

Status of SUSY after LHC Run I

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“BayesFits” et al

Contents...

1. SUSY and the CMSSM
2. Experimental constraints on SUSY
3. Bayesian statistics
4. Favoured regions of CMSSM
5. Viability & naturalness

Lots of topics =>
Lots of details and
caveats missing...

Flavour of our
work...

SUSY Motivation

- Solves hierarchy problem between Planck and EW scales (stabilises Higgs sector). Radiative top loops cancel with new stop loops
- Lightest neutralino (if lightest SUSY particle) is a WIMP; explains dark matter
- Predicts light Higgs at $\lesssim 130$ GeV
- Unification of gauge couplings (by extended Higgs sector)
- “Predicts” heavy top quark (big Yukawa required for REWSB)
- Explains anomalous magnetic moment of muon, by neutralino/smuon or chargino/muon-sneutrino loops

Summary of CMSSM

- SUSY is broken. General breaking MSSM has ~ 100 free parameters

- **CMSSM**: Four free continuous parameters at GUT:

m_0 = universal soft scalar mass

$m_{1/2}$ = universal soft gaugino mass

A_0 = universal soft trilinear

$\tan \beta$ = the ratio of the two Higgs vevs

Also looking at pheno pMSSM
with 8 params at SUSY scale

- Neutralino-1:

$$m_{\chi} = 0.4m_{1/2}$$

- Neutralino-2:

$$m_{\chi} = 0.8m_{1/2}$$

- Gluino:

$$m_{\tilde{g}} = 2.7m_{1/2}$$

- Stau-1:

$$m_{\tilde{\tau}} = \sqrt{0.15m_{1/2}^2 + m_0^2}$$

Bayesian statistics

- Consider **posterior probability** - probability density of the CMSSM's parameter space given the experimental data
- cf. frequentist statistics (e.g. χ^2) - probability of data given the theory
- Posterior proportional to likelihood times prior (**Bayes' theorem**):

$$p(p_1, p_2, \dots | D) = \mathcal{L}(D | p_1, p_2, \dots) \times \pi(p_1, p_2, \dots)$$

- **Likelihood contains experimental information, often Gaussian:**

$$\mathcal{L} = \exp \left[-\frac{(x - \mu)^2}{2\sigma^2} \right]$$

- Prior contains belief in parameter space before seeing data
- **Bayesian stats is a calculus for beliefs –won't tell us what our prior beliefs ought to be, but how to update them once we see experiments**

Experimental constraints...

Theory error now > exp't

SUSY constrained by:

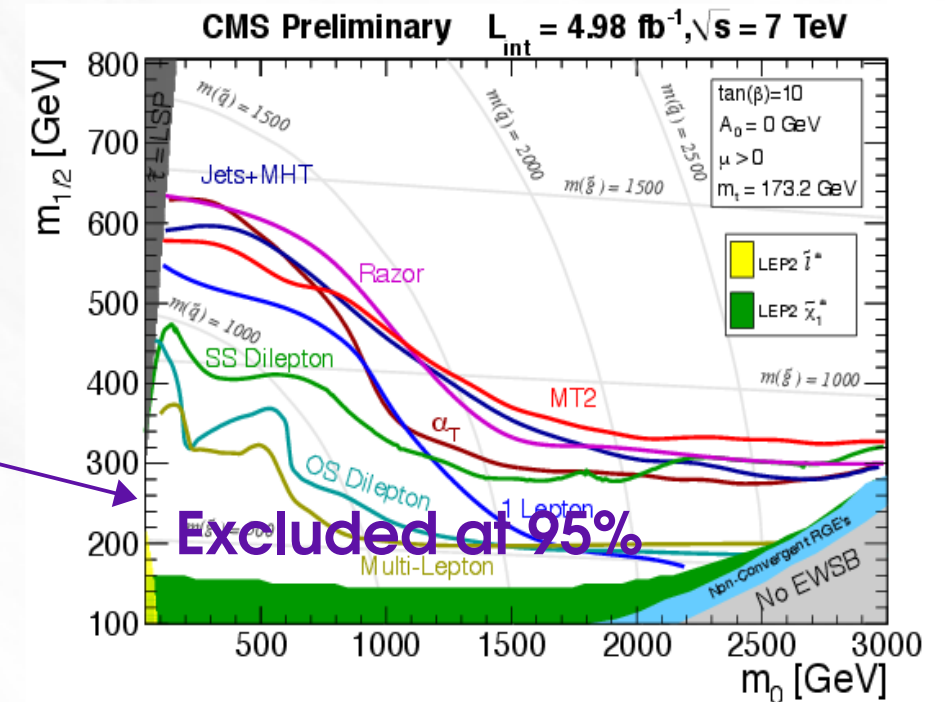
- Direct searches \longleftrightarrow
- Higgs mass \longleftrightarrow
- Dark matter \longleftrightarrow
- EWPO \longleftrightarrow
- g-2 anomaly \longleftrightarrow
- B-physics
- Perhaps also non-SM Higgs rates...

Measurement	Mean	Error: Exp., Th.	Likelihood
CMS razor 4.4/fb		Explained later	Poisson
m_h (GEV)	125	2, 2	Gaussian
$\Omega_\chi h^2$	0.1120	0.0056, 10%	Gaussian
$\sin \theta_{\text{eff}}$	0.23116	0.00013, 0.00015	Gaussian
m_W (GeV)	80.399	0.023, 0.015	Gaussian
$\delta(g-2)_\mu^{\text{SUSY}} \times 10^{10}$	28.7	8.0, 1.0	Gaussian
$\text{BR}(\bar{B} \rightarrow X_s \gamma) \times 10^4$	3.60	0.23, 0.21	Gaussian
$\text{BR}(B_u \rightarrow \tau \nu) \times 10^4$	1.66	0.66, 0.38	Gaussian
ΔM_{B_s} (GeV)	17.77	0.12, 2.40	Gaussian
$\text{BR}(B_s \rightarrow \mu^+ \mu^-)$	$< 4.5 \times 10^{-9}$	0, 14%	Error Fn.

Updated to LHCb measurement in recent BayesFits papers

Including direct LHC SUSY searches

- Signature jets and missing energy, from cascade decay of heavy coloured sparticle, with 2 neutralinos in final state
- No statistically significant excess of events. Takes a large bite out of CMSSM
- We simulated expected SUSY events by MC, including detector efficiency and acceptance, across the m_0, m_{12} plane of the CMSSM



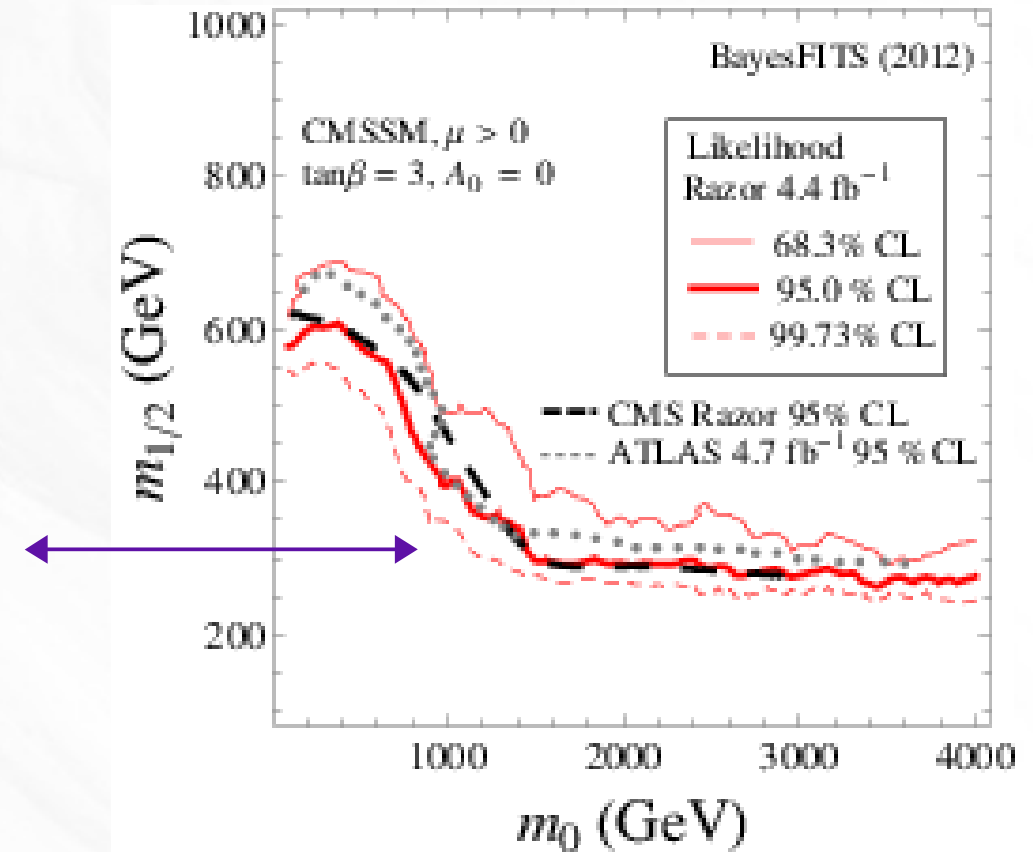
Now a little out-of-date...
Exclusion slightly higher

Likelihood from LHC SUSY search

- Calculated our likelihood with a Poisson, plus systematics on background predictions

$$\mathcal{L} = \text{Po}(o|s + b) \times \text{Syst}(b'|b)$$

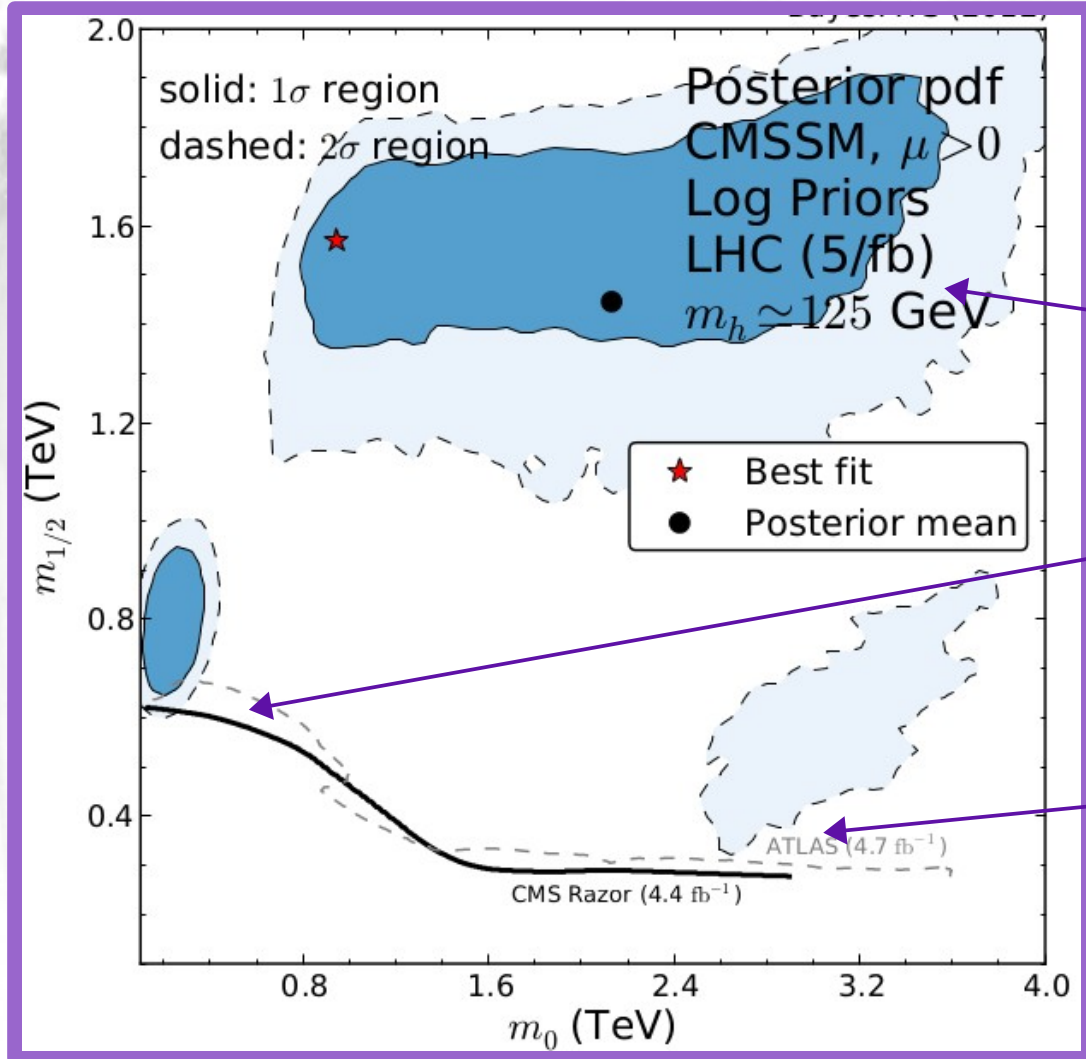
- Official likelihood not published
- Our 95% exclusion contour agrees well with official result for CMSSM ($m_0, m_{1/2}$); this validates method
- We can re-interpret SUSY searches in ANY model! Including ph'cal pMSSM, NMSSM etc



Results...

- Scan parameter space with MultiNest“MC algorithm
- Now present results as 68% and 95% two-dimensional “credible regions”
- These regions contain 68% and 95% of the posterior pdf
- **Reflect degree of belief**
- But don't indicate any (frequentist) coverage (cf. Confidence interval)
- We choose **non-informative log or flat priors** for the CMSSM parameters (no “correct”choice, but some choices are bad...)
- Prior dependence \Leftrightarrow “weak”data

Results (m_0 , $m_{1/2}$)



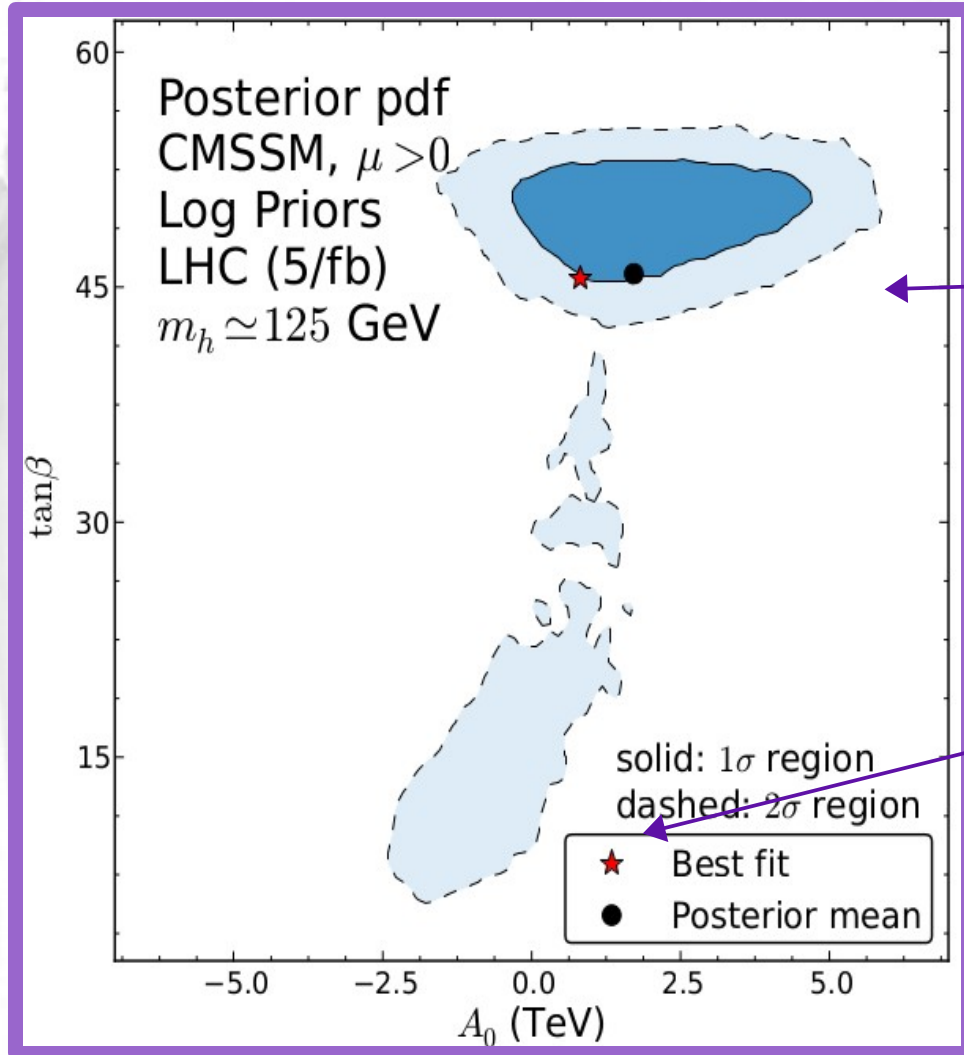
- DM annihilation mechanism shapes plot:

A-funnel region. Neutralinos annihilate via heavy-Higgs resonance

Stau co-annihilation region. Neutralinos co-annihilate with staus to reduce relic density,

Focus point at 2 sigma (sizable Higgsino component)

Results (A_0 , $\tan\beta$)



- DM annihilation again shapes plot:
 - **A-funnel** prefers large $\tan\beta$, to lower m_A and open Higgs resonance. At the expense of flavour physics, which likes small $\tan\beta$.
 - **Stau co-annihilation** prefers smaller $\tan\beta$.

“Viability”

- Frequentist: Find minimum chi-squared \Rightarrow find the probability of obtaining a chi-squared that large (p-value)
- If less than, e.g., 5%, reject model
- Problem: don't know the distribution of the chi-squared, so guess?!
- Problem: frequentist quantities are properties of experiment
- “Let data speak for itself?! \Leftrightarrow frequentist statistics properties of hypothetical, unrealised experiments -information that is NOT in data
- We (+fivals) so far find CMSSM p-value $> 5\%$, even with g-2, we are making a (sensible) guess for chi-squared dist'n

“Viability”

- Bayesian: find probability of model given the data!

$$p(M|D) = p(D|M)p(M)/p(D) \propto \mathcal{Z}$$

- Problem: don't know normalisation $p(D)$ or prior for model $p(M)$, so compare evidence $Z=P(D | M)$ with a reference model, and these factors cancel
- If evidence of, e.g., CMSSM much less than reference model (e.g. SM augmented with DM candidate?!), reject model
- e.g. if $\mathcal{Z}(M1)/\mathcal{Z}(M2) >$ critical value, model 1 is significantly favoured over model 2
- Common interpretation is Jeffreys scale
- This is ongoing...

Naturalness

- Naturalness – fine-tuning of EW scale wrt SUSY parameters.

Barbieri & Giudice measure:

$$\Delta = \max_i \frac{p_i}{M_Z} \frac{\partial M_Z}{\partial p_i}$$

- Failed SUSY searches => large radiative corrections to EW scale => "unnatural"
- Naturalness is a statistical argument (Strumia '99)- formalised by Bayesian stats. Bayesian evidence measures naturalness:

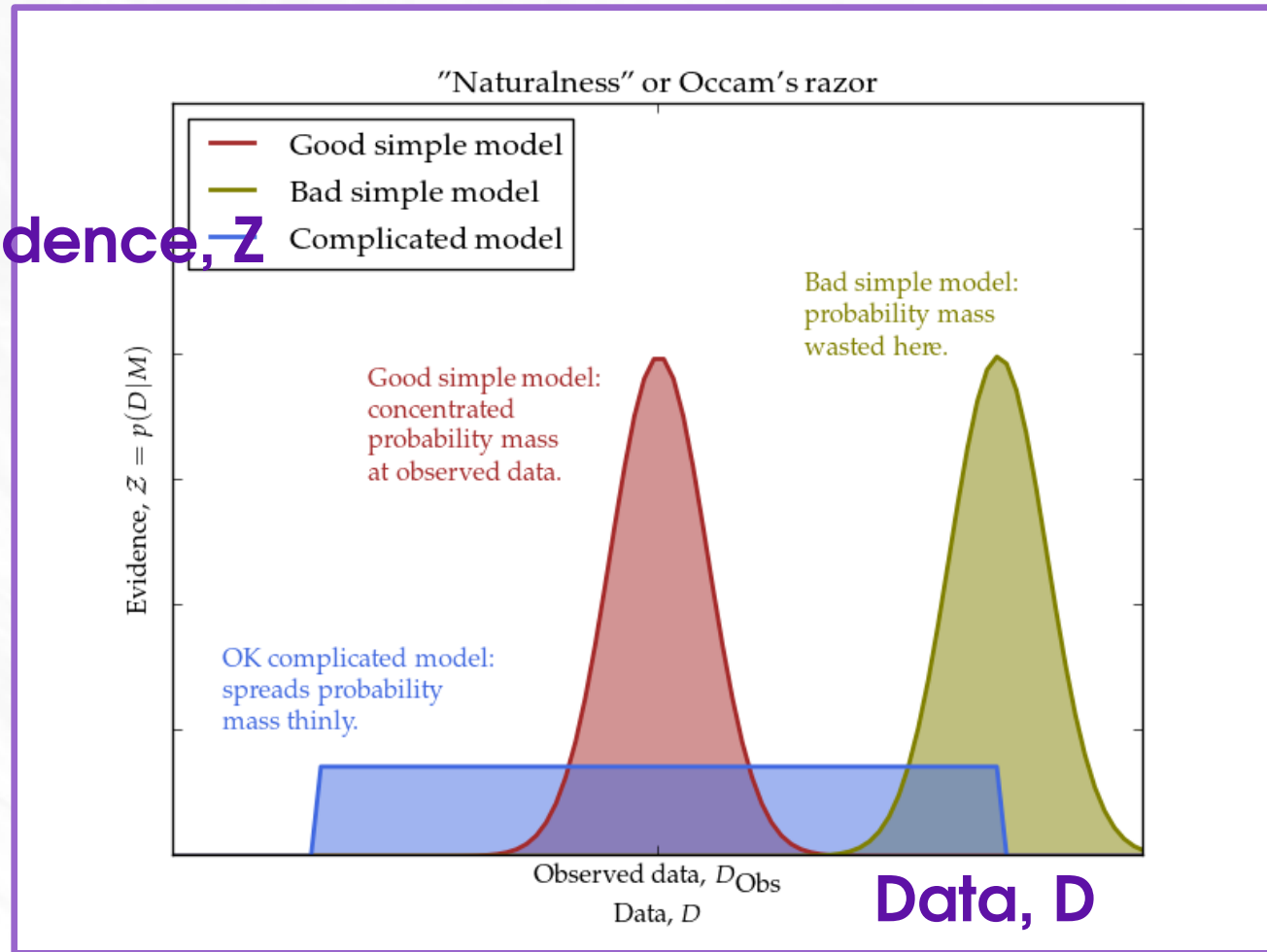
$$\mathcal{Z} = p(D|M) = \int \mathcal{L}(D|p_1, p_2, \dots) \pi(p_1, p_2, \dots) dp_1 dp_2 \dots$$

BIG CAVEAT: "CMSSM" $b, \mu \mapsto \tan \beta, M_Z$
To measure naturalness wrt M_Z , need "fair" priors in b, μ
rather than $\pi(M_Z) = \delta(M_Z - 91.18 \dots)$

Naturalness (cont.)

- Naturalness – Ill-defined? Aesthetic principle? No! Statistical argument, formalised with Bayesian stat.
- Links with Occam's razor & falsifiability – lots of insights

Evidence, Z



Summary

1. Have powerful statistical tools to explore rich parameter space.
2. Even simplest CMSSM viable
3. Though mass scales now high
4. Typical masses: neutralino ~ 0.5 TeV, squarks & gluinos ~ 3 TeV.
5. Naturalness formalised with Bayesian stats

Lots of topics =>
Lots of details and caveats missing...

Search Fowlie or Roszkowski on arXiv for full picture