### nuSTORM meeting

chaired by Kenneth Long (Imperial College London), Elena Wildner (CERN)

from Tuesday, March 26, 2013 at 12:30 to Wednesday, March 27, 2013 at 19:00 (US/Central) at CERN

866-2-D05 & 864-1-C02

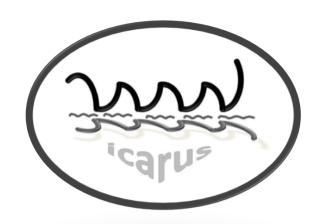
## The NESSiE concept for Sterile Neutrinos

- 1) The Sterile Neutrino Issue with SPS beam at CERN
- 2) Which Detector(s)?

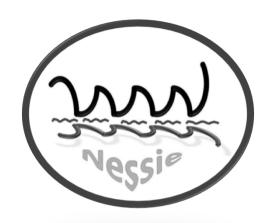


Luca Stanco (INFN-PD)

March 27, 2012



## Proposal SPSC-P347 (March 2012) (ICARUS + NESSIE)



- 1. Physics ground & Different Scenarios
- 2. Latest Beam & new simulation
- 3. Spectrometers
- 4. CERN Schedule disantengling

#### "Tensions" in the Standard Model between Quark and Lepton sectors:

- No right handed neutrino (in classical SM)
- Neutrinos are neutral
- Majorana masses (the HIERARCHY problem)
- Very different from quarks: large mixings and tiny masses (the FLAVOUR puzzle)
- Non coherent picture: discrepancies at ≈ 3-4 sigma level, in several measurements...

As Higgs is going to be where expected and with expected properties ... As many open questions stay (e.g. Dark Matter)...

Neutrinos are an excellent place for looking Beyond Standard Model

Neutrino oscillations have established a picture through a large number of experiments consistent with the mixing of three physical neutrino  $v_e$ ,  $v_\mu$  and  $v_\tau$  with mass eigenstates  $v_1$ ,  $v_2$  and  $v_3$ .

#### BUT

Anomalies appear in neutrino data in the region of  $\Delta m^2 \sim eV^2$ 

Predominantly from single detector experiments...

Are they measurement "errors" or are they due to unaccounted oscillations between EW active neutrinos and other non-active neutrinos?

### Sterile neutrinos

- O The possible presence of oscillations into sterile neutrinos was proposed by B. Pontecorvo, (*JETP*, 53, 1717, 1967), but so far without conclusion.
- $\circ$  "Sterile" means "No Standard Model Interactions" (i.e think to anti- $v_R$ , light neutrinos which can oscillate with "active" neutrinos)
- O Smoking Gun: Neutral Current Deficit
- O Counterchecked Smoking Gun: NC/CC ratios
- O Two distinct classes of anomalies have been analyzed, namely
  - the apparent <u>disappearance signal</u> in the anti- $v_e$  events detected from (1) near-by nuclear reactors and (2) the from Mega-Curie k-capture calibration sources in the Gallium experiments to detect solar  $v_e$
  - bobservation for <u>excess signals</u> of anti- $v_e$  electrons from neutrinos from particle accelerators (LNSD/MiniBooNE)
- O At least a fourth non-standard neutrino state can oscillate at small distances,  $\Delta m_{new}^2 \approx 1 \text{ eV}^2$  ( SHORT BASELINE projects)

# PHYSICS

What do you want to discover?

What do you need/want to disprove ?

What do you need to measure?

different answers to different questions:

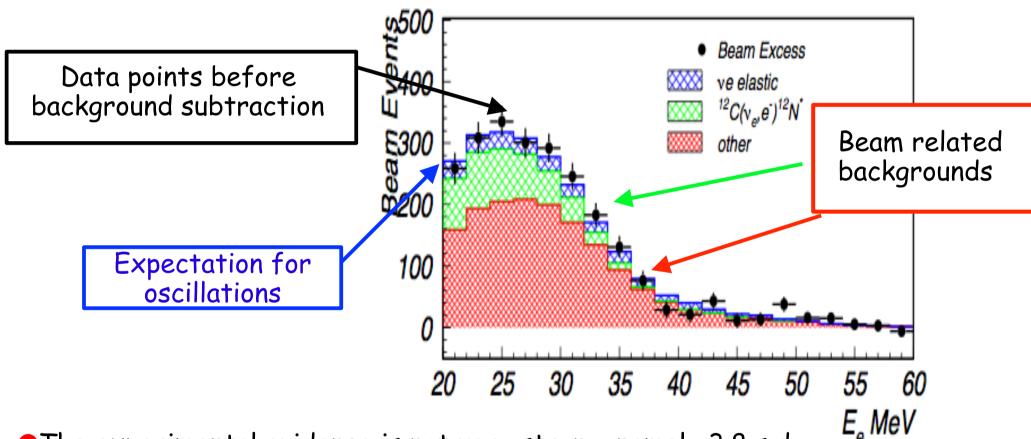
You want to discover New Physics

You want to outcome Negative Physics

You want to use New Technologies

## LSND: Evidence for $(v_{\mu} \rightarrow v_{e})$

Excess of events:  $87.9 \pm 22.4 \pm 6.0$ 



- The experimental evidence is not very strong, namely 3.8 s.d.
- The experimental result so far has not been challenged experimentally

MiniBooNE (USA, 15 years of work/money/people) was inconclusive

## Going on and next Future

ANOMALIES/Steriles

Radioactive sources Reactors Beams

None of these experiments can be considered the definitive one:, which is

- measure v and anti-v, muons ed electrons
- in different sites (≥ 2)
- provide a 5 sigma result

We need a Superior Class Experiment: 3 Kton Fe + 1 Kton LAr

The ultimate 3+1 experiment!

## CERN

- CERN, towards creating an infrastructure (actually a MULTI-structure) at North Area for beams/R&D/multi-experiments
- Underway
- Multistep project
   Begin with SBL and R&D neutrino detectors,
   proceed with nuStorm and LBNO
- What? (decide on tecnology)
   we are proposing LAr+Fe

Mandatory: measure v-muon disappearance (see yesterday Stephen Parke presentation and following discussion)

# CERN DG requested a detailed document in November, ready on February 7<sup>th</sup>.

#### Letter of Intent for the New CERN Neutrino Facility

**Draft 0 - v1.** 

**January 11, 2013** 

Leader Project: Marzio Nessi (Cern)

Deputy: Rende Steerenberg (Cern)

#### **Neutrinos CERN Community**

#### **ABSTRACT:**

The nature and characteristics of neutrinos still represent one of the most intriguing questions of modern particle physics. This Letter of Intent presents a plan to build at CERN a Short Baseline Neutrino Facility capable to meet this challenge. The new beam line will make use of the SPS beam, it will be located in the CERN North Area of the Prevessin site and will comprise in a first step 2 experimental halls (far and near) which will host the detector facilities. The beam layout will be compatible with a long neutrino baseline which might be constructed in a second stage.

KEYWORDS: Neutrino, Short-baseline, Long-baseline, Sterile Neutrinos, ICARUS, NESSIE, LAGUNA, LBNO, CERN.

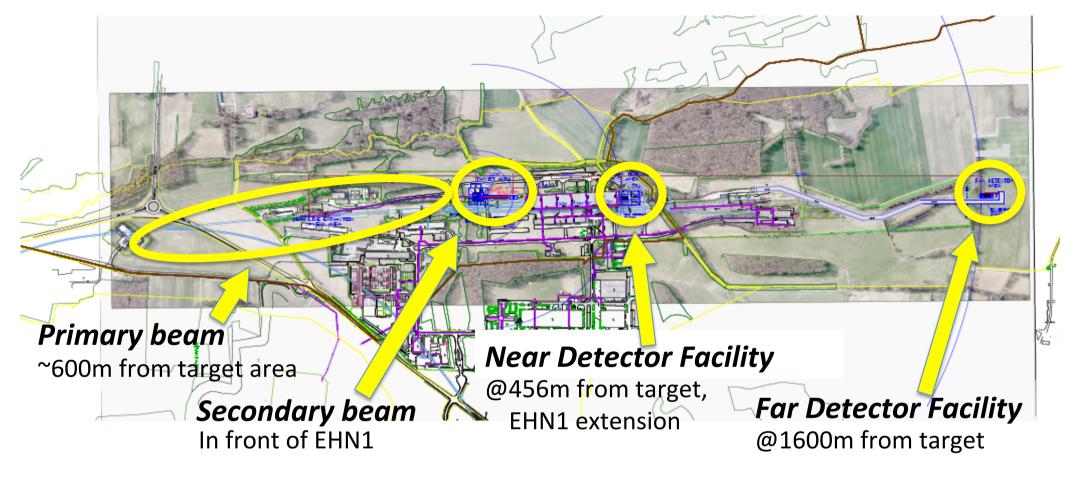
- > 200 pages for 5 Work-Packages,
- > a TDR for Beam, Infrastructures and Experiments

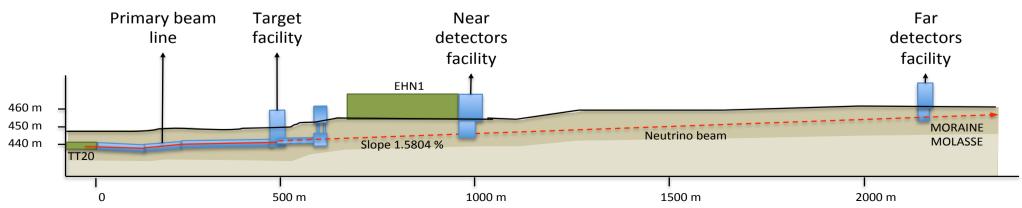


## In practice (existing requests)

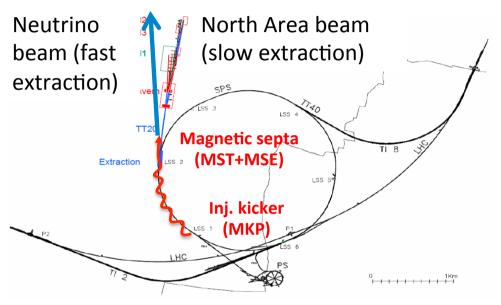
- Interesting experimental proposal by ICARUS-NESSiE (SPSC-P-347)
   for the search of sterile neutrinos
- Demonstrate a new generation of neutrino detectors of doublephase LAr TPC prototype by the LAGUNA-LBNO consortium (SPSC-E-007)

Provide a neutrino beam at CERN in the medium term (~4 years) with 2 experimental facilities ( near and far ). Satisfy the existing requests and allow new detectors/ideas to come forward



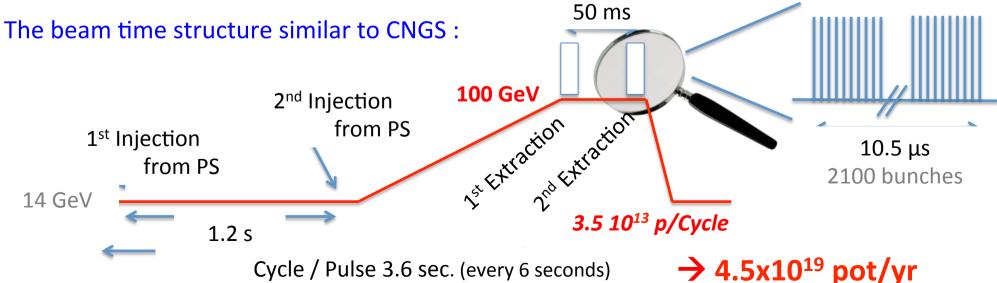


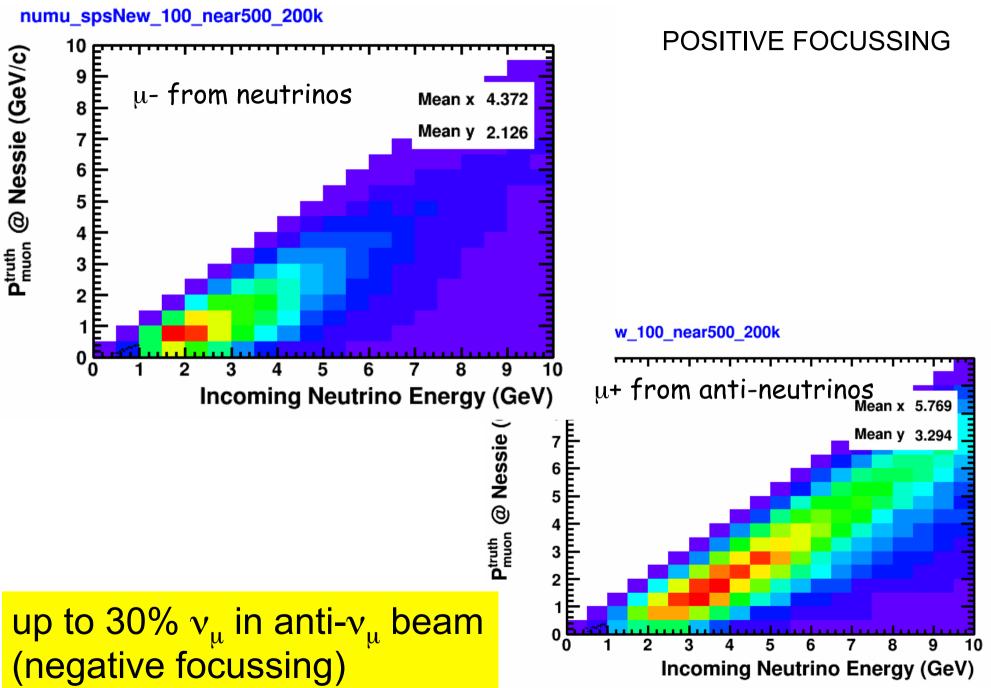
## Which type of beam can CERN offer?



### Fast extraction

- Beam excitation via injection kicker in LSS1 + extraction in LSS2 via existing septa (incompatibility with simultaneous north area slow extracted beam!)
- Solution tested for low intensities during recent beam tests

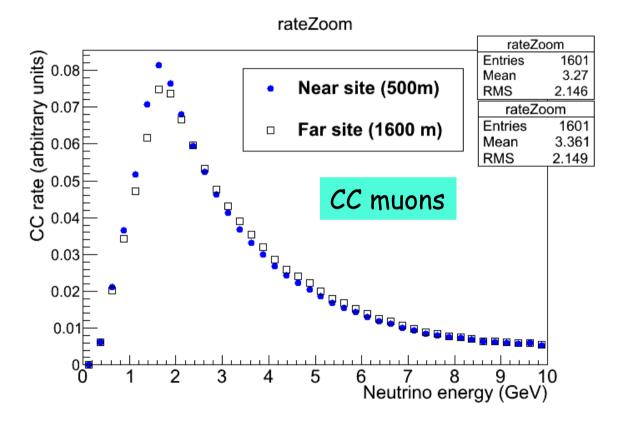




### Latest Beam Studies at SPS

100 GeV proton, Fast Extraction (10.5 μs), Luminosity as at CNGS

On-axis configuration Event rates 4.5 10¹9 pot (≈1 year)



Unoscillated  $v_e$  fluxes are ~ identical  $\rightarrow$  N/F deviations = oscillations The oscillated signals are clustered below 6 GeV of visible energy

Scenario defined for DATA Taking: 2 years of anti-v followed by 1 year v

## Start of the game: this year (OR NOT VALUABLE)

The project has to be approved this year, so to allow CERN to start construction

MANDATORY FOR US

## Search for "anomalies" from neutrino and anti-neutrino oscillations at $\Delta m^2 \approx 1 eV^2$ with muon spectrometers and large LAr–TPC imaging detectors.

#### Technical proposal.

(CERN-SPSC-2012-010 and SPSC-P-347)

2 collaborations 30 institutions ~140 people

#### **ICARUS Collaboration**

M. Antonello¹, D. Bagliani², B. Baibussinov⁵, H. Bilokon⁶, F. Boffelli³, M. Bonesiniゥ, E. Calligarich³, N. Canci¹, S. Centro⁴,⁵, A. Cesana¹⁰, K. Cieslik¹¹, D. B. Cline¹², A. G. Cocco¹⁴, D. Dequal⁴,⁵, A. Dermenev¹⁶, R. Dolfini⊓, M. De Gerone², S. Dussoni², C. Farnese⁴, A. Fava⁵, A. Ferrari¹, G. Fiorillo¹₃¹⁴, G. T. Garvey¹³, F. Gatti², D. Gibin⁴,⁵, S. Gninenko¹⁶, F. Guber¹⁶, A. Guglielmi⁵, M. Haranczyk¹¹, J. Holeczek¹ゅ, A. Ivashkin¹⁶, M. Kirsanov¹⁶, J. Kisiel¹ゅ, I. Kochanek¹ゅ, A. Kurepin¹⁶, J. Łagoda²⁰, G. Lucchiniゥ, W. C. Louis¹³, S. Mania¹ゥ, G. Mannocchi⁶, S. Marchini⁵, V. Matveev¹⁶, A. Menegolli⊓, G. Meng⁵, G. B. Mills¹³, C. Montanari³, M. Nicoletto⁵, S. Otwinowski¹², T. J. Palczewski²⁰, G. Passardi¹, F. Perfetto¹₃¹⁴, P. Picchi⁶, F. Pietropaolo⁵, P. Płoński²¹, A. Rappoldi³, G. L. Raselli³, M. Rossella³, C. Rubbia¹, P. Sala¹⁰, A. Scaramelli¹⁰, E. Segreto¹, D. Stefan¹, J. Stepaniak²⁰, R. Sulej²⁰, O. Suvorova¹⁶, M. Terrani¹⁰, D. Tlisov¹⁶, R. G. Van de Water¹³, G. Trinchero⁶, M. Turcato⁵, F. Varanini⁴, S. Ventura⁵, C. Vignoli¹, H. G. Wang¹², X. Yang¹², A. Zani³, K. Zaremba²¹

(a) Contact Person

#### **NESSiE Collaboration**

M. Benettoni<sup>5</sup>, P. Bernardini<sup>26,27</sup>, A. Bertolin<sup>5</sup>, C. Bozza<sup>31</sup>, R. Brugnera<sup>4,5</sup>, A. Cecchetti<sup>6</sup>, S. Cecchini<sup>25</sup>, G. Collazuol<sup>5,6</sup>, P. Creti<sup>27</sup>, F. Dal Corso<sup>5</sup>, I. De Mitri<sup>26,27</sup>, G. De Robertis<sup>23</sup>, M. De Serio<sup>23</sup>, L. Degli Esposti<sup>25</sup>, D. Di Ferdinando<sup>25</sup>, U. Dore<sup>29,30</sup>, S. Dusini<sup>5</sup>, P. Fabbricatore<sup>3</sup>, C. Fanin<sup>5</sup>, R. A. Fini<sup>23</sup>, G. Fiore<sup>27</sup>, A. Garfagnini<sup>4,5</sup>, G. Giacomelli<sup>24,25</sup>, R. Giacomelli<sup>25</sup>, G. Grella<sup>31</sup>, C. Guandalini<sup>25</sup>, M. Guerzoni<sup>25</sup>, U. Kose<sup>5</sup>, G. Laurenti<sup>25</sup>, M. Laveder<sup>4,5</sup>, I. Lippi<sup>5</sup>, F. Loddo<sup>23</sup>, A. Longhin<sup>6</sup>, P. Loverre<sup>29,30</sup>, G. Mancarella<sup>26,27</sup>, G. Mandrioli<sup>25</sup>, A. Margiotta<sup>24,25</sup>, G. Marsella<sup>27,28</sup>, N. Mauri<sup>6</sup>, E. Medinaceli<sup>4,5</sup>, A. Mengucci<sup>6</sup>, M. Mezzetto<sup>5</sup>, R. Michinelli<sup>25</sup>, M. T. Muciaccia<sup>22,23</sup>, D. Orecchini<sup>6</sup>, A. Paoloni<sup>6</sup>, A. Pastore<sup>22,23</sup>, L. Patrizii<sup>25</sup>, M. Pozzato<sup>24,25</sup>, R. Rescigno<sup>31</sup>, G. Rosa<sup>29</sup>, S. Simone<sup>22,23</sup>, M. Sioli<sup>24,25</sup>, G. Sirri<sup>25</sup>, M. Spurio<sup>24,25</sup>, L. Stanco<sup>5,b</sup>, S. Stellacci<sup>31</sup>, A. Surdo<sup>27</sup>, M. Tenti<sup>24,25</sup>, V. Togo<sup>25</sup>, M. Ventura<sup>6</sup> and M. Zago<sup>5</sup>.

(b) Contact Person



# The ICARUS-NESSiE P-347 proposal at the CERN-SPS

SPSC-P-347 (arXiv:1203.3432)

- $^{\circ}$  L/E oscillation path lengths to ensure appropriate matching to the  $\Delta m^2$  window for the expected anomalies
- NEAR and FAR sites
- "Imaging" LAr-TPC detector capable of identifying unambiguously <u>all</u> reaction channels
- Magnetic spectrometers to determine muon charge and momentum
- Interchangeable v and anti-v beams
- $^{\bullet}$  High rates due to detector large masses , in order to record relevant effects at the percent level (>10^6  $\nu_\mu$  , ~ 10^4  $\nu_e$ )
- Both initial  $\nu_e$  and  $\nu_\mu$  components cleanly identified.

Couple the Best and Biggest Lar-TPC presently working with a relatively cheap but <u>essential</u> Muon detector

Charge and momentum measurements in Neutrino Interactions for the Charge Current mode: <u>essential</u> and <u>challenging</u>

#### Essential because:

- increase the active target mass by measuring the muon momentum
- increase the range of  $\Delta m^2$  (at higher values, especially in the eV<sup>2</sup> range)
- · calibration of the beam with a clean muon measurement at high p
- normalization point for the NC/CC rates
- $\bullet$  clean separation of v and vbar interactions
- disentangle the  $\nu$  and  $\nu$ bar reverse contaminations in the beams

#### Challenging because:

• find best compromise between passive and active materials

#### GOALS/CAVEATS

#### Perform a DESIGN STUDY for a two-module Detector

- cheap
- almost no R&D
- fast
- fully compatible with the LAr proposal(s)
- standalone
- maximizing the physics output

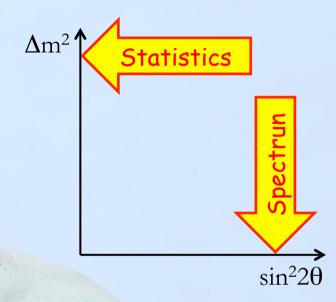
#### SpectrometerS are essential to allow

- full Charge ID
- systematics and calibration controls
- clean separation v and anti-v

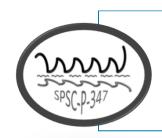
### NESSiE detector concept

#### NESSiE (Neutrino Experiment with SpectrometerS in Europe)

- Goal:
  - Allow charge separation and momentum measurement of as many muons as possible escaping from LAr (large statistics ↔ low sin²20)
  - Go as low as possible in muon momentum (low momenta  $\leftrightarrow$  low  $\Delta m^2$ )
  - Possibility to also study (NESSiE) internal events (coarser resolution w.r.t. LAr)



- Solution:
  - Air-core magnets for low momentum muons escaping from LAr  $(E_u < 0.5 \text{ GeV/c in NESSiE} \leftrightarrow \langle E_v \rangle < 1 \text{ GeV in LAr})$
  - Downstream massive iron dipolar magnets for higher momenta extension



## **ICARUS-NESSIE**

A coupled system of LAr detectors and Muon Spectrometers

NEAR SITE @300 m:

LAr mass = 119 t

Iron magnet mass = 840 t

5 interactions/spill

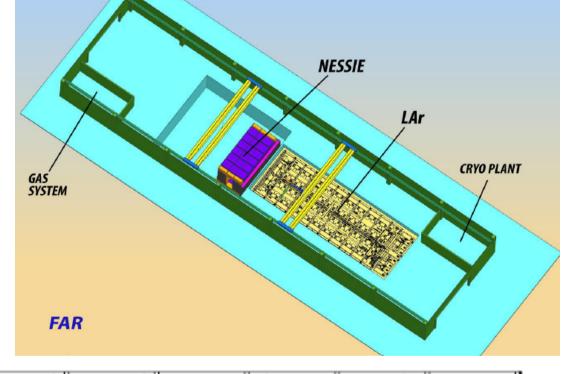
2·10<sup>13</sup> p/spill (with "+" polarity)

**FAR SITE** @1600 m:

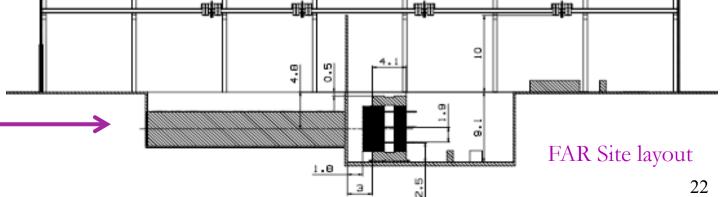
LAr mass = 476 t

Iron magnet mass = 1515 t

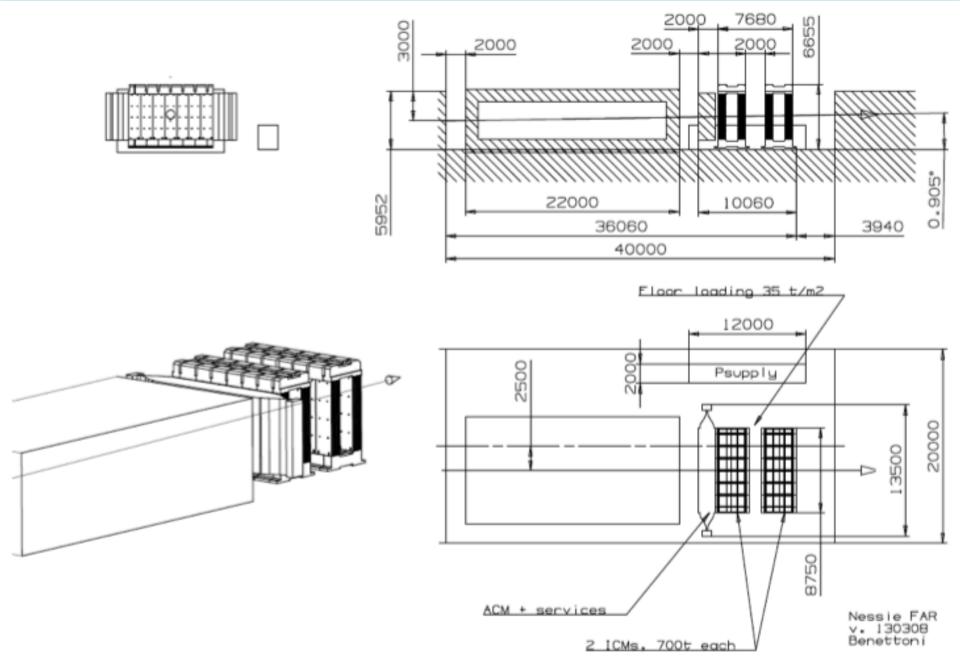
0.65 interactions/spill



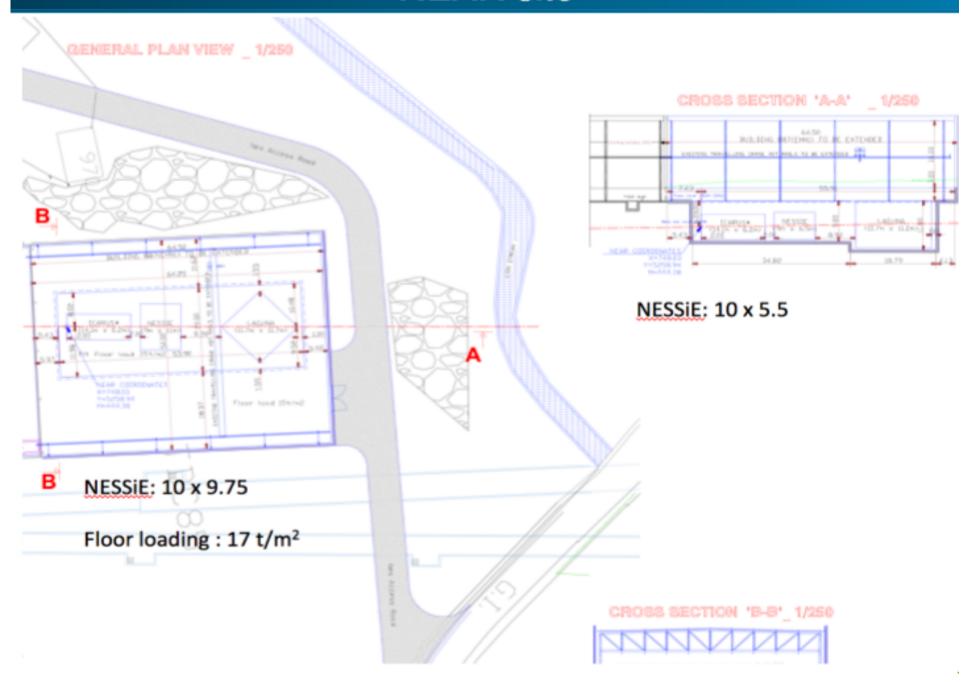
Observation  $v_{\mu}$ ,  $v_{e}$  CC, NC channels Charge separation and muon momentum



## **FAR site**



## **NEAR site**



## e.g. Expected Events in 1 year of Running ( $v_{\mu}$ or $v_{\mu}$ -bar)

To reconstruct: 5.3 M muons in LAr (Near), 0.67 M muons in Lar (Far) pos.foc. 5.2 M muons in Nessie (Near), 0.42 M in Nessie (Far) (with factor 2 in overhead of triggers, positive focussing)

	NEAR (v-bar)	NEAR(v)	FAR(v-bar)	FAR(v)
$v_e + v_e$ -bar (LAr)	35 K	54 K	4.2 K	6.4 K
$v_{\mu}$ + $v_{\mu}$ -bar (LAr)	2000 K	5250 K	270 K	670 K
Appear. test point	590	1900	360	910
$v_{\mu}$ (LAr+NESSiE)	230 K	1200 K	21 K	110 K
$v_{\mu}$ (NESSiE)	1150 K	3600 K	94 K	280 K
ν <sub>μ</sub> -bar (Lar+NESSiE)	370 K	56 K	33 K	6.9 K
$v_{\mu}$ -bar (NESSiE)	1100 K	300 K	89 K	22 K
Disappear. test point	1800	4700	1700	5000

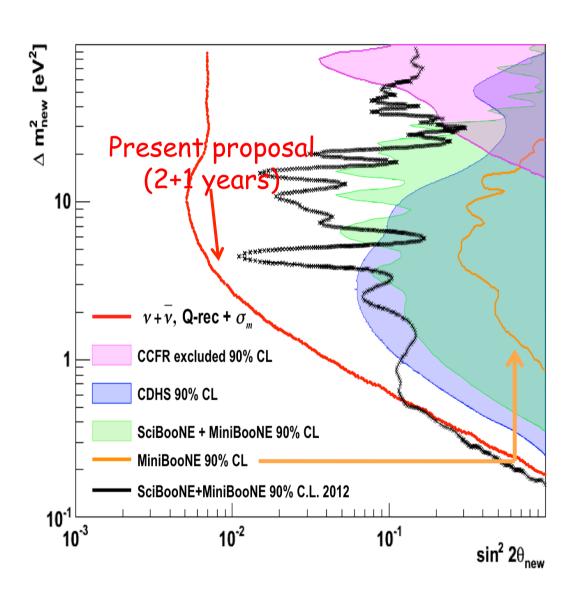
NOTE: v "contamination" in anti-v negative polarity beam

detected

 $\triangleright$  Values for  $\Delta m^2$  (sterile model) around 2 eV<sup>2</sup> are reported as example

## Sensitivity to $v_{\mu}$ disappearance

90% C.L. sensitivity for 2 years anti-v + 1 year v Exclusion limits: CCFR, CDHS, SciBooNE + MiniBooNE

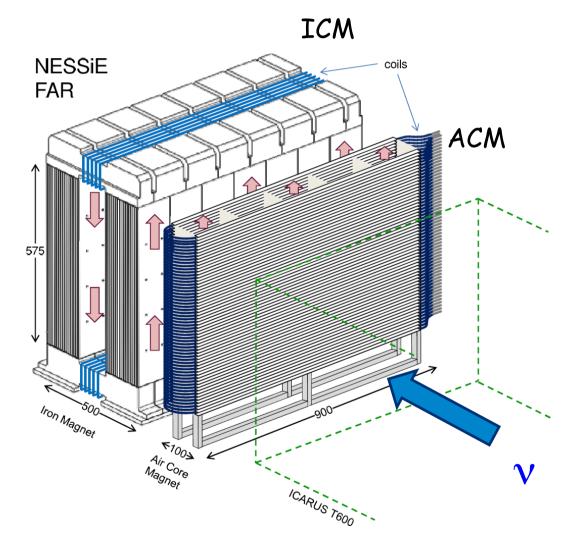


(only LAR+NESSiE events)

## Nessie (initial design)

- Two Iron spectrometers (ICM), 1500 + 800 t, composed by:
- 48 yoke blocks, , 4.5 x 0.6 x 1 m, 25t
- 480 slabs, 2 3 t, 1.25m x 3.5–6 m
- 1800 + 700 m<sup>2</sup> of RPC
- «sandwich style» assembly to be made in situ, one piece per time
- 20,000+12,000 digital channels
- Two ACM preassembled and installed in one shot
- Precision Trackers preassembled and installed in one shot
- Near Nessie movable aside on air-pad

 1 + 0.5 MW , 10 kA, power (summed up for ACM and ICM)



Optimized! Reduced by almost a factor of 2!!!

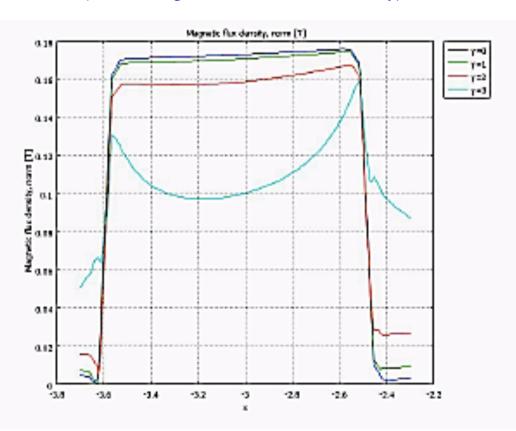
## Air-core magnet (initial design)

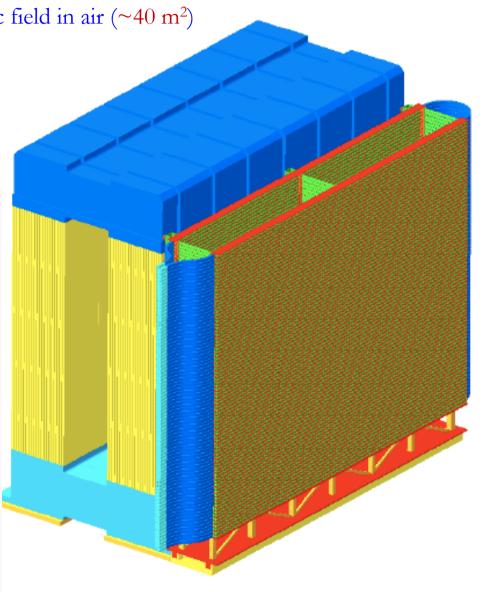
• New concept for a large transverse area magnetic field in air (~40 m²)

 $\bullet$  B = 0.17 T

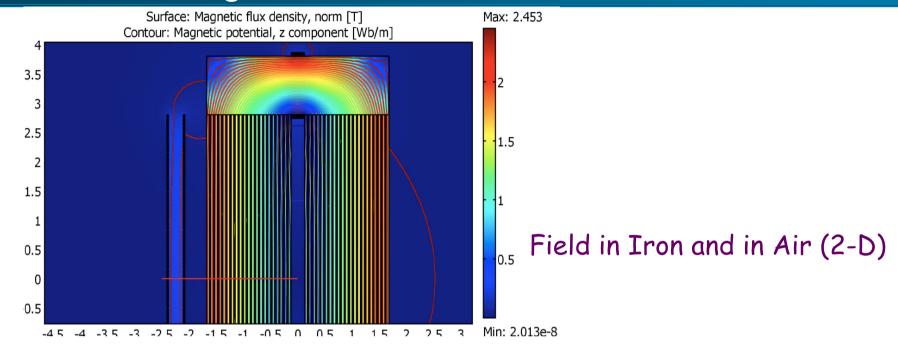
• Power < 2 MW

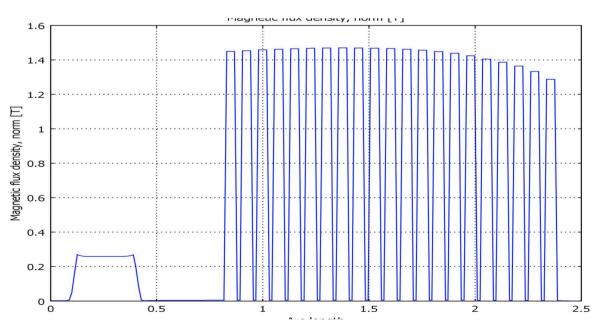
 To be coupled to a mm resolution detector (different possibilities under study)



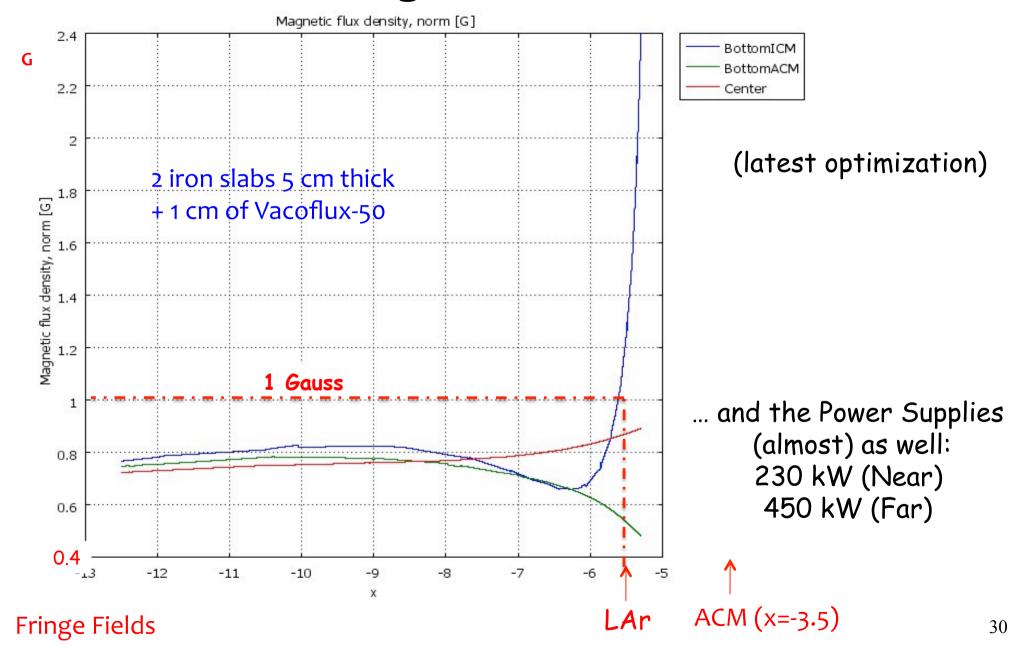


## Magnetic Field Simulation

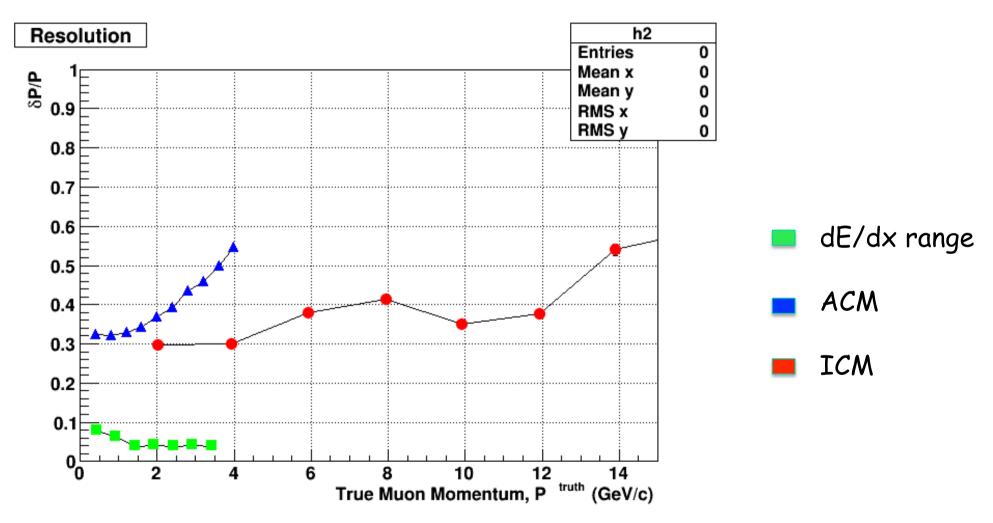




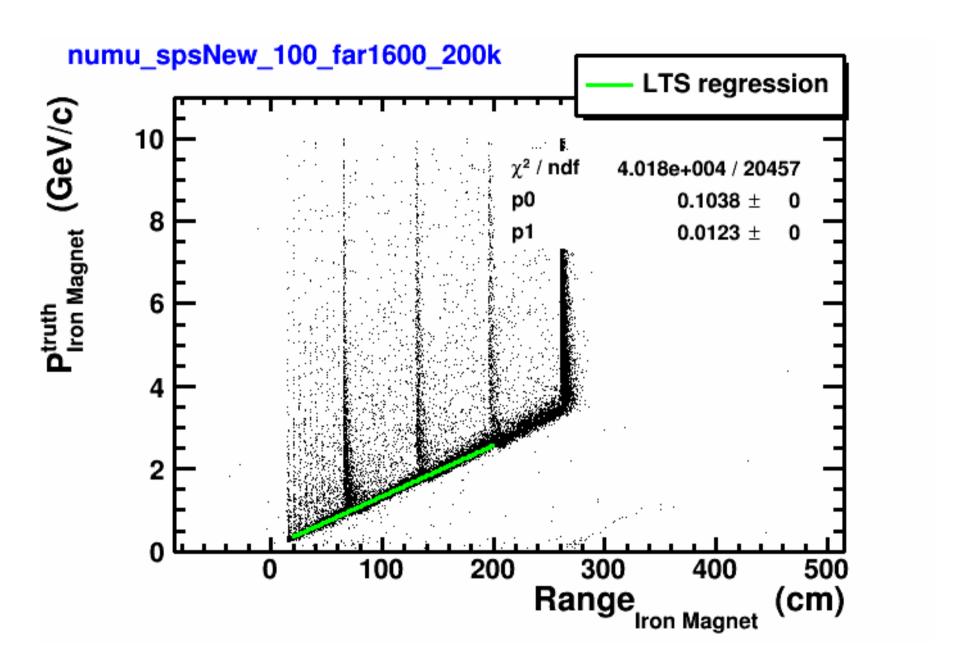
## ... and the fringe field is under control



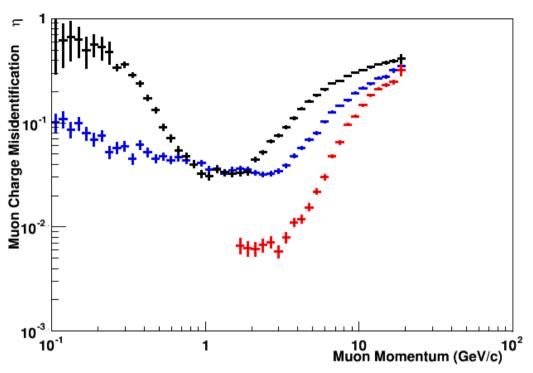
Momentum measurement at 4% with dE/dx up to 3.5 GeV At 30% with Prec.Tracker above 3 GeV At 30% with ACM below 1.5 GeV



Latest configuration (OPERA-2), algorithms not optimized ...



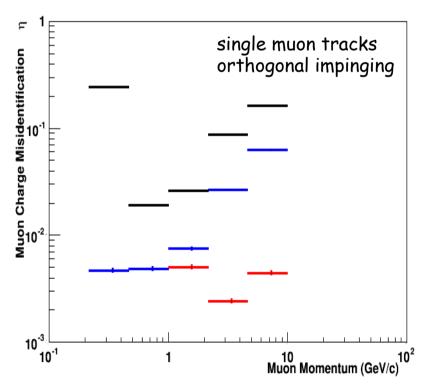
v interaction in the LAr, measured by dE/dx-range in the Spectrometer



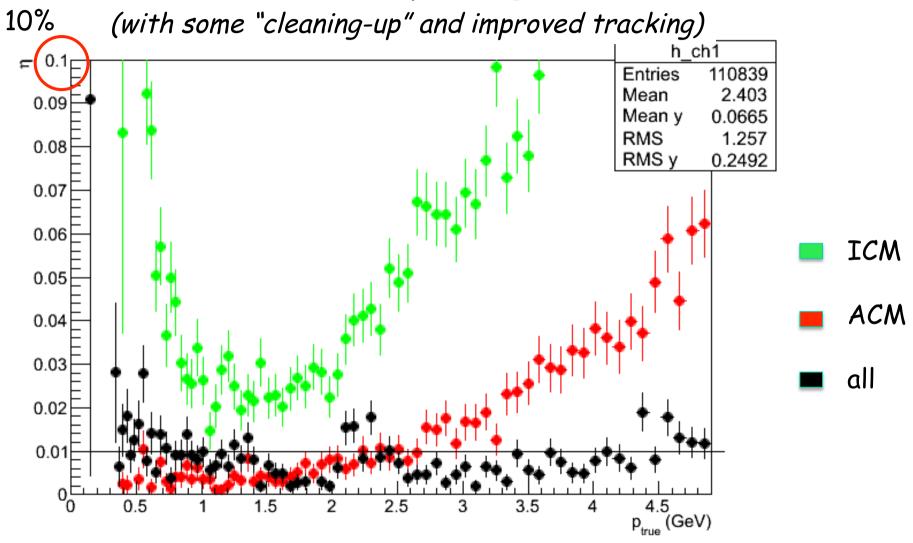
Charge mis-reconstruction at few percent level in all dynamical range of interest (full simulation including selection, efficiency and reconstruction)

## Muon mis-ID





## Muon mis-ID



(no way without ACM)

## Opera re-use

- 2 Spectrometers "available", with Detectors and Servicing
- Possibility to full re-use for Far AND Near ICM
- Need two new sets of Yokes (Top & Bottom)

new Electronics for RPC

Mechanical Tools

PT detectors

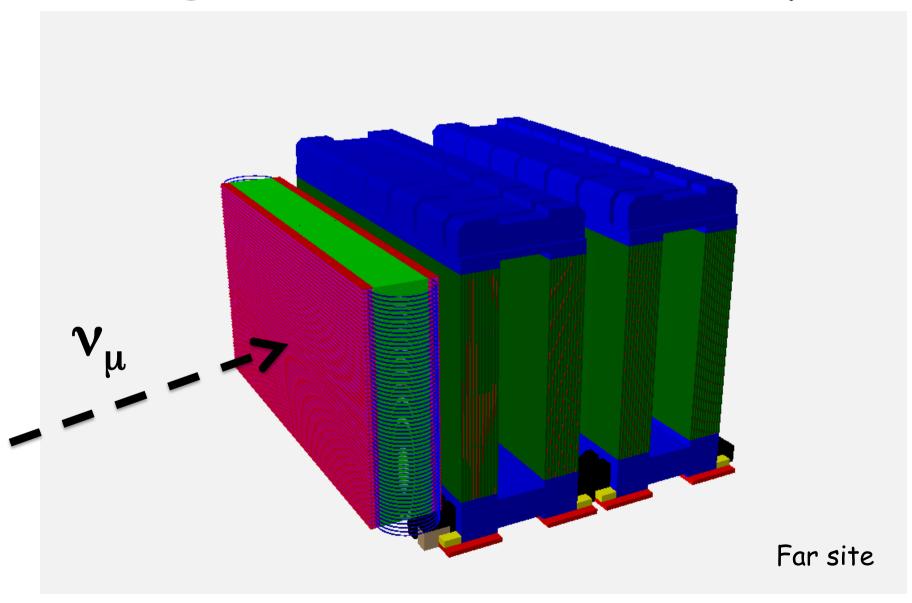
Scintillators

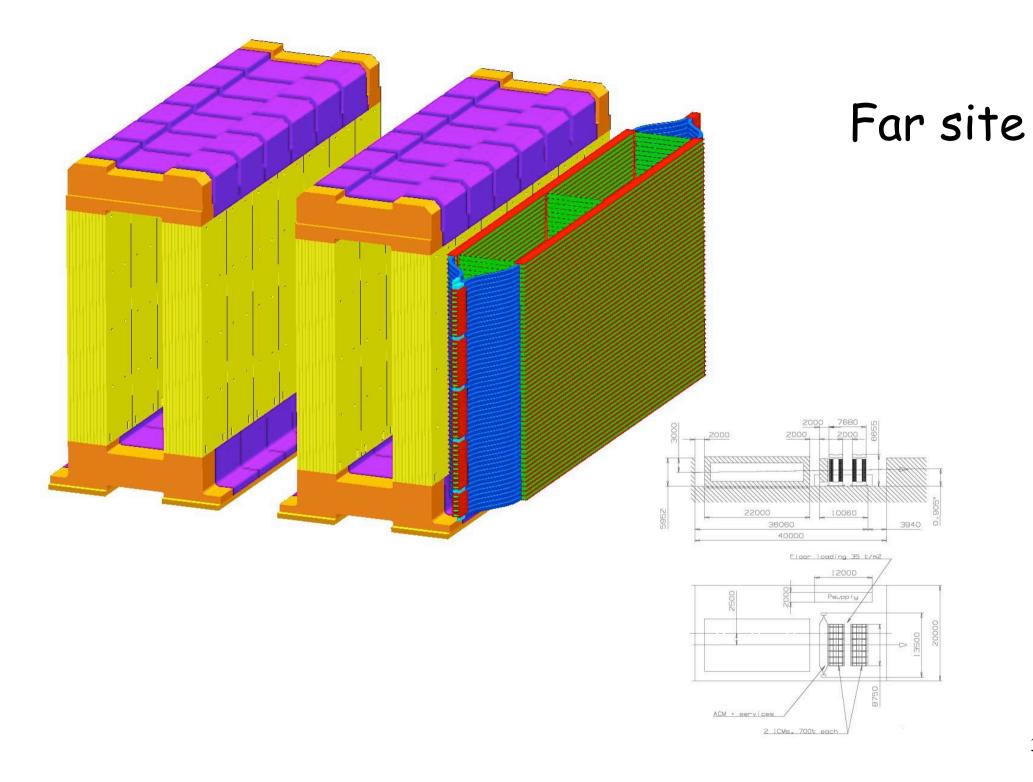
Other: ACMs

TO availability for dismantling and transportation at CERN: Autumn 2014

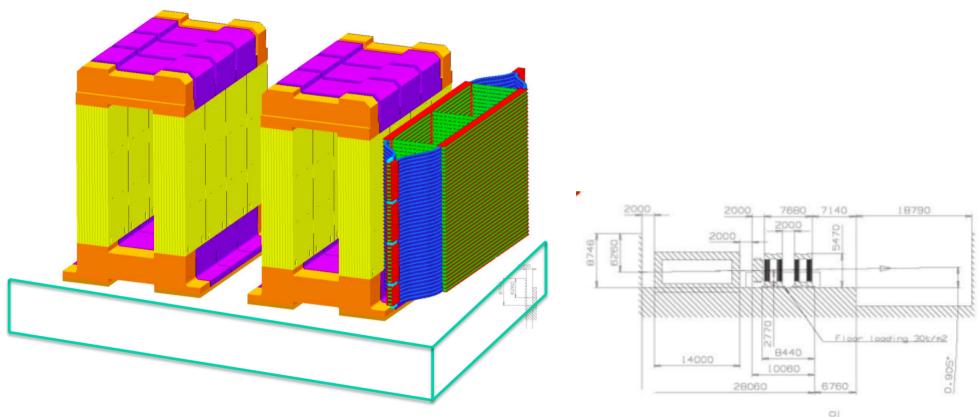
OPERA discussion: possible start dismantling July-December 2014

## Arrangement with OPERA Spectrs

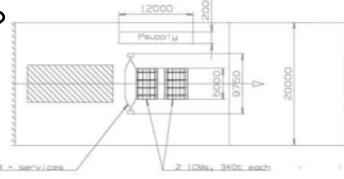


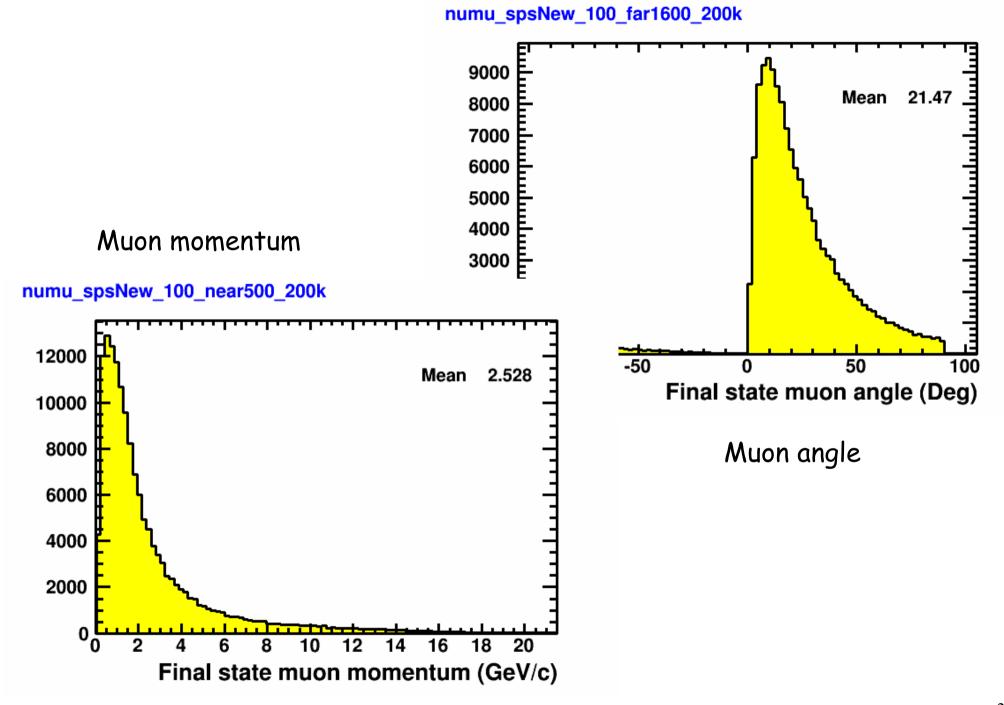


### Near site

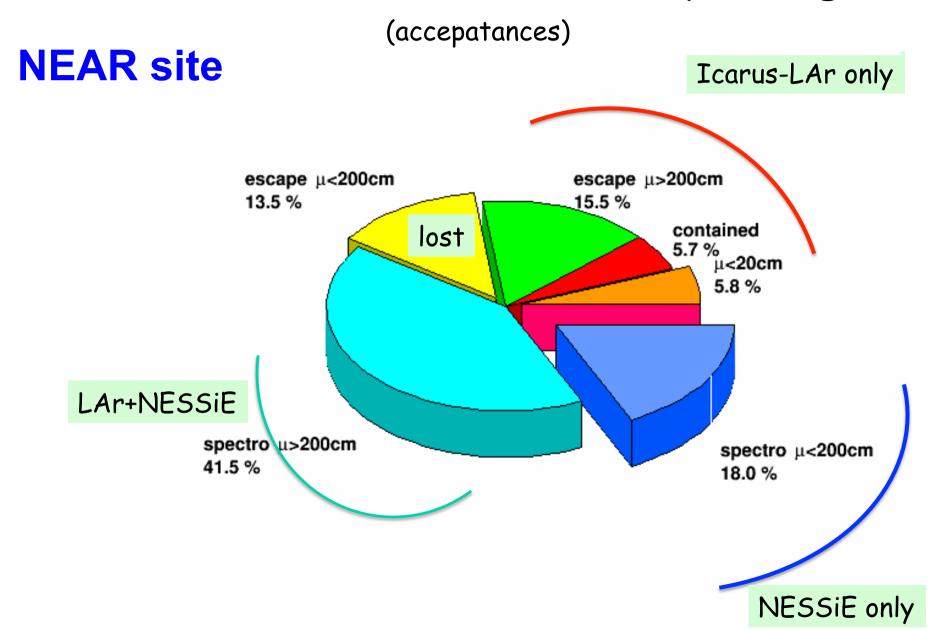


movable platform with air-pads à la ATLAS/CMS?



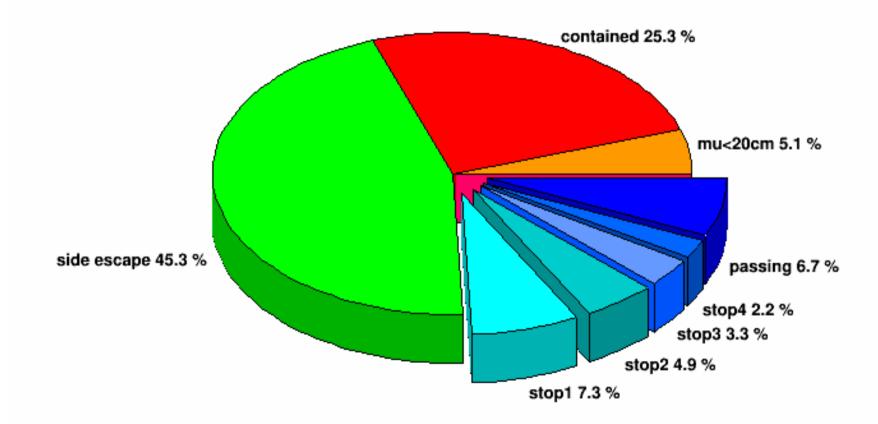


## Neutrino Interactions in the Liquid Argon



## muon stopping in the iron spectrometer

#### **FAR** site



#### The OPERA Magnetic Spectrometer

M. Ambrosio, R. Brugnera, S. Dusini, B. Dulach, C. Fanin, G. Felici, F. Dal Corso, A. Garfagnini, F. Grianti, C. Gustavino, P. Monacelli, A. Paoloni, L. Stanco, M. Spinetti, F. Terranova, and L. Votano

Abstract—The OPERA neutrino oscillation experiment foresees the construction of two magnetized iron spectrometers located after the lead-nuclear emulsion targets. The magnet is made up of two vertical walls of rectangular cross section connected by return yokes. The particle trajectories are measured by high precision drift tubes located before and after the arms of the magnet. Moreover, the magnet steel is instrumented with resistive plate chambers that ease pattern recognition and allow a calorimetric measurement of the hadronic showers. In this paper, we review the construction of the spectrometers. In particular, we describe the results obtained from the magnet and RPC prototypes and the installation of the final apparatus at the Gran Sasso laboratories. We discuss the mechanical and magnetic properties of the steel and the techniques employed to calibrate the field in the bulk of the magnet. Moreover, results of the tests and issues concerning the mass production of the resistive plate chambers are reported. Finally, the expected physics performance of the detector is described: estimates rely on numerical simulations and the outcome of the tests described above.

Index Terms—Gas detectors, neutrinos, particle measurements, particle tracking.

#### I. INTRODUCTION

PERA is a long-baseline neutrino experiment currently under construction at the Gran Sasso underground laboratories (LNGS) [1]. Its aim is the observation of  $\nu_{\mu} \rightarrow \nu_{\tau}$  oscillations in the parameter region indicated by Super-Kamiokande [2] through direct observation of  $\nu_{\tau}$  charged current interactions. The detector design is based on a massive lead/nuclear emulsion target (ECC) complemented by electronic detectors (scintillator bars) that allow the location of the event and drive the scanning of the emulsions. This

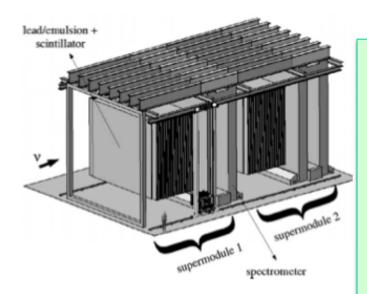
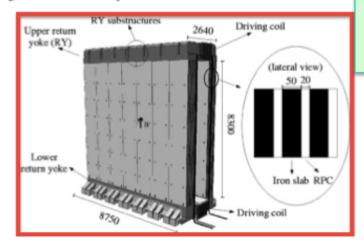


Fig. 1. The OPERA experiment at LNGS.



TWO spectrs with: 12+12 Iron walls (instead of 21+21)

7 rows 3-RPC (instead of 5)

→ use 4 rows for FAR 3 rows for NEAR

7 vertical slabs:

- → ok for FAR
- → use 4 for NEAR

Need: new NEAR Yokes

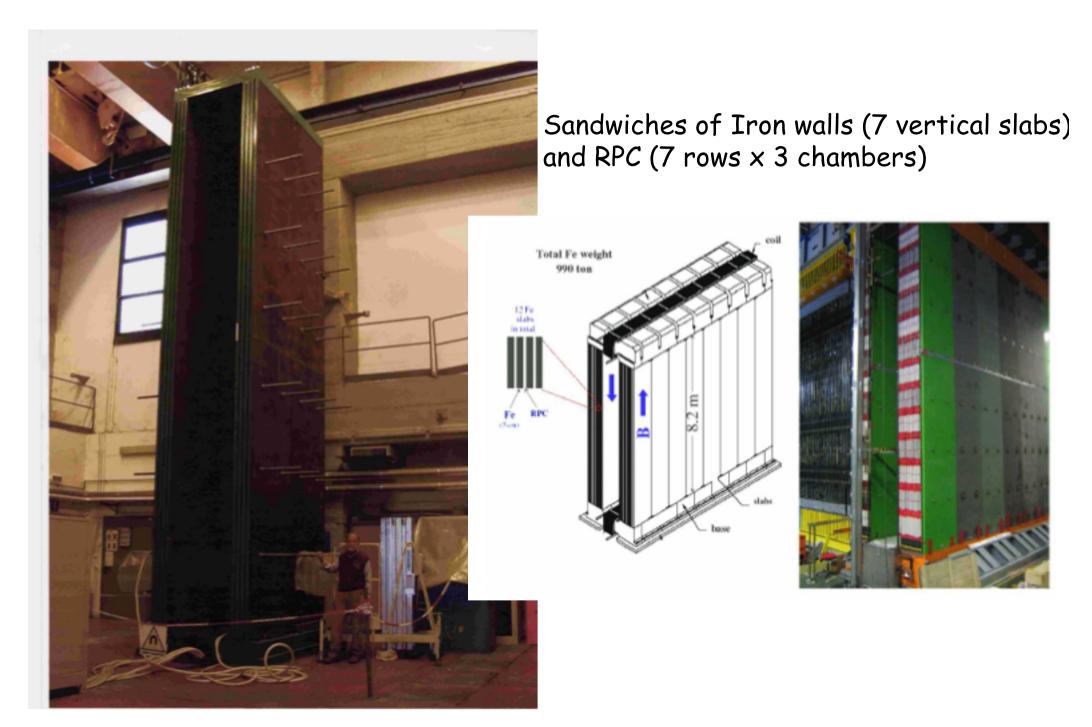
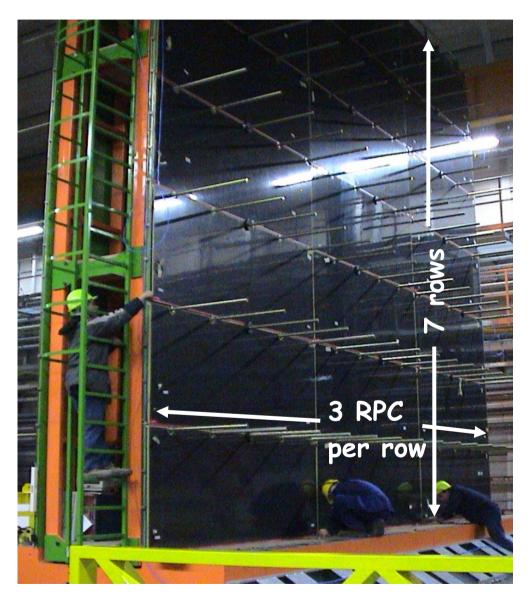
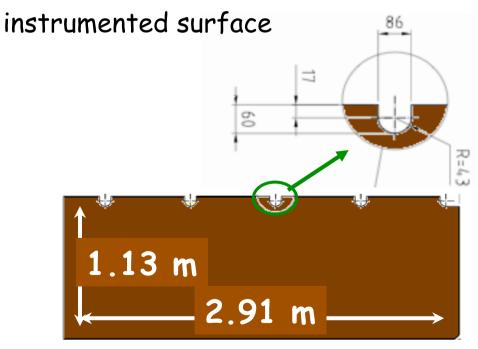


Figure 1: The OPERA dipolar magnet prototype.

#### ■ RPC mechanical structure



- 3 x 7 RPC per plane
- $\cdot \sim 70 \text{ m}^2$  instrumented surface per plane
- plane geometrical acceptance ~96%
- 22 + 2 RPC planes per spectrometer
- 1008 RPC for a total of  $\sim$  3200  $m^2$



## RPC strips

Vertical strips (2.6 cm pitch) and Horizontal strips (3.5 cm pitch)

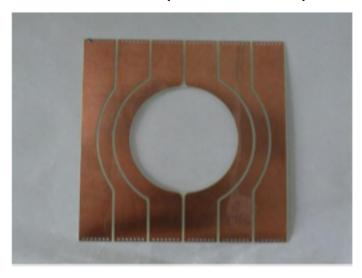
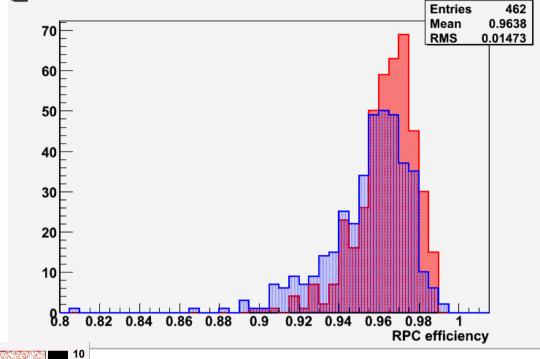
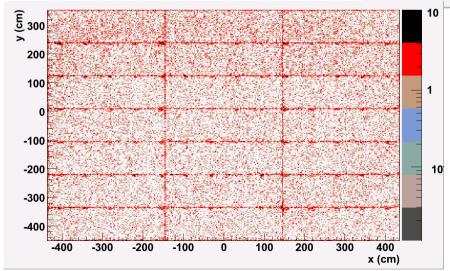


Fig. 6. Readout strips near the bolt.



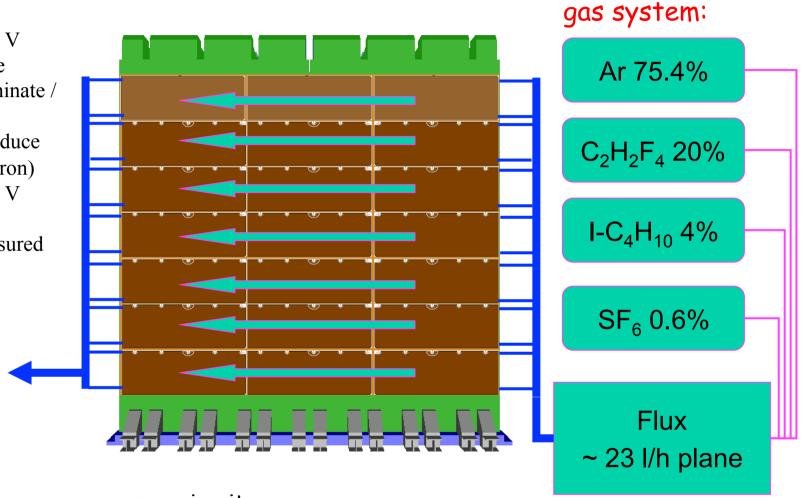


inefficient points (summed over all the planes)

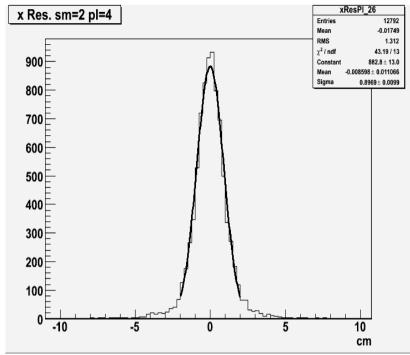
#### ■ RPC HV and gas system



- running at about 5800 V
- 2 mm gap between the electrodes (plastic laminate / "bakelite")
- opposite polarity to reduce discharge to ground (iron)
- running at about 5800 V
- 144 HV channels
- RPC row current measured with nA resolution

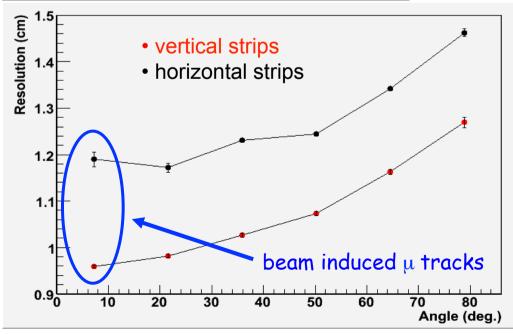


- · open circuit
- RPC rows flushed in parallel with needle at the input to equalize the flux and bubbler on the output line
- · in total ~ 1 m<sup>3</sup>/h



#### ■ RPC tracking resolution

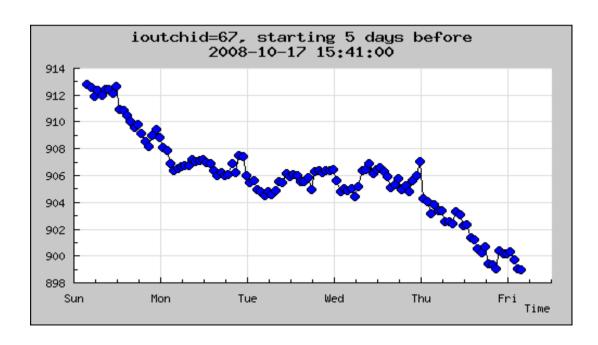
- tracks residuals for a typical plane
- mean well centered at 0 cm hence alignment is fine
- $\cdot$  RMS = 1.3 cm
- $\sigma = 0.9 \text{ cm}$

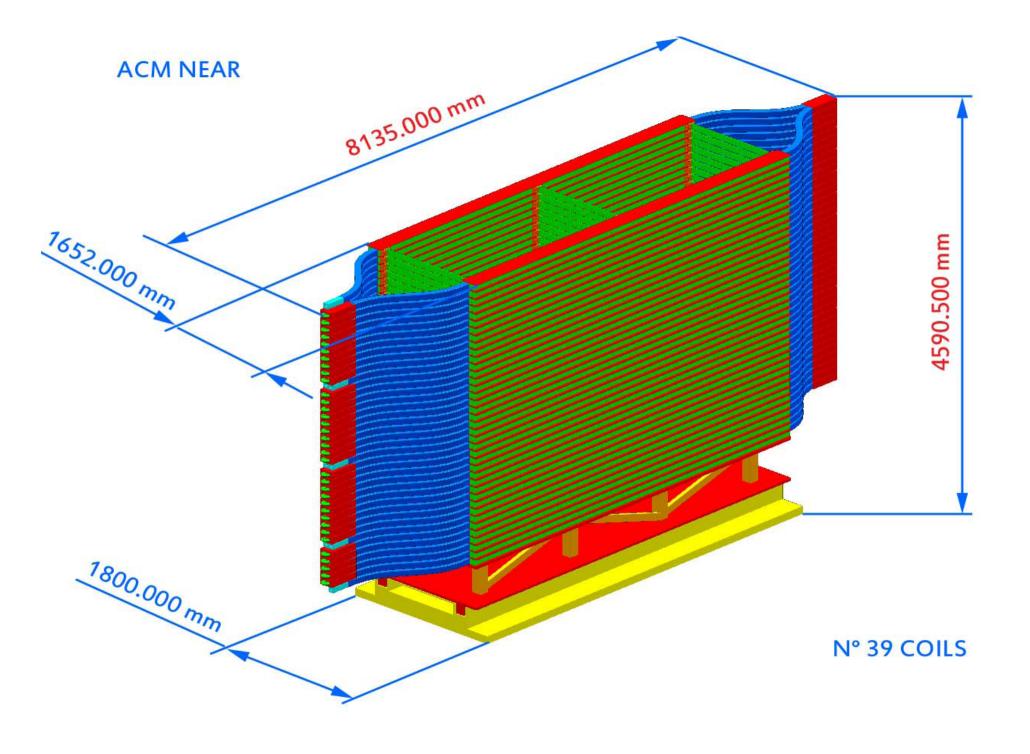


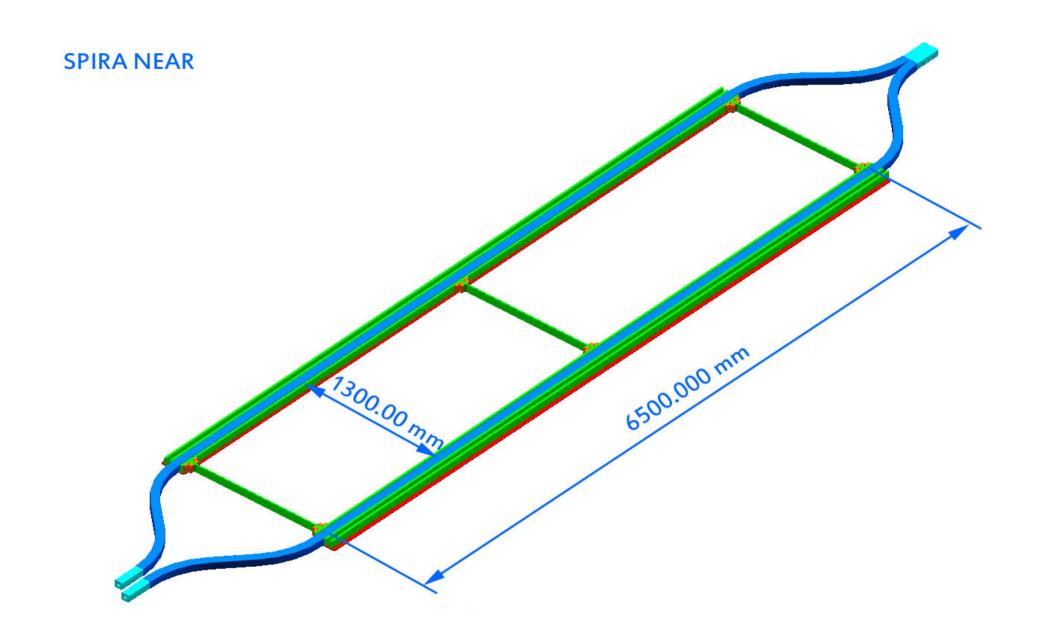
- resolution averaged over all 44 RPC planes
- higher resolution of the vertical view (lower strips size)
- better resolution of orthogonal tracks

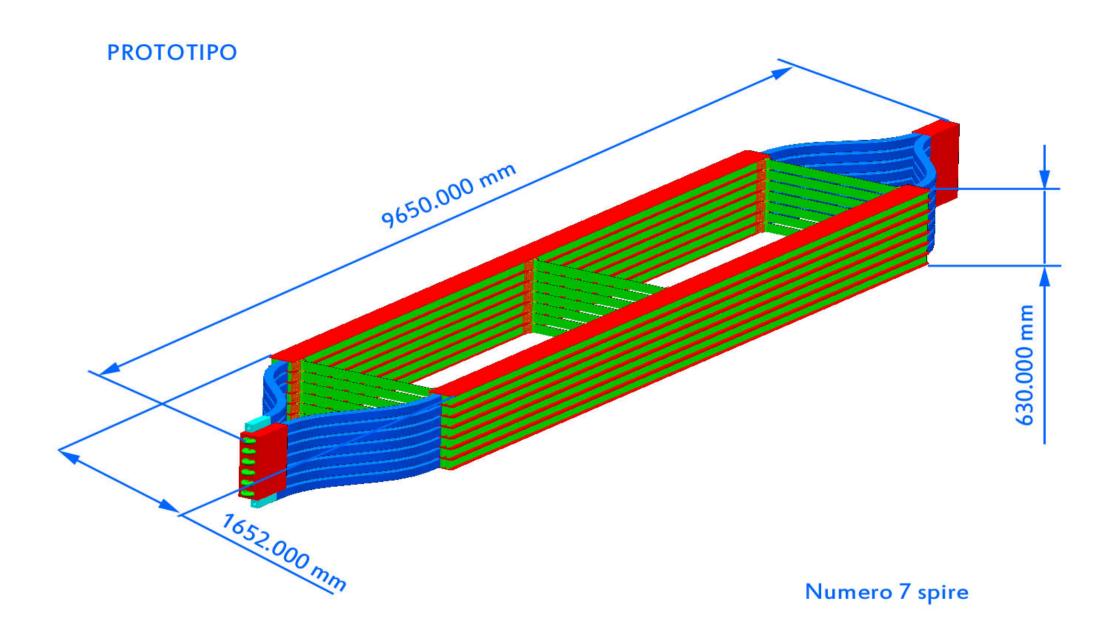
#### ■ RPC monitoring

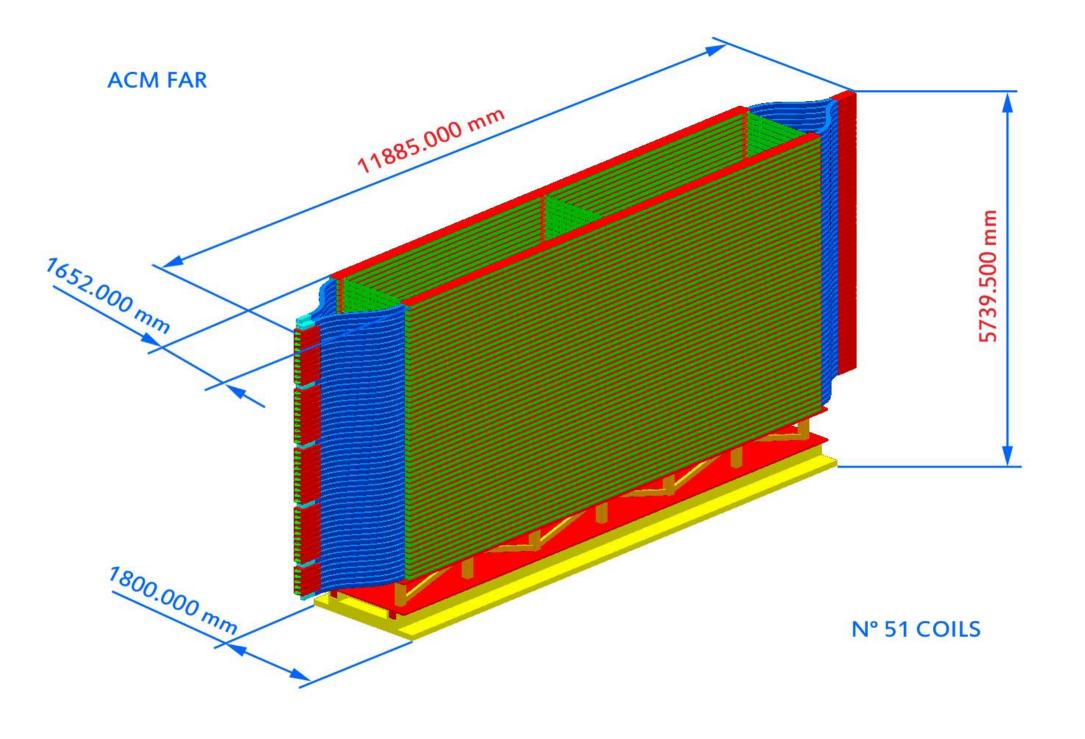
- operating pressure and temperatures recorded and shown via a web interface, HV adjusted accordingly
- · HV and currents, at the single channel level, recorded and shown via web interface
- running conditions monitored via web interface
- running conditions can be monitored / changed via vcn-viewers
- alarm conditions on HV / currents / running conditions broadcasted to the relevant experts automatically via e-mail and sms messages



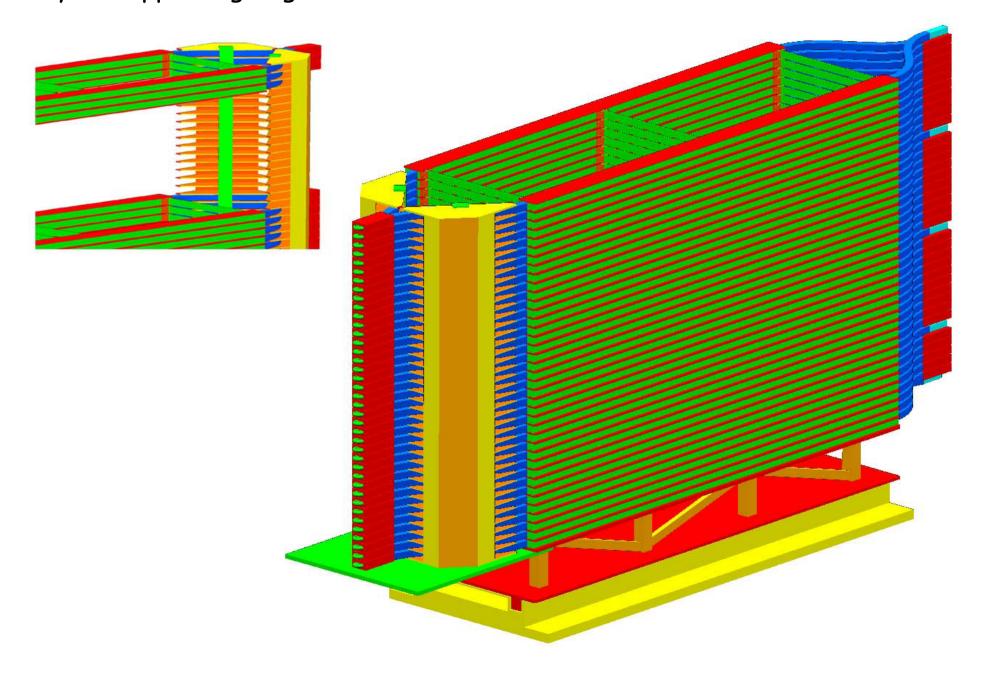






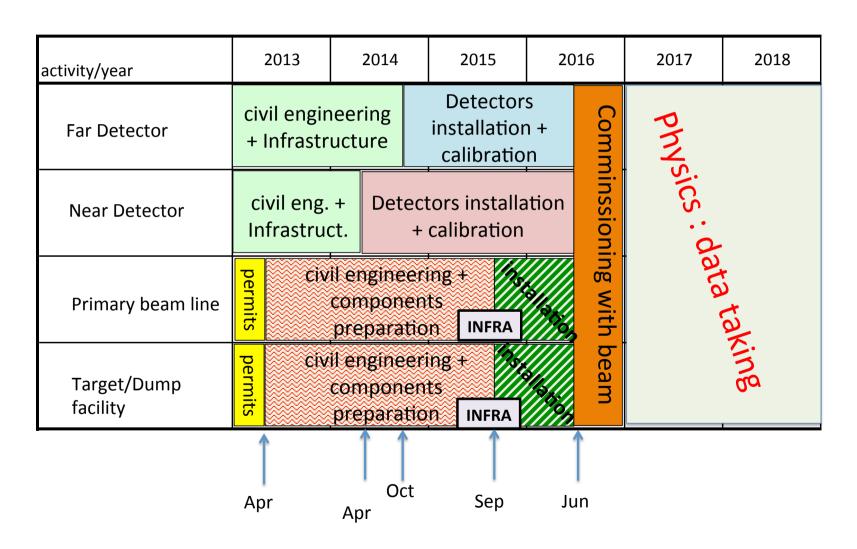


Sudy of supporting edges



#### **CERN Schedule**

#### Overall Planning Proposal (nord area)



#### Nessie Schedule

2013: ACM prototype construction,

iron SM, ACM, HPT and ancillaries design

2013 - end: start tenders process

2014 - early: issue tenders for production/modification

(copper. yokes, slabs, rpc, strips, tools/frames)

2014: parts and tools production/modification

2014 - end: first deliveries at Cern

(copper coils, yokes, slabs, strips, tools/frames, RPC)

2015 - early: start assembly

2016 - mid: finish assembly

2016 - fall: commissioning and start of data taking

operation	work days	work days	weeks	crane 5 t	crane 40 t	crane	
	min	max	(average)	occupancy	occupancy	drivers	
Measuring , tracking, drilling , stud bolts fixing with resins	15	20	4			0	
Bottom Copper coil, lower bars + thermal shield installation	5	5	1	50		o	
internal formwork - preparation	1	2	0	50		0	
Bottom Jokes installation	10	15	3		100	3	
jokes delivery and lowering inside pit							
Jokes preparation							
Jokes mounting on the stud bolts							
Checks, leveling, bolt tightning							
external formwork - preparation	1	2	0	20		0	
self-leveling concrete deposal	1	2	0	20		0	
#3 provisional carpentries delivery, installation, test	5	10	2	20	50	11	
personnel platforms delivery,installation, test:	5	10	2	20	50	11	
Mounting of studs on central support carpentry	1	2	0	20		0	
SLABS/STRIPS/RPCs sandwich installation	100	200	30	100	20	26	
2? Slabs Stacks (14 slabs, 40t) delivery + ancillaries							
First slab layers Installation, checks, fixing							
Delivery of RPCs, strip planes, foam lining							
Slabs, rpc strips preparation							
20+20 Slabs and RPC +strip sandwich installation							
Upper Copper coil, lower bars delivery, installation	3	5	1	50		0	
Top yokes delivery, preparation, installation, fixing	15	20	4		100	4	
jokes delivery and lowering inside pit							
Jokes preparation							
Jokes mounting							
Checks, leveling, bolt tightning		1		i			
Upper Copper coil, upper bars delivery, installation	5	15	2	50		0	
Structural checks	1	1	0			0	
Provisional carpentries removal, extracting aside	5	5	1		100	1	
Structural test and checks	1	1	0			0	
Bottom Copper coil, upper bars + thermal screen delivery,						1.5	
installation	20	20	4	50		0	
tools dismantling	5	10	2	50		0	
mount personnel scaffolding aside/inside	5	10	2	50		0	
ICM coils 50 kW cooling and power connections, test	10	15	3			0	
ICM FINISHED	214	370	58			-	
HPT detector intermediate position delivery, insertion,	10	20	3	50		o	
ACM coil delivery, insertion, fixing, connection, checks, tests	5	20	3		50	1	unless mobil
HPT detector for ACM delivery, insertion, fixing, checks, test?	5	20	3	50		0	
(Mounting of platform on top of Nessie)	5	15	2		50	1	
Detectors PS and electronics installation	15	20	4	50	5244000 V	0	
Detectors cabling, piping, test	15	25	4			0	
ACM and SERVICES	55	120	18				
OVERALL	54	98	76			57	

#### Status of approval and further steps

#### At CERN

- Group established by CERN in order to realize SPS based new short-baseline  $\nu$  beam in the North Area (project leader M. Nessi)
- Scientific Approval (SPSC) middle of January 2013
- Feasibility document submitted to CERN Directorate on February 7<sup>th</sup> 2013
- Research Board evaluation on March 4<sup>th</sup> 2013
- SPC
- CERN Council

#### INFN: Currently Major contributor to the experiments (not beam)

- Scientific approval
- Under evaluation by the Technical Scientific Committee (CTS) as for costs, manpower
- In-Kind contribution of Opera Spectrometers

#### MONEY estimation

Iron magnets: in-kind value 5940 K€ (from OPERA MoU)

Cost for transportation to CERN and refurbishing: 3000 K€

In-kind value of Precision Tracker: 1900 K€

possible refurbishing: 700 K€

In-kind value of Scintillators: 1900 K€

possible refurbishing: 300 K€

Cost ACM: 700 (Near) + 1800 (Far)

TOTAL: 3+1+1+2 = 7 M€

might be staged at 2<sup>nd</sup> phase (after LS2)

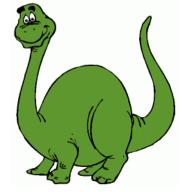
#### **Conclusions**

Possibility of exciting discoveries of BSM Physics with vast consequences or a complete clarification of present anomalies.

Favorable time scale thanks to the use of existing/running detectors (or reasonable extensions).

Opportunity for a revival of neutrino activity in Europe. Possible synergies with the other 3v and R&D programs.

Large room and availability for contributions in NESSiE



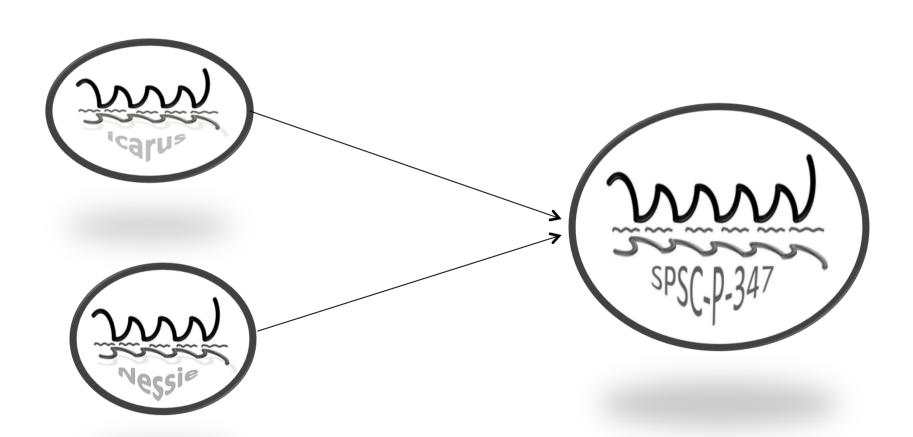


## Summary

- To make it, one has to start now!
- The real heavy load is on the detectors, which need to move very fast
- This effort will provide CERN with a new neutrino facility for the longer-term future
- .... and will represent a strong basis for the European neutrino community towards the future (i.e. long baseline in or outside Europe)

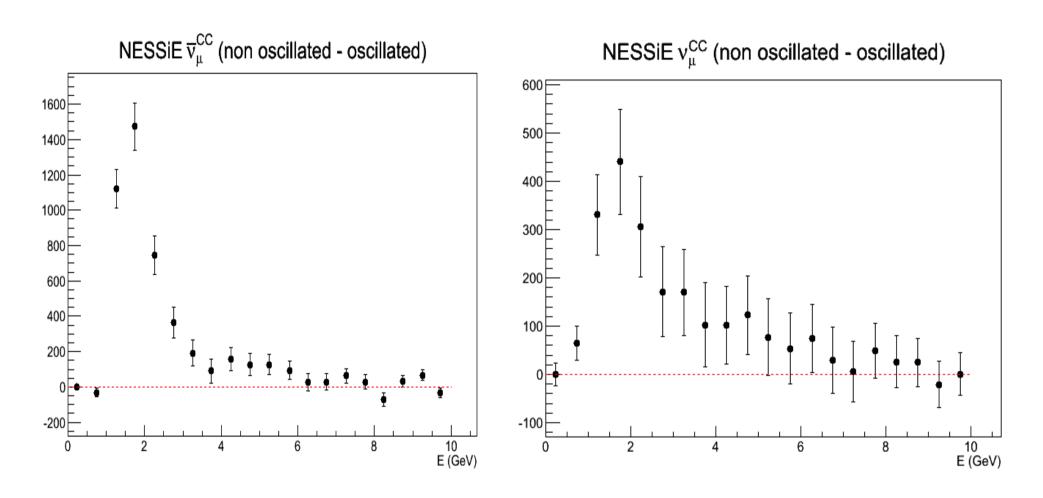
Let's not loose this momentum !!!

# Thank you!

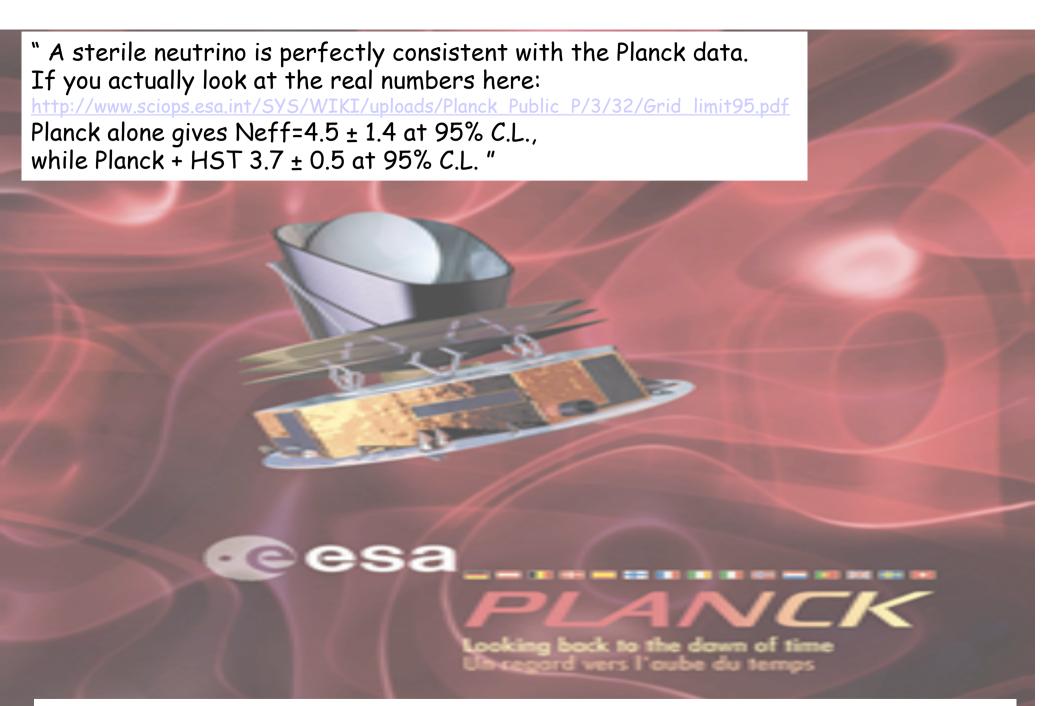


Backup slides

#### Muon disappearance signal

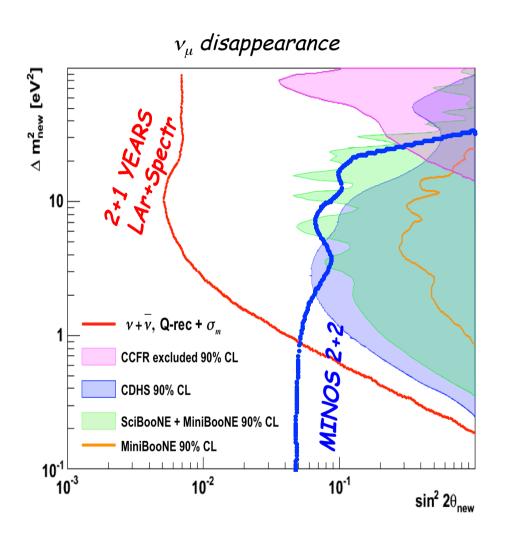


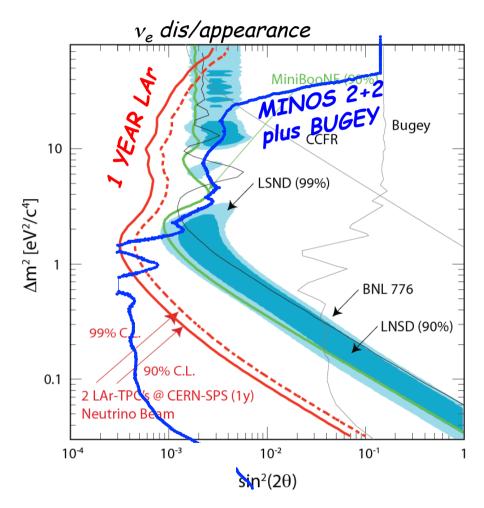
1 year of anti-neutrinos!

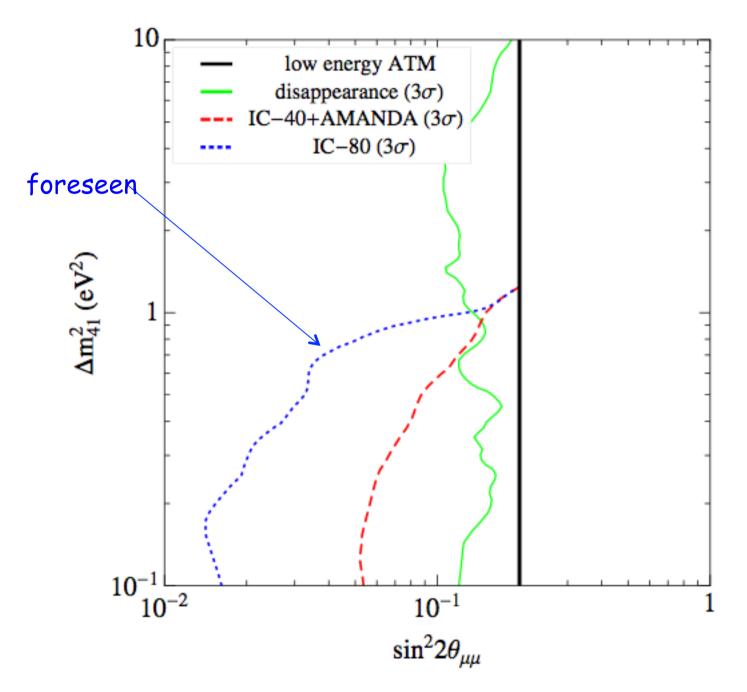


"PLANCK preferred to be conservative and put 3.3  $\pm$  0.3 at 68% C.L. in the Abstract"

### Comparison with MINOS+ limits







JCAP11(2012)041

(not sign by the whole collaboration)

#### Nessie FAR vs Icarus acceptance

