A solvable toy model for QED at finite T

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The Bloch-Nordsieck Model

- Introduction
- The B-N solution

Perturbation Theory

One-loop correction

Resummations

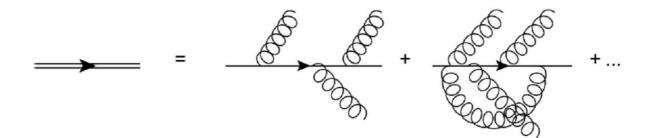
- 2PI resummation
- Exact resummation
- Finite temperature

Conclusion, Outlook

The Bloch-Nordsieck Model

Photon Mass= 0

 Long range interactions →No free charged fermion propagator, "photon cloud, quasi particle"



 From any (finite) EM energy infinite number of photons can be created

$$N = \lim_{\nu \to 0} \frac{I(\nu)}{h\nu} = \infty$$
 since $I(0) \neq 0$

We need to sum over all possible photon contribution! (even $\nu \rightarrow 0$ when!)

The Bloch-Nordsieck Model

B-N solution

- An exact solution by path integral
- Hard to generalize to more complex theory

Perturbation Theory

- 1-loop correction
- Breaks down in IR

Resummations

- 2PI, Schwinger-Dyson
- Works in IR regime

A method to treat infrared physics (?)

The B-N Model

QED $\mathcal{L} = \bar{\psi} \left(i \partial \!\!\!/ - m - e A \right) \psi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu}$ $egin{aligned} & F^{\mu u} = \partial^{\mu}A^{ u} - \partial^{ u}A^{\mu} \ & otin & ot$ *m*: fermion mass, *e*: coupling constant **Bloch-Nordsieck** NO spinflips NO positrons $\mathcal{L} = \psi^{\dagger} (i u^{\mu} \partial_{\mu} - m - e u^{\mu} A_{\mu}) \psi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu}$ $u^{\mu} \in M^4$ and $u_{\mu}u^{\mu} = 1$

BN = scalar field theory

Introduction

The B-N Model

Introduction

Free theory

$$(iu^{\mu}\partial_{\mu} - m)G_0(x - y) = \delta(x - y)$$
 (EOM)

The propagator in momentum space:

(retarded)

$$\hat{G}_0(p) = \frac{1}{u_\mu p^\mu - m + i\epsilon} \iff G_0(x - y) = 0 \text{ if } x^0 < y^0$$

Interacting case

$$[u^{\mu}(i\partial_{\mu} + eA_{\mu}(x)) - m]G(x, y|A) = \delta(x - y)$$
(EOM)
$$G(x, y) = \frac{\int G(x, y|A) \langle T \exp\left\{ie \int \bar{\psi}(z)A(z)\psi(z)dz\right\} \rangle_{F_0} \mathcal{D}A}{\int \langle T \exp\left\{ie \int \bar{\psi}(z)A(z)\psi(z)dz\right\} \rangle_{F_0} \mathcal{D}A}$$

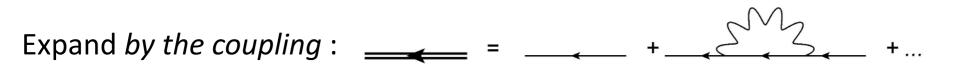
$$\hat{G}(up) = \frac{1}{(up-m)^{1+\gamma}} = \frac{1}{up-m} e^{-\gamma \ln(up-m)}$$

No antifermions

$$\gamma = rac{e^2(3-\xi)}{8\pi^2} = rac{lpha(3-\xi)}{2\pi}$$

 $\boldsymbol{\xi}$ is gauge fixing parameter

1-loop



Divergent loop-integral (in Feynman gauge, ξ =1): k

$$\underbrace{\int \int \frac{d^4k}{(2\pi)^4} u^{\mu} \frac{i}{(p^{\alpha} - k^{\alpha})u_{\alpha} - m} u^{\nu} \frac{-ig_{\mu\nu}}{k^2} = -i\Sigma_{1/2}$$

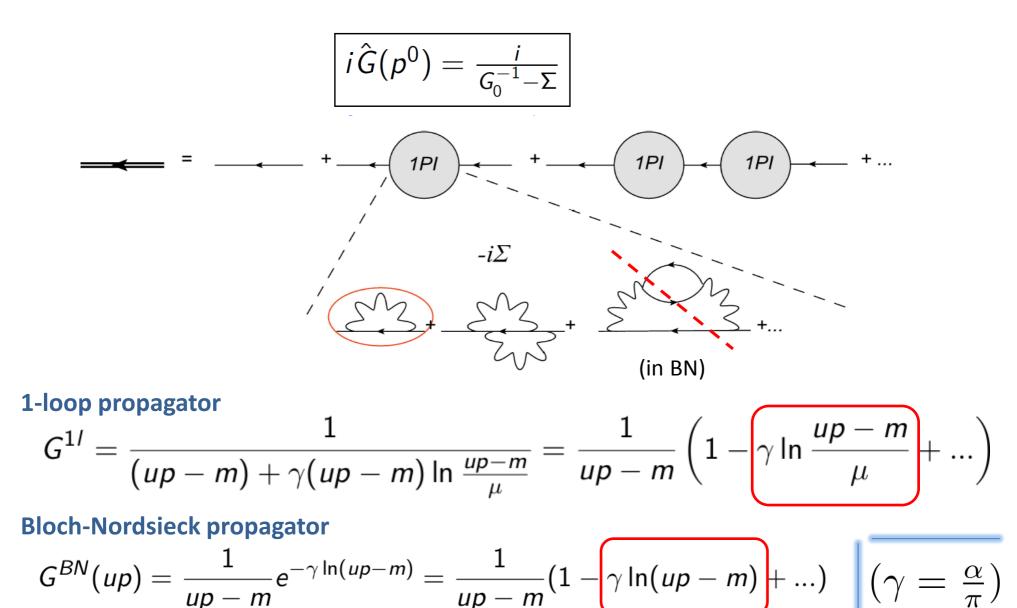
- The self-energy of the fermion
- Dimensional regularization
- Special frame *u*=(1,0,0,0)

(+ UV RENORM.)

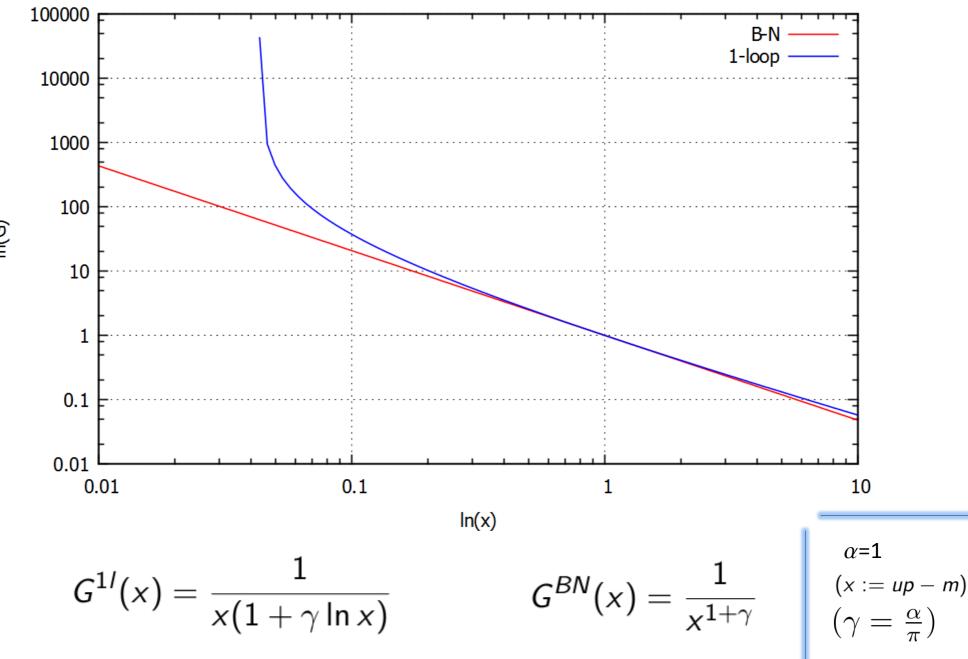
$$\Sigma_{1\prime}^{ren} = \Sigma_{1\prime} - \Sigma_{ct} = -rac{lpha}{\pi} (p^0 - m) \ln rac{m - p^0}{\mu}$$
 ($< \infty$)

1-loop

Dyson-equation \rightarrow geometric series



1-loop



ln Q

Why is that?

$$G^{BN}(up) = \frac{1}{up-m} e^{-\gamma \ln(up-m)} = \frac{1}{\sqrt{p-m}} \sum_{n=0}^{\infty} (-1)^n \frac{(\gamma \ln(up-m))^n}{n!}$$
$$R = \infty$$

1-loop

$$G^{1/} = \frac{1}{(up-m)(1+\gamma \ln(up-m))} = \frac{1}{up-m} \sum_{n=0}^{\infty} (-1)^n (\gamma \ln(up-m))^n$$
$$|\gamma \ln(up-m)| < 1$$

The Dyson-series doesn't converge everywhere PT: BREAKS DOWN IN IR

1PI Effective Action

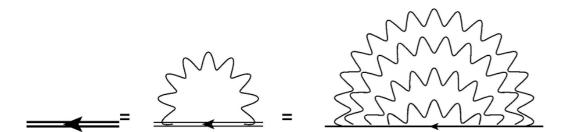
 $Z[J] = \int \mathcal{D}\phi e^{i(S[\phi] + J^i \phi_i)} = e^{iW[J]}$ $i \frac{\delta W[J]}{\delta(i I_i)} = \langle \phi^i \rangle \equiv \Phi^i$ $i\frac{\delta^2 W[J]}{\delta(iI_i)\delta(iI_i)} = \langle T\phi^i \phi^j \rangle_c \equiv G^{ij}$ $\Gamma^{1PI}_{\kappa}[\Phi] = W[J] - J_i \frac{\delta W}{\delta I_i} = W[J] - J_i \Phi^i$ $Z[J] = e^{i(\Gamma^{1PI}[\Phi] + J_i \Phi^i)}$ Gener. functional of 1PI / 2PI diagrams $\left(\phi^{i} = \left\{\phi(x), A^{a}_{\mu}(x), \psi^{a}(x), \ldots\right\}\right)$

2PI Effective Action

 $Z[J,K] = \int \mathcal{D}\phi e^{i(S[\phi] + J^i \phi_i + \frac{1}{2}\phi^i K_{ij}\phi^j)} = e^{iW[J,K]}$ $i \frac{\delta W[J,K]}{\delta(iI_i)} \equiv \Phi^i$ $i\frac{\delta W[J,K]}{\delta(iK:)} \equiv \frac{1}{2}(\Phi^{i}\Phi^{j} + iG^{ij})$ $\Gamma^{2PI} = W[J, K] - J^i \Phi_i - \frac{1}{2} K_{ij} (\Phi^i \Phi^j + i G^{ij})$ $\Gamma^{2PI} = S_0[\Phi] + ic \operatorname{Tr}[\ln G^{-1} + G_0^{-1}G - 1] - i\Gamma_{int}[\Phi, G]$ $\frac{\delta \Gamma^{2PI}(\Phi_s, G_s)}{\delta G} = 0 \rightarrow G_s^{-1} = G_0^{-1} - 2 \frac{\delta \Gamma_{int}[\Phi_s, G_s]}{\delta G}$ $\Sigma[G] = 2 \frac{\Gamma[\Phi, G]}{\delta G}$ Self-consistent Eqs. $G^{-1} = G_0^{-1} - \Sigma[G]$

We need to *handle the IR regime* \longrightarrow new approach: *2PI resummation*

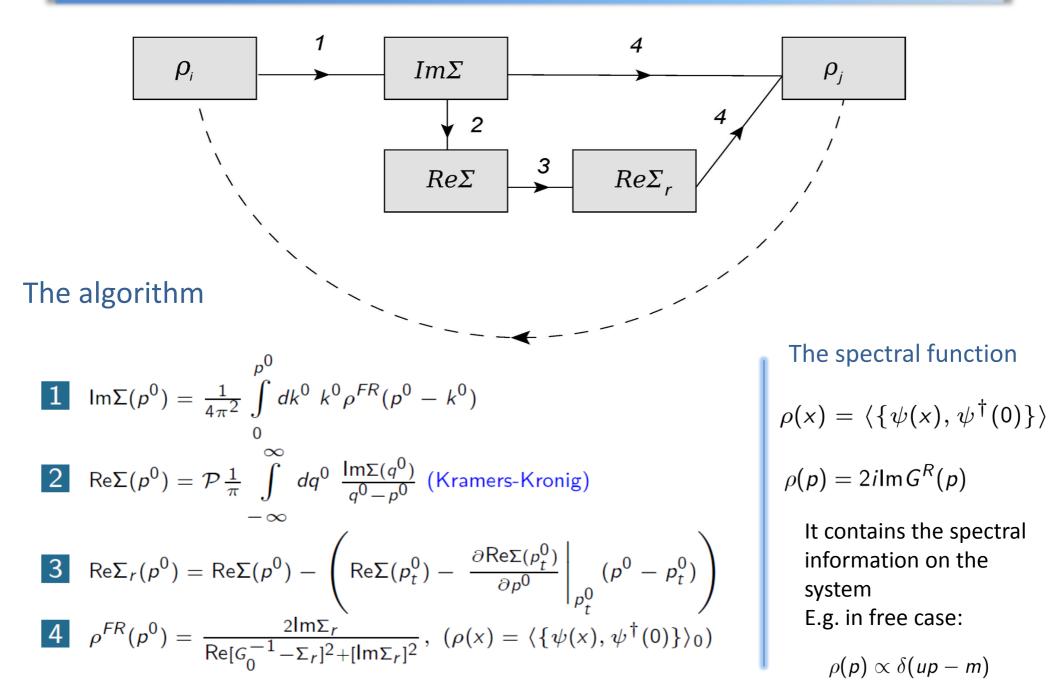
- Summing up the photon-loops (rainbow diagram)
- Treating G as full propagator
- "Quasi particle picture"
- Details: A. Jakovac, Phys. Rev. D76, 125004 (2007). [hep-ph/0612268]



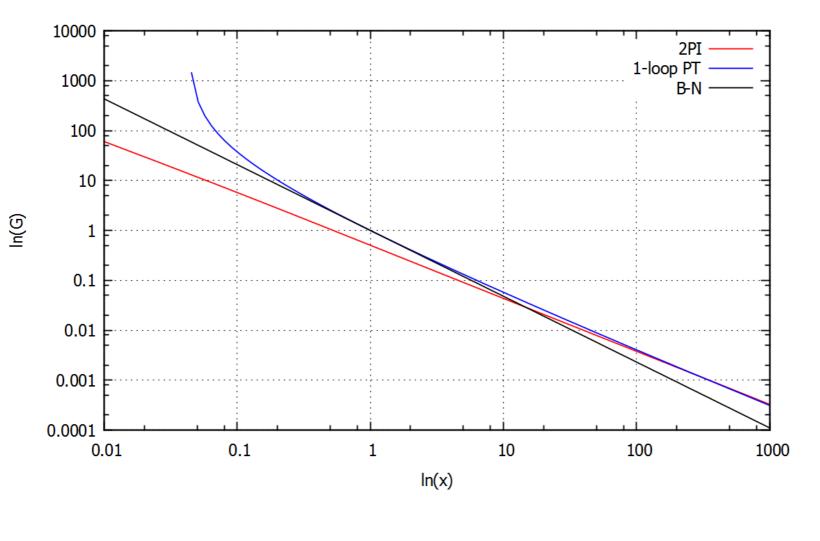
 $G[\Sigma] \Leftrightarrow \Sigma[G]$

Self-consistent equations

 $\Sigma(p) = -ie^2 \int \frac{d^4 k}{(2\pi)^4} G(p-k) \frac{1}{k^2} \qquad G(p) = \frac{1}{G_0^{-1}(p) - \Sigma(p)}$



The propagator

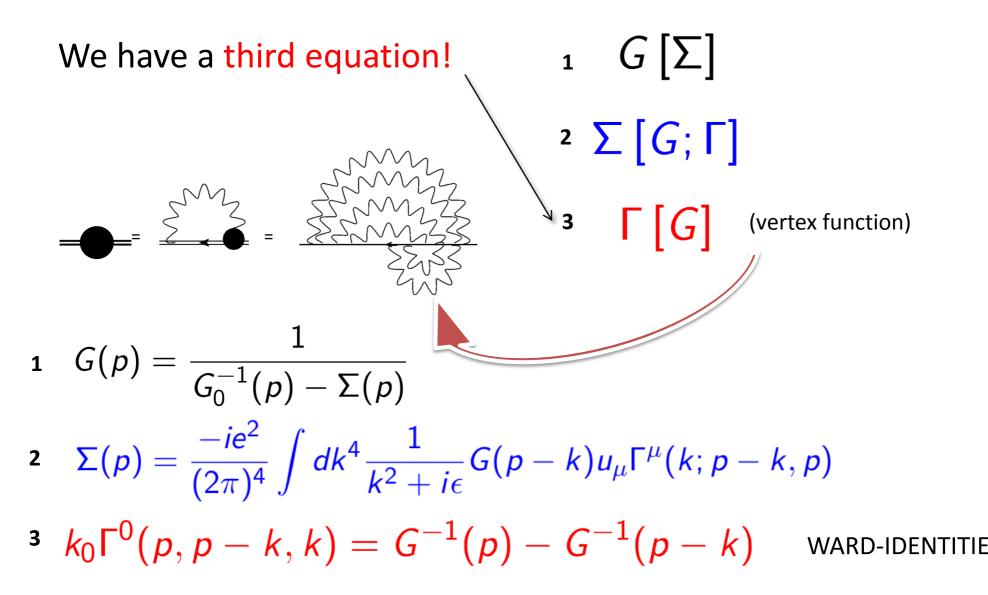


Doesn't fit well

• Works in IR

Resummations Dyson-Scwinger eq.

"Modified 2PI " = 2PI + vertex corrections (D-S eq.)



Dyson-Scwinger eq.

The self-energy:

$$\Sigma(p) = \frac{-ie^2}{(2\pi)^4} \int dk^4 \frac{1}{k^2 + i\epsilon} G(p-k) u_\mu \Gamma^\mu(k; p-k, p)$$

Since $\Gamma^0(p, p - k, k) = \frac{G^{-1}(p) - G^{-1}(p-k)}{k_0}$

And u = (1, 0, 0, 0)

$$\Sigma(p^{0}) = \frac{-ie^{2}}{(2\pi)^{4}}G(p^{0})\int dk^{4}\frac{1}{k^{2}+i\epsilon}\frac{G(p^{0}-k^{0})}{k_{0}}$$

From this we can get

$$(p^{0}-m)G(p^{0}) = \frac{\alpha}{\pi}\int_{p_{0}}^{M}d\omega G(\omega)$$

(+ D-S renorm., A. Jakovac, P. Mati PHYSICAL REVIEW D85, 085006 (2012))

Its solution:

$$G(p^0) = \frac{const.}{(p^0 - m)^{(1 + \frac{\alpha}{\pi})}} \equiv B-N$$

D-S eq. = "Exact Resummation"

Finite Temperature

Let's go to $T \neq 0!$

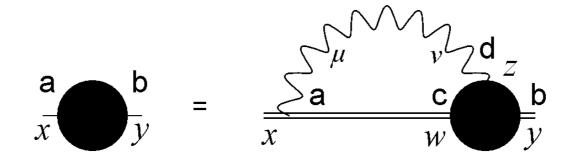




Finite Temperature

What's new?

- The plasma assigns the frame of reference $\rightarrow u = (u_0, u_1, u_2, u_3)$
- F-D and B-E distributions come into picture
- New kind of loop integrals (retarded Self-energy, propagators have matrix structure; R/A & Keldysh formalism!)



Finite Temperature

The path integral formalism:

(A similar matrix structure hold for the self energy)

$$\mathbf{G}^{\mathbf{R}/\mathbf{A}} = \begin{pmatrix} 0 & G_{ar} \\ G_{ra} & G_{rr} \end{pmatrix}$$

Finite Temperature

The propagators:
Keldysh
$$iG_{ab}(x) = \langle T_C O_a(x) O_b^{\dagger}(0) \rangle$$

(a,b=1,2 contours)

(a,b=1,2 contours)

(a,b=1,2 contours)

(b,c)

(c,c)

(

R/A (relations)

$$G_{rr} = \frac{G_{21} + G_{12}}{2}, \quad G_{11} = G_{ra} + G_{12}, \quad \rho = iG_{ra} - iG_{ar}$$

*

The Dirac-Fermi distribution is zero in this model:

- The fermion is a hard probe of the soft photon fields, so it's NOT part of the thermal medium
- It can be checked: at one-loop order it violates causality

contribution is not

singular

The Self sonsistent equations at finite T:

$G[\Sigma] \ 1 \ G_{ra}(p) = G_{ra}^{(0)}(p) + G_{ra}^{(0)}(p) \Sigma_{ar}(p) G_{ra}(p)$	
$\Sigma_{ab}(p) = i(-1)^{a+1}e^2 u_{\mu} \int \frac{d^4k}{(2\pi)^4} \mathcal{G}_{ac}(p\!-\!k) G_{ad}^{\mu u}(k) \Gamma_{ad}(k)$	ν;dcb(k;p−k,p)
$\Gamma[G] \ 3 \ \Gamma_{abc}(k;p,q) = \frac{1}{uk} \left[\delta_{ab} \mathcal{G}_{bc}^{-1}(q) - \delta_{ac} \mathcal{G}_{bc}^{-1}(p) \right]$	
We're only interested in the spectral function	
$war{ ho}(w) = -rac{lpha}{\pi}\int dqf(q,u)ar{ ho}(w-q)$	 Notes: The UV renorm. is the same as at T=0 Finite T

$$f(q) = \frac{u_0^2 - u^2}{2} \int_{u_0 - u}^{u_0 + u} \frac{ds}{us^2} \left(1 + n\left(\frac{q}{s}\right)\right)$$

Solution for the $|\underline{u}|=0$

$$w\bar{\rho}(w) = -\frac{\alpha}{\pi}\int dq \,f(q,u)\,\bar{\rho}(w-q) \longrightarrow w\bar{\rho}(w) = -\frac{\alpha}{\pi}\int dq\,(1+n(q))\bar{\rho}(w-q)$$

By inverse Fourier transformation (convolution ightarrow product) : $i\partial_t \bar{\rho}$

$$\bar{p}(t) = \frac{iT\alpha}{\tanh(\pi tT)}\bar{p}(t)$$

$$ar{
ho}(t)=ar{
ho}_0\,(\sinh\pi tT)^{lpha/\pi}$$

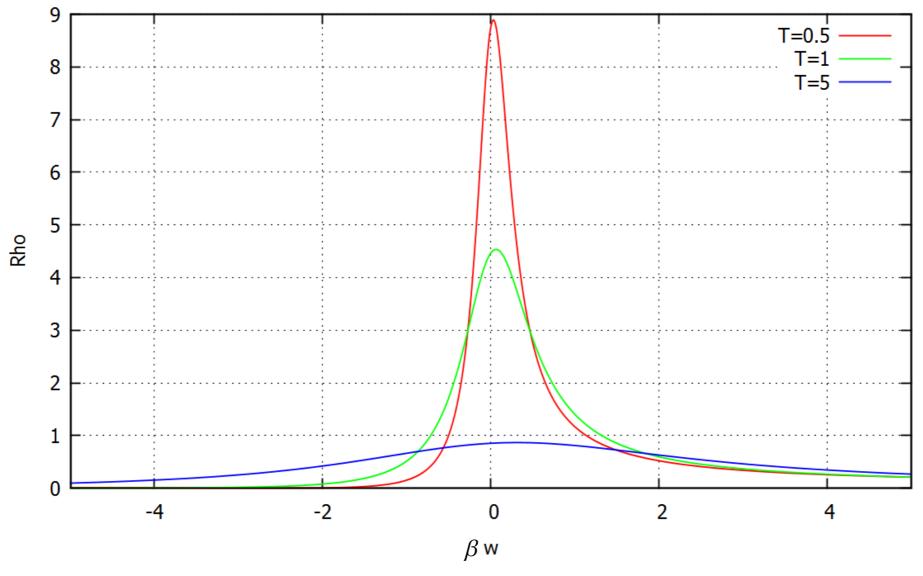
Fourier transform: using analytic continuation

$$\bar{\rho}(w) = \frac{e^{\beta w/2} \sin \alpha}{\cosh(\beta w) - \cos \alpha} \frac{1}{\left|\Gamma\left(1 + \frac{\alpha}{2\pi} + i\frac{\beta w}{2\pi}\right)\right|^2}$$

 $\bar{
ho}(eta w \gg 1) \sim rac{e^{eta w}}{2\cosh(eta w)} rac{1}{w^{1+rac{lpha}{\pi}}} \stackrel{T o 0}{ o} \Theta(w) w^{-1-rac{lpha}{\pi}}$ It give

It gives backl T=0 case!

Finite Temperature



- Spreads out as T increases
- Peak shifted by finite thermal correction

Finite Temperature

Solution for the <u>|u</u> finite case

The asymptotic can be factorized:

$$\overline{\rho}(t) = Z(t)\overline{\rho}_{u=0}(t;\alpha_{eff})$$

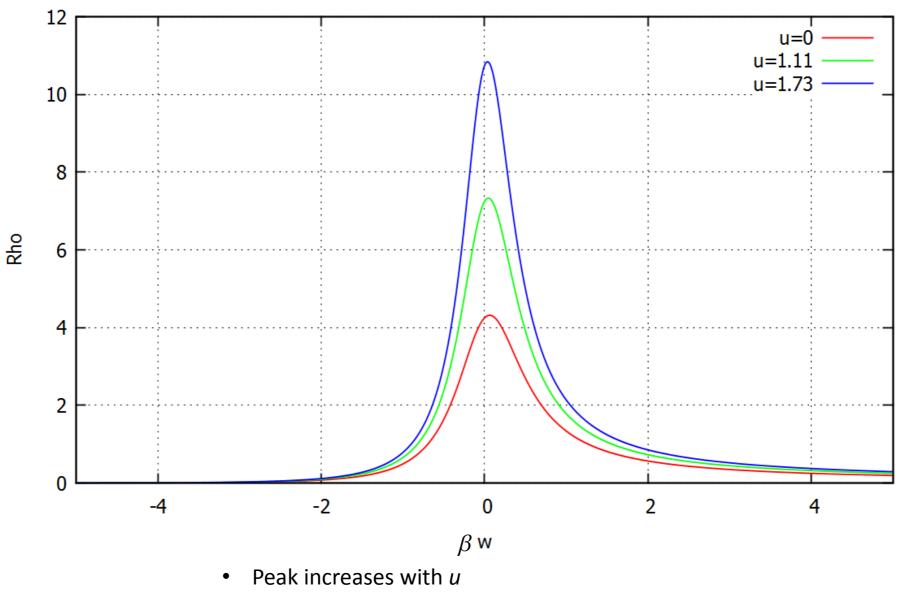
$$\alpha_{eff}(u) = \alpha \frac{u_0^2 - u^2}{2u} \ln \frac{u_0 + u}{u_0 - u} \qquad Z(t) = \exp\left\{\frac{U^2 \alpha}{2\pi u} \int_{u_0 - u}^{u_0 + u} \frac{ds}{s^2} \ln \frac{\sinh \pi t Ts}{(\sinh \pi Tt)^s}\right\}$$

We need the spectral function in the momentum space:

$$\bar{\rho}(t) = Z(t)\bar{\rho}_{u=0}(t;\alpha_{eff}) \xrightarrow{\text{FT}} \bar{\rho}(w) = \int_{-\infty}^{\infty} \frac{dw}{2\pi} Z(w-w')\bar{\rho}_{u=0}(w';\alpha_{eff})$$

This can be solved only numerically

Finite Temperature



• The half-width decreasing with $u \rightarrow$ particle lifetime increases

Conclusion, Outlook

B-N: Exactly solvable model
We made different levels of approximations
The resummations are valid even in IR
We found a new way to solve the model
Possible to extend to finite temperature (closed form for u=0 case!)

<u>Outlook</u>:

- adapting the method to QED
- applications at ELI (strong fields vs. nonperurbative method)
- examine bounded states (IR physics questions)
- adapting the method to QCD (???)



THANK YOU!