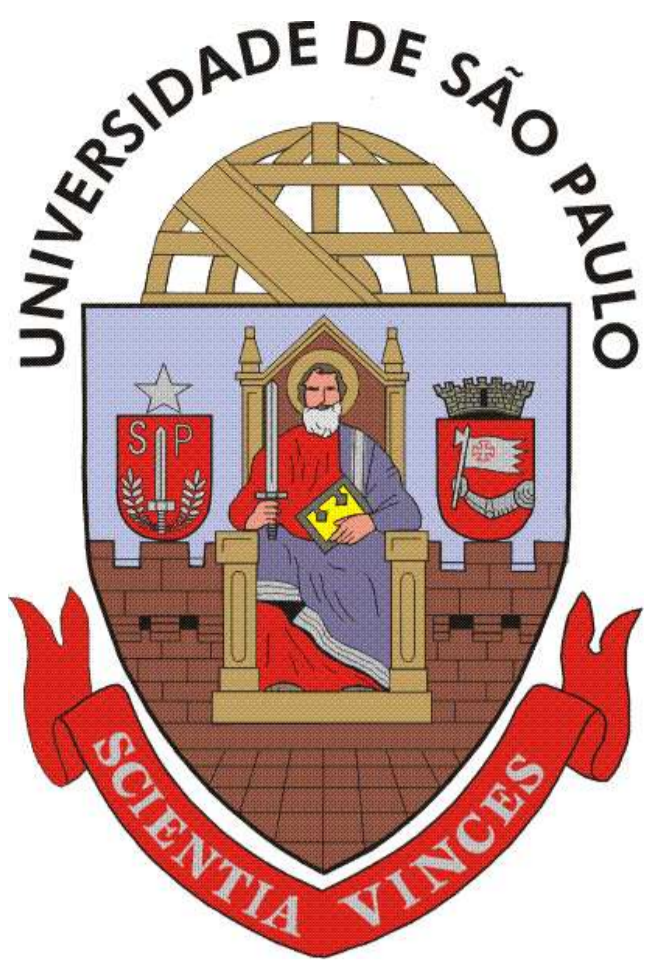


Analytical solutions of 2nd order conformal hydrodynamics



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Introduction

- Second order conformal hydrodynamics (where $\varepsilon = 3p$ and bulk scalar $\Pi = 0$) contains many terms in the equations of motion for the temperature T , flow velocity u^μ , and shear stress tensor $\pi^{\mu\nu}$.
- We have derived the first analytical (and semi-analytical) solutions of 2nd order conformal hydro equations [1,2].
- These solutions are found by mapping a nontrivial flow in flat spacetime to a trivial (static) flow in curved spacetime [3] via a Weyl transformation: $g_{\mu\nu} \rightarrow \Lambda^2 \hat{g}_{\mu\nu}$.

2nd Order Conformal Hydrodynamics

Energy-momentum conservation $\nabla_\mu T^{\mu\nu} = 0$ (∇_μ is the space-time covariant derivative):

$$D\varepsilon + (\varepsilon + p)\vartheta + \pi^{\mu\nu}\sigma_{\mu\nu} = 0 \quad (\varepsilon + p)Du^\mu + \Delta^{\mu\alpha}\nabla_\alpha p + \Delta^\mu_\nu \nabla_\alpha \pi^{\alpha\nu} = 0, \quad (1)$$

where we defined the comoving derivative $D \equiv u^\mu \nabla_\mu$ and fluid expansion rate $\vartheta \equiv \nabla_\mu u^\mu$, while

$$\sigma^{\mu\nu} \equiv \nabla^{\langle\mu} u^{\nu\rangle} \equiv \left(\frac{1}{2} (\Delta^{\mu\alpha} \Delta^{\nu\beta} + \Delta^{\mu\beta} \Delta^{\nu\alpha}) - \frac{1}{3} \Delta^{\mu\nu} \Delta^{\alpha\beta} \right) \nabla_\alpha u_\beta, \quad (2)$$

is the shear tensor. We employ the result of Denicol et al. [4] and generalize it to curved spacetimes [1,2] taking into account the constraints from conformal symmetry [5] to find the most general equation for $\pi^{\mu\nu}$

$$\begin{aligned} \pi^{\mu\nu} = & -2\eta\sigma^{\mu\nu} - \tau_\pi \left(\Delta^\mu_\alpha \Delta^\nu_\beta D\pi^{\alpha\beta} + \frac{4}{3}\pi^{\mu\nu}\vartheta \right) + \lambda_2 \pi^\mu_\lambda \Omega^{\nu\lambda} \\ & + \lambda_1 \pi^\mu_\lambda \pi^{\nu\lambda} + \lambda_3 \Omega^\mu_\lambda \Omega^{\nu\lambda} - \tau_{\pi\pi} \sigma^\mu_\lambda \pi^{\nu\lambda} - \tilde{\eta}_3 \sigma^\mu_\lambda \sigma^{\nu\lambda} - \tilde{\eta}_4 \sigma^\mu_\lambda \Omega^{\nu\lambda} \\ & + \tau_\sigma \left(\Delta^\mu_\alpha \Delta^\nu_\beta D\sigma^{\alpha\beta} + \frac{1}{3}\sigma^{\mu\nu}\vartheta \right) + \kappa \left(\mathcal{R}^{\langle\mu\nu\rangle} - 2u_\alpha \mathcal{R}^{\alpha\langle\mu\nu\rangle} u_\beta \right) \end{aligned} \quad (3)$$

where $\Omega^{\mu\nu} \equiv \frac{1}{2} \Delta^{\mu\alpha} \Delta^{\nu\beta} (\nabla_\alpha u_\beta - \nabla_\beta u_\alpha)$ is the vorticity tensor and η is the shear viscosity that appears in the first-order (Navier-Stokes) theory.

Gubser flow in Israel-Stewart theory

Gubser generalized Bjorken's solution by including a nontrivial x_\perp -dependence while retaining boost invariance [3] by performing the Weyl transformation of the metric

$$ds^2 \equiv \frac{ds^2}{\tau^2} = \frac{-d\tau^2 + dx_\perp^2 + x_\perp^2 d\phi^2}{\tau^2} + d\eta^2 = -d\varrho^2 + \cosh^2 \varrho (d\Theta^2 + \sin^2 \Theta d\phi^2) + d\eta^2, \quad (4)$$

where

$$\sinh \varrho = -\frac{L^2 - \tau^2 + x_\perp^2}{2L\tau}, \quad \tan \Theta = \frac{2Lx_\perp}{L^2 + \tau^2 - x_\perp^2}. \quad (5)$$

Eq. (4) shows that Minkowski space is conformal to $dS_3 \times \mathbb{R}$ (dS_3 is the 3-dimensional de Sitter space) up to a Weyl rescaling factor $\Lambda^2 = \tau^2$. The parameter L has dimension of length and is identified with the 'radius' of dS_3 . The flow velocity of the static fluid in $dS_3 \times \mathbb{R}$ is simply $\hat{u}_\varrho = (-1, 0, 0, 0)$. The flow 4-vector in Minkowski space (Milne coordinates) is (with $u_\eta = u_\phi = 0$)

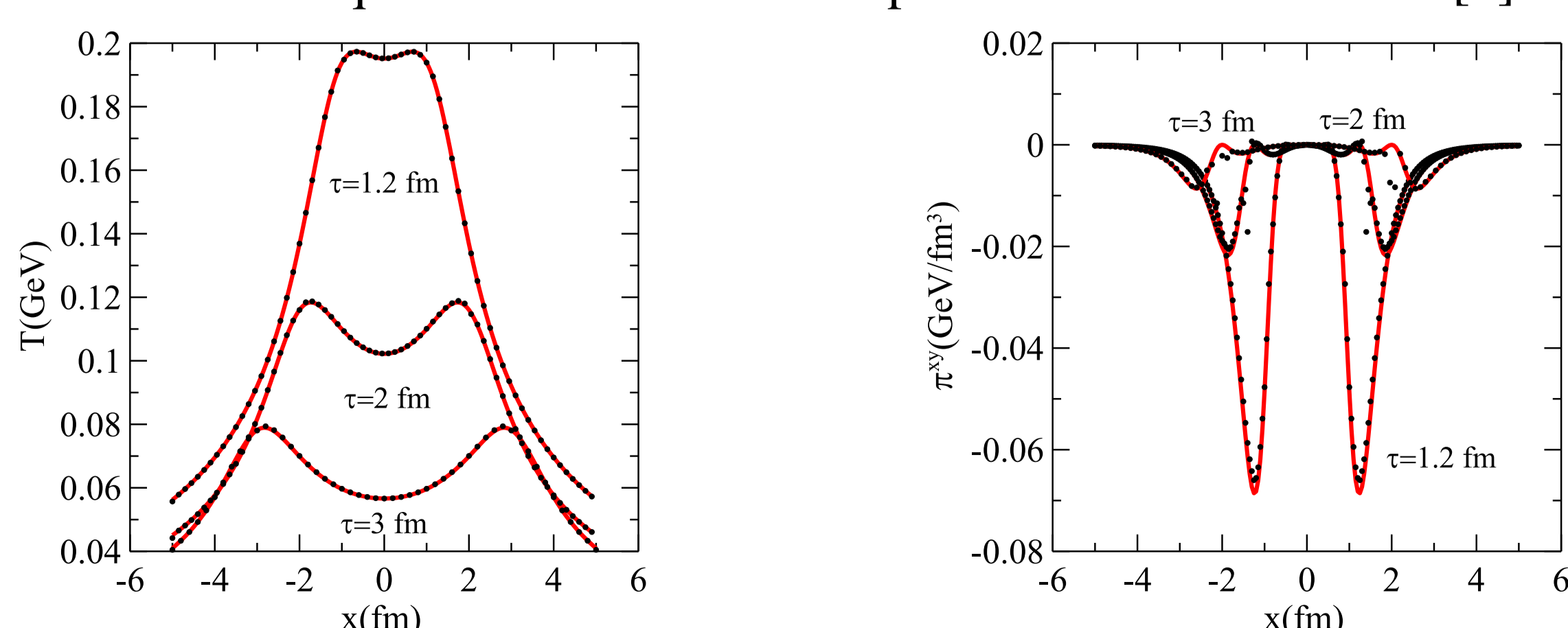
$$u_\tau = -\cosh \left[\tanh^{-1} \frac{2\tau x_\perp}{L^2 + \tau^2 + x_\perp^2} \right], \quad u_\perp = \sinh \left[\tanh^{-1} \frac{2\tau x_\perp}{L^2 + \tau^2 + x_\perp^2} \right]. \quad (6)$$

Comparing semi-analytical solution to MUSIC ($\eta/s = 0.2$)

We take the simplest conformal theory in which only η and $\tau_\pi = 5(\eta/s)/T$ are nonzero. The equations drastically simplify in $dS_3 \times \mathbb{R}$ and we have

$$\frac{1}{\hat{T}} \frac{d\hat{T}}{d\rho} + \frac{2}{3} \tanh \rho = \frac{1}{3} \hat{\pi}_\eta^\eta(\rho) \tanh \rho \quad \frac{5\eta}{\hat{T}s} \left[\frac{d\hat{\pi}_\eta^\eta}{d\rho} + \frac{4}{3} (\hat{\pi}_\eta^\eta)^2 \tanh \rho \right] + \hat{\pi}_\eta^\eta = \frac{4}{3s\hat{T}} \tanh \rho, \quad (7)$$

where $\hat{\pi}_\eta^\eta \equiv \hat{\pi}_\eta^\eta / (\hat{T}\hat{s})$. These equations can then be solved numerically. An analytical solution to these equations valid at low temperatures can be found in [1].



First full analytical solution of 2nd order hydrodynamics

In [2], several solutions of the 2nd order hydro equations were found using the same idea: nontrivial flow in flat spacetime is taken, via a Weyl transformation, to a trivial flow in curved spacetime. The transformation is

$$\begin{aligned} ds^2 & \equiv \frac{ds^2}{x_\perp^2} = \frac{-dt^2 + dz^2 + dx_\perp^2}{x_\perp^2} + d\phi^2 \\ & = -\cosh^2 \rho dT^2 + d\rho^2 + \sinh^2 \rho d\bar{\Theta}^2 + d\phi^2. \end{aligned} \quad (8)$$

where

$$\tan T = \frac{L^2 + r^2 - t^2}{2Lt}, \quad \cosh \rho = \frac{1}{2Lx_\perp} \sqrt{(L^2 + (r+t)^2)(L^2 + (r-t)^2)}. \quad (9)$$

The metric in Eq. (8) is that of $AdS_3 \times S^1$ where AdS_3 is 3-dimensional Anti-de Sitter space. In this space the 4-flow is again trivial but in Minkowski space

$$u_t = -\frac{L^2 + r^2 + t^2}{\sqrt{(L^2 + (r+t)^2)(L^2 + (r-t)^2)}}, \quad \vec{u} = \frac{2t\vec{r}}{\sqrt{(L^2 + (r+t)^2)(L^2 + (r-t)^2)}}. \quad (10)$$

An interesting property of this flow is that $\hat{\sigma}_{\mu\nu} = \hat{\Omega}_{\mu\nu} = 0$ and $\hat{\vartheta} = 0$. Therefore, the full equation for the shear stress tensor $\hat{\pi}^{\mu\nu}$ simplifies to

$$\hat{\pi}^{\mu\nu} = \frac{\lambda_1}{\hat{\varepsilon}} \hat{\pi}^\mu_\lambda \hat{\pi}^{\nu\lambda}. \quad (11)$$

Note that this equation has a trivial solution ($\hat{\pi}^{\mu\nu} = 0$) but there are other nontrivial solutions as well. Assuming $\hat{\pi}^{\mu\nu}$ is diagonal, we find the non-perturbative solutions in λ_1

$$(\hat{\pi}^{\rho\rho}, \sinh^2 \rho \hat{\pi}^{\Theta\Theta}, \hat{\pi}^{\phi\phi}) = \frac{\hat{\varepsilon}}{\lambda_1} \times \begin{cases} (-1, -1, 2), \\ (-1, 2, -1), \\ (2, -1, -1). \end{cases} \quad (12)$$

The corresponding energy density in Minkowski space reads [2]

$$\varepsilon \propto \begin{cases} \frac{1}{(L^2 + (t+r)^2)^2 (L^2 + (t-r)^2)^2} \left(\frac{4L^2 x_\perp^2}{(L^2 + (t+r)^2)(L^2 + (t-r)^2)} \right)^{\frac{9}{2(\lambda_1-3)}}, \\ \frac{1}{(L^2 + (t+r)^2)^2 (L^2 + (t-r)^2)^2} \left(1 - \frac{4L^2 x_\perp^2}{(L^2 + (t+r)^2)(L^2 + (t-r)^2)} \right)^{\frac{9}{2(\lambda_1-3)}}, \\ \frac{1}{(L^2 + (t+r)^2)^2 (L^2 + (t-r)^2)^2} \left(\frac{4L^2 x_\perp^2 ((L^2 + (t+r)^2)(L^2 + (t-r)^2) - 4L^2 x_\perp^2)}{(L^2 + (t+r)^2)^2 (L^2 + (t-r)^2)^2} \right)^{-\frac{9}{2(\lambda_1+6)}}. \end{cases} \quad (13)$$

Note that ε is even under time reversal $t \rightarrow -t$ even though $\pi^{\mu\nu}$ is nonzero !!!

Conclusions and Outlook

- These analytical (and semi-analytical) solutions of 2nd order viscous conformal hydrodynamics can be readily used to check the precision of existing numerical codes that solve the (2+1) and (3+1) relativistic hydrodynamic equations.
- The MUSIC code from the McGill group has been shown to reproduce these analytical (and semi-analytical) solutions to a good approximation in the case of Gubser flow.
- Full analytical solutions of the 2nd order conformal hydrodynamic equations, which include nonzero vorticity, can be found using the method shown here (for details see [2]).
- Due to non-linear structure of the equation of motion for $\pi^{\mu\nu}$ (particular, the term $\frac{\lambda_1}{\hat{\varepsilon}} \hat{\pi}^\mu_\lambda \hat{\pi}^{\nu\lambda}$), one can see that there are solutions in 2nd order hydrodynamics in which $\pi_{\mu\nu} \neq 0$ but time reversal invariance is not broken.
- It would be interesting to see how these solutions may appear in other approaches such as kinetic theory and the fluid/gravity correspondence.

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