

Impact of transverse coupling on the *ATLAS* luminosity calibration in the Gaussian beam limit

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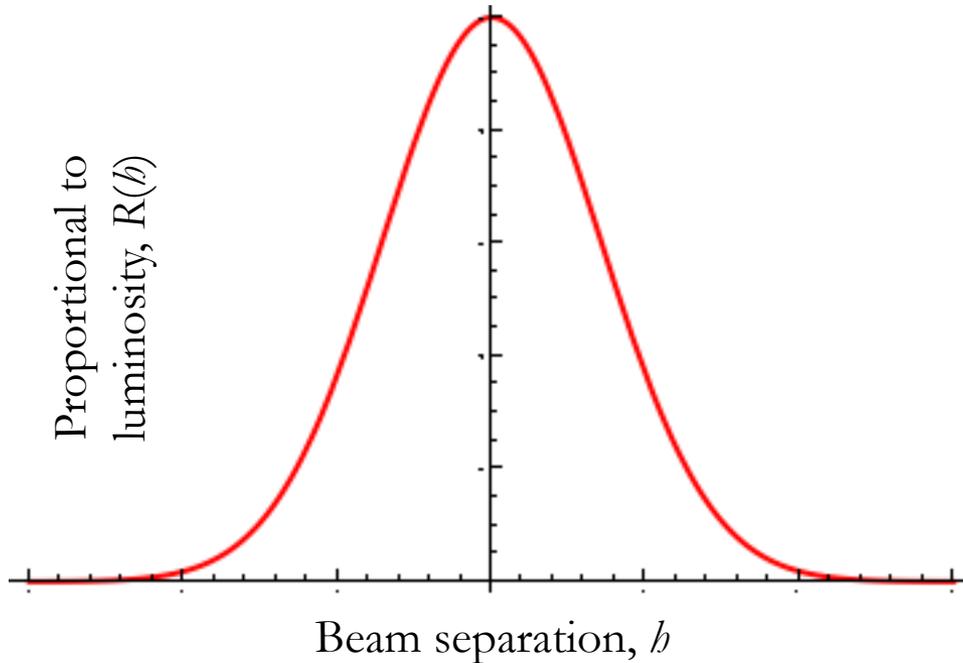
The University of Manchester

IOP Liverpool

8 – 10 April 2013



The van der Meer method



$$\mathcal{L} = \frac{R}{\sigma}$$

The vdM method allows one to calibrate the *absolute* luminosity scale by scanning the beams past one another.

S. van der Meer, *Calibration of the effective beam height in the ISR*.
oai:cds.cern.ch:296752, Tech. Rep. CERN-ISR-PO-68-31, CERN, Geneva, 1968.

$$\Sigma_{x_i} = \frac{1}{\sqrt{2\pi}} \frac{\int dh R_{x_i}(h_{x_i})}{R(0)} \quad \mathcal{L} = \frac{n_b f_r n_1 n_2}{2\pi \Sigma_x \Sigma_y}$$

Implicitly assumes factorisation of x and y components! How does (linear) x - y correlation in the beam affect this method?

The single-Gaussian model

$$\rho_i(x, y, z, t) = \frac{\exp\left(-\frac{1}{2}\vec{x} \cdot \underline{\sigma}_i^{-1} \cdot \vec{x}\right)}{\sqrt{(2\pi)^3 |\underline{\sigma}_i|}}$$

This is the single-Gaussian bunch parameterisation

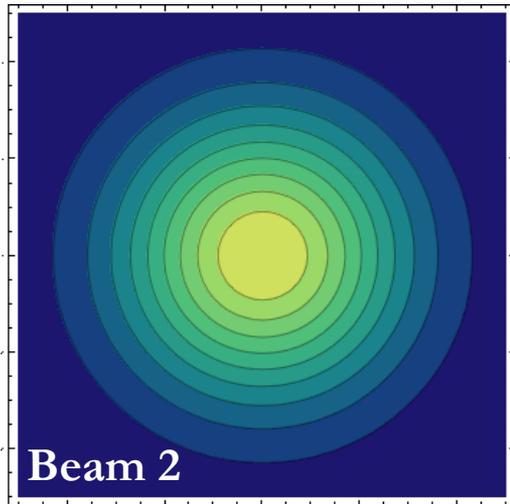
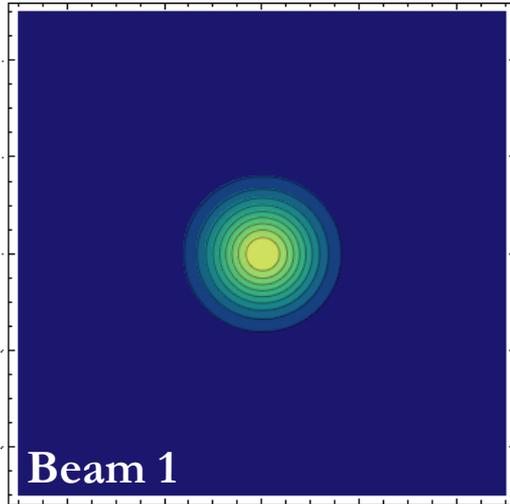
$$\text{where } \underline{\sigma}_i = \begin{pmatrix} \sigma_{x,i}^2 & \kappa_i \sigma_{x,i} \sigma_{y,i} & 0 \\ \kappa_i \sigma_{x,i} \sigma_{y,i} & \sigma_{y,i}^2 & 0 \\ 0 & 0 & \sigma_{z,i}^2 \end{pmatrix} \text{ and } \vec{x} = \begin{pmatrix} x \\ y \\ z \pm z_0(t) \end{pmatrix}$$

The luminosity in this model is proportional to

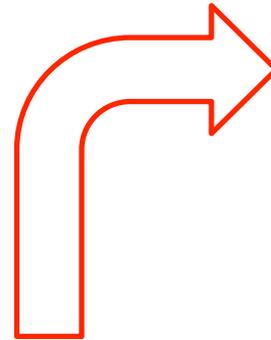
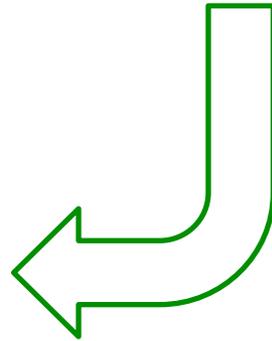
$$\begin{aligned} \mathcal{L}(x, y, z) &\propto \int \rho_1(x, y, z, t) \rho_2(x, y, z, t) dt \\ &= \rho_1(x, y) \rho_2(x, y) \int \rho_1(z, t) \rho_2(z, t) dt \end{aligned}$$

In the absence of explicit beam crossing angle

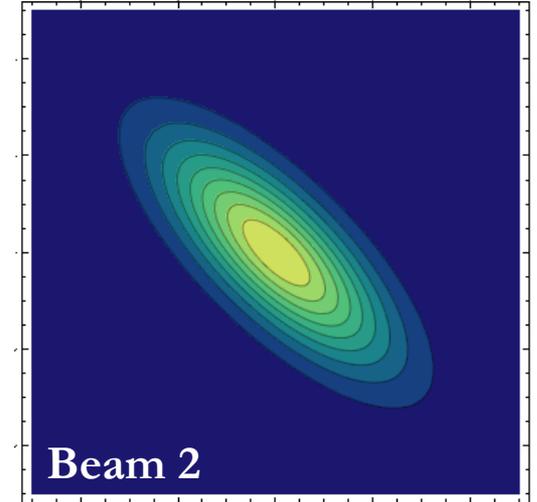
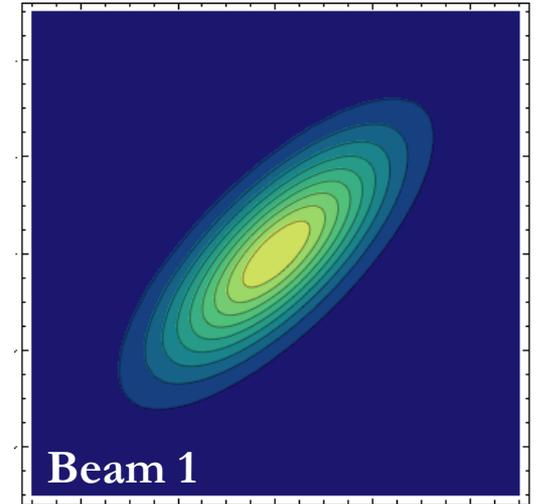
Two illuminating examples

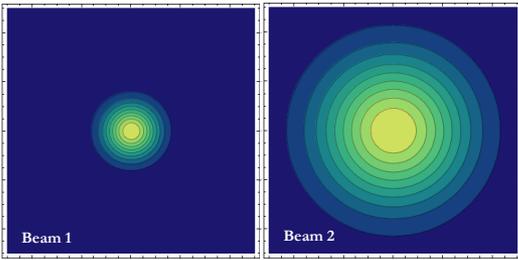


Different beam sizes



x-y correlation

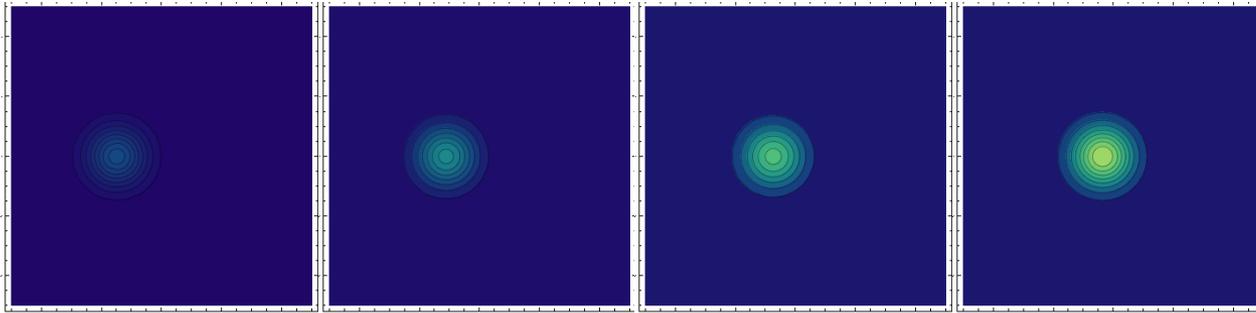




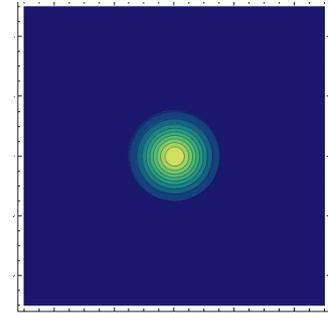
Scanning beams of different size

Luminous region follows the bulk of the **narrow** beam as it scans the wider beam

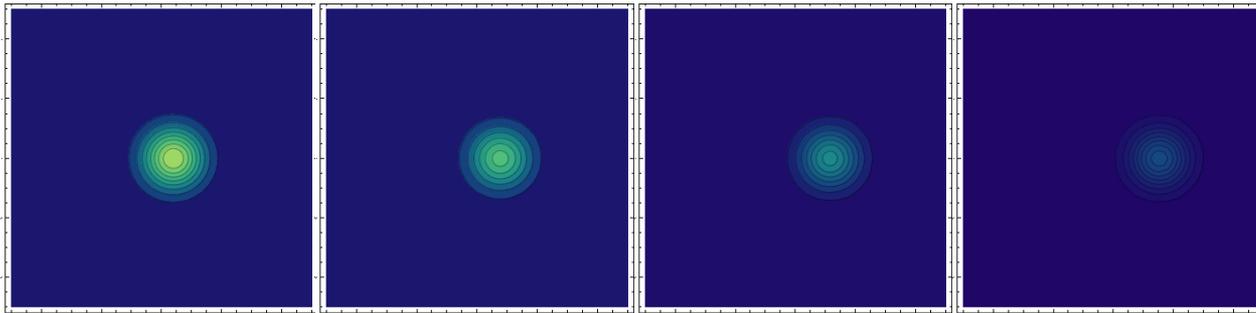
← Higher negative separation

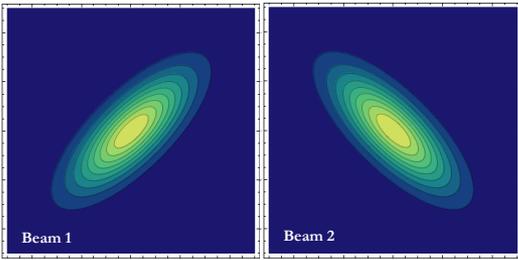


Centred



Higher positive separation →

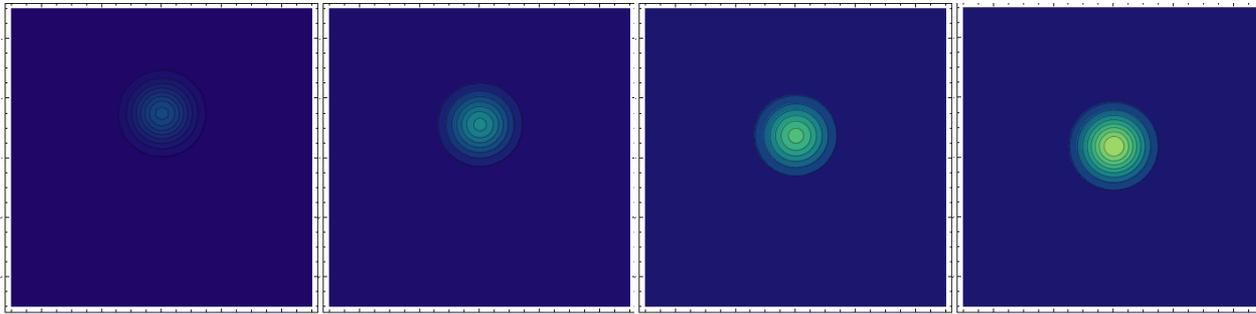




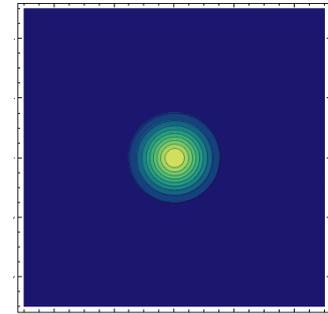
Scanning x-y correlated beams

Movements of luminous region is **transverse** to scan direction

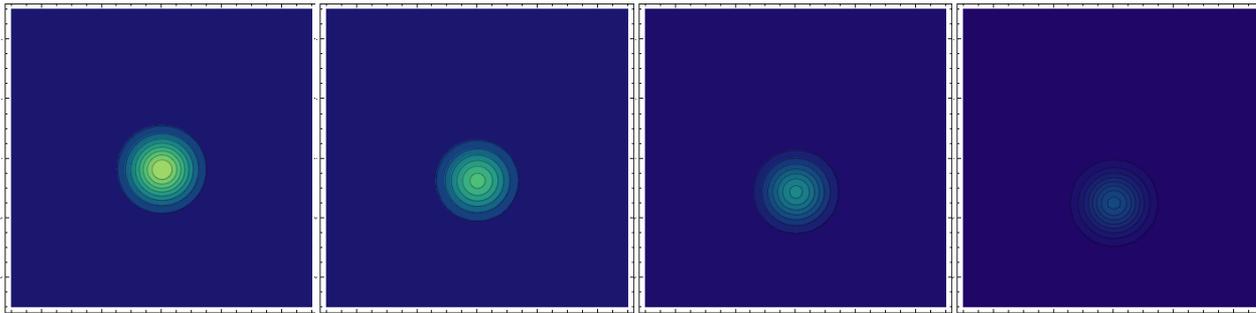
← Higher negative separation



Centred



Higher positive separation →

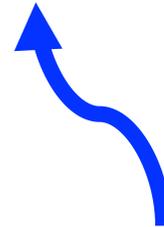
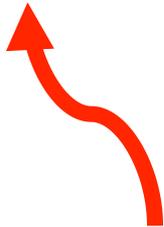
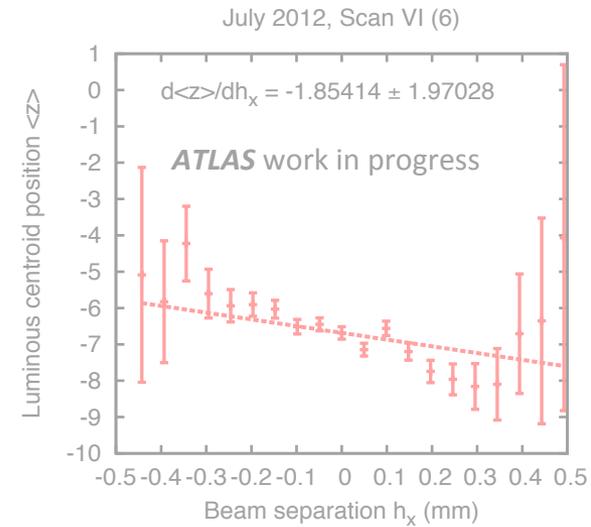
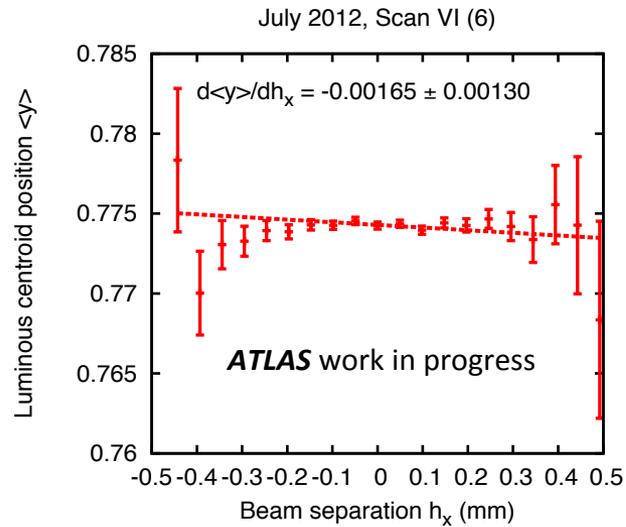
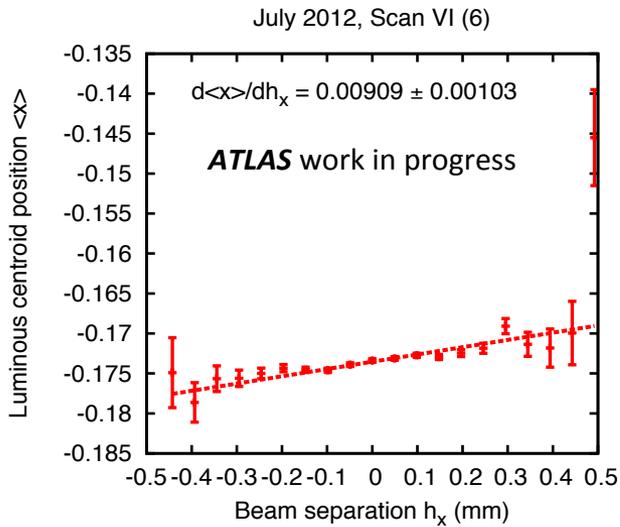


Movements of the luminous centroid

Observed phenomenon	Cause
$\frac{d\langle z \rangle}{dh_{x,y}}$ $\frac{d\langle x \rangle}{dh_x} \neq \frac{d\langle y \rangle}{dh_y} \neq 0$ $\frac{d\langle y \rangle}{dh_x} \sim \frac{d\langle x \rangle}{dh_y} \neq 0$	<p>Beam crossing angle in x-z or y-z plane; α_{xz} and α_{yz}, respectively.</p> <p>Different beam sizes in transverse plane.</p> <p>x-y correlation within each beam.</p>

$$\langle \vec{x} \rangle = h_x \frac{d\langle \vec{x} \rangle}{dh_x} + h_y \frac{d\langle \vec{x} \rangle}{dh_y}$$

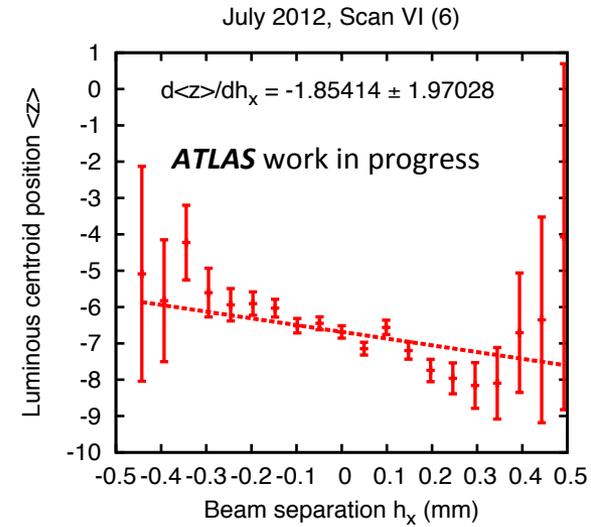
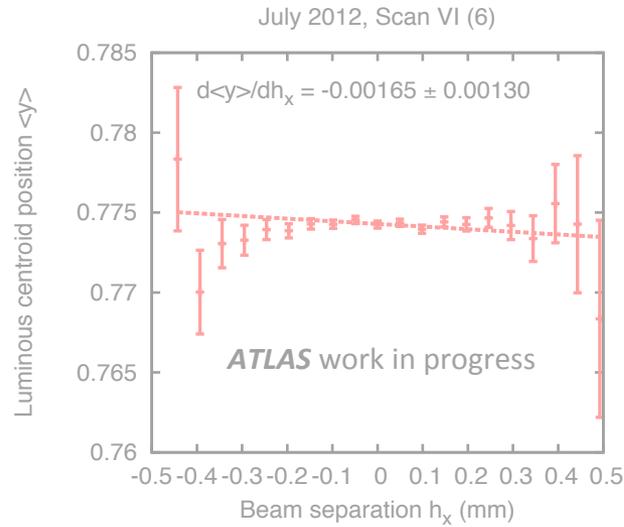
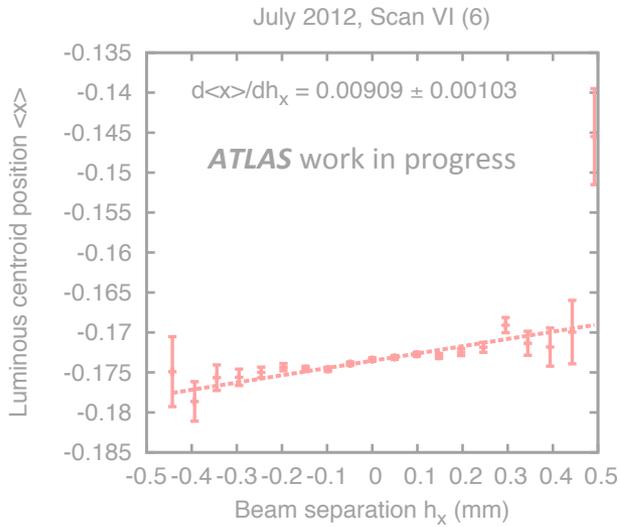
July 2012 scan (6) data



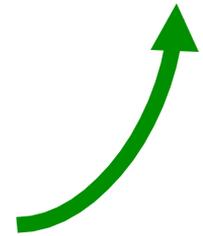
$$\frac{d\langle x \rangle}{dh_x} = \frac{\sigma_{x,1}^2 - \sigma_{x,2}^2}{2(\sigma_{x,1}^2 + \sigma_{x,2}^2)} + \text{h.o.t.}$$

$$\frac{d\langle x \rangle}{dh_y} = \frac{\sigma_{x,1} \sigma_{x,2} (\kappa_2 \sigma_{x,1} \sigma_{y,2} - \kappa_1 \sigma_{x,2} \sigma_{y,1})}{(\sigma_{x,1}^2 + \sigma_{x,2}^2)(\sigma_{y,1}^2 + \sigma_{y,2}^2)} + \text{h.o.t.}$$

July 2012 scan (6) data

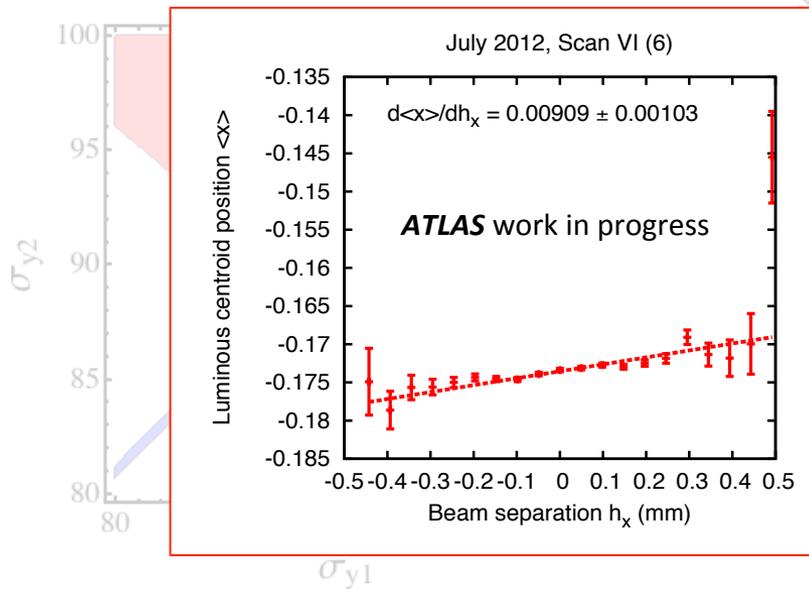
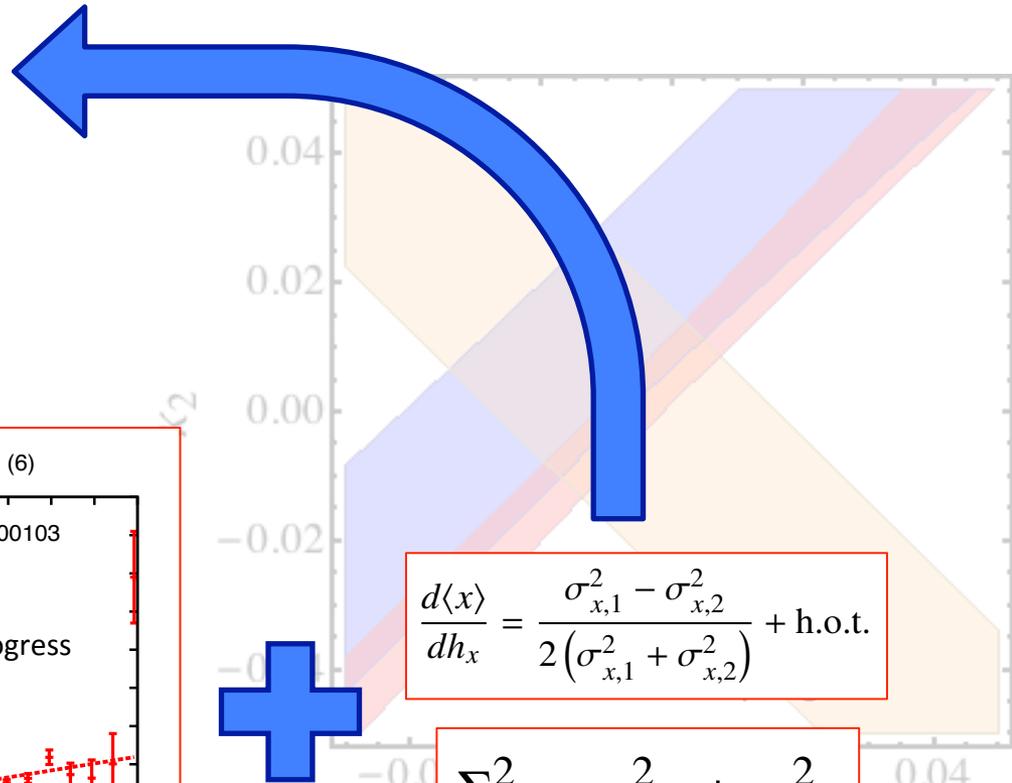
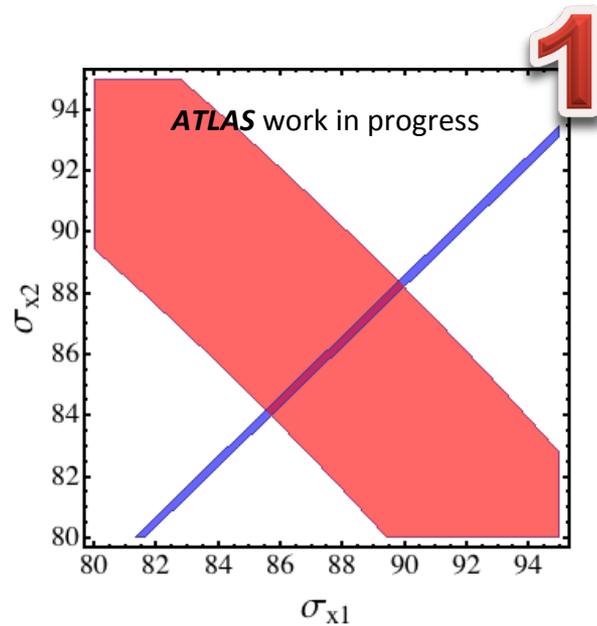


$$\frac{d\langle z \rangle}{dh_x} = \frac{\alpha_{xz} \sigma_z^2}{2(\sigma_{x,1}^2 + \sigma_{x,2}^2)} + \dots$$



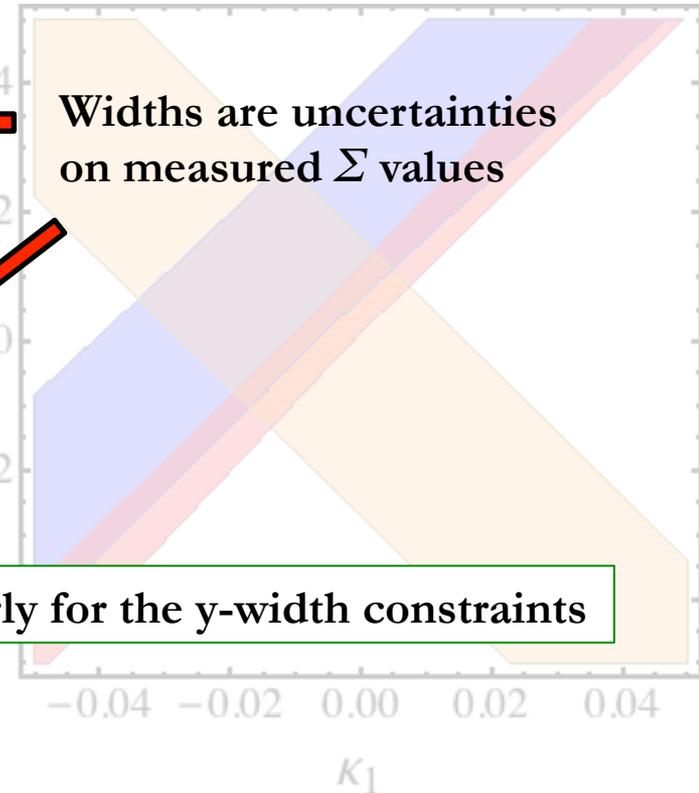
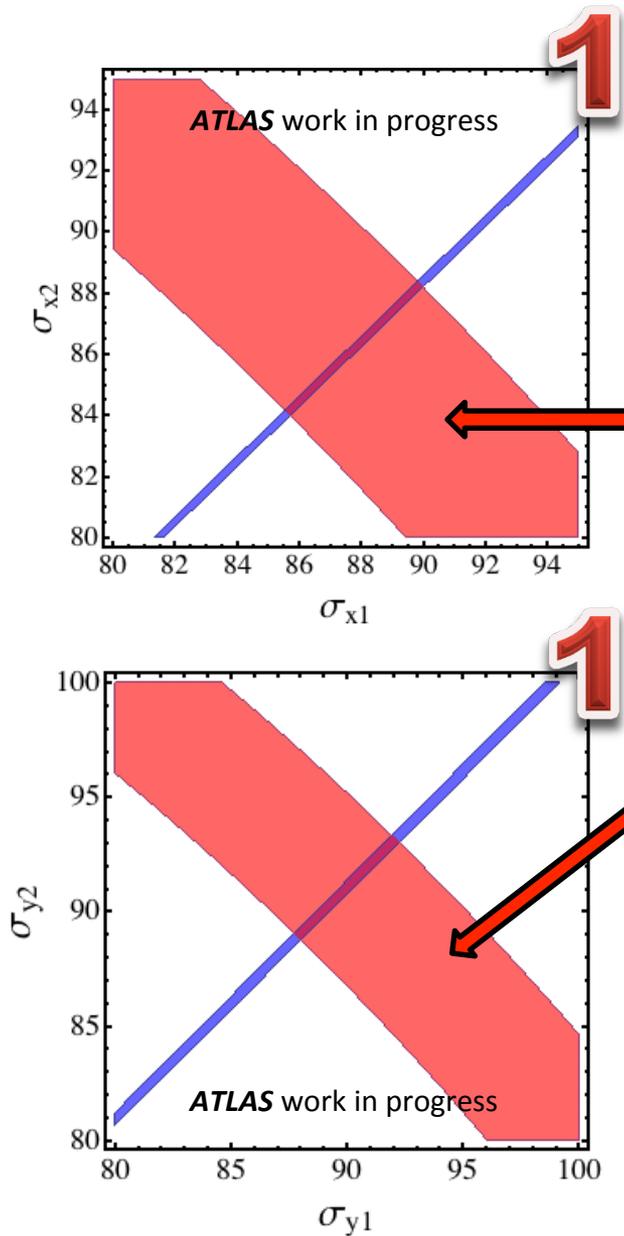
Constraints

(July 2012 scan 6)



Constraints

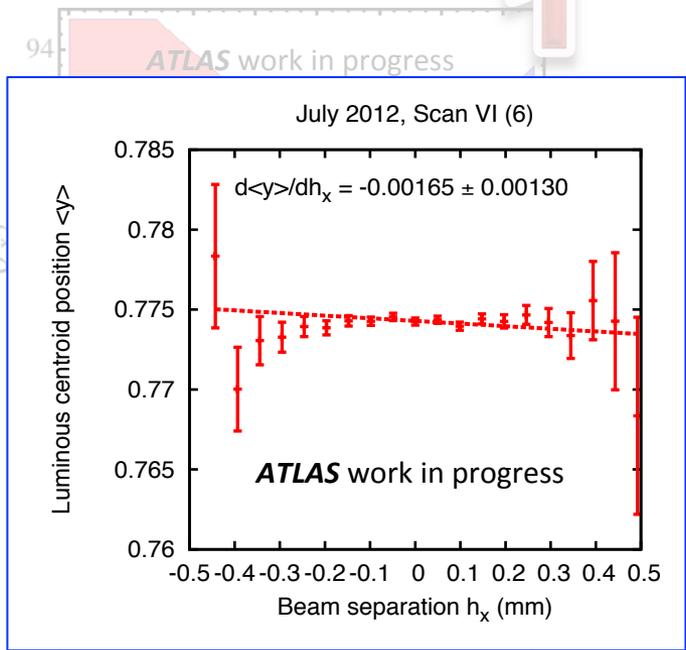
(July 2012 scan 6)



Widths are uncertainties on measured Σ values

Similarly for the y-width constraints

1



$$\frac{d\langle x \rangle}{dh_y} = \frac{\sigma_{x,1} \sigma_{x,2} (\kappa_2 \sigma_{x,1} \sigma_{y,2} - \kappa_1 \sigma_{x,2} \sigma_{y,1})}{(\sigma_{x,1}^2 + \sigma_{x,2}^2)(\sigma_{y,1}^2 + \sigma_{y,2}^2)} + \text{h.o.t.}$$

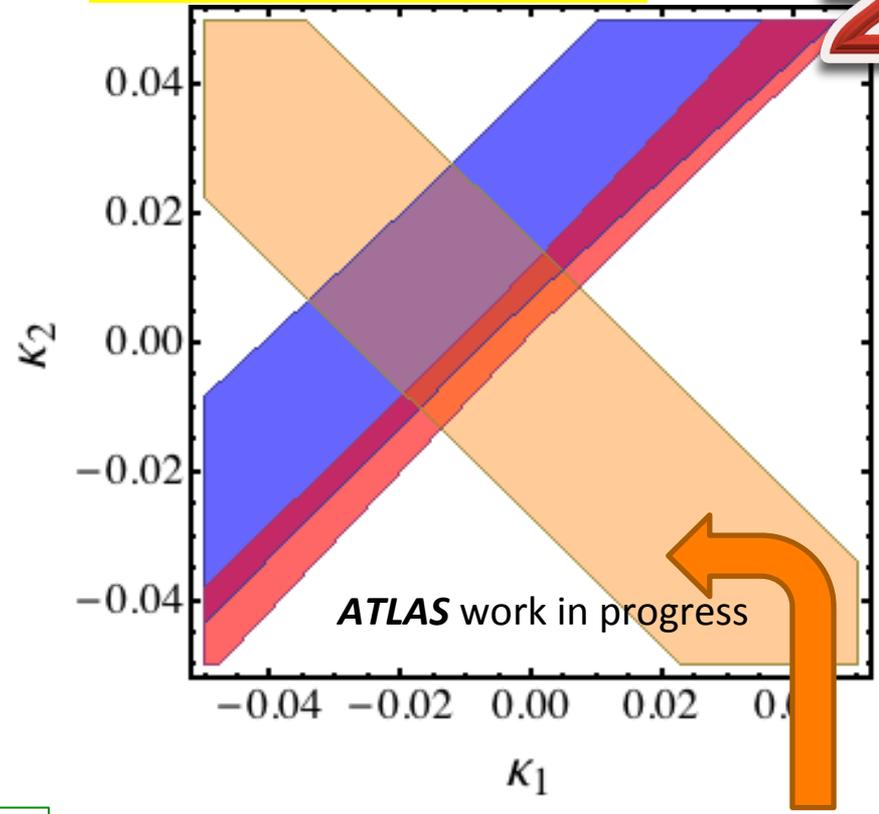
Similarly for the y position during x scan

Constraints

(July 2012 scan 6)

κ is correlation coefficient

2

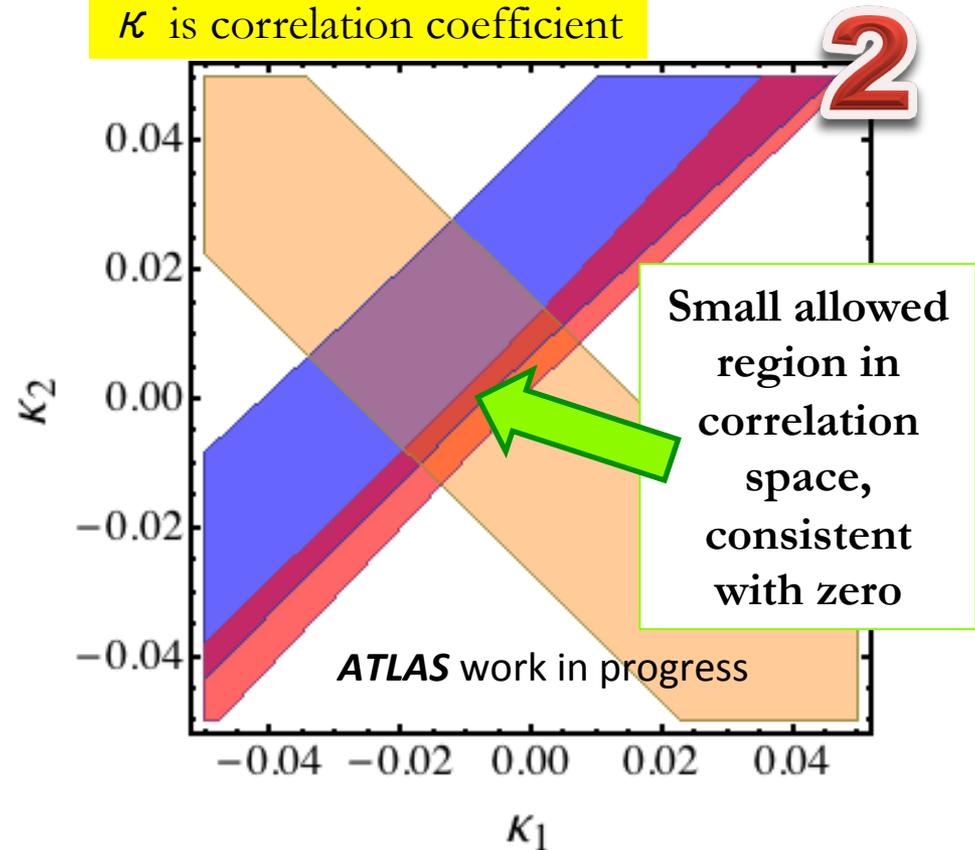
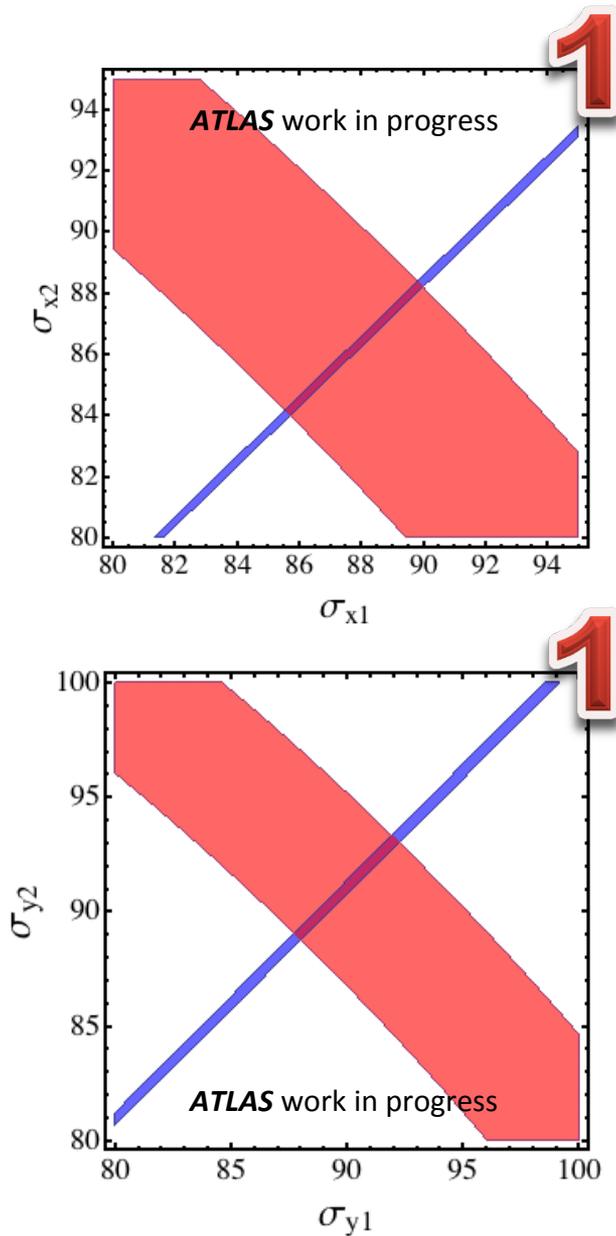


From direct measurement of luminous region correlation

Constraints

(July 2012 scan 6)

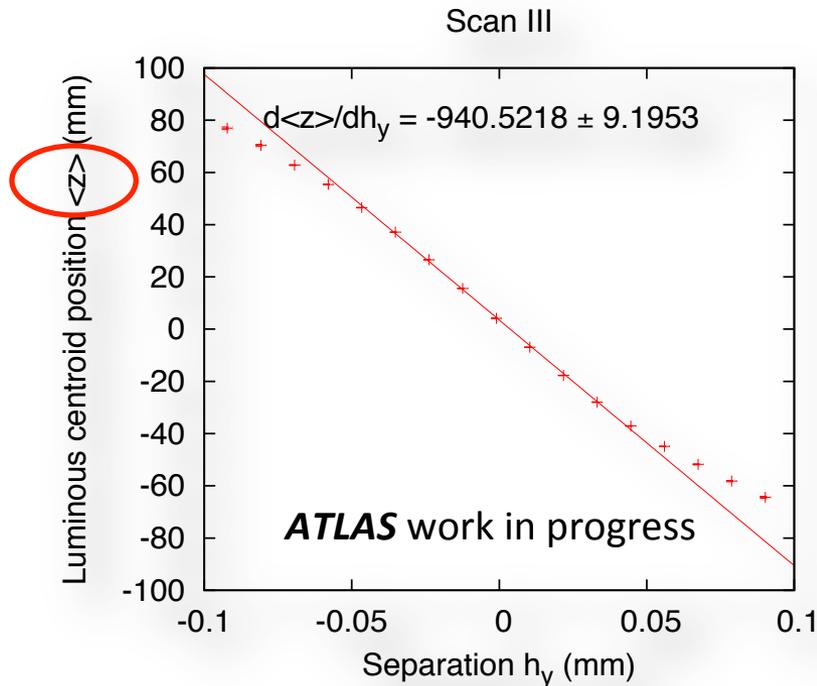
κ is correlation coefficient



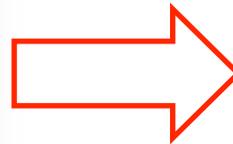
Uncertainties from constraints in (1) are propagated to band widths in (2)

Explicit beam crossing angle

...has been included in the model, essentially functioning as an x - z and a y - z correlation.



...vdM scan III (3) in April 2012



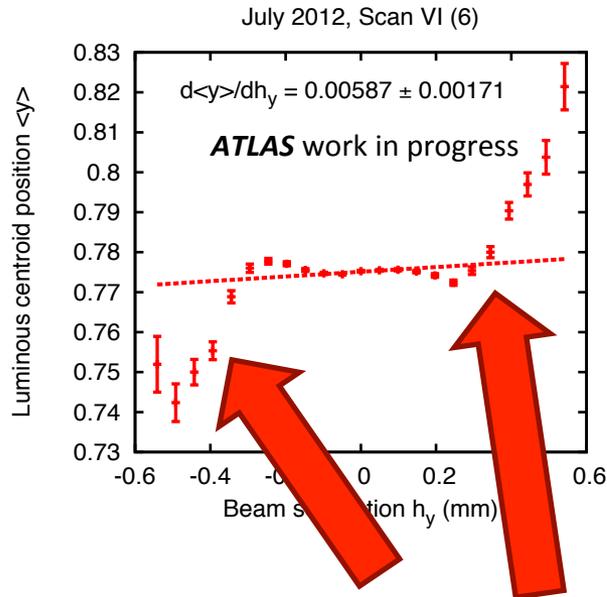
“Significant” crossing angle

Conclusions

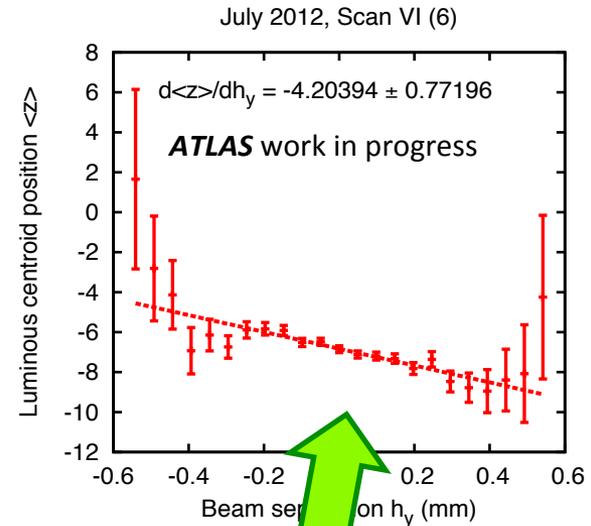
- Permits a fully-analytic approach. Soluble with series solution about small correlation.
- Systematic method for constraining all beam parameters without bootstrapping, iterating, etc.
- Arbitrary crossing-angle easily incorporated, retaining analyticity.
- Applied to vdM scans in *Oct '10, Mar '11, May '11, Apr '12* and *Jul '12*.
- Consistently suggests impact of (linear) transverse coupling on luminosity calibration is negligible.
- One of the systematic contributions that enters an unprecedented 1.8% luminosity error at ATLAS.
- On-going work: Further models have been explored in detail, which successfully account for many of the non-linear features observed in the data.

BACKUP SLIDES

Successes and limitations



**Non-linear tails
cannot be
described using
single-Gaussian
model**



**Crossing angles
are easily
implemented in
single-Gaussian
model**

Further models and work in progress

Double Gaussian

$$\rho(\vec{x}; \dots) = w \rho_A(\vec{x}; \dots) + (1 - w) \rho_B(\vec{x}; \dots)$$

... = model parameters

Further models have been explored in detail, which successfully account for many of the non-linear features observed in the data.

An analytical approach is not always possible, as with the single-Gaussian model. Work is based on solving these as far as possible.

Super Gaussian

$$\rho(x; \dots) = \mathcal{N} \exp \left[-\frac{1}{2} \left(\frac{|x|}{\sigma} \right)^{2+\epsilon} \right]$$

