

# Sterile neutrino as a pseudo-Goldstone fermion

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- introduction and motivations
- theoretical framework
- numerical results: correlations among sterile parameters
- conclusions

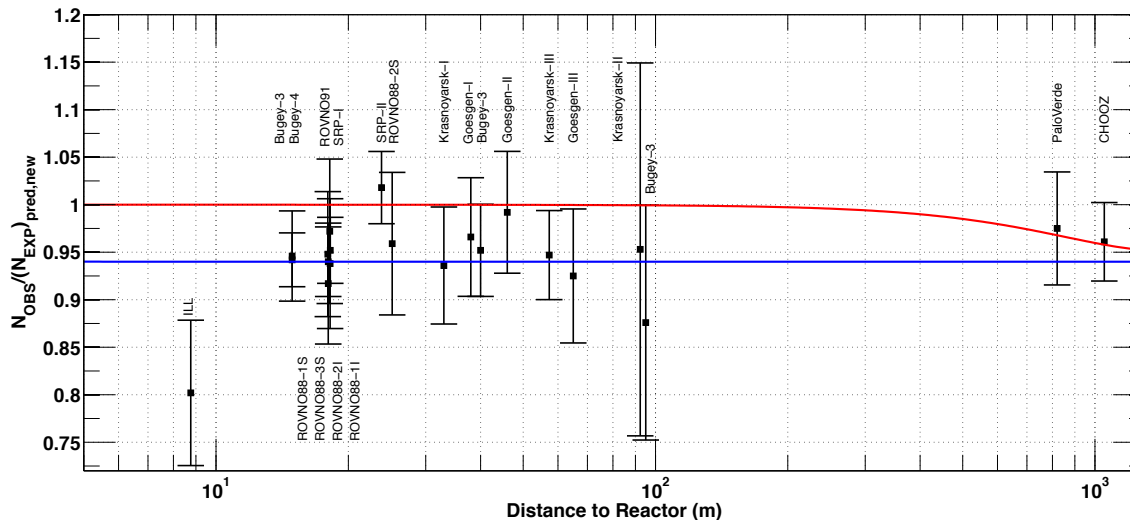
based on a work in progress with Enrico Bertuzzo (IPhT Saclay)

Probing the Standard Model and New Physics at Low and High Energies  
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# Introduction and motivations

Renewed interest in the past few years for sterile neutrinos, mainly driven by experimental anomalies and cosmology:

- the reactor anti-neutrino anomaly (deficit of  $\bar{\nu}_e$  in short-baseline reactor experiments) could be due to oscillations into sterile neutrinos



G. Mention et al.  
(arXiv:1101.2755)

- measurement of CMB anisotropies and other cosmological data are consistent with extra light degrees of freedom

$$N_{\text{eff}} = 4.34^{+0.86}_{-0.88} \quad (68\% \text{ C.L.}) \quad [\text{WMAP 7yr} + \text{BAO} + H_0]$$

$$N_{\text{eff}} = 3.84 \pm 0.40 \quad (68\% \text{ C.L.}) \quad [\text{WMAP 9yr} + \text{eCMB} + \text{BAO} + H_0]$$

Recent Planck data leave less room for a sterile neutrino. One usually quotes:

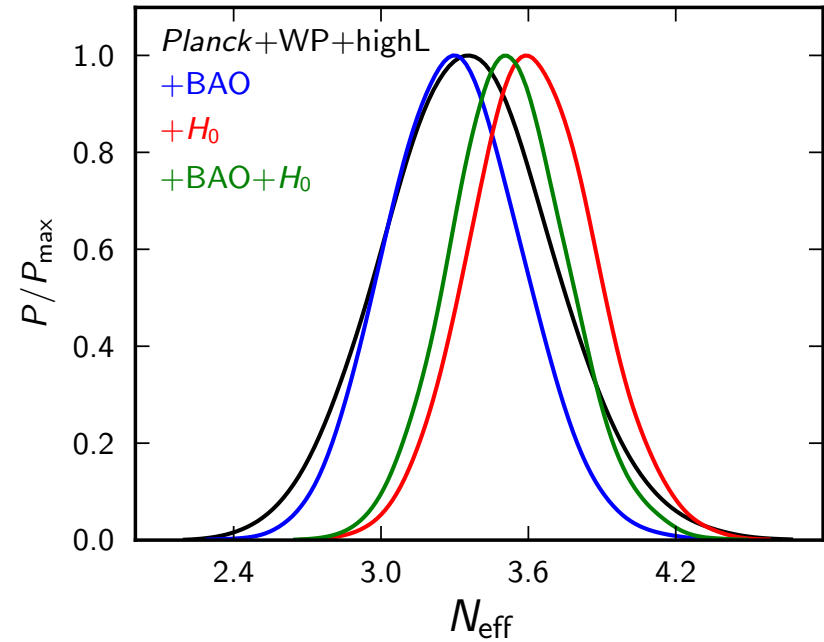
$$N_{\text{eff}} = 3.30^{+0.54}_{-0.51} \quad (95\% \text{ C.L.})$$

[Planck + WMAP + highL + BAO]

However the constraint strongly depends on the set of data used:

$$N_{\text{eff}} = 3.52^{+0.48}_{-0.45} \quad (95\% \text{ C.L.})$$

[Planck + WMAP + highL + BAO +  $H_0$ ]



arXiv:1303.5076

Assuming a fully thermalized massive sterile neutrino, the constraint becomes:

$$N_{\text{eff}} < 3.91, \quad m_{\nu_s} < 0.59 \text{ eV} \quad (95\% \text{ C.L.}) \quad [\text{Planck} + \text{WMAP} + \text{highL}]$$

$$N_{\text{eff}} < 3.80, \quad m_{\nu_s} < 0.42 \text{ eV} \quad (95\% \text{ C.L.}) \quad [\text{Planck} + \text{WMAP} + \text{highL} + \text{BAO}]$$

These bounds are in tension with the sterile neutrino interpretation of the neutrino anomalies, which require  $m_{\nu_s} \sim 1 \text{ eV}$

[see e.g. Mirizzi et al., arXiv:1303.5368]

e.g. a combined analysis of SBL reactor data, gallium calibration experiments and MiniBooNE neutrino data gives [G. Mention et al., arXiv:1101.2755]:

$$|\Delta m_{SBL}^2| > 1.5 \text{ eV}^2, \quad \sin^2 2\theta_{ee} = 0.14 \pm 0.08 \quad (95\% \text{ C.L.})$$

From a theoretical point of view, sterile neutrinos also pose a problem: since they are gauge singlets, their mass is not protected by any symmetry

Sterile neutrinos are present e.g. in the seesaw mechanism:

$$-m \bar{\nu}_L \nu_R - \frac{1}{2} M \nu_R^T C \nu_R + \text{h.c.} = -\frac{1}{2} (\bar{\nu}_L \quad \bar{\nu}_L^c) \begin{pmatrix} 0 & m \\ m & M \end{pmatrix} \begin{pmatrix} \nu_R^c \\ \nu_R \end{pmatrix} + \text{h.c.}$$

the mass eigenstates  $\nu_{L1}$  and  $\nu_{L2}$  are admixtures of the active neutrino  $\nu_L$  and of the sterile neutrino  $\nu_L^c \equiv C \bar{\nu}_R^T$

However, for  $m \ll M$ , the sterile neutrino  $\nu_L^c \simeq \nu_{L2}$  is very heavy and has negligible mixing with the active neutrino

# Theoretical scenarios for naturally light sterile neutrinos

- 1) (very) low-energy seesaw with  $M < 10$  eV [de Gouvêa, Huang, arXiv:1110.6122]  
but the seesaw explanation of small neutrino masses is lost
- 2) singular seesaw mechanism [Glashow '91]  
 $\det M = 0$  leading to a fourth light mass eigenstate  
accidental or due to symmetries of the neutrino sector
- 3) flat extra dimensions [Arkani-Hamed et al. '98; Dienes et al. '98]  
a massless bulk RH neutrino generates a tower of Kaluza-Klein sterile  
neutrinos with masses  $n/R$
- 4) singlet fermions (modulinos) in supersymmetry/string theory  
[Benakli, Smirnov '97; Dvali, Nir '98]
- 5) pseudo-Goldstone fermion [Chun, Joshipura, Smirnov '95; Chun '99]  
supersymmetric partner of the Goldstone boson of a spontaneously  
broken global symmetry (e.g. lepton number or Peccei-Quinn)

# Theoretical framework

Some global symmetry spontaneously broken at a scale  $f \gg M_{\text{SUSY}}$

The supersymmetric effective field theory below  $f$  involves a (pseudo-)Goldstone multiplet

$$A = \frac{s + ia}{\sqrt{2}} + \sqrt{2}\theta\chi + \theta^2 F$$

with a shift symmetry  $A \rightarrow A + ia f$

In the supersymmetric limit and in the absence of explicit global symmetry breaking, all components  $s, a, \chi$  are massless

Supersymmetry breaking can give a mass to  $\chi$  and  $s$ , while some explicit breaking of the global symmetry is needed to give a mass to  $a$  (or the symmetry must be anomalous)

Irreducible  $\chi$  mass from supersymmetry breaking [Cheung, Elor, Hall, 1104.0692]

$$m_\chi \sim m_{3/2} \quad \text{from} \quad \int d^4\theta \frac{1}{M_P} (A + A^\dagger)^2 (X + X^\dagger), \quad \langle X \rangle = F\theta^2$$

$\Rightarrow$  low-scale supersymmetry breaking needed

Assuming no R-parity (but baryon number), the most general Lagrangian compatible with the shift symmetry  $A \rightarrow A + i\alpha f$  is ( $\alpha = 0, 1, 2, 3$ ):

$$W = \mu_\alpha H_u L^\alpha + \frac{1}{2} \lambda_{\alpha\beta k}^e L^\alpha L^\beta \bar{e}^k + \lambda_{\alpha j k}^d L^\alpha Q^j \bar{d}^k - \lambda_{j k}^u H_u Q^j \bar{u}^k$$

$$K = \frac{1}{2} (A + A^\dagger)^2 + H_u^\dagger H_u + L^{\alpha\dagger} L^\alpha + C_u H_u^\dagger H_u \frac{A + A^\dagger}{f} \\ + C_{\bar{\alpha}\beta} L^{\alpha\dagger} L^\beta \frac{A + A^\dagger}{f} + \left( C_{u\alpha} H_u L^\alpha \frac{A + A^\dagger}{f} + \text{h.c.} \right) + \dots$$

$V_{\text{soft}}$  = generic MSSM soft terms with leptonic RPV

After minimization of the scalar potential,  $H_u, L_\alpha$  get vevs (assume  $\langle A \rangle = 0$ )

$\Rightarrow$  non-canonical kinetic terms for  $H_u$  and  $L_\alpha$

$\rightarrow$  redefine  $H_u$  and  $L_\alpha = (H_d, L_i)$  such that

- (i) the charged fields  $H_u^+, H_d^-, e_i^-$  have canonical kinetic terms
- (ii) the sneutrino vevs  $\langle \tilde{\nu}_i \rangle$  vanish
- (iii)  $\lambda_{0jk}^e = \lambda_j^e \delta_{jk}$ ,  $\lambda_j^e$  real

## Neutralino and chargino mass matrices

As a consequence of the bilinear RPV terms ( $\mu_i H_u L_i$ ), leptons mix with charginos and neutralinos.

Furthermore, since the kinetic terms of the neutral fields  $H_u^0, H_d^0, \nu^i, A$  are not canonical, the neutralino mass matrix receives contribution from the Kähler potential and mixes  $\chi$  with the standard neutrinos and neutralinos

Charginos:

$$\begin{pmatrix} \widetilde{W}^- & \widetilde{H}_d^- & e_i^- \end{pmatrix} \begin{pmatrix} M_2 & gv_u & 0_{1 \times 3} \\ gv_d & \mu & 0_{1 \times 3} \\ 0_{3 \times 1} & \mu_i & \lambda_i^e v_d \delta_{ij} \end{pmatrix} \begin{pmatrix} \widetilde{W}^+ \\ \widetilde{H}_u^+ \\ \bar{e}_k^+ \end{pmatrix}$$

2 heavy mass eigenstates (charginos)

3 light mass eigenstates (charged leptons) with masses  $m_i = \lambda_i^e v_d$

chargino / charged lepton mixing suppressed by  $\mu_i / \mu$  or smaller



## Neutralinos:

The neutralino mass matrix, written in the basis  $(\tilde{W}^3, \tilde{B}, \tilde{H}_u^0, \tilde{H}_d^0, \nu_i, \chi)$  is a 8x8 matrix with a seesaw structure

$$M_N = \begin{pmatrix} M_{4 \times 4} & \mu_{4 \times 4} \\ \mu_{4 \times 4}^T & m_{4 \times 4} \end{pmatrix} \quad m, \mu \ll M$$

hence the neutrino masses and mixing are given by the diagonalization of the 4x4 effective neutrino mass matrix

$$M_\nu = m - \mu^T M^{-1} \mu = \begin{pmatrix} A \frac{\mu_i}{\mu} \frac{\mu_j}{\mu} & \left( B \frac{\mu_i}{\mu} + D_i \right) \frac{v}{f} \\ \left( B \frac{\mu_j}{\mu} + D_j \right) \frac{v}{f} & C \frac{v^2}{f^2} + m_\chi \end{pmatrix}$$

where  $A = \frac{\mu^2 M_{11} M_Z^2 \cos^2 \beta}{\det M}$  depends only on MSSM parameters, while

B, D<sub>i</sub> and C depend also on the Kähler parameters  $C_{ud}, C_u, C_{\bar{d}d}, C_{\bar{l}j}$  (Rp-conserving) and  $C_{u\bar{u}}, C_{\bar{d}l}$  (RPV). Assuming the former are of order 1, one has  $B, C = \mathcal{O}(\mu)$

One gets a consistent neutrino phenomenology by assuming that all RPV parameters  $(\mu_i, C_{ui}, C_{\bar{d}i})$  are small, while the Rp-conserving Kähler parameters are of order 1

In practice, need  $\frac{\mu_i}{\mu} \lesssim 10^{-5}$ ,  $C_{\bar{d}i} \lesssim 10^{-6}$ ,  $\frac{v}{f} \lesssim 10^{-6}$ ,  $m_\chi \lesssim 1 \text{ eV}$

Can rewrite the 4x4 neutrino mass matrix in a more compact form:

$$M_\nu = \begin{pmatrix} D\epsilon_\alpha\epsilon_\beta & E\eta_\alpha \\ E\eta_\beta & F \end{pmatrix} \quad \sum_\alpha \epsilon_\alpha^2 = \sum_\alpha \eta_\alpha^2 = 1$$

where we have renamed the indices  $i, j \rightarrow \alpha, \beta = e, \mu, \tau$

This structure implies:

1) the matrix has rank 3, so  $m_1 = 0$

2) the active-sterile mixing is given by  $U_{\alpha 4} \simeq \frac{E}{F} \eta_\alpha$  ( $m_4 \simeq F$ )

3) at order 1 in the active-sterile mixing, the active neutrino parameters are given by the matrix

$$(m_\nu)_{\alpha\beta} = D\epsilon_\alpha\epsilon_\beta - \frac{E^2}{F} \eta_\alpha\eta_\beta$$

$$(m_\nu)_{\alpha\beta} = \sum_{i=2}^3 m_i U_{\alpha i} U_{\beta i} = D \epsilon_\alpha \epsilon_\beta - \frac{E^2}{F} \eta_\alpha \eta_\beta$$

Since we know the active neutrino parameters (neglecting CPV), we can “reconstruct” the sterile neutrino parameters using [when the first term dominates]

$$m_2 \simeq -\frac{E^2}{F} \vec{\zeta}^2, \quad m_3 \simeq D$$

$$U_{\alpha 2} \simeq -(\xi_\mu \xi_\tau, \xi_e \xi_\tau, \xi_e \xi_\mu) / \kappa, \quad U_{\alpha 3} \simeq (\epsilon_e, \epsilon_\mu, \epsilon_\tau)$$

where  $\vec{\zeta} \equiv \vec{\epsilon} \times \vec{\eta}$      $\xi_\alpha \equiv \zeta_\beta \zeta_\gamma + \zeta^2 \epsilon_\beta \epsilon_\gamma$  ( $\alpha, \beta, \gamma$  all different)

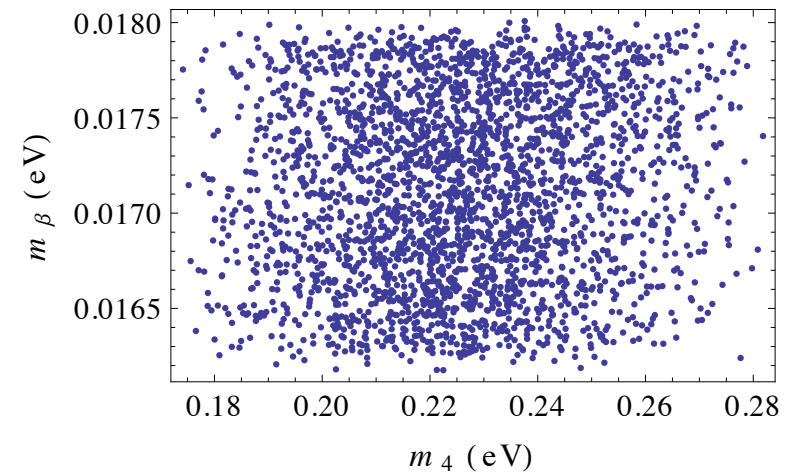
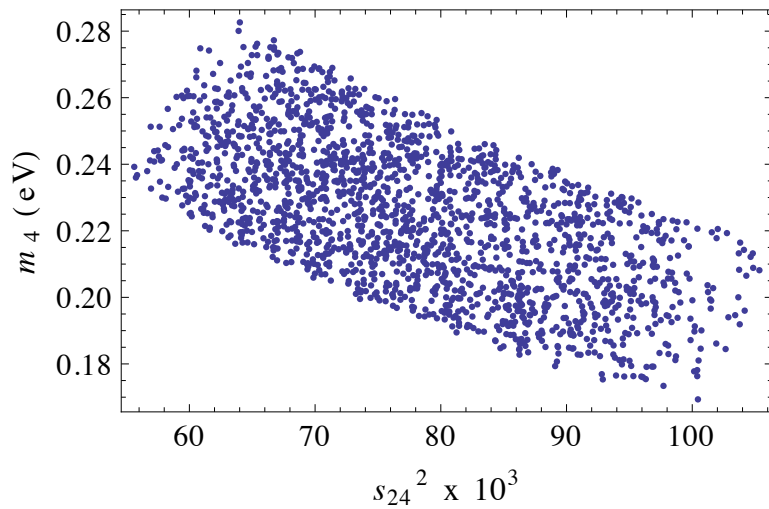
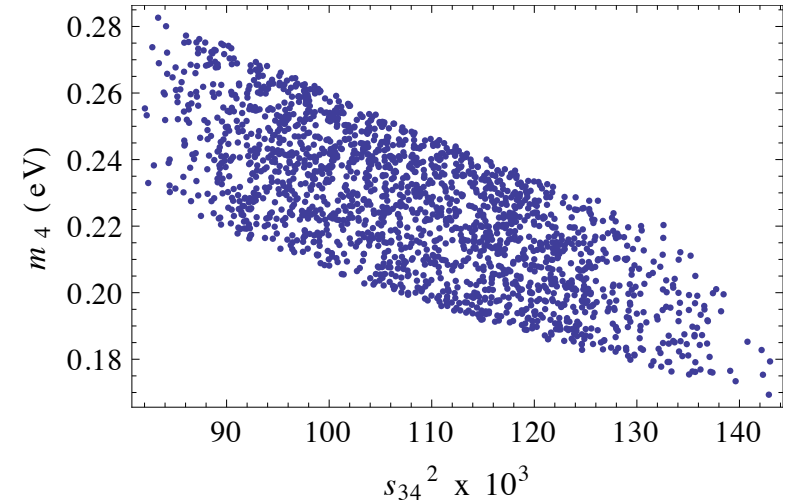
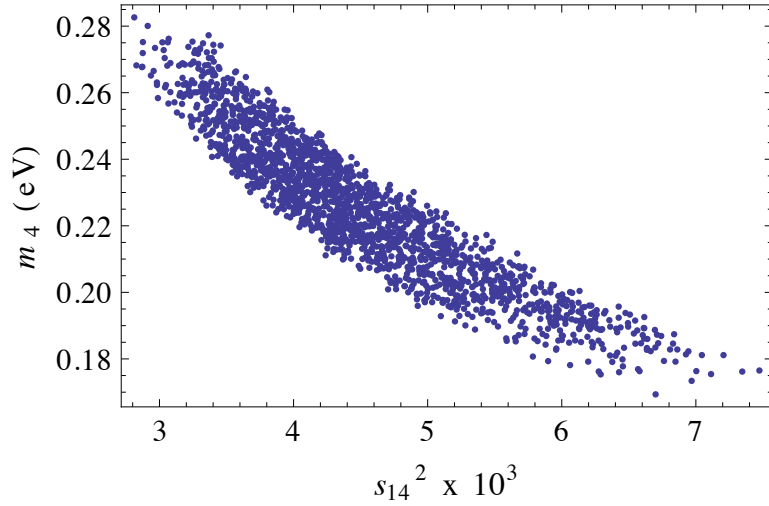
and  $\kappa \equiv (|\xi_\mu \xi_\tau|^2 + |\xi_e \xi_\tau|^2 + |\xi_e \xi_\mu|^2)^{1/4}$

In the reconstruction process, one obtains the  $\zeta_\alpha^2$  as a function of  $\kappa/\vec{\zeta}^2$ , which is the solution of a polynomial of degree 4. Among the solutions, only the ones that satisfy the constraint  $\vec{\zeta}^2 \leq 1$  (if any) are acceptable

Finally, one identifies  $\frac{E^2}{F} \eta_\alpha \eta_\beta \equiv m_4 U_{\alpha 4} U_{\beta 4} \Rightarrow$  correlations

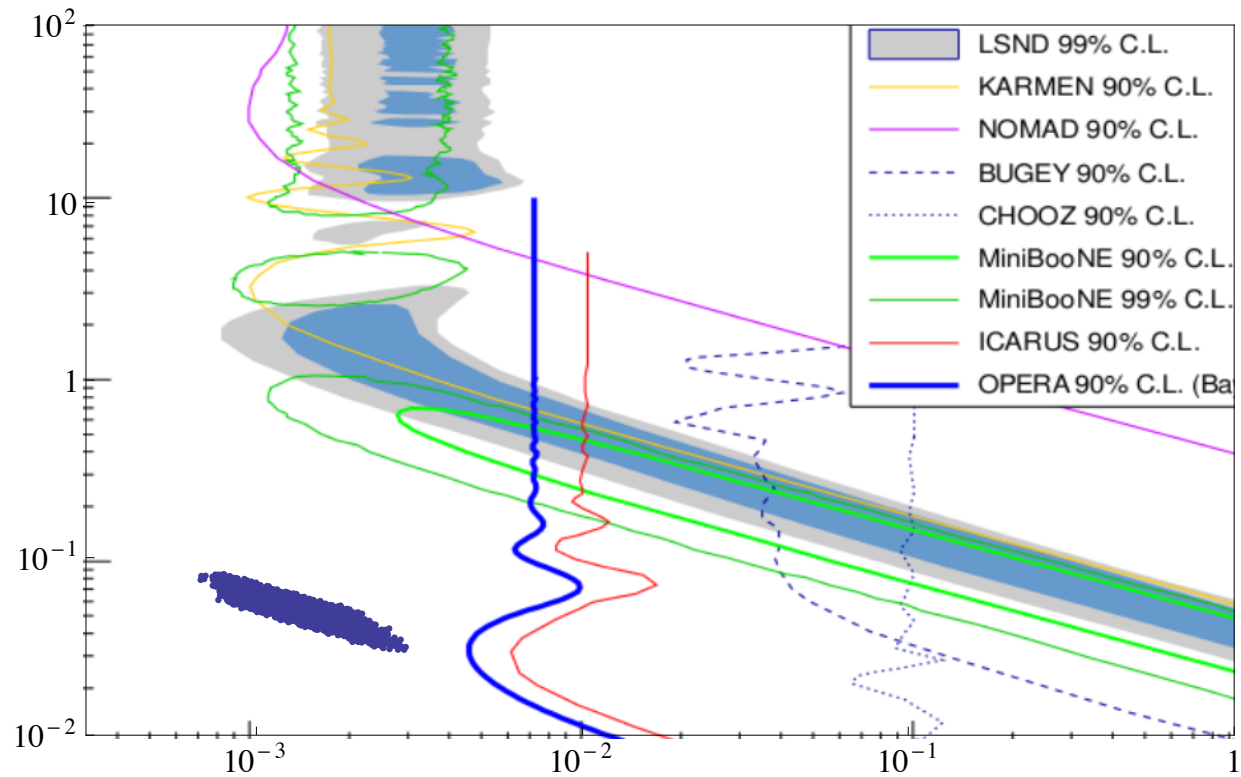
# Numerical results

Normal hierarchy, case 1  $D \epsilon_\alpha^2 \ll (E^2 / F) \eta_\alpha^2$

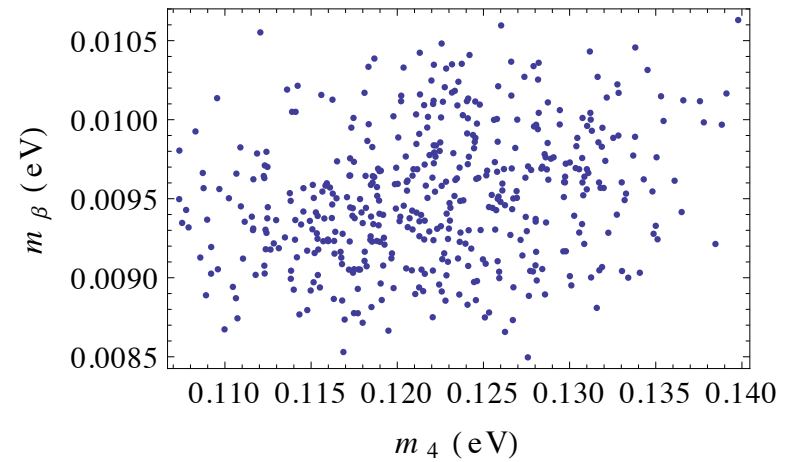
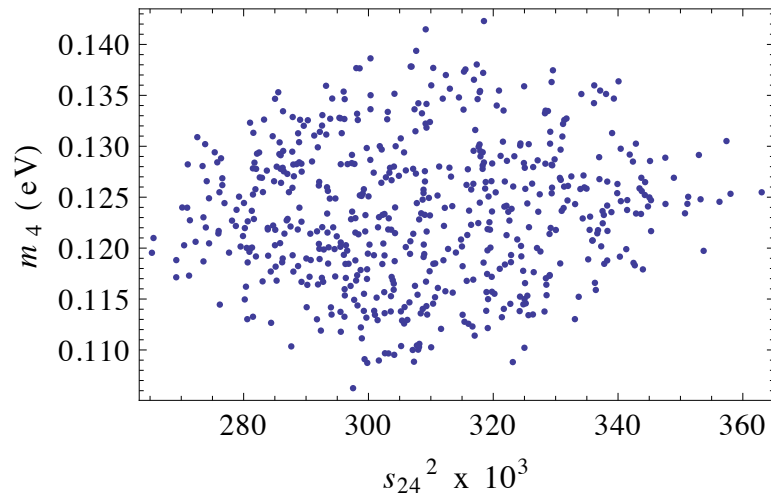
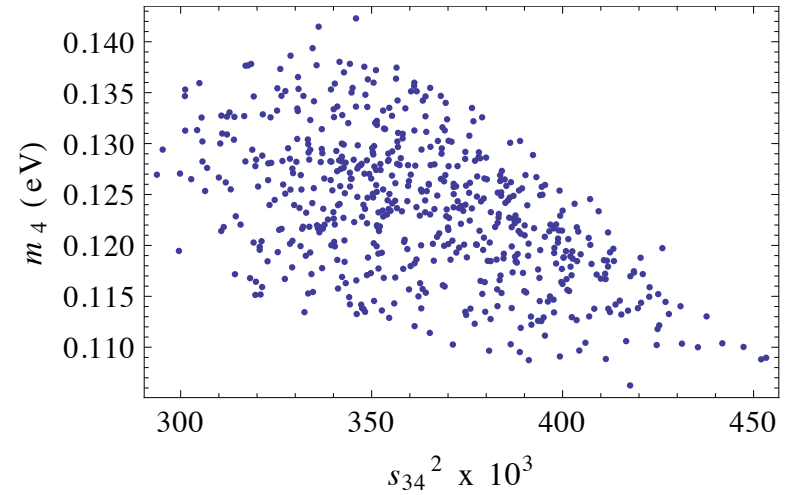
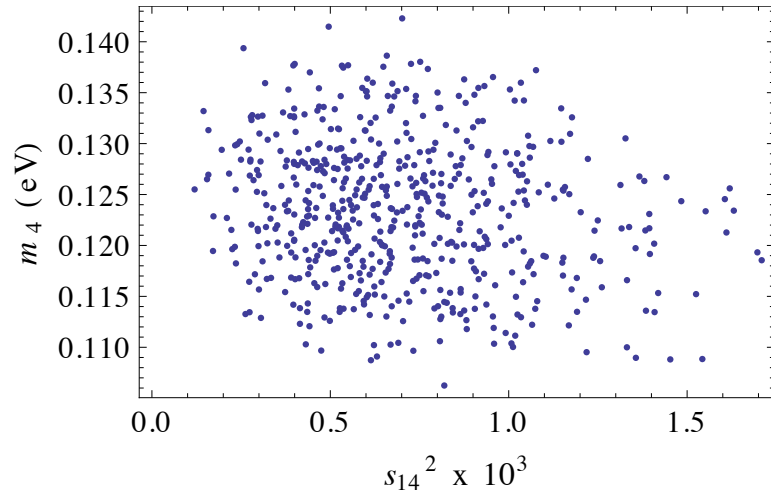


Not relevant to the reactor anomaly

Not relevant to LSND / MiniBooNE

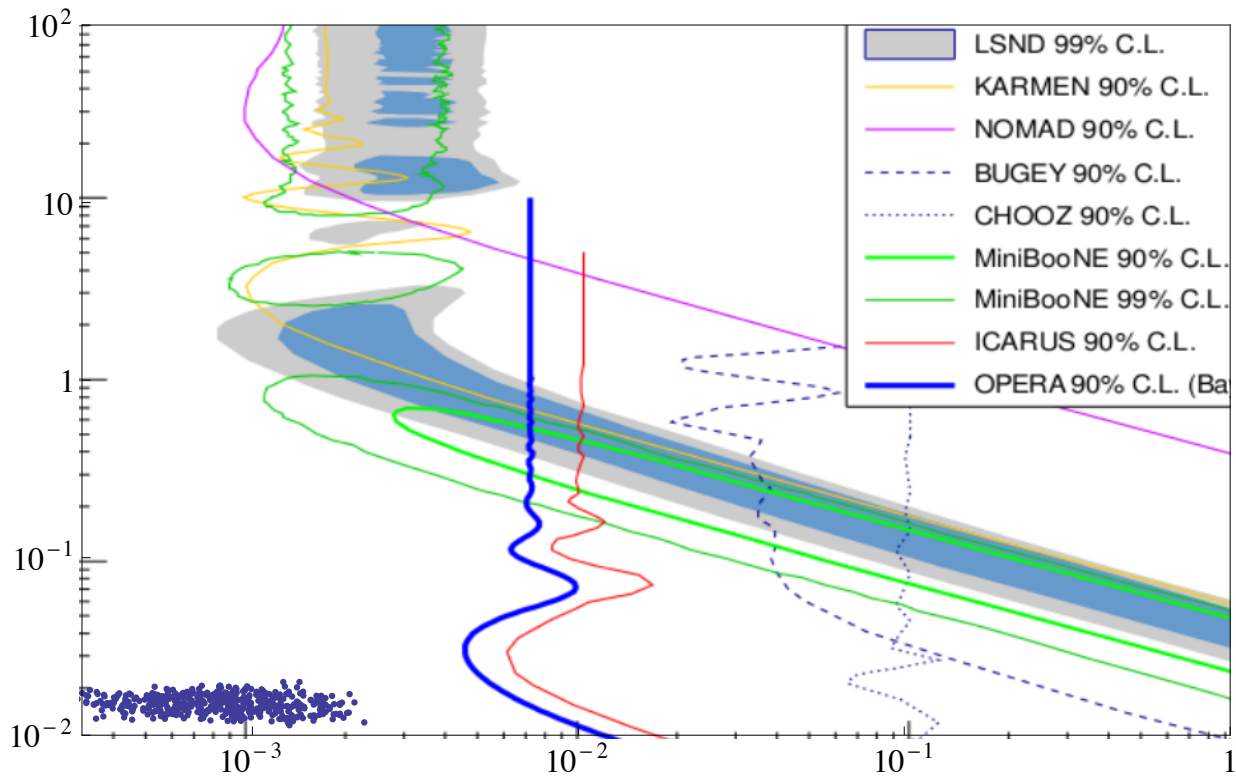


Normal hierarchy, case 2  $(E^2/F)\eta_\alpha^2 \ll D\epsilon_\alpha^2$



Not relevant to the reactor anomaly

Not relevant to LSND / MiniBooNE



## Inverted hierarchy

No numerical solution found so far

Under study



# Conclusions

Sterile neutrino as a pseudo-Goldstone fermion within R-parity violating supersymmetry provides a surprisingly predictive scenario

Correlations between the sterile neutrino mass and the active-sterile mixing (in spite of a large number of parameters)

Does not explain the reactor neutrino anomaly nor LSND/MiniBooNE, but could be tested in future appearance experiments

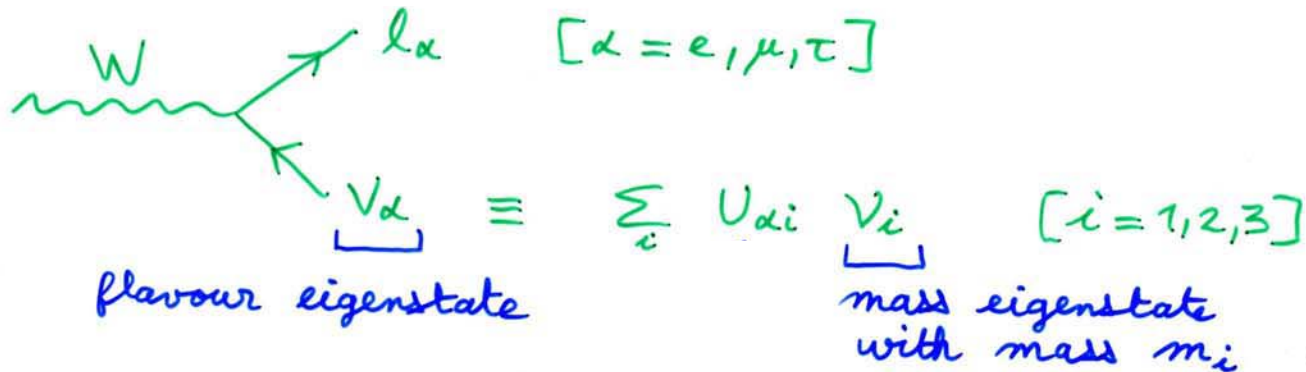
Does not seem to be consistent with an inverted hierarchy in the active neutrino sector (to be confirmed)

In progress...

**BACK UP**

# Active-sterile neutrino mixing

Standard case (3 flavours):



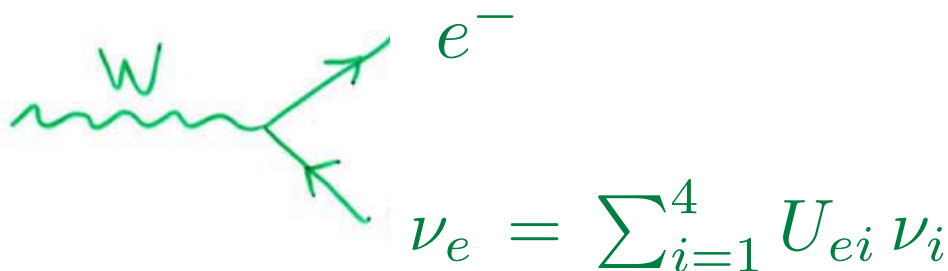
Add a sterile neutrino:

$$\nu_\alpha = \sum_{i=1}^4 U_{\alpha i} \nu_i \quad [\alpha = e, \mu, \tau]$$

$\nu_s$  flavour eigenstate  
 $\nu_4$  mass eigenstate ( $m_4$ )

$U = 4 \times 4$  unitary matrix

Only  $\nu_e, \nu_\mu, \nu_\tau$  couple to electroweak gauge boson, but all four mass eigenstate are produced in a beta decay:



## 2-flavour oscillations:

$$P(\nu_\alpha \rightarrow \nu_\beta) = \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E} \right)$$

$$\begin{pmatrix} \nu_\alpha \\ \nu_\beta \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix} \quad \Delta m^2 \equiv m_2^2 - m_1^2$$

## N-flavour oscillations:

$$P_{\nu_\alpha \rightarrow \nu_\beta} (\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta) = \delta_{\alpha\beta} - 4 \sum_{i < j} \text{Re} (U_{\alpha i} U_{\beta i}^* U_{\alpha j}^* U_{\beta j}) \sin^2 \left( \frac{\Delta m_{ij}^2 L}{4E} \right) \\ \mp 2 \sum_{i < j} \text{Im} (U_{\alpha i} U_{\beta i}^* U_{\alpha j}^* U_{\beta j}) \sin \left( \frac{\Delta m_{ij}^2 L}{2E} \right)$$

### 3+1 case:

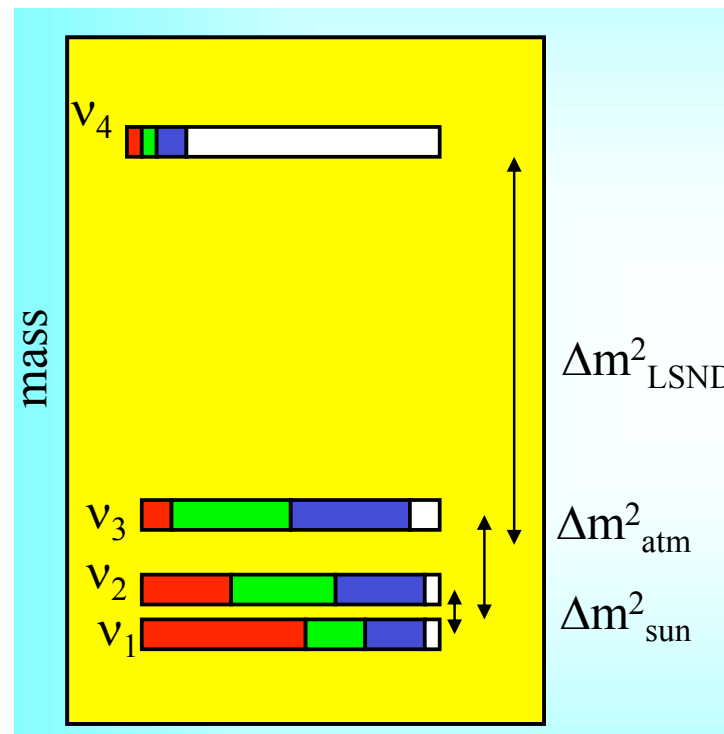
Since  $\Delta m_{SBL}^2 \gg \Delta m_{atm.}^2, \Delta m_{sun}^2$ , it is natural (and cosmologically preferred) to assume  $m_4 \gg m_3, m_2, m_1$

Then  $\Delta m_{SBL}^2 \equiv \Delta m_{41}^2 \simeq \Delta m_{42}^2 \simeq \Delta m_{43}^2 \gg$  all other  $\Delta m_{ij}^2$ 's

All data but short baseline oscillations well described by 3-flavour oscillations

$\Rightarrow \nu_{1,2,3}$  mainly composed of  $\nu_{e,\mu,\tau}$  + small admixture of  $\nu_s$ , and

$\nu_4$  mainly composed of  $\nu_s$  + small admixture of  $\nu_{e,\mu,\tau}$



Smirnov

We are interested in short baseline oscillations with

$$\frac{\Delta m_{41}^2 L}{4E} \lesssim 1 \quad \Longrightarrow \quad \sin^2 \left( \frac{\Delta m_{41}^2 L}{4E} \right) \gg \sin^2 \left( \frac{\Delta m_{31}^2 L}{4E} \right), \sin^2 \left( \frac{\Delta m_{21}^2 L}{4E} \right)$$

$$\begin{aligned} P_{\nu_\alpha \rightarrow \nu_\alpha} &\simeq 1 - 4 (|U_{\alpha 1}|^2 + |U_{\alpha 2}|^2 + |U_{\alpha 3}|^2) |U_{\alpha 4}|^2 \sin^2 \left( \frac{\Delta m_{41}^2 L}{4E} \right) \\ &\equiv 1 - \sin^2 2\theta_{\alpha\alpha} \sin^2 \left( \frac{\Delta m_{41}^2 L}{4E} \right) \end{aligned}$$

where  $\sin^2 2\theta_{\alpha\alpha} \equiv 4 (1 - |U_{\alpha 4}|^2) |U_{\alpha 4}|^2$

$$\begin{aligned} P_{\nu_\alpha \rightarrow \nu_\beta} &\simeq -4 \operatorname{Re} [(U_{\alpha 1} U_{\beta 1}^* + U_{\alpha 2} U_{\beta 2}^* + U_{\alpha 3} U_{\beta 3}^*) U_{\alpha 4}^* U_{\beta 4}] \sin^2 \left( \frac{\Delta m_{41}^2 L}{4E} \right) \\ &\equiv \sin^2 2\theta_{\alpha\beta} \sin^2 \left( \frac{\Delta m_{41}^2 L}{4E} \right) \end{aligned}$$

where  $\sin^2 2\theta_{\alpha\beta} \equiv 4 |U_{\alpha 4} U_{\beta 4}|^2$

### 3+2 case:

Assume  $m_5 \sim m_4 \gg m_3, m_2, m_1$

$\Rightarrow$  two relevant squared mass differences  $\Delta m_{51}^2$  and  $\Delta m_{41}^2$

$\Rightarrow$  CP-violating effects possible due to interference between the two oscillations frequencies

$$P_{\nu_\alpha \rightarrow \nu_\beta}(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta) = 4 |U_{\alpha 4} U_{\beta 4}|^2 \sin^2 \left( \frac{\Delta m_{41}^2 L}{4E} \right) + 4 |U_{\alpha 5} U_{\beta 5}|^2 \sin^2 \left( \frac{\Delta m_{51}^2 L}{4E} \right) \\ + 8 |U_{\alpha 4} U_{\beta 4} U_{\alpha 5} U_{\beta 5}| \sin \left( \frac{\Delta m_{41}^2 L}{4E} \right) \sin^2 \left( \frac{\Delta m_{51}^2 L}{4E} \right) \cos \left( \frac{\Delta m_{54}^2 L}{4E} \mp \eta \right)$$

$$\eta \equiv \arg [U_{\alpha 4} U_{\beta 4}^* U_{\alpha 5}^* U_{\beta 5}]$$

# Experimental situation

Several experimental anomalies suggest the existence of sterile neutrinos

LSND:  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations

Excess of  $\bar{\nu}_e$  events over background at  $3.8 \sigma$  (still controversial)

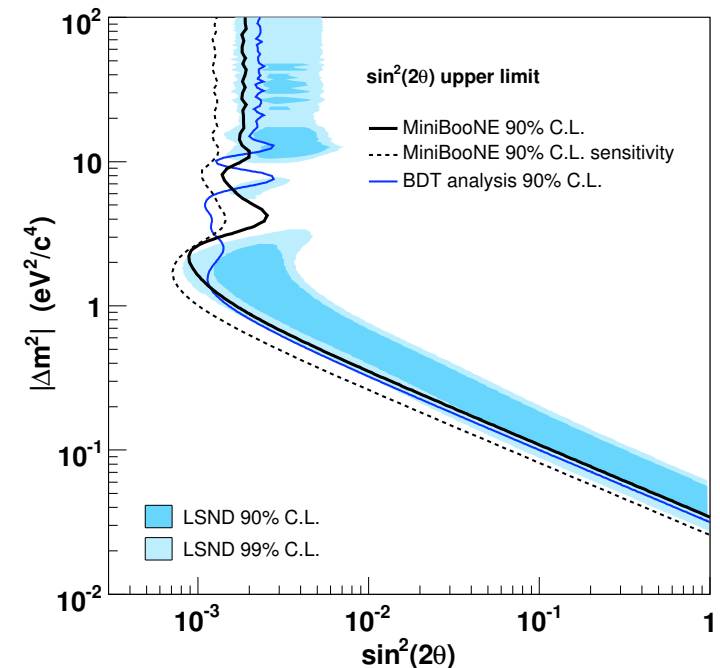
Not observed by KARMEN

MiniBooNE:

$\nu_\mu \rightarrow \nu_e$  data: no excess in the 475-1250 MeV range, but unexplained  $3\sigma$   $\nu_e$  excess at low energy

$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  data:  $\bar{\nu}_e$  excess in the  $E > 475$  MeV region consistent with LSND-like oscillations, but also (after the 2011 update) with a background-only hypothesis

A low-energy  $\bar{\nu}_e$  excess is also seen



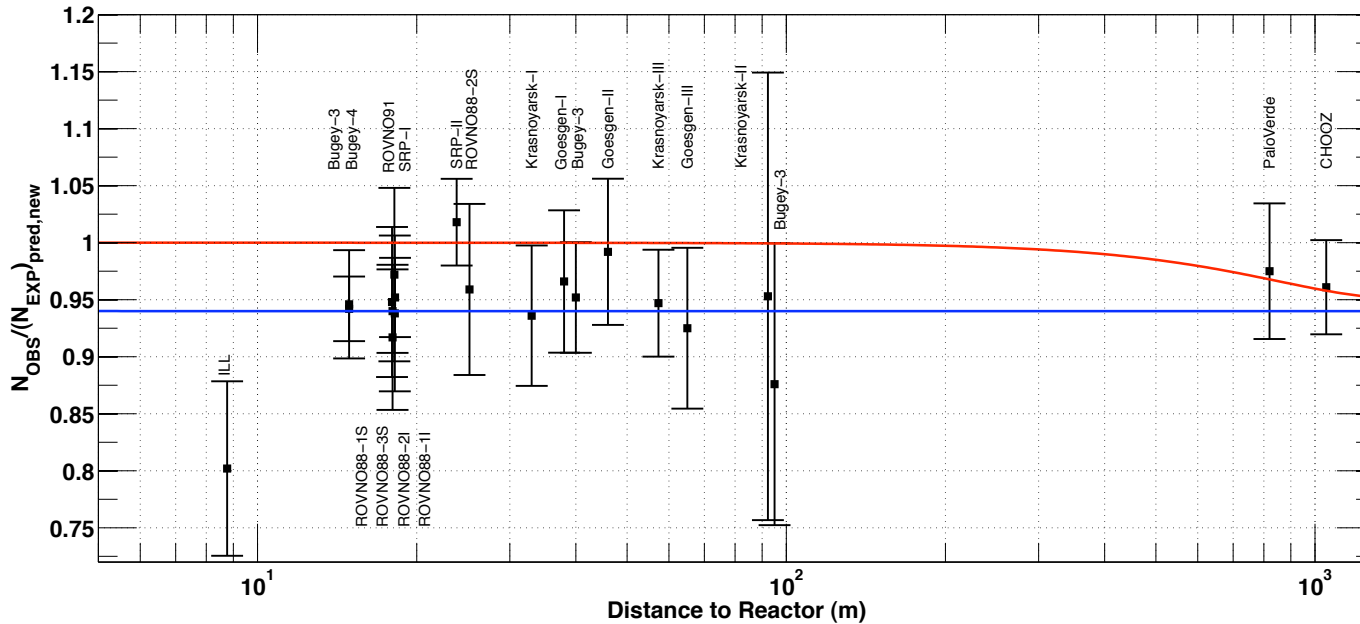


## Reactor antineutrino anomaly:

### New computation of the reactor antineutrino spectra

⇒ increase of the flux by about 3%

⇒ deficit of antineutrinos in SBL reactor experiments  
mean observed to predicted rate  $0.943 \pm 0.023$



G. Mention et al.

FIG. 5. Illustration of the short baseline reactor antineutrino anomaly. The experimental results are compared to the prediction without oscillation, taking into account the new antineutrino spectra, the corrections of the neutron mean lifetime, and the off-equilibrium effects. Published experimental errors and antineutrino spectra errors are added in quadrature. The mean averaged ratio including possible correlations is  $0.943 \pm 0.023$ . The red line shows a possible 3 active neutrino mixing solution, with  $\sin^2(2\theta_{13}) = 0.06$ . The blue line displays a solution including a new neutrino mass state, such as  $|\Delta m_{\text{new,R}}^2| \gg 1 \text{ eV}^2$  and  $\sin^2(2\theta_{\text{new,R}}) = 0.12$  (for illustration purpose only).

## Gallex-SAGE calibration experiments:

Calibration of the Gallex and SAGE experiments with radioactive sources  
⇒ observed deficit of  $\nu_e$  with respect to predictions

$$R = 0.86 \pm 0.05$$

[tension with  $\nu_e$  - Carbon cross-section measurements at LSND and KARMEN, 1106.5552]

Combined analysis of SBL reactor data, gallium calibration experiments and MiniBooNE neutrino data [G. Mention et al.]:

$$|\Delta m_{SBL}^2| > 1.5 \text{ eV}^2, \quad \sin^2 2\theta_{ee} = 0.14 \pm 0.08 \quad (95\% \text{ C.L.})$$

However, no coherent picture of the data with an additional (or even 2) sterile neutrinos (even if the global fit has improved with the new reactor antineutrino flux):

I) tension between appearance (LSND/MiniBooNE antineutrino data) and disappearance experiments (reactors,  $\nu_\mu$  disappearance experiments)

Reactors: 
$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} \simeq 1 - \sin^2 2\theta_{ee} \sin^2 \left( \frac{\Delta m_{41}^2 L}{4E} \right)$$

require relatively small  $\sin^2 2\theta_{ee} \equiv 4(1 - |U_{e4}|^2)|U_{e4}|^2 \simeq 4|U_{e4}|^2$   
(using info from solar neutrino data)

CDHS: 
$$P_{\nu_\mu \rightarrow \nu_\mu} \simeq 1 - \sin^2 2\theta_{\mu\mu} \sin^2 \left( \frac{\Delta m_{41}^2 L}{4E} \right)$$

require relatively small  $\sin^2 2\theta_{\mu\mu} \equiv 4(1 - |U_{\mu4}|^2)|U_{\mu4}|^2 \simeq 4|U_{\mu4}|^2$   
(using info from atm. neutrino data)

# Other implications of sterile neutrinos

Tritium beta decay:

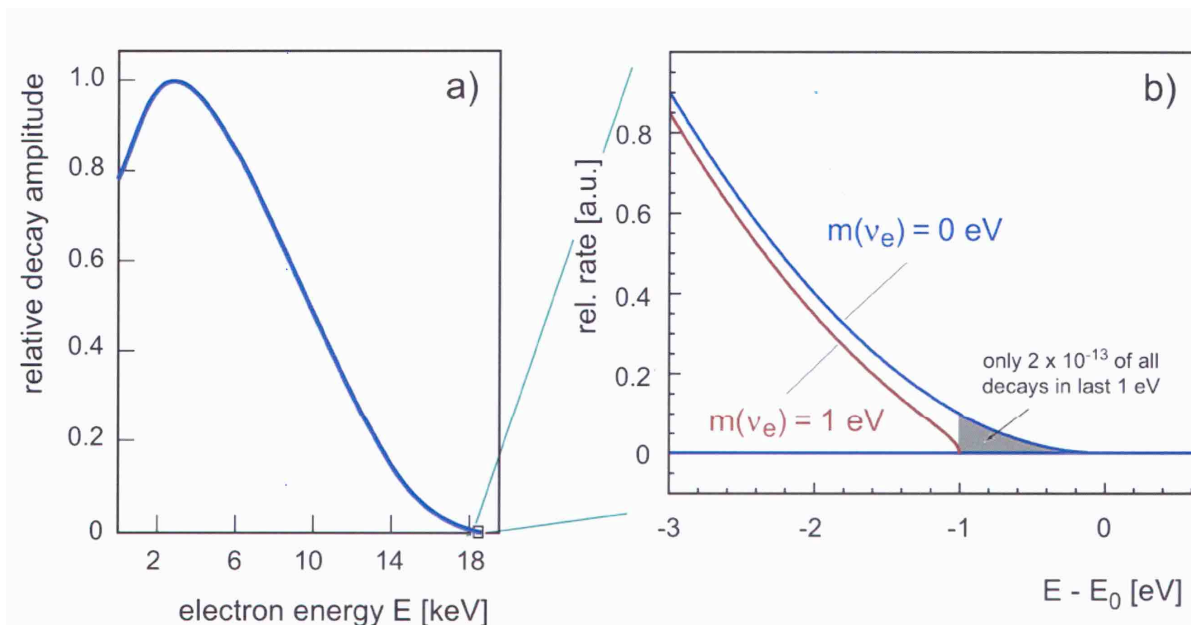


The electron energy spectrum is given by:

$$\frac{dN}{dE_e} = R(E_e) \sqrt{(E_0 - E_e)^2 - m_\nu^2} \quad E_e = E_0 - E_\nu$$

Effect of the non-vanishing neutrino mass:  $E_e^{max} = E_0 \rightarrow E_0 - m_\nu$

$\Rightarrow$  distorsion of the  $E_e$  spectrum close to the endpoint



Present bound (Troitsk/Mainz):  $m_{\nu_e} < 2.2 \text{ eV}$  (95% C.L.)

KATRIN will reach a sensitivity of about 0.3 eV

In practice, there is no electron neutrino mass, but 3 (or more) strongly mixed mass eigenstates, and

$$\frac{dN}{dE_e} = R(E_e) \sum_i |U_{ei}|^2 \sqrt{(E_0 - E_e)^2 - m_i^2} \Theta(E_0 - E_e - m_i)$$

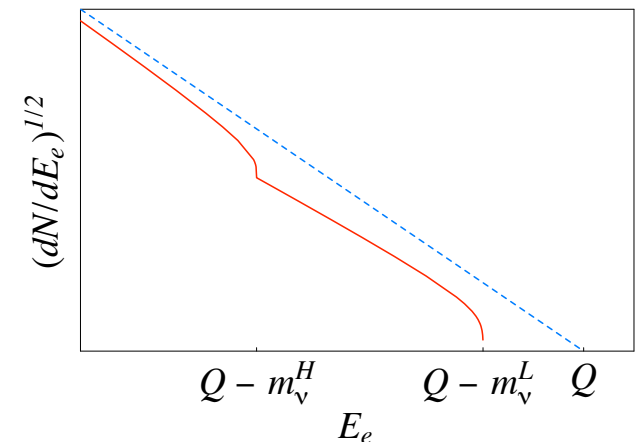
If all  $m_i$  are smaller than the energy resolution, this can be rewritten as:

$$\frac{dN}{dE_e} = R(E_e) \sqrt{(E_0 - E_e)^2 - m_\beta^2} \quad m_\beta^2 \equiv \sum_i m_i^2 |U_{ei}|^2$$

If there is an eV-scale sterile neutrino (comparable to the energy resolution of KATRIN), its mass may be resolved (but difficult measurement):

$$\begin{aligned} \frac{1}{R(E_e)} \frac{dN}{dE_e} &= (1 - |U_{e4}|^2) \sqrt{(E_0 - E_e)^2 - m_\beta^2} \\ &+ |U_{e4}|^2 \sqrt{(E_0 - E_e)^2 - m_4^2} \Theta(E_0 - E_e - m_4) \end{aligned}$$

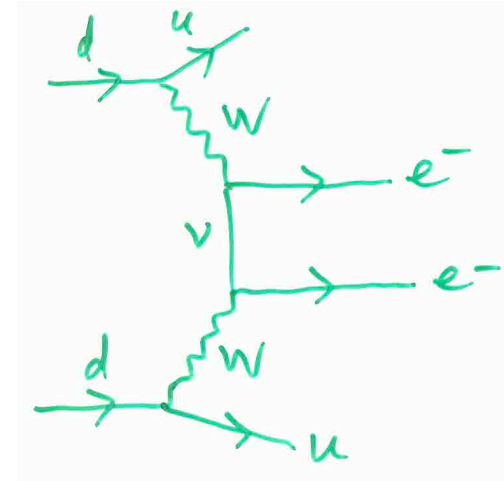
(also: upper bound on  $m_4$  from beta decay)



## Neutrinoless double beta decay:

$$(A, Z) \rightarrow (A, Z + 2) + e^- + e^-$$

Possible if lepton number violated  
(Majorana neutrinos), in nuclei where  
the single beta decay is forbidden

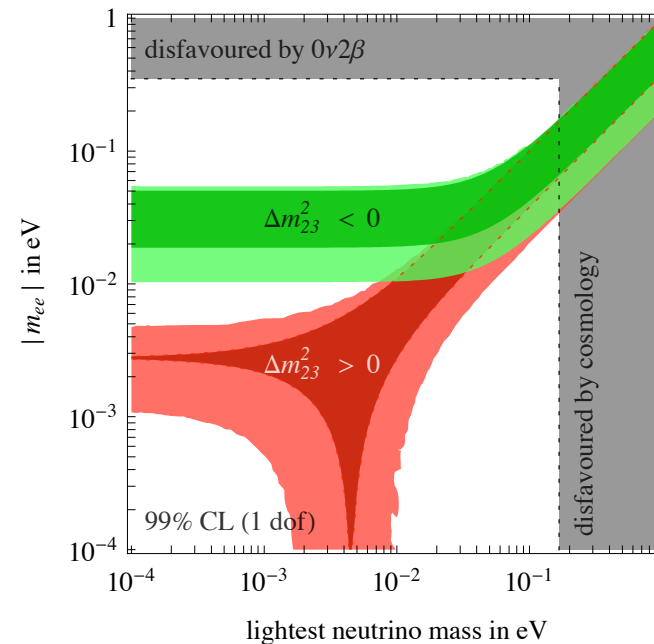


Sensitive to the effective mass parameter:

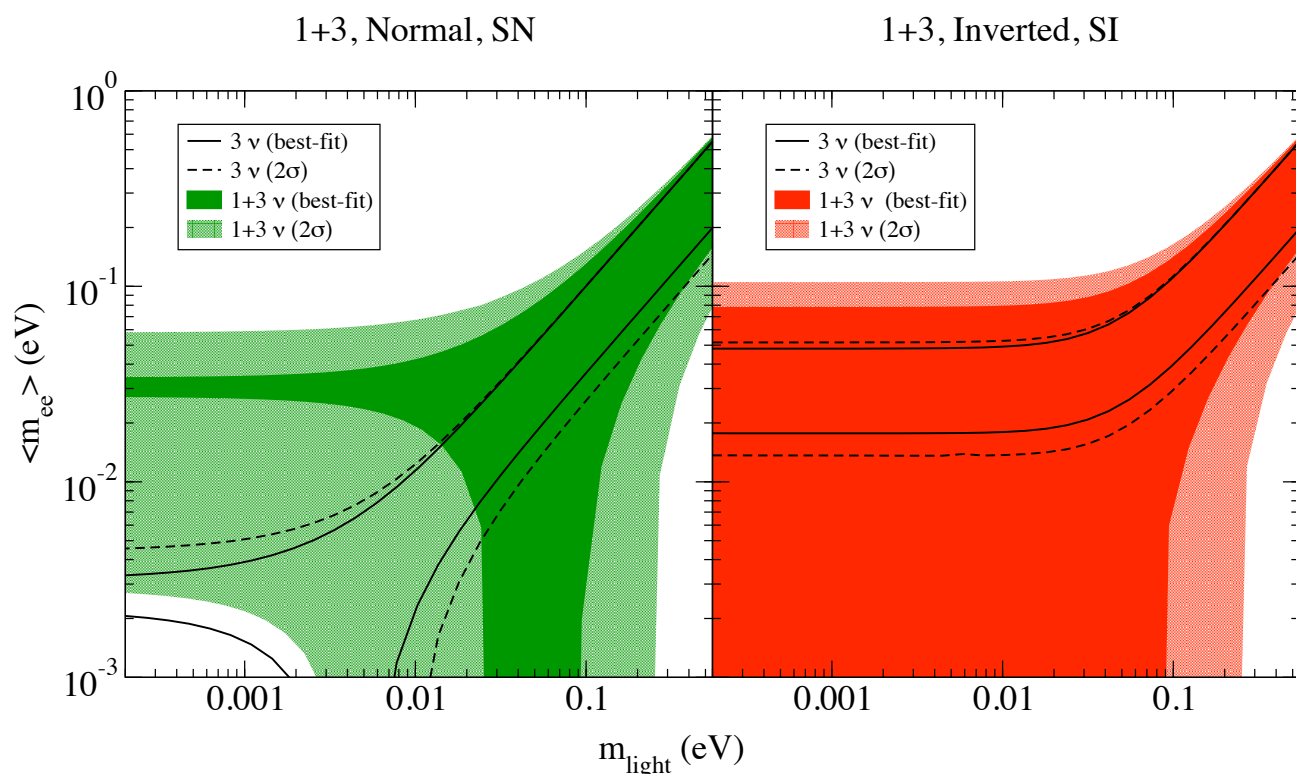
$$m_{\beta\beta} \equiv \sum_i m_i U_{ei}^2$$

possible cancellations in the sum (phases in U)

3-neutrino case  
(Strumia, Vissani)



An additional sterile neutrino will contribute  $m_4|U_{e4}|^2 e^{i\gamma}$  to the effective mass  $m_{\beta\beta} \equiv \sum_i m_i U_{ei}^2$ ; depending on the active neutrino parameters it may dominate or lead to cancellations



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using the fit of Kopp, Maltoni and Schwetz:

		parameter	$\Delta m_{41}^2$ [eV]	$ U_{e4} ^2$
3+1/1+3	best-fit		1.78	0.023
	$2\sigma$		1.61–2.01	0.006–0.040