

Simulation and analysis of the LUCID experiment in the Low Earth Orbit radiation environment

T. Whyntie^{*,†}, M. Harrison^{*}

** Langton Star Centre, † Queen Mary, University of London*

CHEP 2013, Amsterdam

Tuesday 15th October 2013

A brief history of LUCID

In 2008, the Simon Langton Grammar School for Boys entered a satellite experiment design competition run by the British National Space Centre (now UK Space Agency) and Surrey Satellite Technology Limited (SSTL).

- The Langton Ultimate Cosmic ray Intensity Detector (LUCID) would use Timepix detectors, developed by the Medipix Collaboration, to measure the space radiation environment in Low Earth Orbit.
- Designed by students, built by SSTL, now due to launch in February 2014.
- LUCID now part of CERN@school, supported by UK Science and Technology Facilities Council (STFC) Large Award ST/J000256/1.



Overview

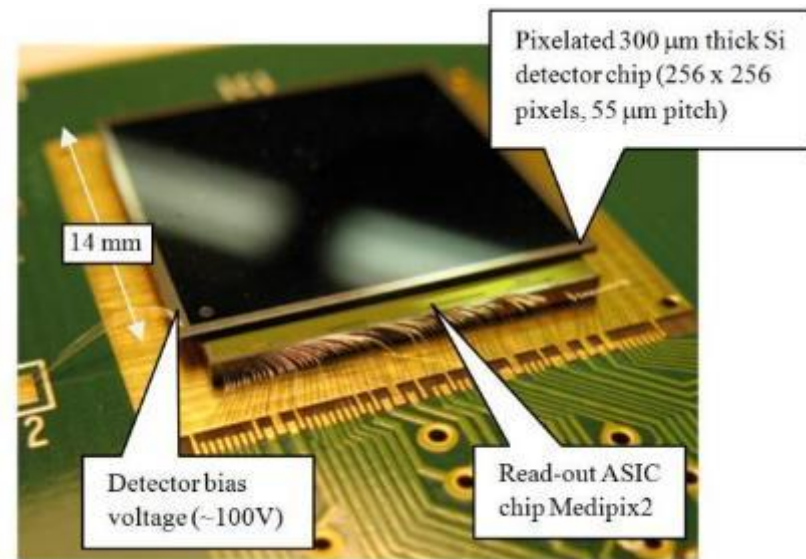
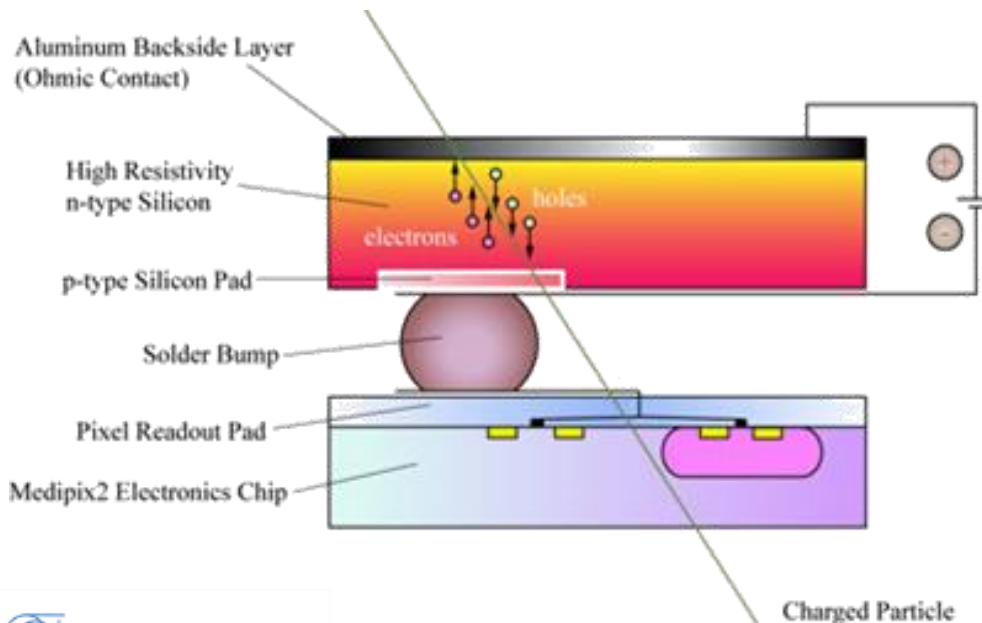
- Introduction:
 - *A brief history of LUCID, the Timepix detector, the LUCID experiment, TechDemoSat-1.*
- Simulating LUCID:
 - *GEANT4, experiment geometry, particle sources from SPENVIS, simulation management.*
- Simulation results:
 - *Trapped protons, trapped electrons, estimated data rates, discussion, further work.*
- Summary and conclusions.

The Timepix detector



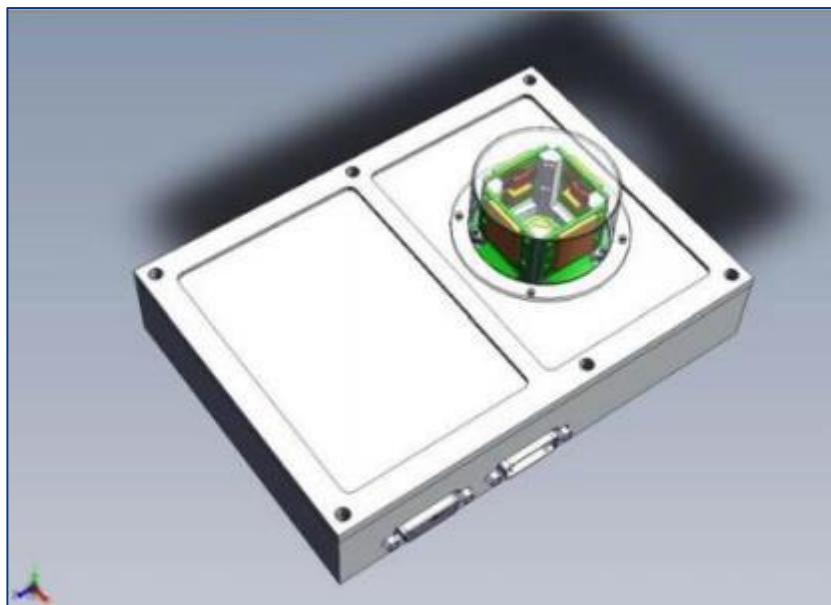
The Timepix hybrid silicon pixel detector ([Llopart et al. 2007](#)), developed by the [Medipix Collaboration](#), features a 300 μm thick silicon sensor bump-bonded to a Timepix readout chip. 256 \times 256 pixels of pitch 55 μm provide 65,536 readout channels from the 1.98 cm^2 sensor element.

It can be used to detect ionising radiation, make energy measurements (with calibration) and perform particle identification (to an extent). It has never been used in *open space*.



The LUCID experiment

The Langton Ultimate Cosmic ray Intensity Detector experiment features five Timepix detectors in an open-faced cube, housed in a $\sim 0.7\text{m}$ aluminium “dome”, to measure the LEO radiation environment.



Data taking/transmitting capabilities:

- Shutter frequency: $\sim 4\text{Hz}$
- Transmission: 80Mbs^{-1} (20Mbs^{-1});
- Storage: 2GB;
- Operational 2 out of every 8 days.

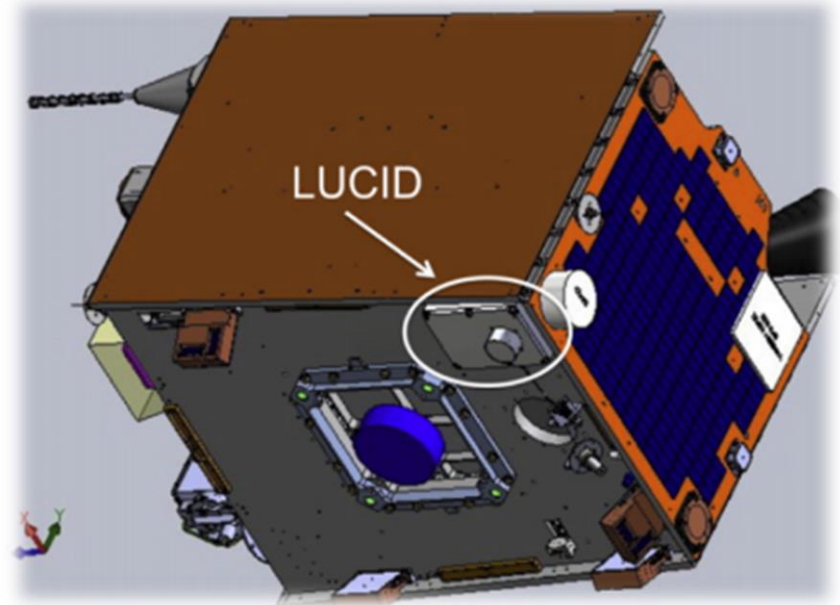
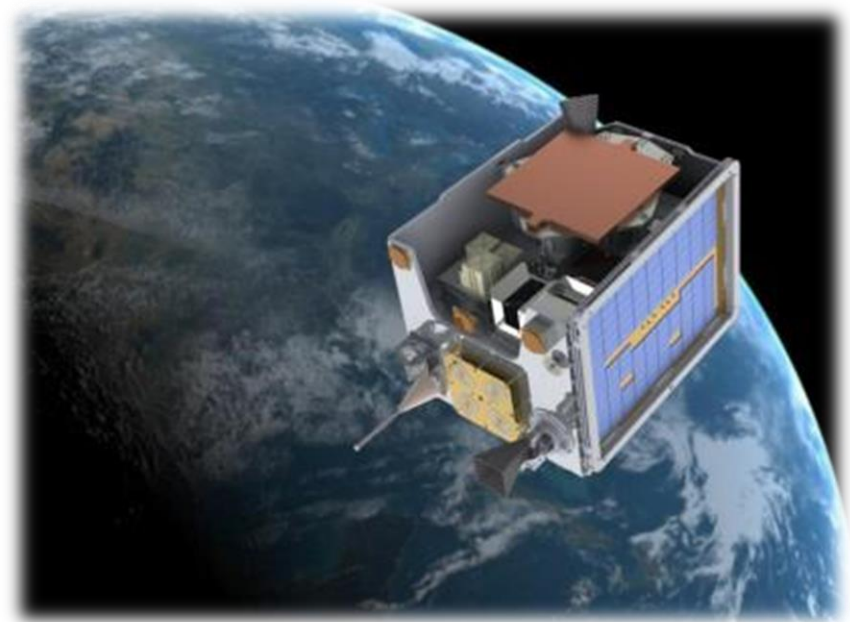
TechDemoSat-1

LUCID will launch aboard **TechDemoSat-1** from Baikonur Cosmodrome aboard a Soyuz 2b launch vehicle in February 2014.

- TechDemoSat-1 is an “in-orbit test facility” from SSTL supported by the UK Technology Strategy Board;
- Many scientific payloads from UK academic institutions/labs, of which LUCID is one.

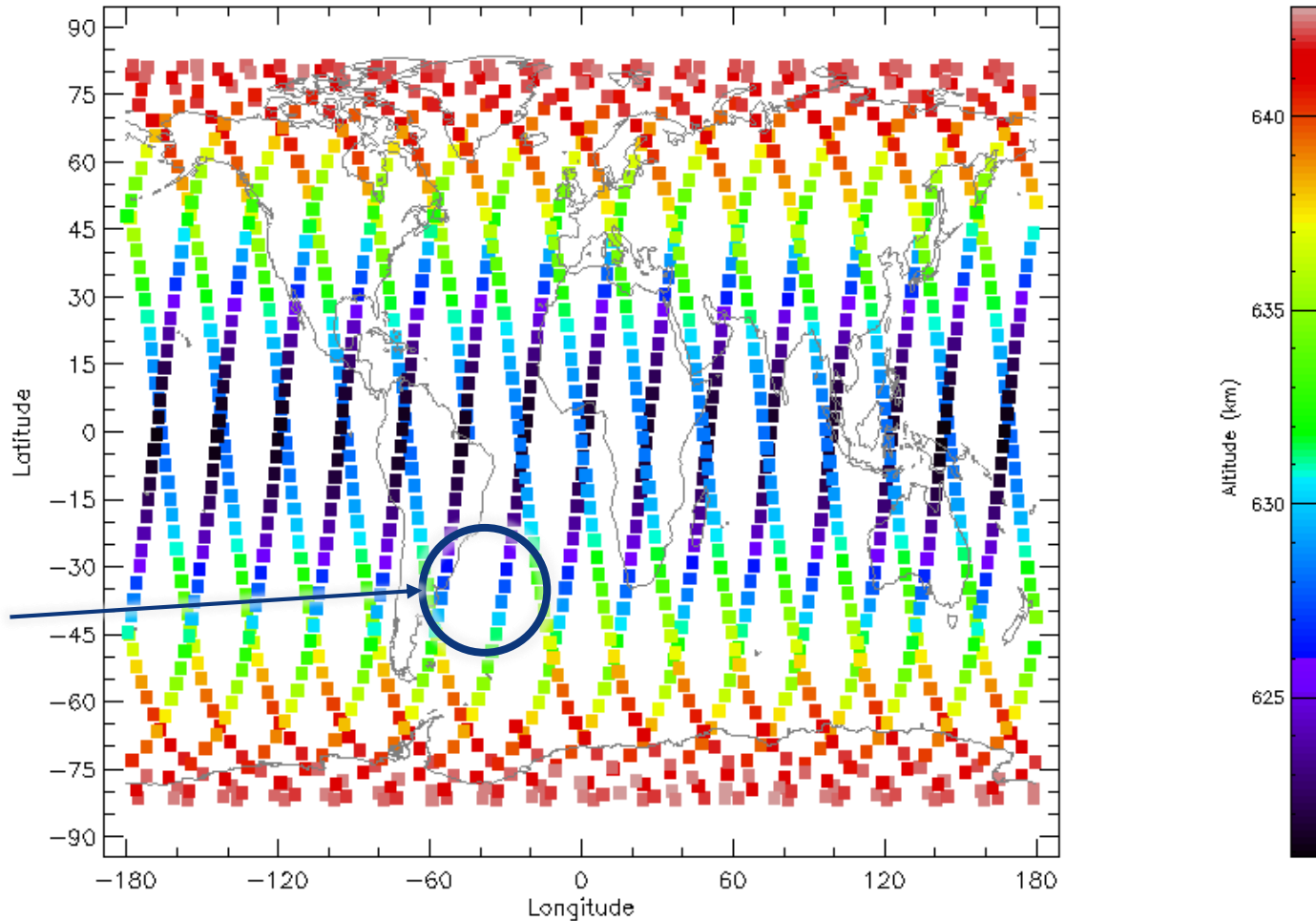
Orbit parameters

- *Altitude*: 635km;
- *Orbit*: sun synchronous;
- *Inclination*: 98.4°;
- *LTDN*: starts at 0900, with drift of 40 minutes per 6 months.
- *Dominant radiation sources*: trapped protons and electrons, outer electron belts (poles).



TechDemoSat-1

South
Atlantic
Anomaly
(SAA)



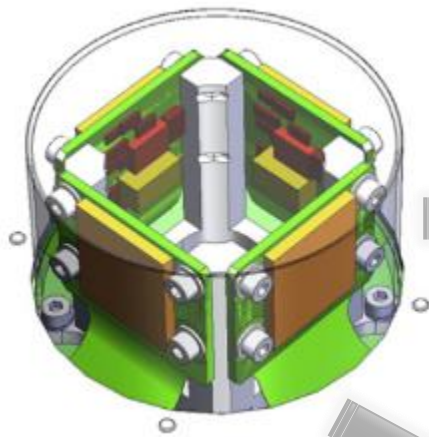
Aims and objectives

- Scientific goals of LUCID:
 - *Directionality of particles* (assumed to be isotropic):
 - Pattern recognition and estimate angle of incidence
 - *Particle spectra*:
 - Measure particle intensities and their change over time
 - Determine dose (J/kg) and LET spectrum
 - *Solar flares/SPEs*:
 - Difference in time between electron and proton surges
 - Log sun activity for mission life time (> 50% solar cycle)
 - Forbrush Decrease.
- Goals of the work presented here:
 - Provide estimates of the expected data rates that LUCID will see;
 - Establish modelling and simulation workflows for the experiment;
 - Groundwork for more complicated/computer intensive simulations.

Simulation with GEANT4

- **GEANT4** has been used to simulate the detector, its surroundings and the particle source(s) in order to obtain data rate estimates.
 - Particular inspiration from J. Idarraga (*CERN/Medipix*):
 - Allpix – full chain simulation of Medipix family (and others);
 - Ultimately, LUCID will be modelled with Allpix, but such detail is not needed for these estimates.
 - Method:
 - Model the experiment **geometry**;
 - Model the Low Earth Orbit **particle source(s)**;
 - Obtain pixel hit rates, combine with flux values to get [bs^{-1}] estimates.
 - The default GEANT4 physics settings have been used.

Geometry



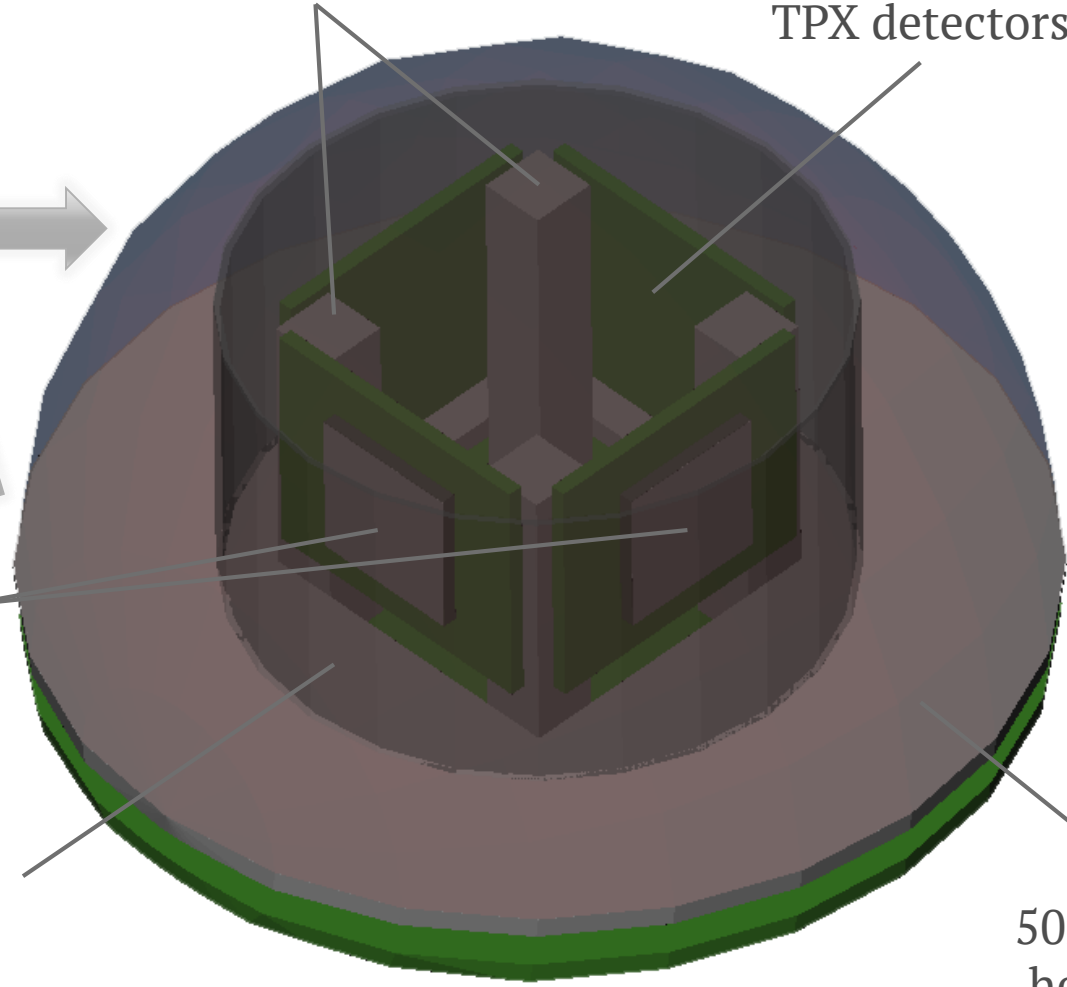
Support posts

PCB mounts for
TPX detectors

Timepix detectors
(including TPX0,
through gap)

Aluminium "dome"
(cylinder with roof),
thickness ~0.7mm

50mm radius
hemisphere



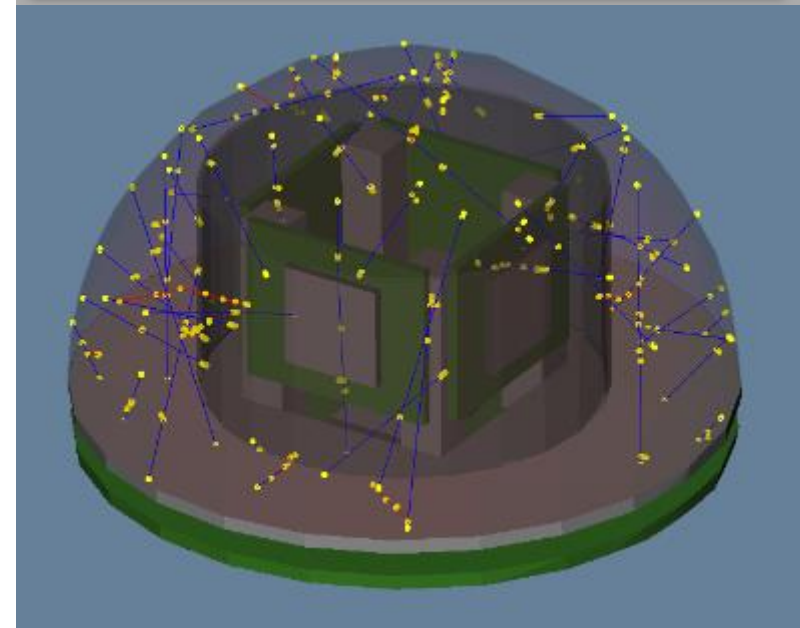
Particle source(s)

SPENVIS

- ESA-backed “Space Environment Information System” web portal.
- *Spacecraft coordinate generators:*
 - Input LUCID orbit details.
- *Trapped radiation models:*
 - AP-8 for protons and electrons;
 - Int. and diff. flux spectra.

GEANT4 General Particle Source (GPS)

- Hemi-spherical surface, energy sampled from flux spectra energy bins;
- *Right:* 50 10-20 MeV protons (“dome” made partially transparent for clarity).



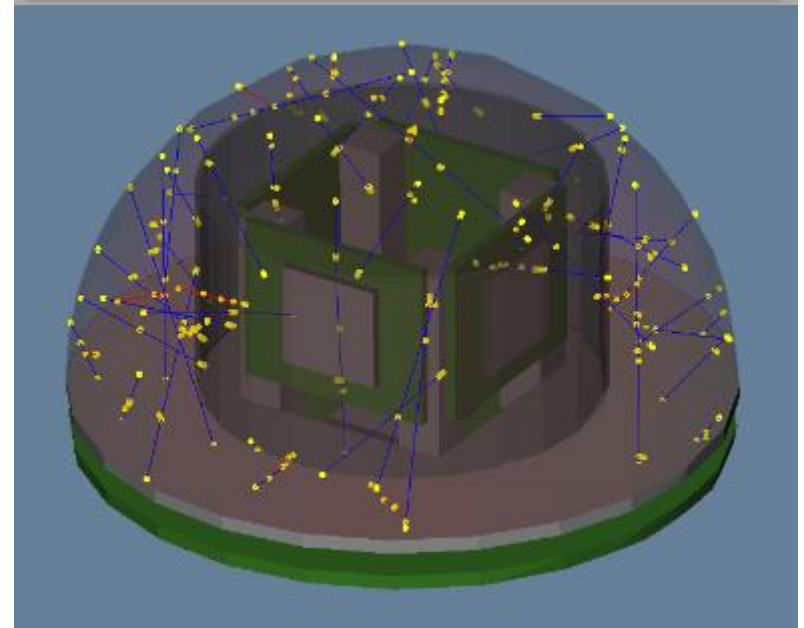
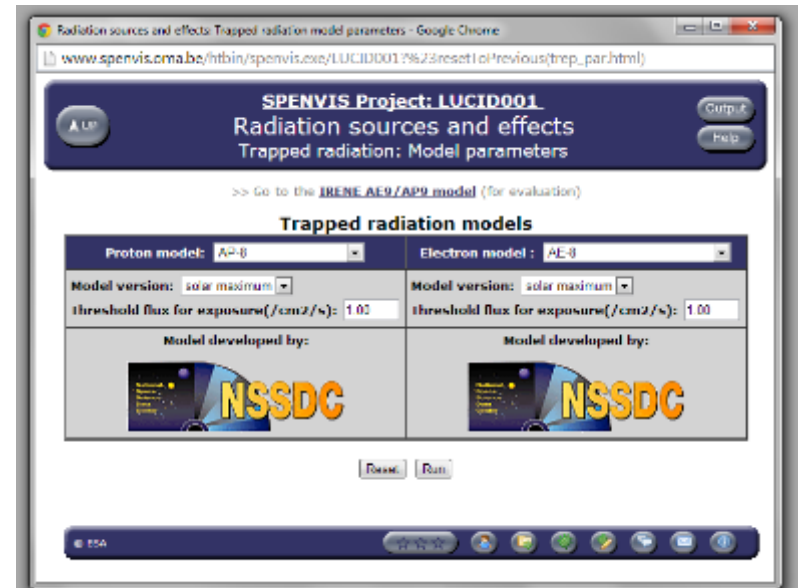
Particle source(s)

```
# Source particle type
/gps/particle proton

# Source particle energy
/gps/ene/type Lin
/gps/ene/min 10.0 MeV
/gps/ene/max 15.0 MeV
/gps/ene/gradient 0.0
/gps/ene/intercept 1.0

# Source particle geometry
/gps/pos/type Surface
/gps/pos/shape Sphere
/gps/pos/centre 0. 0. 0. mm
/gps/pos/radius 49.9 mm
/gps/pos/confine PseudoDetector_phys
/gps/ang/type cos

# Run the simulation.
/run/beamOn 50
```



LUCIDITY

The LUCID Interactive Test sYstem:

- A web portal for managing LUCID simulations, source cards, results.
- **Ruby on Rails** with a **MySQL** backend. Uses **Hobo** – a collection of gems and plug-ins for making fully-functioned web apps *quickly*:
 - <http://www.hobocentral.net/>
- Hosted on an **AWS VM** (Ubuntu 12.04 Server). Deployed with **Phusion Passenger**/Apache. Email with **SES**, storage with **SE2**.
- Import SPENVIS spectra, process, download GEANT4 source cards, upload simulation results.

The screenshot shows the LUCID web portal interface. At the top, there are navigation tabs for 'Home', 'Users', and 'Flux Spectra Files'. The main content area is titled 'Flux Spectra File Starting Orbit'. It includes a 'Process file' button and a text area containing SPENVIS file content. Below this, there is a table of 'Flux Spectra File source cards' with columns for Source Card, Particle type, E min, E max, Integral flux, Differential flux, and Source card file. The table lists several source cards for 'StartingOrbit' with particle type 'Protons' and various energy ranges and flux values.

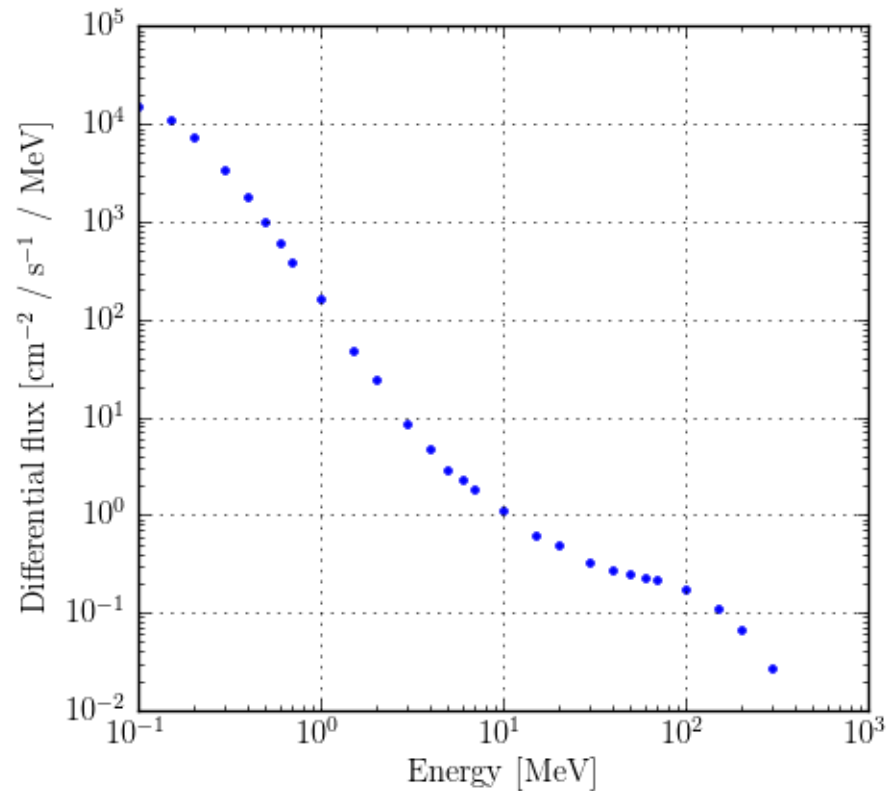
Source Card	Particle type	E min [MeV]	E max [MeV]	Integral flux [cm ⁻² s ⁻¹]	Differential flux [1/m ² s]	Source card file
StartingOrbit_p_0	Protons	0.1	0.15	1.887E+05	1.303E+04	.in Edit Remove
StartingOrbit_p_1	Protons	0.15	0.2	1.541E+05	9.121E+03	.in Edit Remove
StartingOrbit_p_2	Protons	0.2	0.3	8.979E+02	5.404E+03	.in Edit Remove
StartingOrbit_p_3	Protons	0.3	0.4	5.907E+02	2.651E+03	.in Edit Remove
StartingOrbit_p_4	Protons	0.4	0.5	3.677E+02	1.411E+03	.in Edit Remove

Calculating the data rate

Differential data rate	=	Differential flux	Source geometry	Hit pixels per source particle	Bits per pixel
[bs ⁻¹ MeV ⁻¹]		[particles cm ⁻² s ⁻¹ MeV ⁻¹]	[cm ²]	[pixels particles ⁻¹]	[b pixels ⁻¹]
		Obtained from SPENVIS/AP-8 models.	Hemisphere ($R=5.0$ cm) in an isotropic source: $\frac{1}{2}\pi R^2$	Obtained from GEANT4 simulations of simplified LUCID geometry. All five Timepix Sensitive Detector volumes recreated, hits in 55 μ m pixels count only once per event.	16 bits per pixel: 14 for Time-over-Threshold value, 2 for overhead.

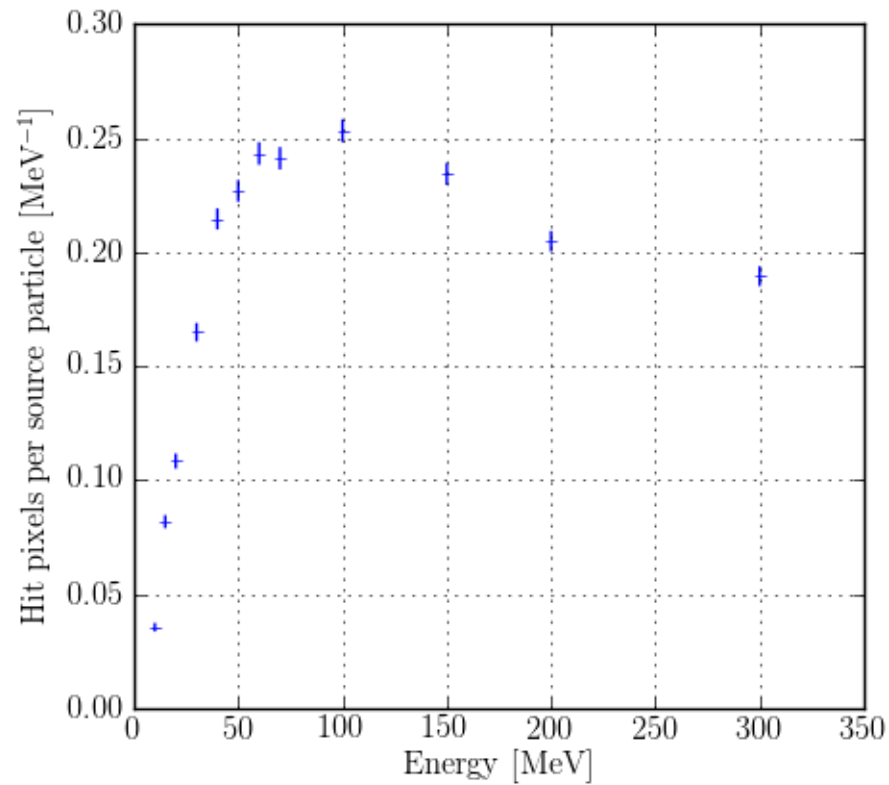


Results - protons



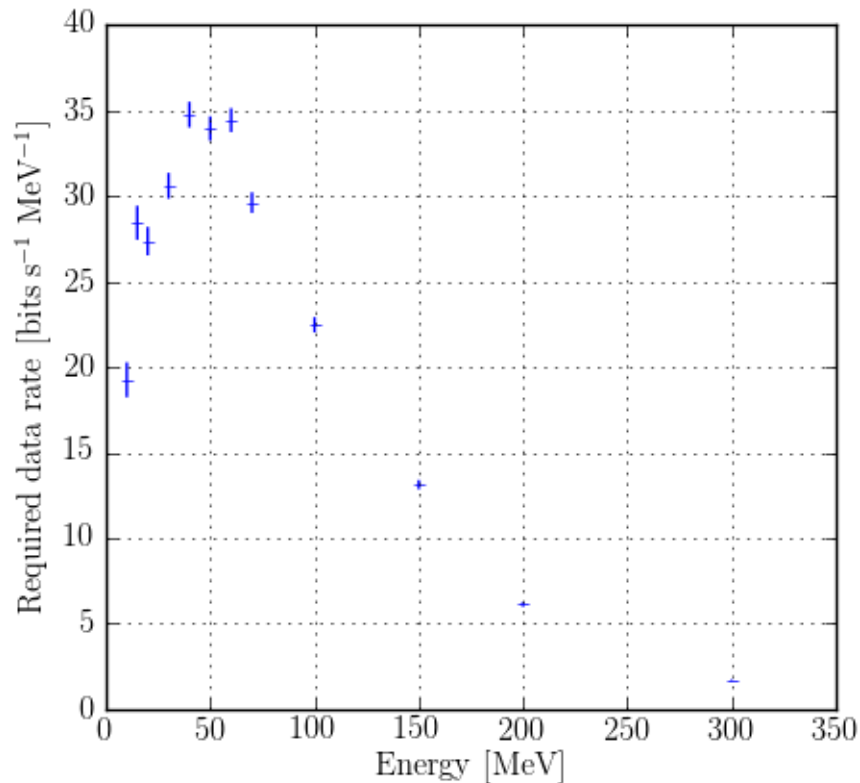
Differential flux (SPENVIS/AP-8 trapped proton model).

Results - protons



Hit pixels per source particle (SimLucid GEANT4 model).

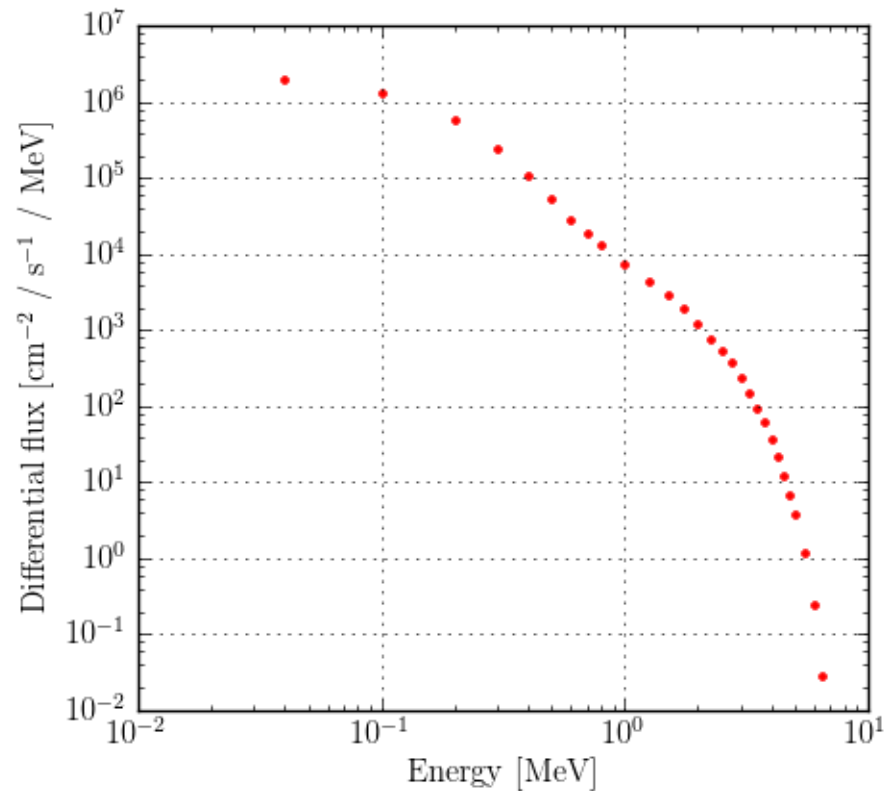
Results - protons



Required data rate summed over all energy bins due to trapped protons in LEO:
= $(5.30 \pm 0.04) \text{ kbs}^{-1}$

Required data rate due to LEO protons.

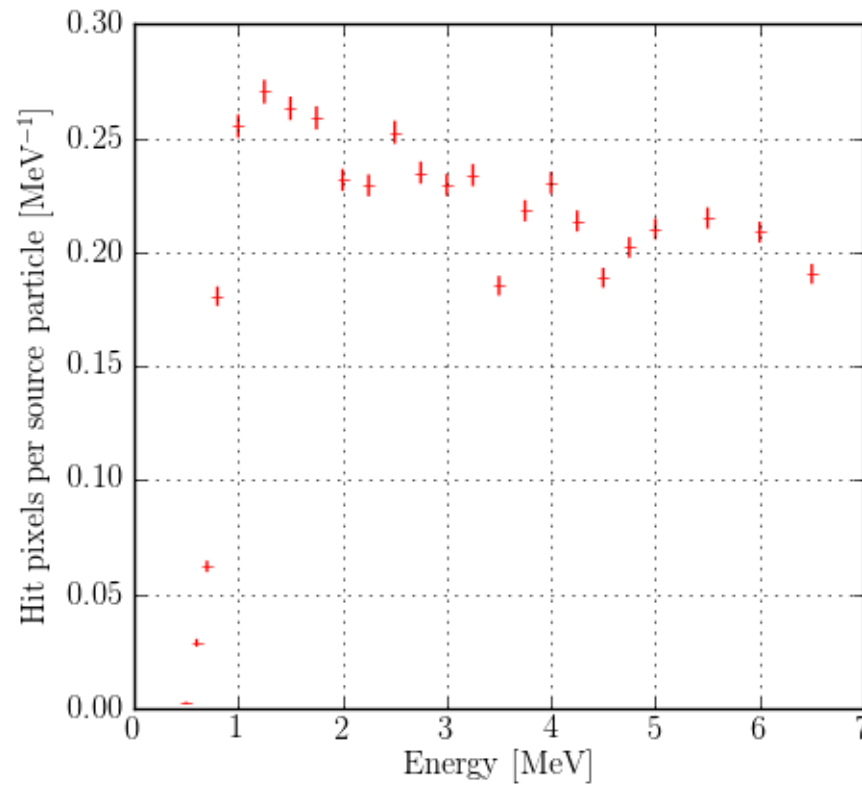
Results - electrons



Differential flux (SPENVIS/AP-8 trapped electron model).

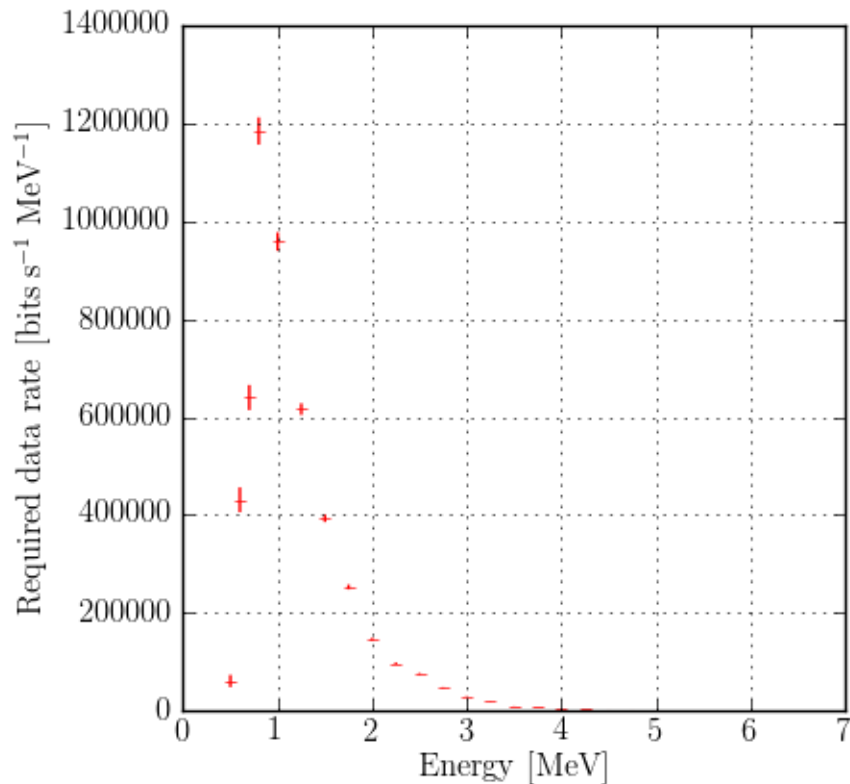


Results - electrons



Hit pixels per source particle (SimLucid GEANT4 model).

Results - electrons



Required data rate summed over all energy bins due to trapped electrons in LEO:
 $= (1.01 \pm 0.01) \text{ Mbs}^{-1}$

Required data rate due to LEO electrons.

Results - discussion

- Observed particle rates:
 - As expected, low energy particles shielded by aluminium dome while keeping energies of interest;
 - About 1 Mbs⁻¹ from trapped radiation in LEO (solar maximum) for all detectors, dominant contribution comes from electrons;
 - Errors: statistical plotted, “factor of 2” from AP-8 models.
- Comparing with LUCID technical capabilities:
 - 1 Mbs⁻¹ suggests 0.2 frame s⁻¹ per detector, so at 4Hz shutter speed the frame occupancy is ~5% (256x256x16 bits per frame);
 - 20 Mbs⁻¹ data transfer, 2GB storage space – sufficient.
- Further work to be done:
 - Allpix full chain simulations (on Grid), other radiation environment (e.g. outer belt electrons, SAA).

Summary and conclusions

- The LUCID experiment, featuring five Timepix detectors, has been modelled in **GEANT4**.
- **SPENVIS** has been used to obtain input flux estimates for proton and electron fluxes in Low Earth Orbit.
- A web portal, **LUCIDITY**, has been developed to manage simulation parameters and results for the LUCID Collaboration.
- Simulation results indicate that the data rates required as a result of trapped protons and electrons should be within LUCID's capabilities.
- There is still plenty to do before **TechDemoSat-1** launches in February 2014!

Further information:
<http://cernatschool.web.cern.ch>
@CERNatschool
@langtonstar @twhyntie @PhysicsatQM

T. Whyntie^{*,†}, M. Harrison^{*}

^{} Langton Star Centre, [†] Queen Mary, University of London*

CHEP 2013, Amsterdam

Tuesday 15th October 2013