### A scaling law of elliptic flow in $\sqrt{S_{NN}} = 7.7 - 39$ GeV Au+Au collisions

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### Content

- Introduction
- Perfect fluid hydrodinamic picture
- Scaling law of  $v_2$  predicted by the Buda-Lund hydro model
- Experimental  $v_2$  data of RHIC-STAR
- Results of model fits to  $v_2$  data (preliminary)
- Summary



### Introduction - Motivation

The Relativistic Heavy Ion Collider (RHIC) showed significant second harmonic moments in transverse momentum distributions (known as elliptic flow).

The elliptic flow shown dependence on particle type, transverse ( $p_t$ ) and logitudinal (*pseodo-rapidity*) parameters, and on the centrality and bombarding energy of the collisions.

In the soft region ( $p_t < 2 \text{ GeV/c}$ ) the measuremet are well described by hydrodynamic model.

(See the Buda-Lund hydro model descriptions, for example, in: arXiv/nucl-th/0512078v4)



## Perfect fluid hydrodynamic picture

It is based on local conservation of entropy  $\sigma$  and four - momentum tensor  $T^{\mu\nu}$  :

$$\partial_{\mu}(\sigma u^{\mu}) = 0,$$
  
$$\partial_{\nu}T^{\mu\nu} = 0,$$

 $u^{\mu}$  stands fo the four-velocity of the matter and it is perfect fluid if the tensor is diagonal:

$$T^{\mu\nu} = (\epsilon + p)u^{\mu}u^{\nu} - pg^{\mu\nu}.$$



## Perfect fluid hydrodynamic picture

The elliptic flow defined as the azimuthal anisotropy or the fourier-coefficient of the one particle momentum distributions  $N_1(p)$ :

$$v_n = \frac{\int_0^{2\pi} N_1(p) \cos(n\varphi) d\varphi}{\int_0^{2\pi} N_1(p) d\varphi},$$



# Scaling law of $v_2$ by Buda-Lund hydro model

The result comes directly from a perfect hydro solution:

$$v_2 = \frac{I_1(w)}{I_0(w)},$$

#### See:

- T. Csörgő et al., Phys. Rev. C67, 034904 (2003).
- T. Csörgő, F. Grassi, Y. Hama, and T. Kodama, Phys. Lett. **B565**, 107 (2003). M. I. Nagy, T. Csörgő, and M. Csanád, Phys. Rev. **C77**, 024908 (2008).
- M. I. Nagy, T. Csörgő, and M. Csanád, Phys. Rev. C77, 024908 (2008).
- Y. M. Sinyukov and I. A. Karpenko, Acta Phys. Hung. A25, 141 (2006).
- T. Csörgő and B. Lörstad, Phys. Rev. C54, 1390 (1996).
- M. Csanád, T. Csörgő, and B. Lörstad, Nucl. Phys. A742, 80 (2004).



# Scaling law of $v_2$ by Buda-Lund hydro model

For the determination of *w* the Buda-Lund model gives the following formula:

$$w = \frac{E_K}{T_*} \varepsilon \qquad \qquad E_K = \frac{p_t^2}{2\overline{m}_t},$$
  
$$\frac{1}{T_*} = \frac{1}{2} \left( \frac{1}{T_x} + \frac{1}{T_y} \right), \qquad T_x = T_0 + \overline{m}_t \dot{X}^2 \frac{T_0}{T_0 + \overline{m}_t a^2},$$
  
$$\varepsilon = \frac{T_x - T_y}{T_x + T_y}. \qquad \qquad T_y = T_0 + \overline{m}_t \dot{Y}^2 \frac{T_0}{T_0 + \overline{m}_t a^2}.$$



### v<sub>2</sub> data

All STAR v<sub>2</sub> data from publication: *arXiv:1206.5528v1 [nucl-ex]* 

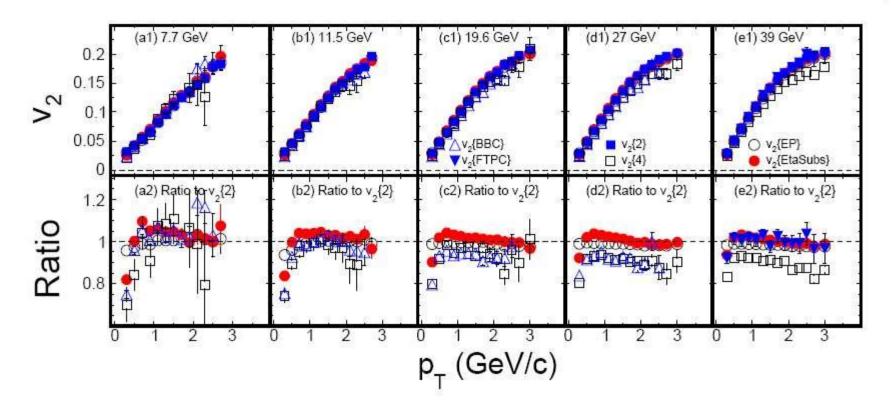


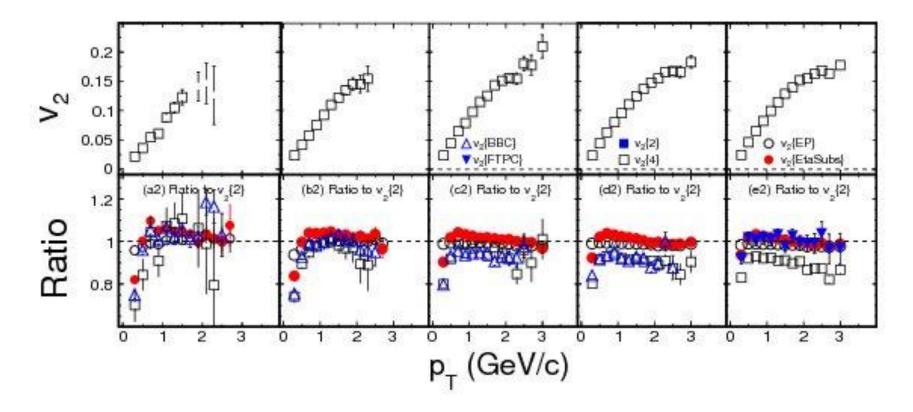
FIG. 5: (Color online) The  $v_2$  as a function of  $p_T$  for 20 - 30% central Au + Au collisions at midrapidity for  $\sqrt{s_{NN}} = 7.7$  GeV (a1), 11.5 GeV (b1), 19.6 GeV (c1), 27 GeV (d1) and 39 GeV (e1). The top panels show  $v_2$  vs.  $p_T$  using various methods as labeled in the figure and discussed in the text. The bottom panels show the ratio of  $v_2$  measured using the various methods with respect to  $v_2\{2\}$ .

8



 $v_2$  data

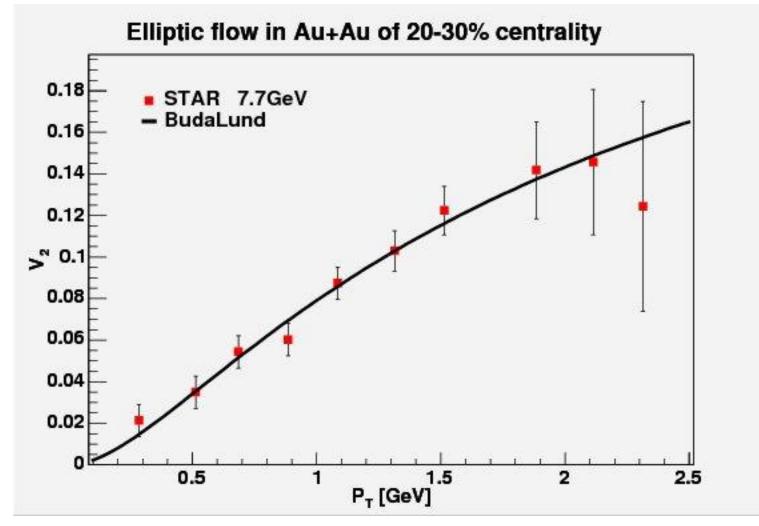
#### The selected $v_2{4}$ data for fits:



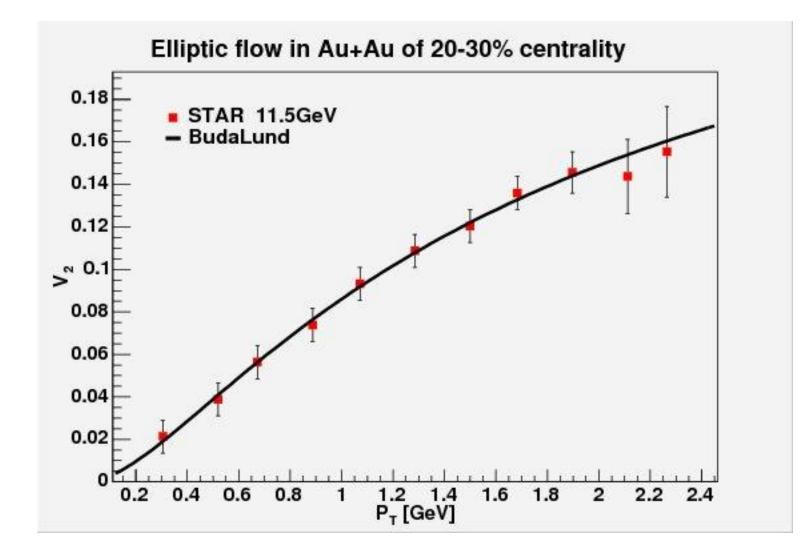
Selection based also on study of  $v_2{4}$  in our pub: *Eur. Phys. J. A (2011) 47: 58* 



#### With the assumptions: m=139MeV, $T_0$ =139MeV, $a^2$ =0.0

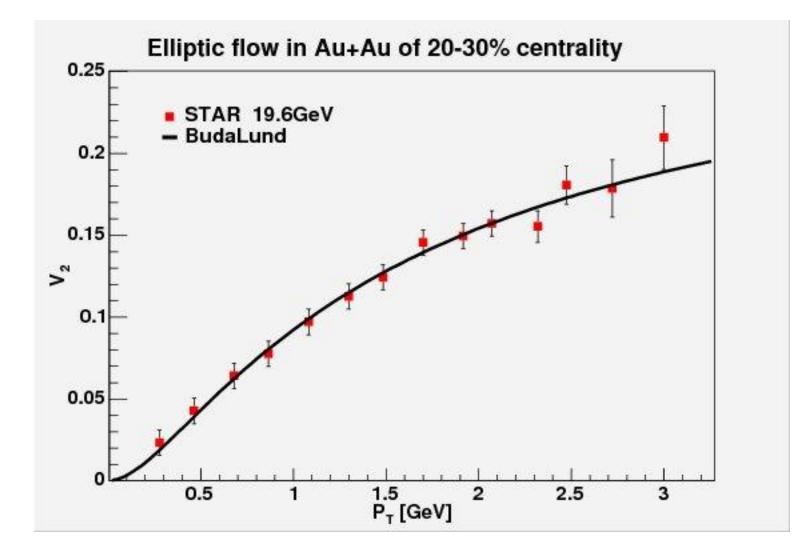




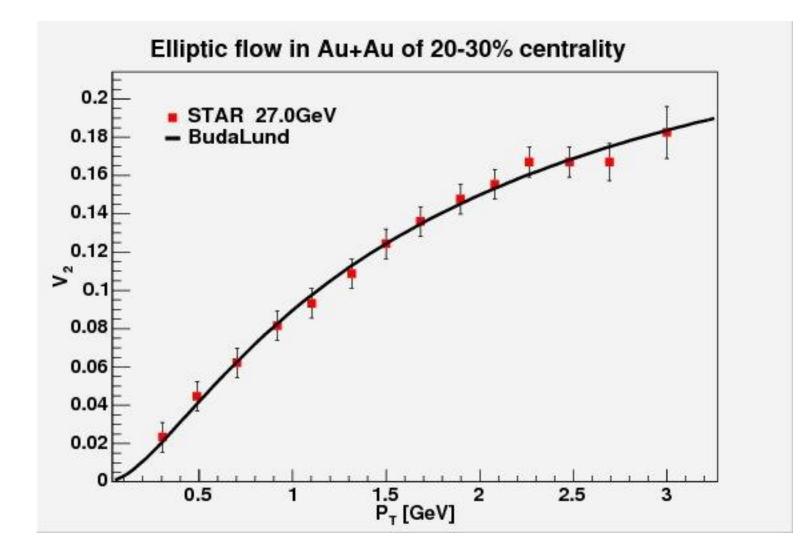


11

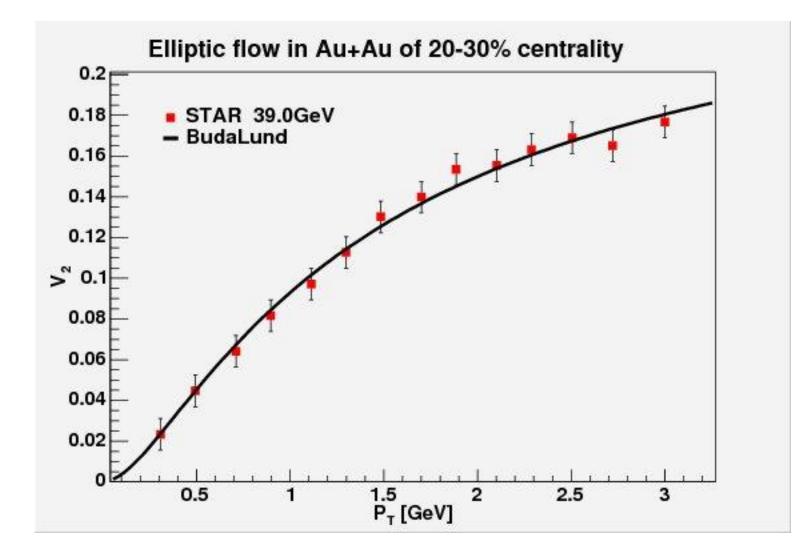




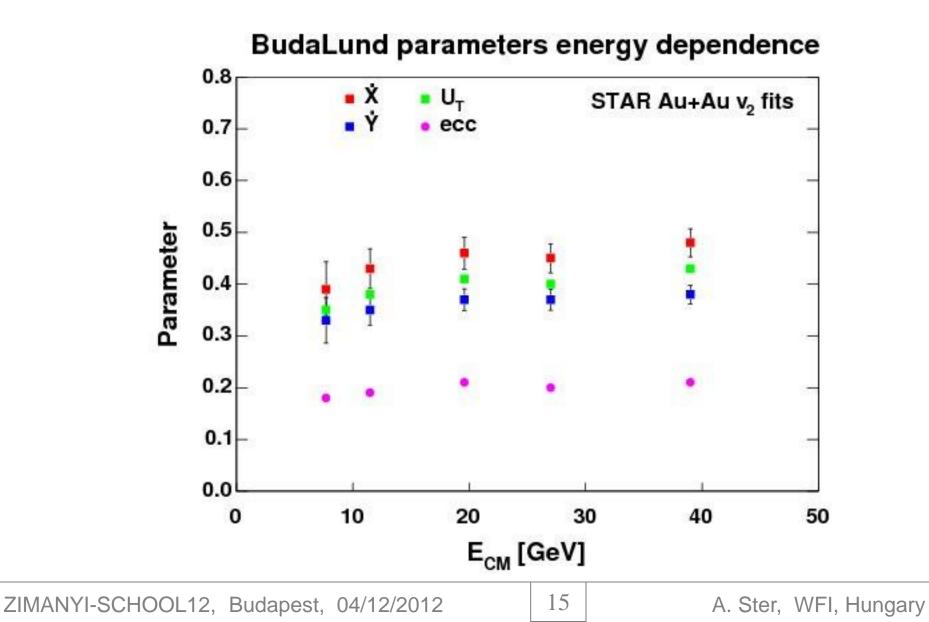




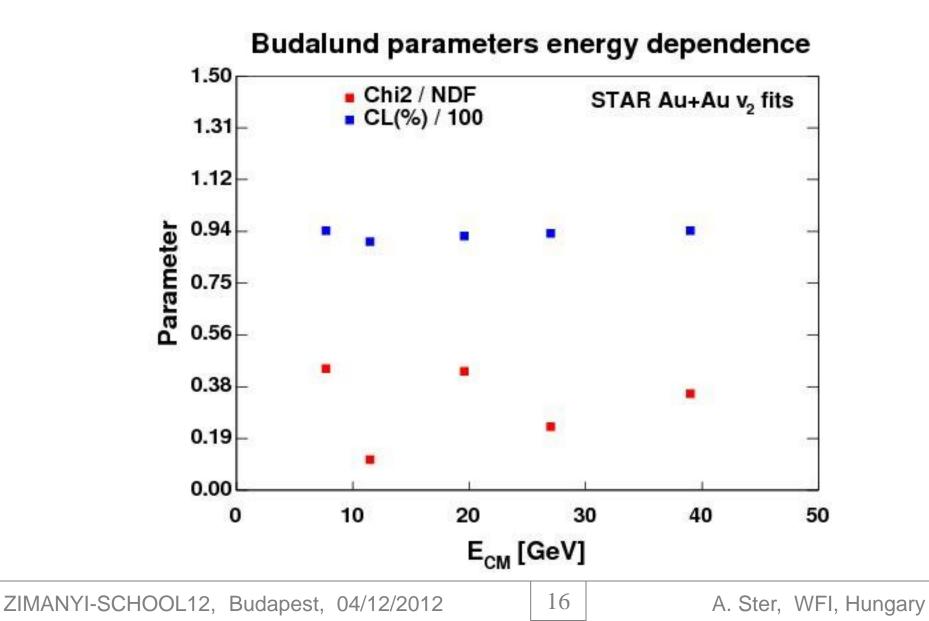




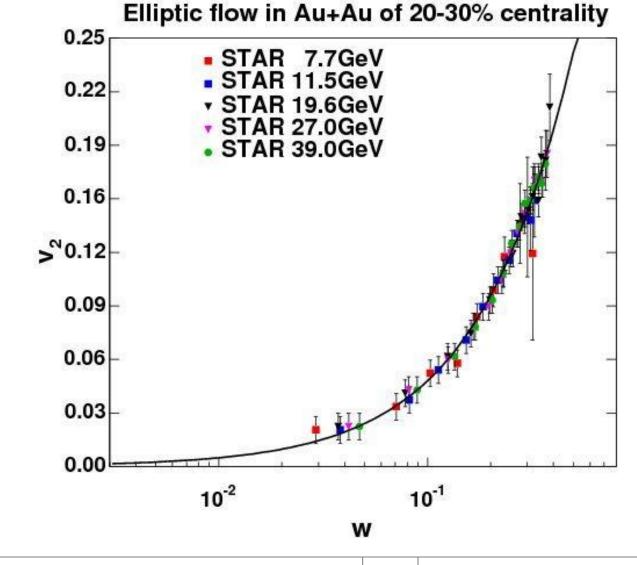














Summary table of the excitation function of the Hubble expansion rates in plane and out-of-plane, as well as, average radial flows, eccentricities and fit quality parameters:

$\sqrt{s_{NN}} ~({ m GeV})$	X	$\dot{Y}$	$\langle u_{\mathrm{T}} \rangle$ (calculated)	$\varepsilon$ (calculated)	$\chi^2/NDF$ (CL %)
7.7	$0.40\pm0.05$	$0.33\pm0.04$	$0.36\pm0.04$	$0.18 \pm 0.02$	3.09/7 (94)
11.5	$0.43\pm0.04$	$0.35\pm0.03$	$0.38\pm0.04$	$0.19\pm0.02$	0.94/8 (93)
19.6	$0.45\pm0.03$	$0.37\pm0.02$	$0.41\pm0.03$	$0.21\pm0.02$	4.78/11 (94)
27.0	$0.45\pm0.03$	$0.37\pm0.02$	$0.41\pm0.03$	$0.20\pm0.02$	2.55/11 (94)
39.0	$0.48 \pm 0.03$	$0.39\pm0.02$	$0.43\pm0.03$	$0.21\pm0.02$	$3.94/11 \ (94)$



## Summary

The Buda-Lund hydrodinamic model succesfully described STAR Au+Au  $v_2$  data with exact perfect fluid solutions.

We found that the elleptic flows consistent with a scaling law  $v_2 = I_1(w)/I_0(w)$  predicted by the model.

Our results show that the radial flow increases gradually but slightly in the studied collision energy range 7.7 - 39 GeV.

