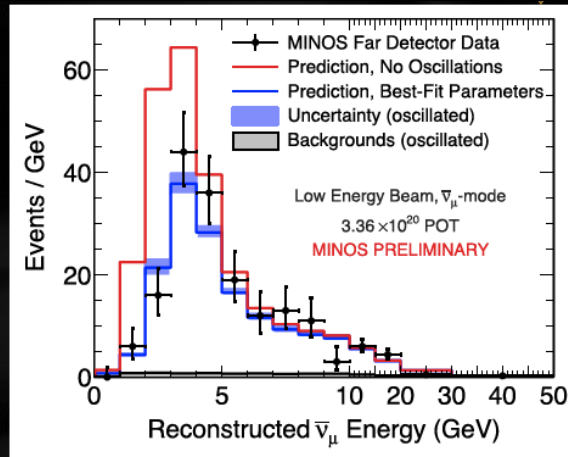
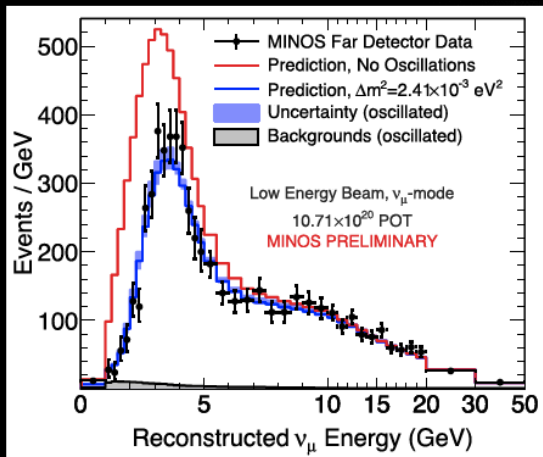
MINOS

almost final
result



2894 events **detected** in MINOS Far Detector
3564 event **expected** in the absence of oscillations

226 events **detected** in MINOS Far Detector
312 events **expected** in the absence of oscillations



ν

$|\Delta m|^2 = (2.35 + 0.11 - 0.08) \times 10^{-3}$
 $\sin^2(2\theta) > 0.91$ (90% CL)

anti- ν

$\Delta m^2 = (2.64 + 0.28 - 0.27) \times 10^{-3}$
 $\sin^2(2\theta) > 0.78$ (90% CL)

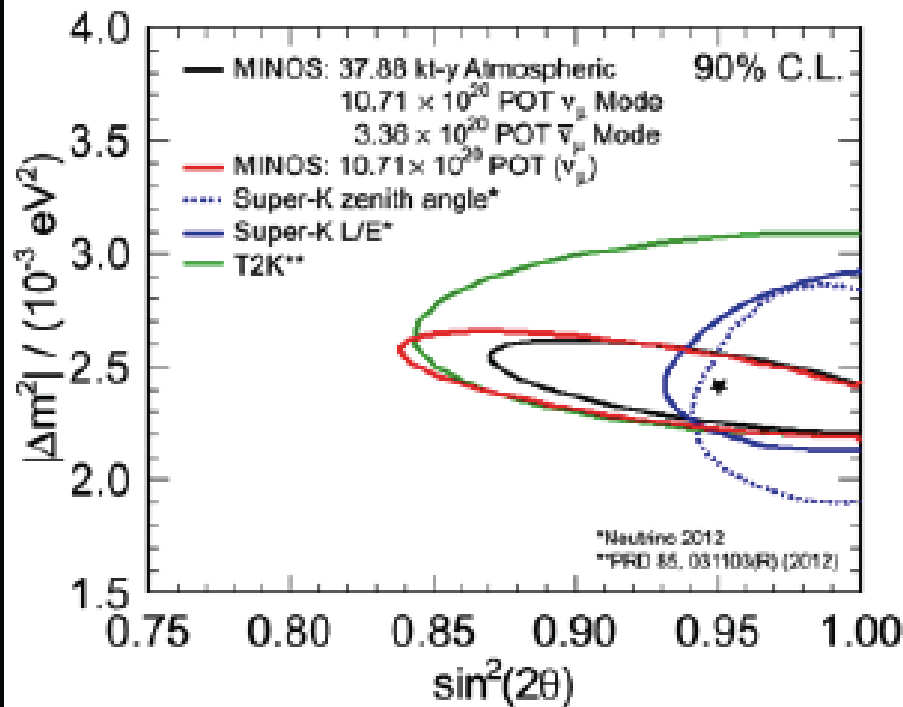
no tension between
neutrinos and
antineutrinos



MINOS: combined ν + anti- ν

All data sets (neutrino, anti-neutrino, atmospheric) combined for final measurement of ν_μ disappearance parameters

arXiv:1304.6335



Combined analysis

$$\sin^2 2\theta = 0.950^{+0.035}_{-0.036}$$

$$|\Delta m^2| = 2.41^{+0.09}_{-0.10} \times 10^{-3} \text{eV}^2$$



a hint on non-maximal
mixing angle θ_{23}



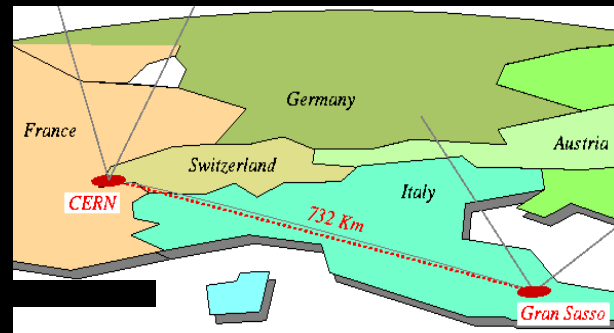
OPERA

$\nu_\mu \rightarrow \nu_\tau$ direct search

$$P(\nu_\mu \rightarrow \nu_\tau) = \cos^4\theta_{13}\sin^2\theta_{23}\sin^2[1.27\Delta m_{23}^2 L(\text{km})/E(\text{GeV})]$$



High energy, long baseline ν beam
($E \approx 17 \text{ GeV}$ $L \sim 730 \text{ km}$)

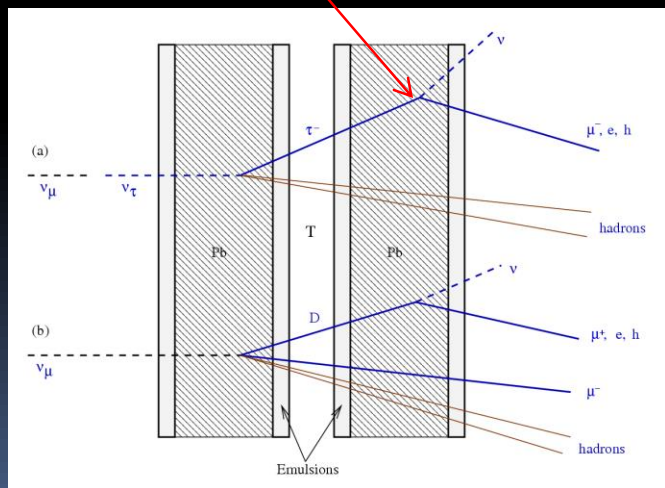


kink

Target mass $\sim 1300\text{t}$

$$E/L \sim 2.3 \times 10^{-2} \sim 10 \Delta m_{23}^2 (\text{atm})$$

pure beam: 2% anti ν_μ ; $<1\%$ ν_e



Main background

Expectation after 5 years
data taking:

~ 22000 ν interactions

~ 120 ν_τ interactions

~ 10 ν_τ reconstructed

<1 background event



OPERA

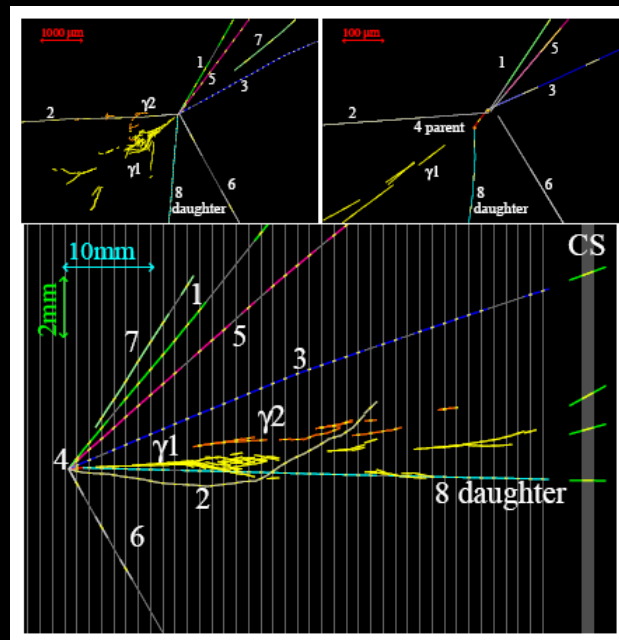
For $|\Delta m|^2 = 2.5 \times 10^{-3} \text{ eV}^2$
 $18.9 \times 10^{19} \text{ POT}$

2.7 tau events expected

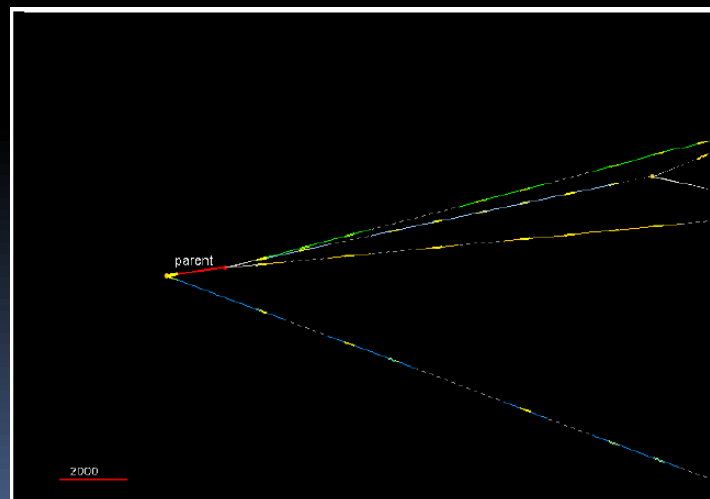
0.3 events background

3 candidates observed

1st event

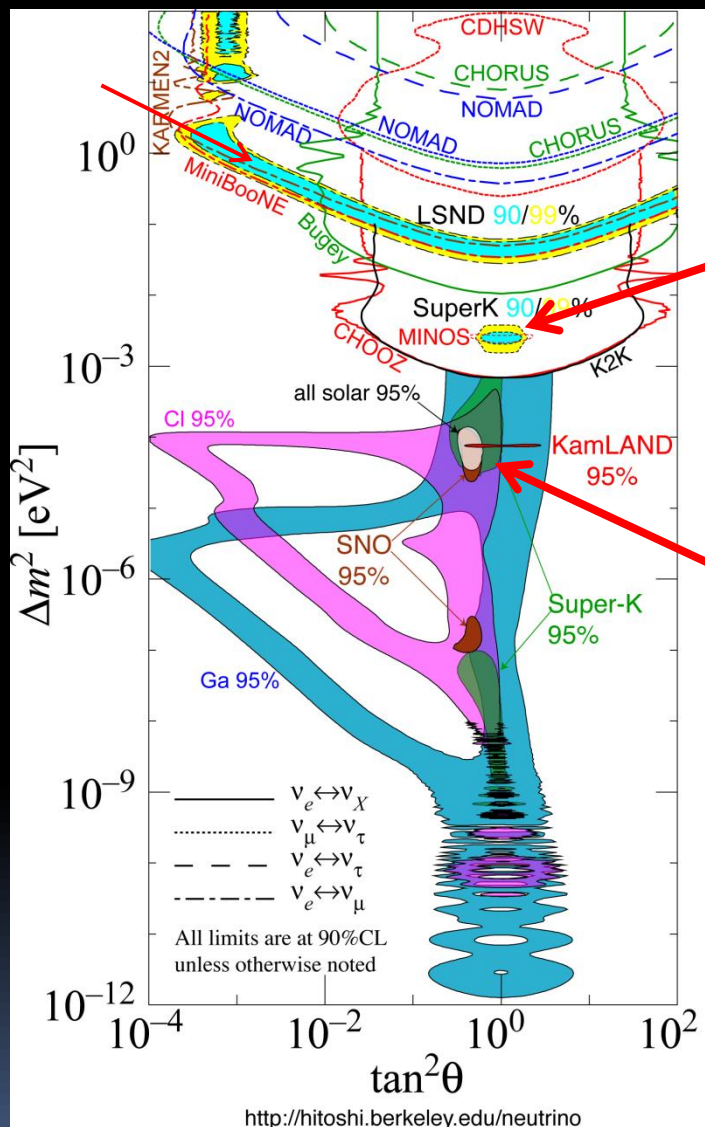


2nd event



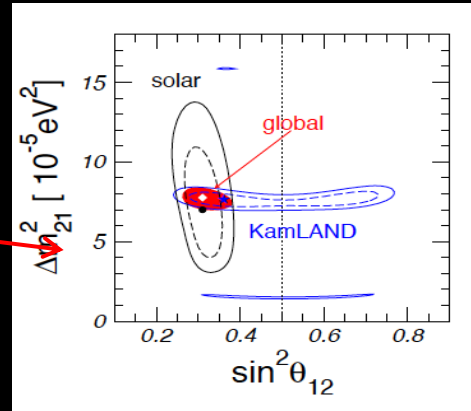


Oscillation results



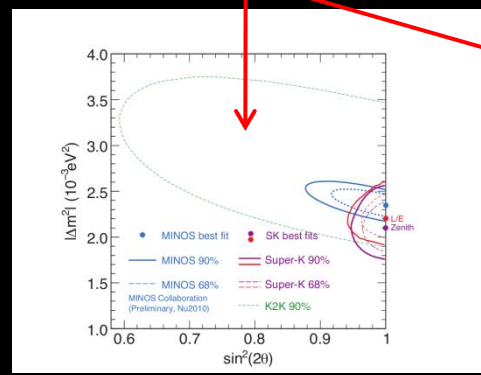
By June 2011

Solar + KamLAND

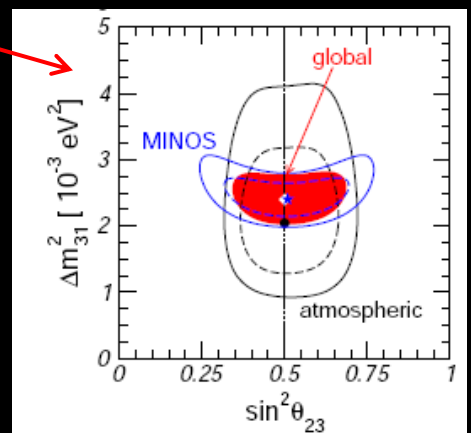


atm

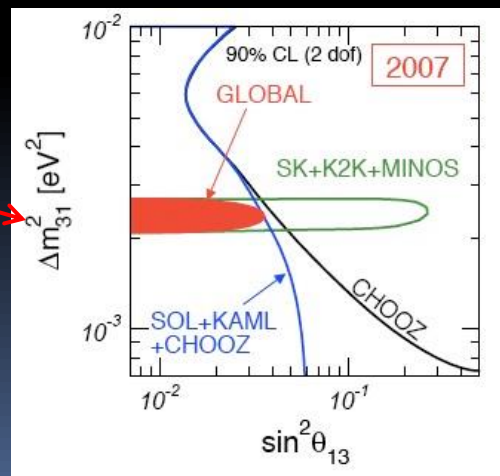
SK + K2K + MINOS



sol



CHOOZ + atm + LBL
 $\sin^2(2\theta_{13}) < 0.11$ (90%CL)





ν oscillations and mixing

Standard Model: neutrinos are *massless* particles

3 families

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix}$$

solar

atmospheric

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \begin{pmatrix} \cos\theta_{13} & 0 & \sin\theta_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin\theta_{13}e^{-i\delta} & 0 & \cos\theta_{13} \end{pmatrix} \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

link between atmospheric and solar

U parameterization:

three mixing angles θ_{12} θ_{23} θ_{13}

CP violating phase δ

$$\Delta m_{ij}^2 = m_i^2 - m_j^2$$

$$\Delta m_{12}^2 + \Delta m_{23}^2 + \Delta m_{31}^2 = 0$$

$$\theta_{23} \sim 45^\circ$$

$$\Delta m_{23}^2 \cong \Delta m_{31}^2 =$$

$$\Delta m_{atm}^2 \approx 2.4 \times 10^{-3} \text{ eV}^2$$

$$\theta_{12} \sim 34^\circ$$

$$\Delta m_{12}^2 = \Delta m_{sol}^2 \approx 7.5 \times 10^{-5} \text{ eV}^2$$

two independent Δm^2

by June 2011

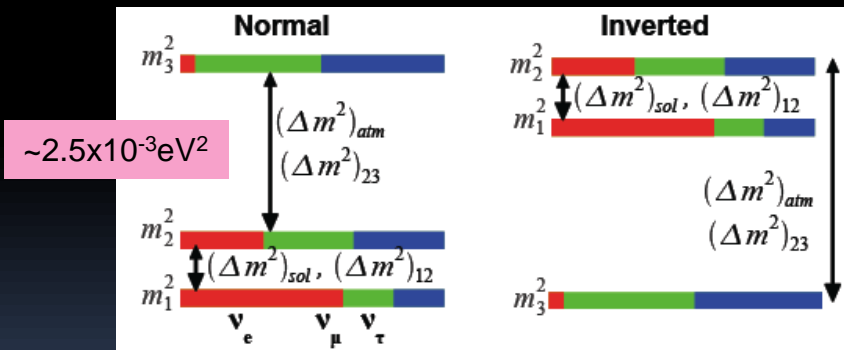
$$\sin^2 2\theta_{13} < 0.15 \text{ at 90\% CL}$$



Before Summer 2011

T.Schwetz, M.Tortola, J.Valle, hep-ph:1103.0734v2

parameter	best fit $\pm 1\sigma$	2σ	3σ
Δm_{21}^2 [10^{-5}eV^2]	$7.59^{+0.20}_{-0.18}$	7.24–7.99	7.09–8.19
Δm_{31}^2 [10^{-3}eV^2]	2.45 ± 0.09 $-(2.34^{+0.10}_{-0.09})$	2.28 – 2.64 $-(2.17 – 2.54)$	2.18 – 2.73 $-(2.08 – 2.64)$
$\sin^2 \theta_{12}$	$0.312^{+0.017}_{-0.015}$	0.28–0.35	0.27–0.36
$\sin^2 \theta_{23}$	0.51 ± 0.06 0.52 ± 0.06	0.41–0.61 0.42–0.61	0.39–0.64
$\sin^2 \theta_{13}$	$0.010^{+0.009}_{-0.006}$ $0.013^{+0.009}_{-0.007}$	≤ 0.027 ≤ 0.031	≤ 0.035 ≤ 0.039



- ✓ only upper limit on θ_{13}
- ✓ θ_{23} maximal?
- ✓ mass hierarchy (sign of Δm_{31}^2)
- ✓ no hint on CP violation

?? θ_{13} , mass hierarchy , δ ??



Importance of θ_{13}

- **Zero value of θ_{13}** would be a hint on a new symmetry (tri-bi-maximal)
- **Zero value of θ_{13}** would eliminate a possibility for the CKM mechanism in neutrino mixing
- **A non-zero value of θ_{13}** opens a door for searching of leptonic **CP violation**
- **A non-zero (and not small) value of θ_{13}** gives good chances for measurement of **mass hierarchy** and **CP violation** in neutrino oscillations using present neutrino beams and detectors

The size of θ_{13} → Future Program of neutrino physics



LONG-BASELINE NEUTRINO OSCILLATION EXPERIMENT



SuperK

Toyama

Kamioka Mine



JPARC

Tokai

Tokyo

JAPAN

Tokyo/Narita Airport



- 12 countries
- 59 institutes
- ≈ 500 collaborators

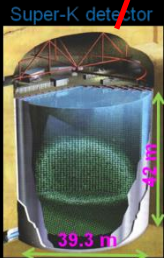
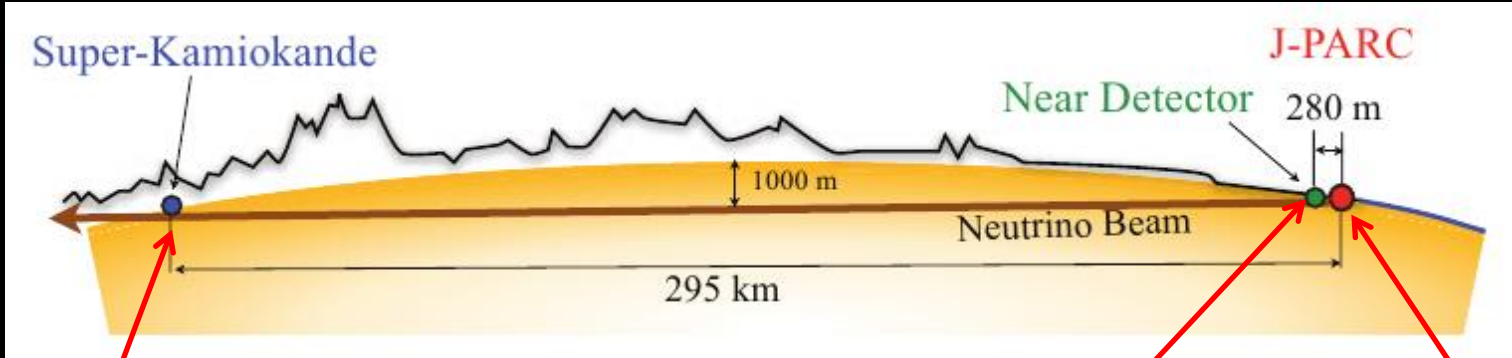
Canada, France, Germany, Italy,
Japan, Korea, Poland, Russia, Spain,
Switzerland, UK, USA.



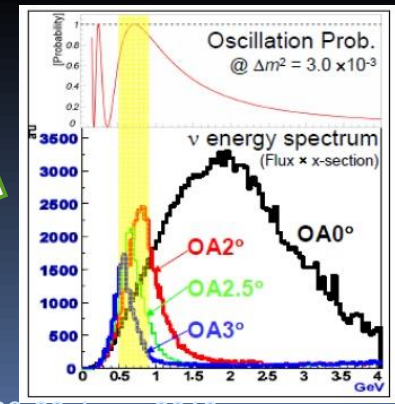
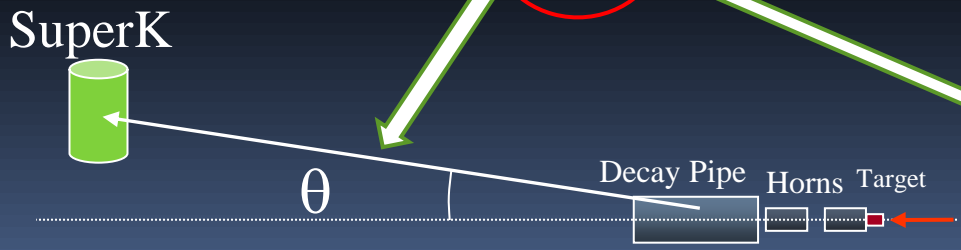
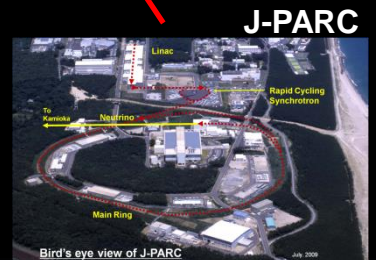
T2K: - search for $\nu_\mu \rightarrow \nu_e$

- measurement of θ_{13}
- CP - violation

- >500 members from 12 countries
- Russia: INR RAS



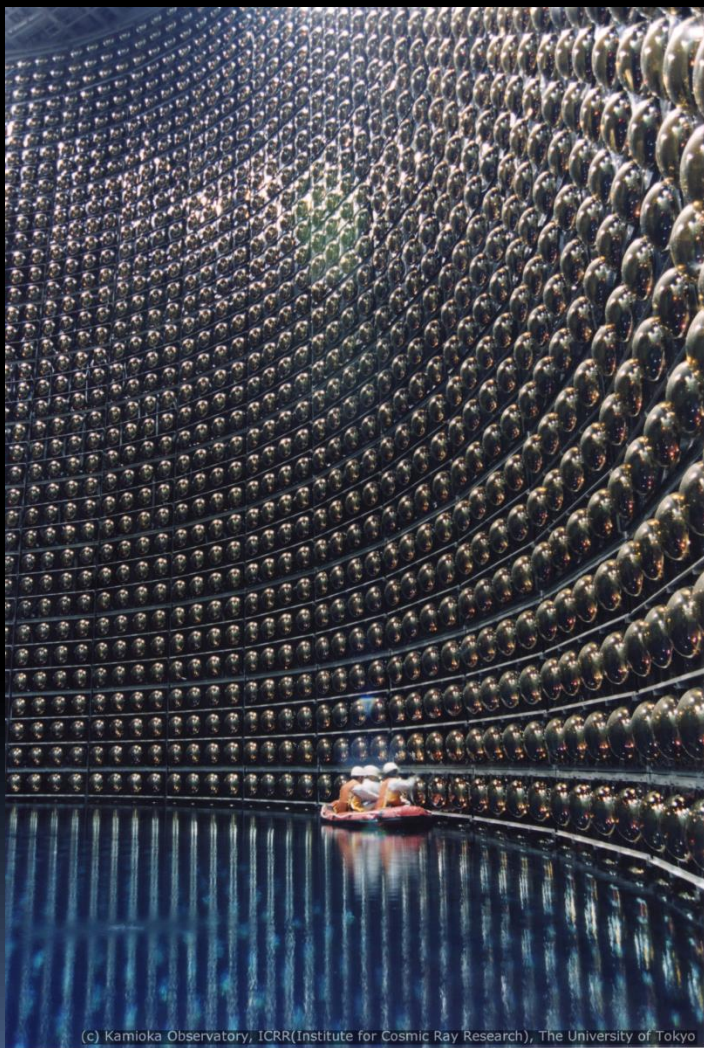
Off-axis beam conception: neutrino detectors are placed at a small angle θ with respect to the proton beam



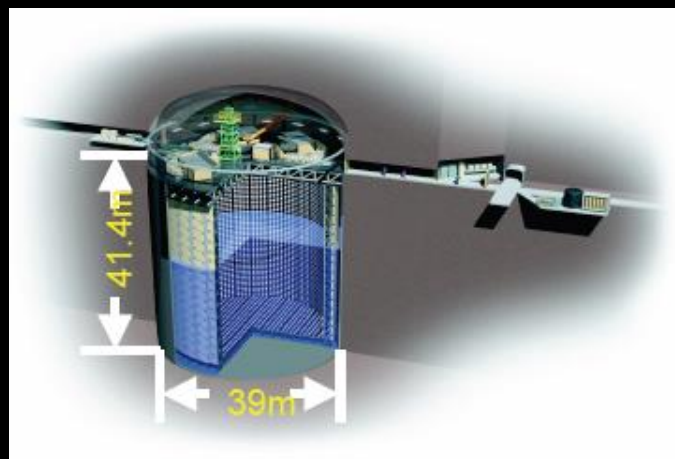
With 2.5 degree off-axis angle the neutrino energy is tuned to the oscillation maximum



SuperKamiokande – T2K Far Detector



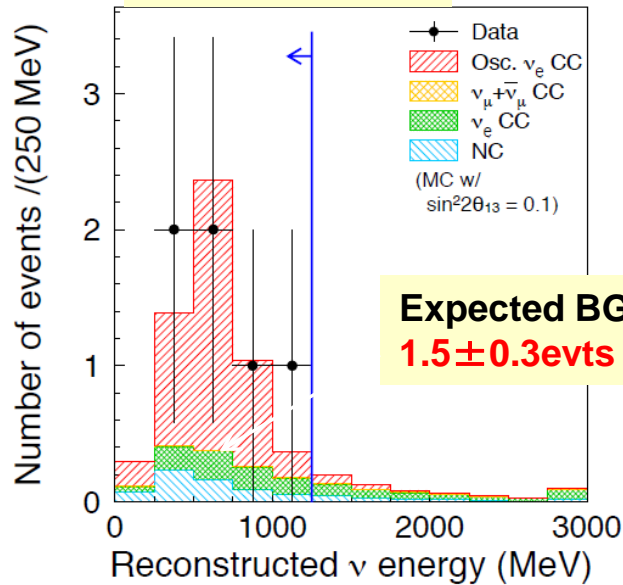
(c) Kamioka Observatory, ICRR(Institute for Cosmic Ray Research), The University of Tokyo



50 kt of pure water;
Fiducial Volume: 22.5 kt
Inner Detector: >11000 PMTs (20")
Outer Detector: ~2000 PMTs (8")
(see talk by M. Smy)



6 ν_e events



First T2K result



published in June 2011

1.43x10²⁰ POT
January 2010 –
March 2011

published **FIRST** clear indication
($\theta_{13} \neq 0$)

About two years ago, T2K published
of **electron neutrino appearance** ($\theta_{13} \neq 0$)

Then,

- 1 - Confirmation from MINOS
- 2 - Precise measurements by Double Chooz
- Daya Bay
- RENO

Selected for a Viewpoint in *Physics*

PRL 107, 041801 (2011)

PHYSICAL REVIEW LETTERS

week ending
22 JULY 2011

Indication of Electron Neutrino Appearance from an Accelerator-Produced Off-Axis Muon Neutrino Beam

K. Abe,⁴⁹ N. Abgrall,¹⁶ Y. Ajima,^{18,†} H. Aihara,⁴⁸ J. B. Albert,¹³ C. Andreopoulos,⁴⁷ B. Andrieu,³⁷ S. Aoki,²⁷ O. Araoka,^{18,†} J. Argyriades,¹⁶ A. Ariga,³ T. Ariga,³ S. Assylbekov,¹¹ D. Autiero,³² A. Badertscher,¹⁵ M. Barbi,⁴⁰ G. J. Barker,⁵⁶ G. Barr,³⁶ M. Bass,¹¹ F. Bay,³ S. Bentham,²⁹ V. Berardi,²² B. E. Berger,¹¹ I. Bertram,²⁹ M. Besnier,¹⁴ J. Beucher,⁸ D. Beznosko,³⁴ S. Bhadra,⁵⁹ F. d.M. M. Blaszczyk,⁸ A. Blondel,¹⁶ C. Bojchko,⁵³ J. Bouchez,^{8,*} S. B. Boyd,⁵⁶ A. Bravar,¹⁶ C. Bronner,¹⁴ D. G. Brook-Roberge,⁵ N. Buchanan,¹¹ H. Budd,⁴¹ D. Calvet,⁸ S. L. Cartwright,⁴⁴ A. Carver,⁵⁶ R. Castillo,¹⁹ M. G. Catanesi,²² A. Cazes,³² A. Cervera,²⁰ C. Chavez,³⁰ S. Choi,⁴³ G. Christodoulou,³⁰ J. Coleman,³⁰

The T2K experiment observes indications of $\nu_\mu \rightarrow \nu_e$ appearance in data accumulated with 1.43×10^{20} protons on target. Six events pass all selection criteria at the far detector. In a three-flavor neutrino oscillation scenario with $|\Delta m_{23}^2| = 2.4 \times 10^{-3} \text{ eV}^2$, $\sin^2 2\theta_{23} = 1$ and $\sin^2 2\theta_{13} = 0$, the expected number of such events is $1.5 \pm 0.3(\text{stat})$. Under this hypothesis, the probability to observe six or more candidate events is 7×10^{-3} , equivalent to 2.5σ significance. At 90% C.L., the data are consistent with $0.03(0.04) < \sin^2 2\theta_{13} < 0.28(0.34)$ for $\delta_{CP} = 0$ and a normal (inverted) hierarchy.

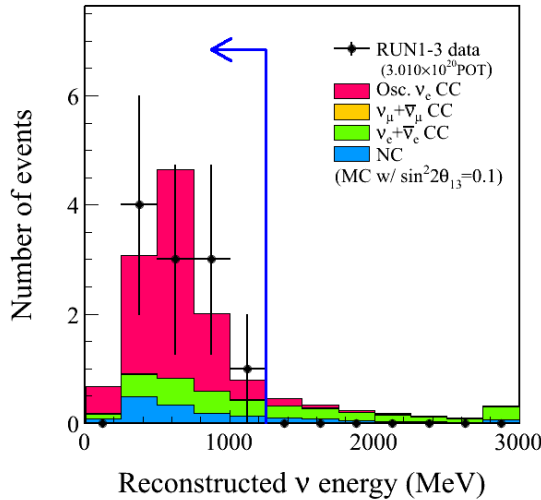
DOI: 10.1103/PhysRevLett.107.041801

PACS numbers: 14.60.Pq, 13.15.+g, 25.30.Pt, 95.55.Vj



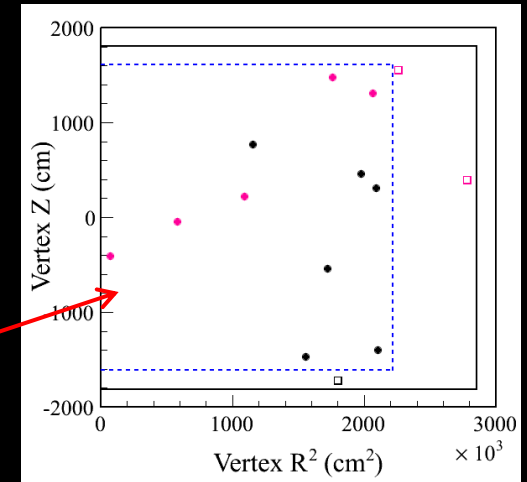
ν_e events

Energy spectra of ν_e events



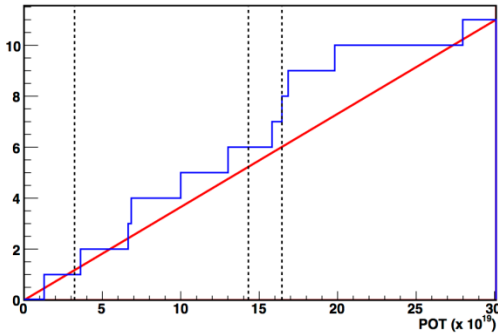
3.01x10²⁰ POT
 Statistics accumulated from January 2010 to July 2012

Run1
 +
 Run2
 +
 Run3

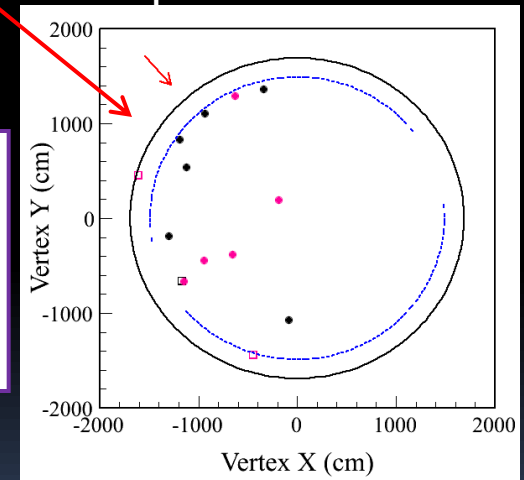


p-value = 6%

FCFV ν_e Candidates RUN1+RUN2+RUN3



Event Category	$\sin^2 2\theta_{13} = 0$	$\sin^2 2\theta_{13} = 0.1$
Total	3.22 ± 0.43	10.71 ± 1.10
ν_e Signal	0.18	7.79
Intrinsic ν_e Background	1.67	1.56
ν_μ Background (mostly π^0)	1.12	1.12
$\bar{\nu}_e + \bar{\nu}_\mu$ Background	0.16	0.16



Detected **11** events
 Expected **3.3 ± 0.4(syst)** events
 for $\theta_{13}=0$, NH and $\delta = 0$



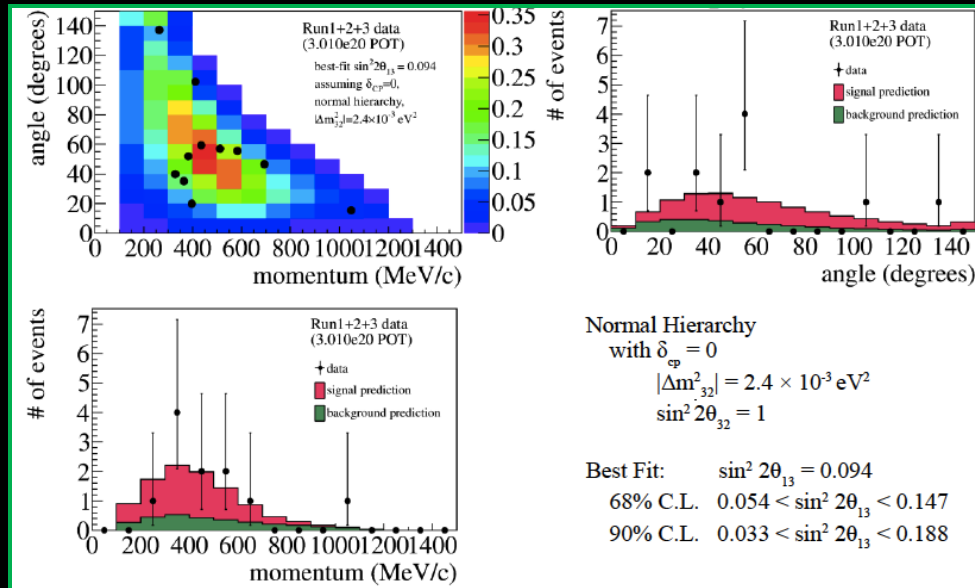
3.1 σ observation of $\nu_\mu \rightarrow \nu_e$

T2K Collaboration, arXiv:1304.0841



T2K: ν_e events

July 2012



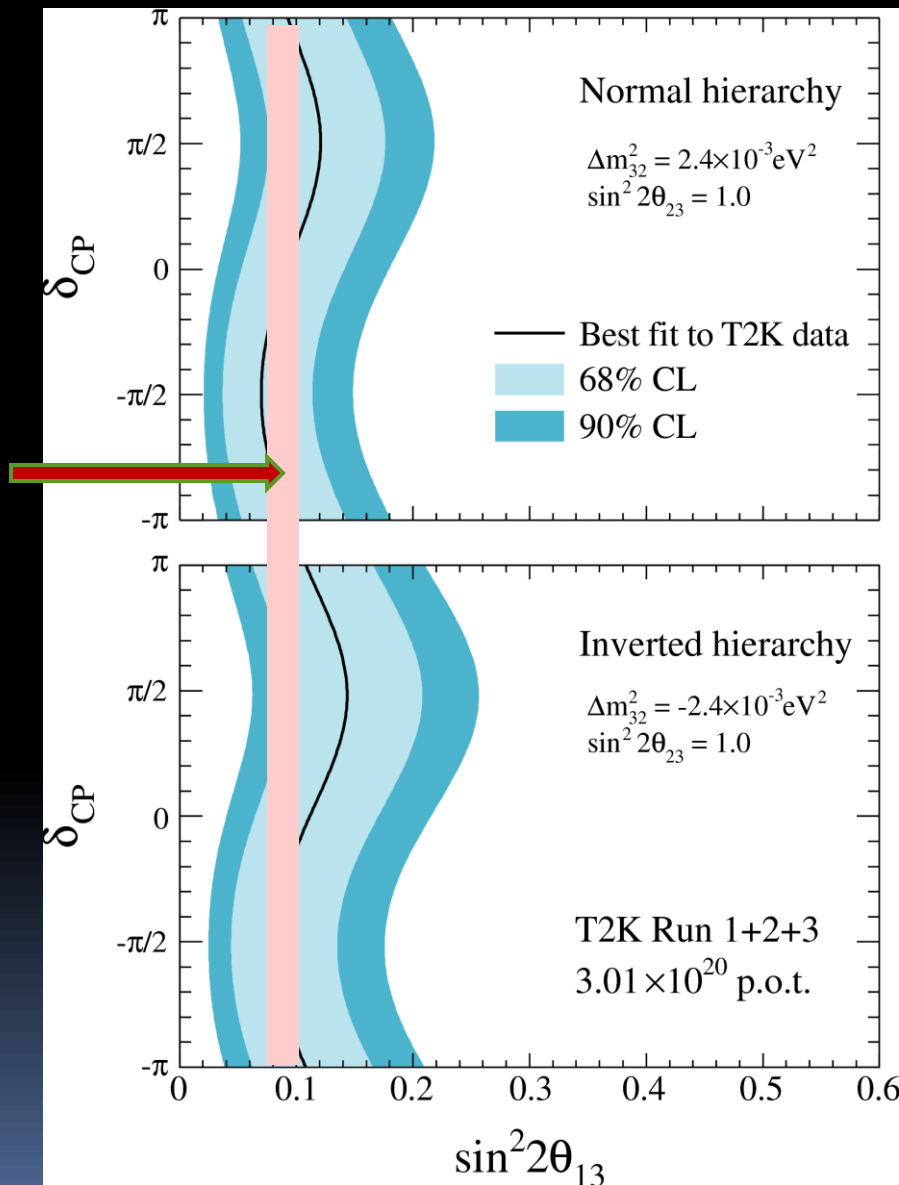
Probability (p-value) to observe 11 events for $\theta_{13} = 0 \rightarrow 0.08\%$

Event Category	$\sin^2 2\theta_{13} = 0$	$\sin^2 2\theta_{13} = 0.1$
Total	3.22 ± 0.43	10.71 ± 1.10
ν_e Signal	0.18	7.79
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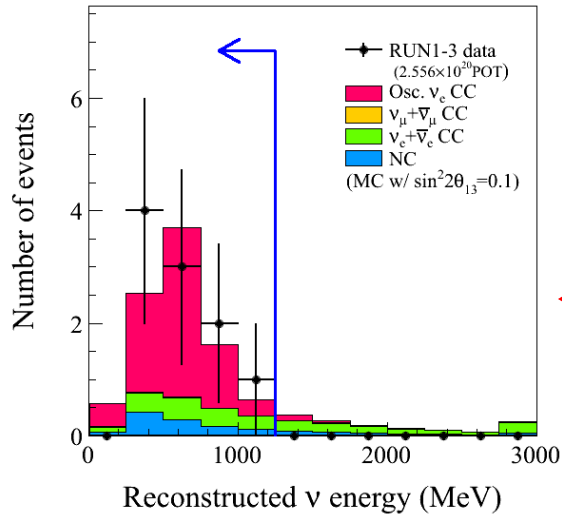
θ_{13} and δ

Daya Bay





T2K

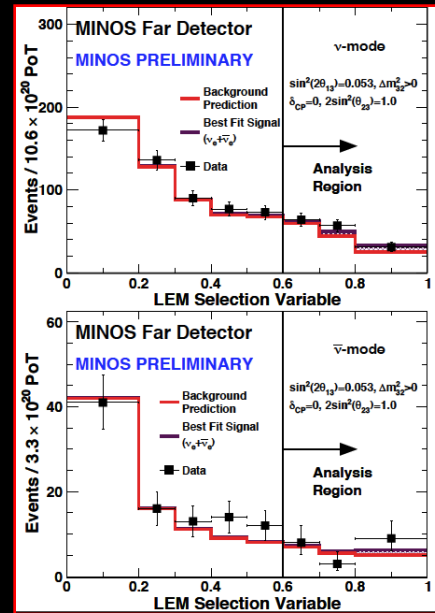


$$\theta_{13} \neq 0$$

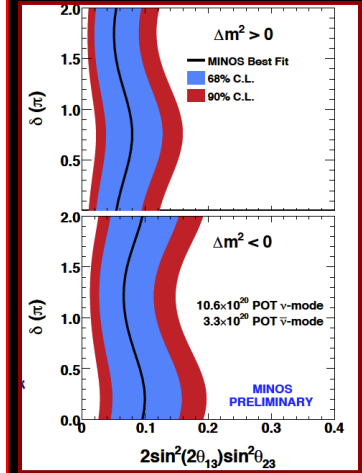
Accelerator experiments



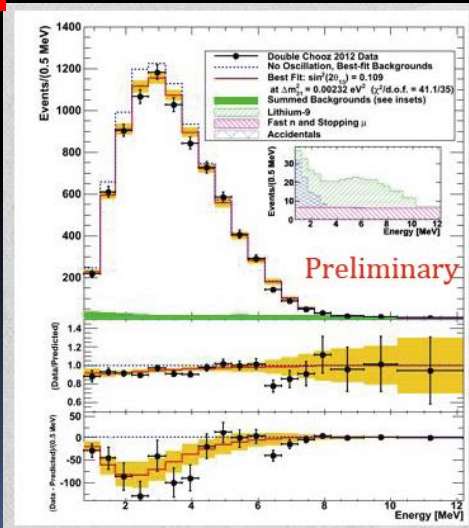
Reactor experiments



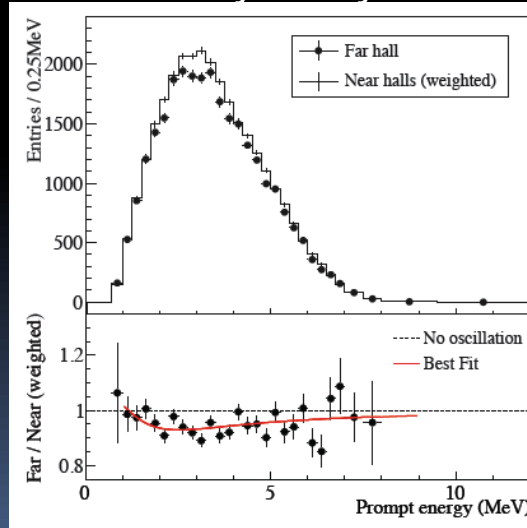
MINOS



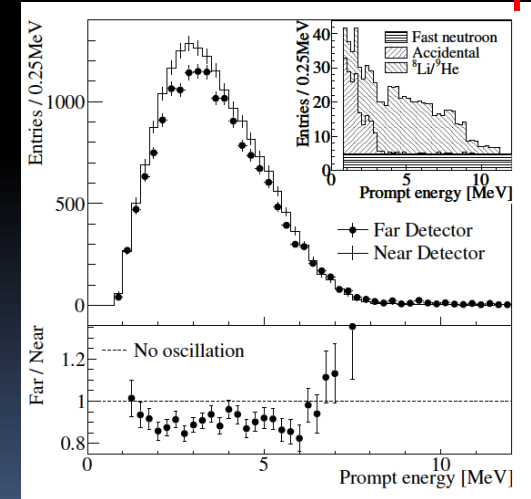
DChooz



Daya Bay



Reno



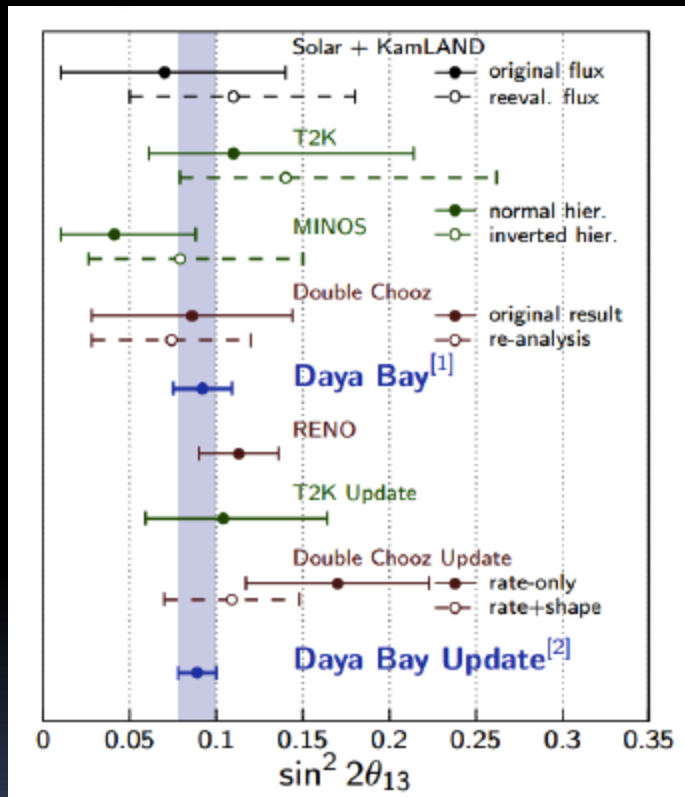


θ_{13} : one year story

from upper limit to precise measurement !

➤ θ_{13} has been well measured by 5 experiments

R.Barbiery ICHEP2012



$\theta_{13} \approx 9 \pm 1 \text{ deg}$



T2K $\nu_{\mu} \rightarrow \nu_e$

first indication of $\theta_{13} \neq 0$ 3.1 σ significance

MINOS $\nu_{\mu} \rightarrow \nu_e$

$\theta_{13} = 0$ disfavored at 1.7 σ

Double Chooz $\bar{\nu}_e \rightarrow \bar{\nu}_e$

$\theta_{13} \neq 0$ at 3.1 σ

Daya Bay $\bar{\nu}_e \rightarrow \bar{\nu}_e$

$\sin^2 2\theta_{13} = 0.089 \pm 0.010(\text{stat}) \pm 0.005(\text{syst})$
~ 8 σ significance

RENO $\bar{\nu}_e \rightarrow \bar{\nu}_e$

$\sin^2 2\theta_{13} = 0.113 \pm 0.013(\text{stat}) \pm 0.019(\text{syst})$
4.9 σ significance



T2K: ν_μ disappearance

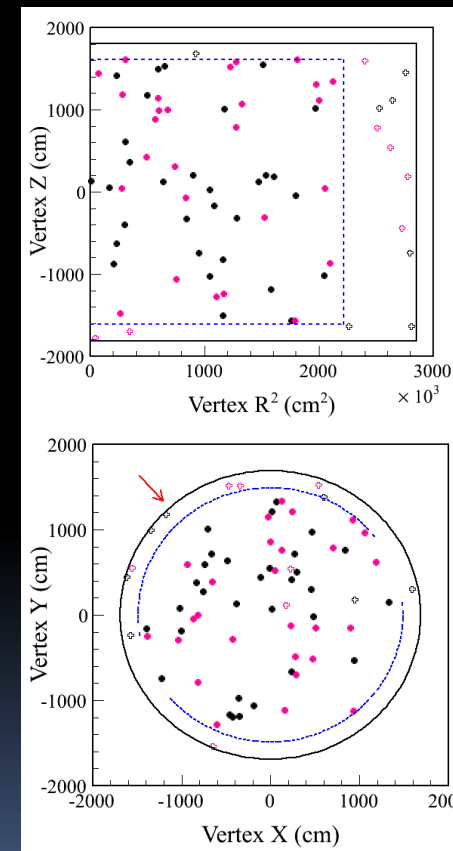
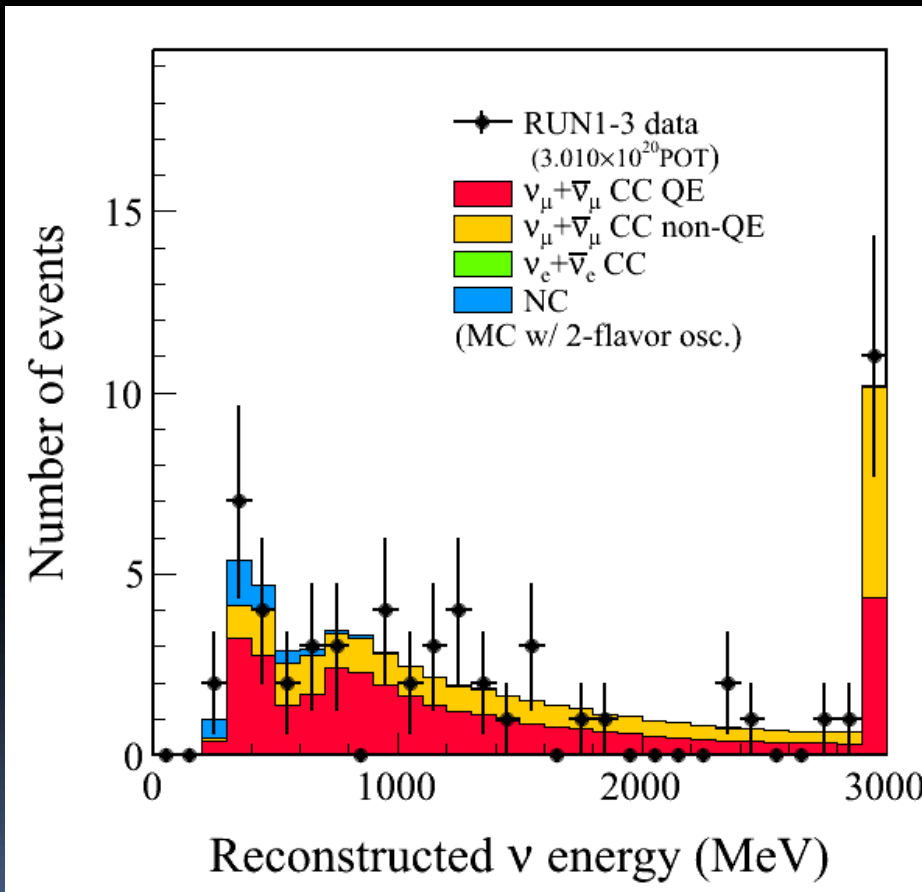


3.1×10^{20} POT

Measurement: 58 events observed

Monte Carlo: 196.2 events no oscillations

Monte Carlo: 57.8 events with oscillations

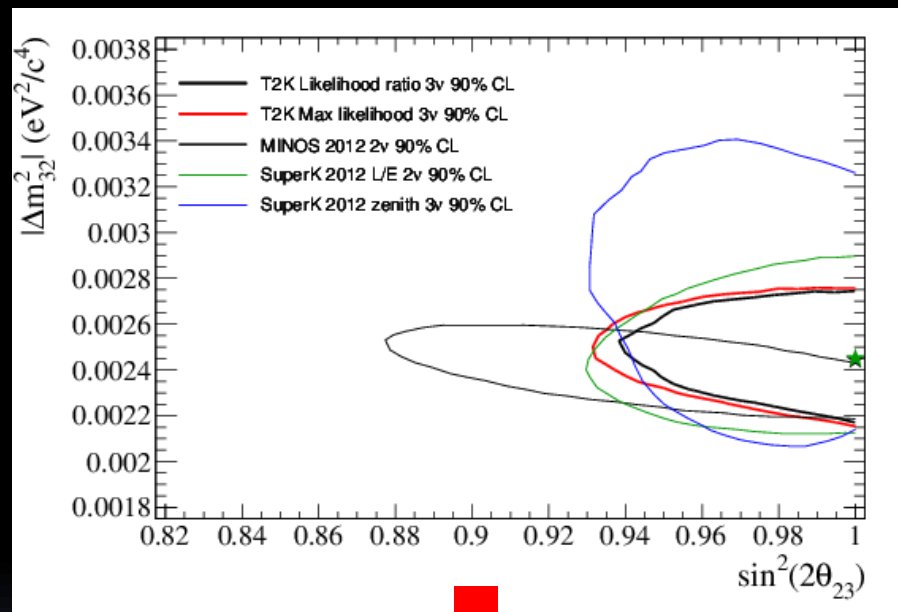
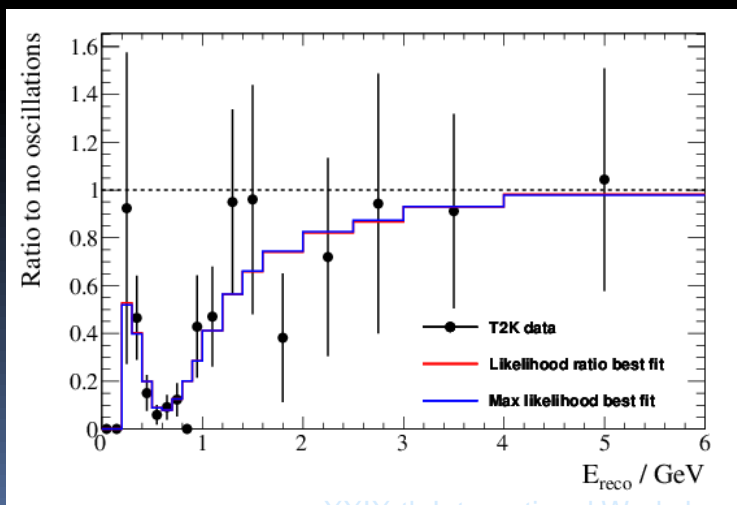
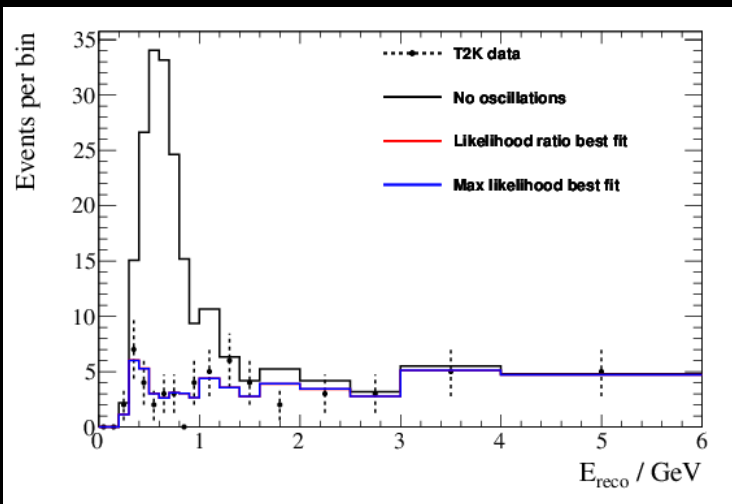




T2K: $\nu_\mu \rightarrow \nu_\mu$

Maximum Likelihood fit

Best fit results: $\sin^2 2\theta_{23} = 1.00$ $\Delta m_{32}^2 = 2.45 \times 10^{-3} \text{eV}^2$



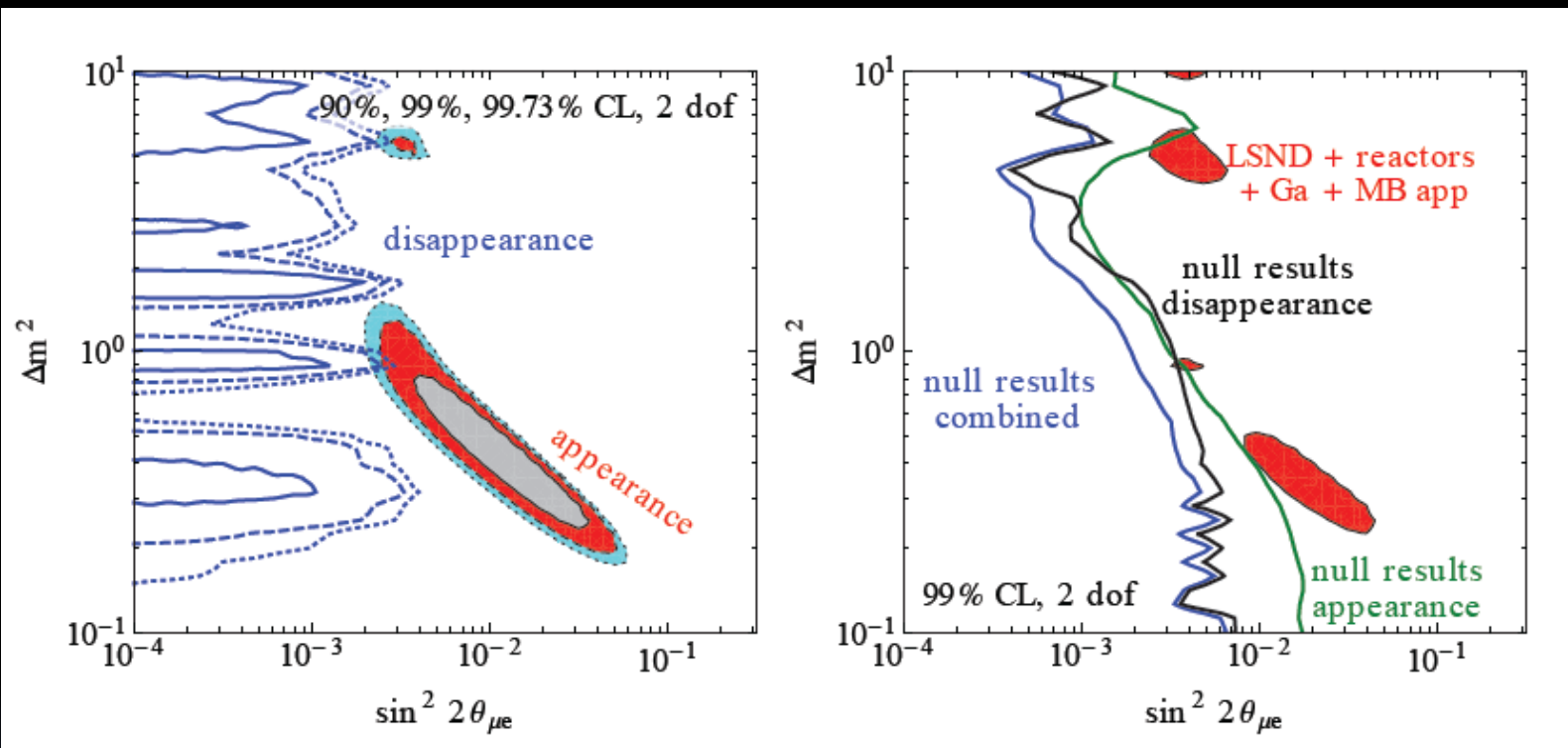
T2K obtained best sensitivity to mixing angle θ_{23}



Sterile neutrinos

MiniBooNe/LSND, reactor, Ga anomalies

Kopp, Machado, Maltoni, Schwetz, 1303.3011



strong tension between appearance and disappearance data



tension between signals and negative results



$\nu_\mu \rightarrow \nu_e$ in matter

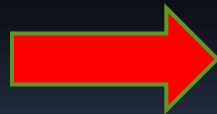
$$\begin{aligned}
P(\nu_\mu \rightarrow \nu_e) = & 4c_{13}^2 \boxed{s_{13}^2} s_{23}^2 \sin^2 \frac{\Delta m_{13}^2 L}{4E_\nu} \times \left[1 + \frac{2a}{\Delta m_{13}^2} (1 - 2s_{13}^2) \right] & \longrightarrow \theta_{13} \\
& + 8c_{13}^2 s_{12} s_{13} s_{23} (c_{12} c_{23} \cos \delta - s_{12} s_{13} s_{23}) \cos \frac{\Delta m_{23}^2 L}{4E_\nu} \sin \frac{\Delta m_{13}^2 L}{4E_\nu} \sin \frac{\Delta m_{12}^2 L}{4E_\nu} & \longrightarrow \text{CP-even} \\
& - 8c_{13}^2 c_{12} c_{23} s_{12} s_{13} s_{23} \sin \delta \sin \frac{\Delta m_{23}^2 L}{4E_\nu} \sin \frac{\Delta m_{13}^2 L}{4E_\nu} \sin \frac{\Delta m_{12}^2 L}{4E_\nu} & \longrightarrow \text{CP-odd} \\
& + 4s_{12}^2 c_{13}^2 (c_{13}^2 c_{23}^2 + s_{12}^2 s_{23}^2 s_{13}^2 - 2c_{12} c_{23} s_{12} s_{23} s_{13} \cos \delta) \sin^2 \frac{\Delta m_{12}^2 L}{4E_\nu} & \longrightarrow \text{Solar} \\
& - 8c_{13}^2 s_{13}^2 s_{23}^2 \cos \frac{\Delta m_{23}^2 L}{4E_\nu} \frac{aL}{4E_\nu} \sin \frac{\Delta m_{13}^2 L}{4E_\nu} (1 - 2s_{13}^2), & \longrightarrow \text{Matter} \quad (30)
\end{aligned}$$

$$s_{ij} = \sin \theta_{ij}$$

$$c_{ij} = \cos \theta_{ij}$$

$$a [eV^2] = 2\sqrt{2} G_F n_e E_\nu = 7.6 \times 10^{-5} \rho \left[\frac{g}{cm^3} \right] E_\nu [GeV]$$

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$$



$$a \rightarrow -a \quad \delta \rightarrow -\delta$$

change sign for NH \rightarrow IH



J_{CP}

$$\theta_{13} \neq 0$$

The strength of CP violation in neutrino oscillations

Jarlskog invariant J_{CP}

$$J_{CP} = \text{Im}(U_{e1}U_{\mu2}U_{e2}^*U_{\mu1}^*) = \text{Im}(U_{e2}U_{\mu3}U_{e3}^*U_{\mu2}^*) = \\ = \cos\theta_{12}\sin\theta_{12}\cos^2\theta_{13}\sin\theta_{13}\cos\theta_{23}\sin\theta_{23}\sin\delta$$

all mixing angles $\neq 0 \rightarrow J_{CP} \neq 0$ if $\delta \neq 0$

Quark sector $J_{CP} \approx 3 \times 10^{-5}$

Lepton sector $J_{CP} \sim 0.02 \times \sin\delta$

neutrinos

$$V_{MNS} \sim \begin{pmatrix} 0.8 & 0.5 & 0.2 \\ 0.4 & 0.6 & 0.7 \\ 0.4 & 0.6 & 0.7 \end{pmatrix}$$

quarks

$$V_{CKM} \sim \begin{pmatrix} 1 & 0.2 & 0.001 \\ 0.2 & 1 & 0.01 \\ 0.001 & 0.01 & 1 \end{pmatrix}$$

Real chance to test CP violation in neutrino oscillations



CP measurements

If $\theta_{13} \neq 0$ and not too small



measurement of δ
in LBL accelerator experiments

(1) ν and anti- ν narrow beams tuned to 1st oscillation maximum

$$A_{CP} = \frac{P(\nu_{\mu} \rightarrow \nu_e) - P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e)}{P(\nu_{\mu} \rightarrow \nu_e) + P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e)} \simeq \frac{\Delta m_{12}^2 L}{4E_{\nu}} \cdot \frac{\sin 2\theta_{12}}{\sin \theta_{13}} \cdot \sin \delta$$

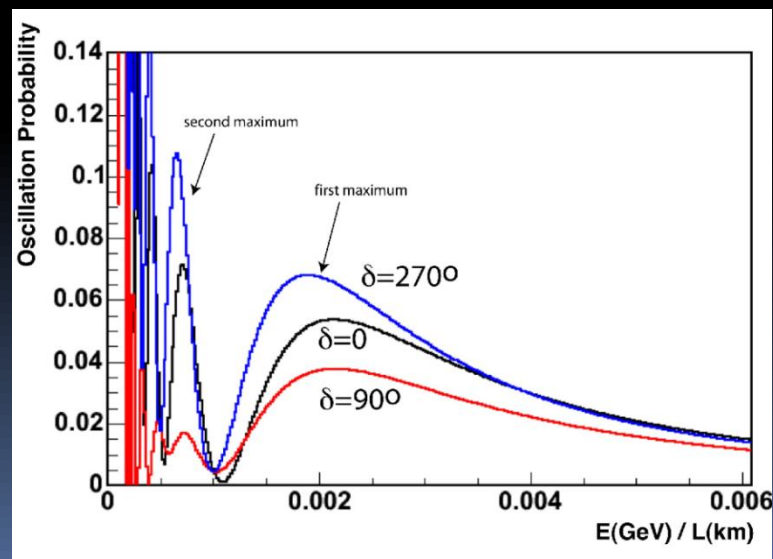
$$P(\nu_{\mu} \rightarrow \nu_e) \sim \sin^2 \theta_{13}$$

- neutrino and antineutrino beams
- massive far detector

(2) wide ν (anti- ν) beam to cover
1st and 2nd oscillation maxima

- wide energy muon neutrino beam
- measurements of two oscillation maxima

$P(\nu_{\mu} \rightarrow \nu_e)$ for $\sin^2 2\theta_{13}=0.1$

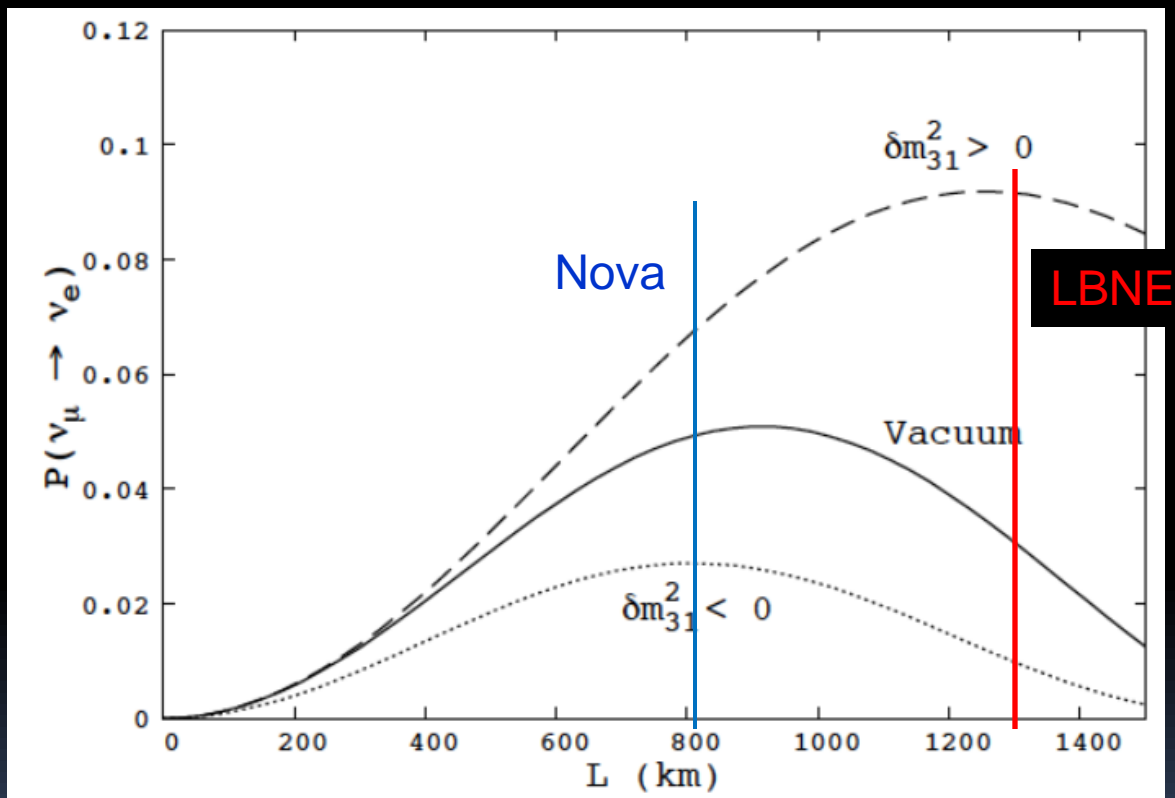


LBL experiments: near and far future



Matter effect in LBL experiments

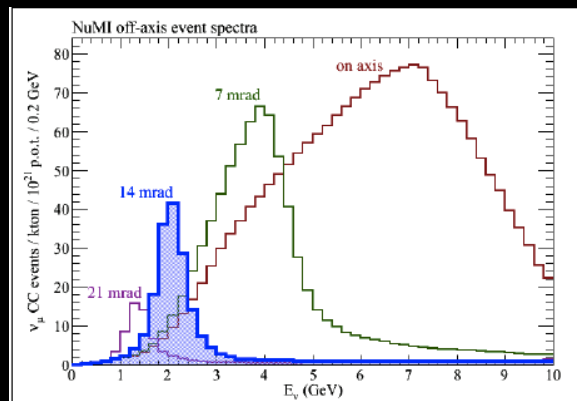
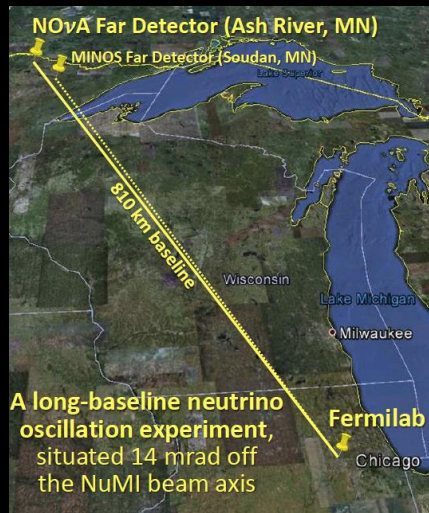
$E_\nu \sim 2 \text{ GeV}$



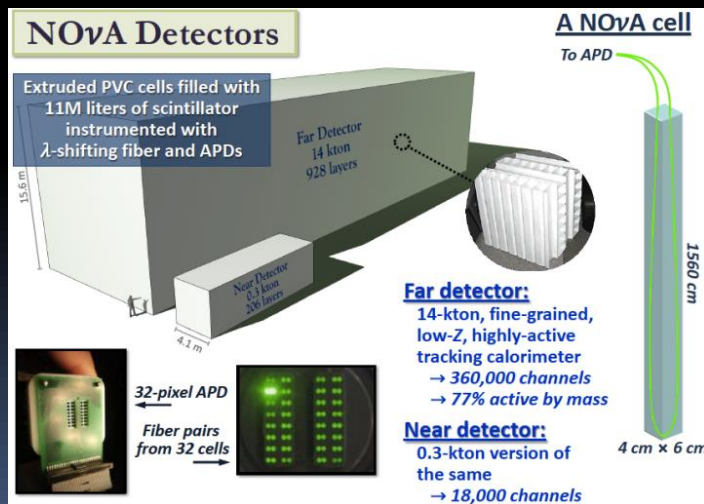
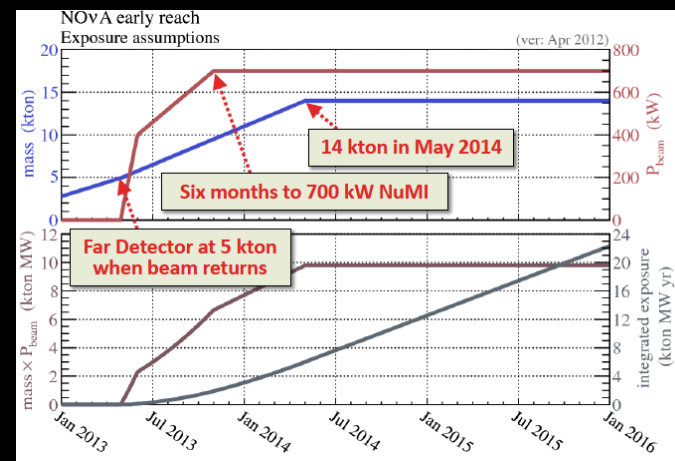


Nova

Neutrino off-axis narrow-band beam from FNAL
 $L = 810$ km, $E \sim 2$ GeV, 700 kW beam power



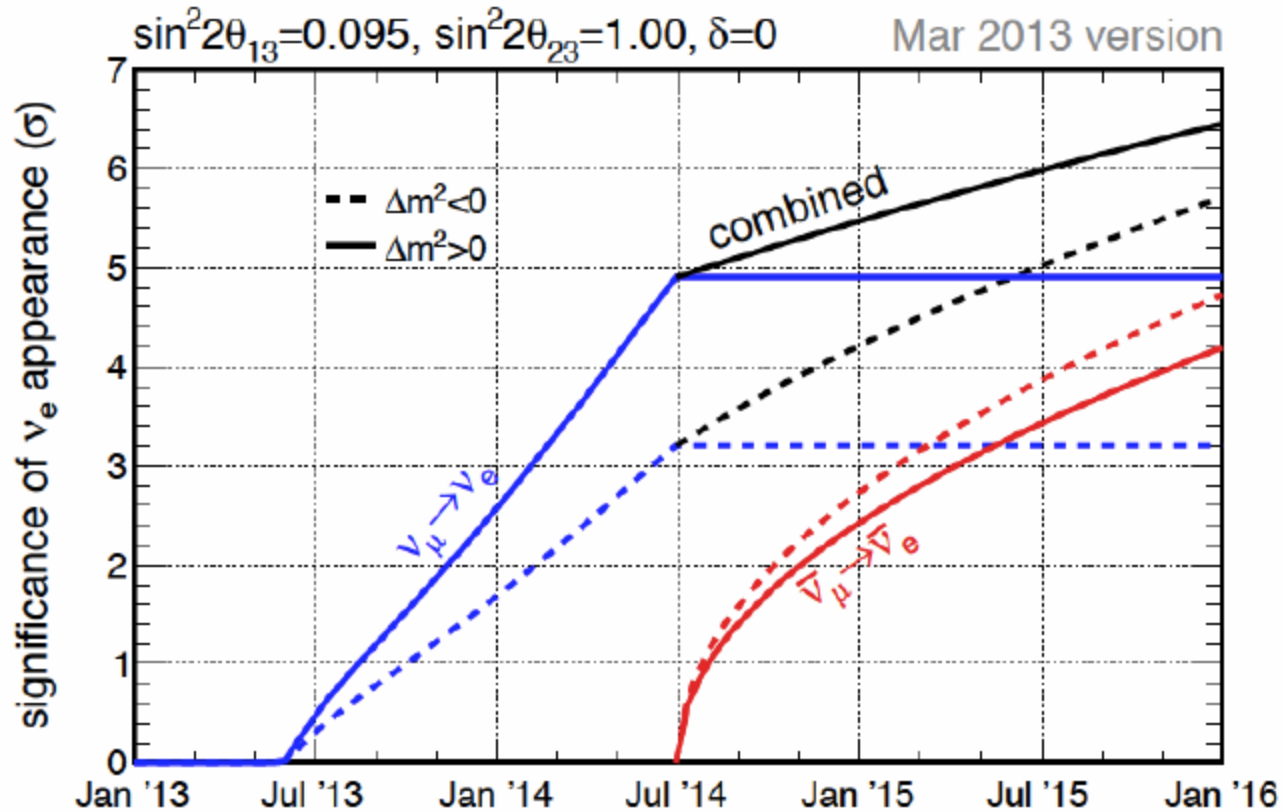
R.Patterson, Neutrino 2012



Experiment will start data taking in late 2013

Nova: early sensitivity

M.Messier talk Prague CP violation Colloquium, May 2013

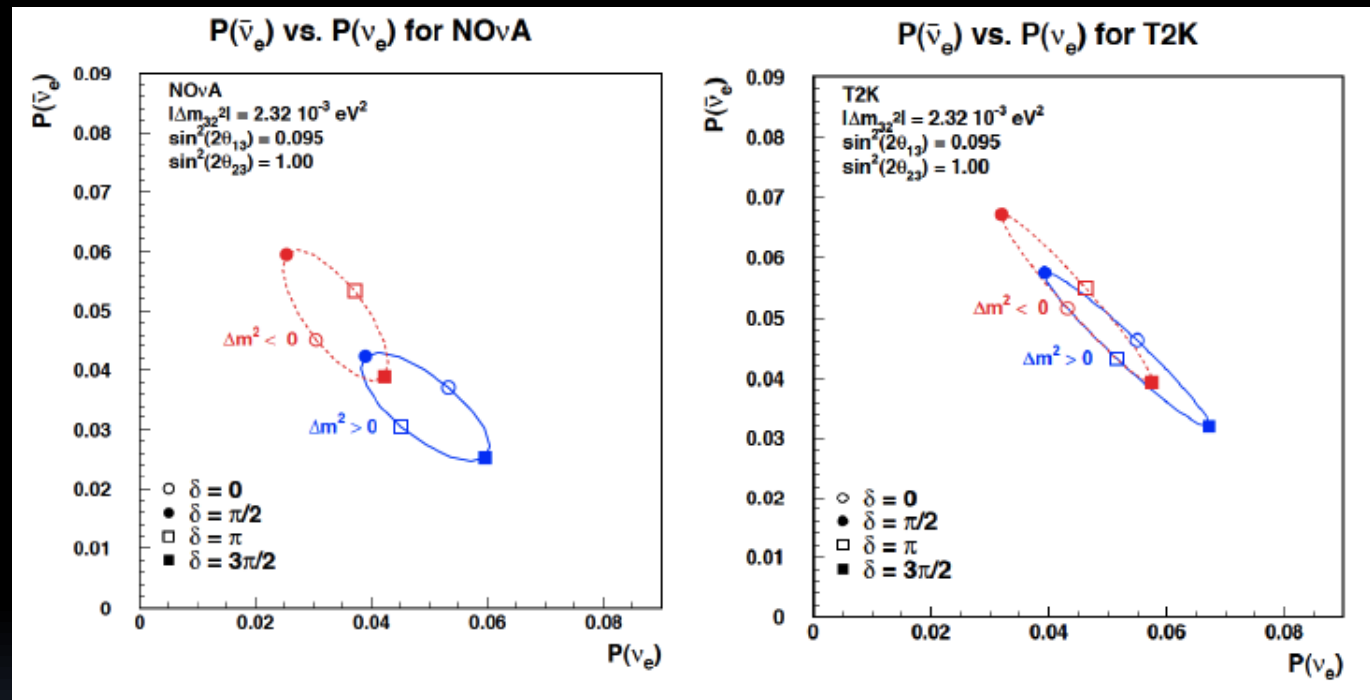




T2K and Nova

C.Backhouse, NNN Workshop, 4-6 October 2012

Possible measurement of mass hierarchy and CP violation



For $\sin^2 2\theta_{13}=0.1$, approximately (at 90% C.L.):

- MH: $\approx 50\%$ coverage
- CPV: $\approx 30\text{-}40\%$ coverage



LBNE

$$\nu_{\mu} \rightarrow \nu_e$$

The US based LBL project

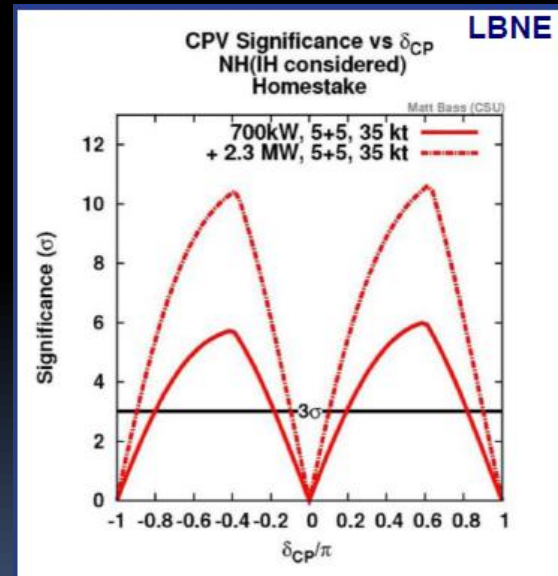
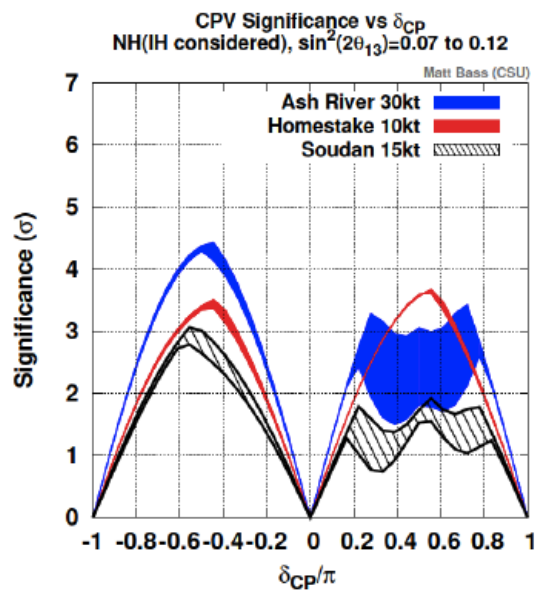
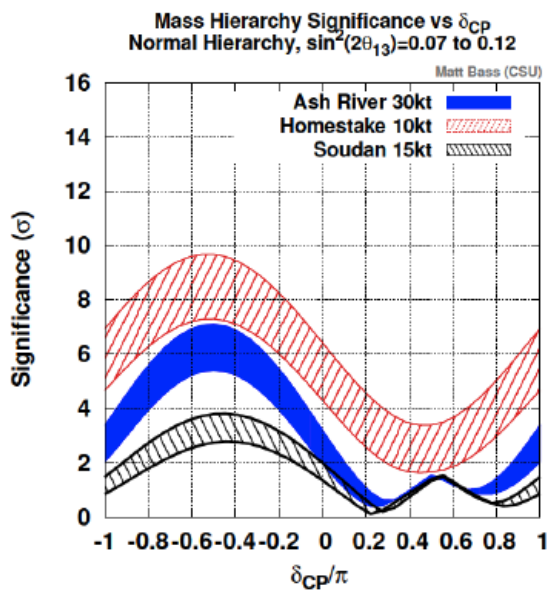
Neutrino beam from FNAL to Homestake

$L = 1300$ km, $E_p = 120$ GeV, 700 kW NuMI beam,

$E_{\nu} = 0.5 - 5$ GeV

Far detector 10 kt LAr TPC, on surface
No near detector

Sensitivity to MH and CP phase

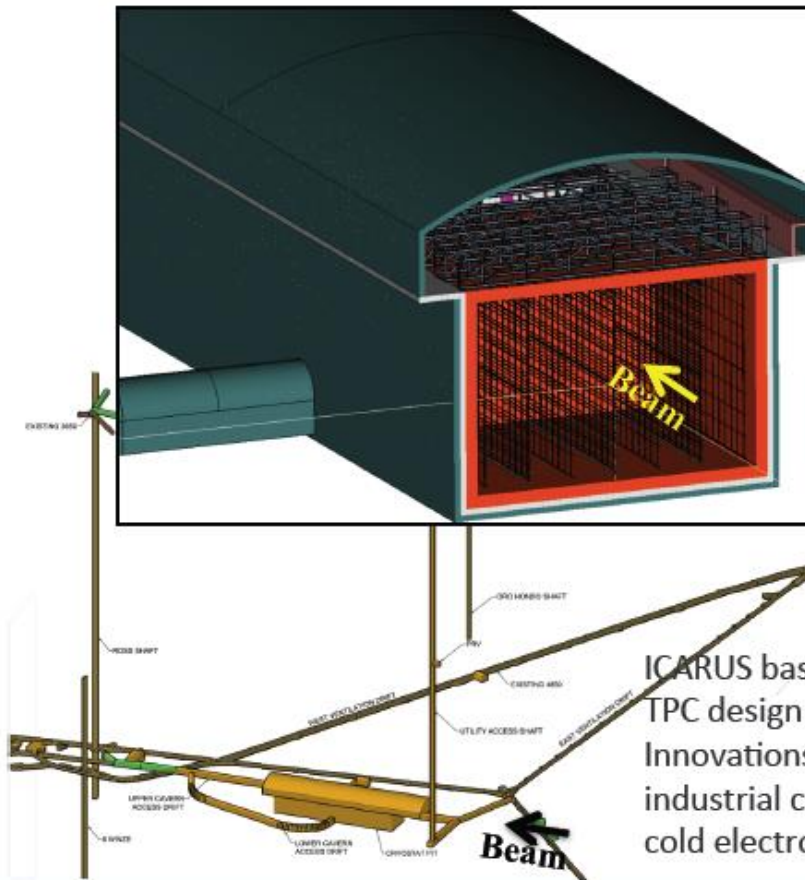




LBNE: Far Detector concept

M.Diwan, talk at ISOUP13

Later phase: 34 kt LAr TPC underground

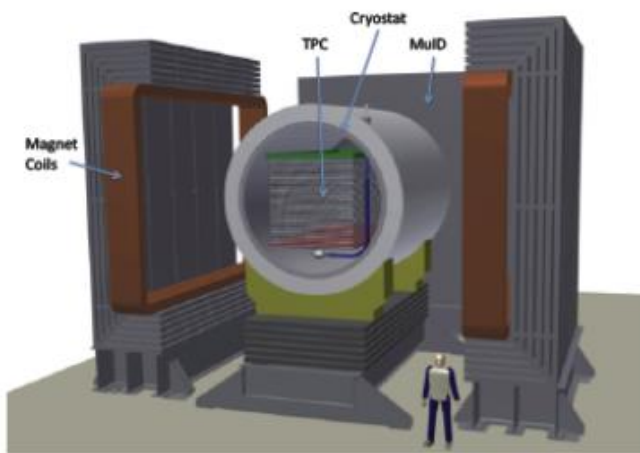


- Two detectors in a common cavern at 4850 ft. depth
- Active volume of each detector:
 $22.4 \times 14 \times 45.6 \text{ m}^3$
- 34 kt fiducial mass
- TPC design:
 - 3.7 m drift length
 - 5 mm wire spacing
 - three stereo views
 - 2X108 anode chambers
 - 2 X 275k channels
 - S/N ~ 10

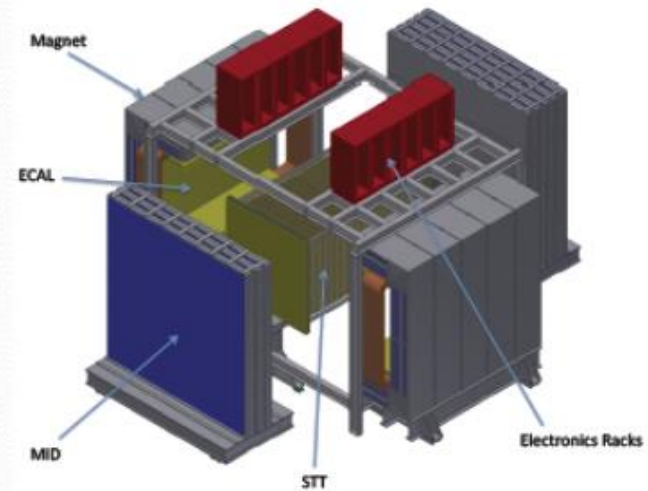
ICARUS based
TPC design
Innovations:
industrial cryostat,
cold electronics



LBNE: Near Detector options



Liquid Argon TPC Tracker
(~18 ton LAr)



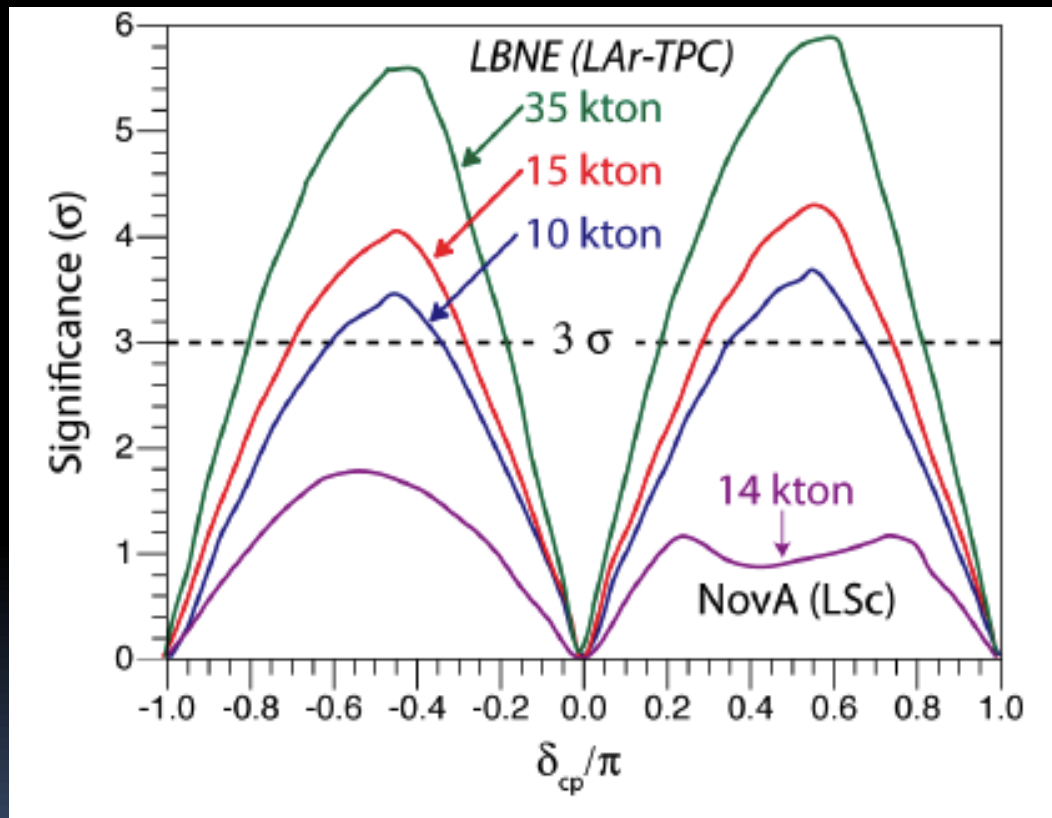
Fine-Grained Tracker
(~0.14 ton Ar@140atm)



LBNE: sensitivity to δ

arXiv:1110.6249

Several options for LBNE far detector





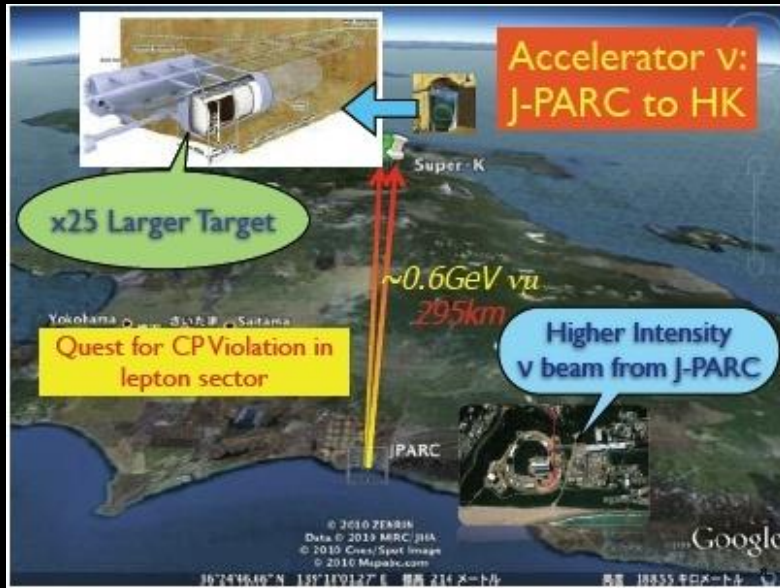
T2HK

$$\nu_{\mu} \rightarrow \nu_e$$

arXiv:1109.3262 [hep-ex]

The LBL project in Japan

T.Nakaya, NNN2012



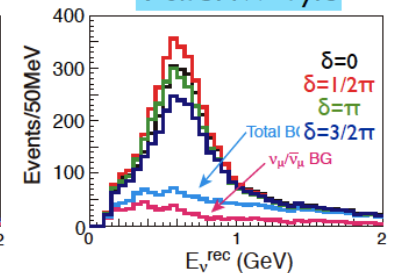
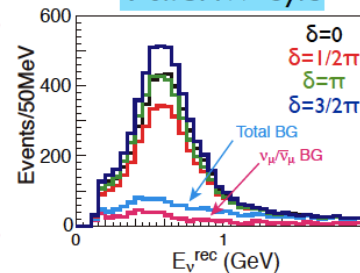
Example: case of Hyper-K

$$\sin^2 2\theta_{13} = 0.1$$

ν 0.75MW \times 3yrs

$\bar{\nu}$ 0.75MW \times 7yrs

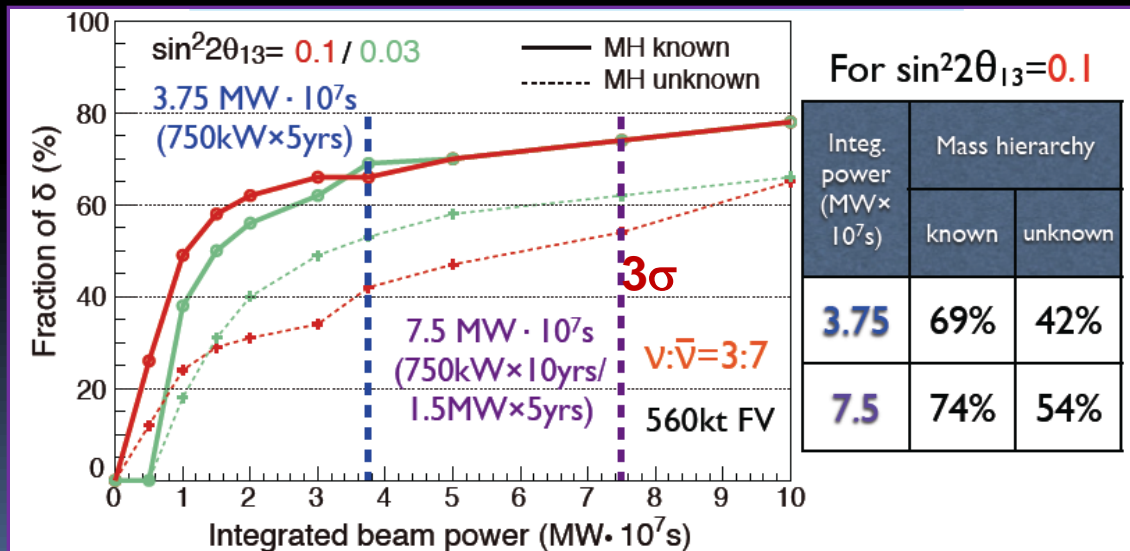
ν_e candidates
(full simulation)



Expected number of ν_e candidate events for $\delta=0$

$\nu_{\mu} \rightarrow \nu_e$ CC	3606
$\nu_{\mu} + \nu_{\mu}$ CC	35
$\nu_e + \nu_e$ CC	880
NC	649

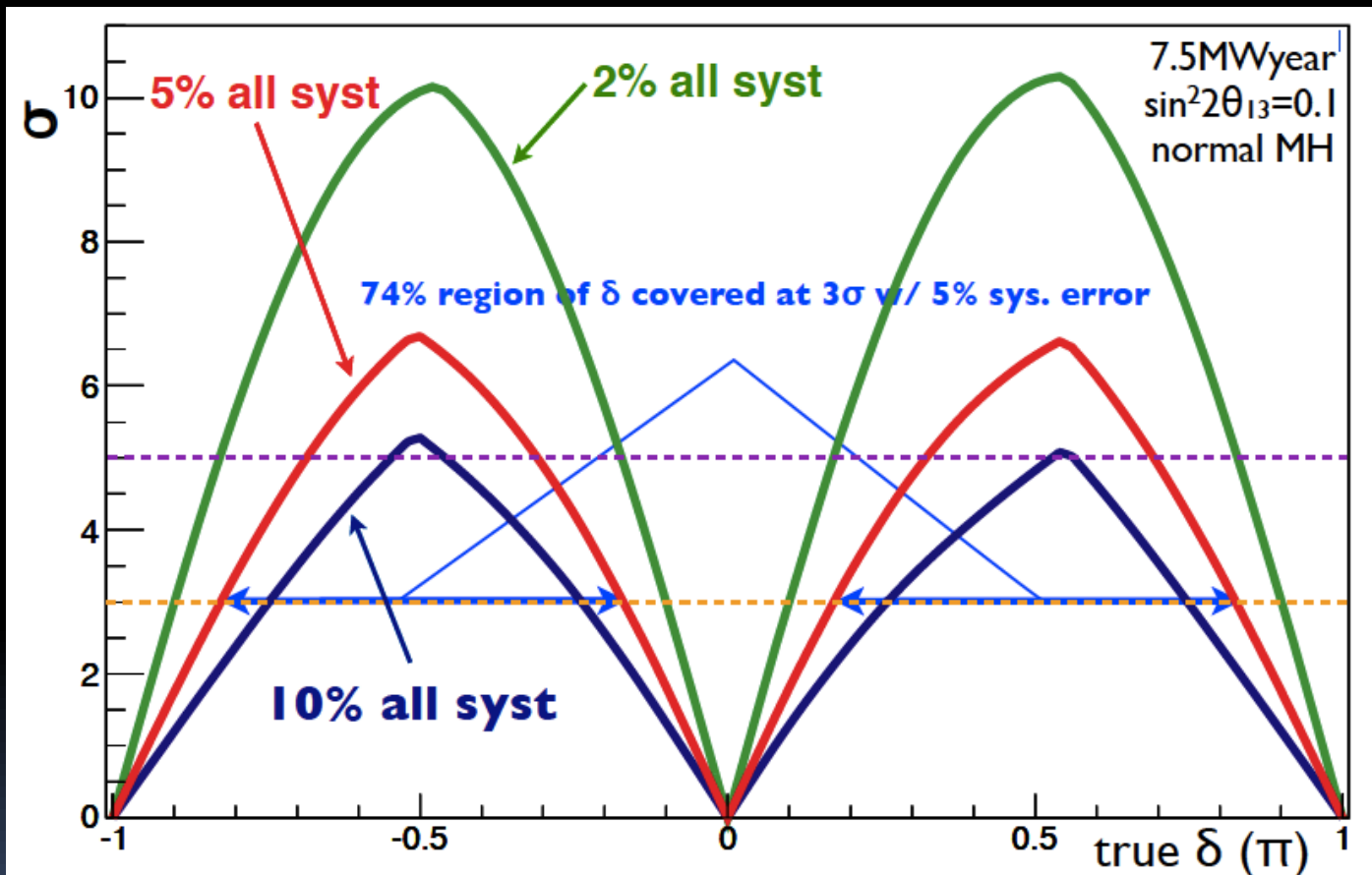
$\nu_{\mu} \rightarrow \nu_e$ CC	2339
$\nu_{\mu} + \nu_{\mu}$ CC	23
$\nu_e + \nu_e$ CC	878
NC	678





T2HK: CPV discovery potential

MH is known !



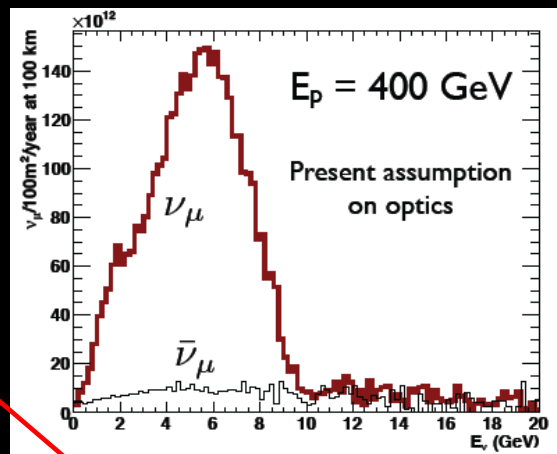
High sensitivity to CP phase for systematics $< 5\%$



LAGUNA-LBNO

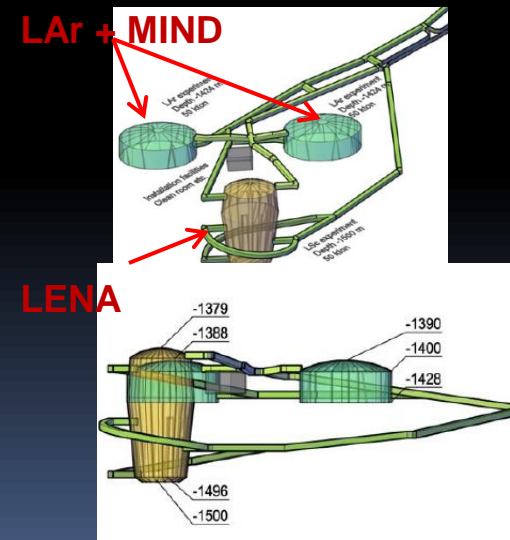
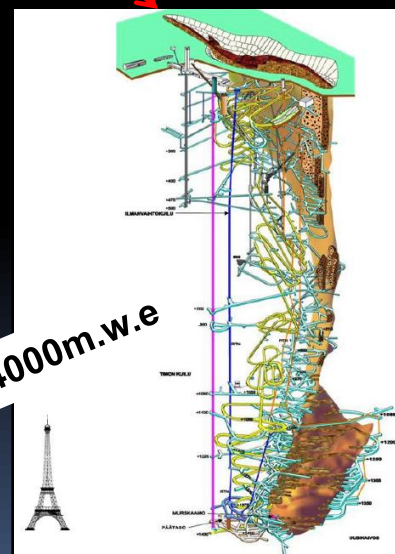
$$\nu_{\mu} \rightarrow \nu_e$$

Wide-band neutrino beam from CERN to Pyhasalmi (Finland)

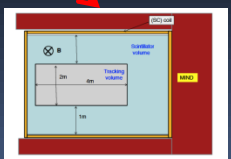


14 countries,
47 institutions
300 members
Russia:
INR
PNPI

Far neutrino detectors



Near detector



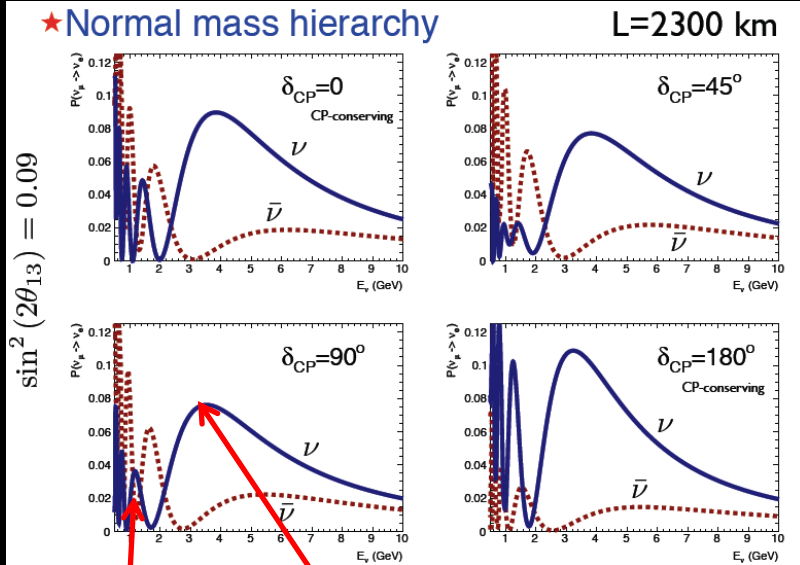


$P(\nu_\mu \rightarrow \nu_e)$

LAGUNA-LBNO

matter effect

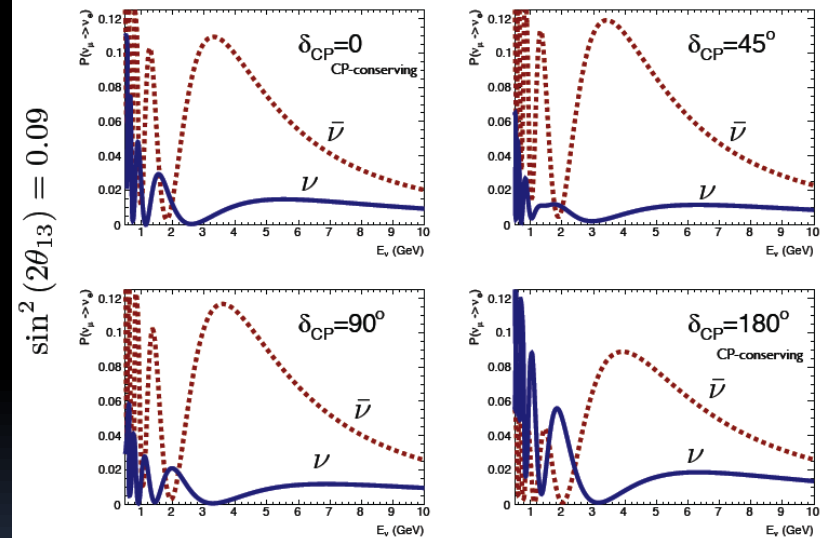
- easy to measure MH
- more difficult CP violation



2nd maximum

1st maximum

★ Inverted mass hierarchy L=2300 km



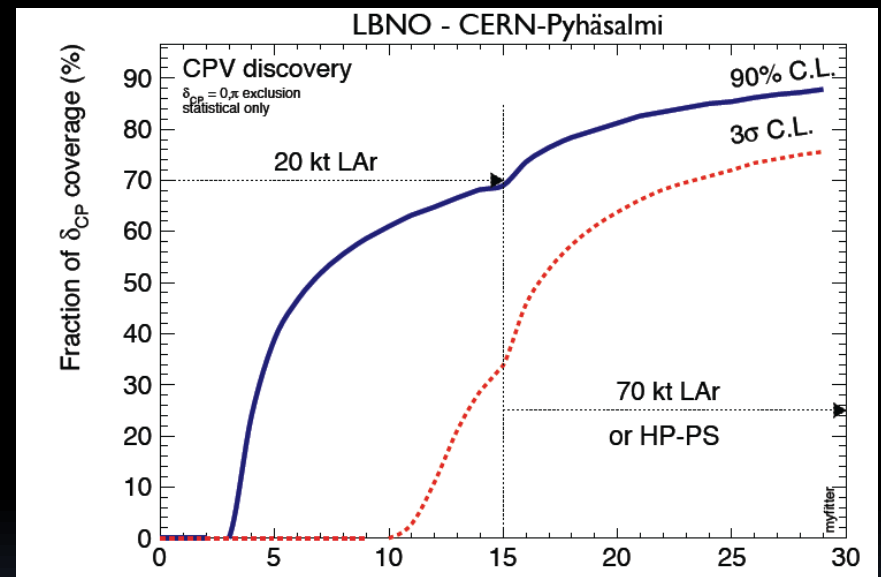
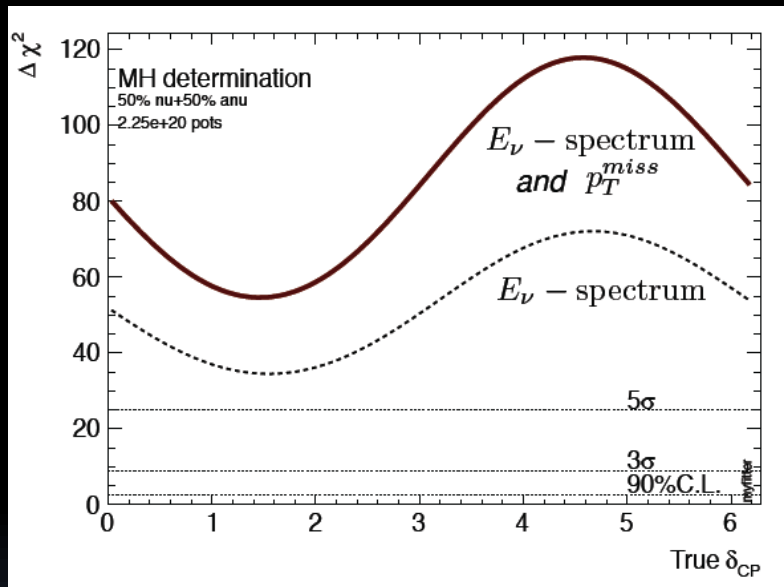


Sensitivity

LAGUNA-LBNO

MH: 100% at $>5\sigma$

CPV: $\sim 60\%$ at 3σ



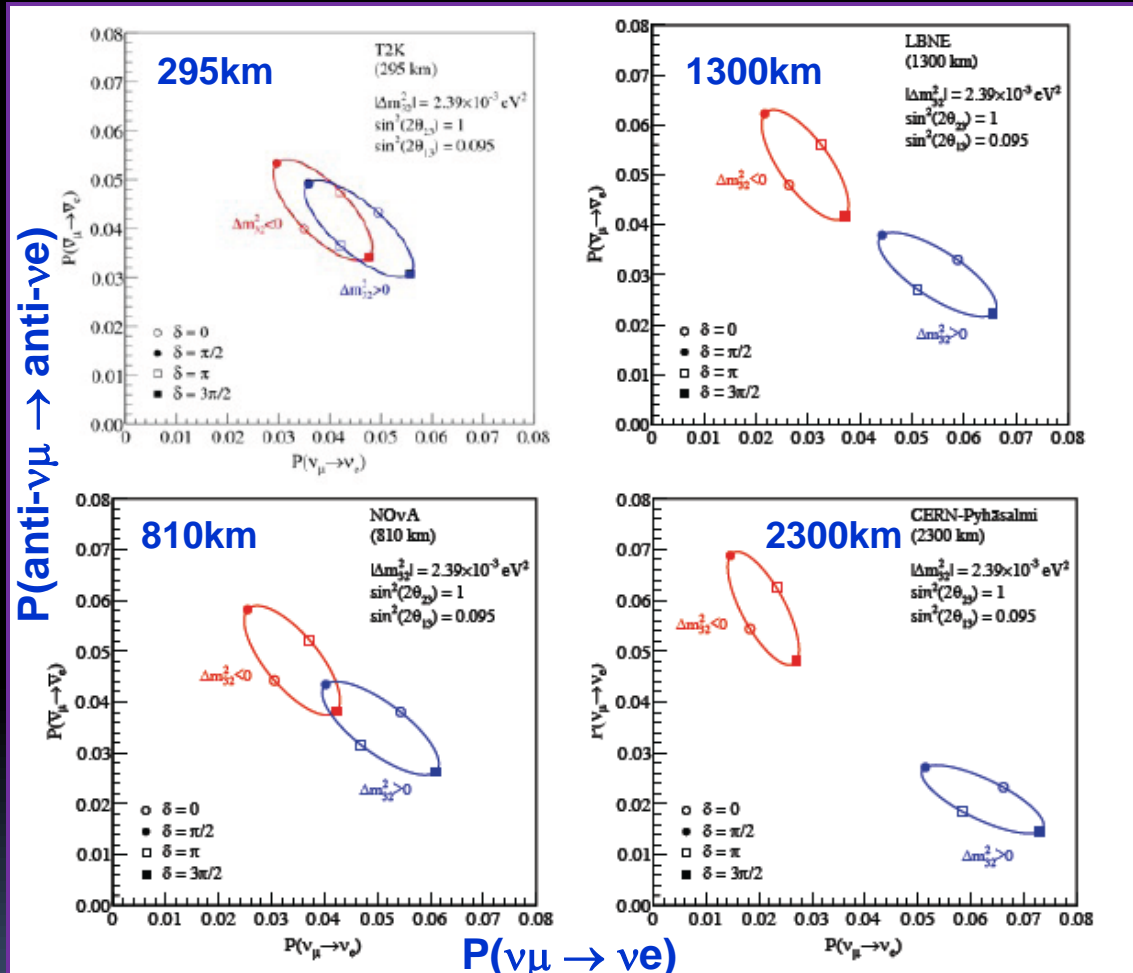
CERN SPS 400 GeV POT($\times 10^{20}$)



δ and MH

LBL experiments

S.Wojcicki, NNN2012





Conclusion

- 3 neutrino mixing angles are measured and non-zero
- Large θ_{13} opens door for searching of CP-violation in lepton sector
- Time to start θ_{12} and δ measurements