



KATRIN experiment: status and perspectives 2013

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1934: Neutrino mass could be observed in the nuclear β -decay

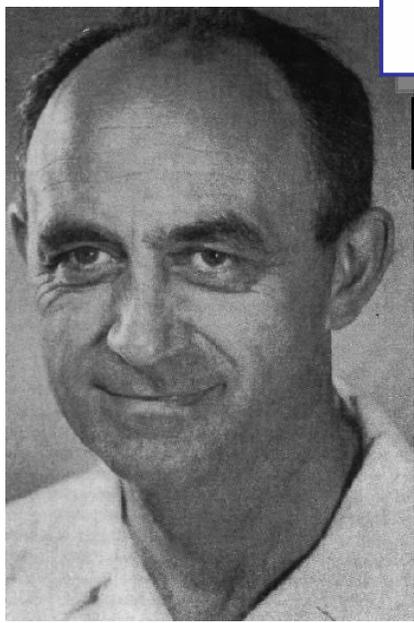


Versuch einer Theorie der β -Strahlen. I¹).

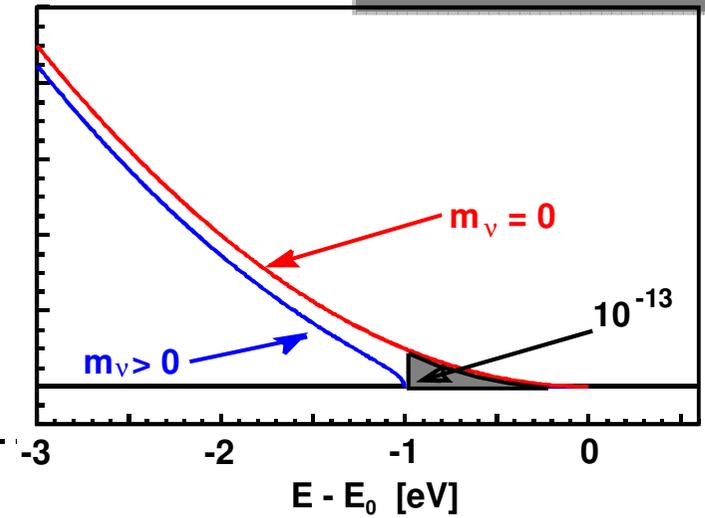
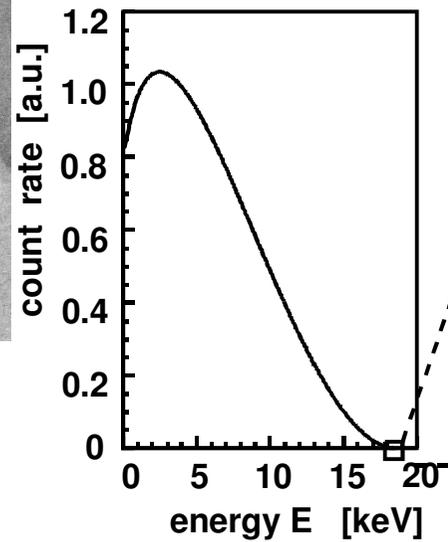
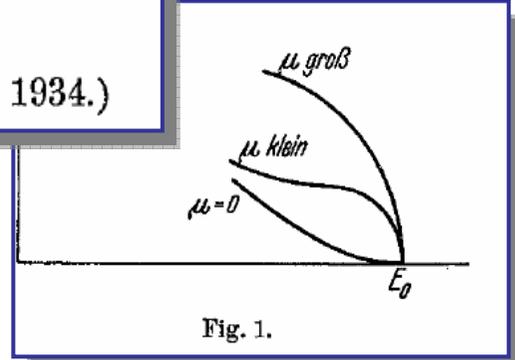
Von E. Fermi in Rom.

Mit 3 Abbildungen. (Eingegangen am 16. Januar 1934.)

E. Fermi, Z. Physik 88 (1934)



Enrico Fermi





1948: First successful experiment with Tritium



Бруно Понтекорво

Bruno Pontecorvo

$$m_\nu < 1 \text{ кеV}/c^2$$

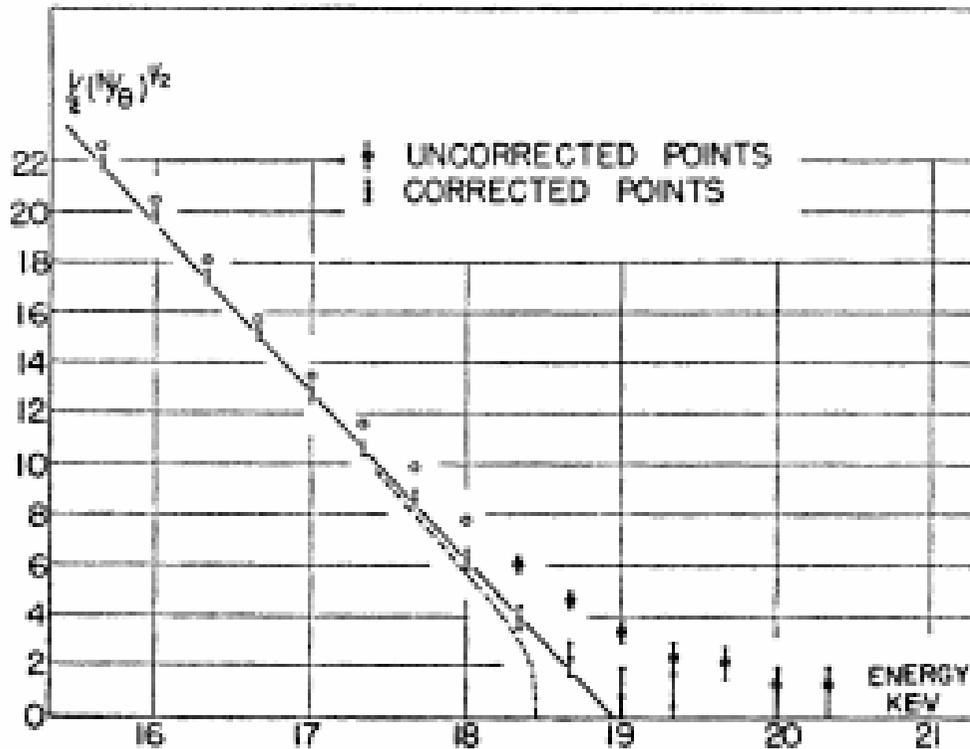


FIG. 2. "Kurie" plot of the end of the H^3 spectrum. The theoretical curve (shown dotted) corresponding to a finite neutrino mass of 500 eV (or 1 keV —see text) has been included for comparison.

Hanna G.C. and Pontecorvo B., Phys. Rev. 75 (1949) 983



1983: Electrostatic spectrometer with adiabatic magnetic collimation *Troitsk ν -mass experiment*



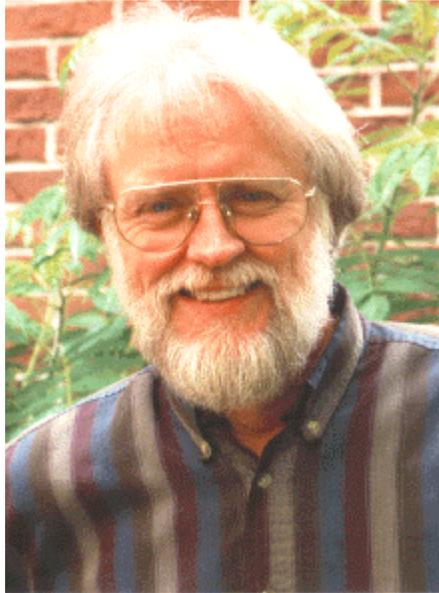
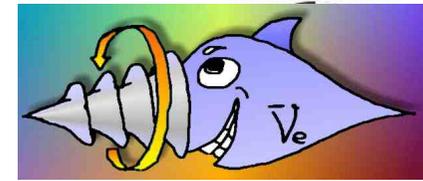
Petr Spivak
24.03.1911 - 30.03.1991



Vladimir Lobashev
29.07.1934 – 3.08.2011



“Great minds think alike”
Mainz Neutrino Mass Experiment



Robert B. Moore
Physics Department,
McGill University
Montreal, Canada



Ernst Otten
Physics Institute
Johannes Gutenberg University
Mainz, Germany



Jochen Bonn

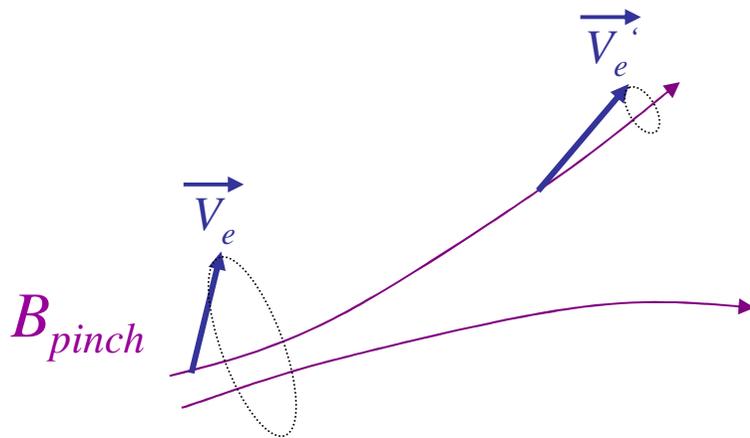


Electrostatic spectrometer with adiabatic magnetic collimation (MAC-E filter) Fundamentals



In the slowly varying magnetic field charged particle motion is *adiabatic*, (*first*) *adiabatic invariant* is conserved.

$$\mu = \frac{E_{per}}{B} = \frac{E_{tot} \cdot \sin^2 \theta}{B} \cong const$$

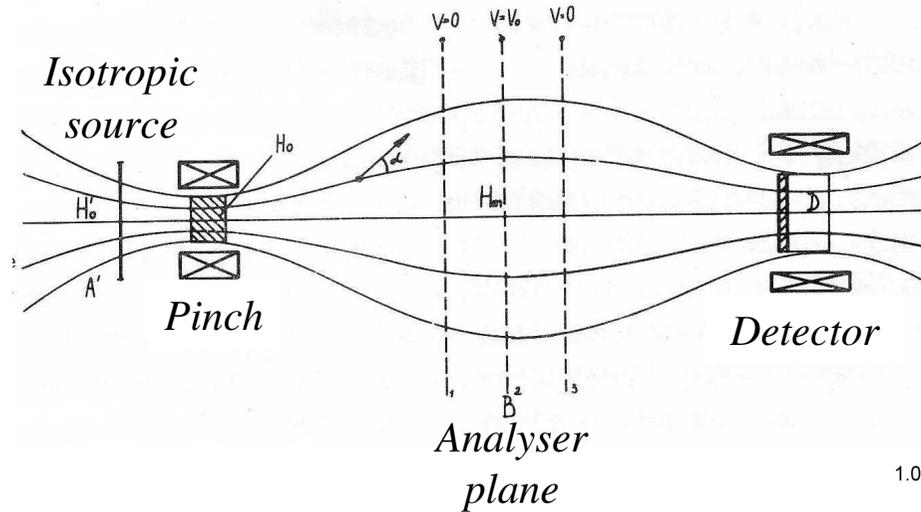


$$B_{analyser} \ll B_{pinch}$$

A particle velocity vector is aligned along the magnetic field direction during transition into weaker magnetic field

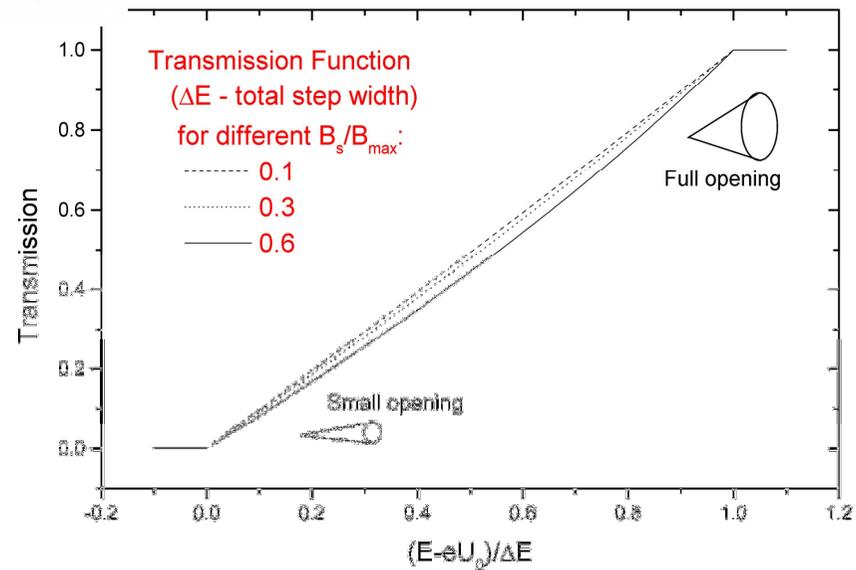


Electrostatic spectrometer with adiabatic magnetic collimation (MAC-E filter) Fundamentals



$$\Delta E = |eU_0| \frac{B_{analyser}}{B_{pinch}}$$

Spectrometer resolution
is decoupled
from the source size!





International project

KATRIN

(KARlsruhe TRItium Neutrino experiment)



Spectrometer \varnothing 10 meters

Resolution $\Delta E = 0,9$ eV (at 18 keV)

Windowless gaseous source \varnothing 90 mm

Column density $5 \cdot 10^{17}$ mol/cm²

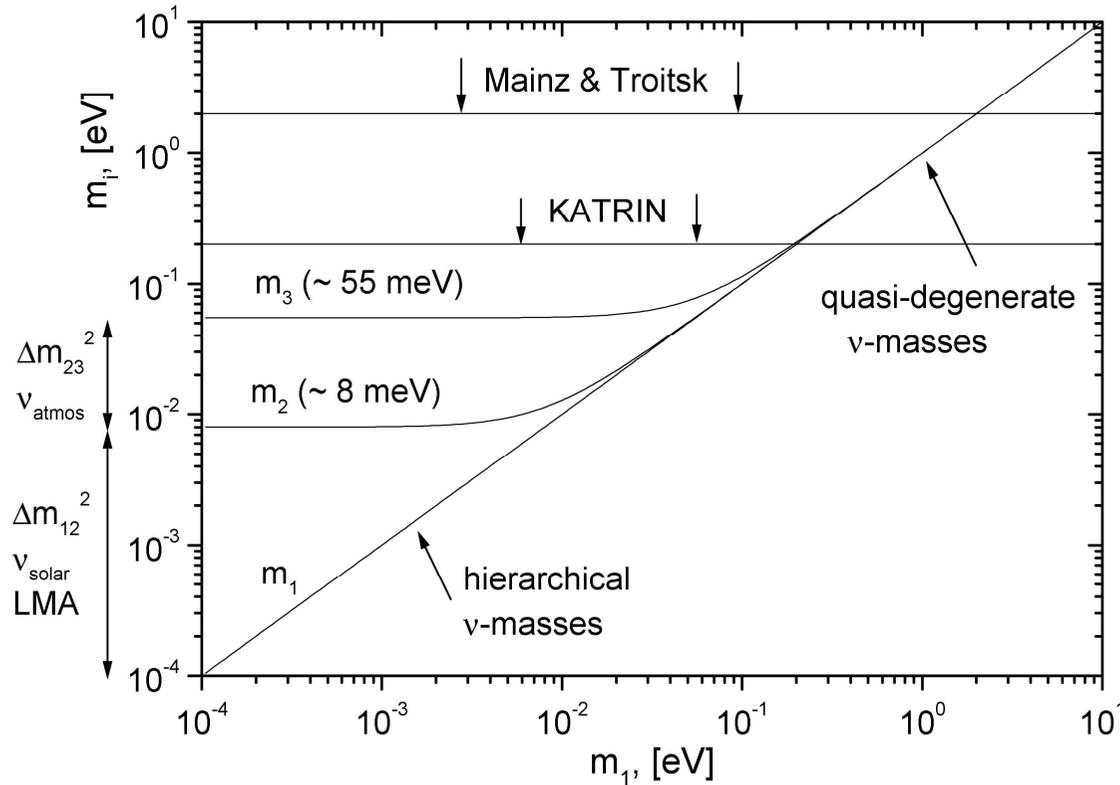
Neutrino mass sensitivity

(after 3 years of data taking):

$$m_\nu < 0,2 \text{ eV}/c^2$$



KATRIN Project

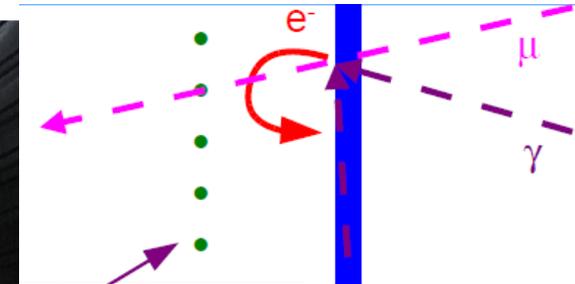
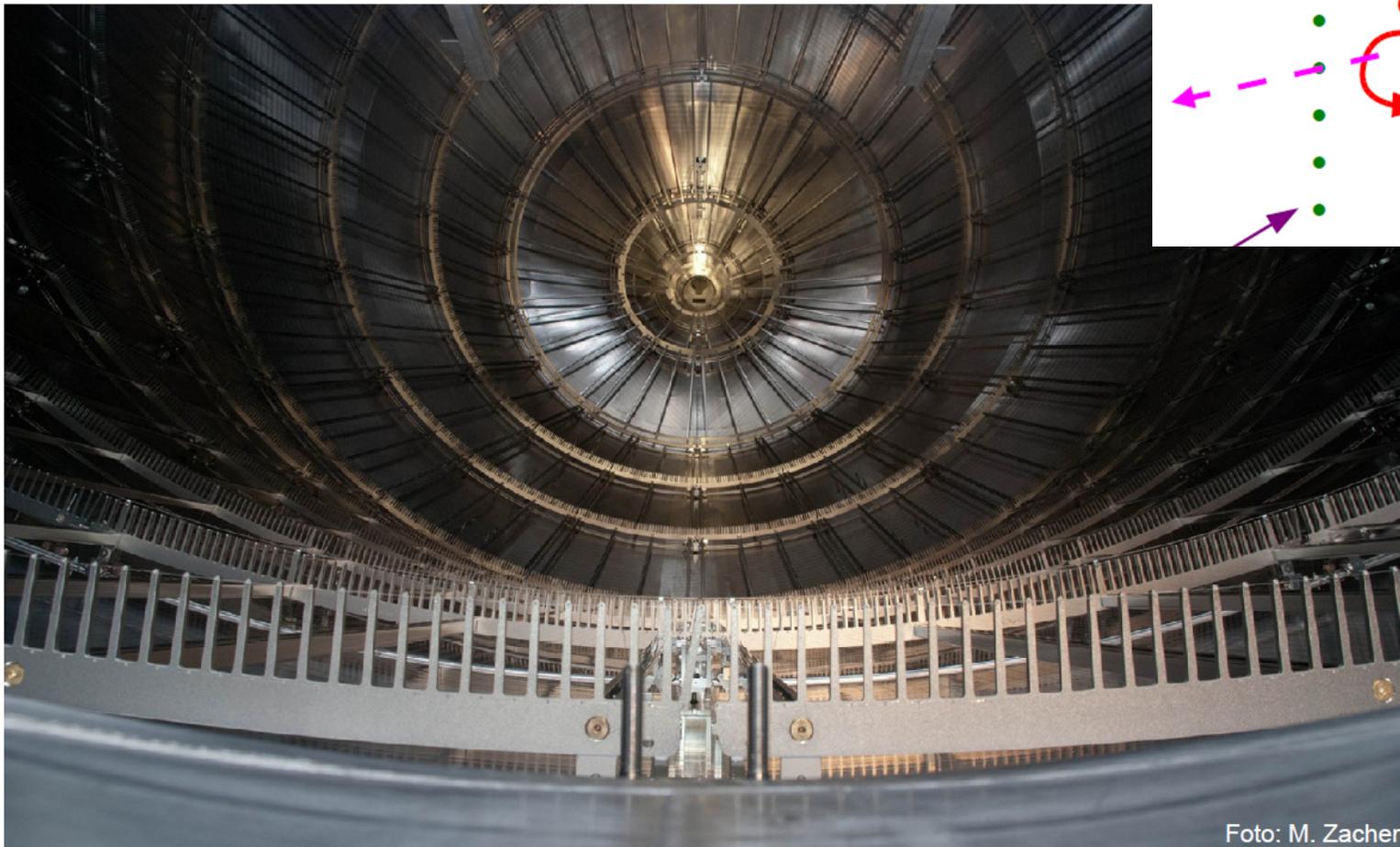


1. Confirms or excludes quasi-degenerate mass regime
2. Tests cosmological neutrino mass evaluation



KATRIN status

Inner electrodes of the main spectrometer are installed (2012)



<http://fuzzy.fzk.de/bscw/bscw.cgi/d736480/95-TRP-5210-S3-CWeinheimer.pdf>

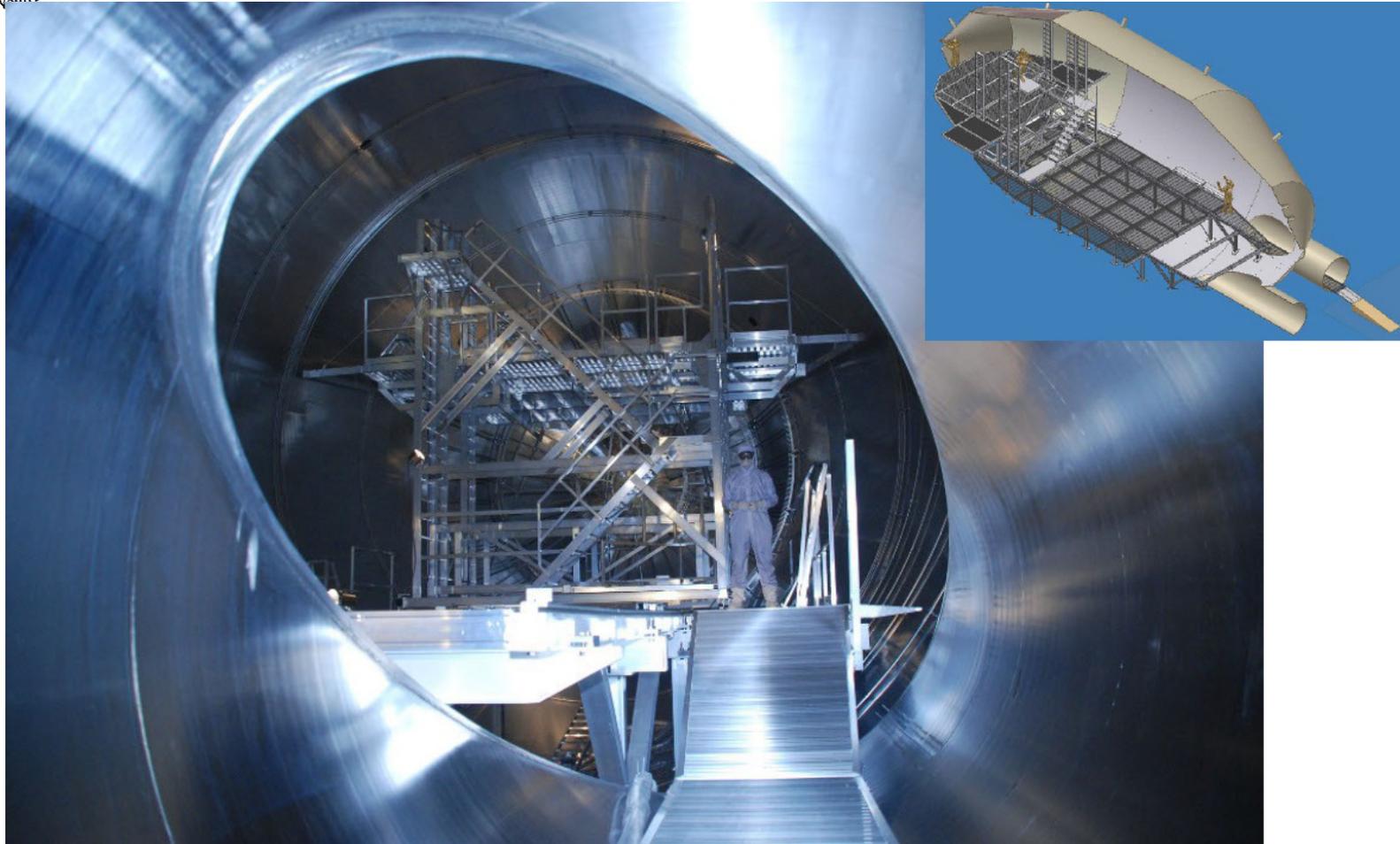
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June 26 –28 of 2013, IHEP, Protvino, Russia



KATRIN status

Inner electrodes installation



Christian Weinheimer

KATRIN CM XXII, March 2012

21

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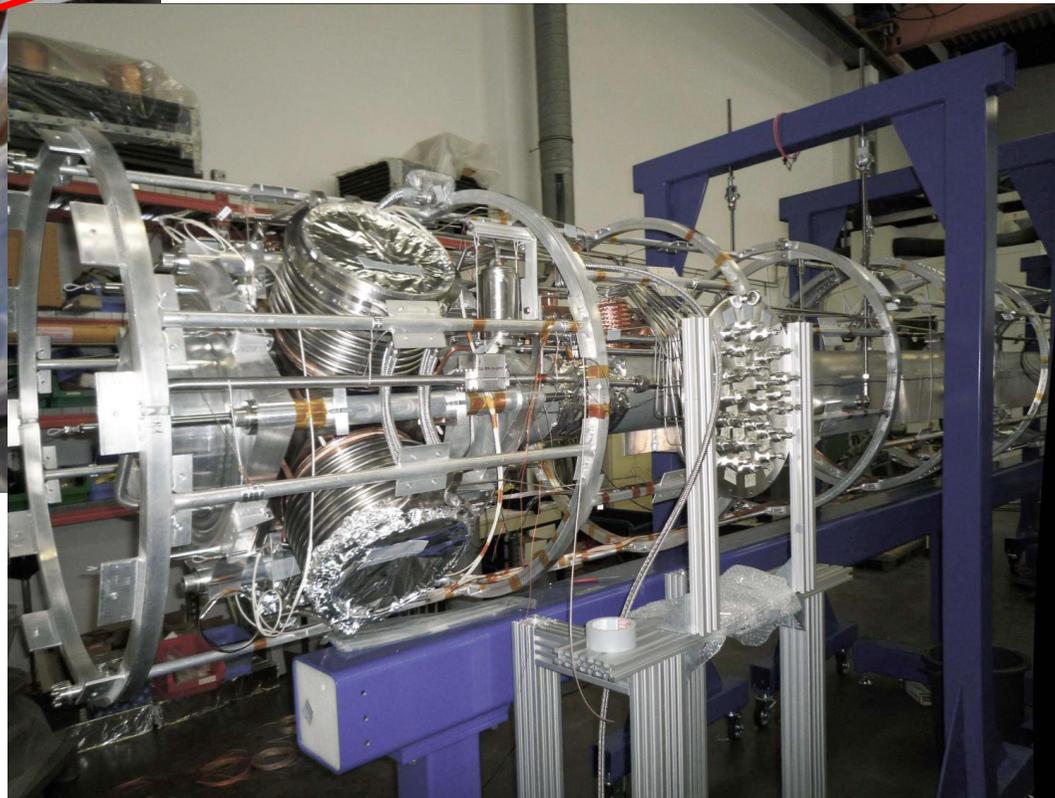
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WGTS demonstrator (2011)



Thermal stabilization by
two phase (gaseous-liquid)
Ne loop up to 30 mK



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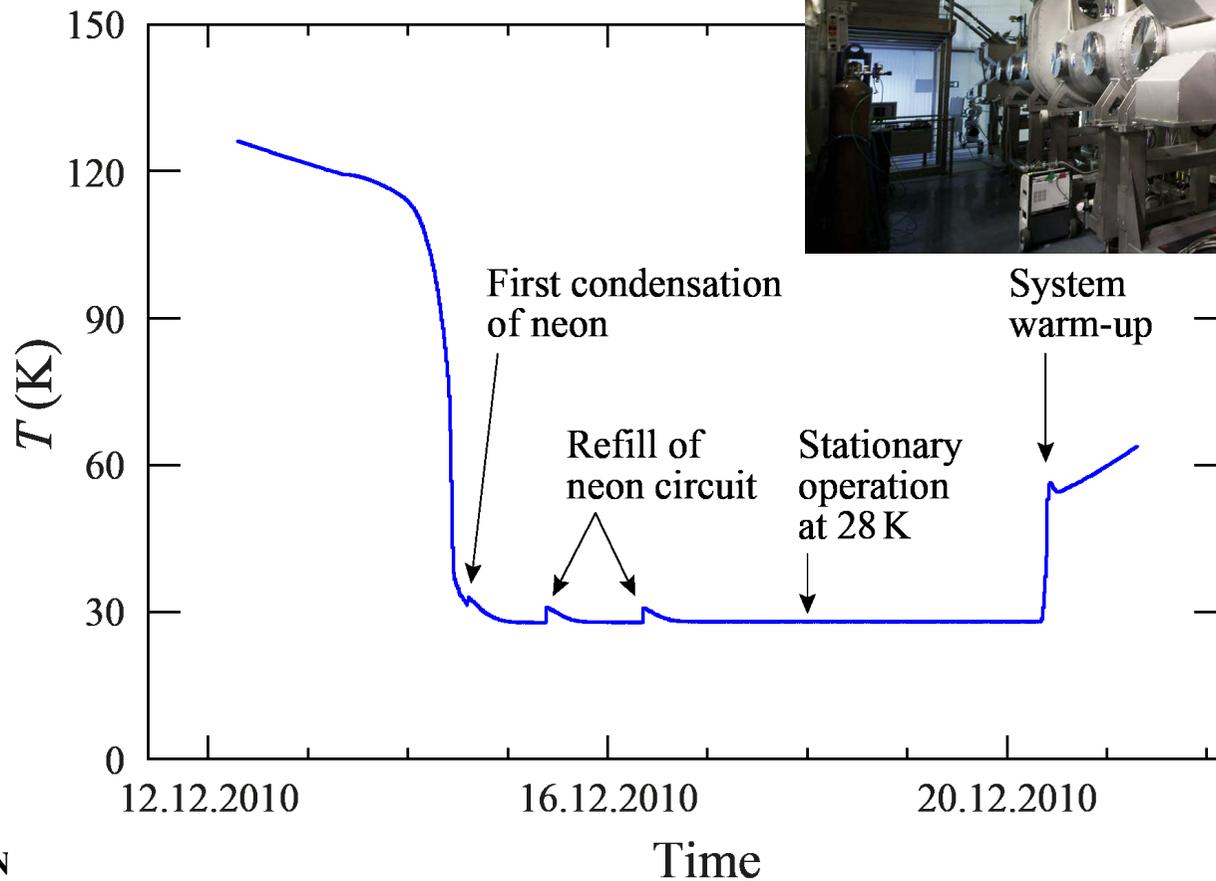
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WGTS demonstrator (2011)



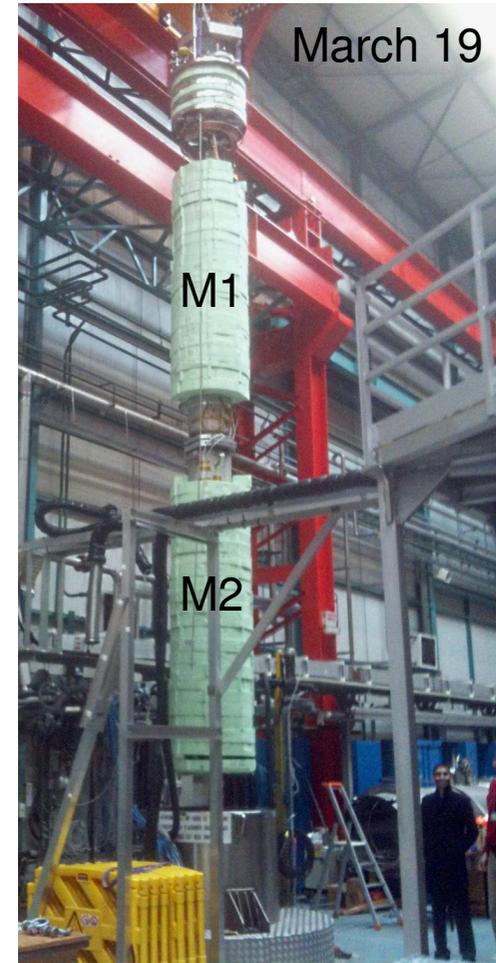
Cool-down of the WGTS-tube



N



WGTS magnets – cold tests March, 2012



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Rear Wall subproject

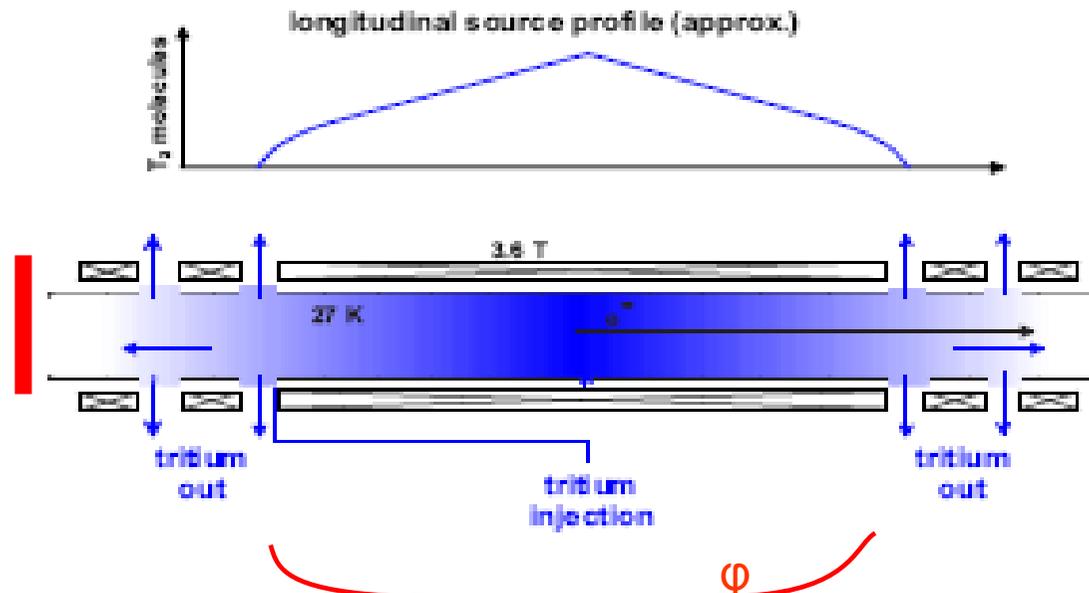
Space charge potential control



1. A space charge potential variation inside the WGTS brings false m_ν^2
2. Plasma is generated in the WGTS due to ionization by β - electrons
3. Ambipolar diffusion in central part (Boltzmann distribution):

$$\frac{n}{n_0} = \exp\left(-\frac{e\phi}{kT_{el}}\right)$$

4. High plasma conductivity along magnetic field lines equals plasma potential to the Rear Wall one.





Rear Wall subproject Space charge potential control



INR program:

1. Potential shift due to the hydrogen adsorption on the side walls.
2. Tritium plasma simulation with ionized hydrogen
 - “Rear Wall concept” check
 - Plasma transport
 - Ion-electron recombination in the WGTS
 - Ion removal by ExB drift

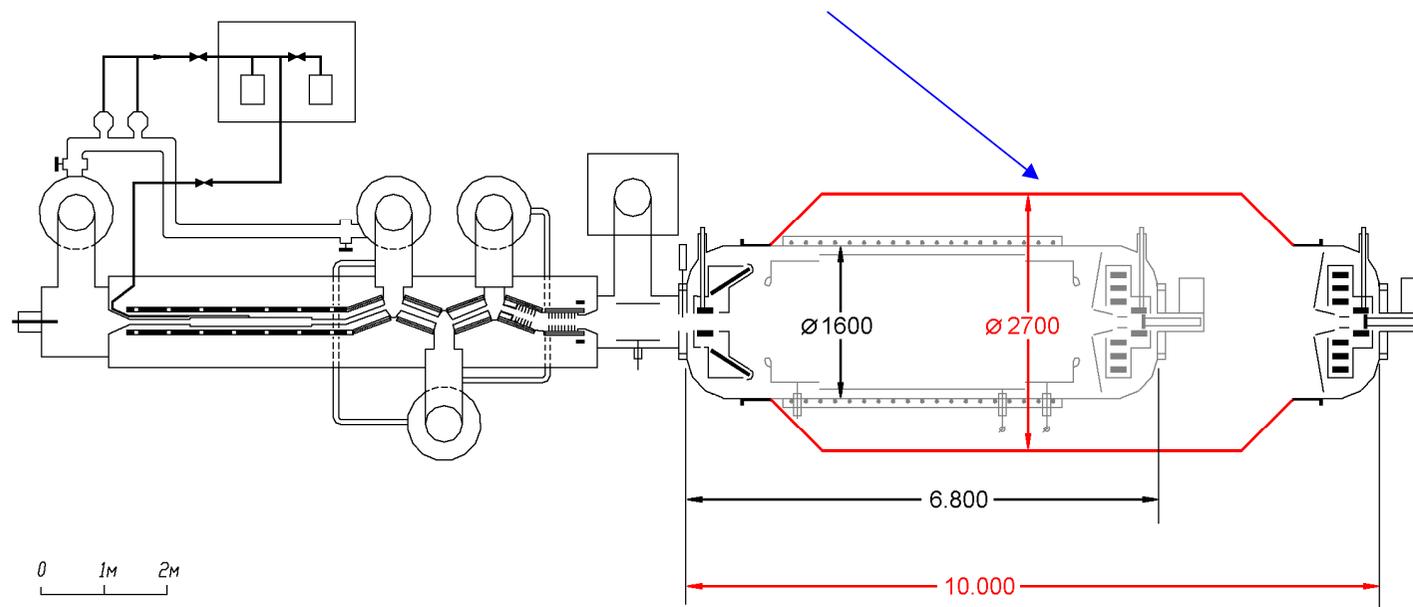
Experimental method – observation of the ^{83m}Kr internal conversion line position shift and line width broadening (when ^{83m}Kr and T_2 are circulating together through WGTS)



2005: Troitsk v-macc spectrometer upgrade



By the increase of spectrometer vessel and central electrode diameters resolution was improved by factor 2!





2009: New INR spectrometer on place!



Spectrometer outer diameter 2.75 m, length 8.10 m.

Central electrode diameter 2.20 m.

Spectrometer vessel bake-up at 200C° and chilling up to 10C°
for vacuum better than 10^{-7} Pa.

Stable operation up to 32 kV.



What's next ?

Heavy sterile neutrino search!

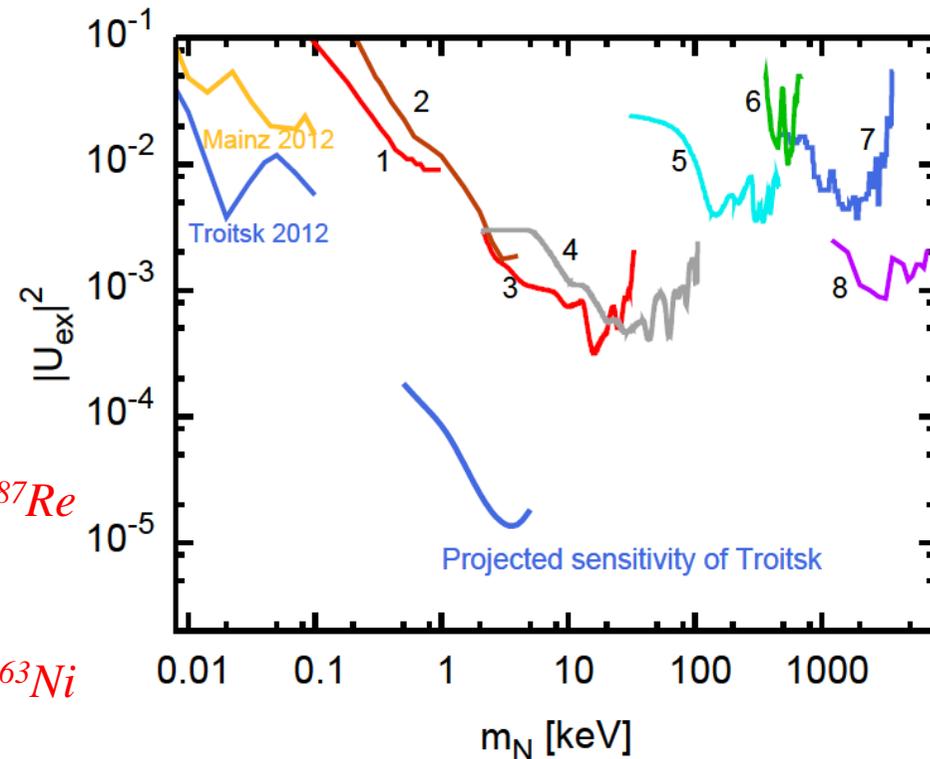


J. Beringer et al. (Particle Data Group)

PR D86, 010001 (2012) (URL: <http://pdg.lbl.gov>)

Heavy Neutral Leptons, Searches for

- [1] M. Galeazzi et al., PRL 86 (2001) 1978 – ^{187}Re
- [2] **K.H. Hidde**mann, **H. Daniel**,
O. Schwentker, JPG 21 (1995) 639 – T
- [3] E. Holzschuh et al., PL B451 (1999) 247 – ^{63}Ni
- [4] E. Holzschuh et al., PL B482 (2000) 1 – ^{35}S
- [5] K. Schreckenbach et al., PL 129B (1983) 265
- [6] M.M. Hindi et al., PR C58 (1998) 2512
- [7] M. Trinczek et al., PRL 90 (2003) 012501
- [8] J. Deutsch, M. Lebrun, R. Prieel NP A518 (1990) 149



Tritium spectrum measurement
up to ≈ 9 keV!



Back up slides

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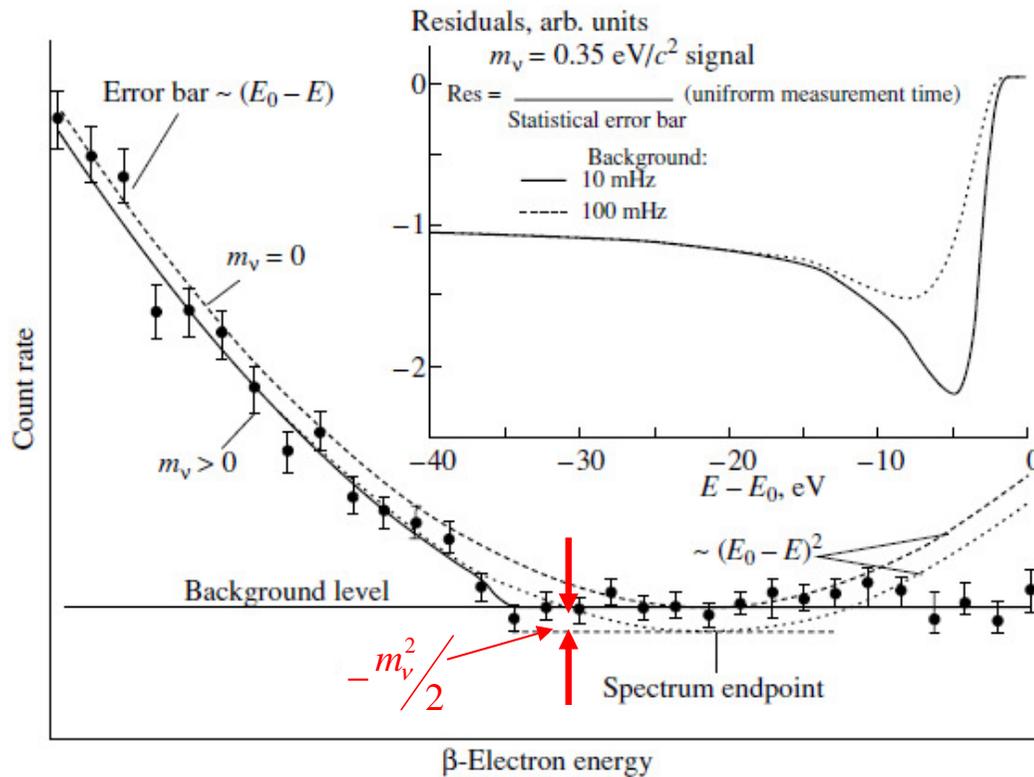
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Форма спектра при $m_\nu > 0$



$$\frac{dN}{dE} = K \times F(Z, E) \times p \times E_{\text{tot}} \times (E_0 - E) \times [(E_0 - E)^2 - m_\nu^2]^{1/2}$$



$$\frac{dN}{dE} \sim (E_0 - E)^2 - m_\nu^2 / 2$$



Electrostatic spectrometer with adiabatic magnetic collimation (MAC-E filter) Fundamentals



Adiabaticity parameter ε :

$$\varepsilon = \frac{|gradB|}{B} r_H \ll 1 \quad \text{or} \quad \varepsilon = \frac{1}{\omega_H} \cdot \left| \frac{\dot{B}}{B} \right| \ll 1$$

where r_H , ω_H – radius and frequency of the Larmor precession

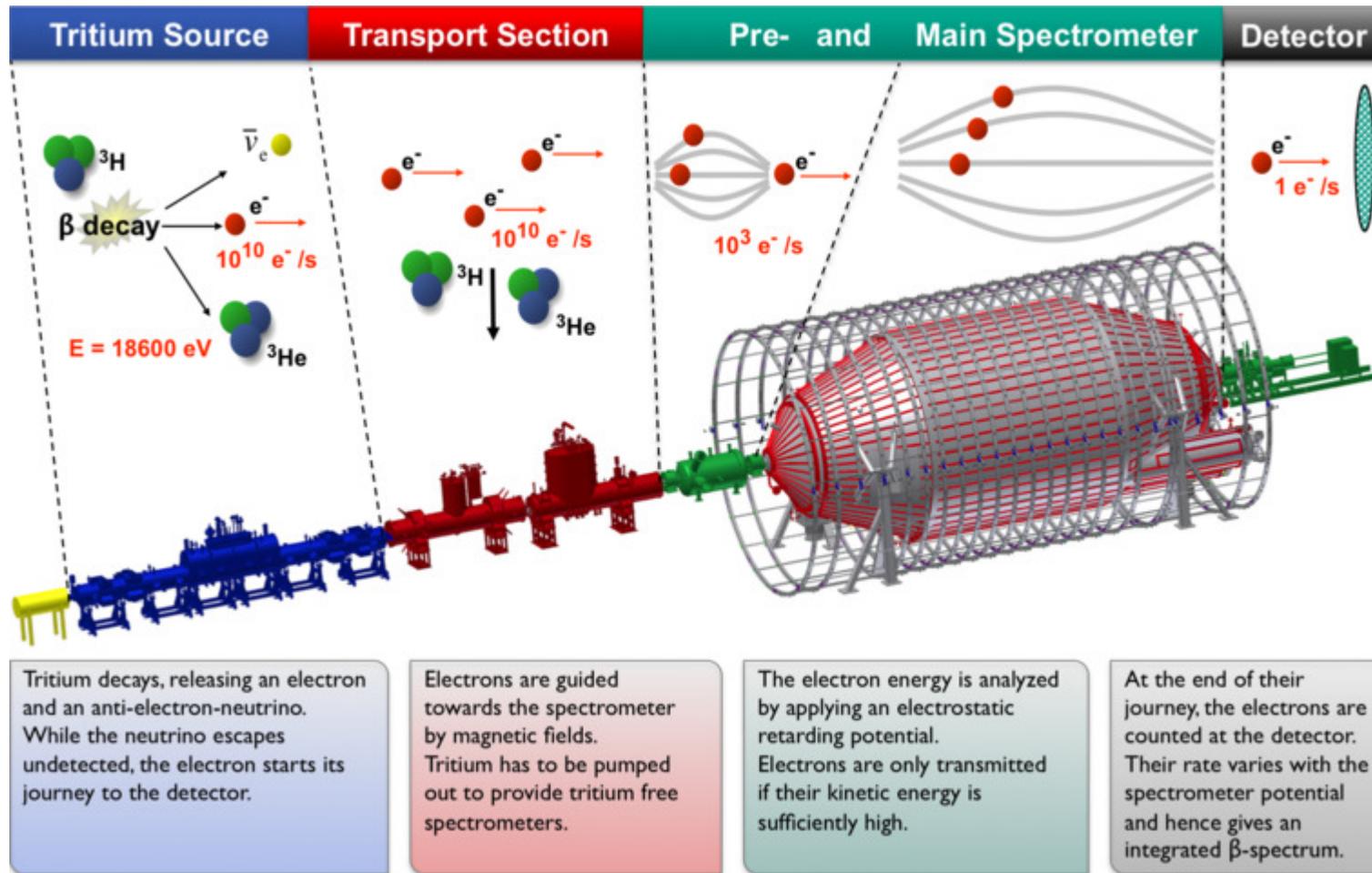
An adiabatic invariant is preserved exponentially:

$$\frac{\Delta\mu}{\mu} \sim e^{-\frac{1}{\varepsilon}}$$

The spectrometer resolution is independent from the trajectory radius and curvature if adiabaticity parameter is small!

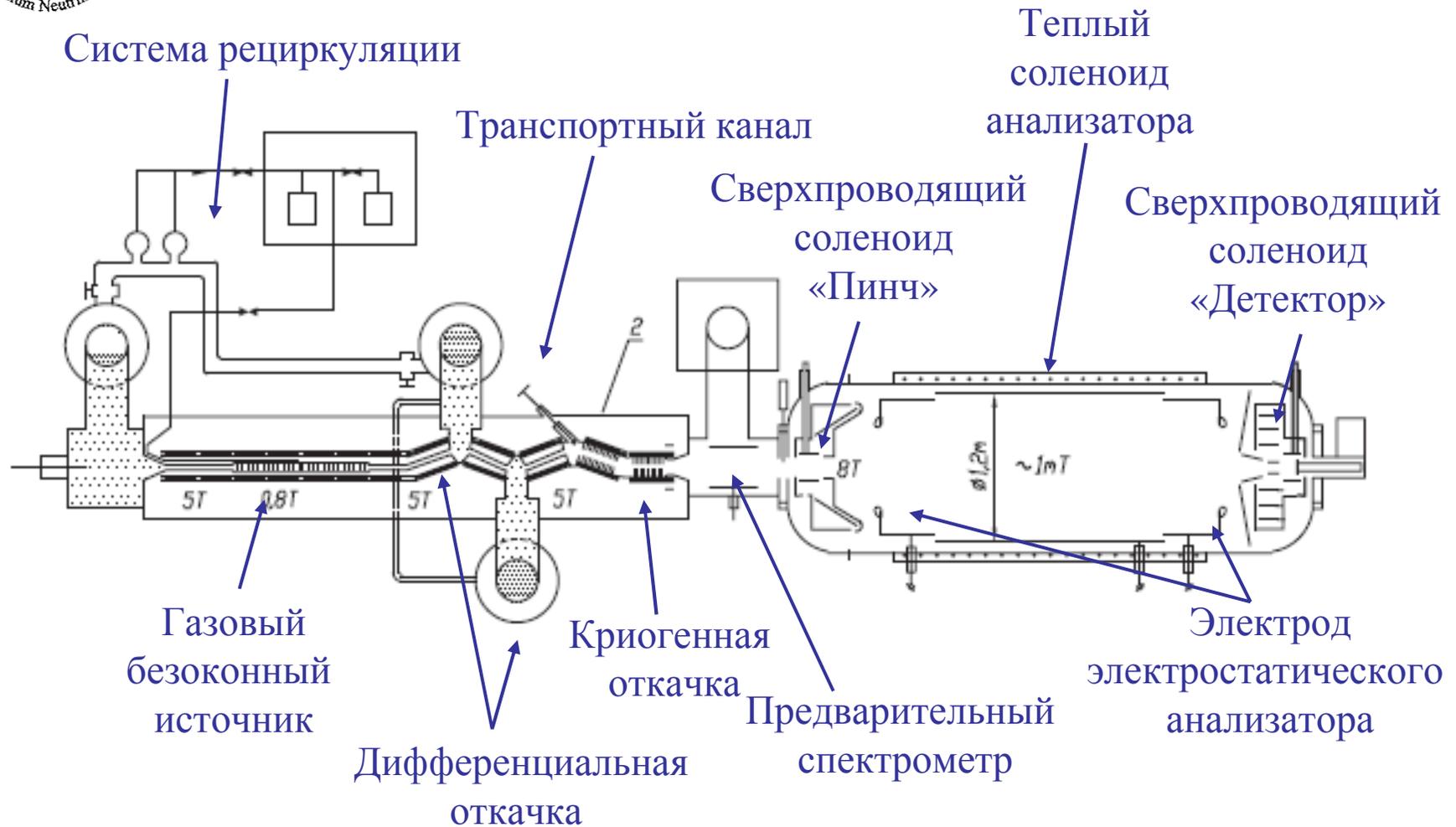


Продолжение поиска массы нейтрино: международный проект КАТРИН в Карлсруэ



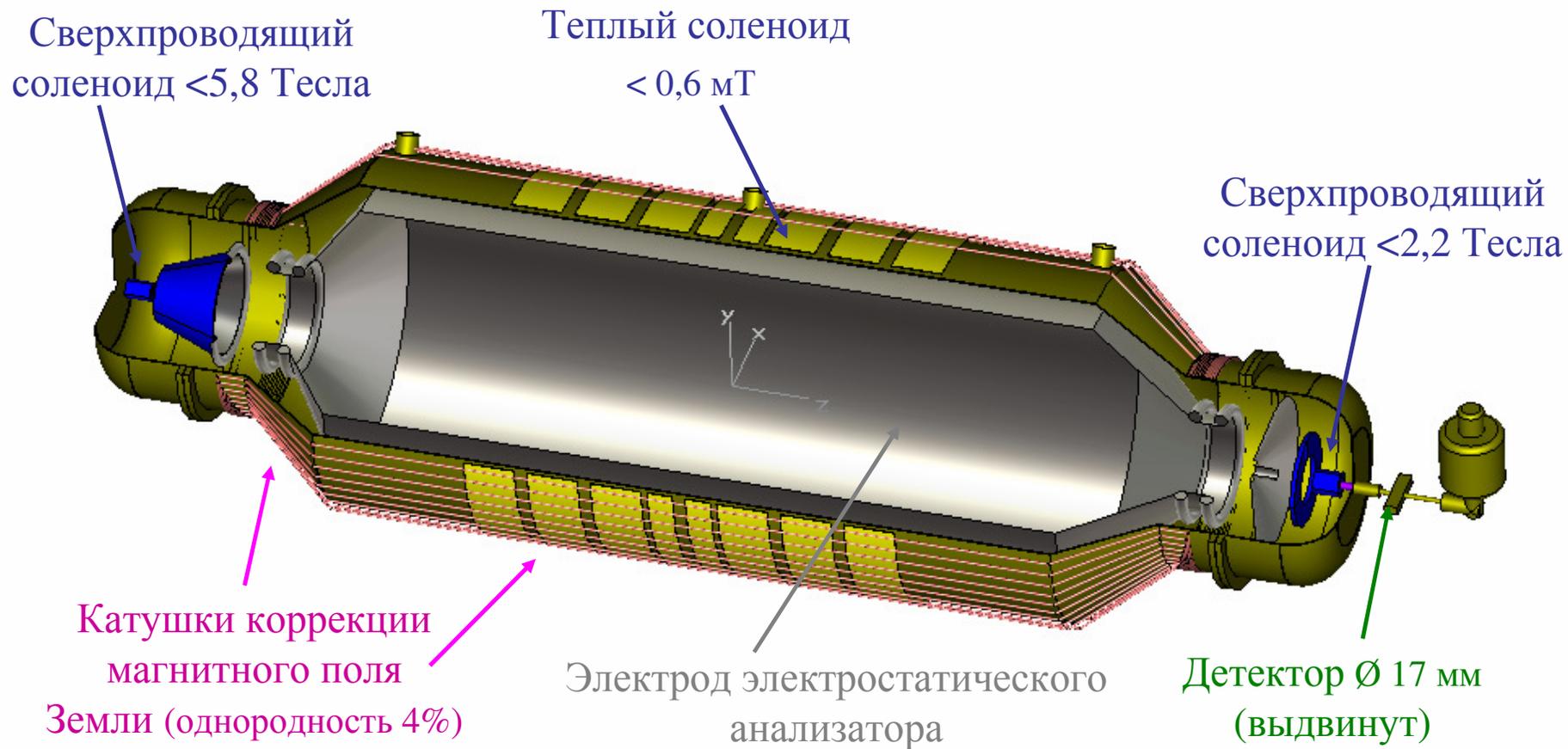


Установка «Троицк V-масс»



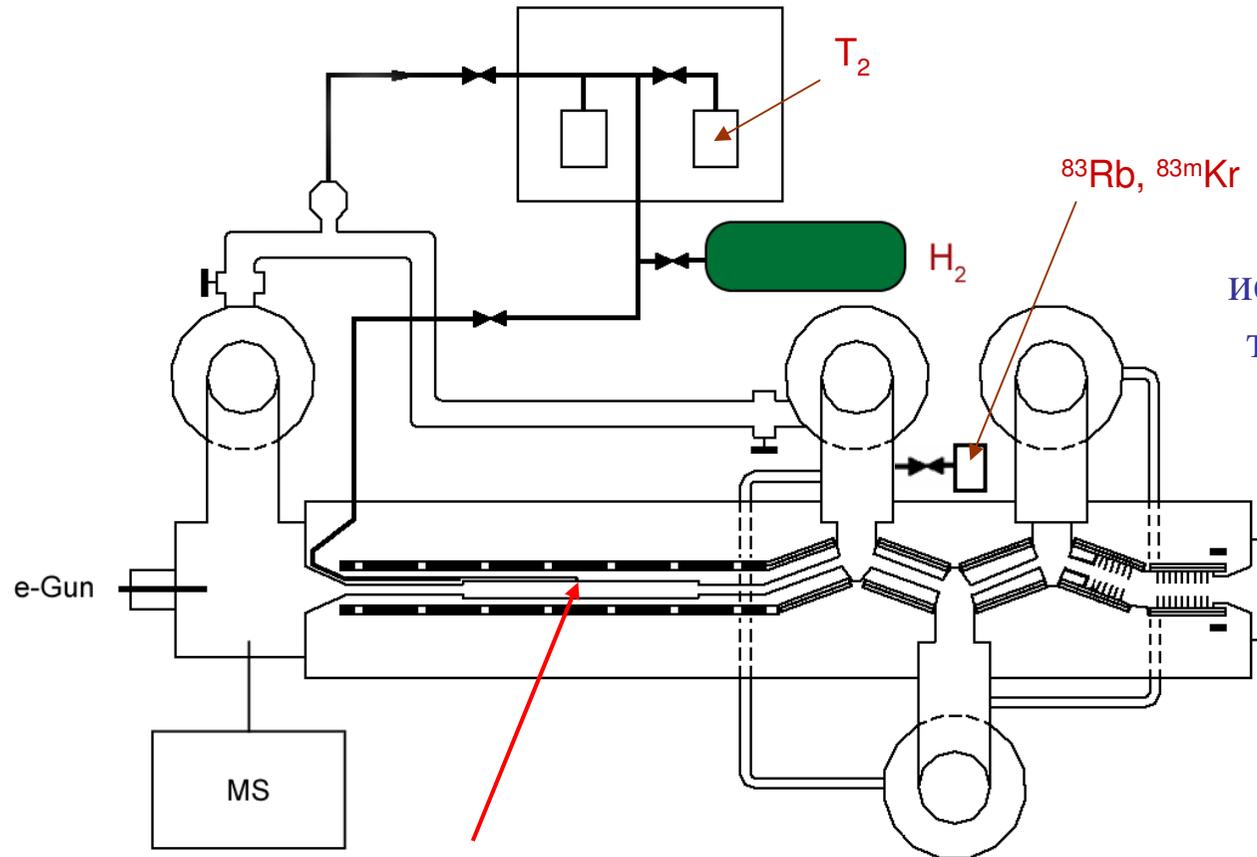


Новый спектрометр: модель





Источник Экспериментальные возможности



Светимость газового
источника $5,4 \text{ стер} \cdot \text{см}^2$,
толщина $1 \cdot 10^{17} \text{ мол/см}^2$

Температура газа
в источнике
 $20\text{K} \dots 110\text{K}$

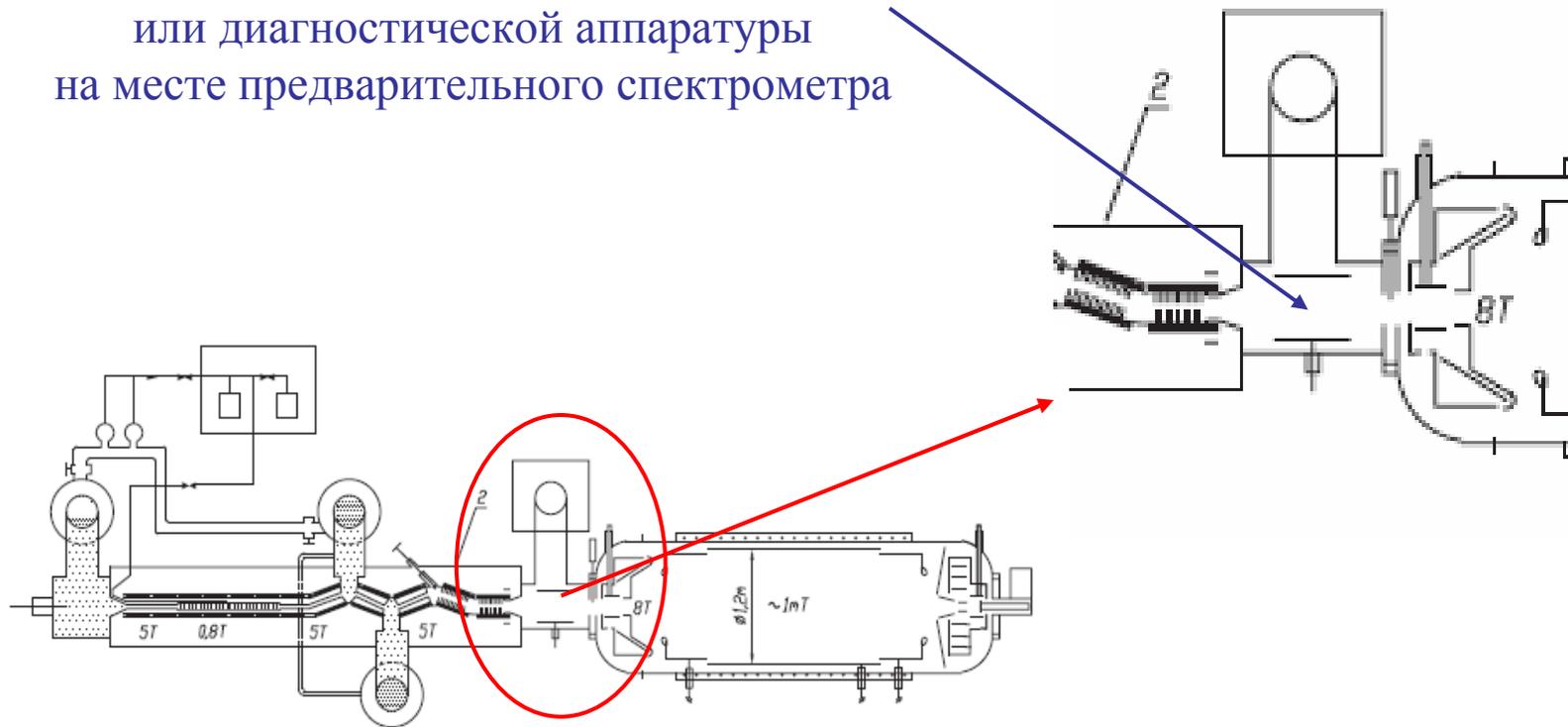
Трубка источника подогревается



Источник Дополнительные возможности



Оборудование станции
для твердых источников
или диагностической аппаратуры
на месте предварительного спектрометра





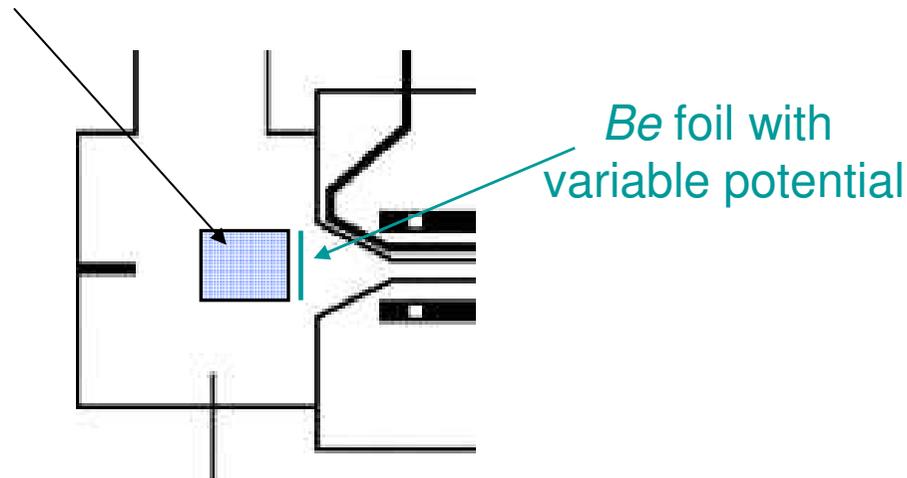
Исследование систематики газового источника



1. Влияние объёмного заряда

Потенциал объёмного заряда исследуется в измерениях спектров электронов внутренней конверсии ^{83m}Kr при совместной циркуляции с тритием или водородом, ионизованным внешним источником электронов:

Электронная пушка с широкой апертурой помещается в вакуумный пост насоса в задней части источника.





Исследование систематики газового источника

2. Рекомбинация ионов в газовом источнике и удаление ионов в скрещенных $E \times B$ полях

