

# Hierarchy and Octant Determination Potential of LBNE and LBNO

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July 18, 2013



# Neutrino Oscillation measurements

- From KamLAND:  $\Delta_{21} = 7.5 \times 10^{-5} \text{ eV}^2$ .
- From MINOS:  $\Delta m_{eff}^2 = \pm 2.4 \times 10^{-3} \text{ eV}^2$ .
- From Solar neutrino data:  $\sin^2 \theta_{12} = 0.3$ .

$\Delta m_{eff}^2$  is related to  $\Delta_{31}$  through the formula

$$\Delta m_{eff}^2 = \Delta m_{31}^2 - \Delta m_{21}^2 (\cos^2 \theta_{12} - \cos \delta_{CP} \sin \theta_{13} \sin 2\theta_{12} \tan \theta_{23}).$$

## Two exciting results from the past year

- From Daya Bay and RENO:  $\sin^2 2\theta_{13} = 0.089 \pm 0.011$  confirming the earlier result of T2K, MINOS and DoubleCHOOZ that  $\theta_{13} \neq 0$ .  
Good for future measurements such as hierarchy and CP violation.
- From MINOS:  $\sin^2 \theta_{23} = 0.94 \pm$ , indicating that  $\theta_{23}$  is not maximal.  
Leading to two degenerate solutions  $\sin^2 \theta_{23} = 0.41$  in lower octant (LO) and  $\sin^2 \theta_{23} = 0.59$  in higher octant (HO).

# Hierarchy Determination at upcoming experiments

Long baseline experiments, such as NO $\nu$ A and T2K, aim to determine hierarchy and octant through measuring  $\nu_\mu \rightarrow \nu_e$  and  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations.

$\nu_\mu \rightarrow \nu_e$  oscillation probability, expanded perturbatively in  $\alpha = \Delta_{21}/\Delta_{31}$ :

$$\begin{aligned} P_{\mu e} = & \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2 \hat{\Delta}(1 - \hat{A})}{(1 - \hat{A})^2} \\ & + \alpha \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23} \cos(\hat{\Delta} + \delta_{CP}) \frac{\sin \hat{\Delta} \hat{A}}{\hat{A}} \frac{\sin \hat{\Delta}(1 - \hat{A})}{1 - \hat{A}} \\ & + \alpha^2 \sin^2 2\theta_{12} \cos^2 \theta_{13} \cos^2 \theta_{23} \frac{\sin^2 \hat{\Delta} \hat{A}}{\hat{A}^2} \end{aligned}$$

where  $\hat{\Delta} = \Delta_{31}L/4E$  and  $\hat{A} = A/\Delta_{31}$ .

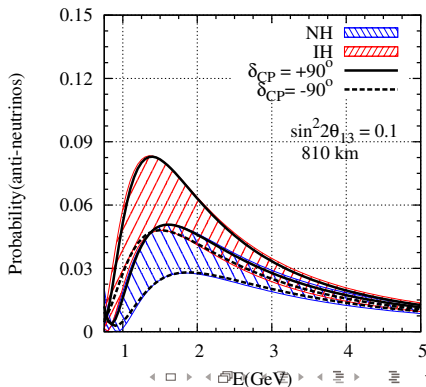
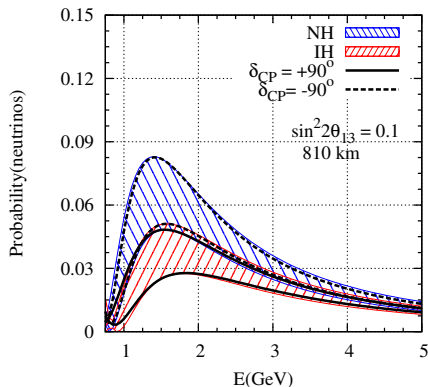
Wolfenstein matter term:  $A(\text{eV}^2) = 0.76 \times 10^{-4} \rho(\text{gm/cc})E(\text{GeV})$ .

# Hierarchy- $\delta_{CP}$ degeneracy

$\hat{A}$  changes sign if the hierarchy is changed. So  $P_{\mu e}$  is sensitive to hierarchy. But,  $P_{\mu e}$  also depends on  $\delta_{CP}$ , of which we know nothing. Thus we are led to hierarchy- $\delta_{CP}$  degeneracy.

$P_{\mu e}$  (as a band in  $\delta_{CP}$ ) vs. Energy for NO $\nu$ A

Left panel: Neutrinos & Right panel: Anti-neutrinos



# Partial resolution of hierarchy- $\delta_{CP}$ degeneracy

Matter term increases  $P_{\mu e}$  for NH and decreases it for IH. Its effect on  $P_{\bar{\mu} \bar{e}}$  is the opposite.

Hierarchy- $\delta_{CP}$  arises because the change due to matter term can be cancelled by taking a wrong value of  $\delta_{CP}$ .

Have to do a careful study of  $\delta_{CP}$  dependence of  $P_{\mu e}$  and  $P_{\bar{\mu} \bar{e}}$  to resolve this.

In  $P_{\mu e}$  plots, the lines for  $\delta_{CP} = \pm 90^\circ$  are emphasized. The line for  $\delta_{CP} = -90^\circ (+90^\circ)$  is close to the top (bottom) for neutrinos and vice-verse for anti-neutrinos.

The lines for  $\delta_{CP} = 0$  and  $180^\circ$  are somewhere in the middle.

# Favourable and Unfavourable Half-planes

$P_{\mu e}$  values for NH and  $\delta_{CP}$  in lower half plane ( $-180^\circ \leq \delta_{CP} \leq 0$ ) are well above the values of  $P_{\mu e}$  for IH and any  $\delta_{CP}$ . Similarly  $P_{\mu e}$  values for IH and  $\delta_{CP}$  in upper half plane ( $0 \leq \delta_{CP} \leq 180^\circ$ ) are well below the values of  $P_{\mu e}$  for NH and any  $\delta_{CP}$ .

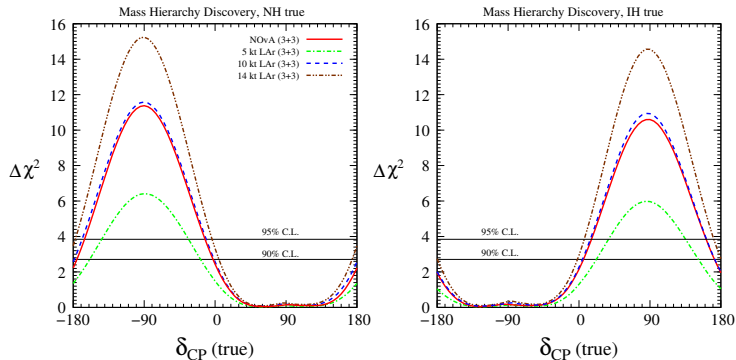
The same clear separation is there in  $P_{\bar{\mu}e}$  values also.

If nature chooses one of the two favourable combinations (NH with  $\delta_{CP}$  in LHP or IH with  $\delta_{CP}$  in UHP) then  $\text{NO}\nu\text{A}$ , by itself, can determine the hierarchy at  $2\sigma$  to  $3\sigma$  depending on  $\delta_{CP}$ .

For unfavourable combinations, NH with  $\delta_{CP}$  in UHP and IH with  $\delta_{CP}$  in LHP,  $\text{NO}\nu\text{A}$  has **NO** sensitivity to hierarchy. This occurs because the data gives two degenerate solutions: One with correct hierarchy and correct  $\delta_{CP}$  and one with wrong hierarchy and wrong  $\delta_{CP}$ .

S. Prakash, S. Raut and SUS, Phys. Rev. D86 (2012) 033012.

# Hierarchy sensitivity of NO $\nu$ A



From: S. Agarwalla, S. Prakash, S. Raut and SUS, JHEP 1212 (2012) 075.

# Does addition of T2K data help?

T2K has a shorter baseline and smaller energy compared to  $\text{NO}\nu\text{A}$  with only one third the matter effect.

Since the matter effects at T2K are small, can it determine the half plane of  $\delta_{CP}$  independent of hierarchy?

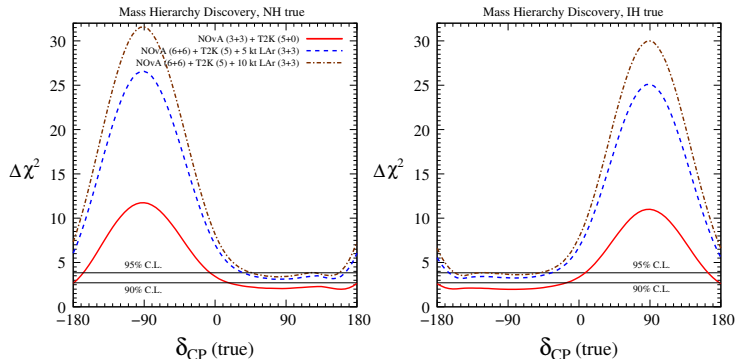
**NO.** Now matter how small the matter effect, in the unfavourable regions, we get degenerate fake hierarchy solutions.

But T2K data will help a little bit. The degenerate solutions with wrong hierarchy, occur at different values of  $\delta_{CP}$  for  $\text{NO}\nu\text{A}$  and for T2K. So the combined data has some moderate hierarchy sensitivity.

(O. Mena and S. J. Parke, hep-ph/040807, PRD70 (2004) 093011.)



# Mass hierarchy sensitivity with $\text{NO}\nu\text{A}$ and T2K

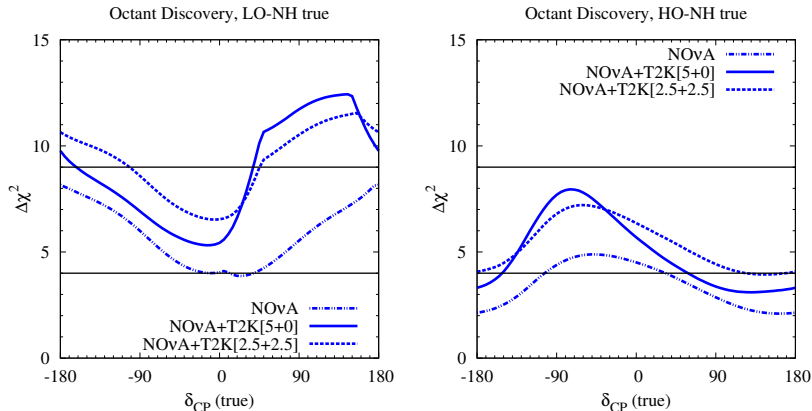


From: S. Agarwalla, S. Prakash, S. Raut and SUS, JHEP 1212 (2012) 075.  
We have something close to 90% C. L. hierarchy discrimination for unfavourable combinations.

# $\theta_{23}$ -octant resolution with $\text{NO}\nu\text{A}$ and T2K

## Octant discovery plots for reoptimized $\text{NO}\nu\text{A}$ and T2K (modified)

Left panel: LO true & Right panel: HO true



From: S. Agarwalla, S. Prakash and SUS, arXiv:1301.2574 (in JHEP)

See also the poster by Suprabh Prakash

# Need for future facilities

- Current experiments have less than 90% C.L. hierarchy sensitivity if the hierarchy- $\delta_{CP}$  combination is unfavourable.
- They have just about 95% C.L. for  $\theta_{23}$  octant resolution.
- Their ability to rule out  $\delta_{CP} \neq 0$  or  $180^\circ$  is also quite limited (can do so at 90% C.L. for a range of about  $\pm 30^\circ$  around  $\delta_{CP} = \pm 90^\circ$ )  
S. Agarwalla, S. Prakash, S. Raut and SUS, JHEP 1212 (2012) 075.

Hence there is a need to consider future facilities, which can make robust measurements of the above three quantities.

Hierarchy and octant of  $\theta_{23}$  resolutions should be considered the first steps of these future facilities. CP non-conservation and measurement of  $\delta_{CP}$  are more difficult measurements and can be considered at later stages.

# Physics constraints on future facilities

- Matter effects at these facilities should be large enough so that  $P_{\mu e}$  and  $P_{\bar{\mu} \bar{e}}$  are well separated for NH and for IH, *even for  $\theta_{23}$  in lower octant.*
- Since the matter effect is proportional to energy, the  $\nu$  ( $\bar{\nu}$ ) beam energy must be reasonably large.
- Since the oscillation probabilities depend on  $L/E$ , the baseline should be correspondingly larger.
- As the flux falls off at longer distances, there must be high intensity sources.
- The detectors should have very good particle identification so that the neutrino events can be kinematically reconstructed. This is needed to suppress the neutral current single  $\pi^0$  background to the electron appearance events.

# LBNE and LBNO

Characteristic	LBNE	LBNO
Baseline	1300 km	2300 km
Location	Fermilab- Homestake	CERN-Pyhasalmi
Proton energy	120 GeV	400 GeV
Beam	on - axis	wideband
Beam power	0.7 MW	0.75 MW
Protons of Target	$6 \times 10^{20}$ per year	$1.5 \times 10^{20}$ per year
Detector	LArTPC (10 kton)	LArTPC (20 kton)
Runtime	5 in $\nu$ + 5 in $\bar{\nu}$	5 in $\nu$ + 5 in $\bar{\nu}$

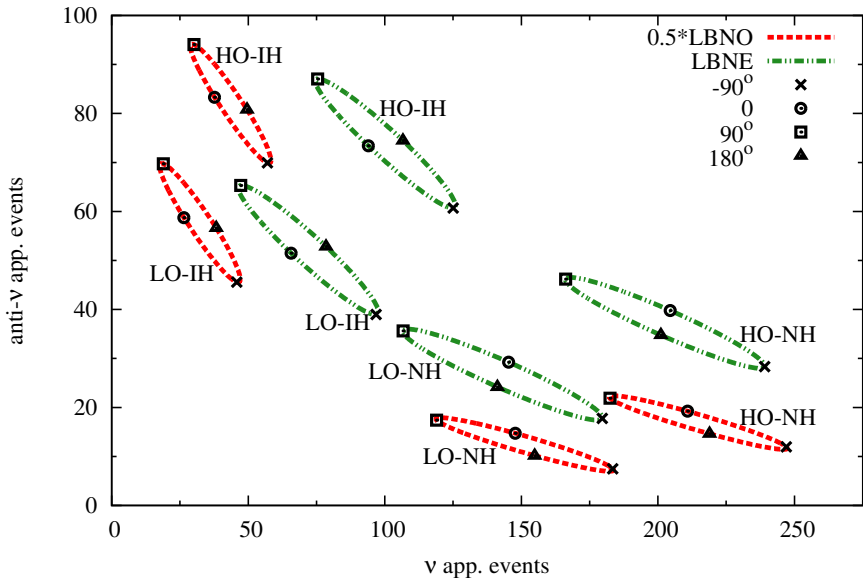
# Simulation Details

- We have used best fit neutrino parameter values quoted in slide-1 and the detector properties quoted in slide-13.
- We have taken  $\sigma(|\Delta m_{eff}^2|) = 4\%$ ,  $\sigma(\sin^2 2\theta_{13}) = 5\%$  and  $\sigma(\text{density}) = 5\%$ .
- In computing  $\Delta\chi^2$ , we marginalized over the above parameters in their  $2\sigma$  ranges and  $\sin^2 \theta_{23}$  in the range  $0.36 - 0.67$  ( $3\sigma$  range from the global best fits).
- We have computed the hierarchy and the octant determination capability of LBNE, LBNO and also  $0.5*LBNO$  (a hypothetical case where the detector mass is the same as LBNE).

The full details are given in

S. Agarwalla, S. Prakash and SUS, arXiv:1304.3251

# Electron appearance events for 0.5\*LBNO and LBNE

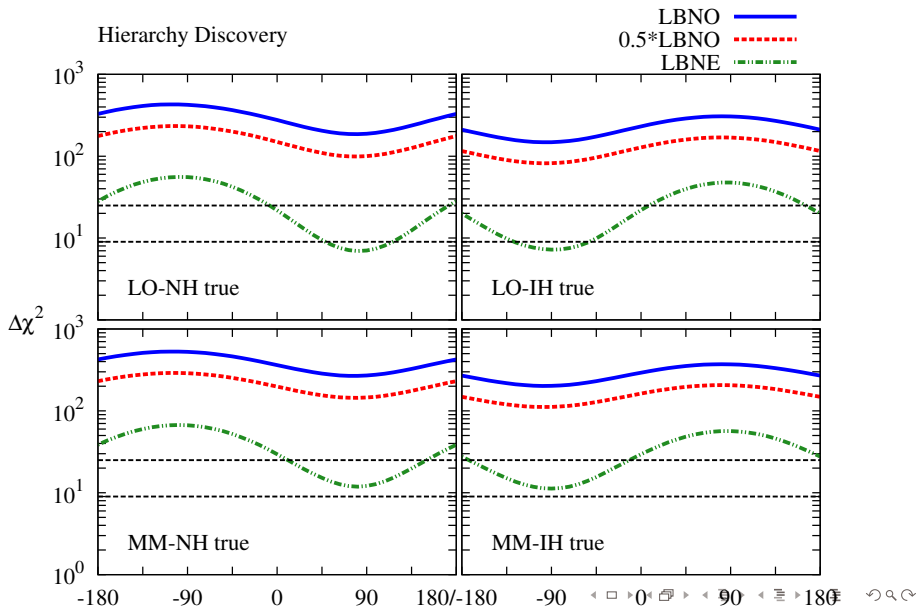


# Physics from event contour plots

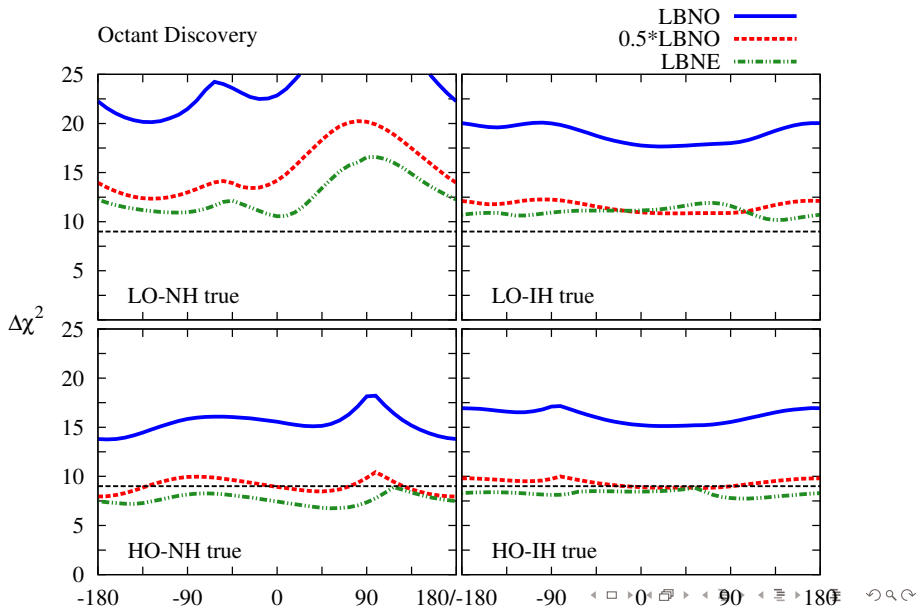
- We plot the  $\nu_e$  appearance events vs  $\bar{\nu}_e$  appearance events for LBNE and 0.5\*LBNO. Varying  $\delta_{CP}$  in its full range leads to the ellipses.
- The ability of an experiment to distinguish hierarchy/octant depends on the **separation between the ellipses** for different hierarchies and different octants.
- 0.5\*LBNO has excellent hierarchy determination capability, just from its neutrino run. This is because it is close to being the **bimagic** baseline.  
S. Raut, R. S. Singh and SUS, Phys. Lett. B696 (2011) 227 and  
A. Dighe, S. Goswami and S. Ray, Phys. Rev. Lett. 105 (2010) 261802.
- Anti-neutrino run is required to obtain octant discrimination.
- If the true octant is LO, then the hierarchy discrimination of LBNE is limited, especially for the most unfavourable values of  $\delta_{CP}$ .



# Hierarchy sensitivity of LBNE and LBNO



# Octant sensitivity of LBNE and LBNO



# Summary

- LBNE and LBNO are two planned future very long baseline facilities to explore the next stage of neutrino oscillation measurements.
- Physics goals are hierarchy determination (should  $\text{NO}\nu\text{A}$  find the hierarchy- $\delta_{CP}$  combination to be unfavourable),  $\theta_{23}$  octant determination and search for CP violation.
- LBNO has excellent hierarchy determination capability for all octant-hierarchy combinations, whereas that LBNE, as currently planned, is limited if the true octant is LO and  $\delta_{CP}$  is most unfavourable.
- Octant determination capabilities of both experiments are similar.

# Thank You